

GrAHal-CAPP for Axion Dark Matter Search with Unprecedented Sensitivity in the 1-3  $\mu eV$  Mass Range

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Grenoble Axion Haloscopes LNCMI

# Outline

- Introduction
  - Weakly Interacting Slim Particles (WISPs)\* as a possible component(s) of the Cold Dark Matter (CDM)
  - & QCD DM Axion
  - The need of high field / high flux magnets

\* Complementary to the "better-known" WIMPs, i.e. Weakly Interactive Massive Particles

- The 43 T Grenoble Hybrid Magnet at LNCMI (CNRS, Univ. Grenoble-Alpes)
  - Overview
  - Development of the superconducting conductor & in-house assembly
  - On-going commissioning phase
- The Grenoble Axion Haloscope (GrAHal)
  - BabyGrAHal
  - Probing QCD Axion Dark Matter with the Grenoble Hybrid Magnet up to 43 T
  - QCD DM axion search at high frequency & high field, i.e. around 12.78 GHz (52.5  $\mu eV)$  & 43 T
  - Focus on QCD DM axion search in the low frequency range (200-600 MHz *i.e.* 1-3  $\mu$ eV)

### Particle Physics Beyond the Standard Model – Oversimplified Picture



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### Why High Magnetic Fields & Flux for QCD-DM Axion/ALPs searches?

To maximize the conversion of this hypothetical weakly interacting particle to photons, via the inverse Primakoff effect



The key ingredient of most of the experiments  $P_{LSW} \sim g_{a\gamma\gamma}^{4} B^{4} L^{4}$  $P_{Haloscope} \sim g_{a\gamma\gamma}^{2} B^{2} V$ 

This "non-trivial" interaction is related to the chiral anomaly, i.e. a purely quantum phenomenon first studied in particle physics in 1969 (Adler, Bell and Jackiw) to explain the neutral pion decay in 2 photons (  $\pi^0 \rightarrow \gamma\gamma$  ) anticipated and observed by Primakoff in 1951.

The puzzle was the anomalous nonconservation of a chiral current, which is today "rejuvenated" in condensed matter physics...

## Hybrid Magnets Worldwide Producing the highest DC-field











Specially developed RCOCC conductor 13 x 18 mm<sup>2</sup>

Key point Modularity High field/high flux

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Cooling by forced flow supercritical He ~11g/s @ 5 bars



FIELD LABORATORY

Grenoble, France	Nijmegen, Netherlands	Hefei, China	Tallahassee, FL
8.5 (9) + 34.5 (36+) = <b>43 (45+) T</b>	12 + 33 = <b>45 T</b>	11 + 34 = <b>45 T</b>	11.5 + 33.5 = <b>45 T</b>
34 mm, 24 (30) MW	32 mm, 24 MW	32 mm, 32 MW	32 mm, 32 MW
RCOCC Nb-Ti, 1.8 K	CICC Nb3-Sn, 4.2 K	CICC Nb3-Sn, 4.2 K	CICC Nb3-Sn, 4.2 K
7.1 kA, 1100/1826 mm dia.	20 kA, 720/1286 mm dia.	13.4 kA, 680/1650 mm dia.	10 kA
2025	In construction	45.22 T Aug. 12, 2022	45.17 T June 26, 2000

### Technological Choices

# The cryogenic system principle: Pressurized superfluid He bath at 1200 hPa, 1.8 K



#### Industrial Production Line Developed, Built, Installed & Operated at LNCMI



Innovative developments have been achieved based on induction heating to strictly control  $R_a$  & AC losses

P. Pugnat, R. Pfister *et al., IEEE Trans. Appl. Supercond.* **26**, 4302405 (2016)





Crimping, soft-soldering, calibration & winding in single pancakes 4 m hight for delivery to the magnet manufacturer https://www.youtube.com/watch?v=cp5NIR2cN5s

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### Integration of the Grenoble Hybrid Magnet with its Cryogenic Plant



# Magnet Seismicity\* during Powering up to the Nominal Field of 8.5 T <u>IEEE TASC 34, 1 (2023)</u>

\*Pioneering work in 2001: DOI 10.1109/77.920111



No training quenches but Kaiser effect observed

Most of the spikes are located at the interface between flanges and first & last DP

Sc magnet fully operational at 8.5 T GrAllal

### **Grenoble Axion Haloscopes**



Grenoble Alpes

agence nationale de la recherche

an







Théorie

R. Ballou P. Camus T. Grenet

P. Perrier

- A. Talarmin
- J. Vessaire

P. Pugnat R. Pfister S. Krämer J. Quevillon C. Smith K. Martineau A. Barrau 

# Few Words from P. Sikivie (Haloscopes proposed in 1983, Rev. Mod. Phys. 93, 015004)



Visit of Olympie during 2<sup>nd</sup> Patras Workshop in 2006 at Patras



**Axion electrodynamics** 

 $\nabla \cdot \mathbf{E} = g_{a\gamma\gamma} \mathbf{B} \cdot \nabla a$  $\nabla \times \mathbf{B} - \partial_t \mathbf{E} = g_{a\gamma\gamma} \left( \mathbf{E} \times \nabla a - \mathbf{B} \partial_t a \right)$  $\nabla \times \mathbf{E} + \partial_t \mathbf{B} = 0$  $\nabla \cdot \mathbf{B} = 0$ 



"Most importantly, the cavity experiment uses a variety of technologies microwave engineering, ultra-low noise receivers in a high magnetic field environment, cryogenics - which are not typically used by high energy physicists and which had to be specially developed.

... Feynman's advice to young scientists aspiring to great discoveries. He said: "You have to develop your own tools". "

https://ep-news.web.cern.ch/content/qa-pierre-sikivie



Grenoble Axion Haloscopes

#### Key expertise at CNRS-Grenoble for High magnetic fields, Extreme Low Temperatures & Quantum Detectors





European Magnetic Field Laboratory LN

Dresden/LNCMI-Toulouse, pulsed up to 95/91 T, 1-10 ms Nijmegen/**LNCMI-Grenoble**, DC up to 38/36 T, Projects 45/43+ T

https://emfl-users.lncmi.cnrs.fr/SelCom/proposals.shtml



**Grenoble Alpes** ÉEL **European Microkelvin Platform** JPA Achievements 20 leading ultralow temperature physics & technology Institutes in Europe including 7 submilliK facilities https://www.cnrs.fr/cnrsinnovationlalettre/actus.php?numero=743  $1 \text{ GHz} < f_o < 10 \text{ GHz}$ http://emplatform.eu/about/facilities  $G \ge 20 \text{ dB}$  $BW \sim 2 \text{ GHz}$ 10 µK nuclear stage  $T_N \gtrsim \frac{hf_o}{2k_{\rm B}}$ temperature 6 μK nuclear stage  $P_{\rm 1dB} \sim -100 \ {\rm dBm}$ 600 nm temperature Quantum limited Josephson parametric amplifiers Nicolas Roch QuantECA Team nstitut Néel, Grenoble, 100 µK nuclear France demagnetization stage **Expertise for dilution fridges & cryostats** 

(Planck, Edelweiss, CUT, SuperCDMS ...)

# Baby-GrAHal 1: 1<sup>st</sup> Experimental Runs Ended







The RF-cavity resonant frequency was tuned & scanned by varying the GHe pressure around the cavity :

- For the range 1-1200 mbar, excursion  $\Delta f$  = 20 MHz, i.e. ~ 0.1  $\mu$ eV
- Sensitivity in the range of 20-25 x KSVZ @ 4.4 K
- Detailed data analysis close to completion (to be published)





T. Grenet et al. https://arxiv.org/abs/2110.14406

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Grenoble Axion Haloscopes



► The key element : The modular Grenoble Hybrid Magnet combining sc and resistive technologies (ongoing commissioning up to 43 T)



Field	Warm dia.	Power	RF-cavity dia.	<b>f</b> <sub>тмо10</sub>	Axion mass	B <sup>2</sup> V (T <sup>2</sup> m <sup>3</sup> )
43 T	34 mm	25.4 MW	20 mm	11.5 GHz	47.2 μeV	0.5
40 T	50 mm	25.4 MW	34 mm	6.76 GHz	27.8 μeV	0.6
27 T	170 mm	19 MW	86 mm	2.67 GHz	11 µeV	3.5
17.5 T	375 mm	12.9 MW	291 mm	0.79 GHz	3.2 μeV	6.6
9 T	800 mm	0.4 MW	675 mm	0.34 GHz	1.4 μeV	40



Operation end of 2024-25 with HTS RF cavity in collaboration with CAPP/IBS-KAIST (see O. Kwon presentation)

# GrAHal-CAPP ► Focus on 1-3 µeV axion mass (200-600 MHz)





Grenoble Axion Haloscopes

Center for Axion and Precision Physics Research



#### GrAHal-CAPP : Phase 1 @ 4K

- 50 K cryo-stage operational
   @ t<sub>0</sub>+18 months
- 4 K cryo-stage operational
  @ t<sub>0</sub>+24 months

 $\rightarrow 1^{st} run$ 

<u>GrAHal-CAPP : Phase 2 @ 50 mK</u> - Operational @  $t_0$  + 42 months

 $\rightarrow 2^{nd}$  run reaching DFSZ, in 2-year integration time

# ► Focus first on 1-3 μeV axion mass (200-600 MHz)



https://doi.org/10.3389/fphy.2024.1358810

# More Information / Outline

#### Few references

- "High magnetic fields for fundamental physics": <u>https://arxiv.org/pdf/1803.07547.pdf</u>
- OSQAR: <u>https://ep-news.web.cern.ch/content/osqar-experiment-sheds-light-hidden-sector-cerns-</u> <u>scientific-heritage</u>, <u>https://arxiv.org/abs/1506.08082</u>
- GrAHal: <u>https://bib-pubdb1.desy.de/record/395493</u>; <u>https://arxiv.org/abs/2110.14406</u>; <u>https://www.frontiersin.org/journals/physics/articles/10.3389/fphy.2024.1358810/full</u>
- VMB@CERN: https://cds.cern.ch/record/2649744

#### CERN PBC Study Group defining the European strategy of Particle Physics

- https://pbc.web.cern.ch/
- <u>https://indico.stfc.ac.uk/event/268/attachments/522/909/Vallee\_PBC\_RAL.pdf</u>
- <u>https://www.nature.com/articles/s41567-020-0838-4</u>
- https://indico.cern.ch/event/1369776/contributions/5795144/attachments/2827635/

#### New EU COST Action : COSMIC WISPers in the Dark Universe: Theory, astrophysics and experiments

- <u>https://www.cost.eu/actions/CA21106/</u> (Chairman/Co-Chair, MoU, Objectives)
- You can apply to working groups of the network from <u>https://www.cost.eu/actions/CA21106/#tabs+Name:Working%20Groups%20and%20Membership</u>
- Kick-off Meeting at Rome 23-24 February 2023 https://agenda.infn.it/e/CosmicWispersKickOff

High Field Magnet Proposal submission open twice a year: <u>https://emfl.eu/apply-for-magnet-time/</u>



Grenoble Axion Haloscopes





