Fundamental Physics Search with SRF Cavities

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Electromagnetic Detection of Ultralight Bosons and HFGW

 \vec{J}_{eff} in a cavity or shield room: $\Box \vec{A} = \vec{J}_{\text{eff}}$.

Dark photon A': $\vec{J}_{eff} = \epsilon m_{A'}^2 \vec{A'};$

No background field. Background \vec{B}_0 .

Axion a: $\vec{J}_{eff} = g_{a\gamma} \vec{B}_0 \partial_t a;$

GW h $\vec{J}_{eff} \sim \partial(h F_0);$ Background \vec{E}_0 or \vec{B}_0 .

- Resonant cavity: $\omega_{\rm rf} \sim \omega_{J_{\rm eff}}$.
- **Circuit/magnetometer**: $B \sim |\vec{J}_{eff} V^{1/3}|$ from \vec{J}_{eff} where $1/\omega_{J_{eff}} \gg V^{1/3}$.



e.g. ADMX, HAYSTAC, CAPP, ORGAN, DM radio...



[Jiang et al, Nature Commun, 2305.00890]

Superconducting Radio-Frequency (SRF) Cavity

- **SRF cavities** are widely used for accelerators.
- Significant $Q_0 > 10^{10}$ compared to copper cavity with $Q_0 \le 10^6$.
- ► High Q_0 boosts dark photon searches [SERAPH 22', SHANHE 23']: $\epsilon \approx 10^{-16} \left(\frac{10^{10}}{Q_0}\right)^{\frac{1}{4}} \left(\frac{4 \text{ L}}{V}\right)^{\frac{1}{2}} \left(\frac{100 \text{ s}}{t_{\text{int}}}\right)^{\frac{1}{4}} \left(\frac{\text{GHz}}{f_0}\right)^{\frac{1}{4}} \left(\frac{T_{\text{amp}}}{3 \text{ K}}\right)^{\frac{1}{2}}.$
- Heterodyne upconversion [Berlin et al 19']: $\omega_{\rm rf} \omega_0 \approx \omega_a$ or ω_h .
- Both EM and mechanical coupling from GW [Berlin et al 21' 23'].



▶ Niobium superconductor requires $B_0^{\text{max}} < 0.2$ T, still $g_{a\gamma}$ or $h \sim 1/Q_0^{1/4}$.

International SRF Campaigns

Fermilab SQMS

•SERAPH:

Dark photon dark matter searches.

•Dark SRF:

Light-shining-wall search for dark photon.



DESY/SQMS:

•MAGO 2.0

HFGW searches with mechanical couplings.







DPDM Scan Search [SHANHE Collaboration]

► TM₀₁₀ of 1-cell elliptical cavity: largest overlapping with DPDM.



- Cavity and amplifier positioned in 2 K liquid helium.
- Mechanical turner scans resonant frequency f₀.
- Each scan is followed by calibration of f₀ and its stability range Δf₀.
- Drift $\delta f_d \leq 1.5 \text{ Hz and}$ microphonics $\sigma_{f_0} \approx 4 \text{ Hz}$ $\rightarrow \Delta f_0 \approx 10 \text{ Hz}.$



Tuner arm

Motor Piezo

Data Analysis and Constraints

- ► Total 1150 scan steps with each 100 s integration time.
- Group every 50 adjacent bins and perform a constant fit to address small helium pressure fluctuation.
- Normal power excess shows Gaussian distribution:



[SHANHE, PRL 2305.09711]

► Scan search with SRF and most stringent constraints in most exclusion space near $f_0 \approx 1.3$ GHz.

Cavity as Radio Telescope for Dark Photon

Galactic boosted dark photon from dark matter decay:

- Perturbative cascade decay from standard halo. [ADMX Dror et al 23']
- Parametric resonant production from scalar clump?

Polarization-dependent production:

- Longitudinal mode from a dark higgs.
- Transverse mode from axion-photon-type coupling.

Cavity as radio telescope for dark photon.

 Diurnal modulation to distinguish direction and polarization.



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Diurnal modulation constraints:

• L-mode enhanced from $|\vec{A'}| \sim \omega_{A'}/m_{A'} \gg 1$.



Response Width for Multi-mode Resonators

Broadened response in multi-mode resonators [YC et al, PRR 2103.12085, 2309.12387]:



Realization via Josephson junctions [Wurtz et al 21', Jiang et al 23', CEASEFIRE].

► New sensitivity limit for multi-mode resonators, optimal for SRF cavity: $\frac{\Delta \omega_r^{\text{MM}}}{\Delta \omega_r^{\text{SM}}} \propto \left(\frac{g}{\gamma n_{\text{occ}}}\right)^{\frac{2N}{2N+1}} \rightarrow \frac{Q_0}{n_{\text{occ}}} \text{ as } N \gg 1, \ g \rightarrow \omega_{\text{rf}}, \ \Delta \omega_r^{\text{MM}} \rightarrow \omega_{\text{rf}}.$

Simultaneous Resonant and Broadband Detection

- *e*-fold time: 10^7 s.
- DC cavity and LC circuits: SNR²_{MM}/SNR²_{SM} ~ Q₀/n_{occ}
 High n_{occ} of LC circuits at low frequency made enhancement ineffective.



Simultaneous scan $N_e = 6$ orders of ω_{Ψ} with significant response:

 $rac{\mathrm{SNR}^2_{\mathrm{MM}}}{\mathrm{SNR}^2_{\mathrm{SM}}}\simeq N_e rac{\overline{\omega}_\Psi \, Q_0}{\omega_{\mathrm{rf}} \, n_{\mathrm{occ}}}$



Summary

 Resonant cavities or circuits are powerful detectors for ultralight bosons and HFGW.

► SRF cavity with significant Q_0 has significant sensitivity: $\epsilon/g_{a\gamma}/h \propto 1/Q_0^{1/4}$.

 SRF cavity as radio antenna for dark photon: dissection of polarization and angular distribution.

Multi-mode resonantors have broadened response.

 \rightarrow Simultaneous resonant and broadband detection for SRF upconversion.





Thank you!

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Appendix

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$$-\frac{1}{2}\nabla^{\mu}\mathbf{a}\nabla_{\mu}\mathbf{a}-\frac{1}{2}\nabla^{\mu}\phi\nabla_{\mu}\phi-\frac{1}{4}\mathbf{\textit{F}}^{\prime\mu\nu}\mathbf{\textit{F}}_{\mu\nu}^{\prime}-\mathbf{\textit{V}}(\Psi),\quad\Psi=\mathbf{a},\phi\text{ and }\mathbf{\textit{A}}_{\mu}^{\prime}.$$

- Axion: hypothetical pseudoscalar motivated by strong CP problem.
- Prediction from fundamental theories with extra dimensions:

e.g. $g_{MN}(5D) \rightarrow g_{\mu\nu}(4D) + A'_{\mu}(4D), \quad A'_{M}(5D) \rightarrow A'_{\mu}(4D) + a(4D).$ String axiverse/photiverse: logarithmic mass window, $m_{\Psi} \propto e^{-\mathcal{V}_{6D}}$.

• Coherent wave dark matter candidates when $m_{\Psi} < 1$ eV:

$$\Psi(\mathrm{x}^\mu)\simeq \Psi_0(\mathbf{x})\cos\omega t; \qquad \Psi_0\simeq rac{\sqrt{
ho}}{m_{W}}; \qquad \omega\simeq m_\Psi.$$

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Axion Photon Coupling and Cavity Haloscope

• Axion photon coupling: $\propto g_{a\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$,

mixture with π_0 and anomaly generation.

$$\rightarrow \nabla \times \mathbf{B} = \partial_t \mathbf{E} + \mathbf{J} - g_{a\gamma} \left(\mathbf{E} \times \nabla a - \mathbf{B} \partial_t a \right).$$



► Background static B_0 → resonant when $\omega_{\rm rf} = m_a \sim V^{-1/3} \sim \mathcal{O}(1)$ GHz.

e.g. ADMX, HAYSTAC, CAPP, ORGAN ···



Resonant LC circuit

- Resonant conversion happens when $m_a \simeq \omega_{\rm rf} = \frac{1}{\sqrt{LC}}$ [Sikivie et al 13'].
- Scanning the mass from 100 Hz to 100 MHz by tuning the capacitor C.



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e.g. DM radio, ADMX-SLIC

Heterodyne Upconversion with SRF Cavity

$$\left(\partial_t^2 + \gamma \partial_t + \omega_{\rm rf}^2\right) \mathbf{E}_{\rm rf} = g_{a\gamma} \partial_t \left(\mathbf{B}_{\mathbf{0}} \partial_t a\right).$$

Heterodyne upconversion [Berlin et al 19']:

injecting AC pump mode

 $\partial_t(\mathbf{B_0}) = i\omega_0\mathbf{B_0}, \quad \omega_{\mathrm{rf}} \simeq \omega_0 + m_a.$



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- Large overlapping between B_0 and $E_{\rm rf}$ is required.
- Tune ω_{rf} ω₀ from Hz to GHz:
- Sensitivity benefits from superconducting nature: Q₀ > 10¹⁰.



• A new U(1) vector couples in different portals with SM particles:

 $\epsilon F'_{\mu\nu}F^{\mu\nu} + A'_{\mu}\bar{\psi}\gamma^{\mu}(g_{V} + g_{A}\gamma_{5})\psi + F'_{\mu\nu}\bar{\psi}\sigma^{\mu\nu}(g_{M} + g_{E}\gamma_{5})\psi.$

- Cavity/circuits for kinetic mixing, optomechanics for hidden U(1), spin sensors for dipole couplings...
- Similar to axion: extra dimensions, misalignment production (or during inflation), coherent wave.
- Novel aspects: three polarization degrees of freedom:

Longitudinal mode: $\vec{\epsilon}_0(\vec{k}) \propto \vec{k}$.

Transverse modes: $\vec{\epsilon}_{R/L} \perp \vec{k}$.

Signals projected to the sensitive direction of a vector sensor: $\sim \vec{\epsilon} \cdot \hat{l}$.

Kinetic Mixing and Hidden U(1) Dark Photon

• Effective currents from $\epsilon F'_{\mu\nu}F^{\mu\nu}$: $A'_{\mu} \rightarrow \vec{J}_{eff}$. **Kinetic mixing U(1)** dark photon shows up in circuit/cavity. [Chaudhuri et al 15'] or geomagnetic fields [Fedderke et al 21'];





► Force from $g_V A'_{\mu} \bar{\psi} \gamma^{\mu} \psi$: $A'_{\mu} \rightarrow \vec{F}$. U(1) B-L & B shows up in optomechanics [Graham et al 15', Pierce et al 18'] Or astrometry [Graham et al 15', PTA, GAIA].



Large Shield Room for Dark Photon Dark Matter

Dark photon dark matter induces

$$\boldsymbol{B} \approx |\vec{\mathbf{J}}_{\rm eff}| \ \boldsymbol{V}^{1/3} \approx 10^{-12} \, \varepsilon \left(\frac{m_{A'}}{10 \, {\rm Hz}}\right) \left(\frac{\boldsymbol{V}^{1/3}}{1 \, {\rm m}}\right) \, {\rm T},$$

Two spatially separated large shield room (8m³) with magnetometers placed on the wall: [Jiang et al 23']



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Long-baseline correlation suppresses common-mode noise.

High-Frequency Gravitational Waves

 Gravitational waves (GW) above 10 kHz have no known astrophysical origins.



Inverse Gertsenshtein effect:

$$\frac{1}{2}h^{\mu\nu}T^{\rm EM}_{\mu\nu} \rightarrow J^{\mu}_{\rm eff} = \partial_{\nu}\left(\frac{1}{2}hF^{\mu\nu} + h^{\nu}{}_{\rho}F^{\rho\mu} - h^{\mu}{}_{\rho}F^{\rho\nu}\right).$$

э.

Mechanical resonance: cavity deformation and mode transition, [MAGO 2.0].

Galactic Boosted Dark Photon

Galactic boosted dark photon from dark matter decay:

- Perturbative cascade decay from standard halo.
 [ADMX Dror et al 23']
- Parametric resonant production from scalar clump?

Polarization-dependent production:

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Diurnal Modulation from Earth Rotation

 X^{μ} Angular-dependent sensitivity to relativistic dark photon characterized by overlapping factor $C(\theta)$. [ADMX Dror et al 23' for galactic axion] TM₀₁₀ Detector is rest at Earth frame while Earth is rotating in galactic frame. 0.3F — L–mode - T-mode 0.2 $C_P(\theta)$ Diurnal modulation of the signals 0.1 in cavity. 0 $\pi/4$ $\pi/2$ Longitudinal and transverse modes θ

show opposite variation.

SRF Constraints for Galactic Dark Photon

Same dataset as dark photon dark matter searches:

- Total scan range of ~ 1 MHz: within bandwidth of galactic dark photon.
- Total experimental time of ~ 60 hours: daily modulation tests.
- Cosmology requires $\rho_{A'} \leq 1000 \, \rho_{\gamma}$ on Earth.
- Constraints for longitudinal modes are more stringent due to its spatial wavefunction is $\sim \omega_{A'}/m_{A'}$.

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Simultaneous Resonant and Broadband Detection for Dark Sectors

based on

arxiv: 2103.12085, Phys. Rev. Res. **4** (2022) no.2, 023015 arxiv: 2309.12387

in collaboration with Jiang, Li, Liu, Ma, Shu, Yang, Zeng

Quantum noise limit for resonant detection

Standard quantum limit for power law detection: [Chaudhuri et al 18']

Noise PSD: resonant intrinsic noise S_{int} + flat readout noise S_r .

Sensitivity to S_{sig} and S_{int} is the same. $SNR^2 \propto \Delta \omega_r \ (S_{int} \gg S_r).$

Beyond standard quantum limit:

Squeezing *S*_r, e.g., HAYSTAC.

Increasing the sensitivity to S_{sig} , e.g., white light cavity in optomechanics/GW detection [Miao et al 15'].



 $S_{
m int} \propto$ Cauchy distribution

White Light Cavity for Axion [Li et al 20']



- **Beam-splitting**: $\hbar g(\hat{a}\hat{b}^{\dagger} + \hat{a}^{\dagger}\hat{b})$.
- Non-degenerate parametric interaction: $\hbar G(\hat{b}\hat{c} + \hat{b}^{\dagger}\hat{c}^{\dagger})$.
- ► \mathcal{PT} -symmetry $(\hat{a} \leftrightarrow \hat{c}^{\dagger})$ emerges when g = G. $(\dot{a} + \dot{c}^{\dagger}) = -i(g - G)\hat{b} - i\alpha\Psi + \cdots;$ $\dot{\hat{b}} = -\gamma_r\hat{b} - ig(\hat{a} + \hat{c}^{\dagger}) + \cdots.$
- Coherent cancellation leads to **double resonance**. S_{sig} is largely enhanced when $g \gg$ intrinsic dissipation γ :

$$S_{\rm sig}^{\rm WLC}(\Omega) = \frac{2\gamma_r \alpha^2 S_{\Psi}(\Omega)}{(\gamma + \gamma_r)^2 + \Omega^2} \left(\frac{g^2}{\gamma^2 + \Omega^2}\right).$$
 Readout coupling γ_r

 $\hbar \alpha (\hat{a} + \hat{a}^{\dagger}) \Psi$

Response Width for Multi-mode Resonators

Signal response width can be significantly broadened in a multi-mode system compared to single-mode ones:



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New sensitivity limit for multi-mode resonators.

Simultaneous Resonant and Broadband Detection

- *e*-fold time: 10^7 s.
- DC cavity and LC circuits: SNR²_{MM} SNR²_{SM} ~ <u>n_{occ}</u> of LC circuits at low frequency made enhancement ineffective.



Simultaneous scan $N_e = 6$ orders of ω_{Ψ} with significant response.

High Q_0 and constant n_{occ} enable reaching QCD axion with $m_a > kHz$.



Property of Ultralight Dark Matter

Galaxy formation: virialization $\rightarrow \sim 10^{-3}c$ velocity fluctuation, thus kinetic energy $\sim 10^{-6}m_{\Psi}c^2$. Effectively coherent waves:

$$\Psi(ec{x},t) = rac{\sqrt{2
ho_{\Psi}}}{m_{\Psi}} \cos\left(\omega_{\Psi}t - ec{k}_{\Psi}\cdotec{x} + \delta_0
ight).$$

• Bandwidth:
$$\delta \omega_\Psi \simeq m_\Psi \left< v_{\rm DM}^2 \right> \simeq 10^{-6} m_\Psi$$
, $Q_\Psi \simeq 10^6$.

- Correlation time: τ_Ψ ≃ ms 10⁻⁶eV/m_Ψ.
 Power law detection is used to make integration time longer than τ_Ψ.
- ► Correlation length: $\lambda_d \simeq 200 \text{ m} \frac{10^{-6} \text{eV}}{m_{\Psi}} \gg \lambda_c = 1/m_{\Psi}$. Sensor array can be used within λ_d .

Quantization of Cavity Modes

• Quantized EM modes with wavefunctions $\vec{\epsilon}_n(\vec{r})$ In Coulomb gauge:

$$\vec{A} = \sum_{n} \frac{1}{\sqrt{2\omega_{\mathrm{rf}}^{n}}} \hat{a}_{n}^{\dagger} \vec{e}_{n}(\vec{r}) e^{-\mathrm{i}\omega_{\mathrm{rf}}^{n}t} + h.c..$$

The Hamiltonian for each mode reduces to harmonic oscillator:

$$H_0 = \frac{1}{2} \int_V \left(\vec{E}^2 + \vec{B}^2 \right) \, \mathrm{d}V = \sum_n \omega_{\mathrm{rf}}^n \left(\hat{a}_n^\dagger \hat{a}_n + \frac{1}{2} \right),$$

Interaction with effective currents:

$$H_{\rm int} = \int_{V} \vec{A} \cdot \vec{J}_{\rm eff} \, \mathrm{d}V = \alpha \Psi \left(\hat{a} \, e^{\mathrm{i}\omega_{\rm rf} t} + \hat{a}^{\dagger} \, e^{-\mathrm{i}\omega_{\rm rf} t} \right) / \sqrt{2},$$

where α contains geometric overlapping factor $\eta_n \propto \int_V \vec{\epsilon}_n \cdot \vec{J}_{eff} dV$.

Quantization of Circuit Modes

Energy stored in an inductor and a capacitor:

$$\mathcal{H}_0 = \frac{\Phi^2}{2L} + \frac{Q^2}{2C} = \omega_{\rm rf} \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right).$$

• Interaction with external Φ_{Ψ} :

$$H_{\rm int} = rac{\Phi \, \Phi_{\Psi}}{L} = lpha \Psi \left(\hat{a} \, e^{\mathrm{i} \omega_{
m rf} t} + \hat{a}^{\dagger} \, e^{-\mathrm{i} \omega_{
m rf} t}
ight) / \sqrt{2}.$$

 Circuit representation of cavity modes with an antenna: Cavity
 LC circuit

$$\Phi = \int_{Ant} \vec{A}(\vec{r}, t) \cdot d\vec{l}.$$

A system interacting with environment:



System mode \hat{a} couples to infinite degrees of freedom \hat{w}_{ω} :

$$i\hbar\sqrt{2\gamma_r}\int_{-\infty}^{+\infty}rac{d\omega}{2\pi}[\hat{a}^{\dagger}\hat{w}_{\omega}-\hat{a}\hat{w}_{\omega}^{\dagger}]+\int_{-\infty}^{+\infty}rac{d\omega}{2\pi}\hbar\omega\hat{w}_{\omega}^{\dagger}\hat{w}_{\omega}.$$

Fourier transformation: 0-dim localized mode â couples to an 1-dim bulk w_ξ (transmission line):

$$i\hbar\sqrt{2\gamma_r}\hat{a}^{\dagger}\hat{w}_{\xi=0}+\mathrm{h.c.}+i\hbar\int_{-\infty}^{+\infty}d\xi\hat{w}^{\dagger}_{\xi}\partial_{\xi}\hat{w}_{\xi}.$$

• Equations of motion for \hat{a} and outgoing mode \hat{w}_{0_+} :

$$\dot{\hat{a}}=-\gamma_r\hat{a}+\sqrt{2\gamma_r}\hat{w}_{0_-};\qquad \hat{w}_{0_+}=\hat{w}_{0_-}-\sqrt{2\gamma_r}\hat{a}$$

Single-mode Resonator as Quantum Sensor

- For a resonator \hat{a} probing weak signal Ψ : $\alpha \left(\hat{a} + \hat{a}^{\dagger} \right) \Psi$
- Readout for outgoing mode $\hat{v}_r \equiv \hat{w}_{0_+}$:

$$\hat{v}_r = rac{\Omega - i\gamma_r}{\Omega + i\gamma_r}\hat{u}_r + rac{\sqrt{2\gamma_r}lpha}{\Omega + i\gamma_r}\Psi.$$



- Fluctuations in incoming mode û_r ≡ ŵ₀ with quantum limited power spectral density S_r = 1.
- Resonant signal spectrum $S_{sig} = \frac{\gamma_r \alpha^2}{\gamma_r^2 + \Omega^2} S_{\Psi}(\Omega)$.
- Trade-off between peak sensitivity and bandwidth by tuning γ_r.

Intrinsic loss and fluctuation

Intrinsic loss ∝ γ exists, characterized by quality factor Q_{int} ≡ ω/γ.



Fluctuation-dissipation theorem predicts intrinsic loss fluctuations

$$S_{
m int}(\Omega) = rac{4\gamma\gamma_r}{(\gamma+\gamma_r)^2+\Omega^2}n_{
m occ}.$$

Using scattering matrix elements:

$$S_{\rm sig} = |S_{0r}|^2 \frac{\alpha^2}{4\gamma} S_{\Psi}, \qquad S_{\rm noise} = |S_{0r}|^2 n_{\rm occ} + |S_{rr}|^2 \frac{1}{2} + \frac{1}{2}$$

 Standard quantum limit for power law detection: resonant S_{int}+ flat S_r. [Chaudhuri et al 18']

Binary Tree Haloscope



► Fully \mathcal{PT} -symmetric setup with $\hat{a}_{ij} \leftrightarrow \hat{c}_{ij}^{\dagger}$ brings strong robustness.

Multi-probing sensors leads to coherent enhancement:

$$S_{\mathrm{sig}}^{\mathrm{BT}}(\Omega) = 2^{2n-2} S_{\mathrm{sig}}^{\mathrm{RC}}(\Omega).$$

Quantum Limit for Multi-mode resonators

Scan bandwidth can be significantly increased in a multi-mode system.



Far beyond the one of single-mode resonators.

New quantum limit for multi-mode resonators.

Beam splitting coupling



Use an additional capacitor to couple two LC circuits:

$$H = \frac{1}{2}C_1\dot{\Phi}_1^2 + \frac{1}{2}C_2\dot{\Phi}_2^2 + \frac{1}{2L_1}\Phi_1^2 + \frac{1}{2L_2}\Phi_2^2 + \frac{1}{2}C_0(\dot{\Phi}_1 - \dot{\Phi}_2)^2.$$

Conjugate momentum to Φ_i involves mixing. Interaction potential:

$$eta\hbar\sqrt{\omega_1\omega_2}(\hat{a}_1-\hat{a}_1^\dagger)(\hat{a}_2-\hat{a}_2^\dagger)\sim \hat{a}_1\hat{a}_2^\dagger+h.c.,$$

Non-Degenerate Parametric amplifier coupling



Use a DC voltage and a Josephson junction to couple two LC circuits:

$$V = -\frac{\hbar I_J}{2e_0} \cos(\omega_0 t + \frac{2e_0}{\hbar} (\Phi_2 + \Phi_3))$$

= $-\frac{\hbar I_J}{2e_0} \cos(\omega_0 t + \kappa_2 (a_2 + a_2^{\dagger}) + \kappa_3 (a_3 + a_3^{\dagger}))$
 $\sim \frac{\hbar I_J}{4e_0} \kappa_2 \kappa_3 [a_2 a_3 + a_2^{\dagger} a_3^{\dagger}],$