# The Daily Modulations And Broadband Strategy In Axion Searches:

An Application With CAST-CAPP Detector

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on behalf of the CAST Collaboration Sep 5, 2024

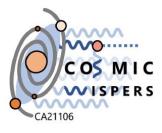
https://arxiv.org/abs/2405.10972

2nd General Meeting of COST Action COSMIC WISPers (CA21106)



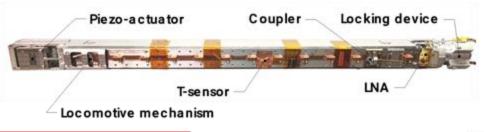


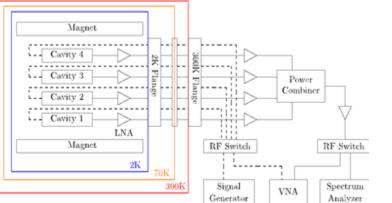


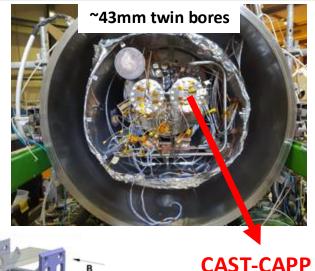


#### **CAST-CAPP** Detector

Four identical stainless steel tunable cavities electroplated with  $^{\sim}30\mu m$  of copper installed in one of the two twin bores of CAST magnet with the split plane parallel to the magnetic field.







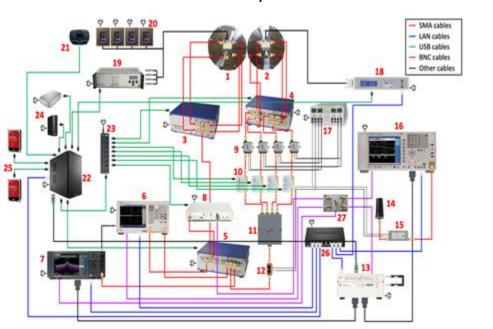
<u>Tuning mechanism</u>: 2 dielectric sapphire bars symmetrically placed parallel to the longitudinal sides, moving simultaneously towards the center and activated by a piezoelectric motor.

Sapphire strips

Locomotive mechanism

### CAST-CAPP Data Acquisition

- 1-min measurements
- Bandwidth = 5 MHz
- Tuning step size = 200 kHz
- Size = ~ 3 GB / file !!



#### **QUALITY CHECKS:**

#### Main source→Vibrations

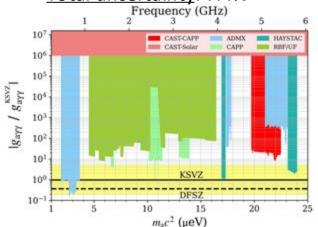
| Nr. | Parameters               | Criteria   |
|-----|--------------------------|--|
| 1   | Frequency stability      | $\Delta \nu_0 < 100  \mathrm{kHz}$                       |
| 2   | Amplitude variation      | $\Delta A_0 < 3\mathrm{dB}$                              |
| 3   | Quality factor           | $10^3 < Q_L < 4 \times 10^4$                             |
| 4   | Quality factor shift     | $\Delta Q_L < 7 \times 10^3$                             |
| 5   | Temperature variation    | $\Delta T_{ m cav} < 3   m K$                            |
| 6   | Temperature              | $1{ m K} < T_{ m cav} < 273{ m K}$                       |
| 7   | Magnetic field variation | $\Delta \vec{B} < 0.1\mathrm{T}$                         |
| 8   | Frequency mismatch       | $< 20\mathrm{kHz}$ (before) & $< 80\mathrm{kHz}$ (after) |
| 9   | Amplitude mismatch       | $< 1  \mathrm{dB}$                                       |
| 10  | Temperature mismatch     | < 3 K  |

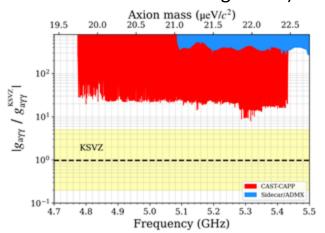
#### **RESULTS:**

- Data-taking time: 4124 h (172 d)
- Frequency range: 660.15 MHz
- Axion Masses: 4.77 5.43 GHz
- Data size: ~ 650 TB !!

### CAST-CAPP Results → Alternative analyses

- Total data-taking time: 4124 h (172 d)
- Data size: ~ 650 TB
- Frequency range: 660.15 MHz
- Axion masses: 19.74 μeV 22.47 μeV
- Present sensitivity:  $g_{avv} = 8 \times 10^{-14} \,\text{GeV}^{-1}$
- Confidence level: 90 %
- Total uncertainty: 7.4%





#### New analysis options:

- 1. Dark photons
- Transient events (high resolution time series measurements)
- 3. AQN daily modulations (24h, Low Q, fixed frequency measurements, temperature isolation of complete datataking chain.)

#### 

CAST-CAPP was not designed for daily modulation analysis. We present an analysis model on retrospective data that will pave the way for future experiments.

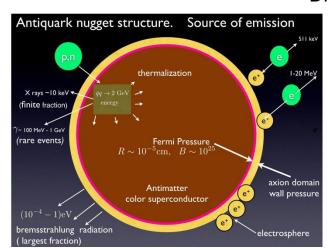
CAST Collab., Search for Dark Matter Axions with CAST-CAPP. Nat Commun 13, 6180 (2022).

### Axion Quark Nugget (AQN) Model

- Originally proposed by Ariel Zhitnitsky (2003) to explain  $\Omega_{DM} \sim \Omega_{visible}$ .
- Also explain other mysteries (core-cusp, solar corona etc).
- Composite particles with axion domain walls.
- Relativistic axions with <u<sub>a</sub>> ~ 0.6c are emitted from AQNs
  - as they penetrate the Earth.
- Production mechanism would cause:
  - 1. Daily modulation (~10-20%).
  - 2. Seasonal phase shift.



**DFSZhitnitsky** 



### Axion Quark Nugget (AQN) Model

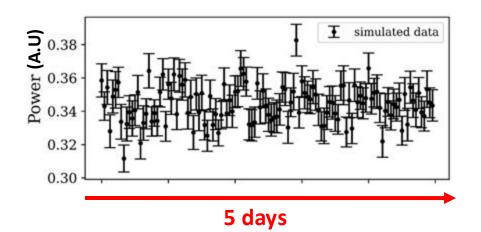
- AQN induced Axions have  $\langle v_a \rangle$  ~0.6c opposed to  $\langle v_a \rangle$  ~10<sup>-3</sup>c of galactic DM axion  $\Longrightarrow$  broadband detection strategy for haloscopes.
- AQNs interact with Earth and lose its mass  $\Rightarrow$  Daily modulation (~10-20%)

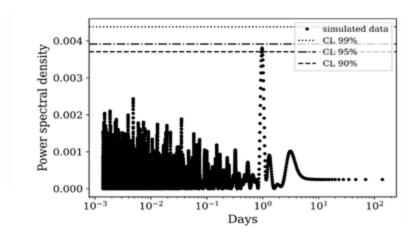
$$ho_{
m a}^{
m AQN}(t)\equiv A_{
m (d)}(t)\langle
ho_{
m a}^{
m AQN}(R_{\oplus})
angle \ A_{
m (d)}(t)\equiv [1+\kappa_{
m (d)}\cos(\Omega_d t-\phi_0)]$$
 Seasonal phase shift  $\phi_0\simeq\pi$  For opposite seasons

Any haloscope experiment can analyze for AQN induced modulations without interfering with conventional searches

#### Simulations

- AQN power= 0.7% of the mean spectrum power
- 1-min spectra created by real background + random noise + AQN power
- 5 days / 1 min = 7200 simulated spectra
- Create a periodogram using Lomb Scargle method
- ML algorithm to find the best decision threshold

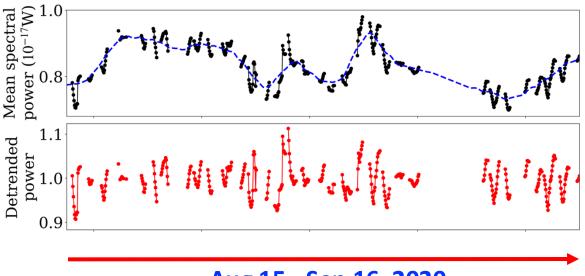




#### Data Treatment

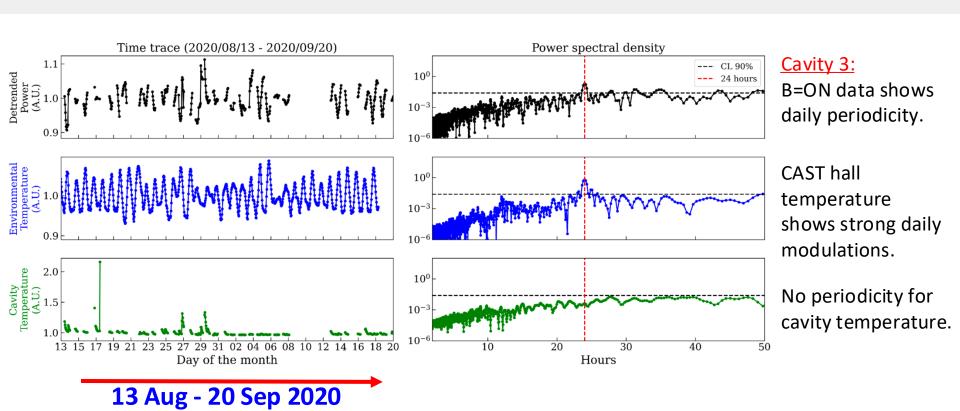
- Phase-matched data can not be used (attenuation added).
- Fast-tuning data can not be used.

- Single cavity data in fixed-frequency measurements.
- Use of high-quality data (clean from mechanical vibrations).
- ❖ Use of both B=ON & B=OFF data.
- **♦** All data are detrended before analysis

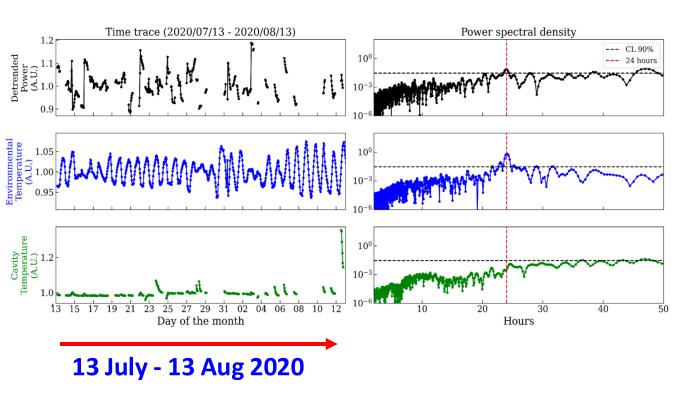


Aug 15 - Sep 16, 2020

### Data Analysis - B=ON



### Data Analysis - B=OFF



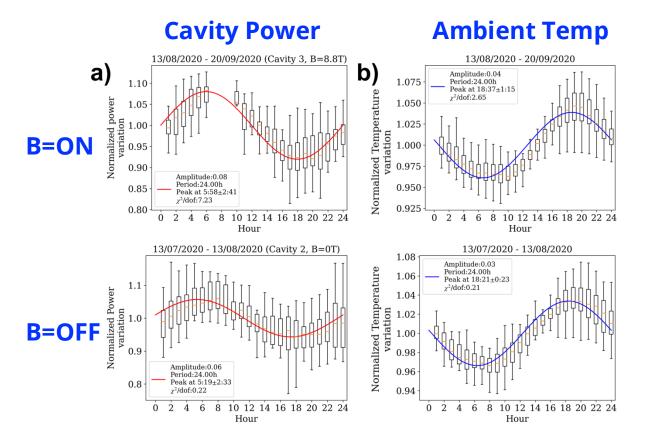
#### Cavity 2: Both B=ON and B=OFF show daily variations.

Origin may not be due to temperature variations OR temperature dependency might be more complex than anticipated

## Further investigation required.

We cannot verify the AQN induced daily modulations without first understanding the daily variations included in B=OFF data.

### Data Analysis – Hourly Distributions



#### ~12 hours phase shift

A similar dependence is expected for the electronic gain of the amplifier chain of the DAQ system

### Discussion / Conclusions

CAST-CAPP data analyzed for signatures (daily modulation) as expected from the AQN model:

- Daily modulation observed in both B=ON and B=OFF data.
- Temperature dependency of the gain of 2<sup>nd</sup> stage (room temp) LNAs is the most probable scenario.
- Voltage drifts less probable due to the existence of a line-interactive UPS.
- Other parameters need to be checked. But CAST is closed.
- **❖** Proof of principle **→** Done
- No conclusive result.

Any haloscope experiment can analyze for AQN induced modulations without interfering with conventional searches.



