



MAX-PLANCK-INSTITUT
FÜR PHYSIK

STATUS OF THE RADES-MPP ACTIVITIES IN HALOSCOPE EXPERIMENTS

Dr. Jose María García Barceló

2nd General Meeting of COST Action COSMIC WISPers (CA21106) – 3rd Sep/2024

ERC-STG802836



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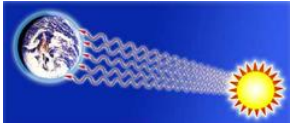
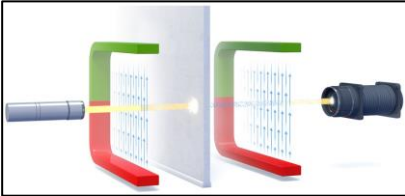
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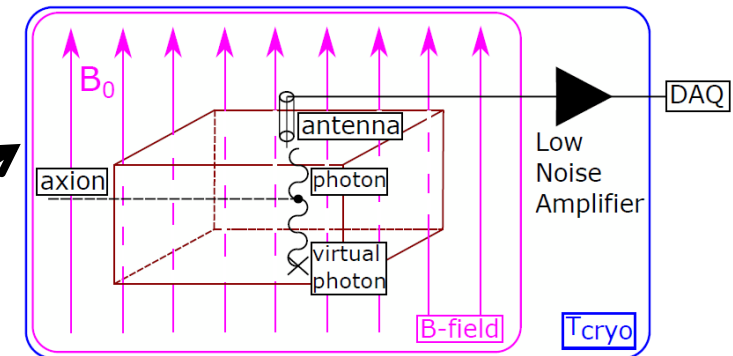
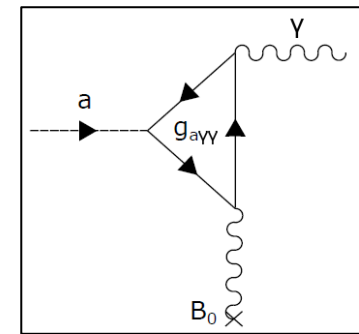
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History of the RADES collaboration

Motivation and goal

- **Axions** elegantly solve 2 problems from one
- Cold Dark Matter between 1-1000 μeV .
- Weak EM interaction. However, potential detection by the **axion-photon coupling**.
- Types of detectors:
 -  Haloscopes
 -  Helioscopes
 -  Light Shining Walls
- Conversion concept: inverse Primakoff effect
- **Haloscope** technology: **microwave cavity** immersed in a high static magnetic field for boosting the axion-photon conversion.



History of the RADES collaboration

RADES team

~30 people from different institutes around the world



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1st RADES collaboration meeting in 2019 at UNIZAR

2nd RADES collaboration meeting in 2024 at MPP

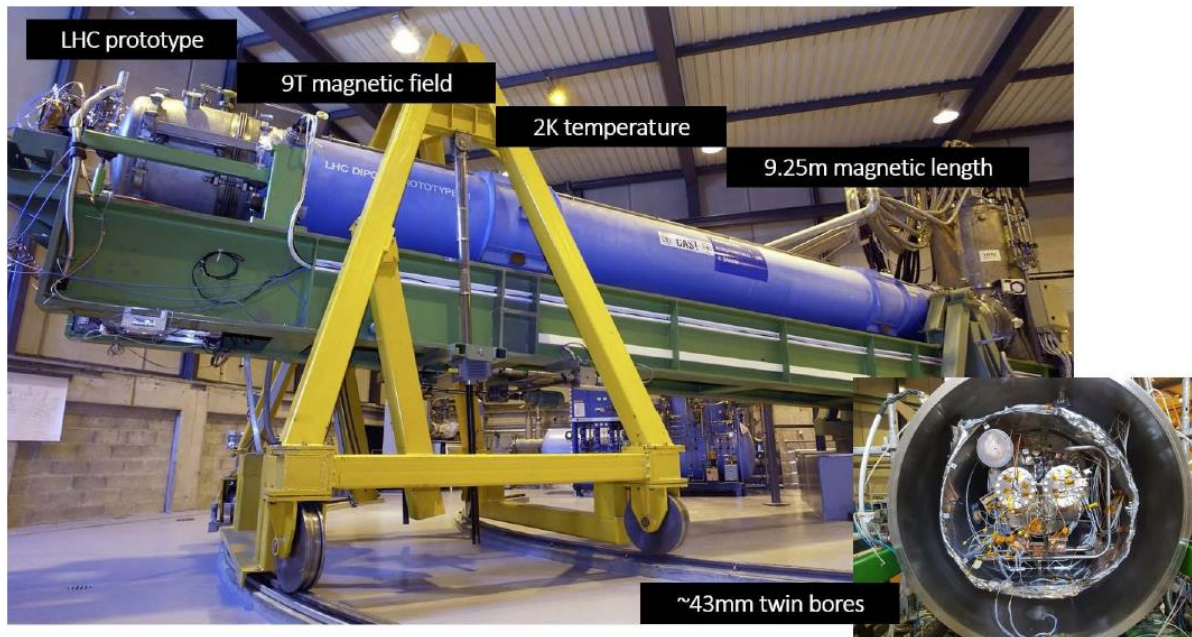


Next RADES collab. meeting → February/2025 in Valencia

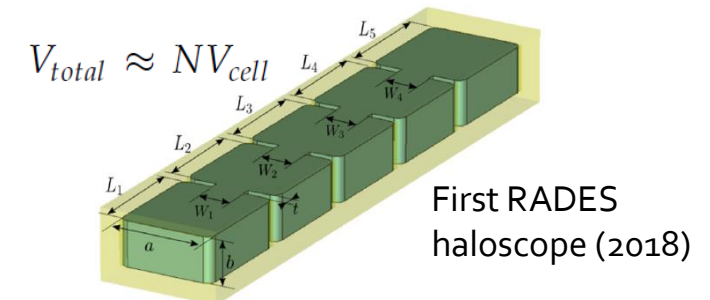
History of the RADES collaboration

RADES origins

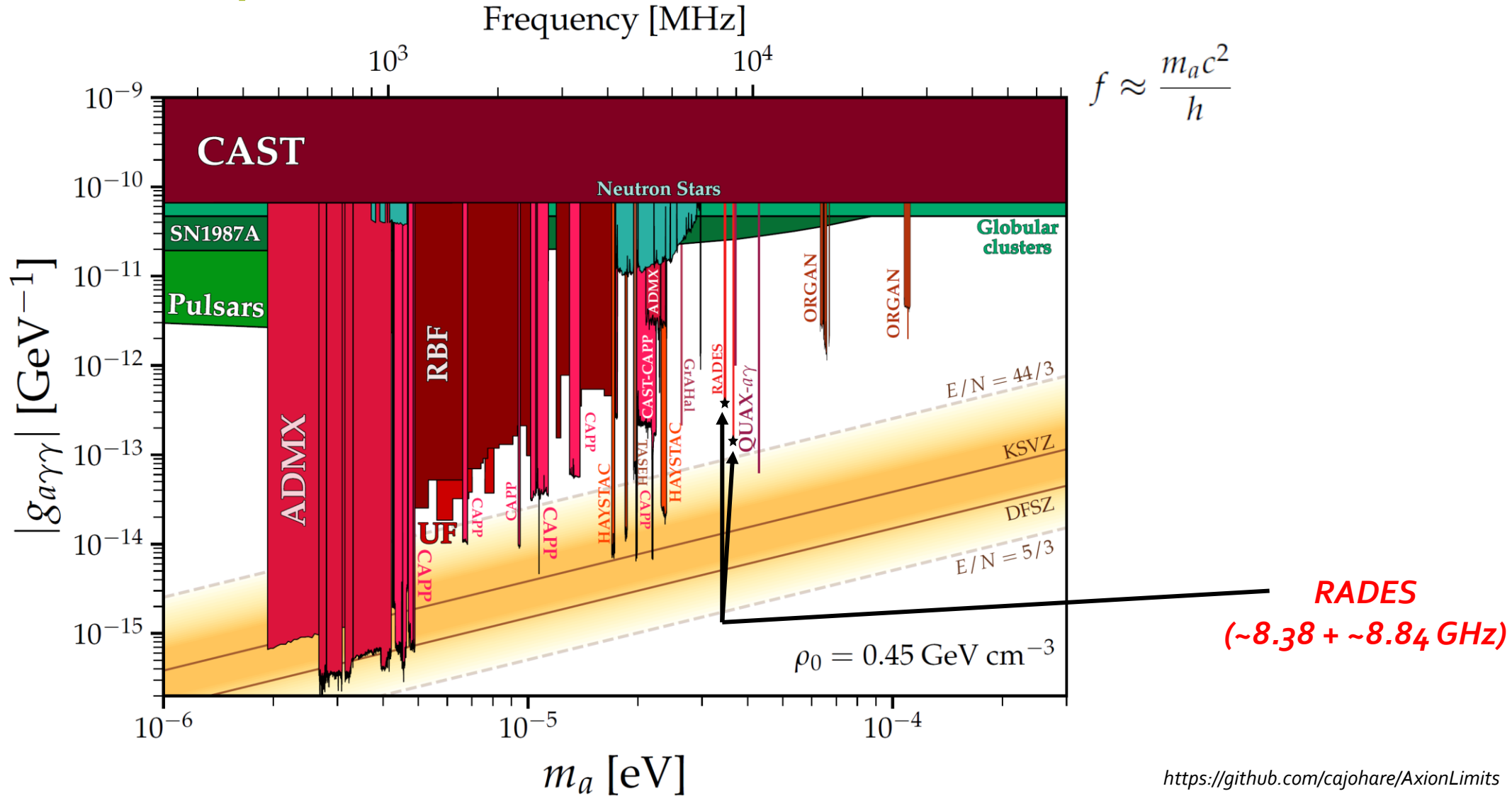
- The RADES (Relic Axion Detection Exploratory Setup) team originated in 2016 within the framework of the CAST (CERN Axion Solar Telescope) experiment, for the search for dark matter axions with haloscopes.



The first haloscope designs were based on the multi-cavity concept:



History of the RADES collaboration



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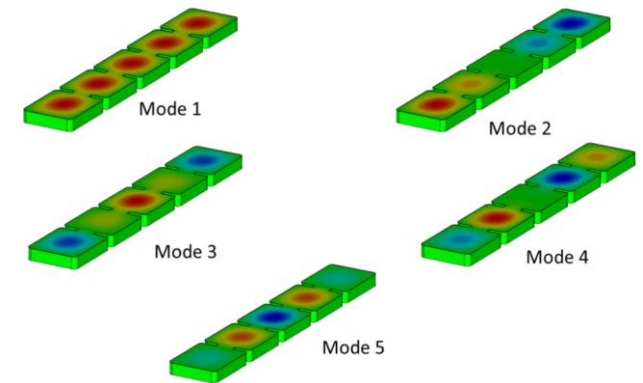
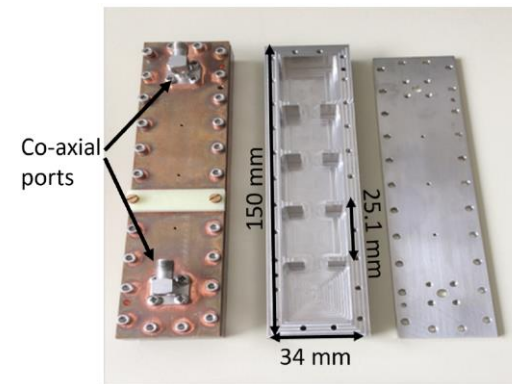
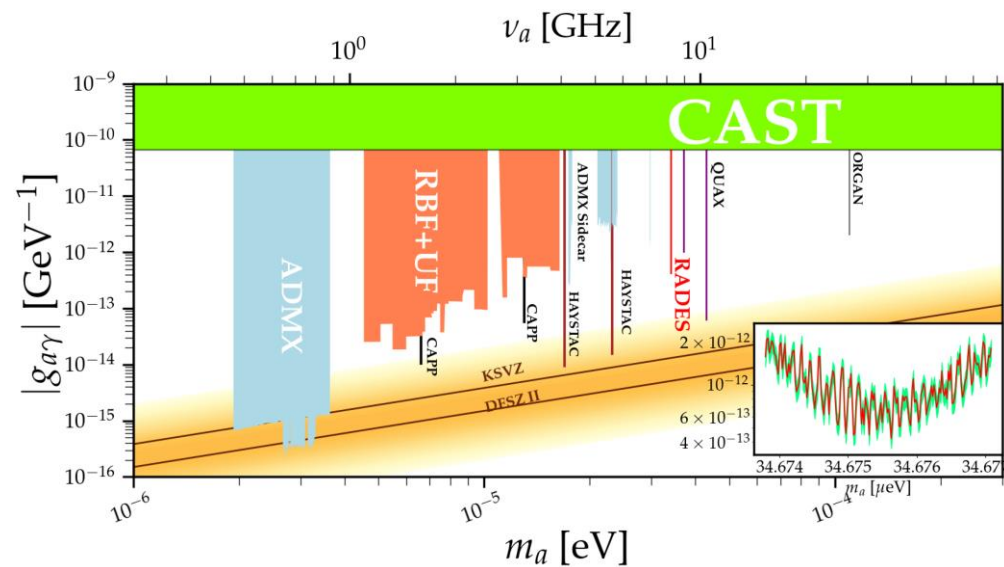
Axion data taking campaigns

- 2 done + 1 in preparation
- 2018 at CAST (CERN) with a multicavity
- 2021 at SM18 (CERN) with a HTS cavity
- 2024 at SM18 (CERN) with a HTS cavity

Axion data taking campaigns

2018 at CAST (CERN) with a multicavity

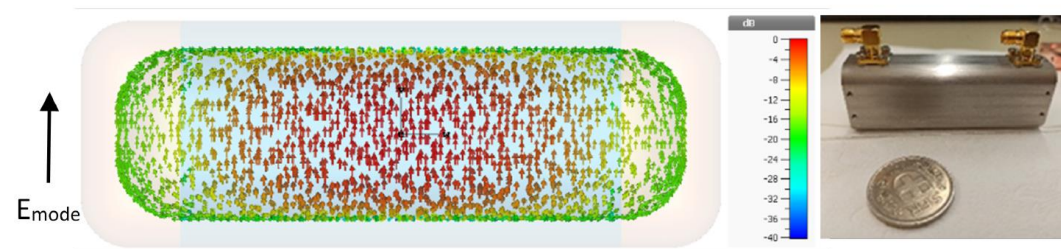
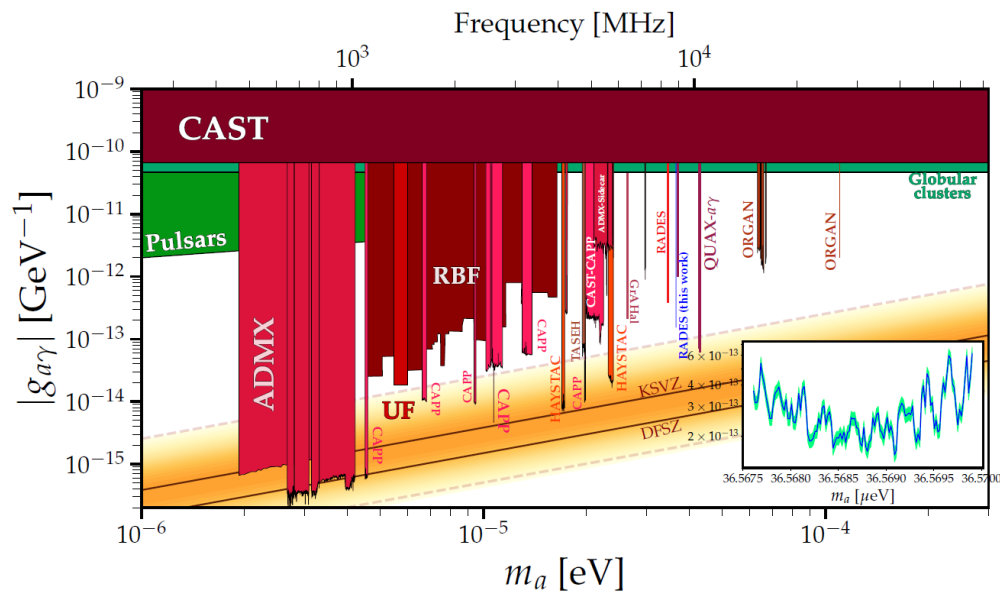
- CAST magnet (9T + 1.9 K)
- Multicavity at 8.38 GHz
- Results (limits) published in: [https://doi.org/10.1007/JHEP10\(2021\)075](https://doi.org/10.1007/JHEP10(2021)075)



Axion data taking campaigns

2021 at SM18 (CERN) with a HTS cavity

- CAST stopped so we explored other activities: ARIES TNA
- SM18 magnet (11.7 T + 1.9 K)
- Single cavity with HTS tapes ($\uparrow Q_0$) at 8.84 GHz (small effective tuning through helium pressure change)
- Results (limits) published in: <https://arxiv.org/abs/2403.07790>



Axion data taking campaigns

2024 at SM18 (CERN) with a HTS cavity

- New window to take axion data at the SM18 magnet (11.7 T + 1.9 K) in November/2024
- Single cavity with HTS tapes ($\uparrow Q_0$)

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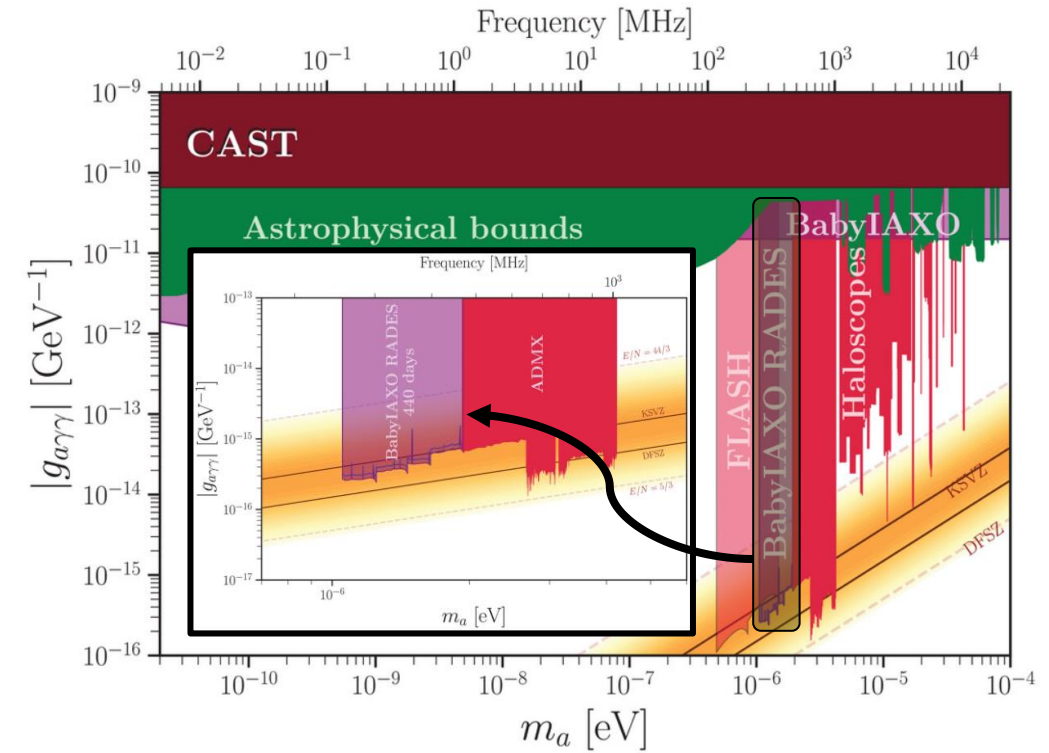
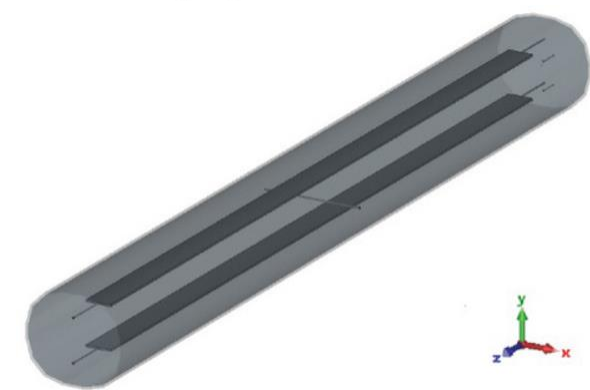
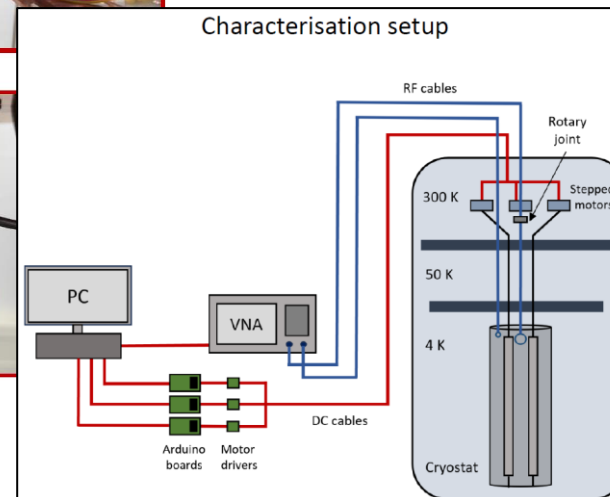
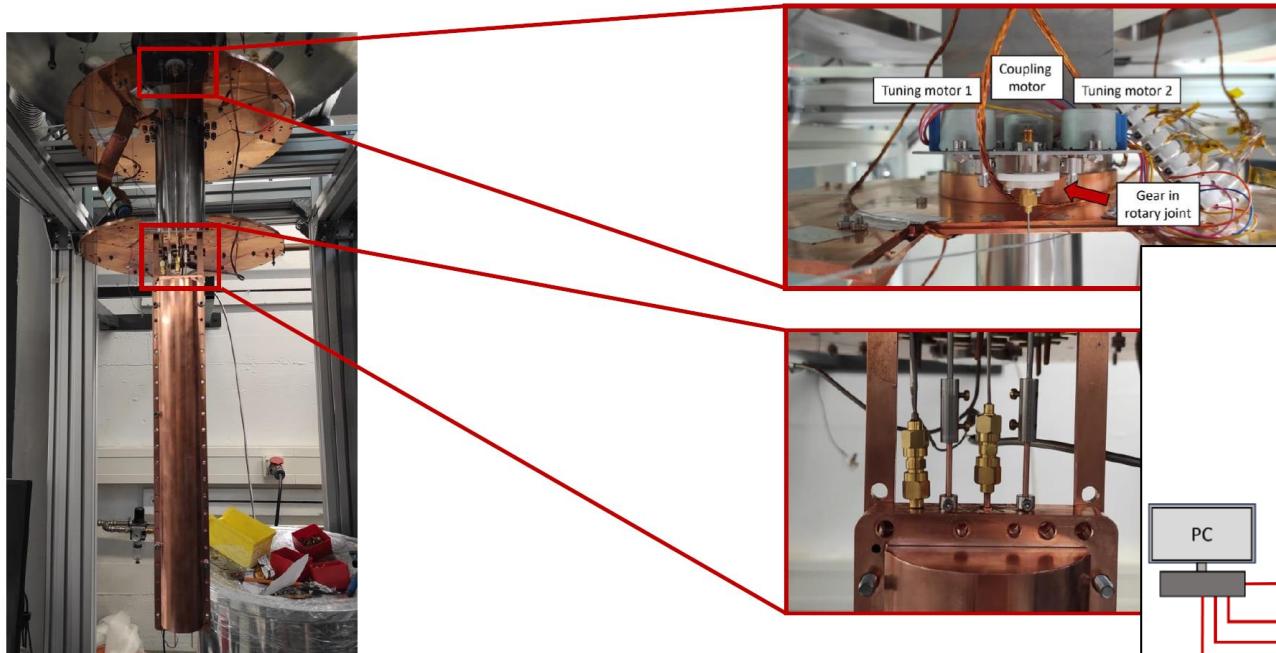
Future plans in BabyIAXO

- The BabyIAXO experiment, at DESY (Hamburg, Germany), is an intermediate step to explore further possible improvements to the final IAXO telescope. Helioscopes and haloscopes could be installed.
- $B = \sim 2.5T$
- Dimensions: $\varnothing = 60 \text{ cm}$, and $L = 10 \text{ m}$.



Future plans in BabyIAXO

- RADES recently suggested (in [Ann.Phys. 535, 12](#)) the use of low-frequency (around [250-470] MHz) axion haloscope configurations (four 5-meter-long cavities) appropriate for operation within the BabyIAXO magnet.
- Preliminary version (10 times smaller, 50 cm in length):

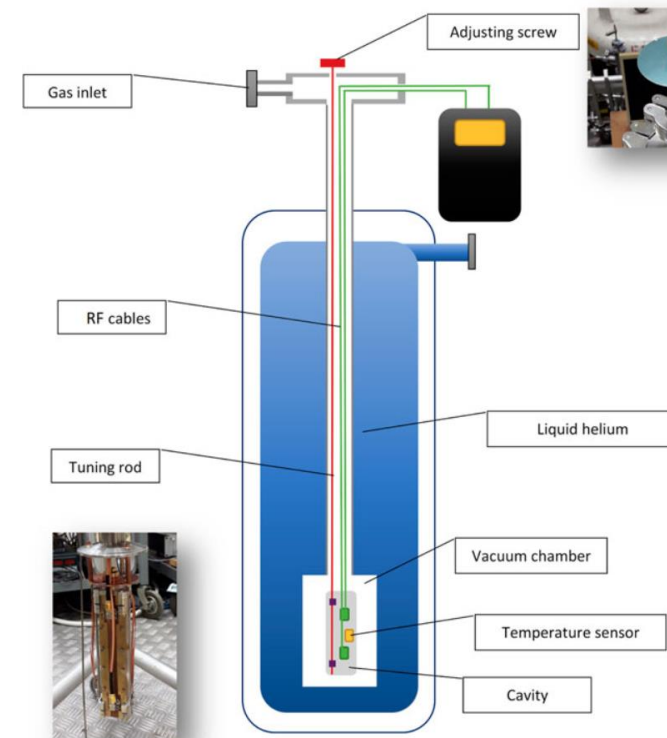
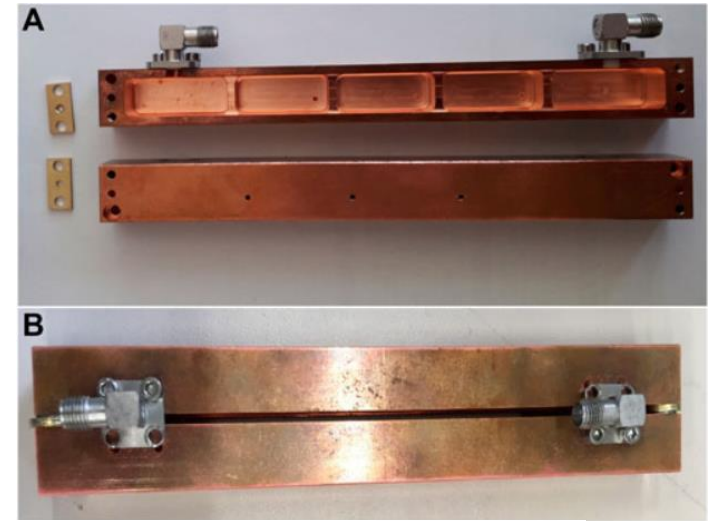


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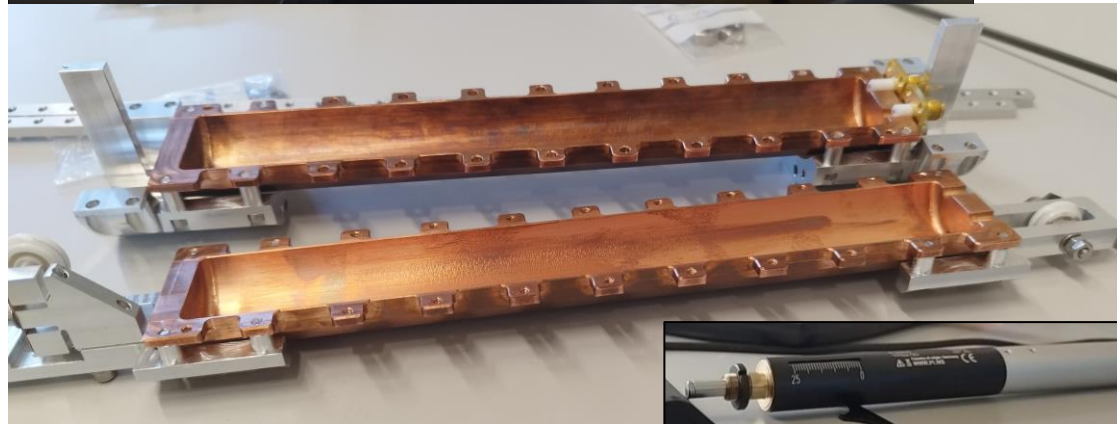
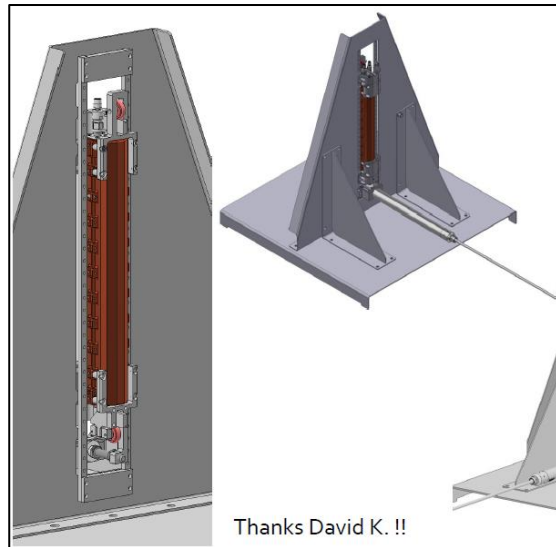
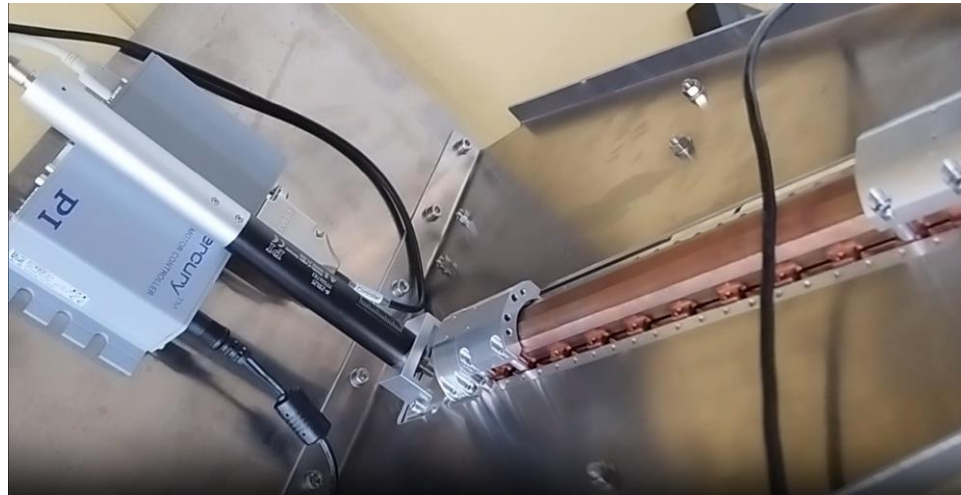
Other R&D activities - Tuning

- Throughout the years of the RADES research, different frequency tuning mechanisms have been explored: ferroelectric, ferromagnetic, vertical cut, rotating plates, helium bath pressure variation, etc.
- The two most explored concepts have been ferroelectric (still without experimental results) and vertical cut tuning. For the latter, an investigation has been carried out in one of the RADES cavities of 5 sub-cavities with a longitudinal cut. The tuning here is based on the controlled opening between the two parts of the haloscope housing in order to modify the width and thus the operating frequency (see [Front. Phys. 12:1372846](#)).



Other R&D activities - Tuning

- At the MPP (Munich, Germany) we are also exploring this concept for a single cylindrical cavity. 4K and 10 mK tests are foreseen.



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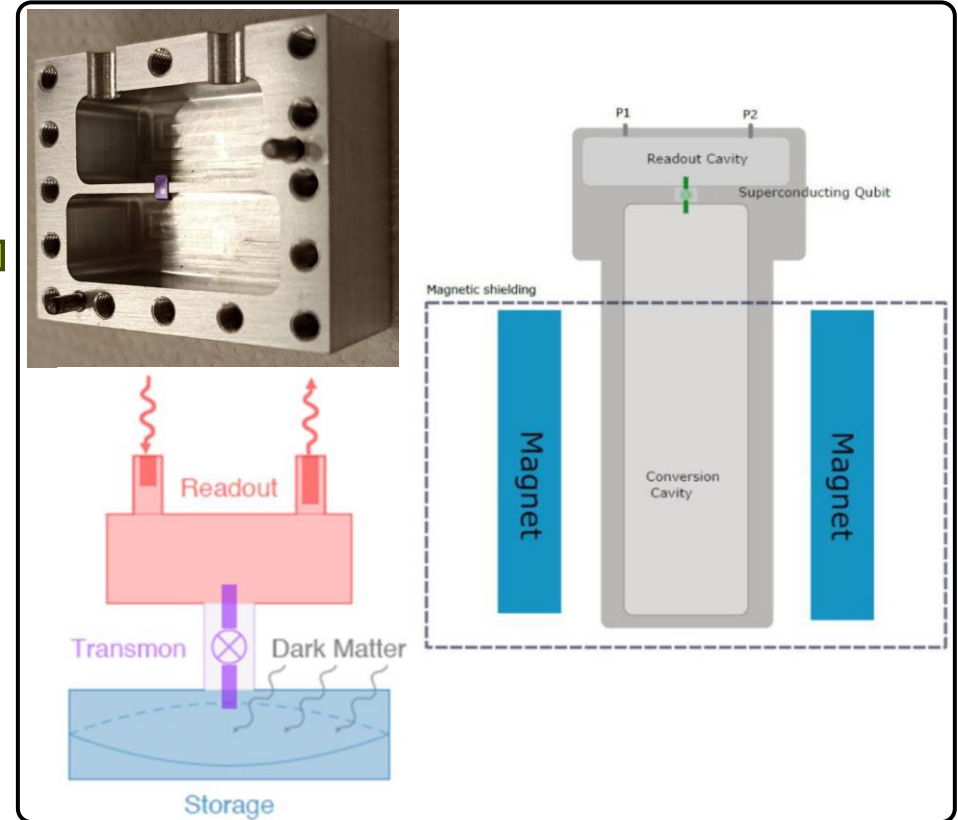
Other R&D activities - Single-photon technologies

- Haloscope sensitivity is determined by the detector's noise: $T_{\text{sys}} = T_{\text{phys}} + T_{\text{sen}}$
- Lowering T_{phys} by cooling down the experiment will only increase sensitivity if T_{sen} is also decreased, so we need:
 - Quantum-limited amplification
 - Single-photon detectors (to overcome standard quantum limits).

Other R&D activities - Single-photon technologies

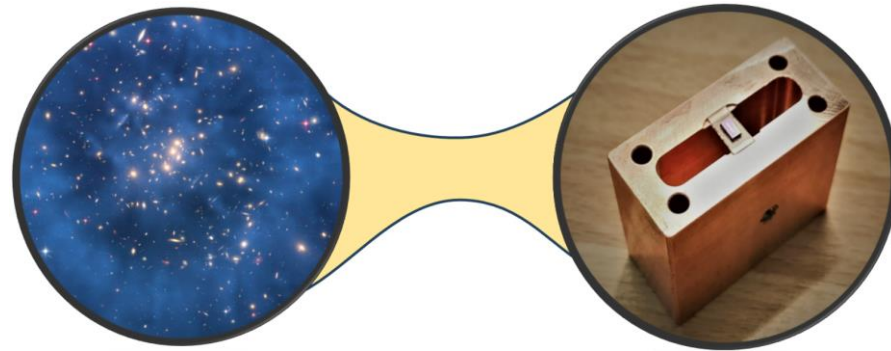
➤ Explored quantum technologies:

- Superconducting QuBits (transmons)
- Nano-TES (Transition Edge Sensor)



Other R&D activities - Single-photon technologies

- As part of DarkQuantum project (ERC Synergy Grants 2023), we plan to develop a qubit-based QSPC and overcome all challenges for implementation in RADES.



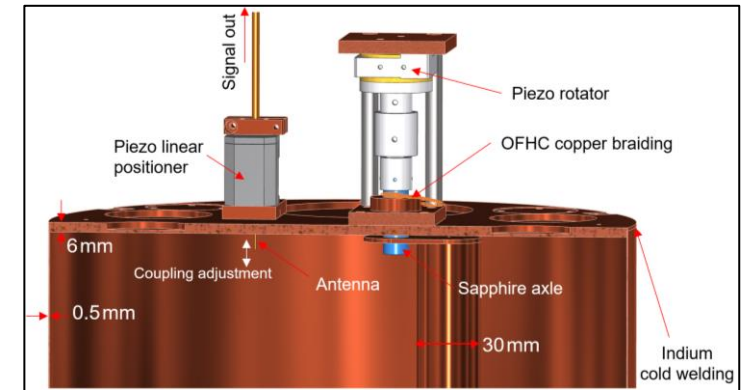
DarkQuantum

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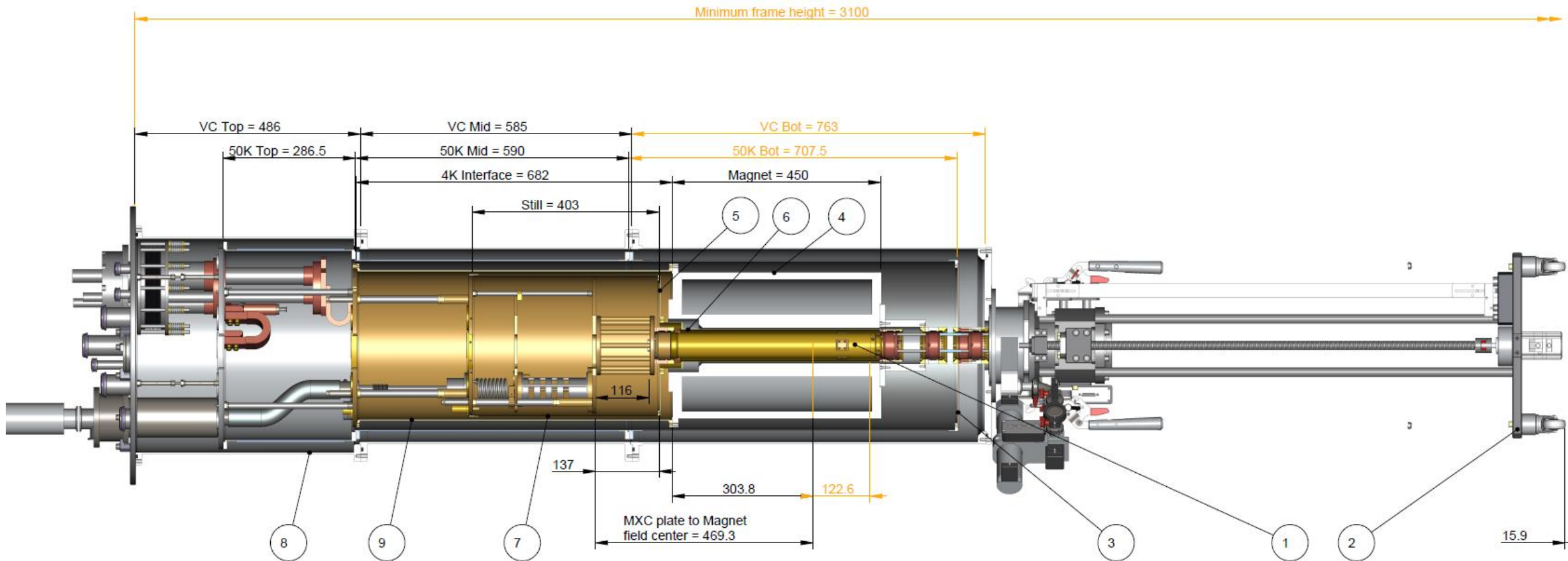
Other R&D activities - MPP-Munich subgroup

- Cylindrical cavity with a frequency tuning system based on the vertical cut concept.
- System for the coaxial cable to be inserted more or less to adjust the beta coupling of the system. Similar to CAPP cavities. →
- We are going to manufacture a new cavity version with oxygen-free copper material ($\uparrow\sigma$, $\uparrow Q_0$).
- Installation of the LD-250 Bluefors dilfridge system (arrives on October/2024).
- Nano-TES studies.
- Preparing data taking protocol for SM18 axion data taking in November/2024.



Other R&D activities - MPP-Munich subgroup

➤ LD-250 Bluefors dilfridge system

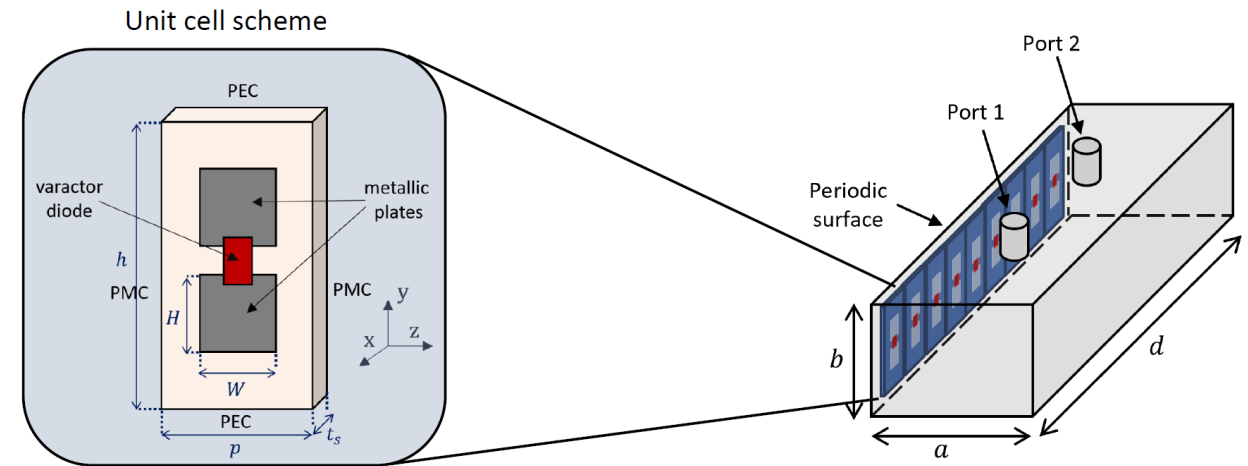
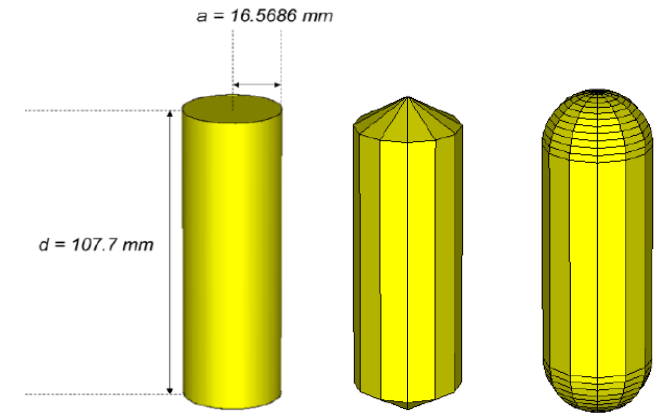


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Other R&D activities - UPCT/UV - Cartagena/Valencia subgroup

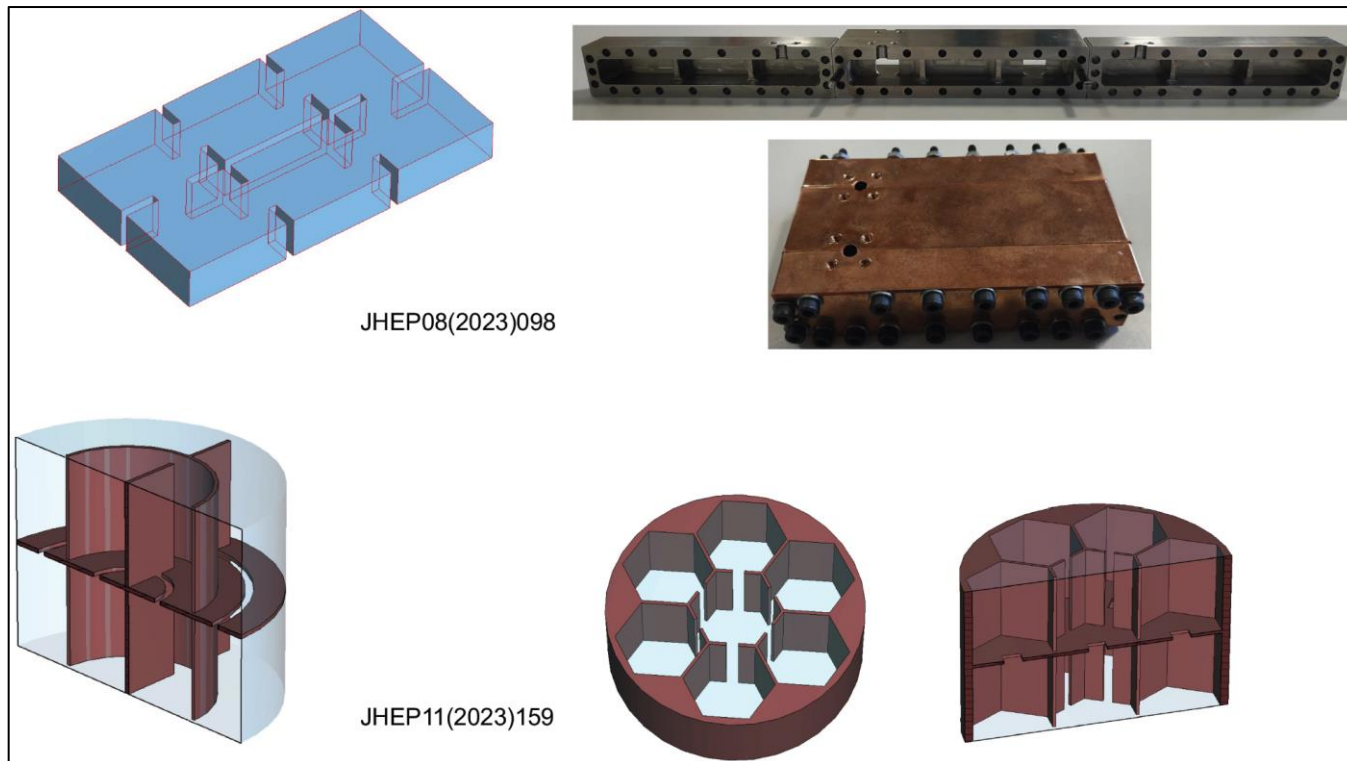
- BabyIAXO haloscopes
- HTS cavities
- Electrical tuning with periodic surfaces based on varactor diodes.
- Ferroelectric and metamaterial studies.
- Other cavity designs to increase V and Q_0



Study of tuning in a 8 GHz (for TE₁₀₁) cavity

Other R&D activities - UPCT/UV - Cartagena/Valencia subgroup

- Rectangular and cylindrical 1D, 2D, and 3D multicavities →



Other R&D activities - UPCT/UV - Cartagena/Valencia subgroup

- Cavity for the detection of GWs at microwave frequencies.

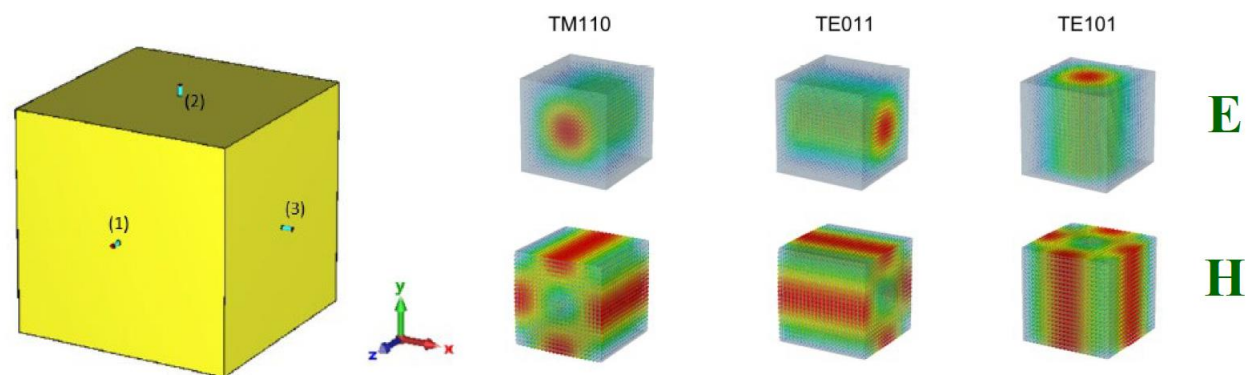
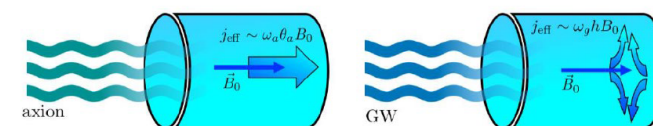
- Ideas and techniques developed for dark matter axions search be adapted to HF-GWs:

“Detecting high-frequency gravitational waves with microwave cavities”,
Asher Berlin et al.,
Physical Review D,
105, 116011 (2022), pp 1-23

Detecting High-Frequency Gravitational Waves with Microwave Cavities

Asher Berlin,^{1,2,3} Diego Blas,^{4,5} Raffaele Tito D’Agnolo,⁶ Sebastian A. R. Ellis,^{7,6}
Roni Harnik,^{2,3} Yonatan Kahn,^{8,9,3} and Jan Schütte-Engel^{8,9,3}

See also Herman, Füzfa, Lehoucq, Clesse, 2012.12189

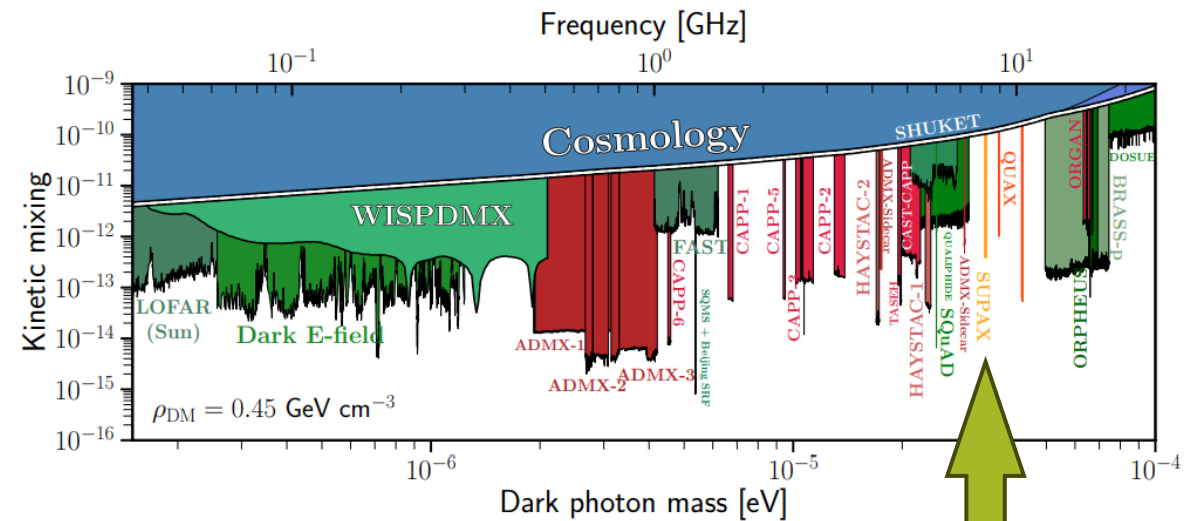
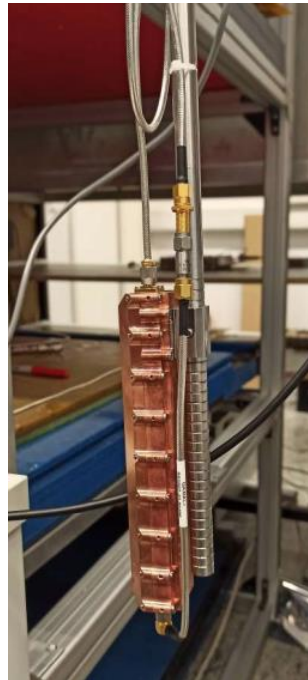
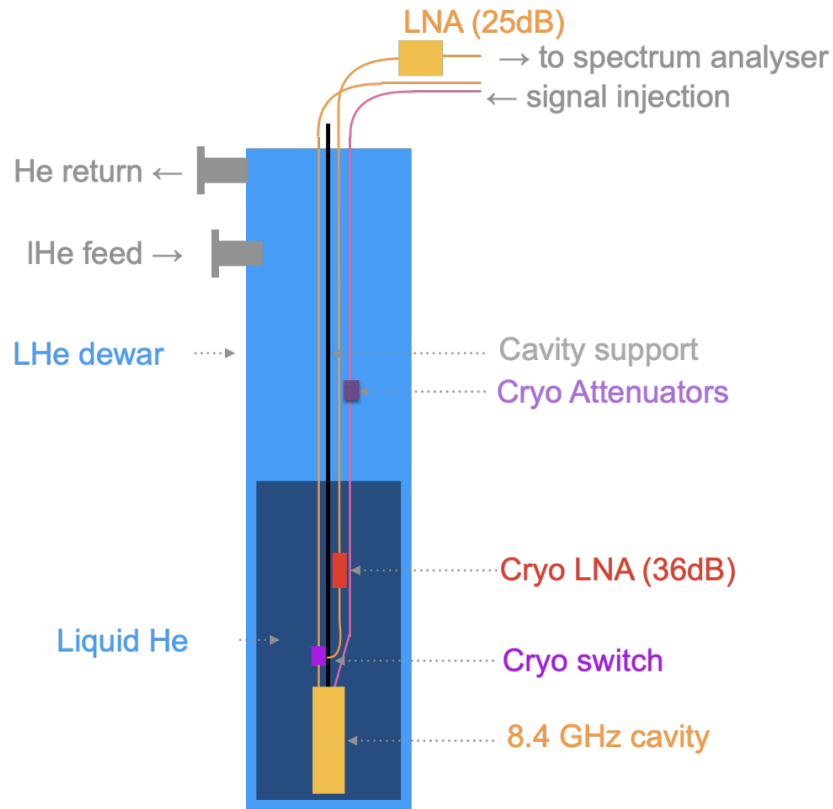


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Other R&D activities - UJG-M - Mainz subgroup

- In Mainz (Germany), the SUPAX experiment is being carried out.
- First results of a new cavity-based haloscope searching for dark photons with masses around $34 \mu\text{eV}$ employing a 8.3 GHz copper cavity at 4K temperatures (see [arXiv:2308.08337](https://arxiv.org/abs/2308.08337)).



Thank you

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