

Back to the origins of brane- antibrane inflation

Gonzalo Villa de la Viña

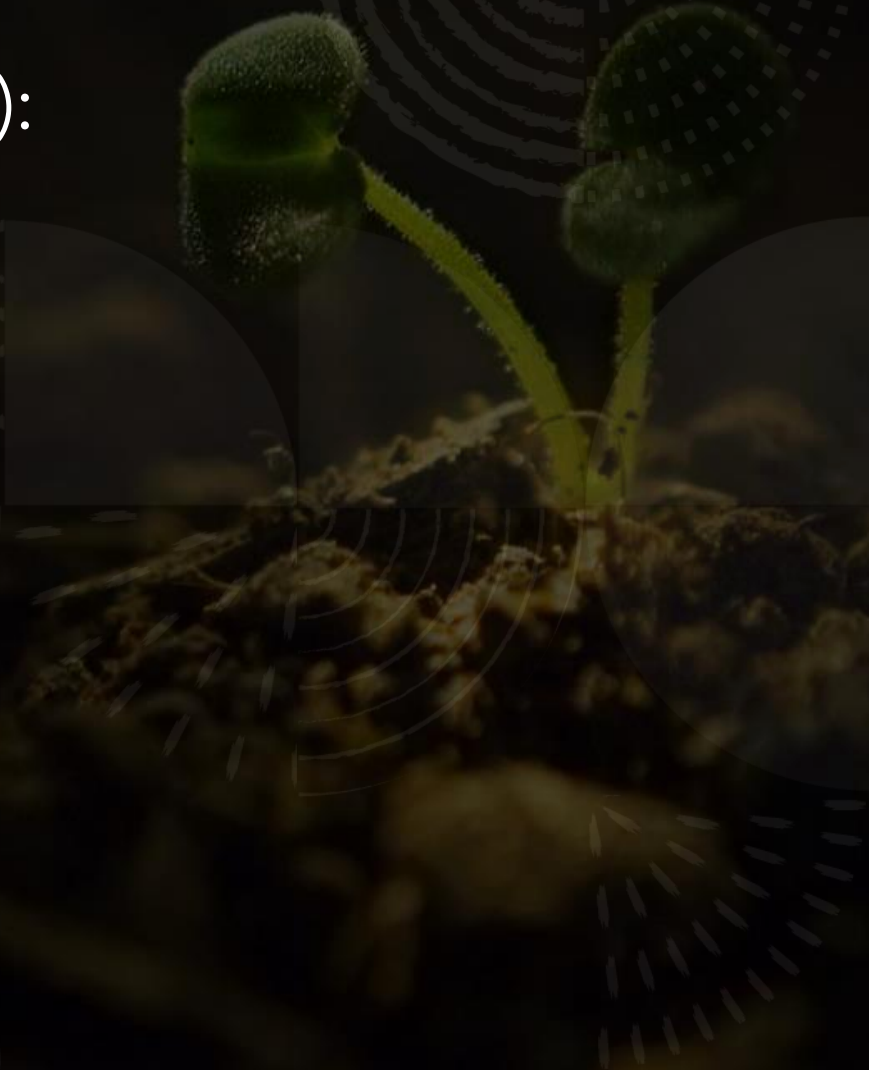
Based on 2409.xxxxx [hep-th]

With M. Cicoli, C. Hughes, A. R. Kamal, F. Marino, F. Quevedo and M. Ramos-Hamud

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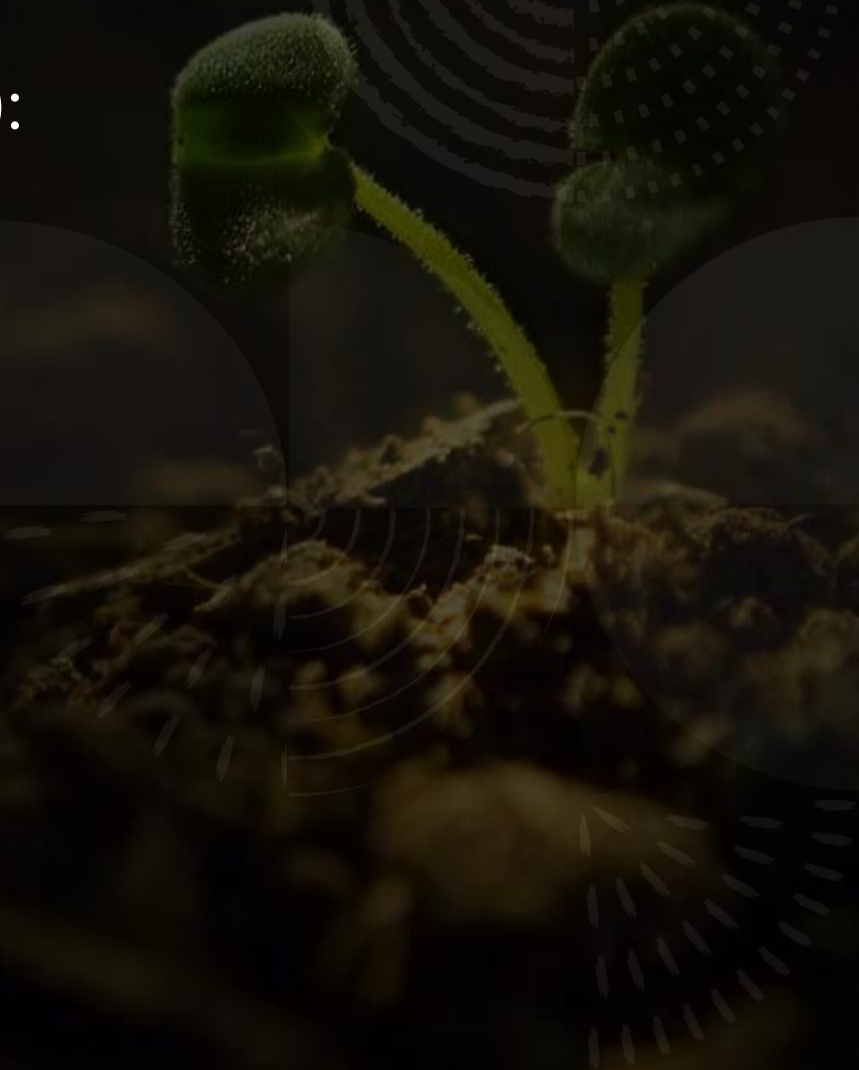
Inflation and the UV

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 - Shift symmetries
 - Scaling symmetries
 - SUSY



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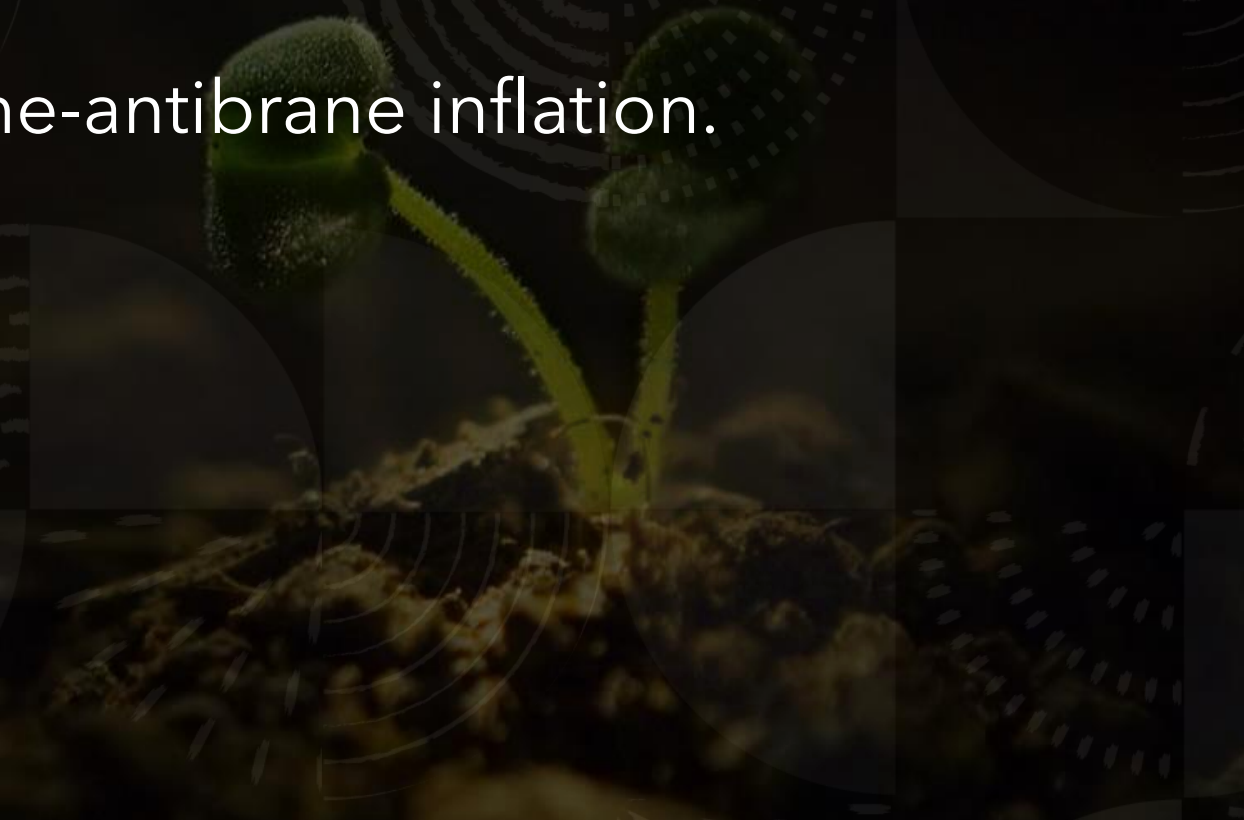
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- Inflation requires understanding of the UV!

Goals

- A consistent description of brane-antibrane inflation.



Dvali, Tye'99
Burgess, Majumdar, Nolte, Quevedo, Rajesh, Zhang'01
Dvali, Shafi, Solganik'01
Kachru, Kallosh, Linde, Maldacena, McAllister, Trivedi'03

Goals

- A consistent description of brane-antibrane inflation.
- Issues tackled in the paper:
 - Eta problem
 - EFT consistency
 - Experimental data
 - Late-time vacuum
- Today: eta problem

Dvali, Tye'99

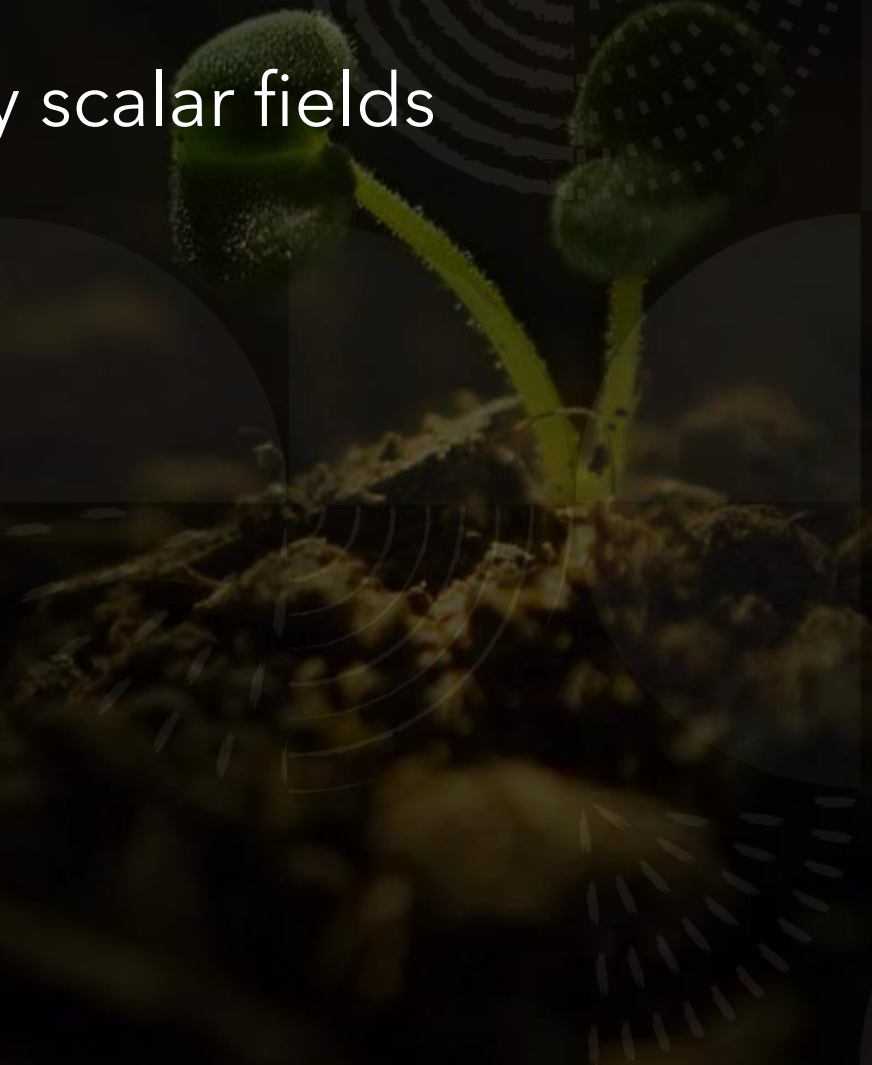
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Brane-antibrane potential

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$$V(\mathcal{V}, \varphi) = C_0 \left(1 - \frac{D_0}{\varphi^4} \right)$$

- When the antibrane sits in a *warped throat*:

$$C_0 \sim D_0 \sim \frac{e^{-8\pi K/g_s M}}{\mathcal{V}^{4/3}} \quad K, M \in \mathbb{Z} \quad \text{Naturally flat with } \mathcal{O}(1) \text{ inputs!}$$

Brane-antibrane inflation: pros and cons

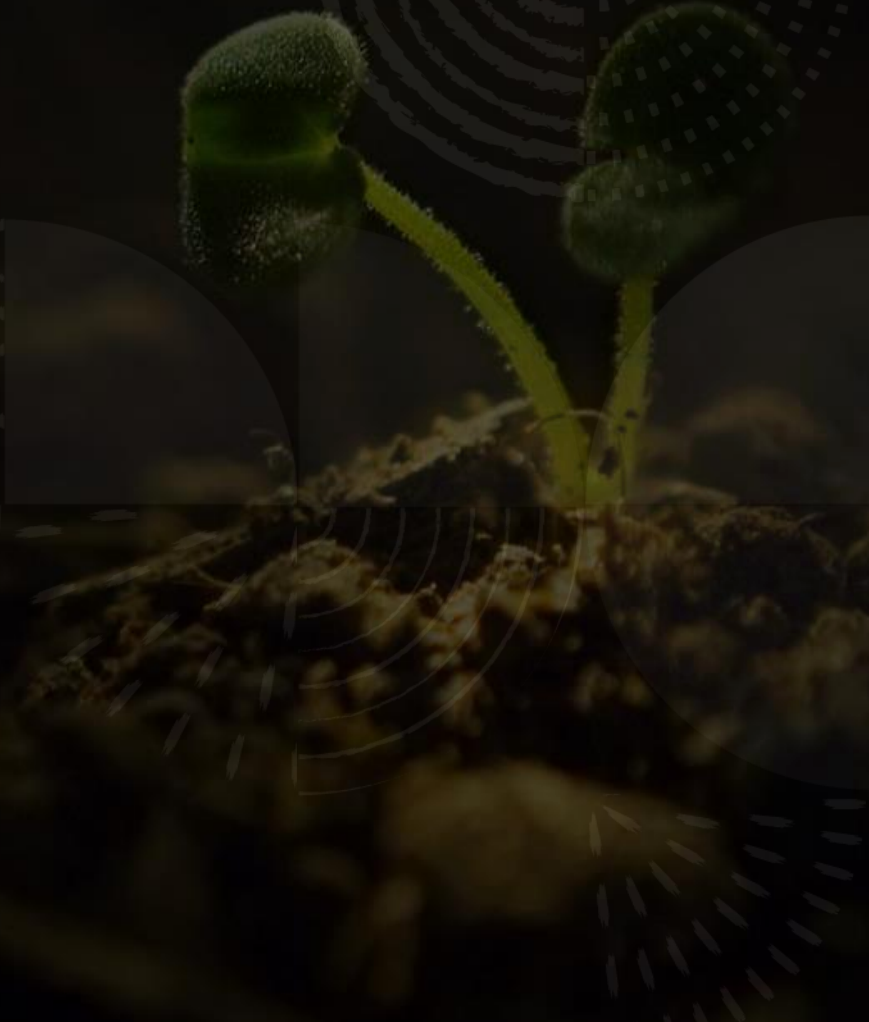


Brane-antibrane inflation: pros and cons

- Pros
 - Naturally stringy ingredients.
 - Order one tuning renders very flat potentials
 - Favoured by observations.
 - Plethora of predictions and stringy features after inflation.

Kofman, Yi'05

Frey, Mazumdar, Myers'05



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- Naturally stringy ingredients.
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• Cons

- Antibrane breaks SUSY.
- Need to stabilize the WISP by excellence in string compactifications: the volume modulus.
- Eta problem.

Kofman, Yi'05

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WISPS in inflation: the volume

- Life is difficult because generically there is at least one further scalar: the volume modulus.



WISPS in inflation: the volume

- Life is difficult because generically there is at least one further scalar: the volume modulus.
- Its potential is generically steep.
- So inflation can never occur far from a minimum of the potential in the volume direction.
- Need to find this minimum (*moduli stabilization*).

Kachru, Kallosh, Linde, Trivedi'03
Balasubramanian, Berglund, Conlon, Quevedo'05

The eta problem

- The volume receives a correction due to brane backreaction:

$$V^{2/3} = \rho - \bar{\varphi}\varphi$$

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KKLMMT'03

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- Which is a large mass term for the inflaton (eta problem)
- Solution: use corrections* that stabilize the whole quantity \mathcal{V} .

*For aficionados: perturbative corrections to the Kahler potential

Nonlinearly realized SUSY

- The antibrane breaks SUSY but its effects can be accommodated into an effective SUSY description.

Nonlinearly realized SUSY

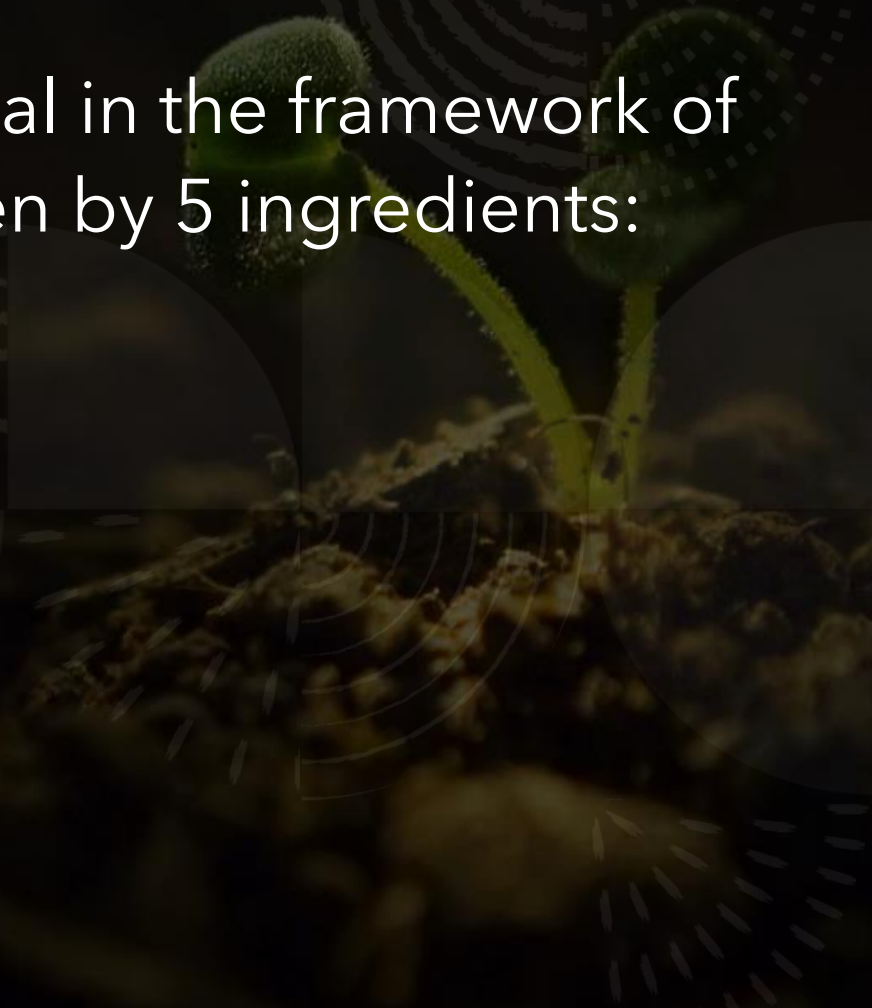
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Nonlinearly realized SUSY

- The antibrane breaks SUSY but its effects can be accommodated into an effective SUSY description.
- A description in terms of SUSY variables allows for a systematic classification of possible corrections.
- We discover a new perfect square structure in the scalar potential which stabilises the volume.

Scalar potential: ingredients

- The (perturbative) scalar potential in the framework of nonlinearly realized SUSY is given by 5 ingredients:



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 - Three functions of the volume

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- A (for current purposes) tunable constant W_0
- A function of φ , but not \mathcal{V} which contains the Coulomb potential W_X

Scalar potential and stabilization

- The most general scalar perturbative potential in the framework of nonlinearly realized SUSY reads*:

$$V = \frac{1}{U} \left[(f'W_X - 3g'W_0)^2 - f'' (fW_X^2 - 6gW_XW_0 - 9hW_0^2) \right]$$

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- Tree level reproduces standard uplift.
- $W_X=g=0$ reproduce standard breaking of no-scale structure (f'' is BBHL)

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Controlled SUGRA for small W_X !!

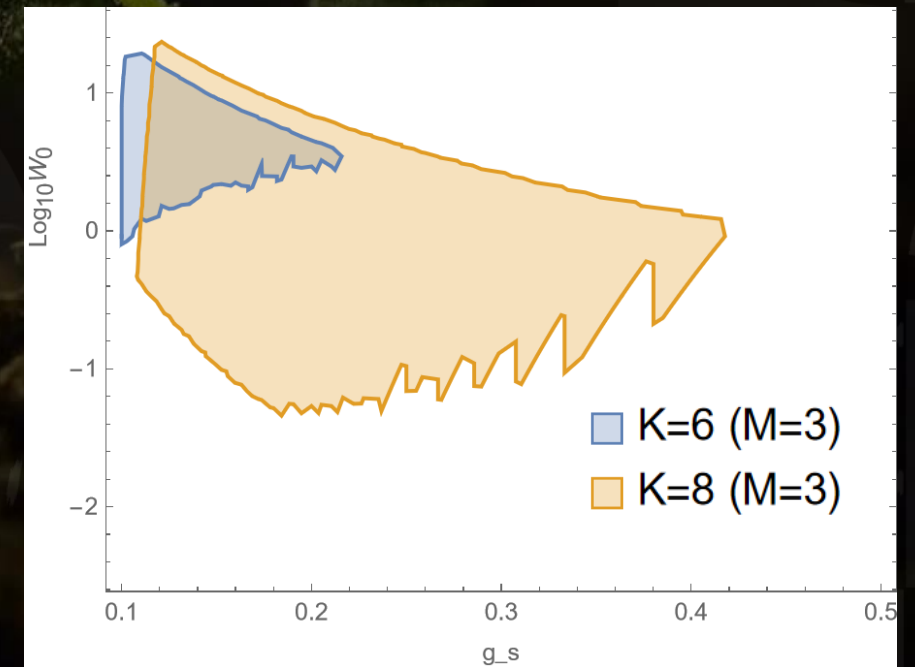
Brane-antibrane inflation reloaded!

- If $g' \gg f''$ (typically BBHL), then the volume is integrated leaving a Coulomb potential*.
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*In the paper we also propose a stabilization mechanism in the opposite case

Brane-antibrane inflation reloaded!

- If $g' \gg f''$ (typically BBHL), then the volume is integrated leaving a Coulomb potential*.
- Brane-antibrane inflation *can now occur*.
- After EFT and experimental constraints,
for $g \sim \log[V]$:



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Conclusions and future directions

- Perturbative corrections to the Kahler potential allow for brane-antibrane inflation.
- Nonlinearly realized SUSY uncovers a perfect-square structure in the scalar potential.
- In this talk, g plays a prominent role. It is a kinetic mixing goldstino-volume. EFT arguments suggest it's there, the question is at what power in $1/V$.
- We find a parameter space of microscopic data where brane-antibrane inflation can take place.

Backup: Kahler and superpotential

- Most general* Kahler and superpotentials with constrained superfields $X^2=0$:

$$K = -3 \log[f(\mathcal{V}^{2/3}) + g(\mathcal{V}^{2/3})(X + \bar{X}) + h(\mathcal{V}^{2/3})X\bar{X}]$$

$$W = W_0 + XW_X$$

*Contributions outside of the log can always be absorbed up to a higher order correction