

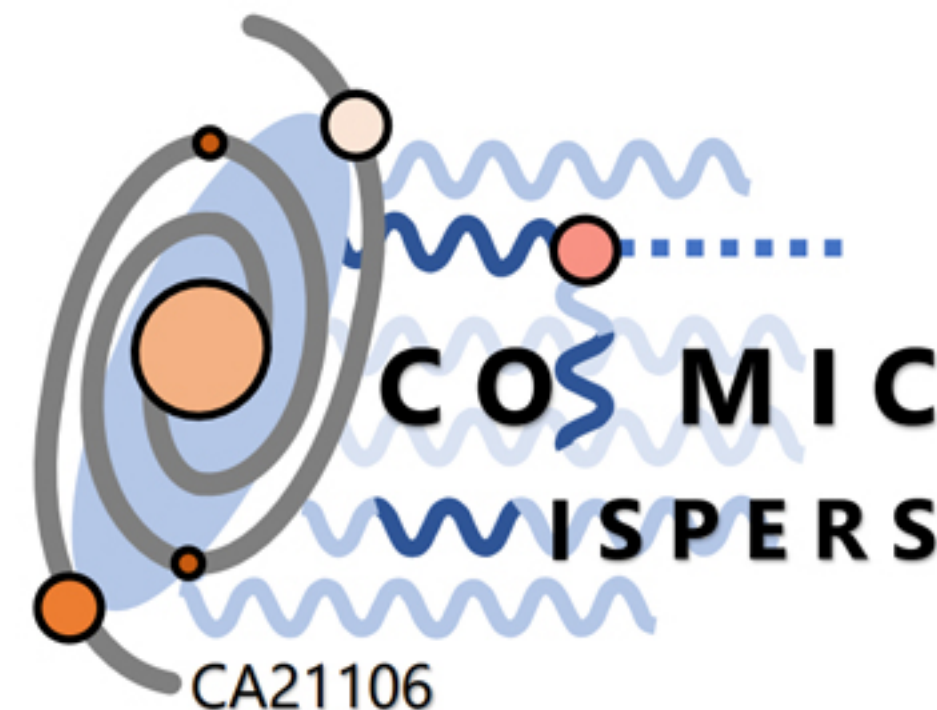
Search for ultralight particles from compact stars through gravitational wave telescopes

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INFN Salerno

Fundamental Physics and Gravitational Wave Detectors, Pollica

September 16, 2024



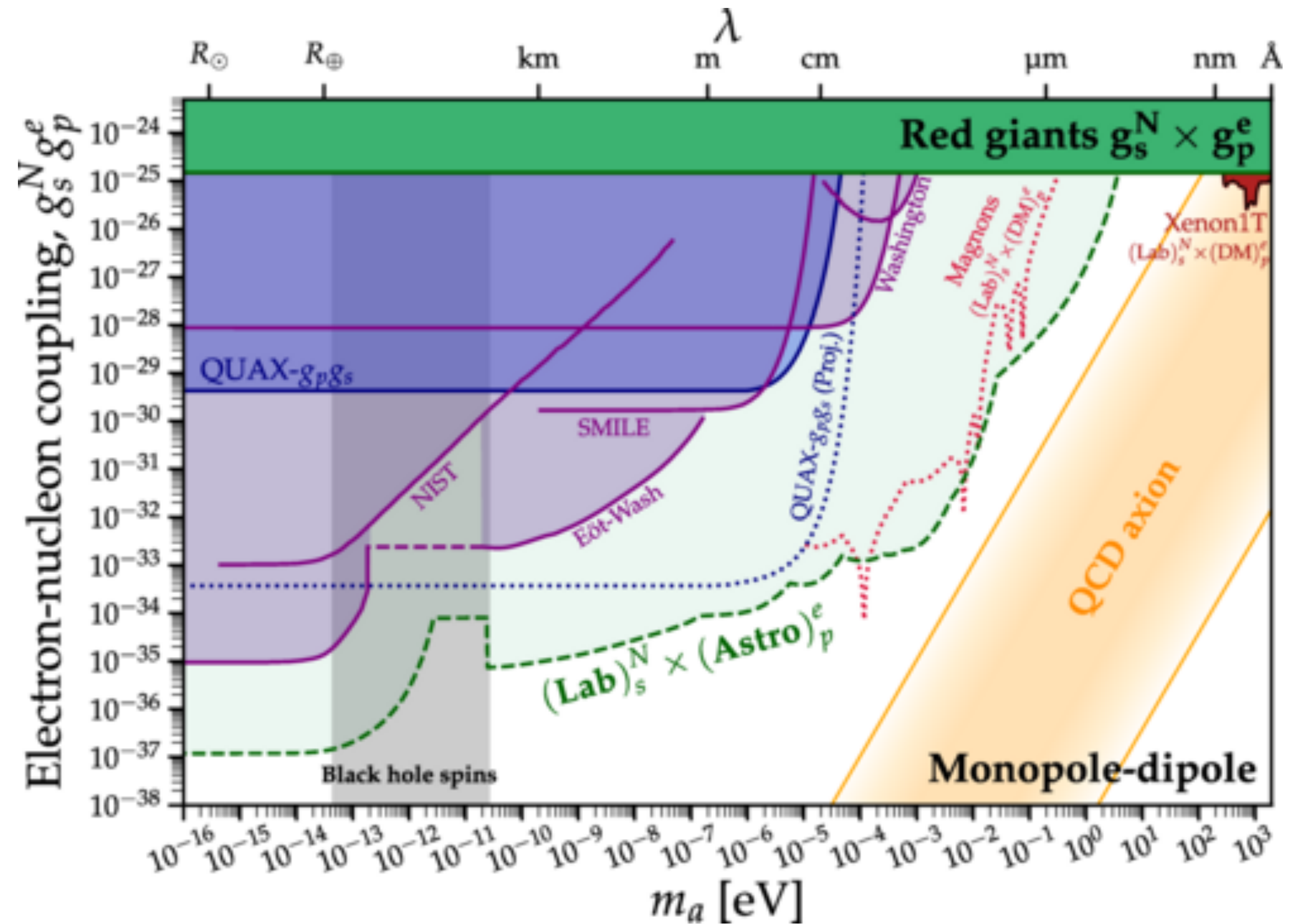
About Me !

PhD @ Physical Research Laboratory, India, 2022

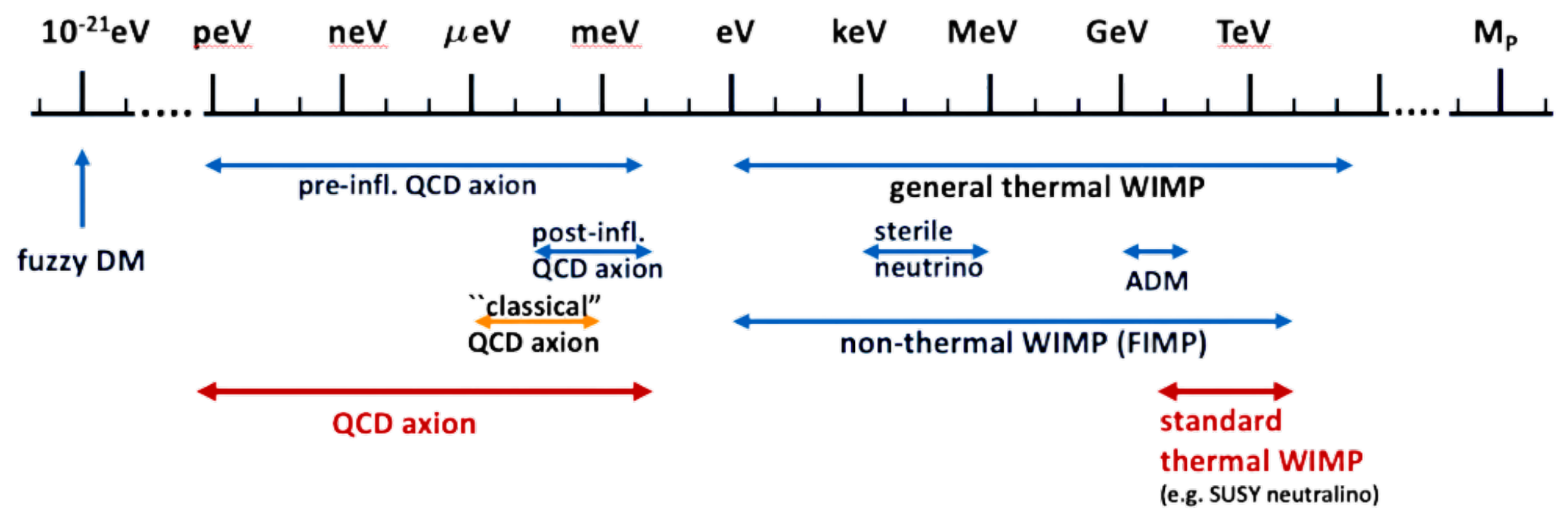
Postdoc @ Tata Institute of Fundamental Research (TIFR), India, 2023

Postdoc @ INFN, Salerno, Italy, 2024

Associated with QUAX gs-gp collaboration

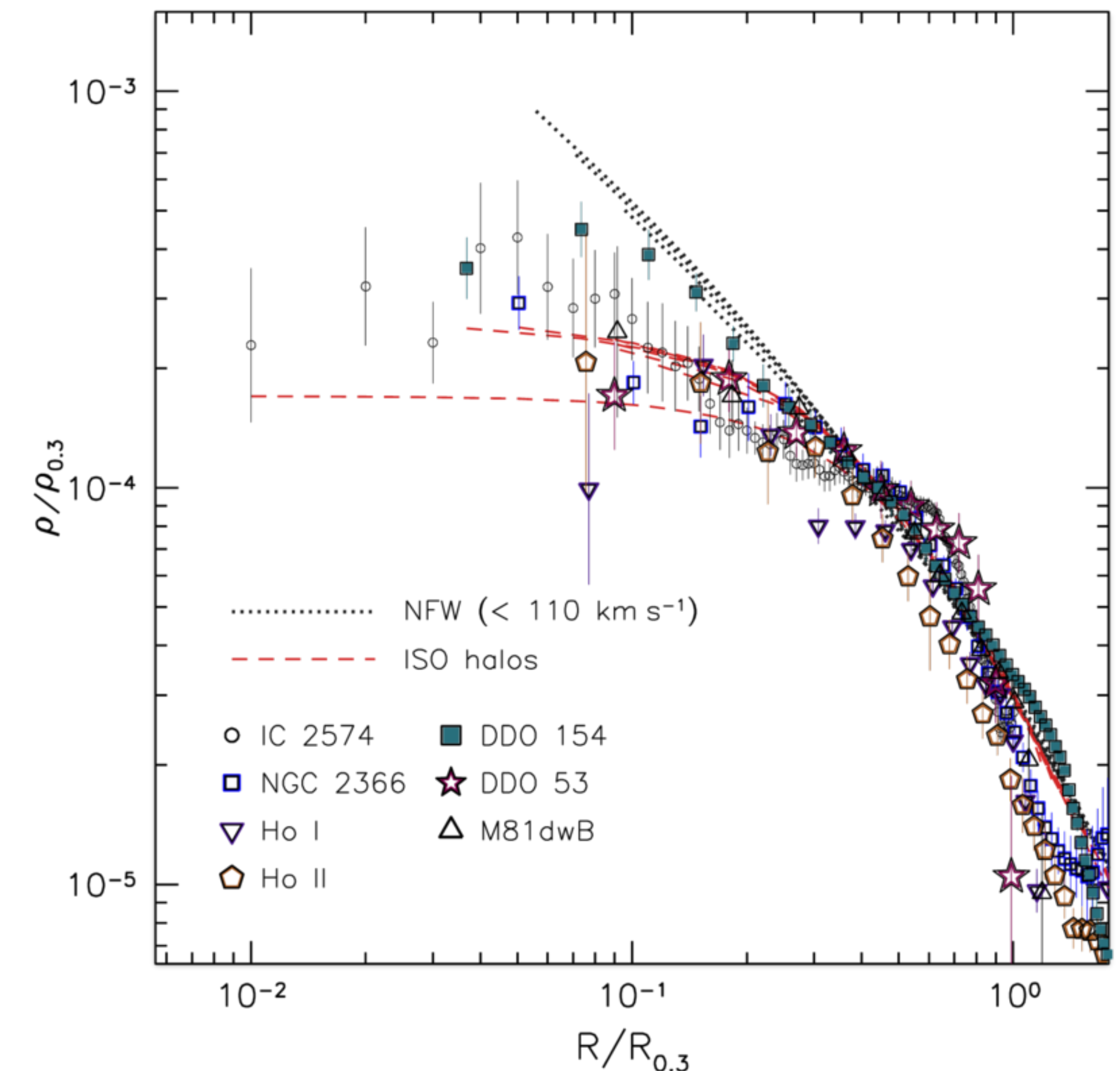
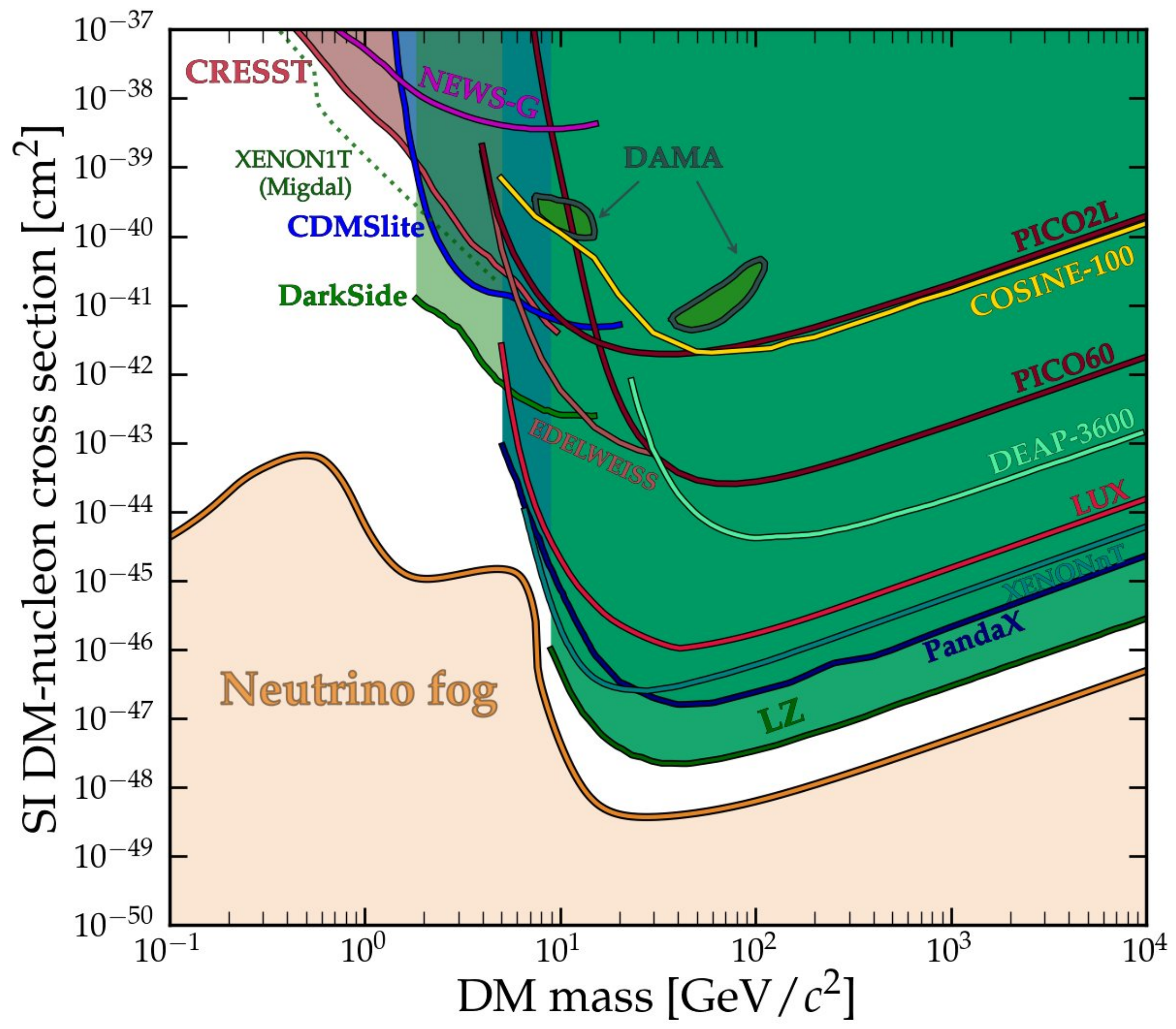


DM: Motivations

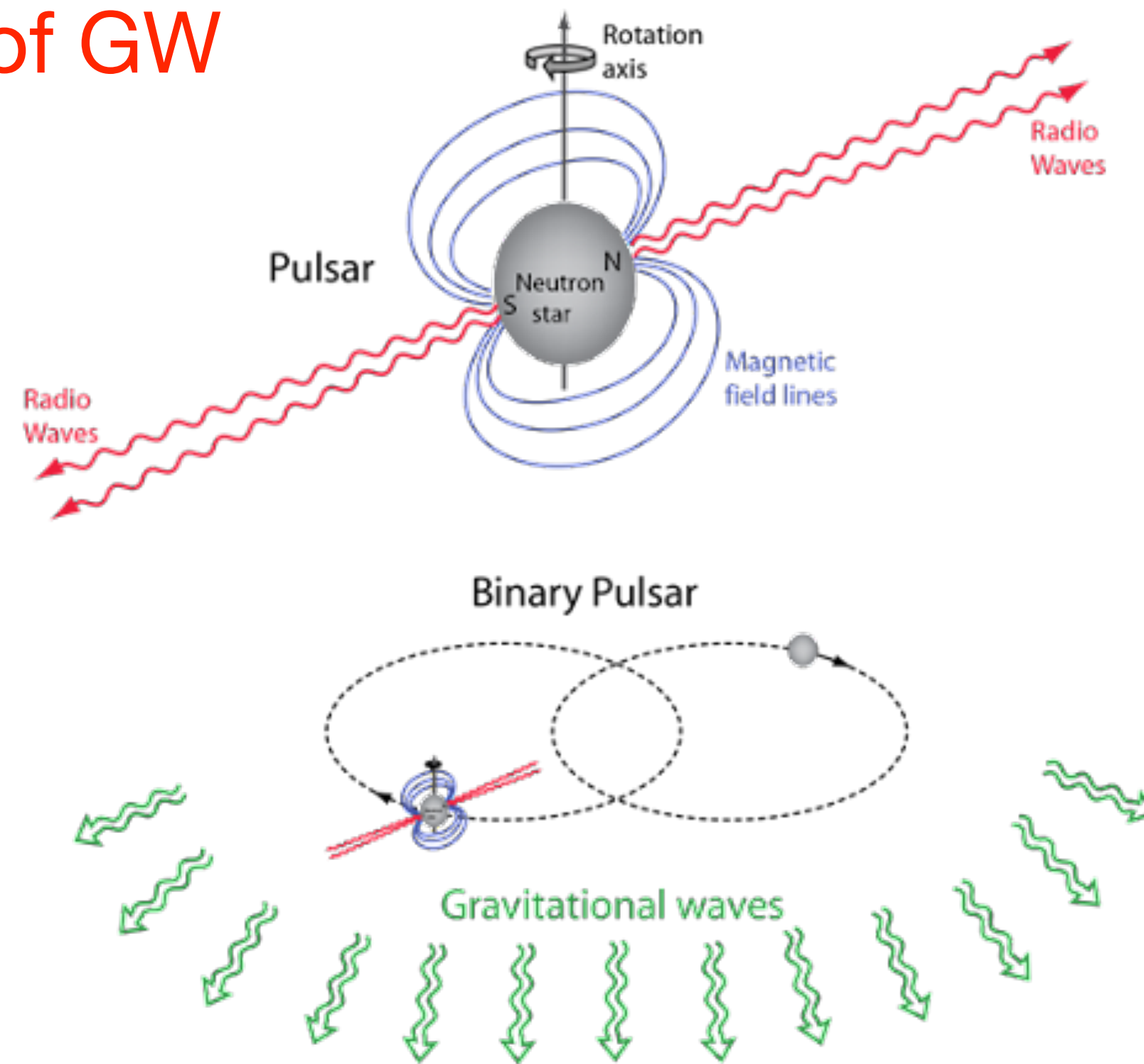


LZ collaboration, 2024

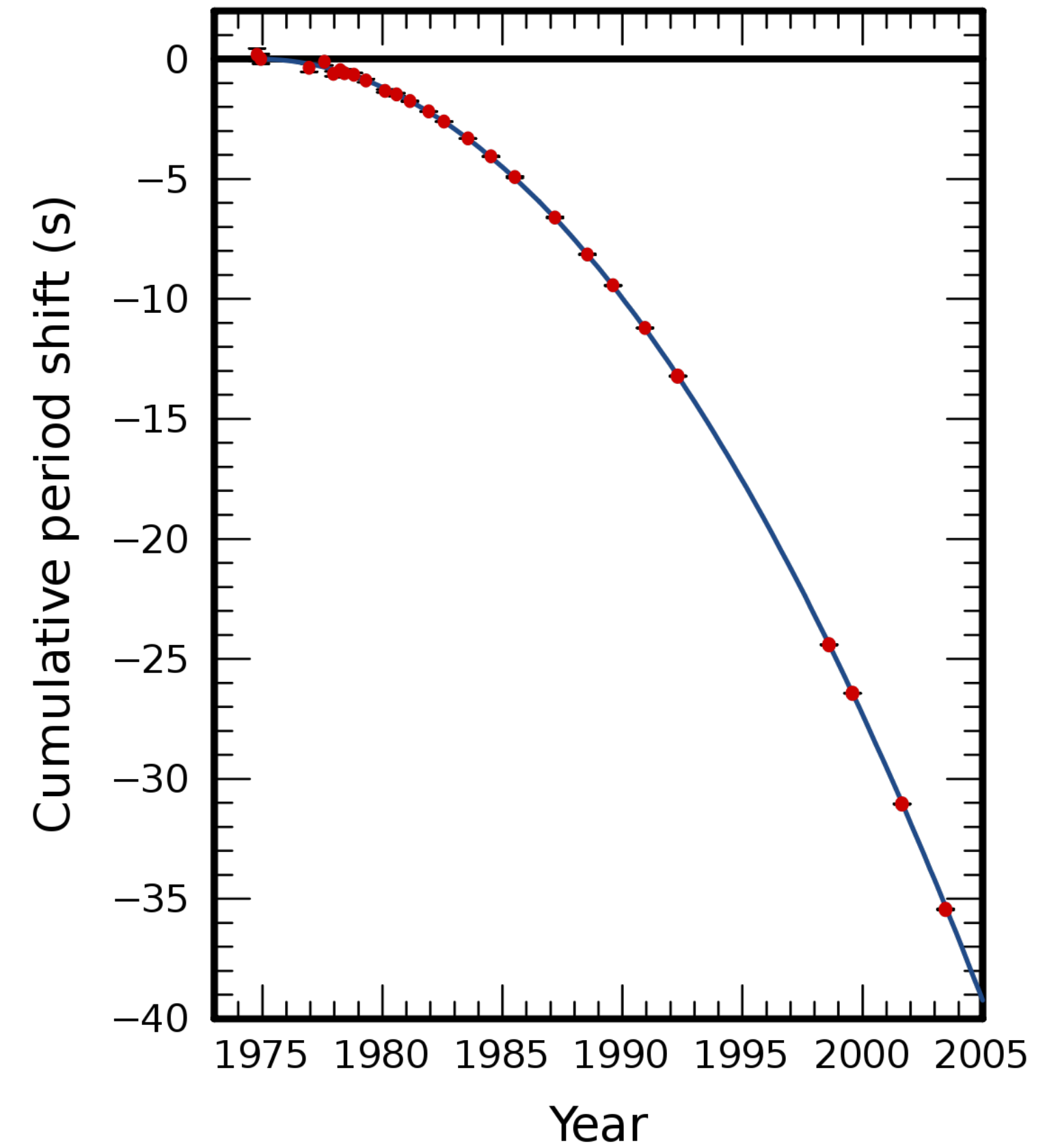
Astron J, 2011



Indirect detection of GW



Weisberg, Taylor, ASP conference series , vol 328, 2005



$$P = \frac{G}{5} \left(\frac{d^3 Q_{ij}}{dt^3} \frac{d^3 Q_{ij}}{dt^3} - \frac{1}{3} \frac{d^3 Q_{ii}}{dt^3} \frac{d^3 Q_{jj}}{dt^3} \right)$$

Peters, Mathews, Phys. Rev, 1963

$$\frac{dE_{GW}}{dt} = \frac{32G}{5} \Omega^6 \left(\frac{m_1 m_2}{m_1 + m_2} \right)^2 a^4 (1 - e^2)^{-7/2} \left(1 + \frac{73}{24} e^2 + \frac{37}{96} e^4 \right)$$

$$\dot{P}_b = 6\pi G^{-3/2} (m_1 m_2)^{-1} (m_1 + m_2)^{-1/2} a^5 \left(\frac{dE}{dt} \right)$$

Spin-1

Ultralight $L_\mu - L_\tau$ vector

Constraining particle physics models

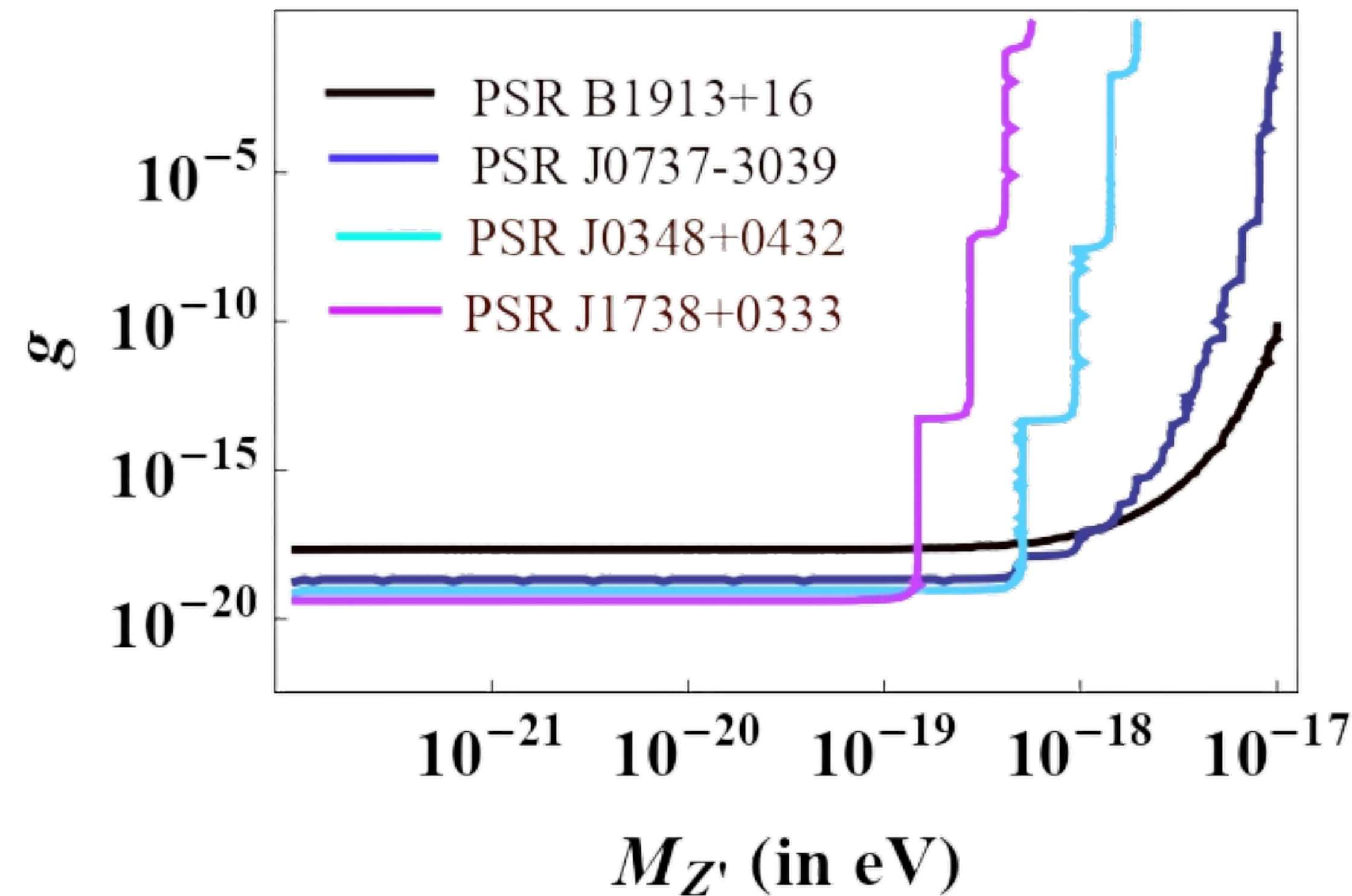
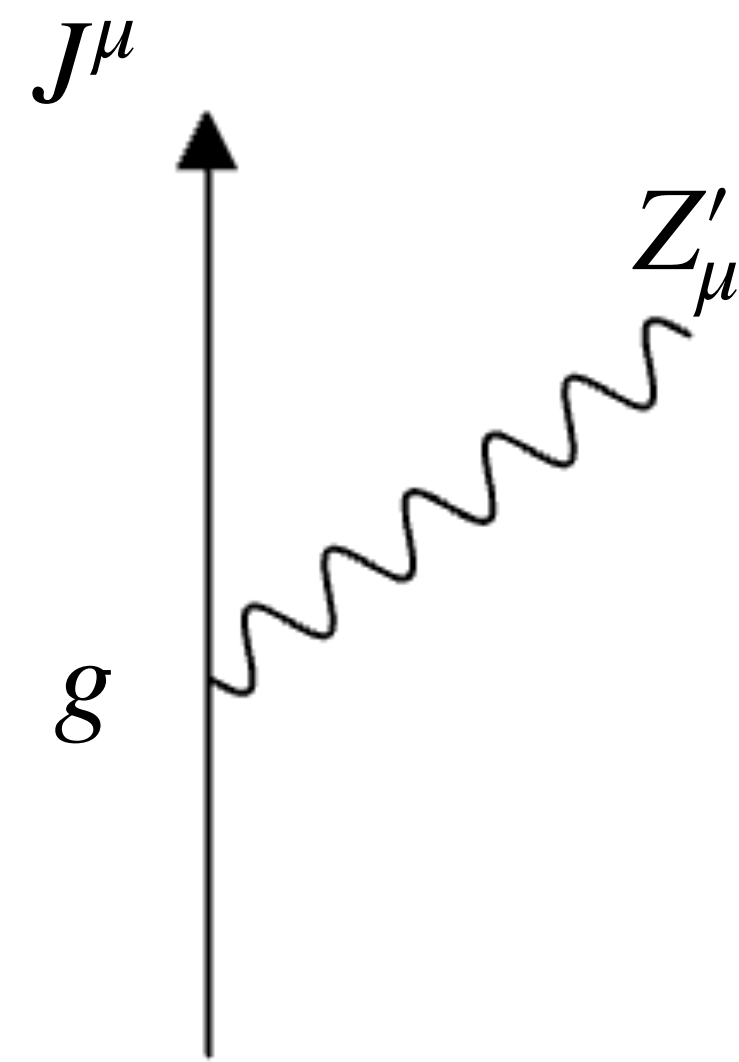
HT: $P \sim 8 \text{ h} \rightarrow \Omega \sim 10^{-19} \text{ eV}$

$Q_\mu \approx 10^{55}$

Garani, Heeck, Phys. Rev. D, 2019

$$\mathcal{L} \supset g' Z'_\mu (\bar{\mu} \gamma^\mu \mu - \bar{\tau} \gamma^\mu \tau + \bar{\nu}_\mu \gamma^\mu L \nu_\mu - \bar{\nu}_\tau \gamma^\mu L \nu_\tau)$$

$$\frac{dE_V}{dt} = \frac{g^2}{6\pi} a^2 M^2 \left(\frac{Q_1}{m_1} - \frac{Q_2}{m_2} \right)^2 \Omega^4 \sum_{n>n_0} 2n^2 \left[J_n^2(ne) + \frac{(1-e^2)}{e^2} J_n^2(ne) \right] \left(1 - \frac{n_0^2}{n^2} \right)^{\frac{1}{2}} \left(1 + \frac{1}{2} \frac{n_0^2}{n^2} \right) \quad n_0 = \frac{M_{Z'}}{\Omega}$$

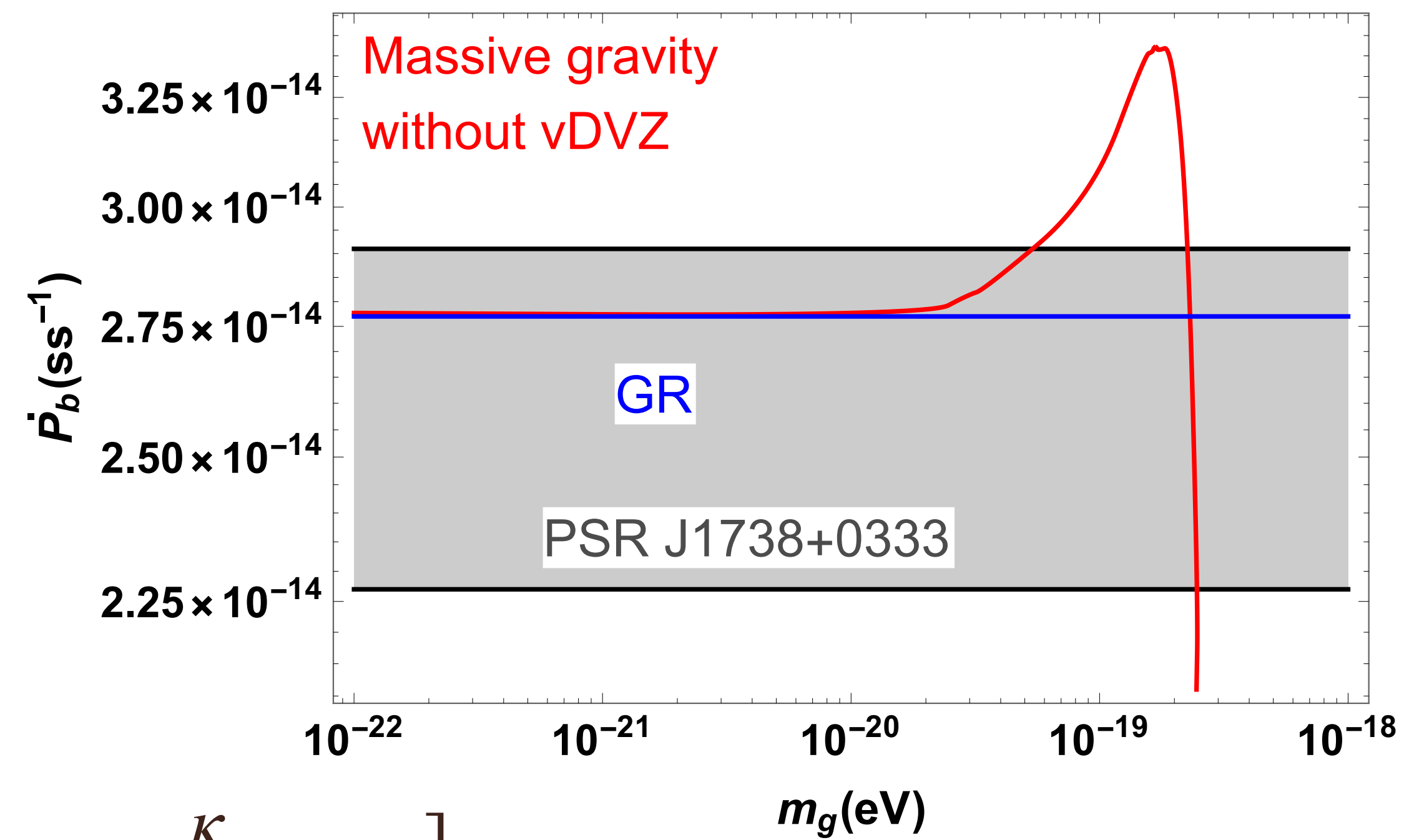
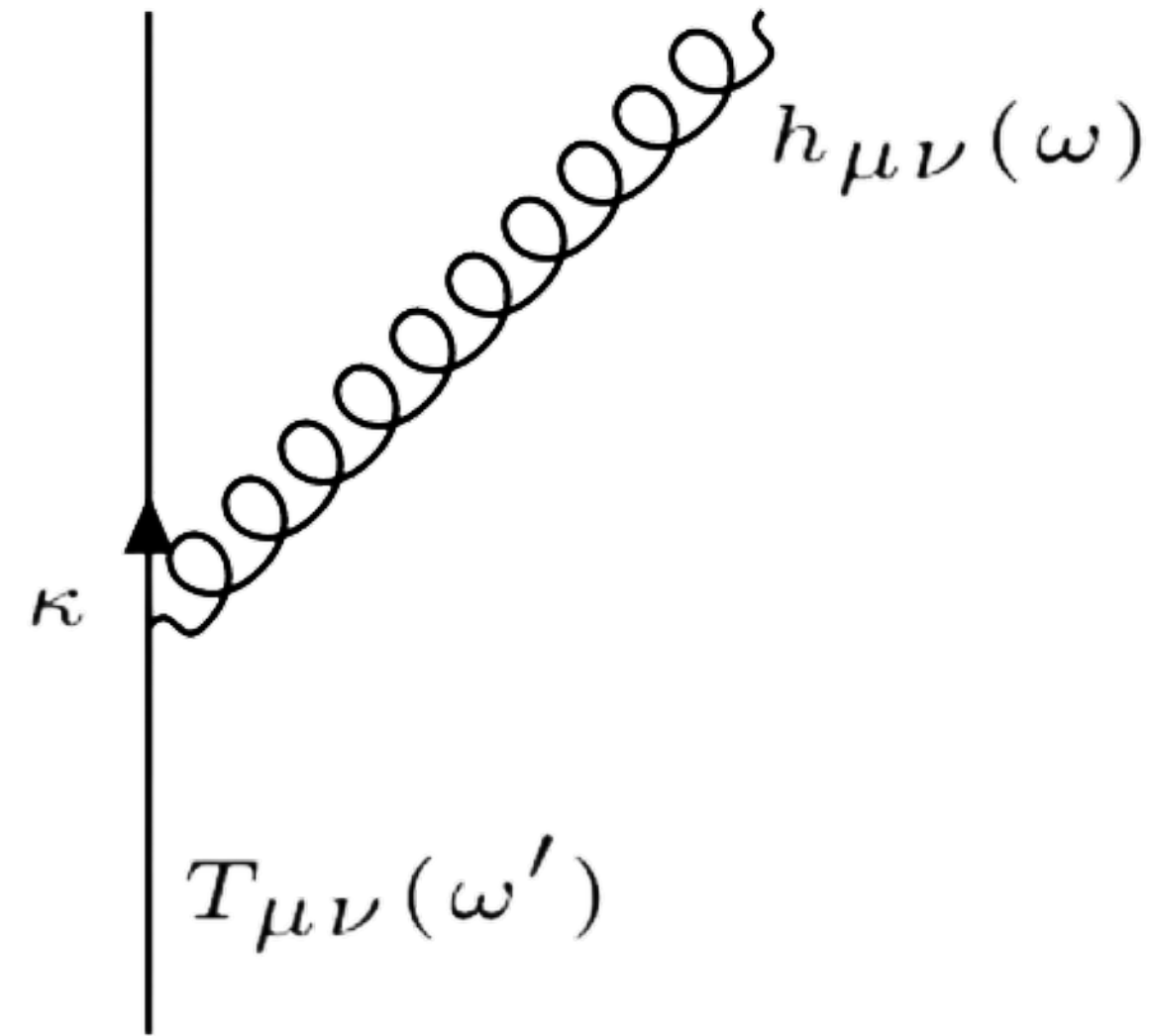


T.K.P, S. Mohanty, S. Jana, Phys. Rev. D, 2019

Spin-2

Massive graviton

$$V^{(1/2)}(r) = \frac{GM_1M_2}{r} e^{-m_g r}$$



$$S = \int d^4x \left[\frac{1}{2} h_{\mu\nu} \epsilon^{\mu\nu\alpha\beta} h_{\alpha\beta} + \frac{1}{2} m_g^2 h_{\mu\nu} (\eta^{\mu(\alpha} \eta^{\beta)\nu} - (1-a) \eta^{\mu\nu} \eta^{\alpha\beta}) h_{\alpha\beta} + \frac{\kappa}{2} h_{\mu\nu} T^{\mu\nu} \right]$$

$$a = \frac{1}{2}$$

$$D_{\alpha\beta\mu\nu}^{1/2}(k) = \frac{1}{-k^2 + m_g^2} \left(\frac{1}{2} (\eta_{\alpha\mu} \eta_{\beta\nu} + \eta_{\alpha\nu} \eta_{\beta\mu} - \frac{1}{2} \eta_{\alpha\beta} \eta_{\mu\nu}) + (k - \text{dependent terms}) \right)$$

$$\frac{dE}{dt} = \frac{32G}{5} \mu^2 a^4 \Omega^6 \sum_{n=1}^{\infty} n^6 \sqrt{1 - \frac{n_0^2}{n^2} \left[f(n, e) \left(1 + \frac{4}{3} \frac{n_0^2}{n^2} + \frac{1}{6} \frac{n_0^4}{n^4} \right) - \frac{5J_n^2(ne)}{36n^4} \frac{n_0^2}{n^2} \left(1 - \frac{n_0^2}{4n^2} \right) \right]}$$

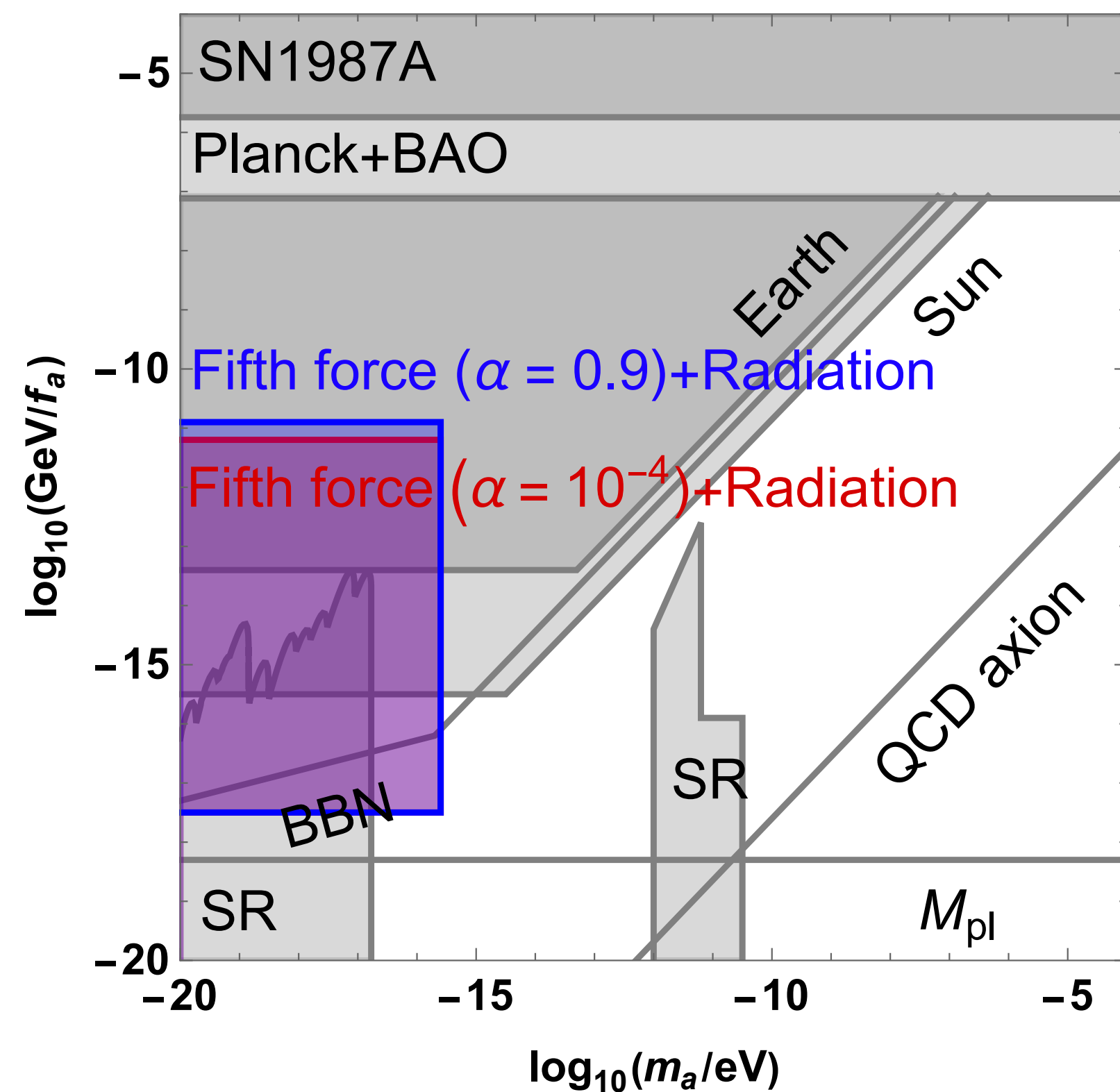
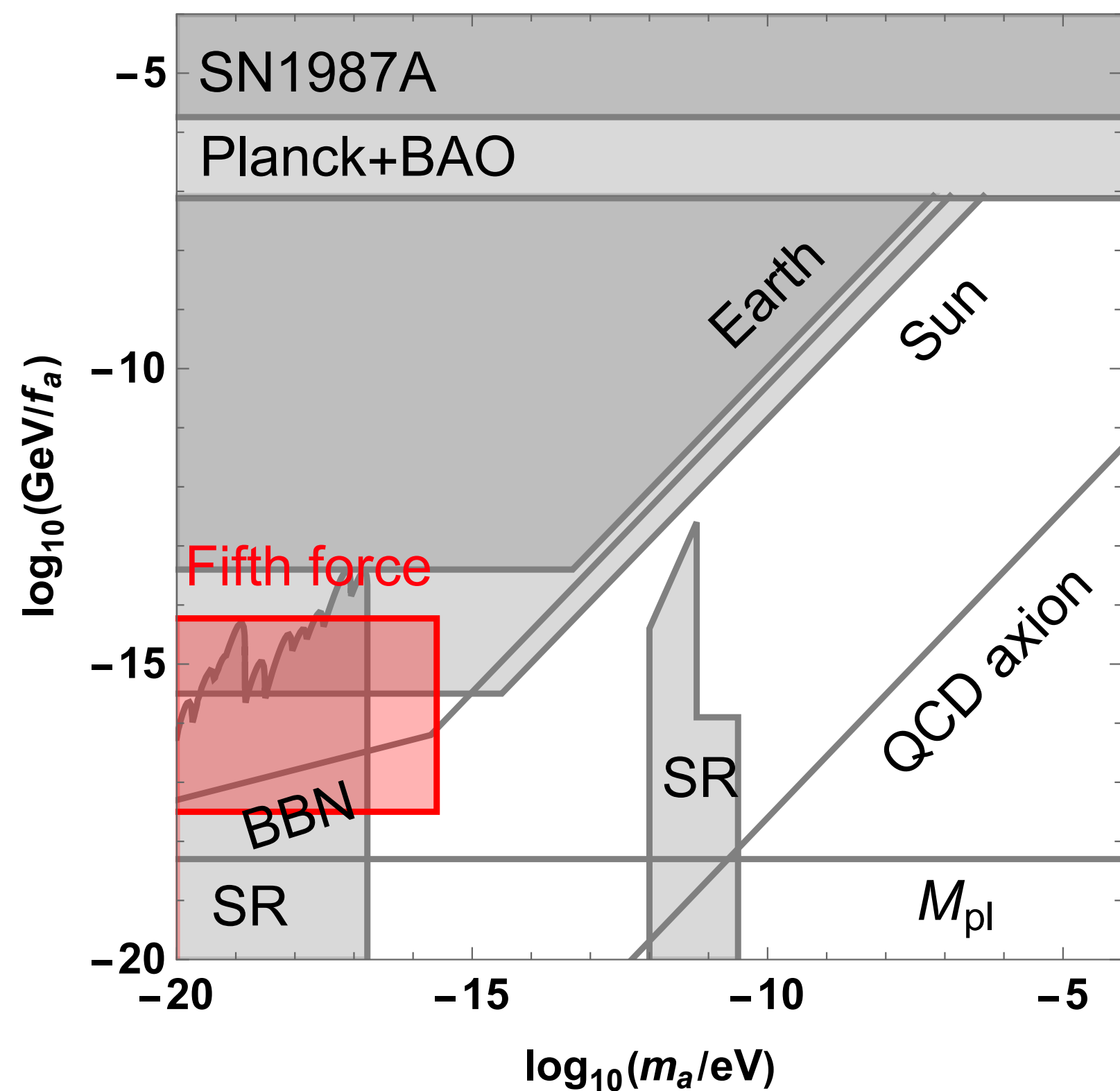
$$m_g < 1.81 \times 10^{-20} \text{ eV}$$

Spin-0

Axions

$$q_{eff} = -\frac{8\pi GMf_a}{\ln\left(1 - \frac{2GM}{r_{NS}}\right)} \quad p = 8\pi Gf_a\mu a \left[\frac{1}{\ln\left(1 - \frac{2Gm_2}{r_{NS}}\right)} - \frac{1}{\ln\left(1 - \frac{2Gm_1}{r_{NS}}\right)} \right] \quad F_5 = \frac{q_1q_2}{4\pi a^2}$$

$$\frac{dE}{dt} = -\frac{32}{5}G\mu^2a^4\omega^6(1 - e^2)^{-\frac{7}{2}}\left(1 + \frac{73}{24}e^2 + \frac{37}{96}e^4\right) - \frac{\omega^4p^2(1 + e^2/2)}{24\pi(1 - e^2)^{\frac{5}{2}}}$$



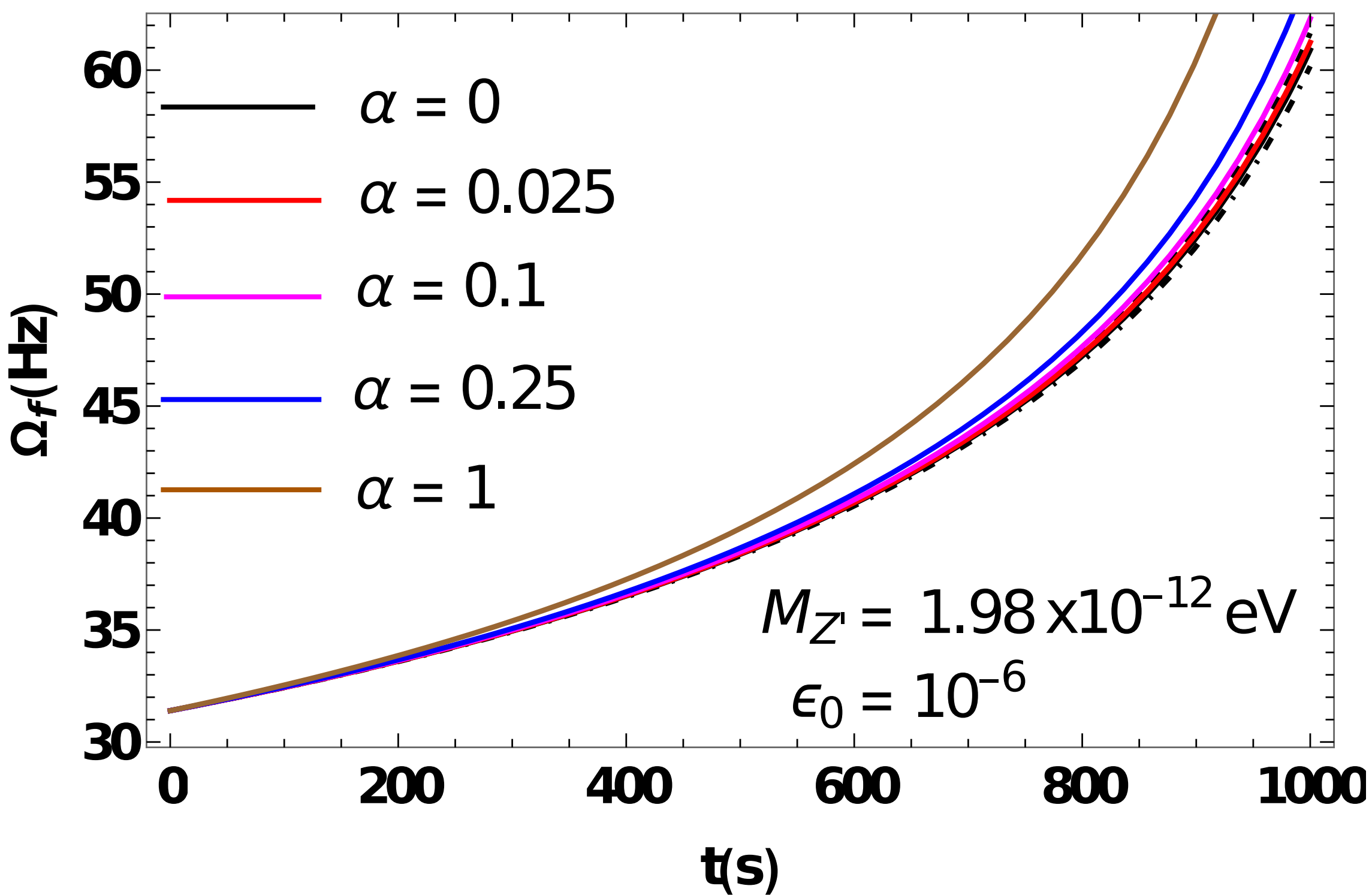
Direct detection: GW170817

See talk by Costantino pacilio

$$\frac{dr}{dt} = -\frac{64 G^3 M m (M + m)}{5 r^3} (1 - \epsilon^2)^{-7/2} \left(1 + \frac{73}{24} \epsilon^2 + \frac{37}{96} \epsilon^4 \right)$$

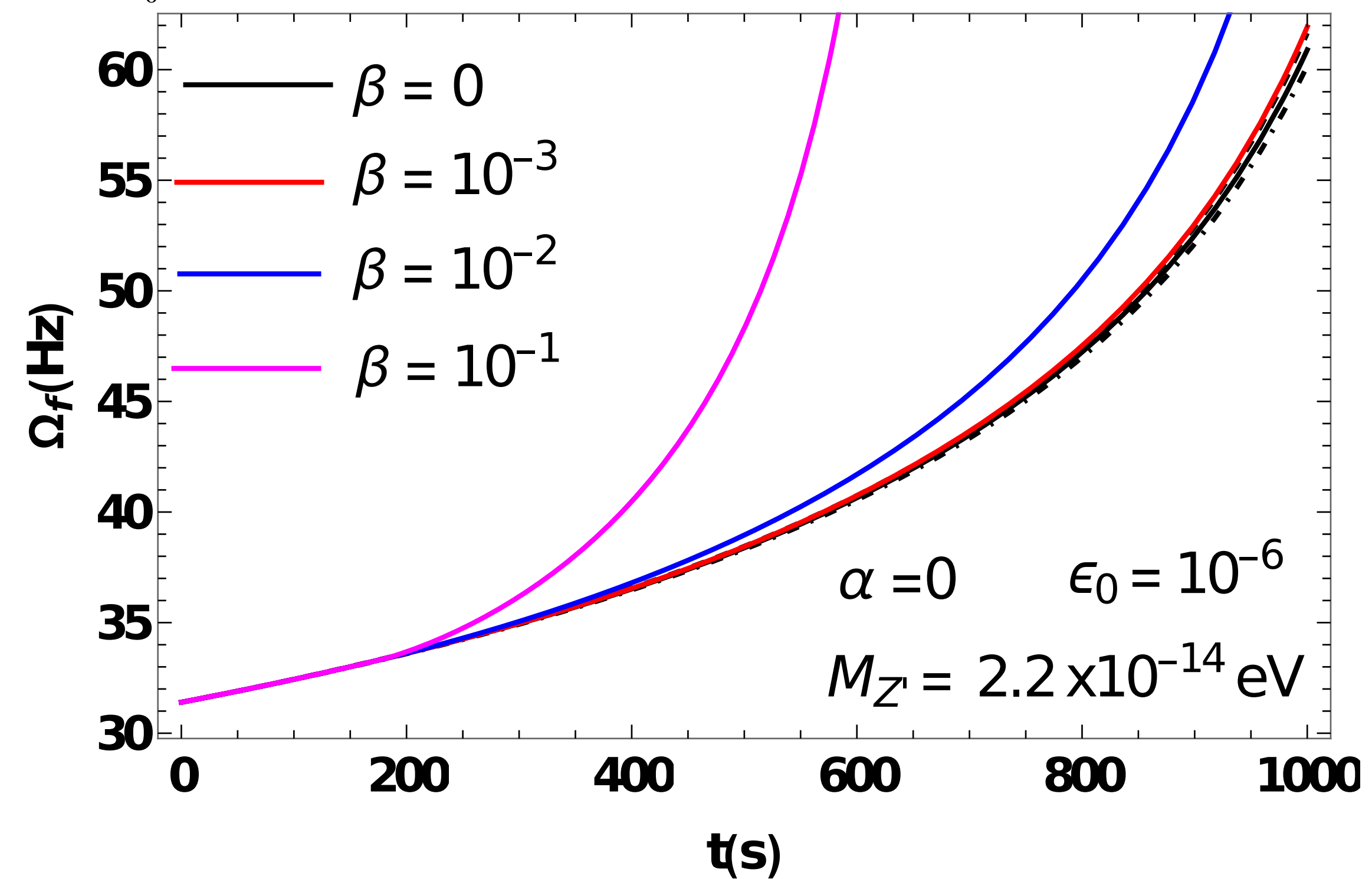
$$\beta = \frac{g_V^2 M m \left(\frac{Q}{M} - \frac{q}{m} \right)^2}{\pi G (M + m)^2} \quad \delta = \left(\frac{Q}{M} - \frac{q}{m} \right)^2$$

$$\dot{\Omega} = \frac{96}{5} G^{5/3} \mathcal{M}_{\text{ch}}^{5/3} \Omega^{11/3} (1 - \epsilon^2)^{-7/2} \left(1 + \frac{73}{24} \epsilon^2 + \frac{37}{96} \epsilon^4 \right) + \frac{g^2 \Omega^3 \mathcal{M}_{\text{ch}}^{5/3}}{2\pi (M + m)^{2/3}} \delta \times \sum_{n > n_0^V} 2n^2 \left[J_n'^2(n\epsilon) + \left(\frac{1 - \epsilon^2}{\epsilon^2} \right) J_n^2(n\epsilon) \right] \sqrt{1 - \frac{n_0^{V2}}{n^2} \left(1 + \frac{n_0^{V2}}{2n^2} \right)}$$



$$\dot{\Omega}_f = L_1 + L_2^{S/V} + L_3 + L_4 + L_5$$

$$L_1 = \frac{96}{5} (G \mathcal{M}_{\text{ch}})^{5/3} \Omega_f^{11/3} (1 - \epsilon^2)^{-7/2} \left(1 + \frac{73}{24} \epsilon^2 + \frac{37}{96} \epsilon^4 \right) (1 + \alpha)^{2/3}$$



Memory signal from pulsar kick

Kusenko, Segre PRL, 1996

Barkovich et al, Phys. Rev. D, 2002

$$Q = - \frac{g'_{ff} \frac{\sqrt{2\rho_{\text{DM}}}}{m_{A'}}}{V_{\nu_f} (h_p^{-1} + h_{V_{\nu_f}}^{-1})}$$

$$|h(t)| \simeq \frac{2G}{r} \alpha L_\nu \mathcal{T} \lesssim 1.06 \times 10^{-19} \left(\frac{|\alpha|}{0.01} \right) \left(\frac{E_{\text{tot}}}{2 \times 10^{53} \text{ erg}} \right) \left(\frac{1 \text{ kpc}}{r} \right)$$

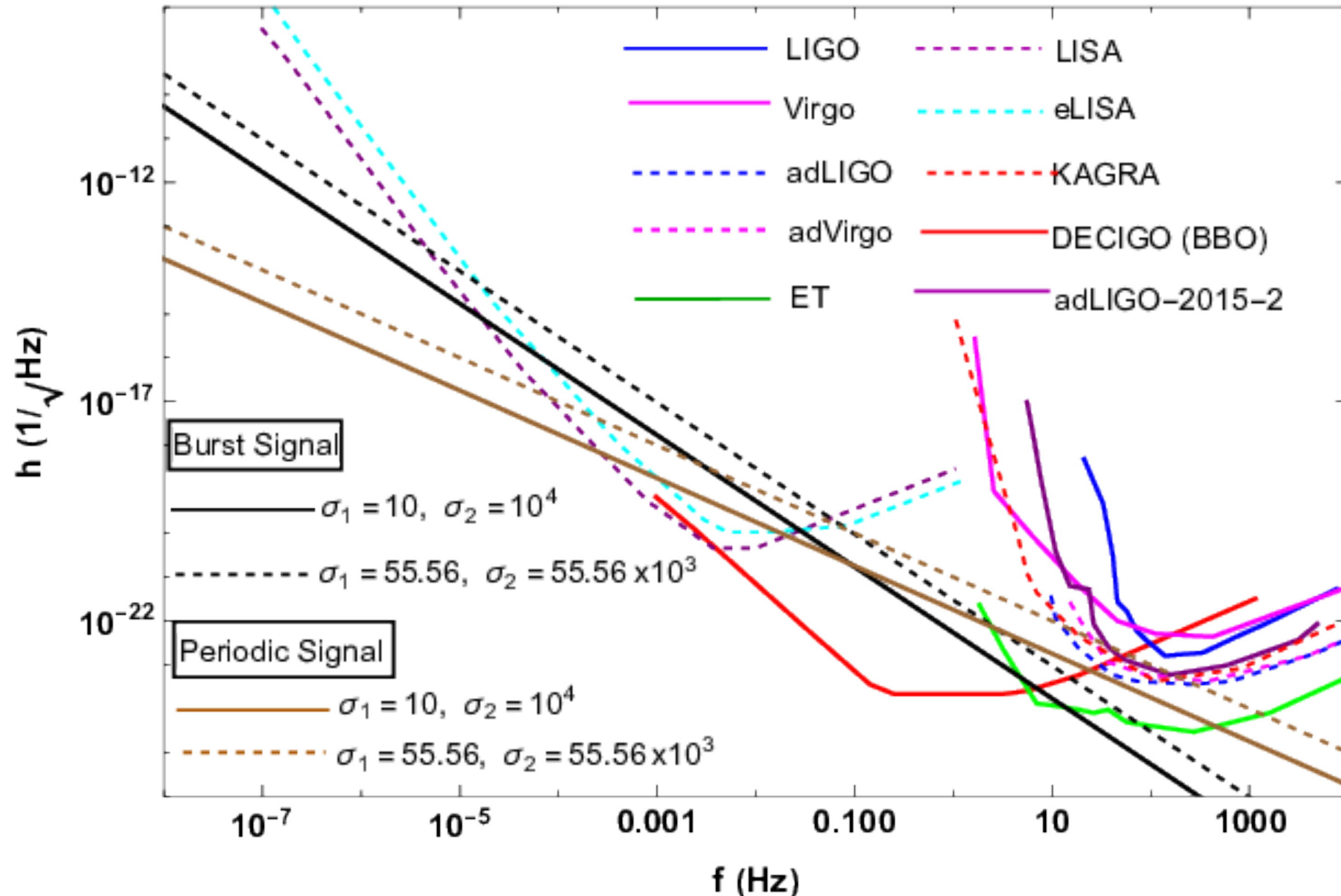
$$\alpha = \frac{S_+ - S_-}{S_+ + S_-} \simeq \frac{\Delta p}{p} = (1/18)(Q/r_{\text{res}})$$

For UVDM

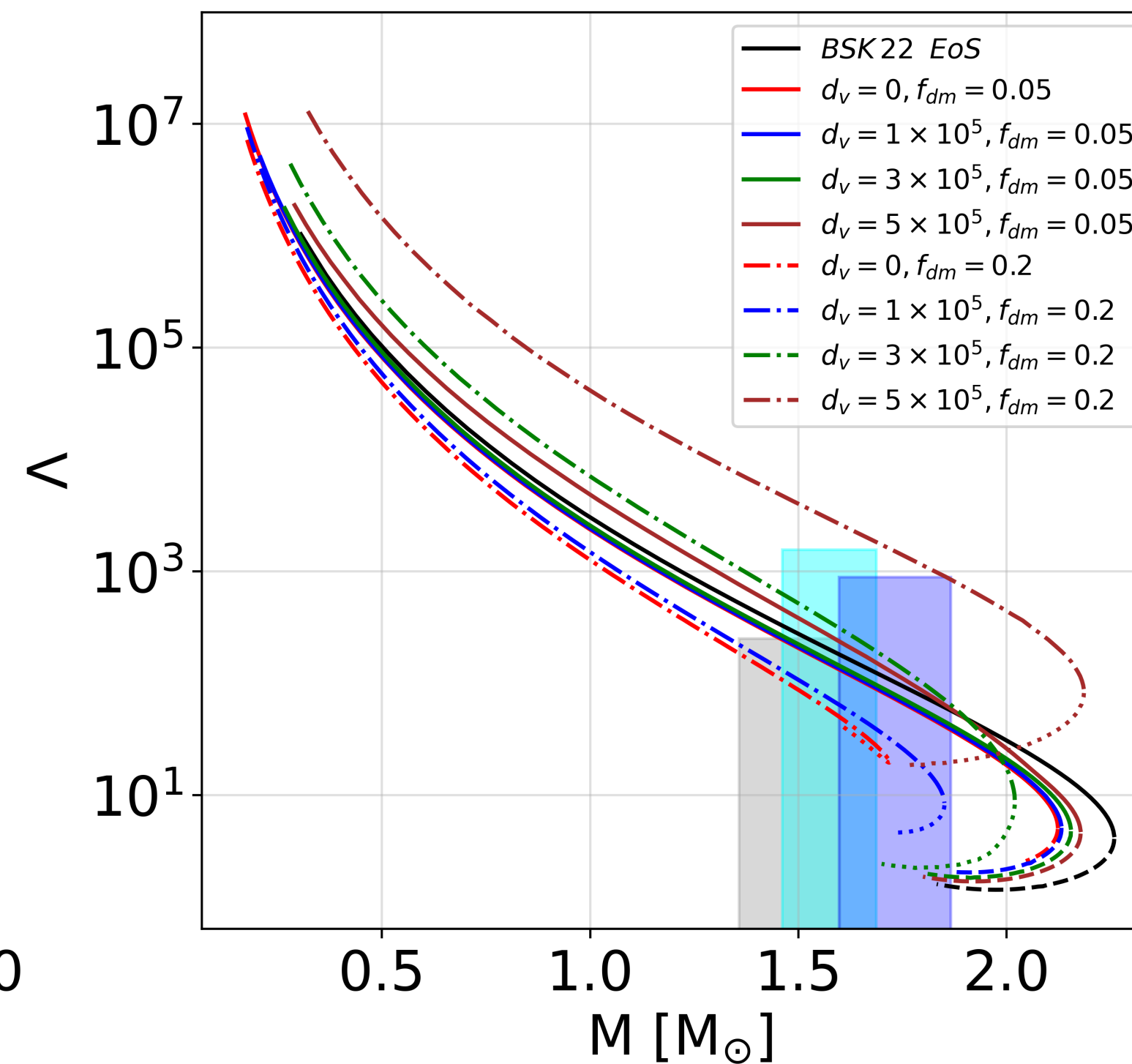
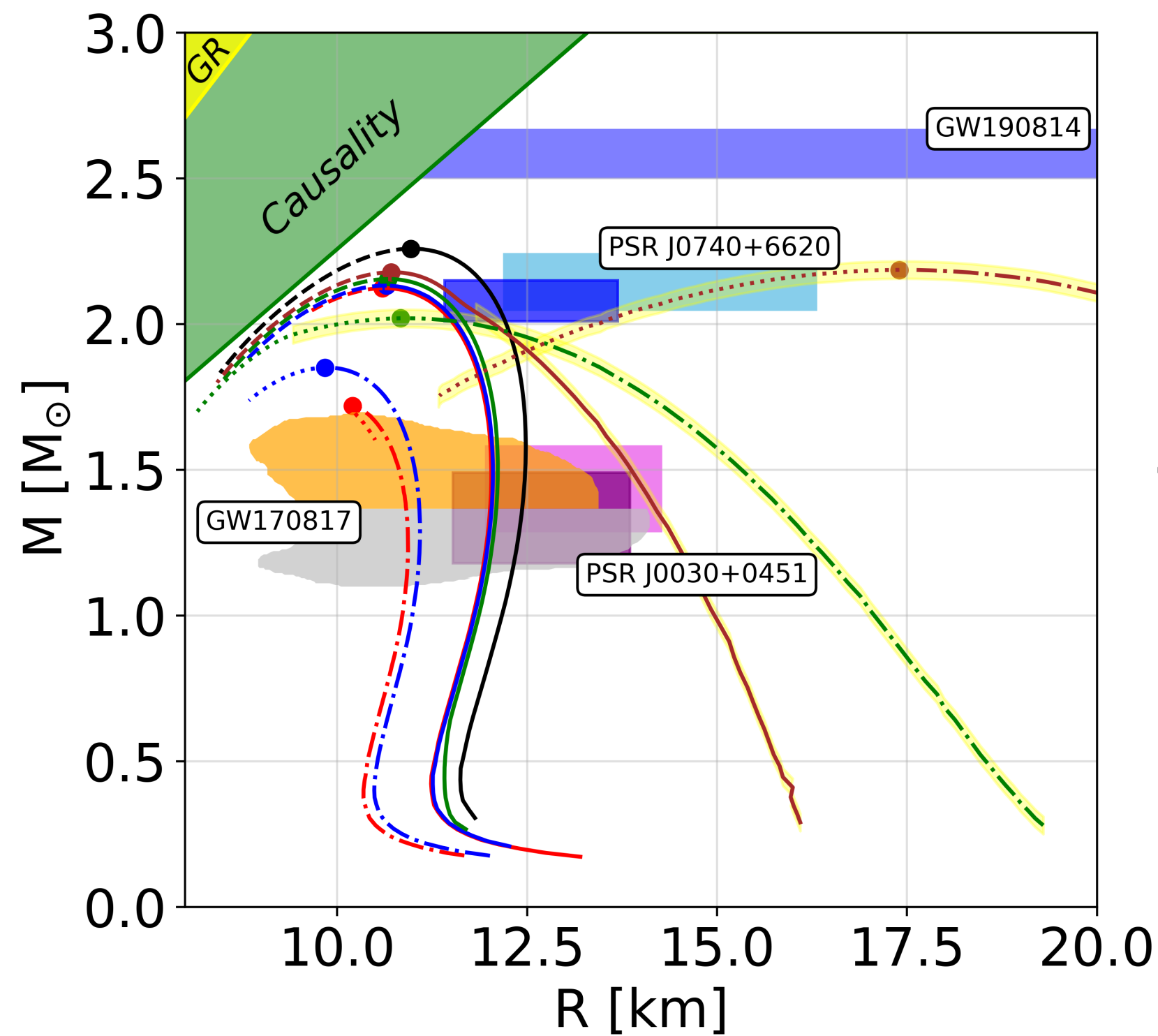
$$\alpha = 1.8 \times 10^{-4} \sigma_1, \quad \sigma_1 = \frac{g'_{ff}/m_{A'}}{x_{\text{res}} \eta}$$

For USDm

$$\alpha = 1.8 \times 10^{-7} \sigma_2, \quad \sigma_2 = \frac{g'_{ff'}}{x_{\text{res}} \eta}$$



$$\mathcal{L} = \bar{\Psi}(x) \left[\gamma_\mu (i\partial^\mu - g_v V^\mu(x)) - (M_D - g_{s_1} \phi(x)) \right] \Psi(x) + \frac{1}{2} (\partial_\mu \phi(x) \partial^\mu \phi(x) - m_{s_1}^2 \phi^2(x)) - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \frac{1}{2} m_v^2 V_\mu(x) V^\mu(x)$$



BM: BSK 22

Thank You !

