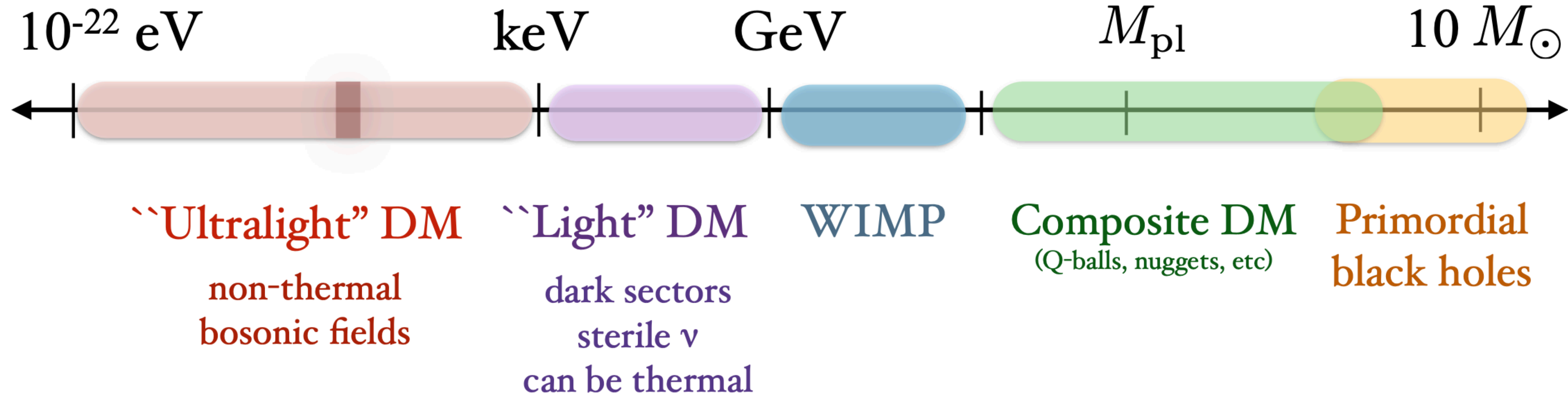


Looking for dark sectors

Fundamental Physics & Gravitational Wave Detectors, Pollica, 10th Sept 2024

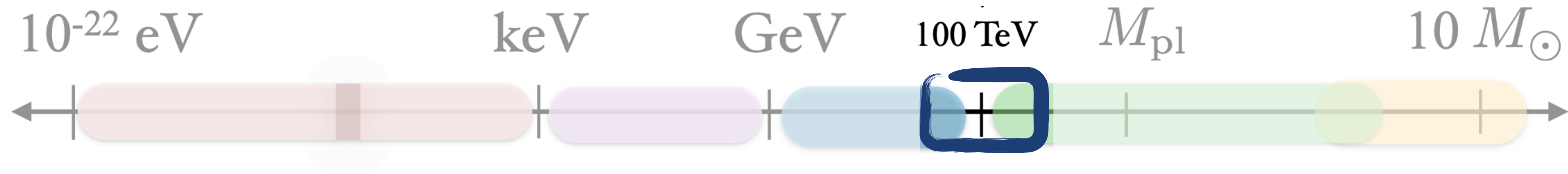
Sonali Verma [ULB Brussels]





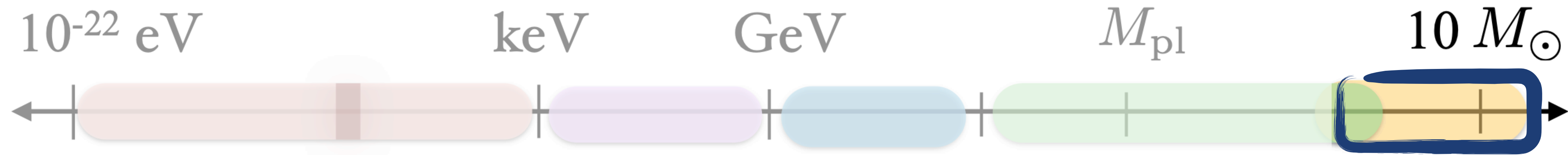
Searching for conformal dark sectors at high-intensity experiments?

[\[M. Costa, R. K. Mishra, SV\]](#)



Could composite dark matter be SU(5) unified?

[\[S. Bottaro, R. Contino, SV\]](#)

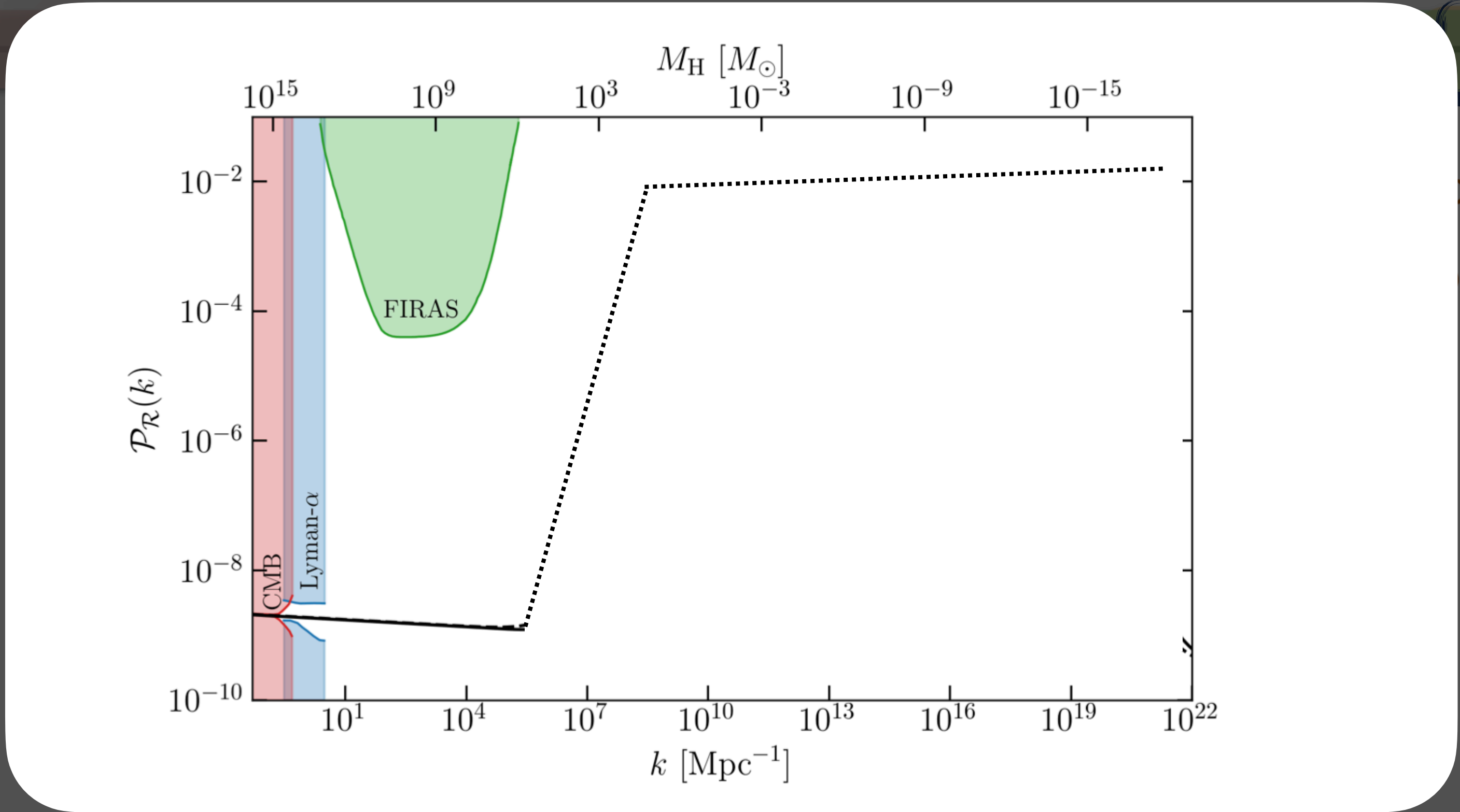


Probing PBH clusters with PTA data?

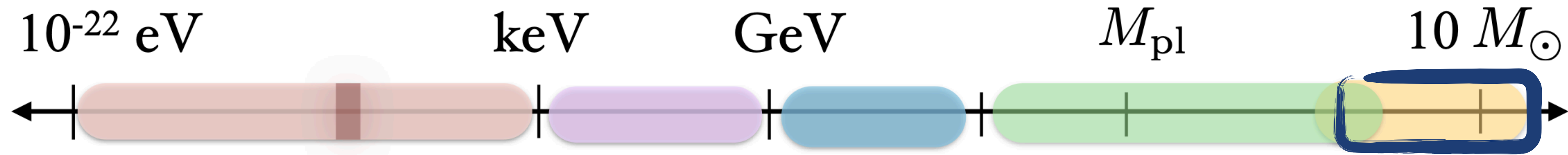
S. Clesse, V. Dandoy, SV [work in progress]

**Primordial
black holes**

10^{-22} eV QCD axion classic window $10^{-6} - 10^{-4}$ eV WDM limit keV GeV unitarity limit 100 TeV M_{pl} $10 M_{\odot}$



Primordial black holes



Probing PBH clusters with PTA data?

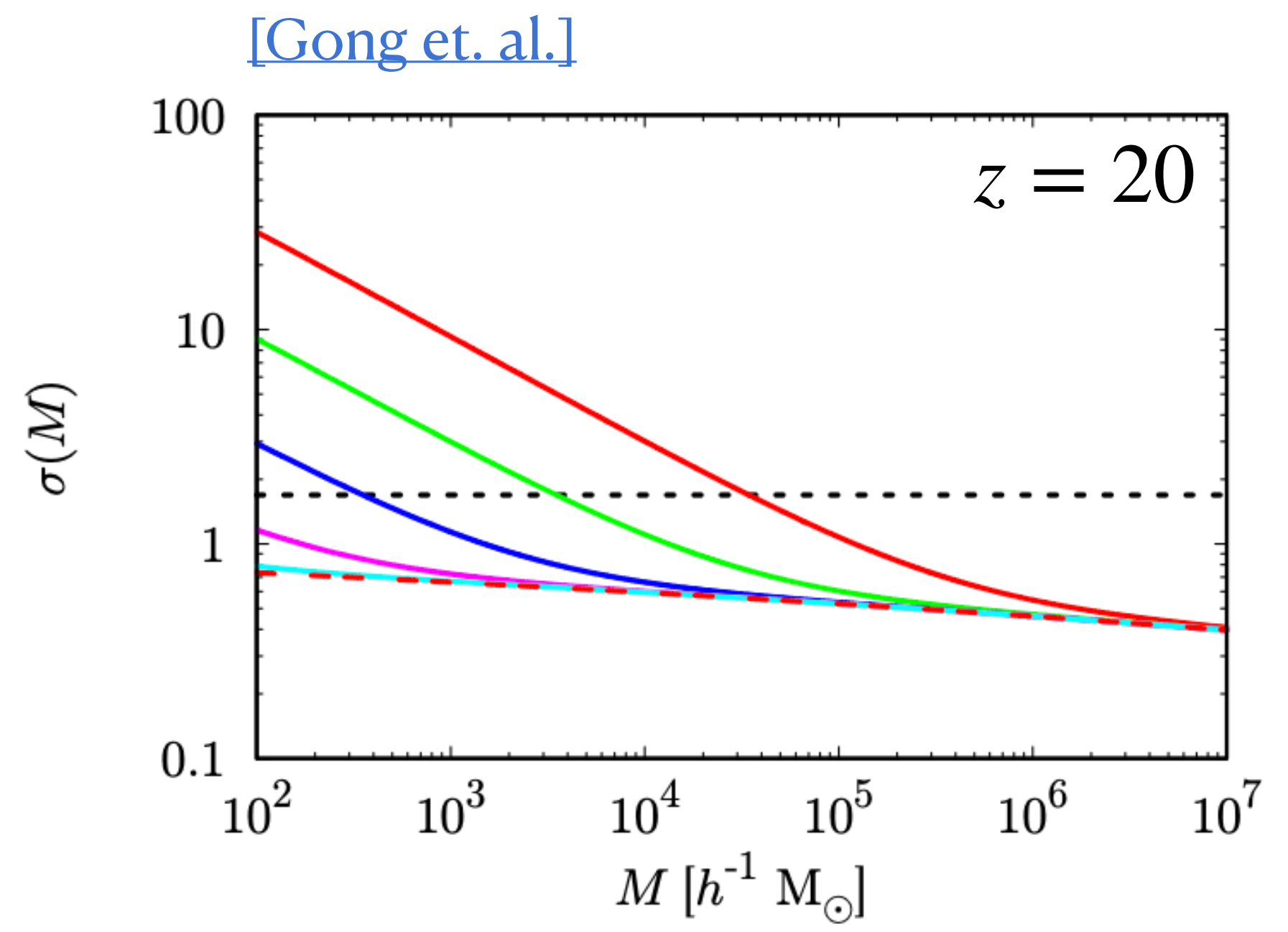
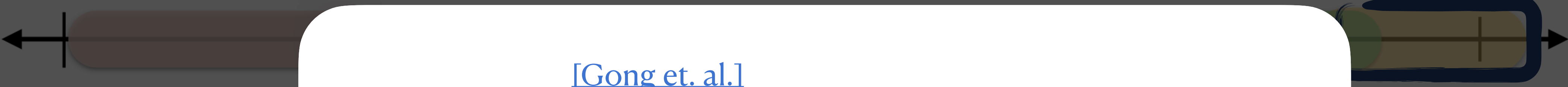
S. Clesse, V. Dandoy, SV [work in progress]

Primordial
black holes

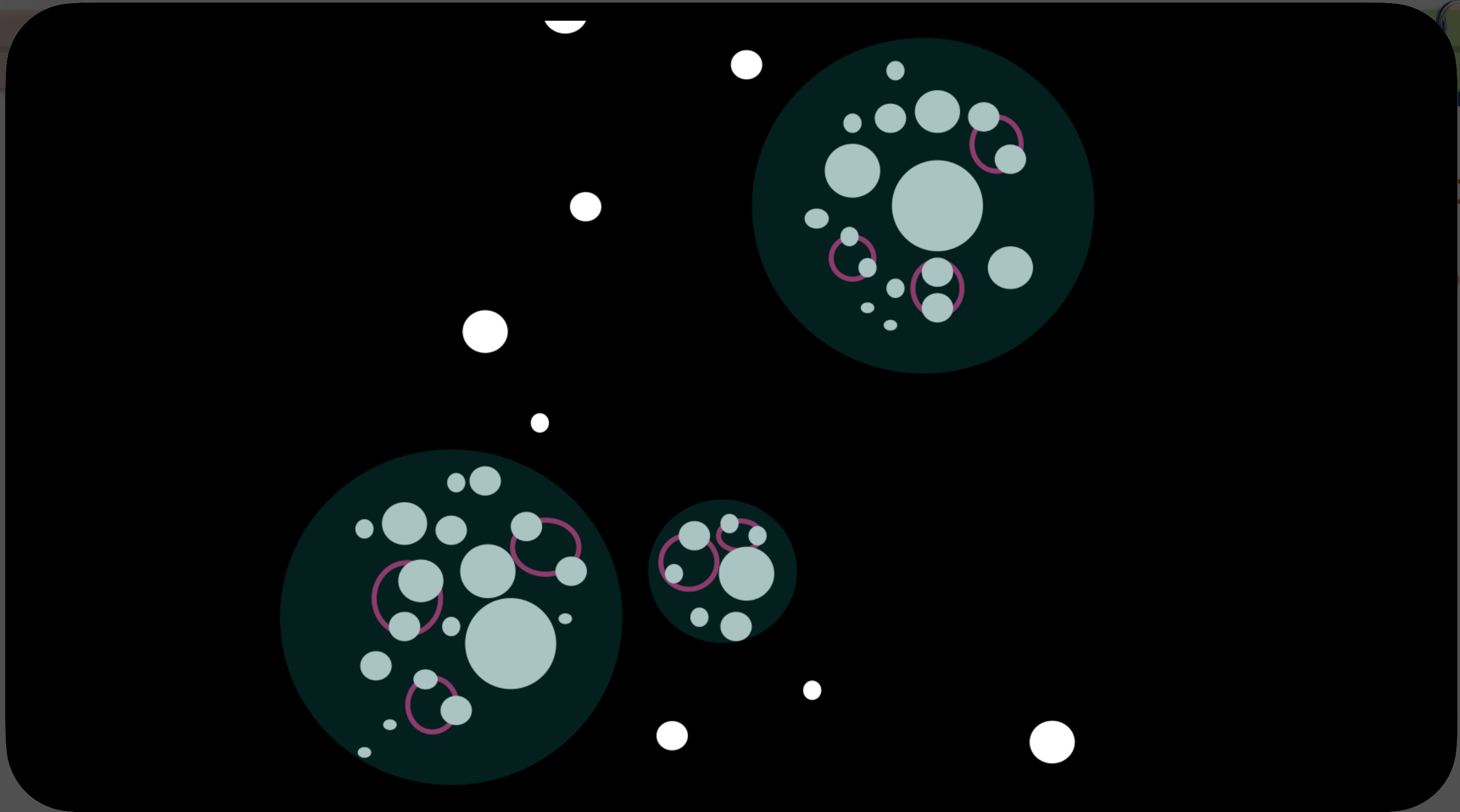
$$P_{poisson} = f_{PBH}^2 n_{PBH}^{-1}$$

Poissonian fluctuations will enhance matter power spectrum at small scales

10^{-22} eV QCD axion classic window $10^{-6} - 10^{-4}$ eV WDM limit keV GeV unitarity limit 100 TeV M_{pl} $10 M_{\odot}$



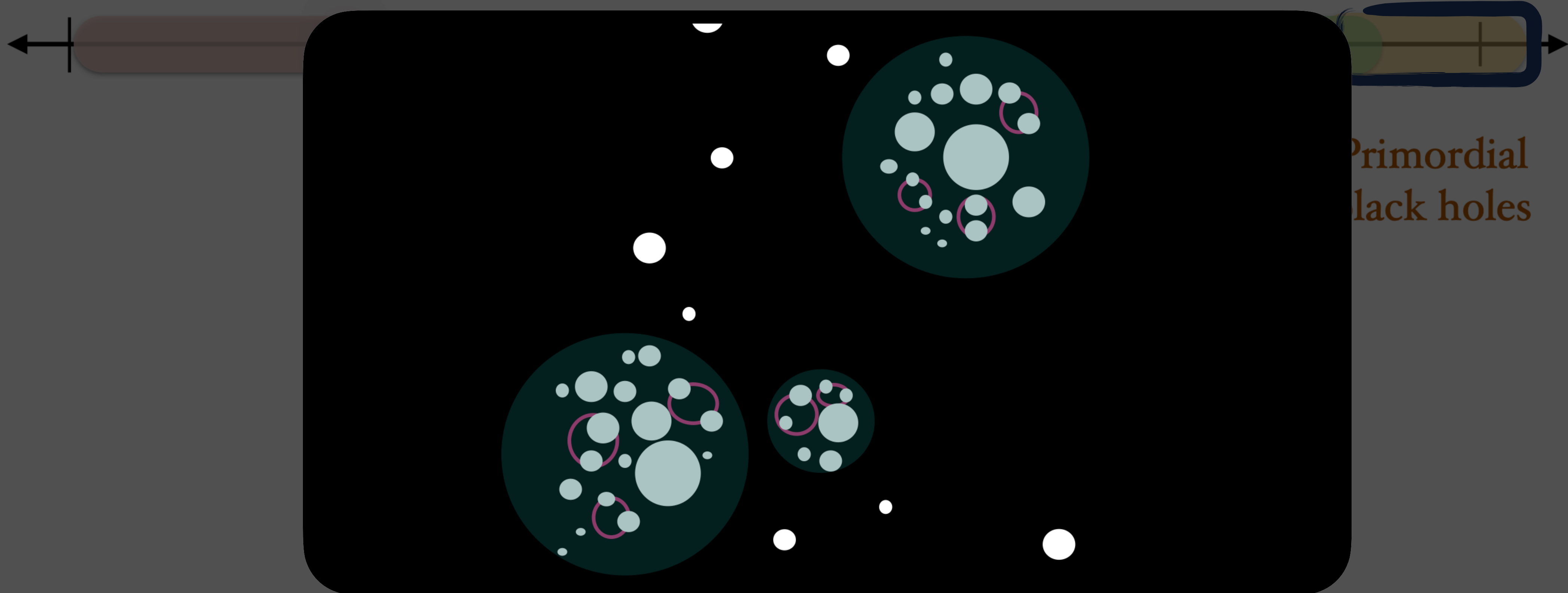
Primordial black holes



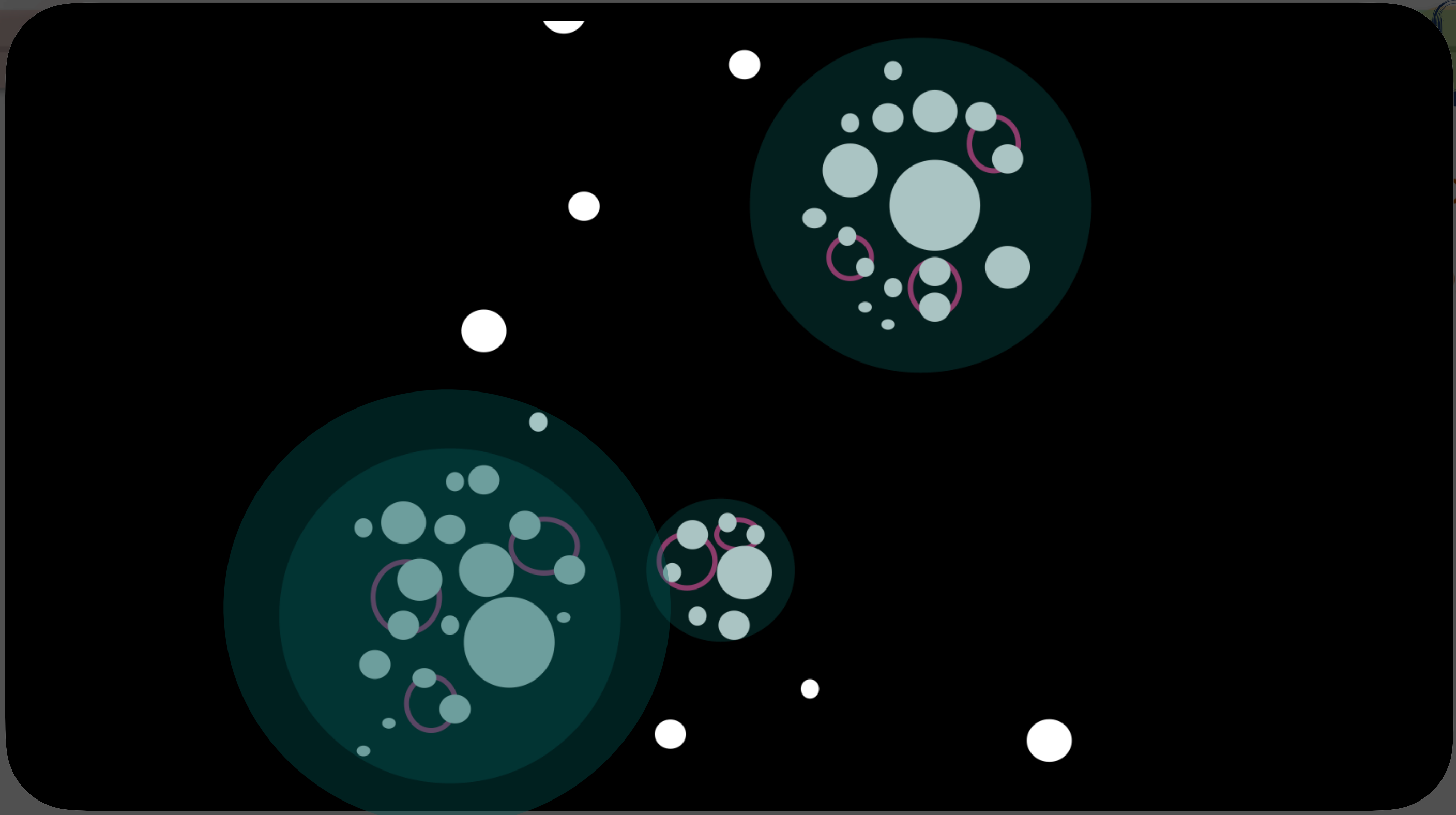
Primordial
black holes

PBH binary formation can take place via capture in PBH clusters

10^{-22} eV QCD axion classic window $10^{-6} - 10^{-4}$ eV WDM limit keV GeV unitarity limit 100 TeV M_{pl} $10 M_{\odot}$

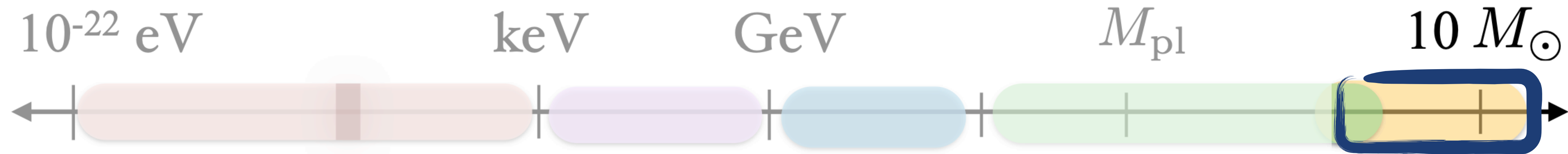


$$\Omega_{GW}(f) = f^{2/3} \frac{(\pi G)^{2/3}}{\rho_c} \int dz d \ln m_1 d \ln m_2 \frac{1}{H(z)(1+z)^{4/3}} \frac{d \mathcal{V}(z)}{d \ln m_1 d \ln m_2} \mathcal{M}_c^{5/3}$$



Primordial
black holes

PBH clusters can also grow in radius due to dynamical heating



Probing PBH clusters with PTA data?

S. Clesse, V. Dandoy, SV [work in progress]

**Primordial
black holes**

GW spectrum from PBH mergers will be sub-leading wrt SIGWs triggered at PBH formation

PBH merger rates required to fit PTA data $\sim \mathcal{O}(10^4) \text{ yr}^{-1} \text{ Gpc}^{-3}$ wrt LIGO/VIRGO events

[\[Bellido et. al\]](#)