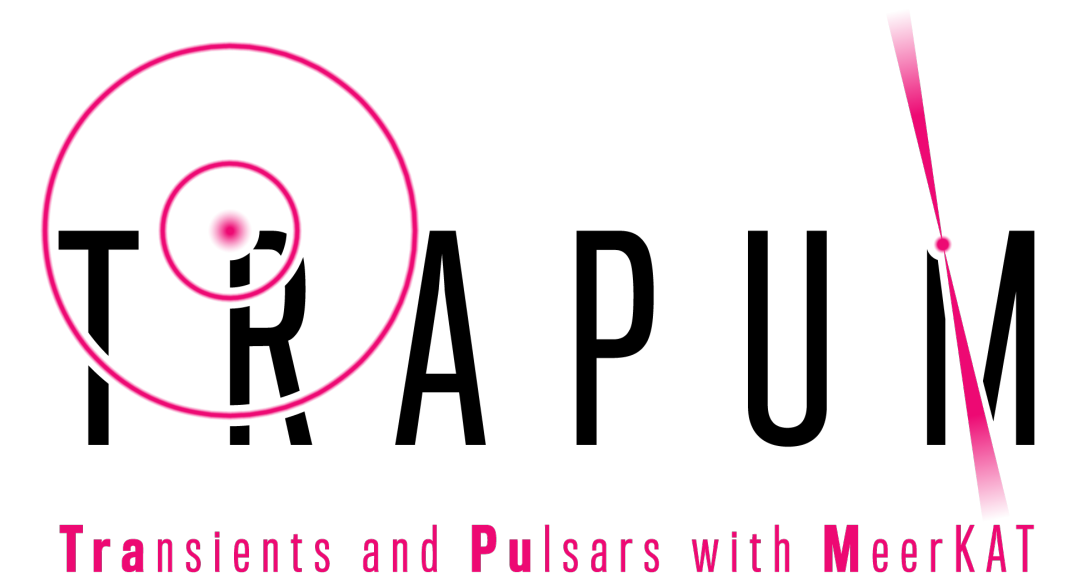
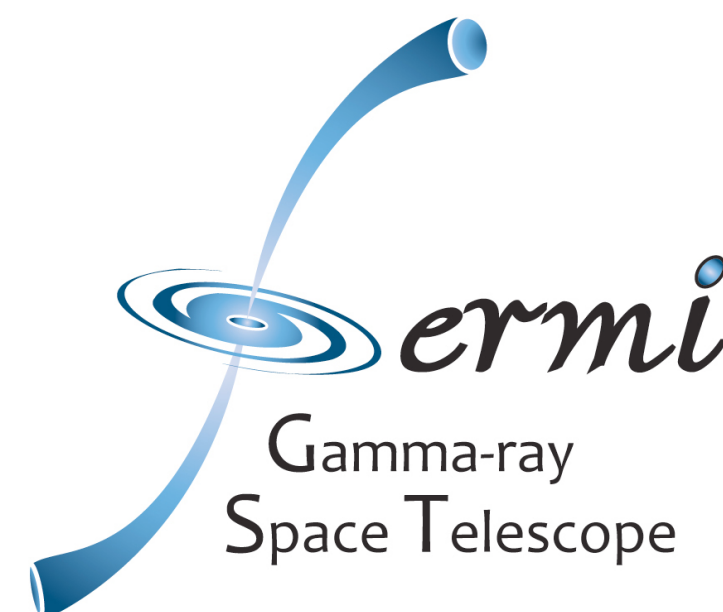
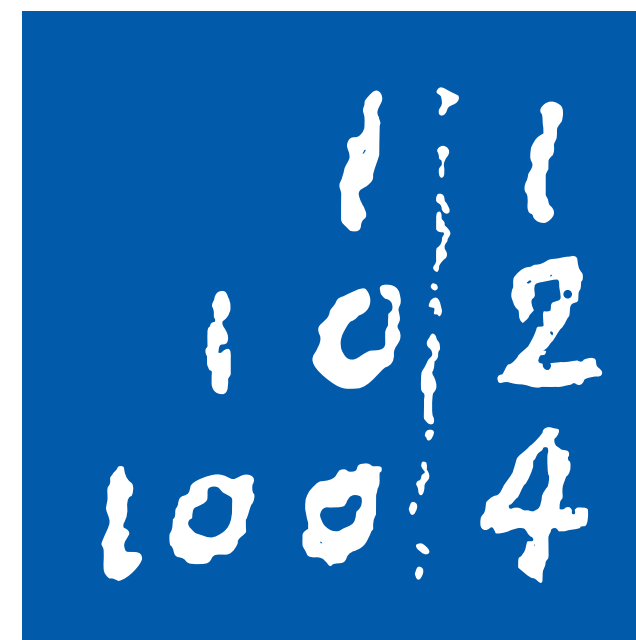


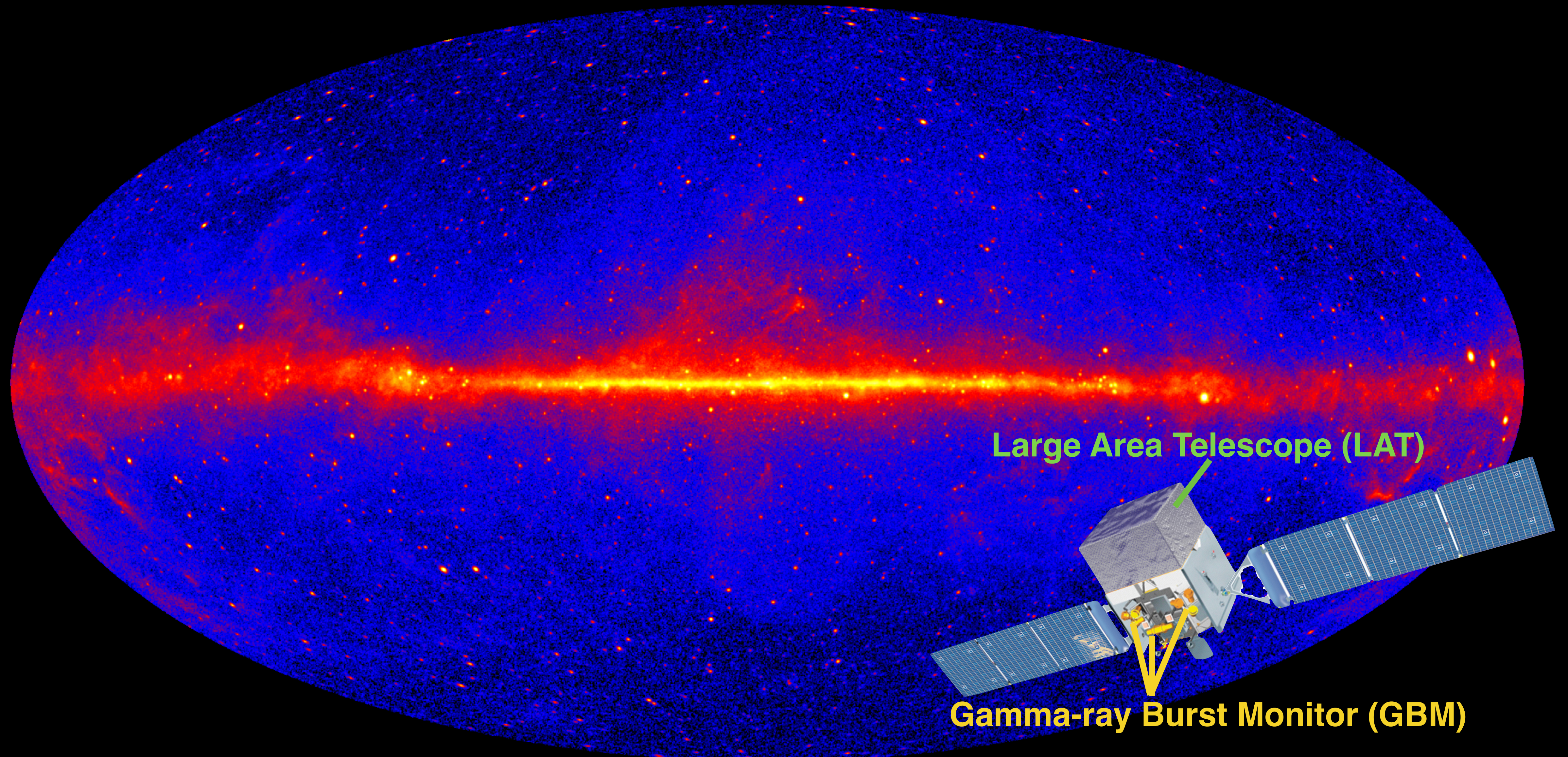
Searching for Radio and Gamma-ray Millisecond Pulsars with MeerKAT

Colin J. Clark, on behalf of the TRAPUM
and *Fermi*-LAT collaborations

GEMMA2 - Sapienza University, Rome
18/09/2024

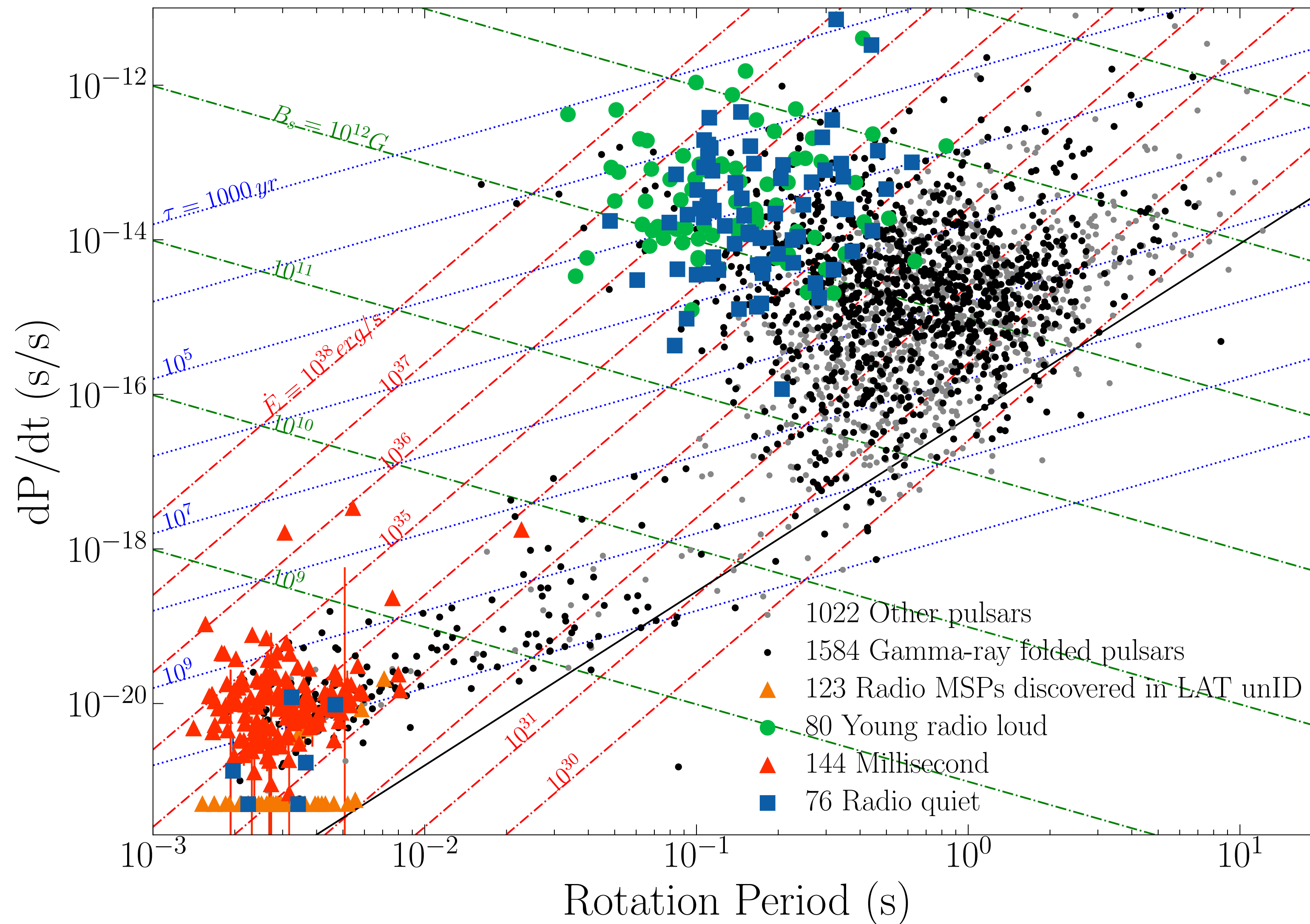


Fermi Gamma-ray Space Telescope

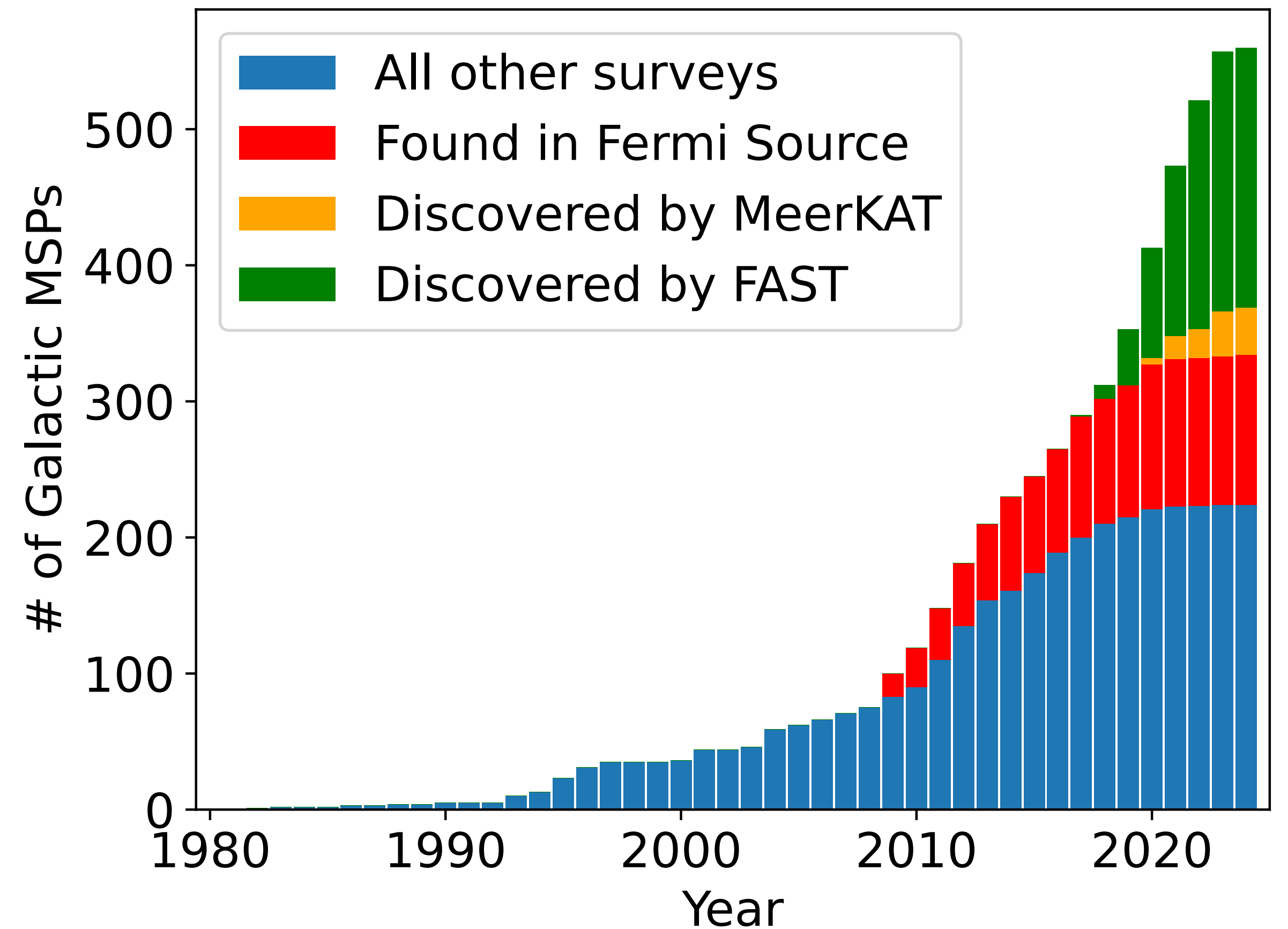
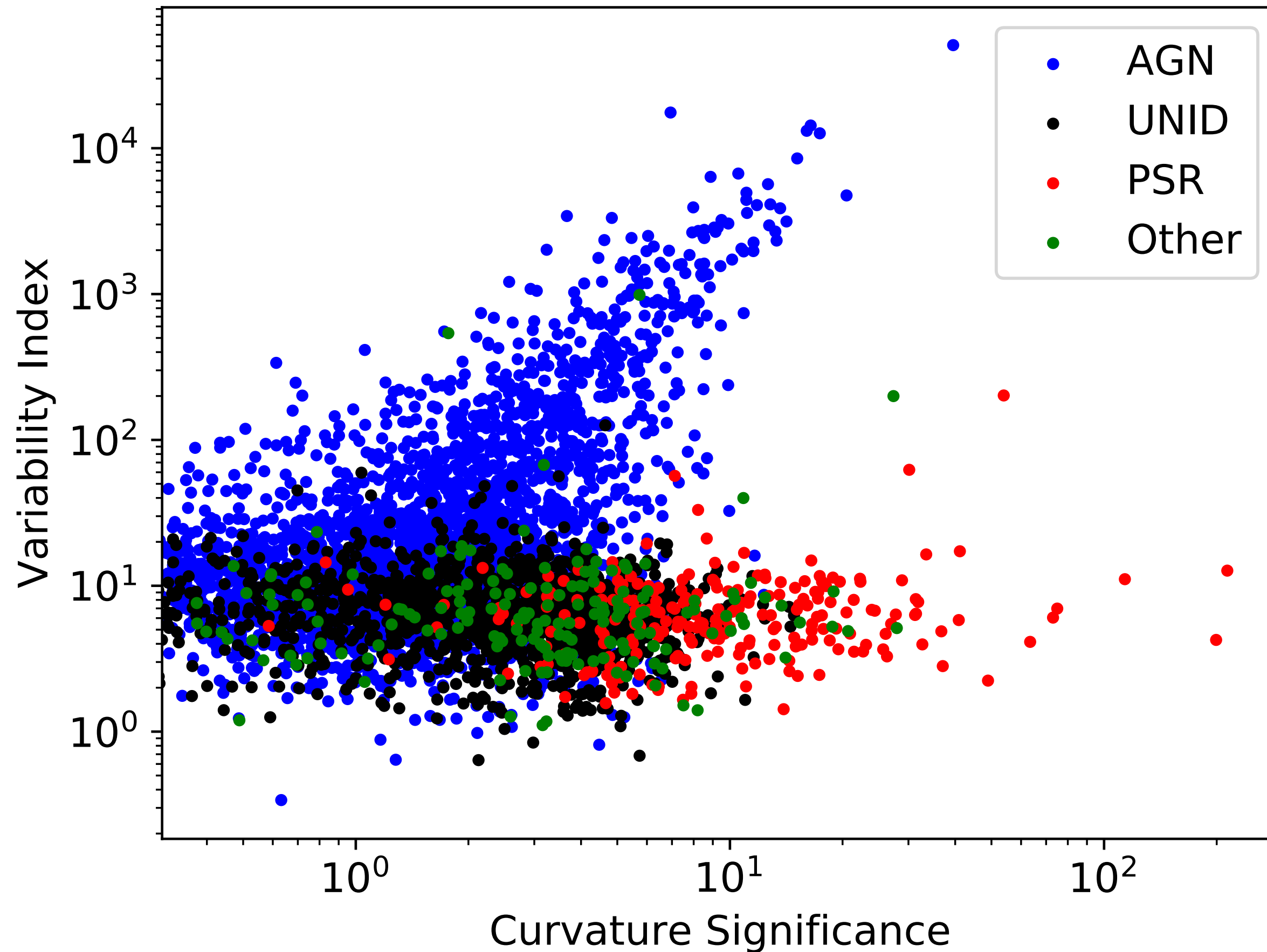


The 3rd Fermi-LAT Catalogue of Gamma-ray Pulsars

Smith, D., et al., 2023, *ApJS*, 958, 191



The Fermi “Treasure Trove”



**> 120 new MSPs found in Fermi sources by the Pulsar Search Consortium
(See Ray et al. 2012, arXiv:1205.3089) + FAST + TRAPUM**

Gamma-ray MSPs and Dark Matter

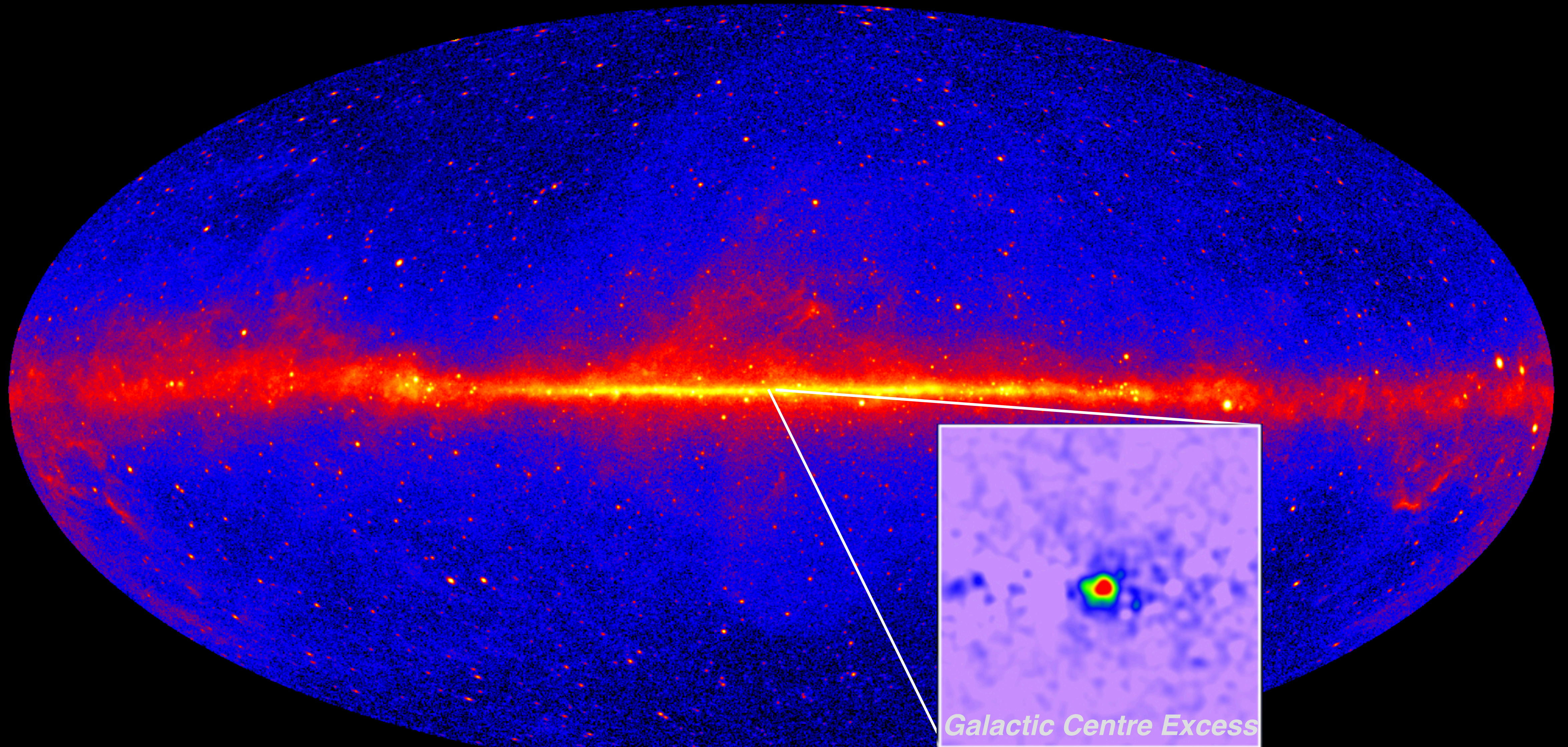
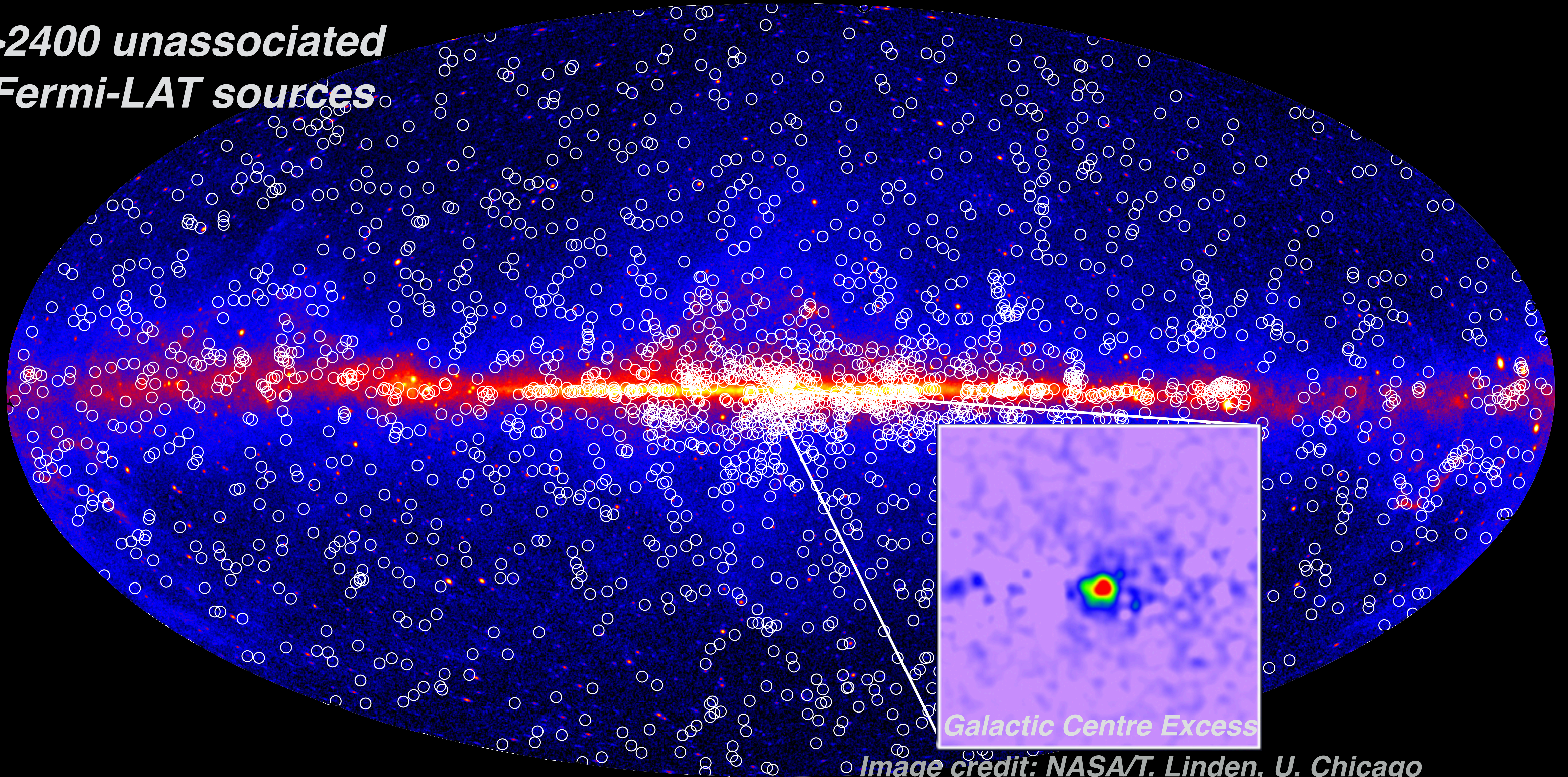


Image credit: NASA/T. Linden, U. Chicago

Gamma-ray MSPs and Dark Matter

**>2400 unassociated
Fermi-LAT sources**

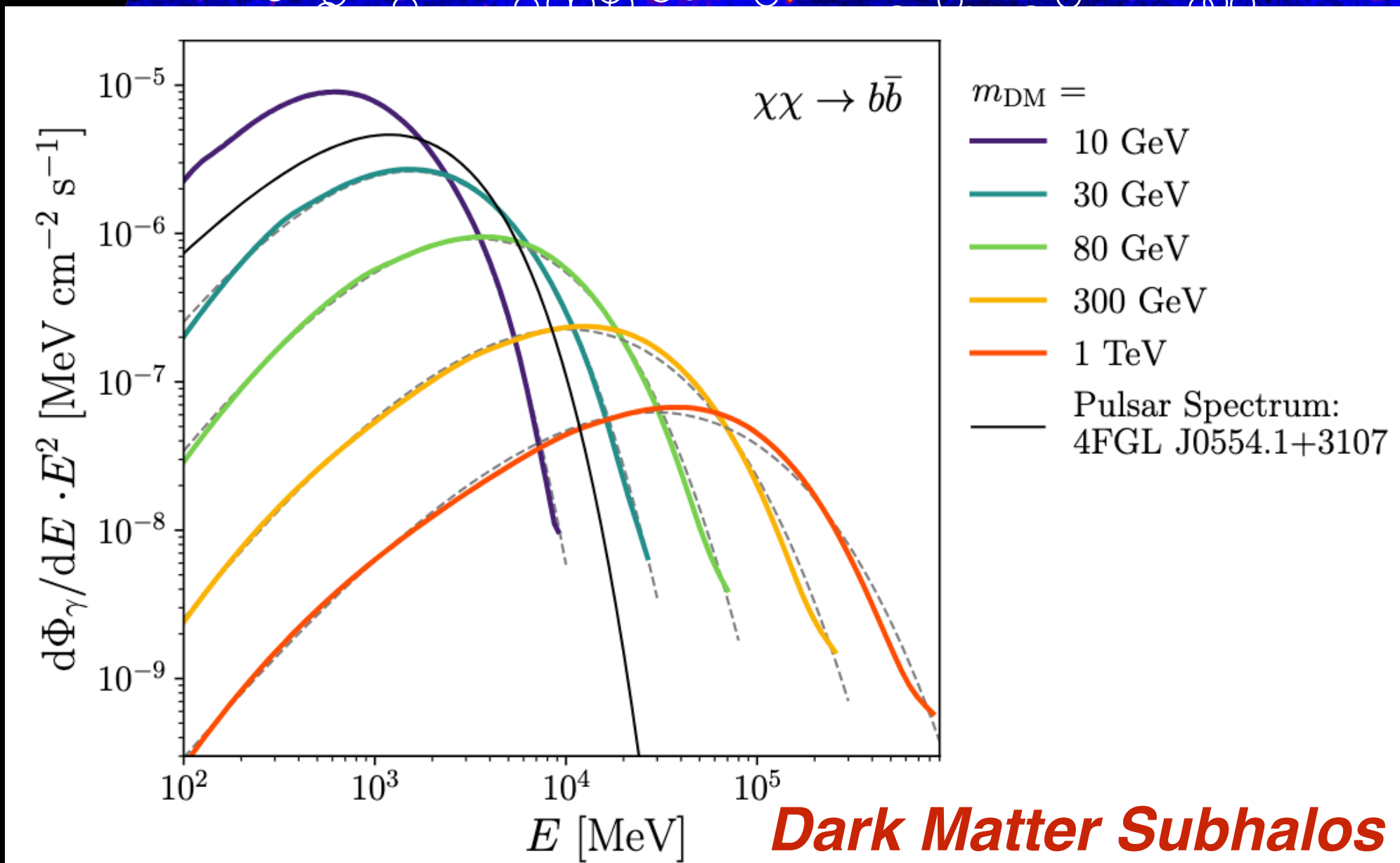


Galactic Centre Excess

Image credit: NASA/T. Linden, U. Chicago

Gamma-ray MSPs and Dark Matter

>2400 unassociated
Fermi-LAT sources



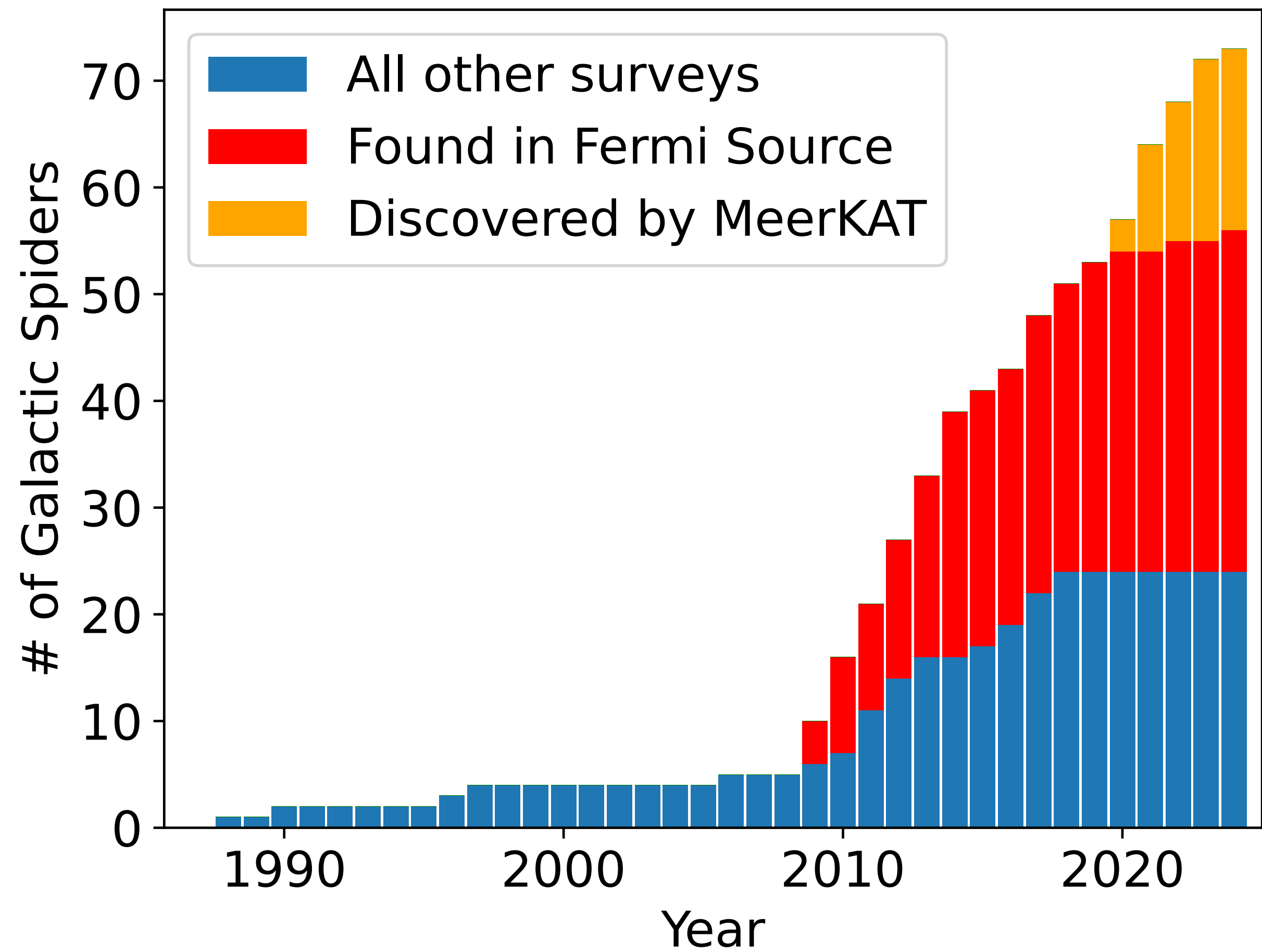
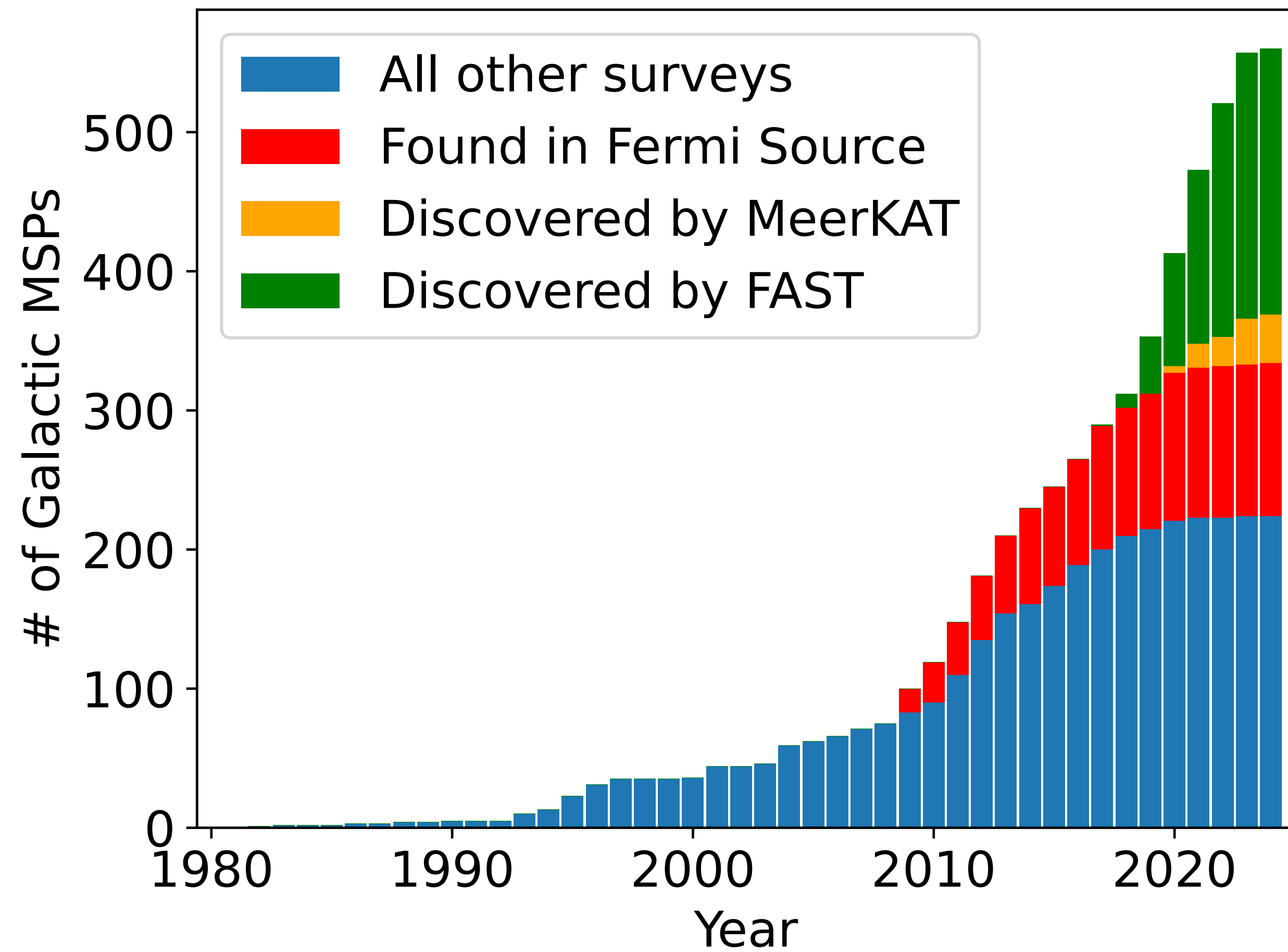
Butter, et al., JCAP07(2023)033

Image credit: NASA/T. Linden, U. Chicago



*Artwork by Soheb Mandhai (JBCA)
@TheAstroPhoenix*

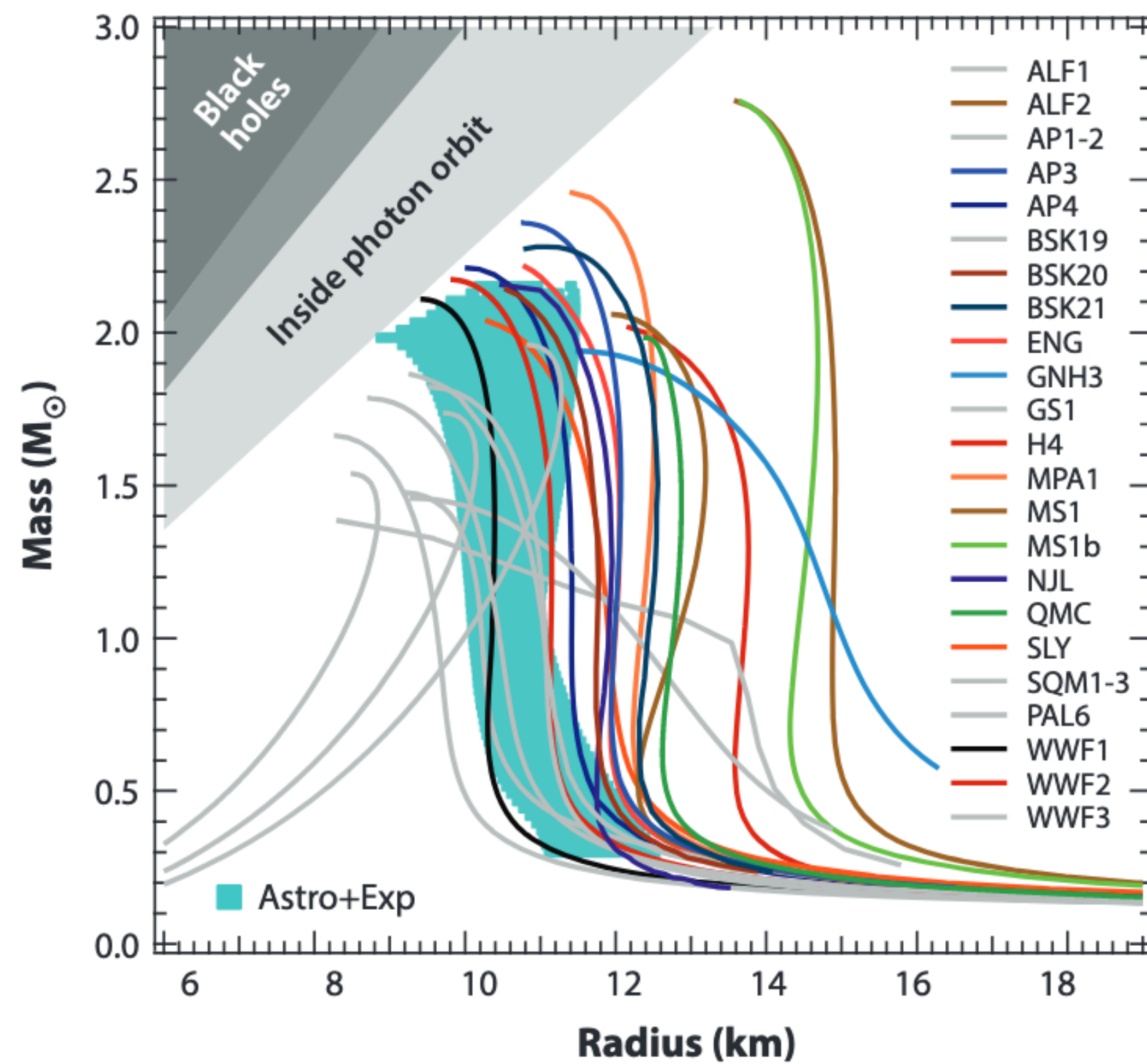
The Fermi “Treasure Trove”



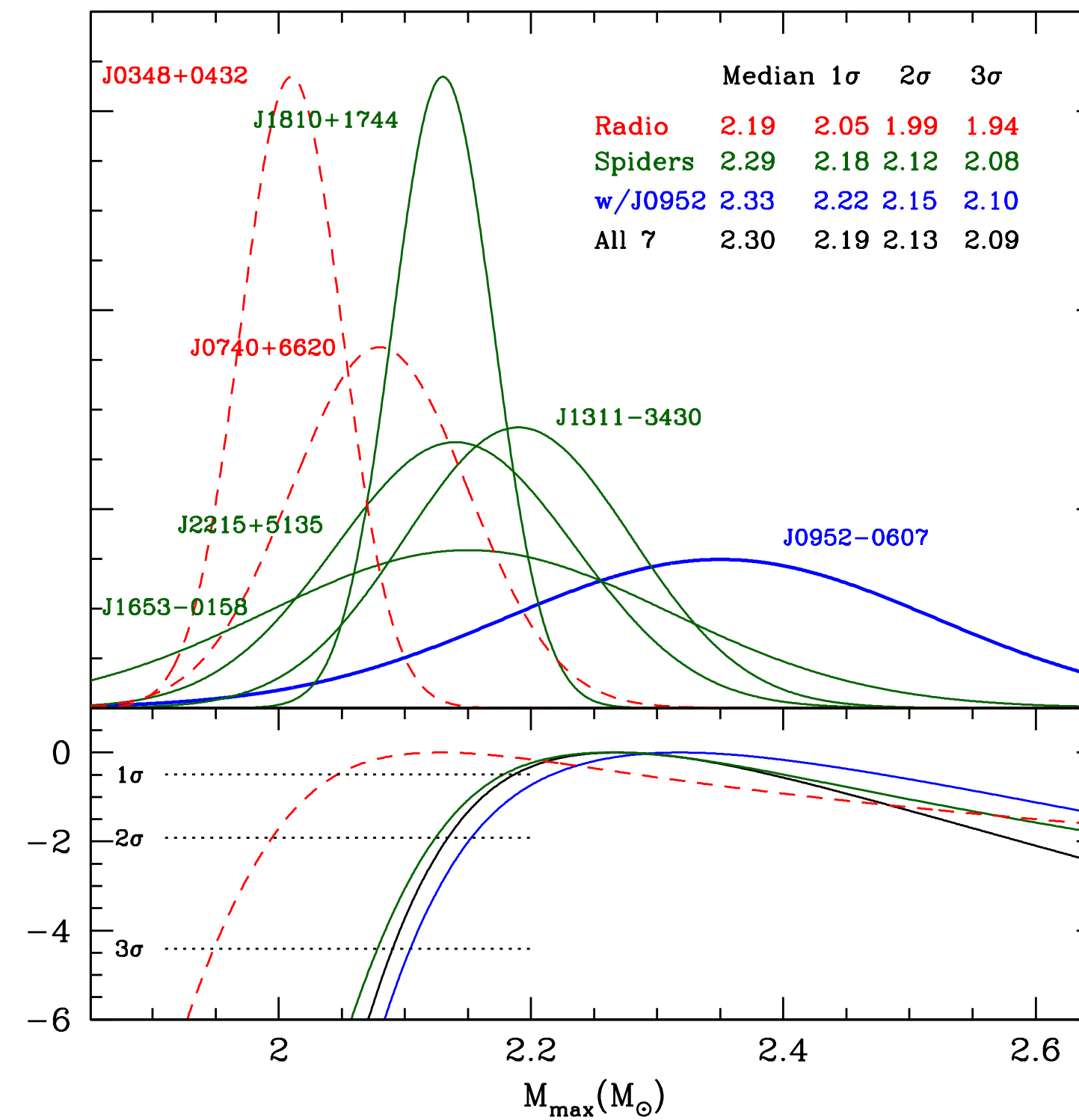
Spider Pulsar Masses and the EoS

Are spider pulsars heavier than other types of neutron star?

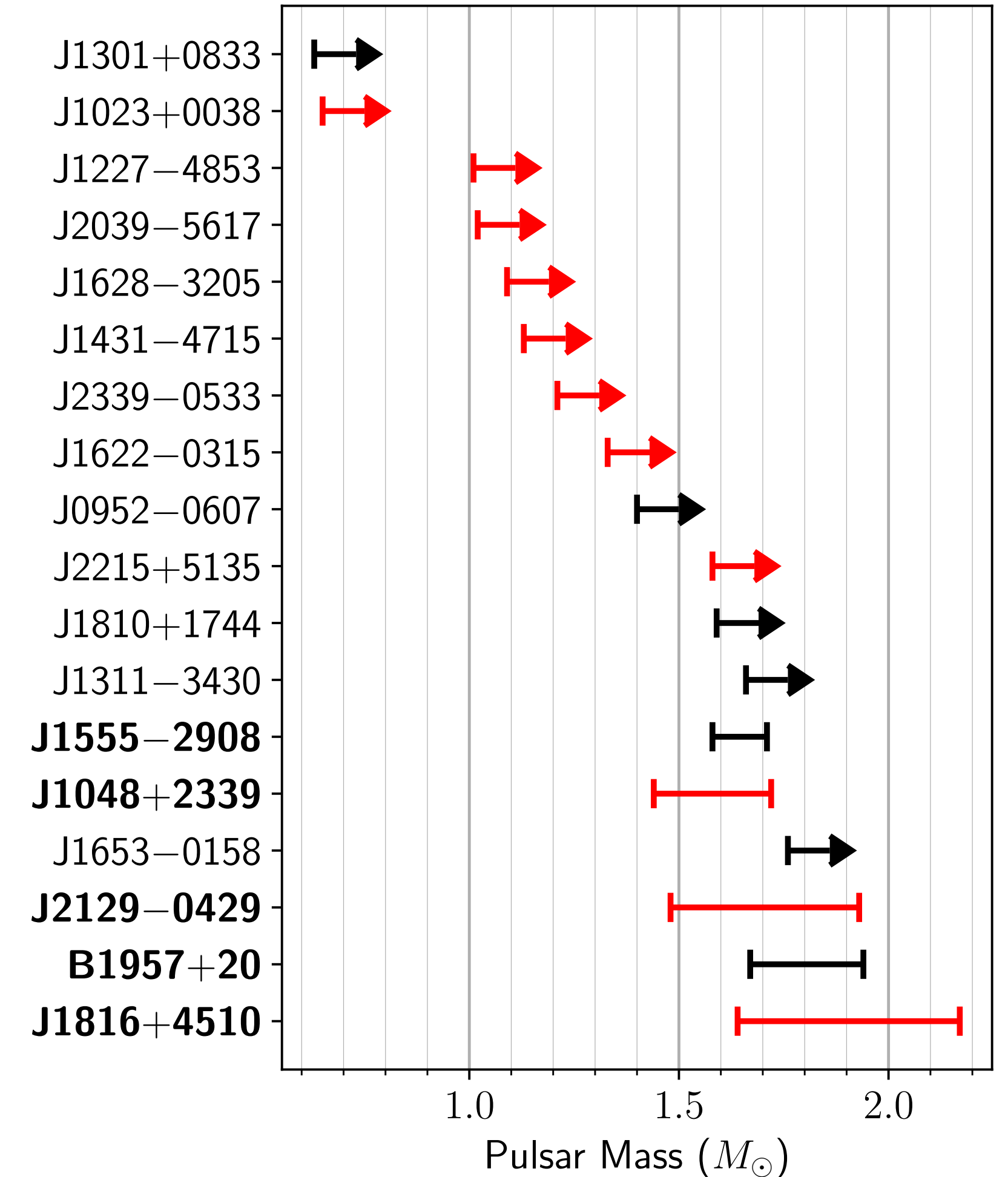
Özel & Freire, 2016, ANA&A, 54, 401



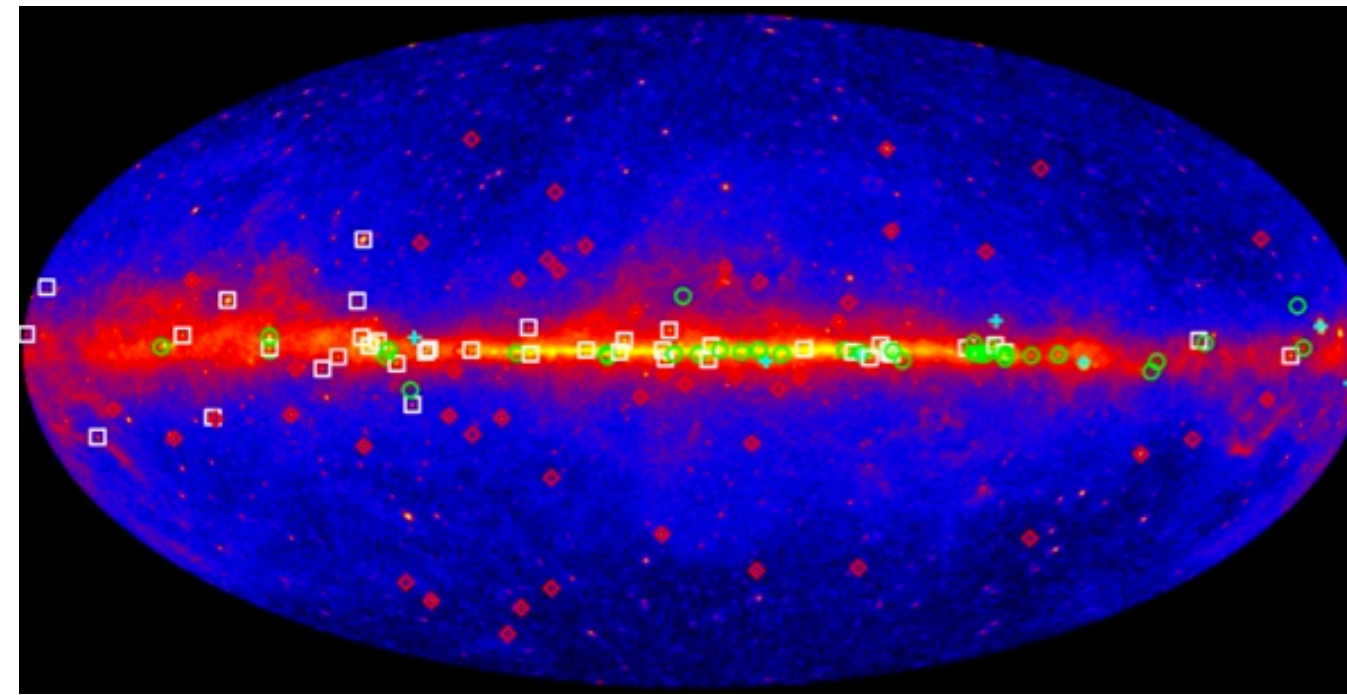
Romani, R.W. et al., 2022, ApJL, 934, L17



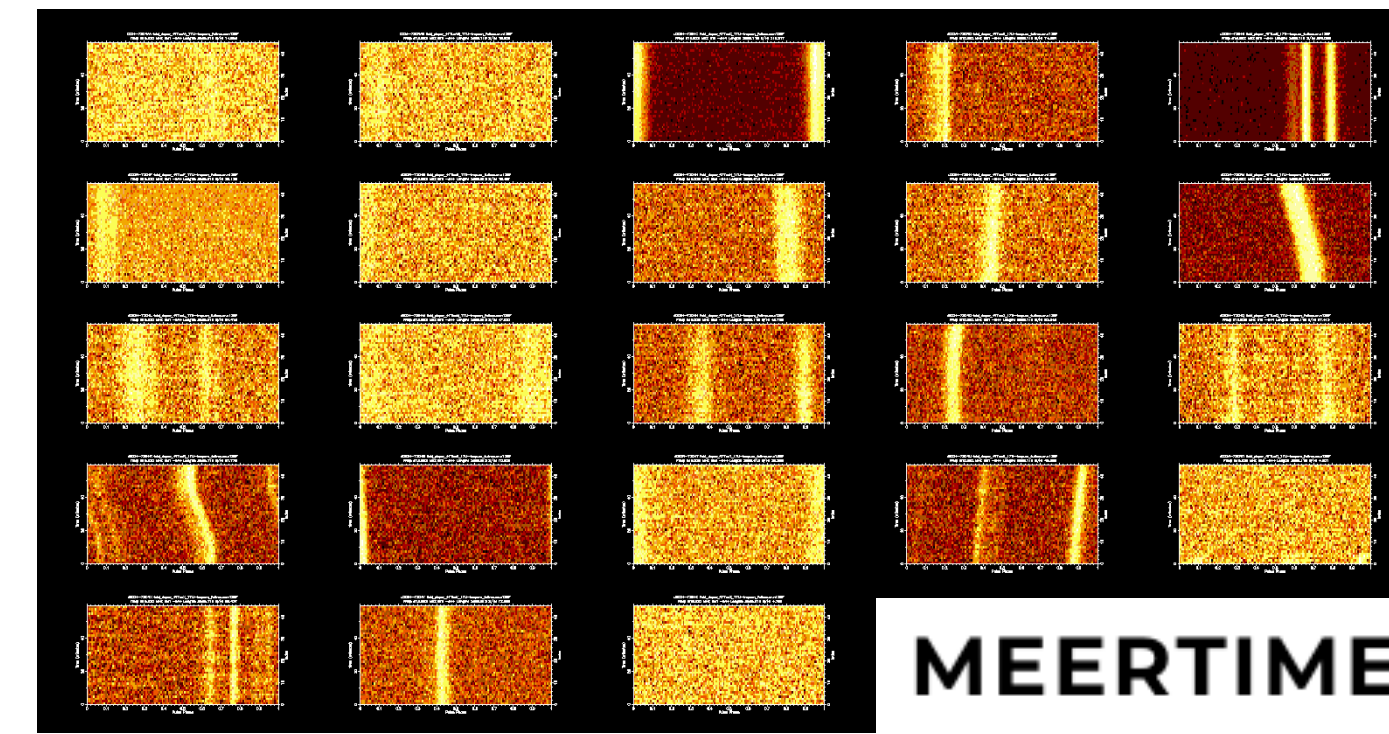
C.J.C., et al., 2023, Nature Astronomy, 7, 451



**Unidentified Fermi
LAT sources**
Chairs: C.J.C.
& R. Breton



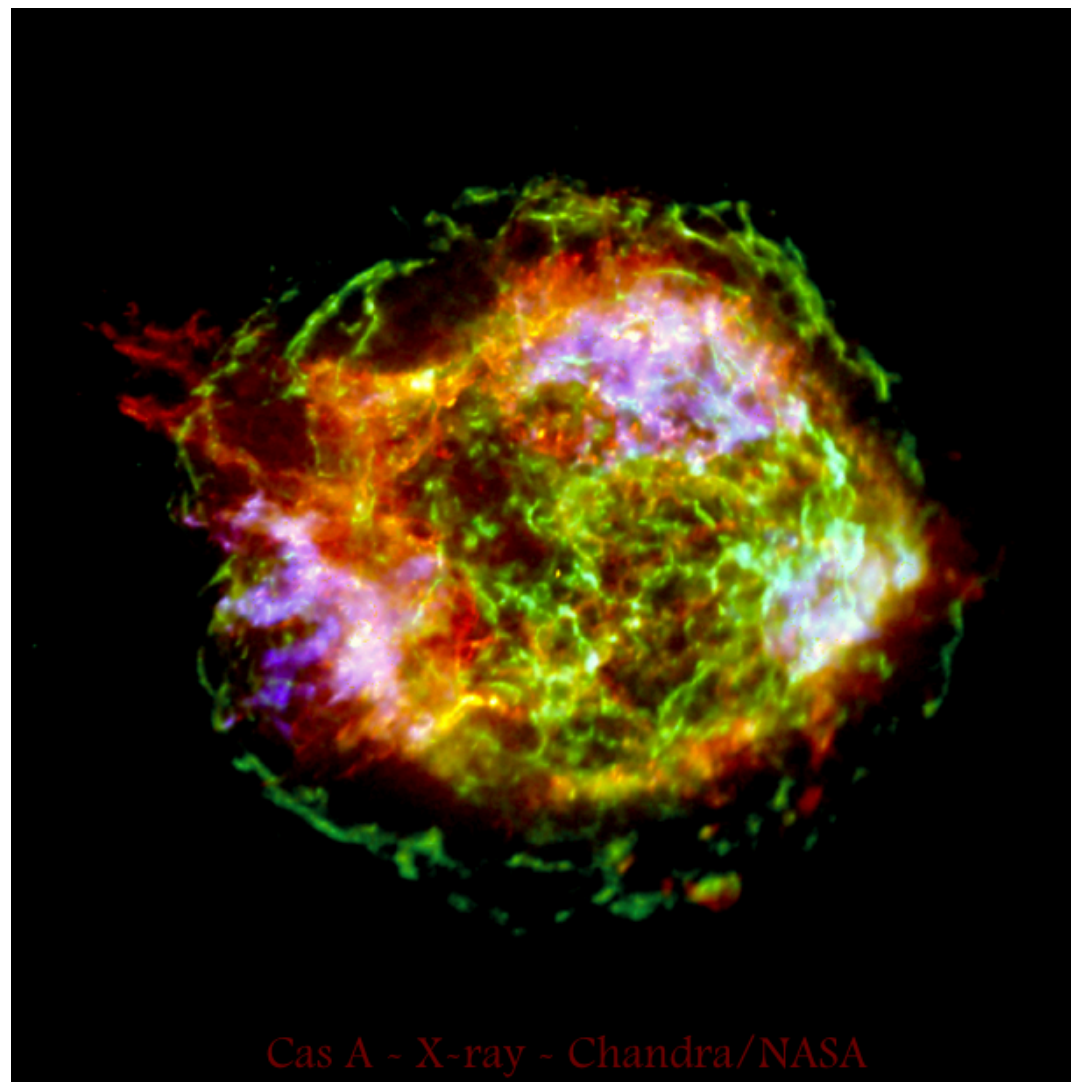
Credit: FERMI/NASA



MEERTIME

Globular clusters
Chair: A. Ridolfi

Credit: Chandra/NASA



Cas A - X-ray - Chandra/NASA

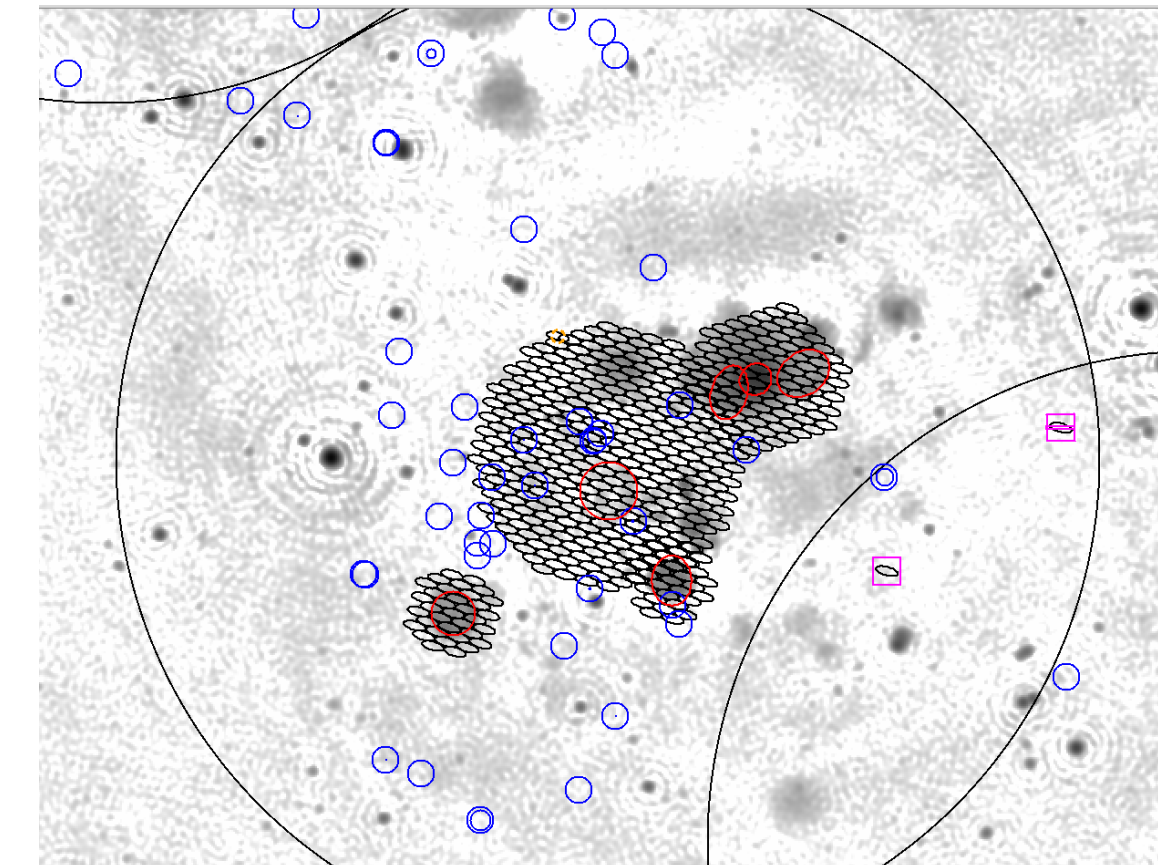


Transients and Pulsars with MeerKAT

Principal Investigators: M Kramer & B. W. Stappers

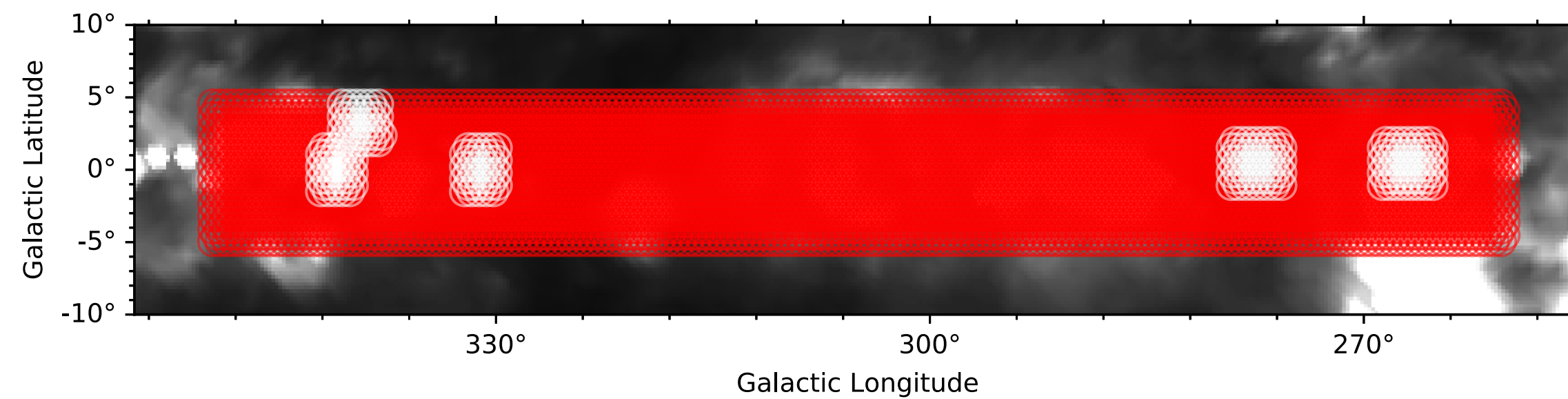
Project Scientist: E. Barr

Follow-up Coordinator: M. Burgay



Nearby galaxies
Chair: L. Levin-Preston

SNR/PWNe/TeV sources
Chair: B. W. Stappers



**MPIfR MeerKAT Galactic Plane Survey
(MMGPS)**

MeerKAT

Number of antennas 64

Dish diameter 13.5 m

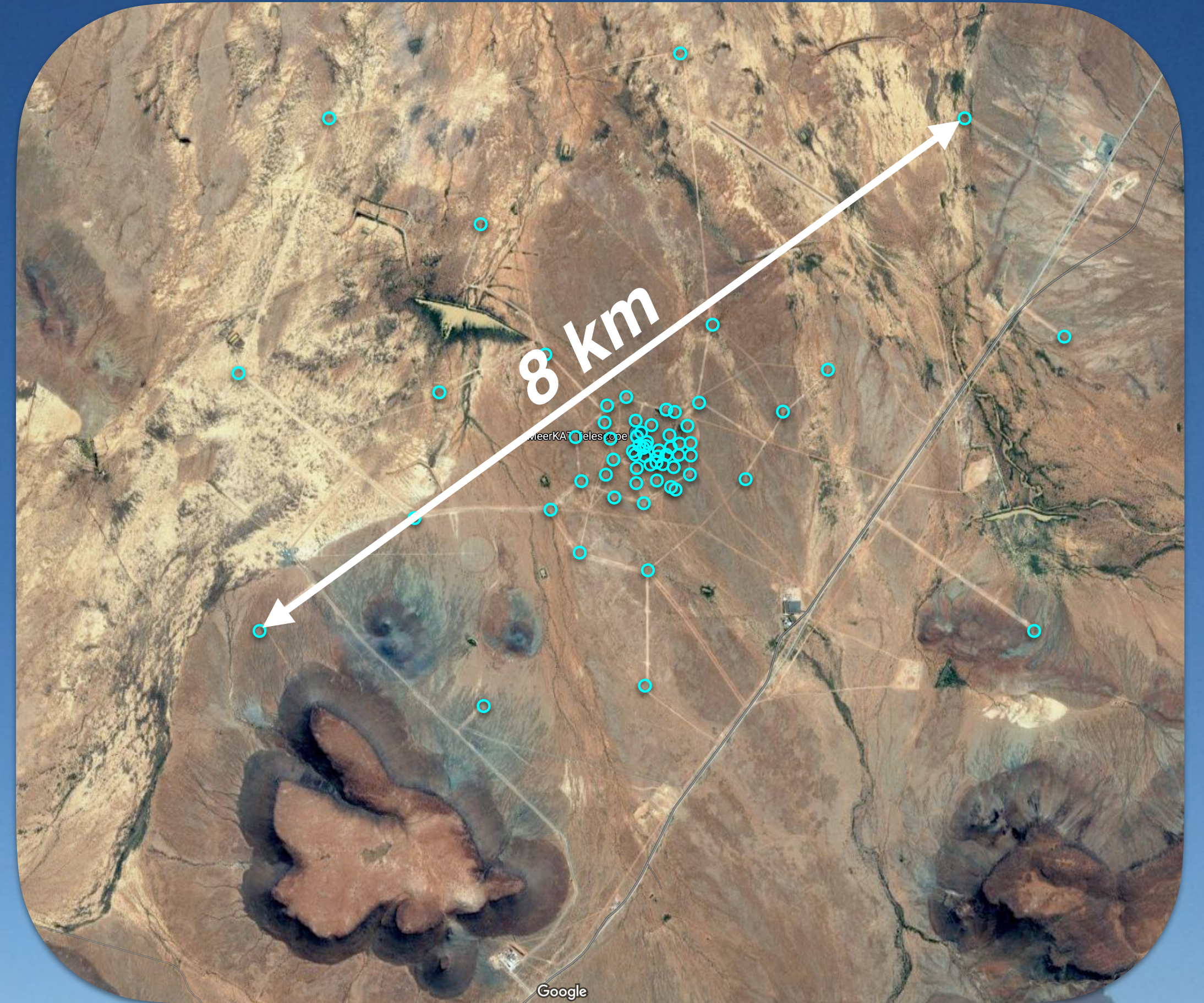
Minimum baseline 29 m

Maximum baseline 8 km

Frequency bands
 544 – 1088 MHz (U-band)
 856 – 1712 MHz (L-band)
 1.75 – 3.5 GHz (S-band)

Instantaneous bandwidth 856 MHz

Sensitivity
(0.58 – 1.67GHz) 220 m²/K



TRAPUM Pulsar Searches

FBFUSE
Multi-beam
beamformer

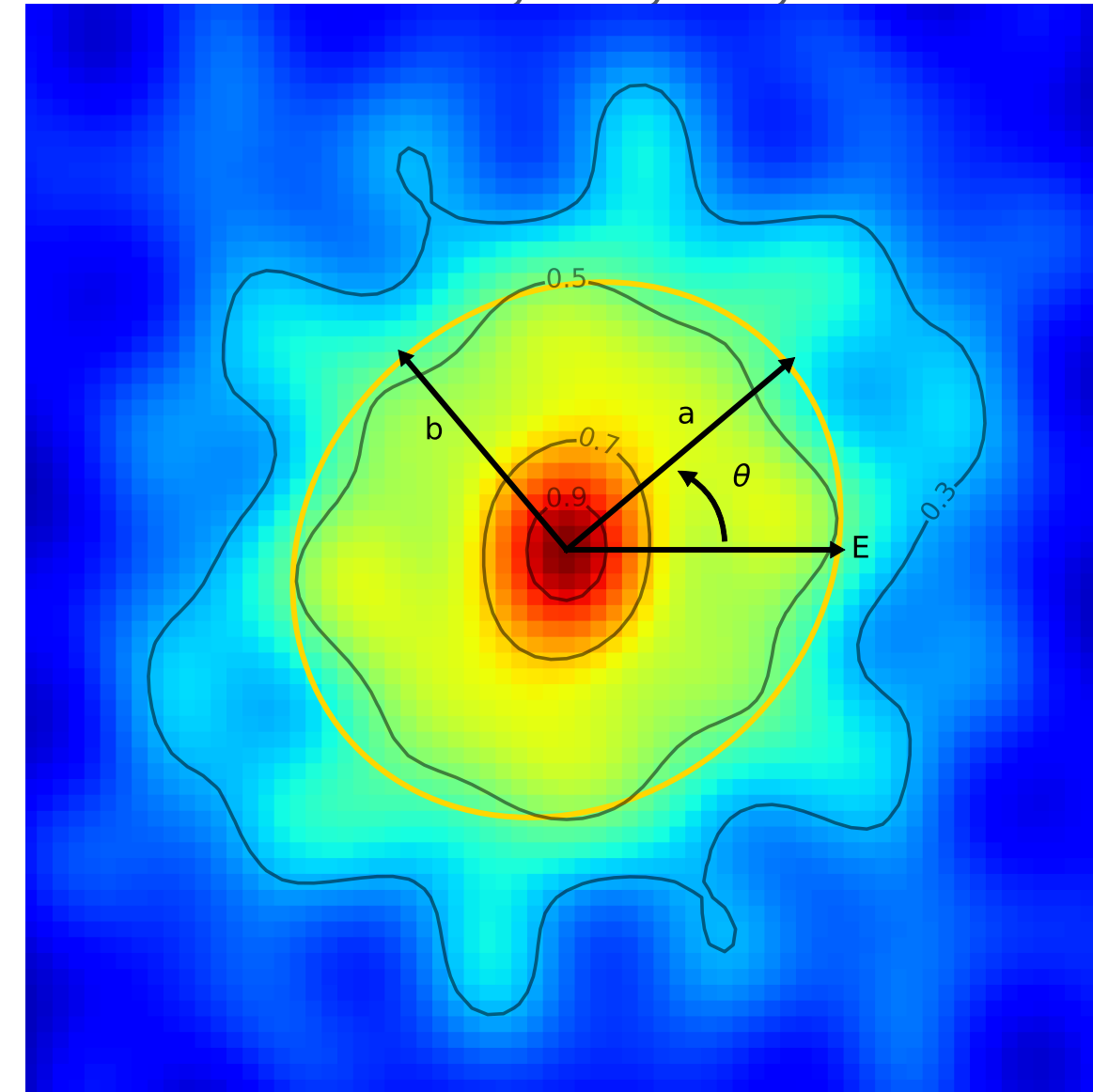
480 coherent
beams

APSUSE
Binary pulsar search

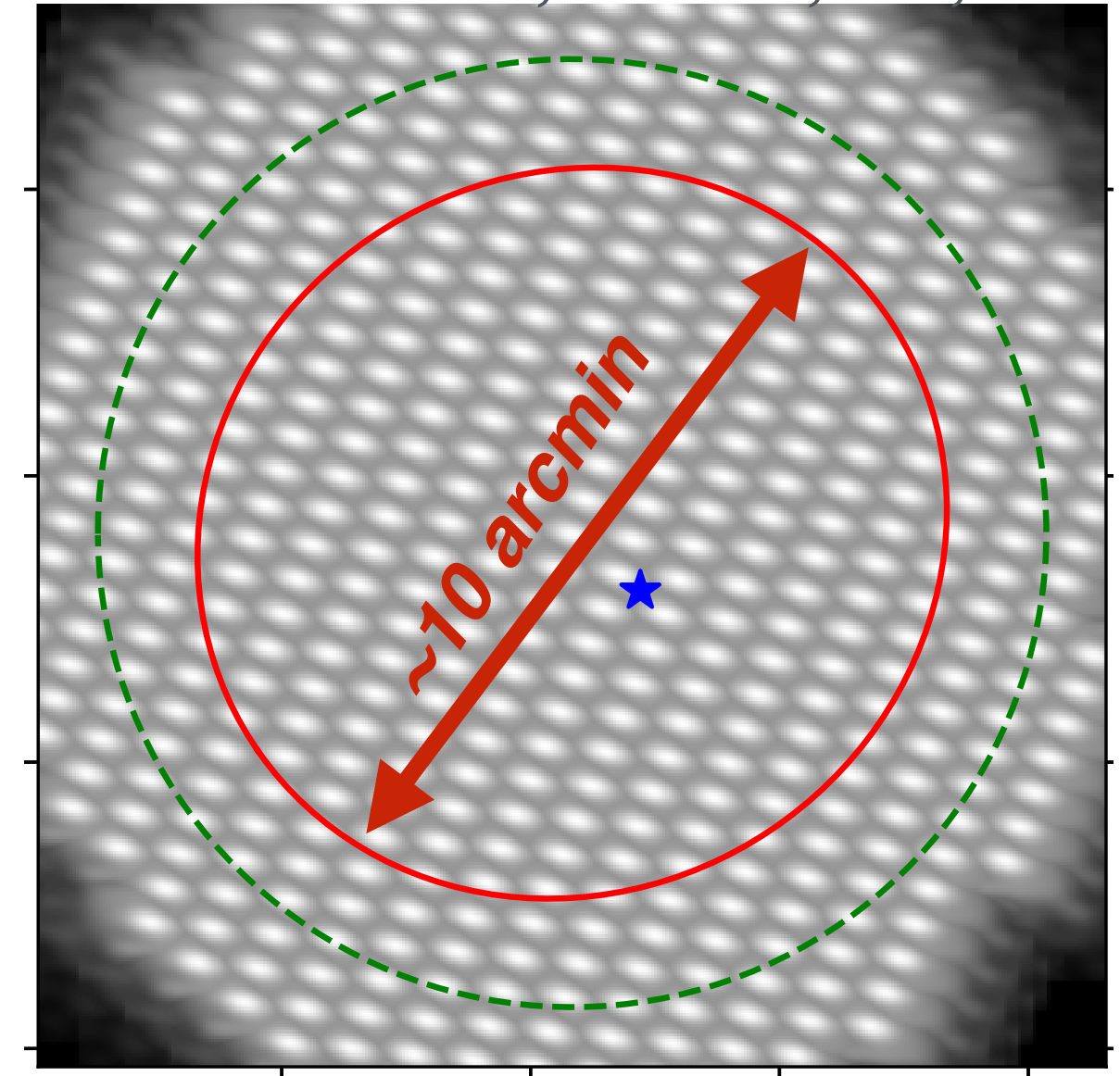
Folded
candidates
ranked by ML
classifier



Chen et al. 2021, JAI, 10, 2150013



C.J.C. et al. 2023, MNRAS, 519, 5590



**Wide field-of-view ($\sim 2^\circ$ at UHF)
x high angular resolution (\sim few arcsec)
= lots of data!**

Project	Beams	Typical integration time (s)	Data volume (per hour)	Acceleration range (m/s/s)	Real-time ratio
Fermi	480	600	46 - 71 TB	+/- 50	x39
Nearby Galaxies	480	1800/7200	46 TB	+/- 50	x54
MGPS-L	480	600	46 TB	+/- 15	x24

TRAPUM+Fermi Pulsar Searches

“Shallow” survey:

repeated 10min observations of ~180 unidentified sources at L-band and UHF

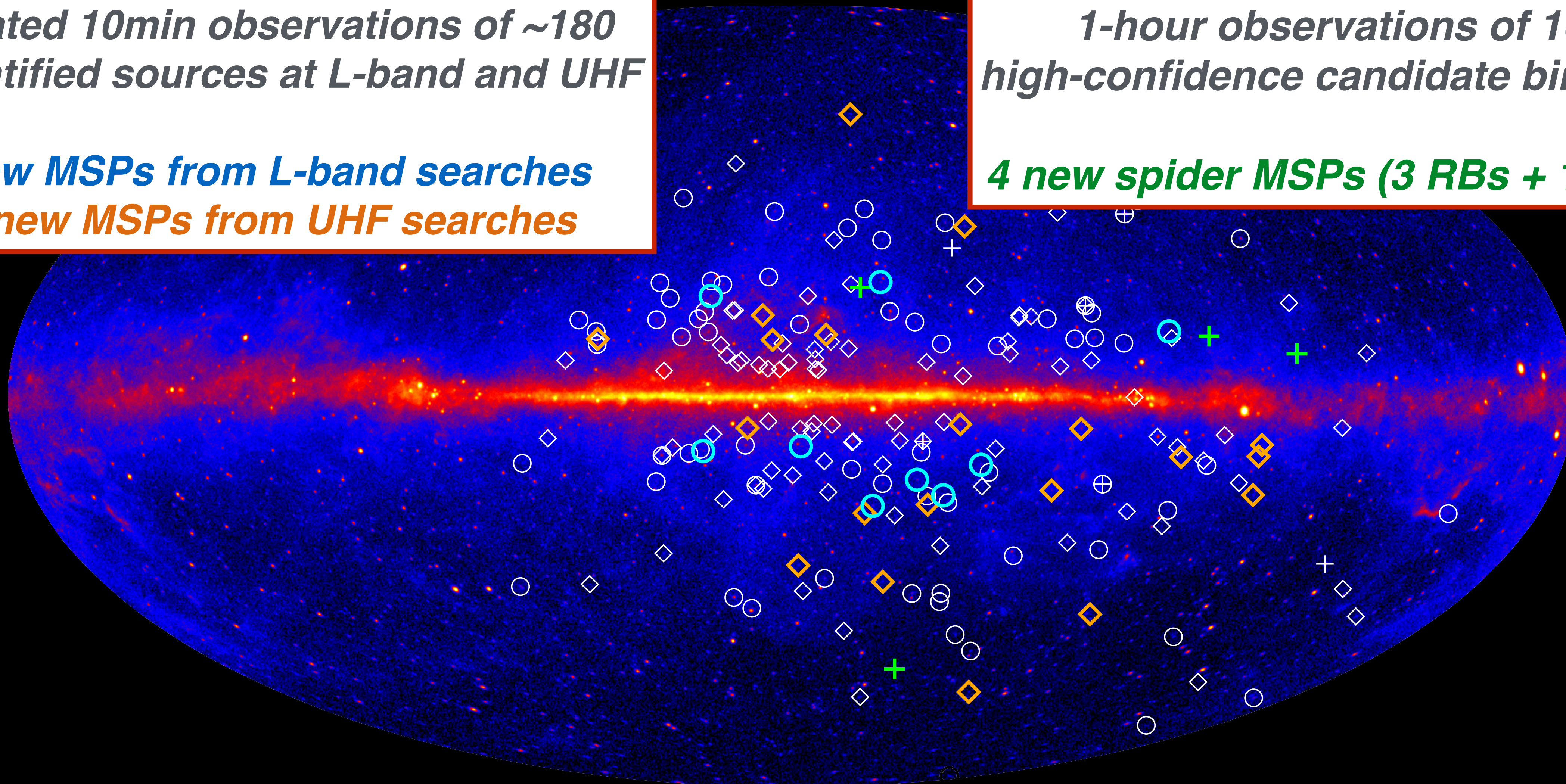
9 new MSPs from L-band searches

21 new MSPs from UHF searches

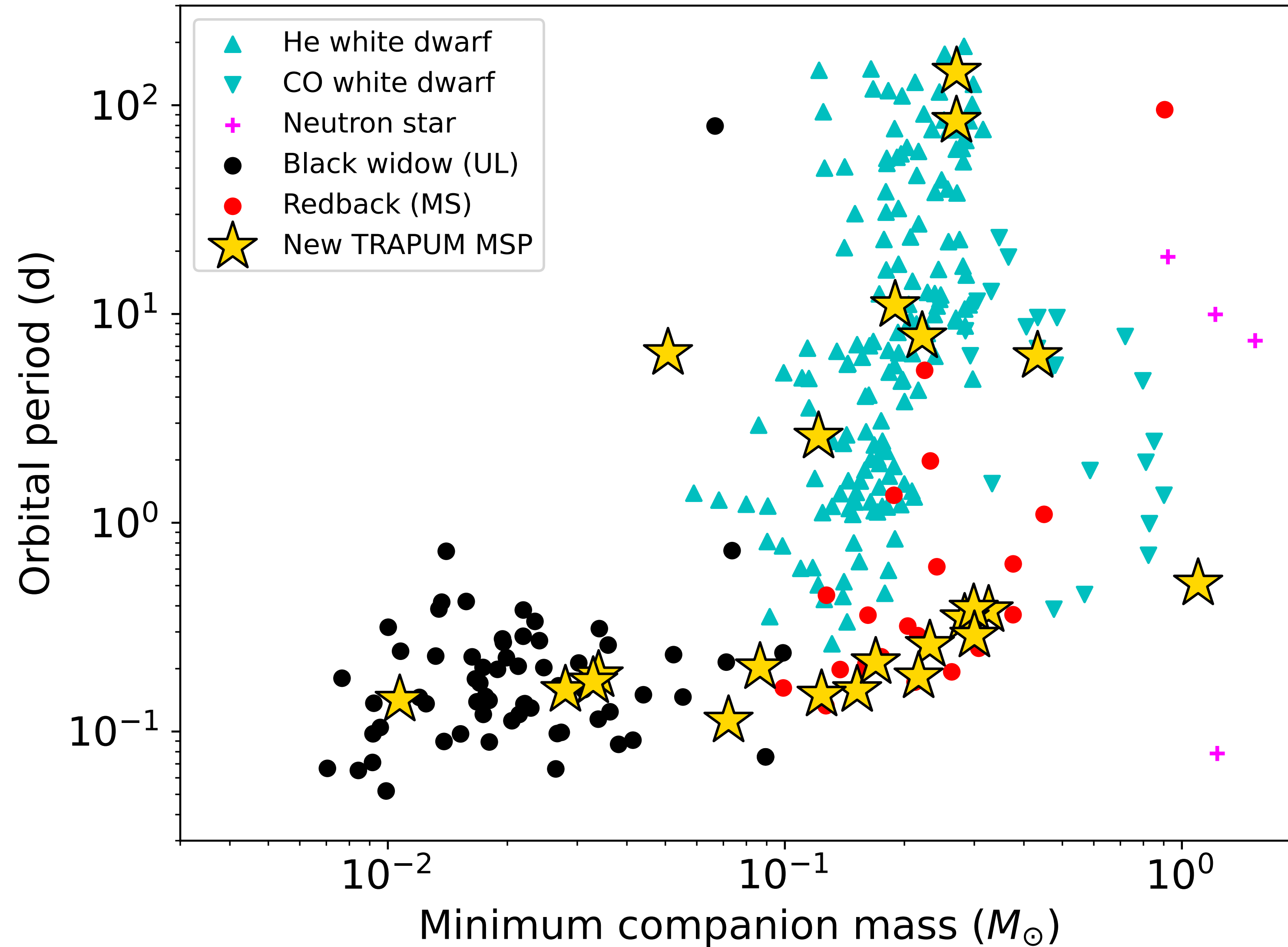
“Targeted” survey:

1-hour observations of 10 high-confidence candidate binaries

4 new spider MSPs (3 RBs + 1 BW)



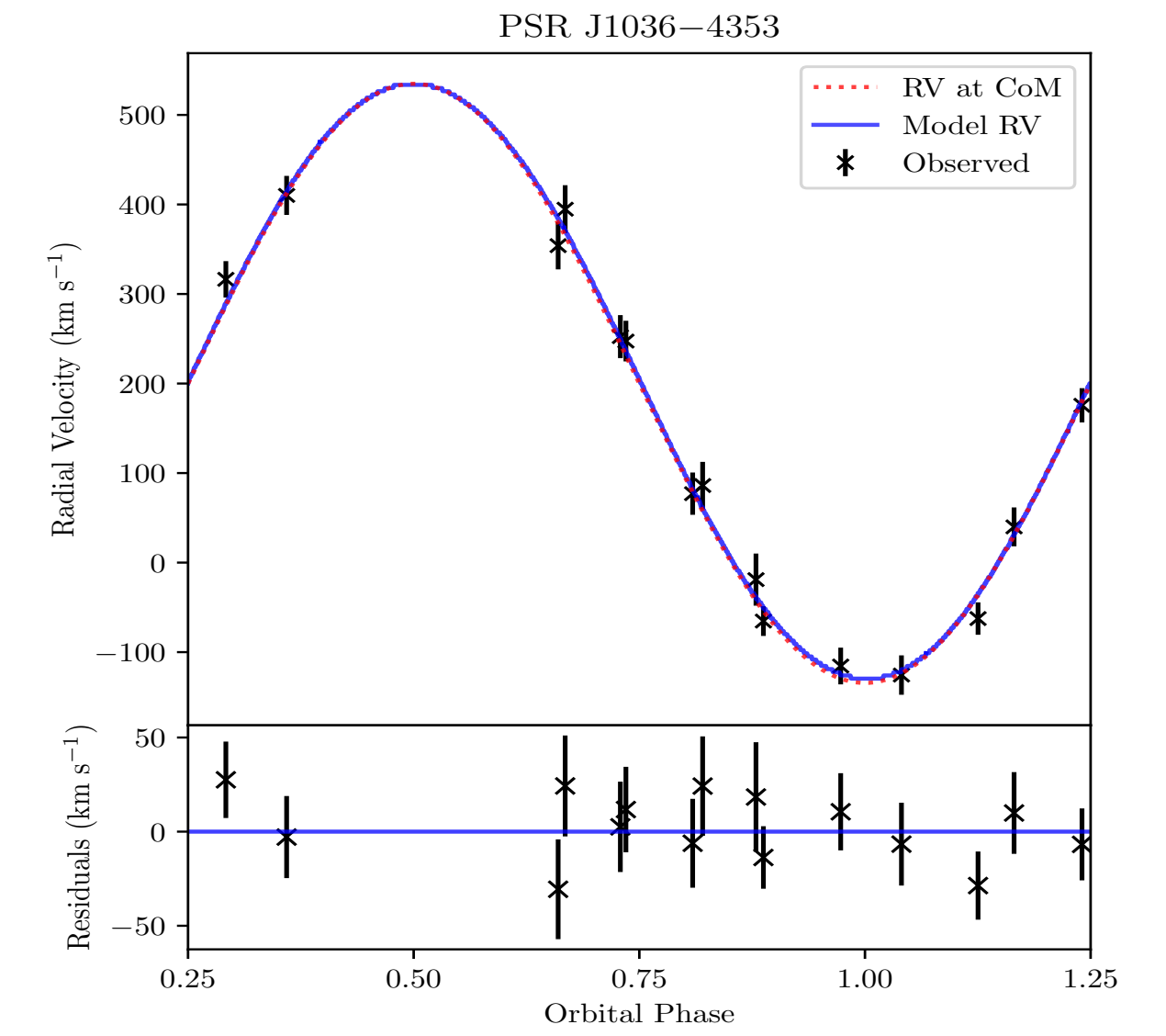
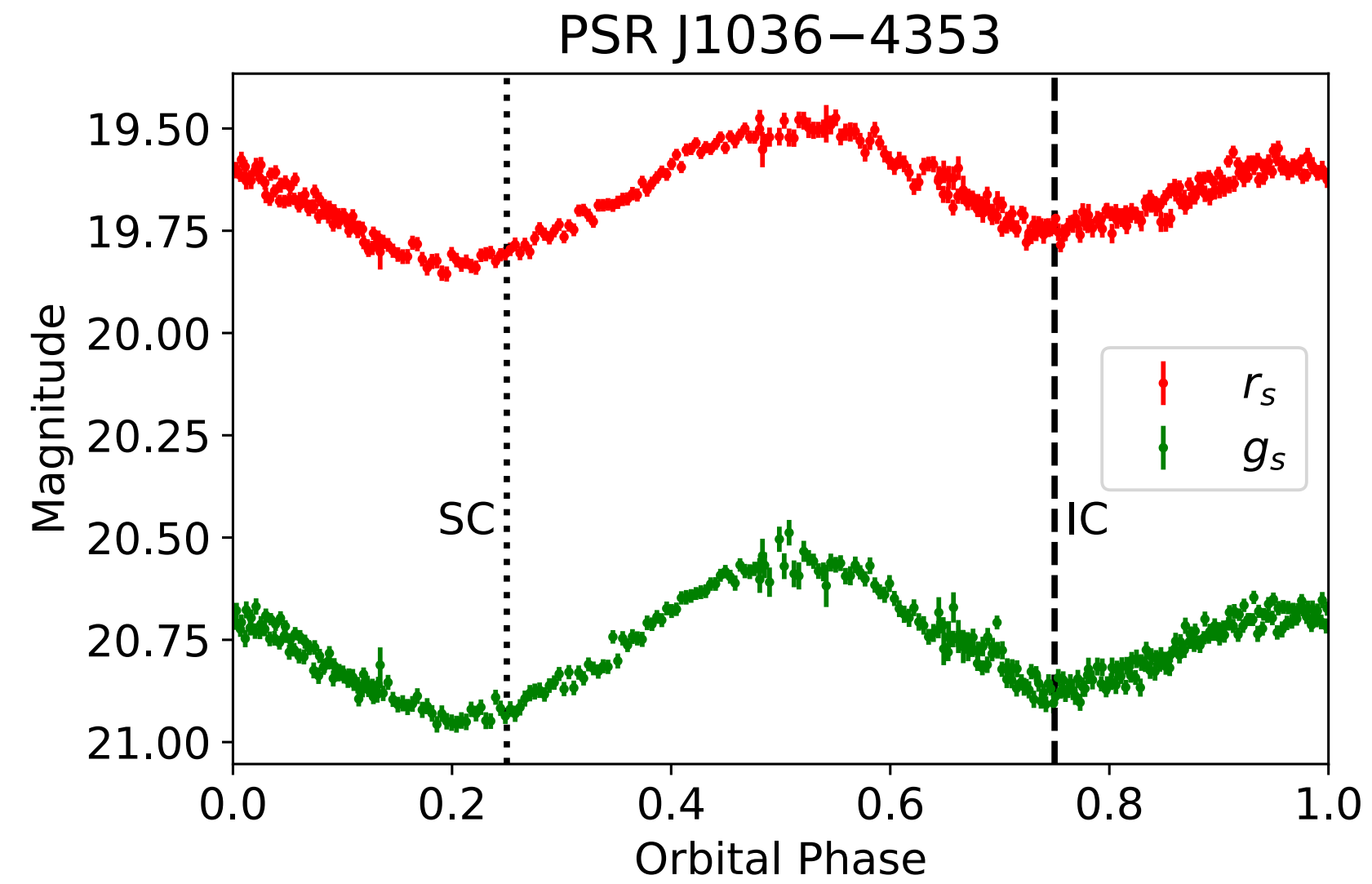
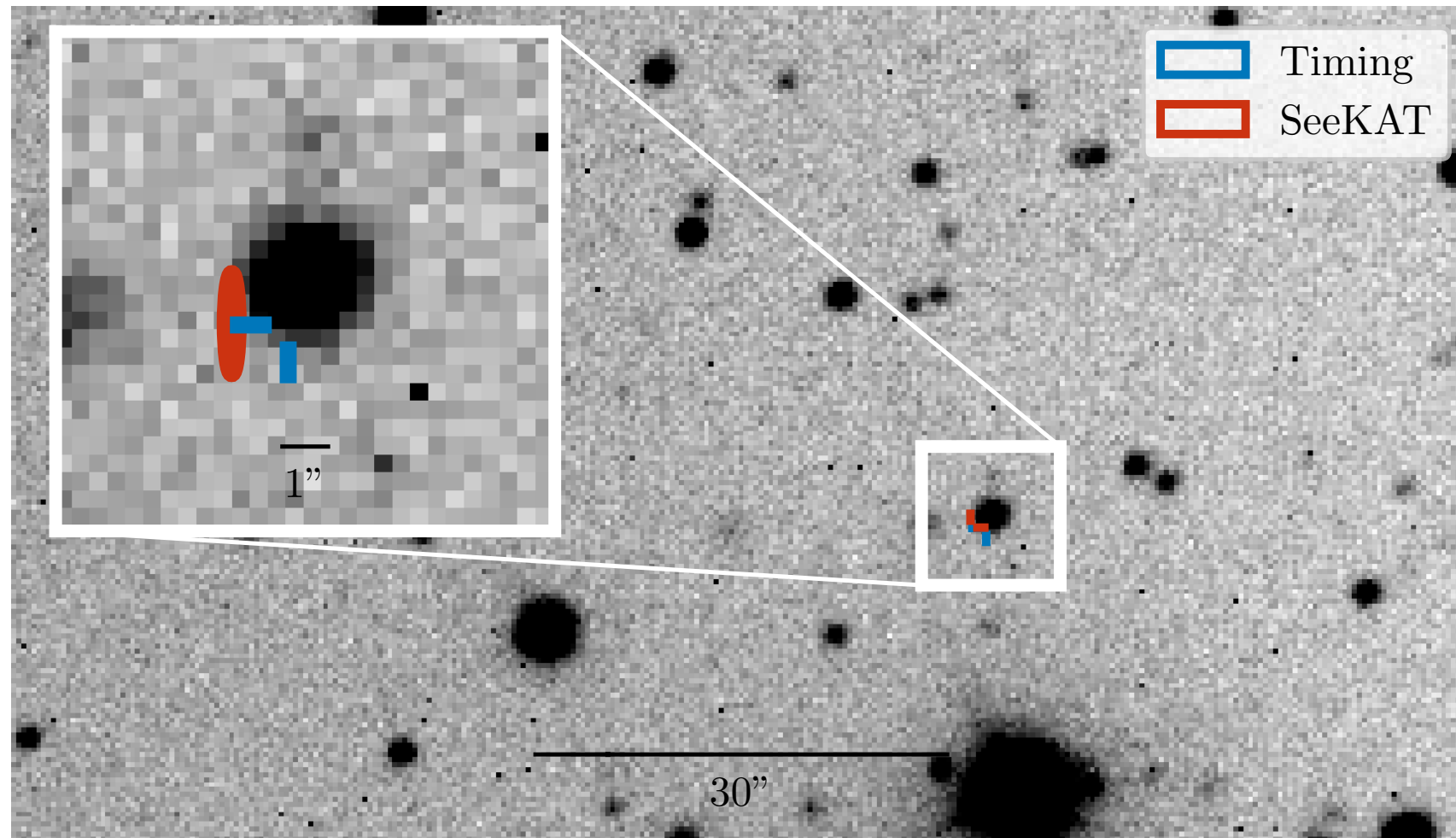
TRAPUM+Fermi discoveries



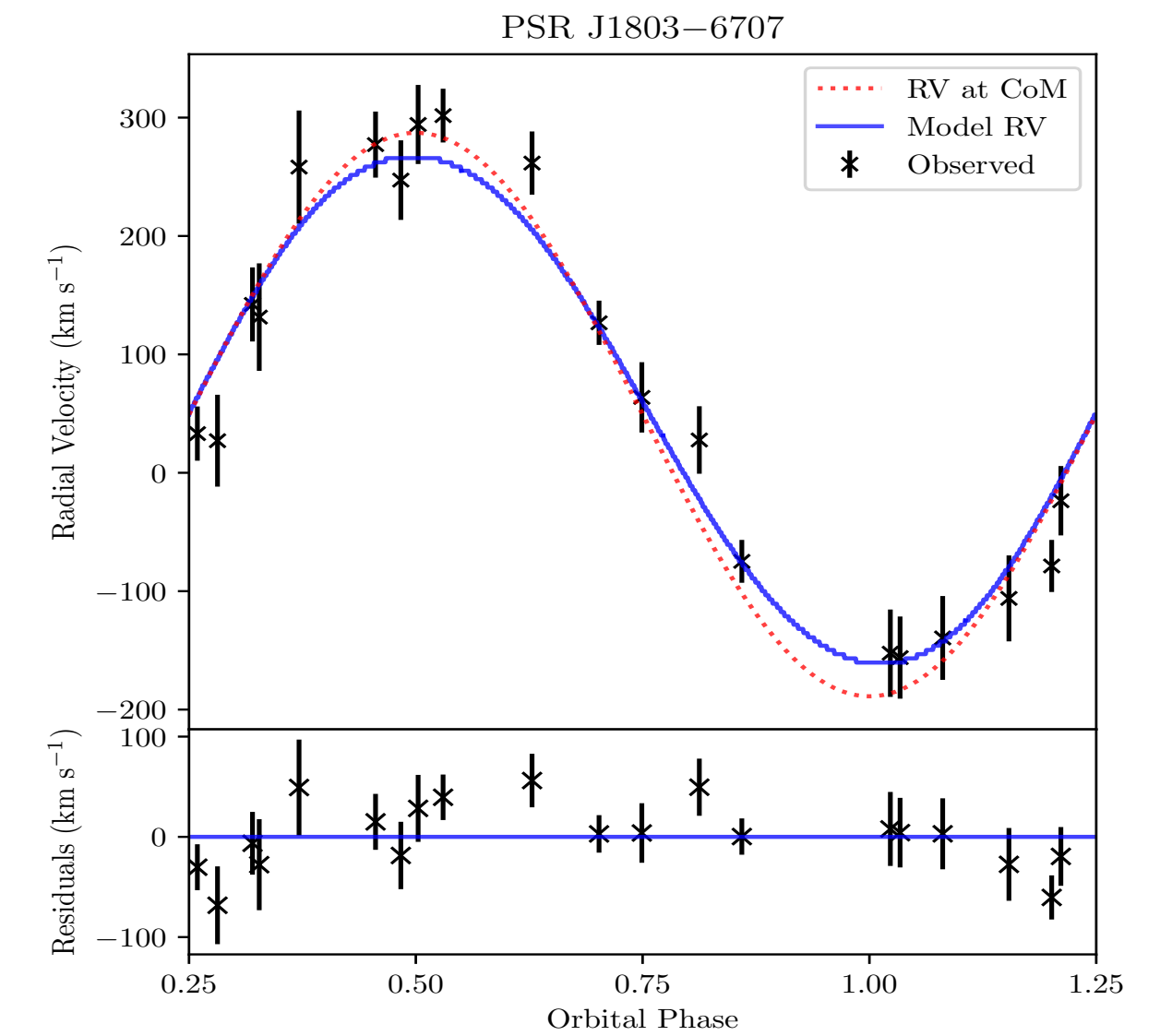
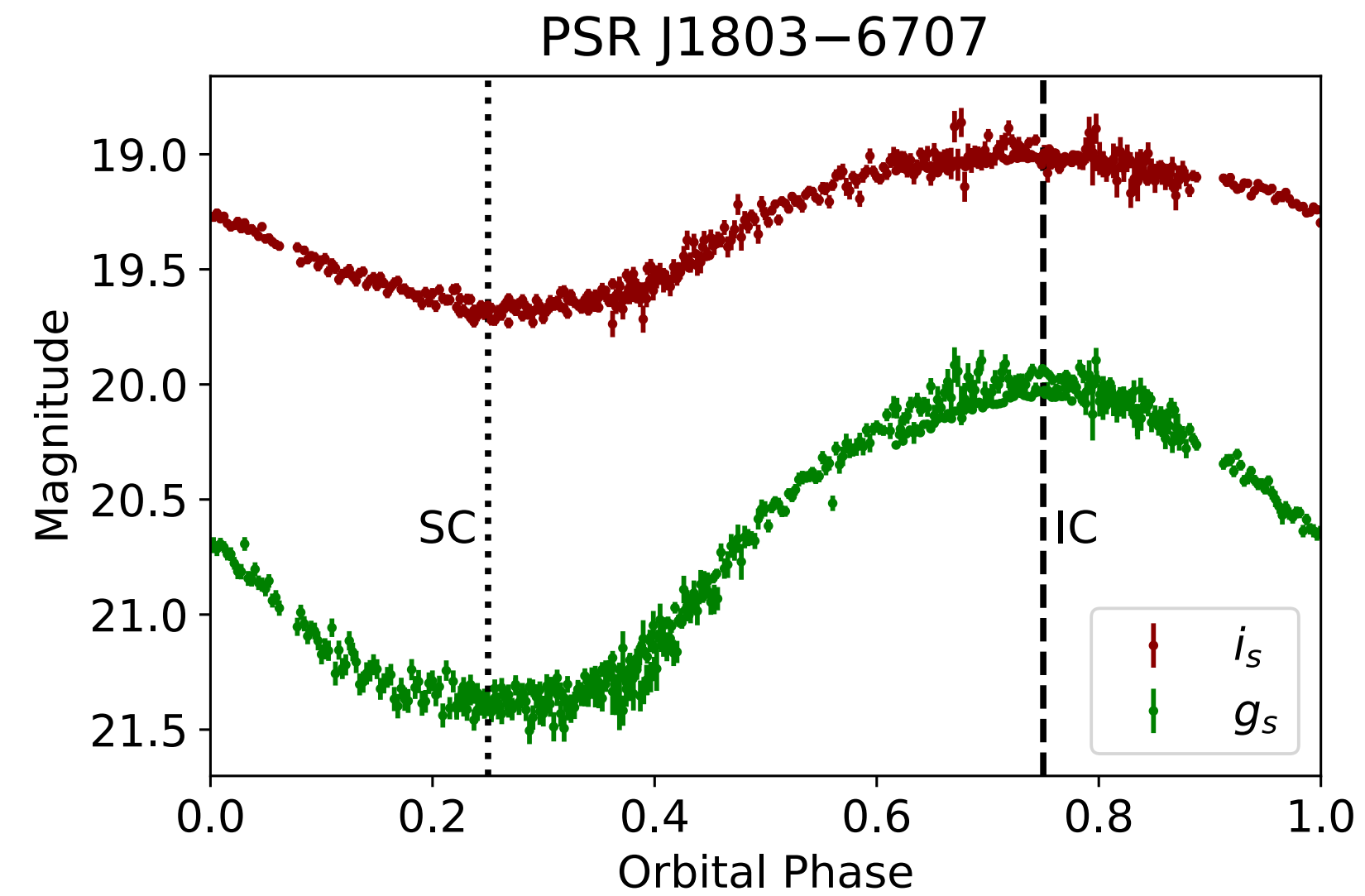
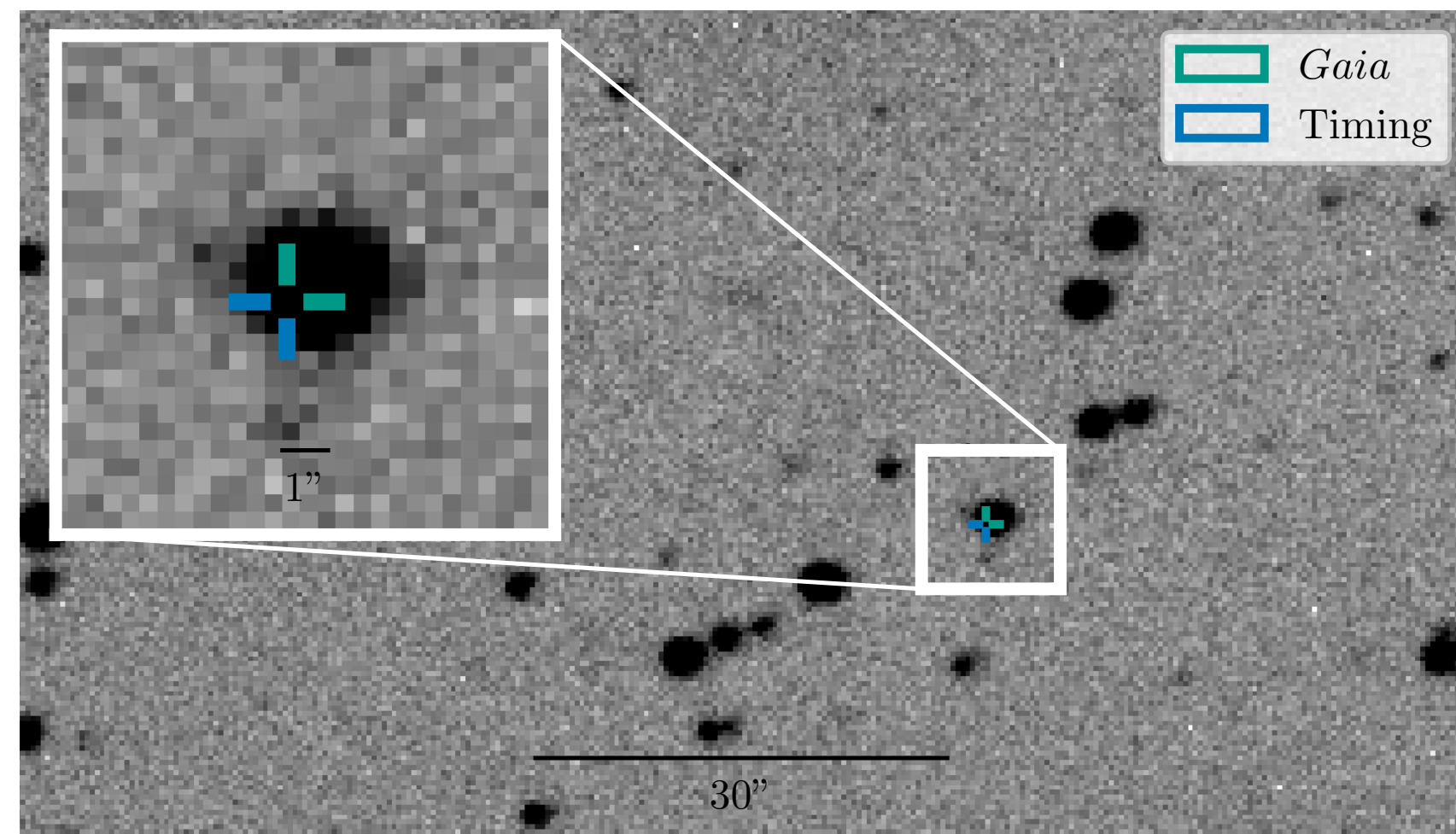
Multi-wavelength Follow-up

Phosrisom, A. et al., in prep

PSR J1036–4353



PSR J1803–6707



Redback Optical Modelling

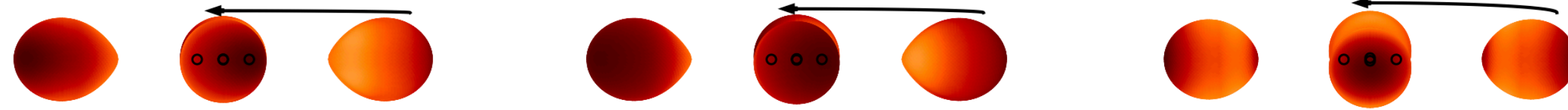
Phosrisom, A. et al., in prep

Radius variations, a “failed” tMSP?

Convection, Pre-IGD

D+C, Pre-IGD

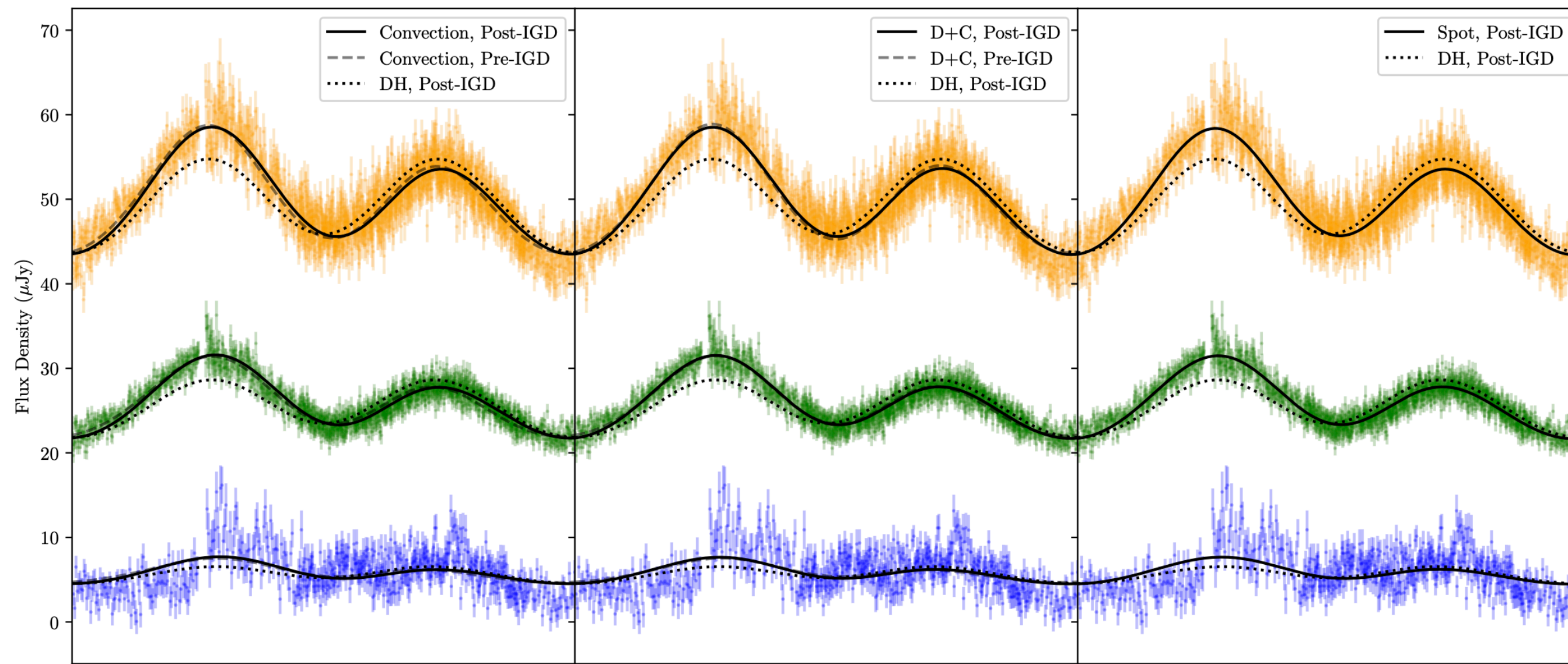
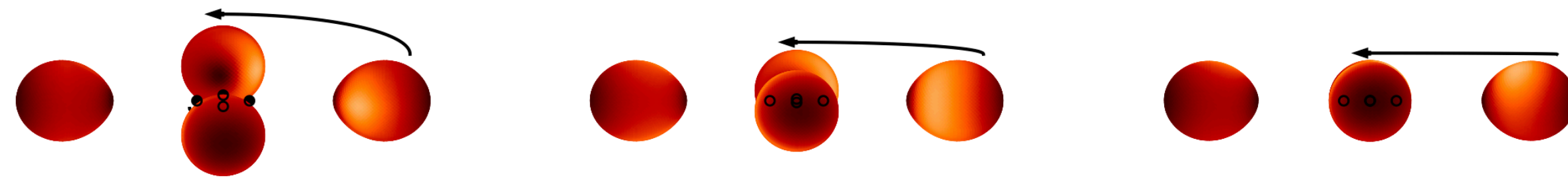
DH, Post-IGD



Convection, Post-IGD

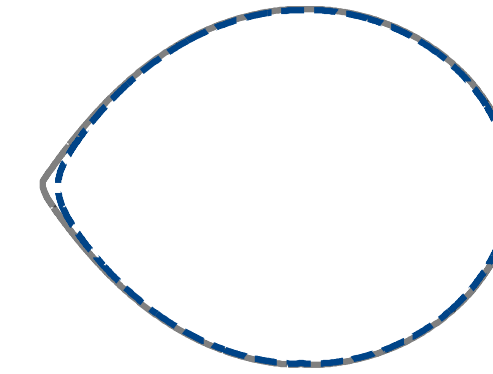
D+C, Post-IGD

Spot, Post-IGD



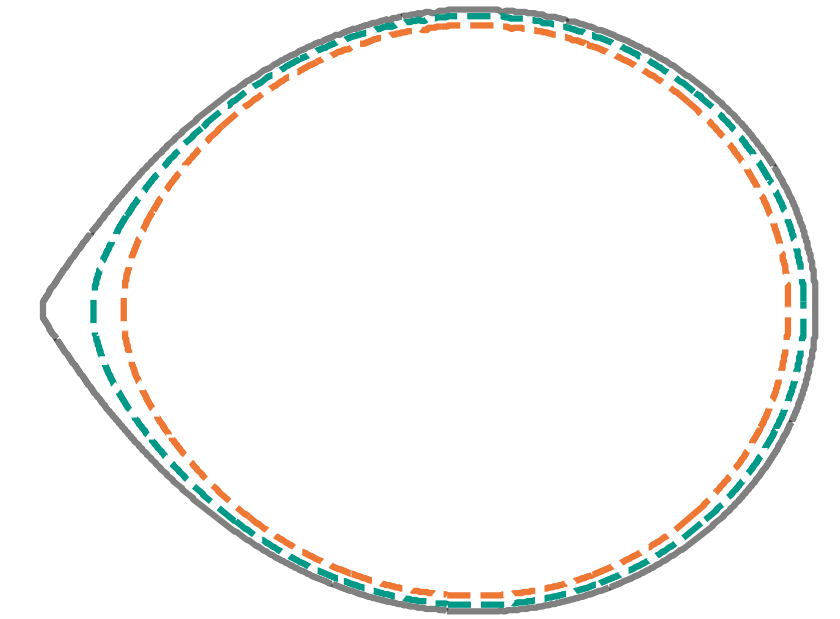
PSR J1036–4353
 $f_{VA}=0.992$

• ×



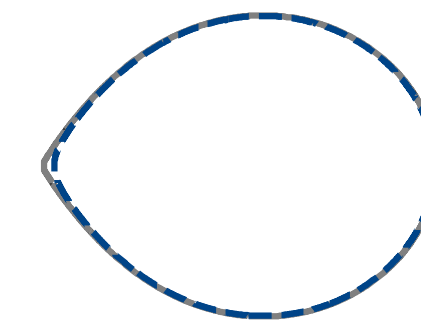
PSR J1803–6707
 $f_{VA}=0.936$ (June 2021)
 $f_{VA}=0.972$ (July 2021)

• ×



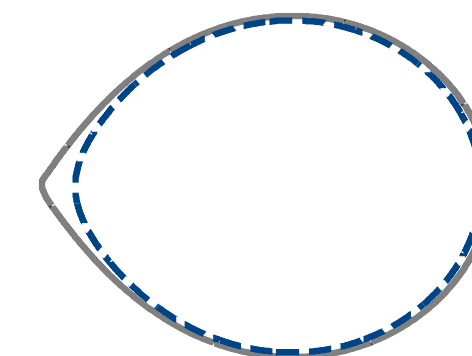
PSR J1023+0038
 $f_{VA}=0.994$

• ×



PSR J1227–4853
 $f_{VA}=0.966$

• ×

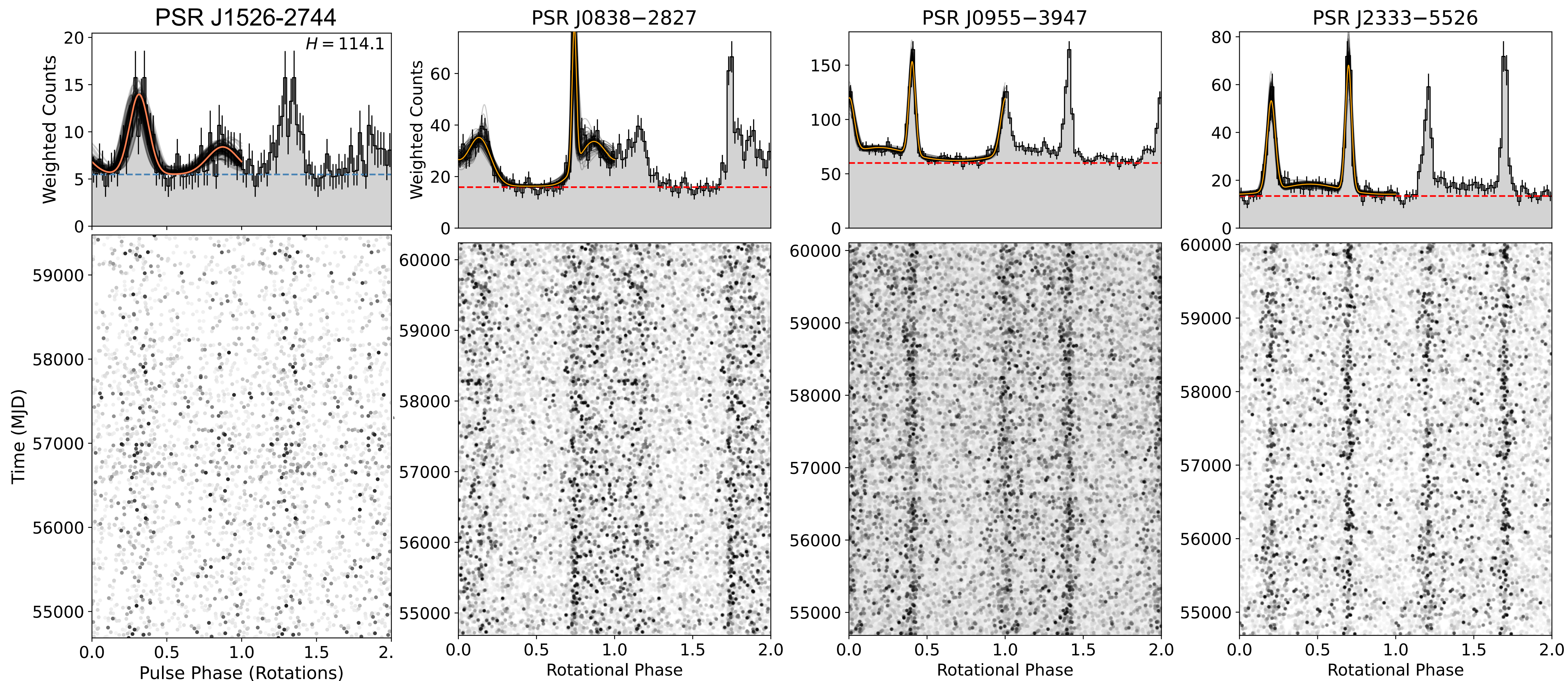


— June 2021
— July 2021
- - - Perimeter
— Roche Lobe
• Pulsars
× Barycentre

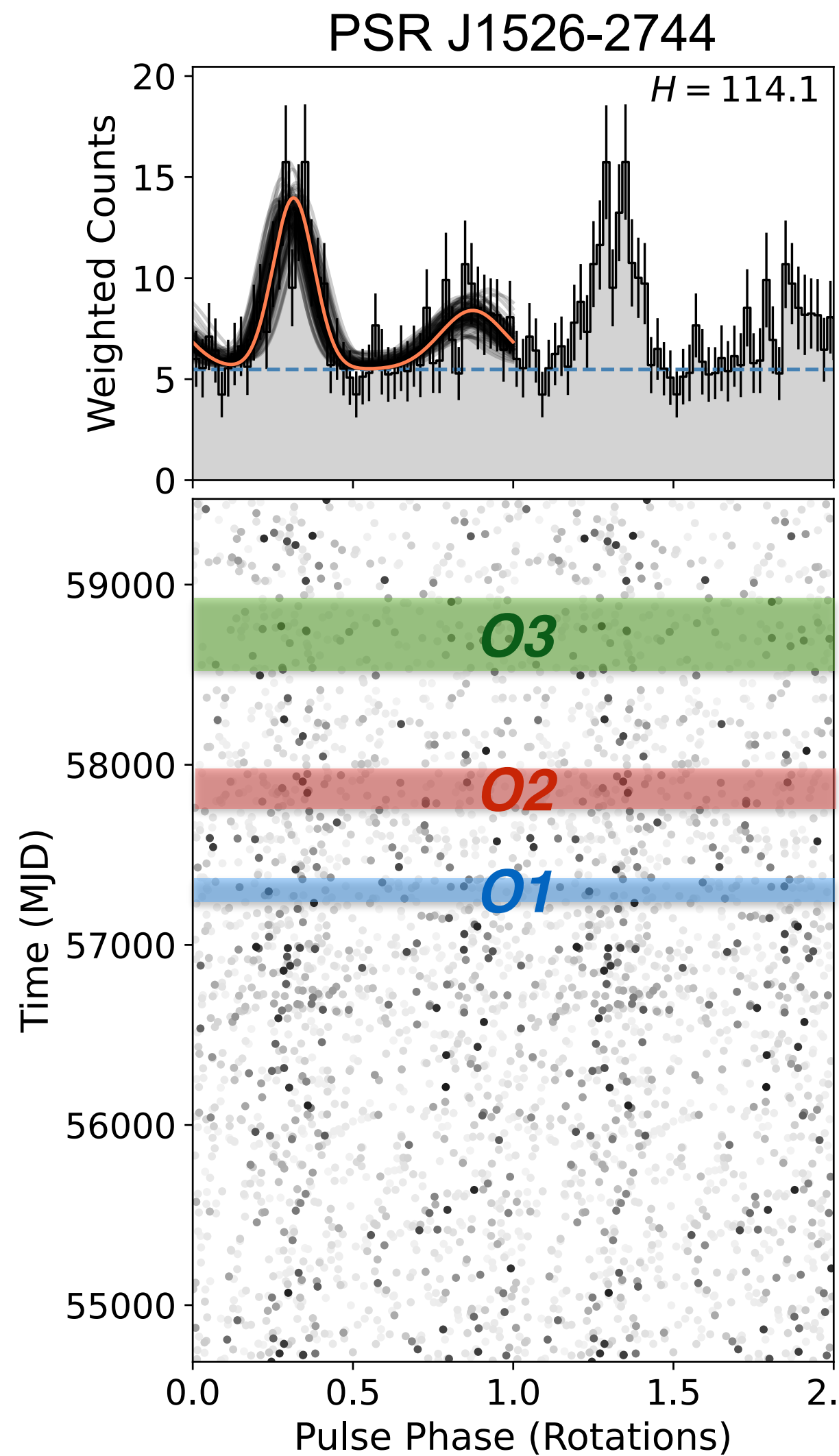
0 1 2 3 4
Distance (R_{\odot})

Gamma-ray Pulsar Timing

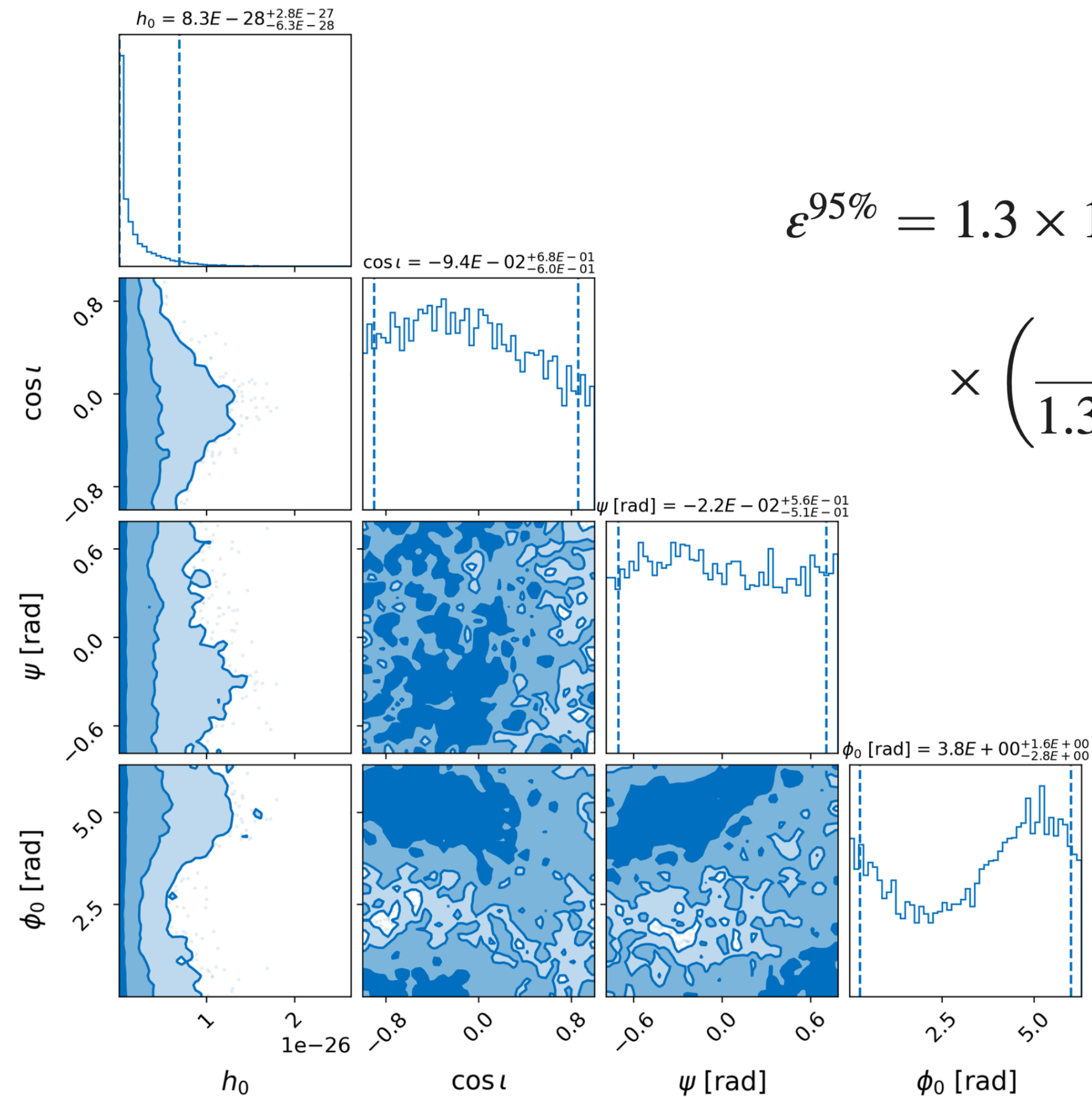
16-year Fermi-LAT all-sky data provides long-term timing solutions



Gamma-ray Pulsar Timing



Retrospective gamma-ray timing solutions enable searches for continuous gravitational waves in public LIGO data



$$\epsilon^{95\%} = 1.3 \times 10^{-8} \left(\frac{h_0^{95\%}}{6.7 \times 10^{-27}} \right) \times \left(\frac{d}{1.3 \text{ kpc}} \right) \left(\frac{803.5 \text{ Hz}}{f} \right)^2 \left(\frac{10^{38} \text{ kg m}^2}{I_{zz}} \right)$$

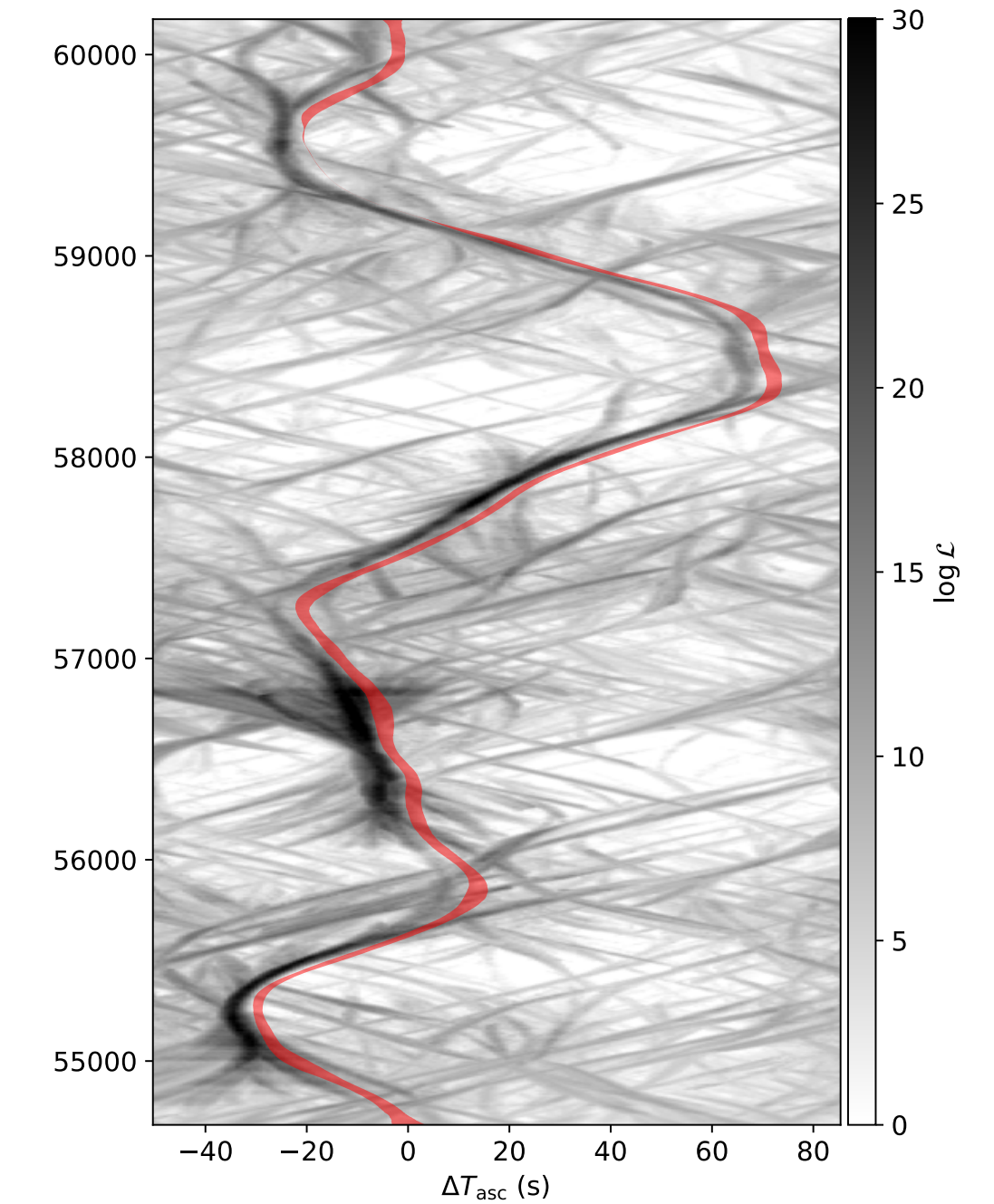
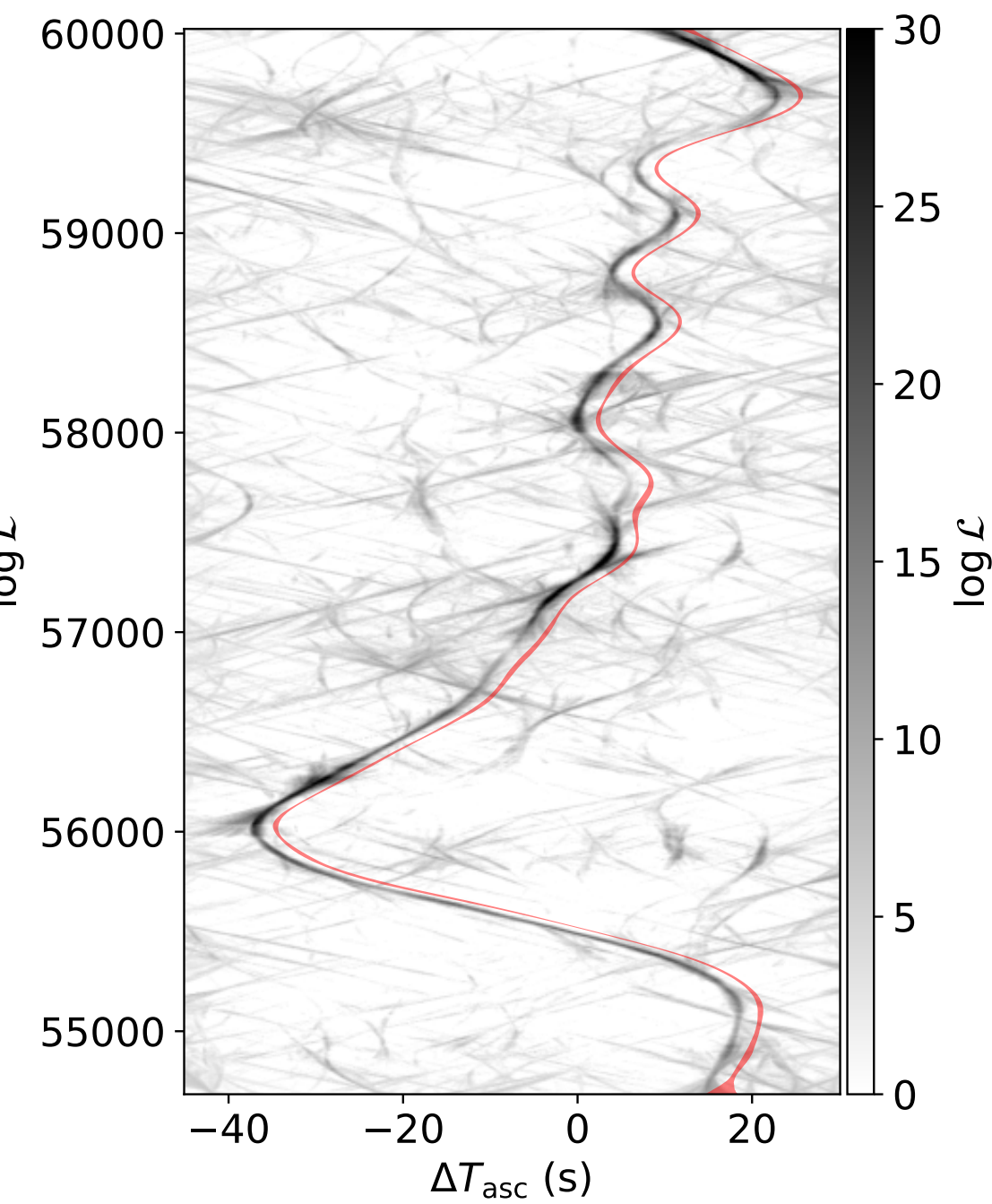
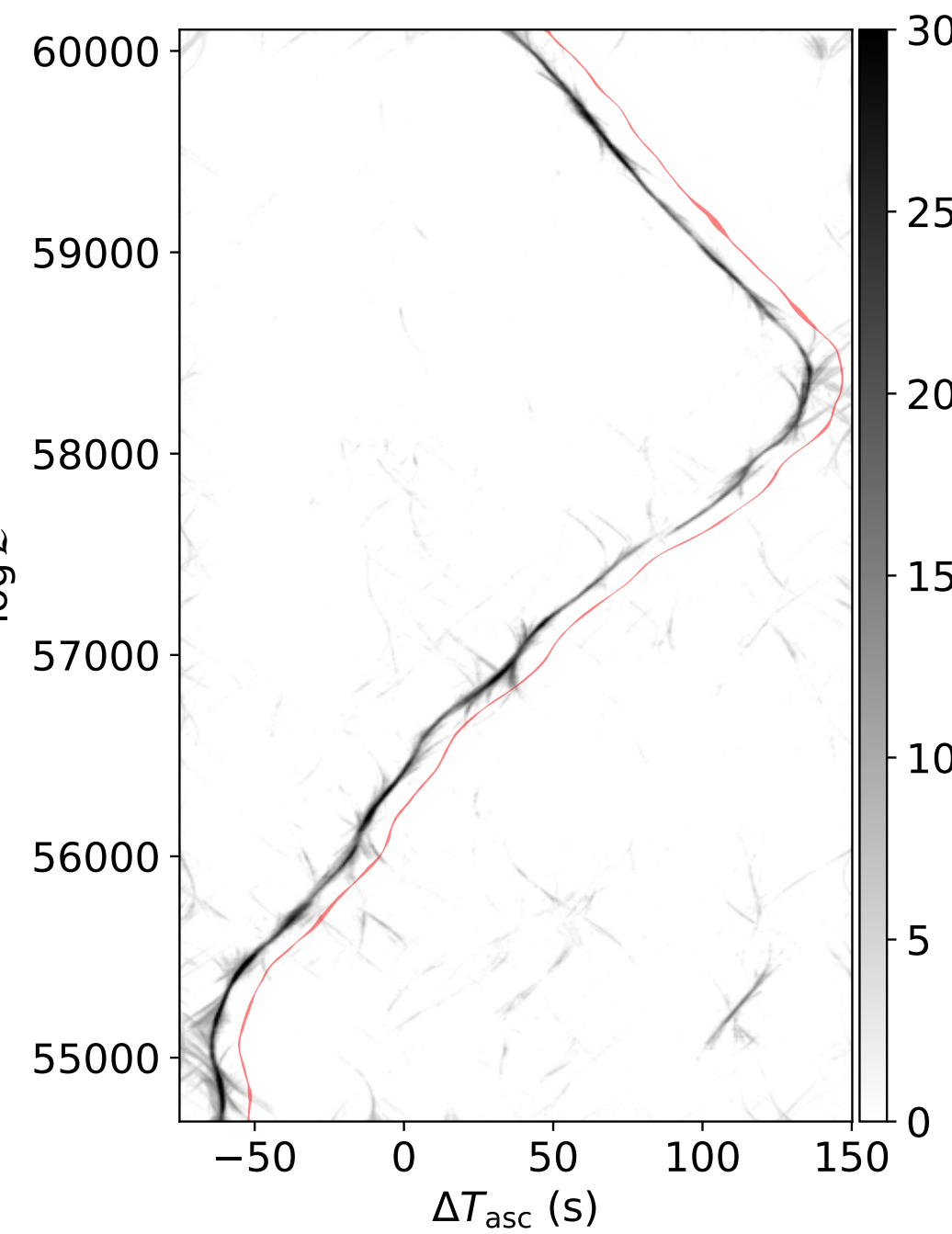
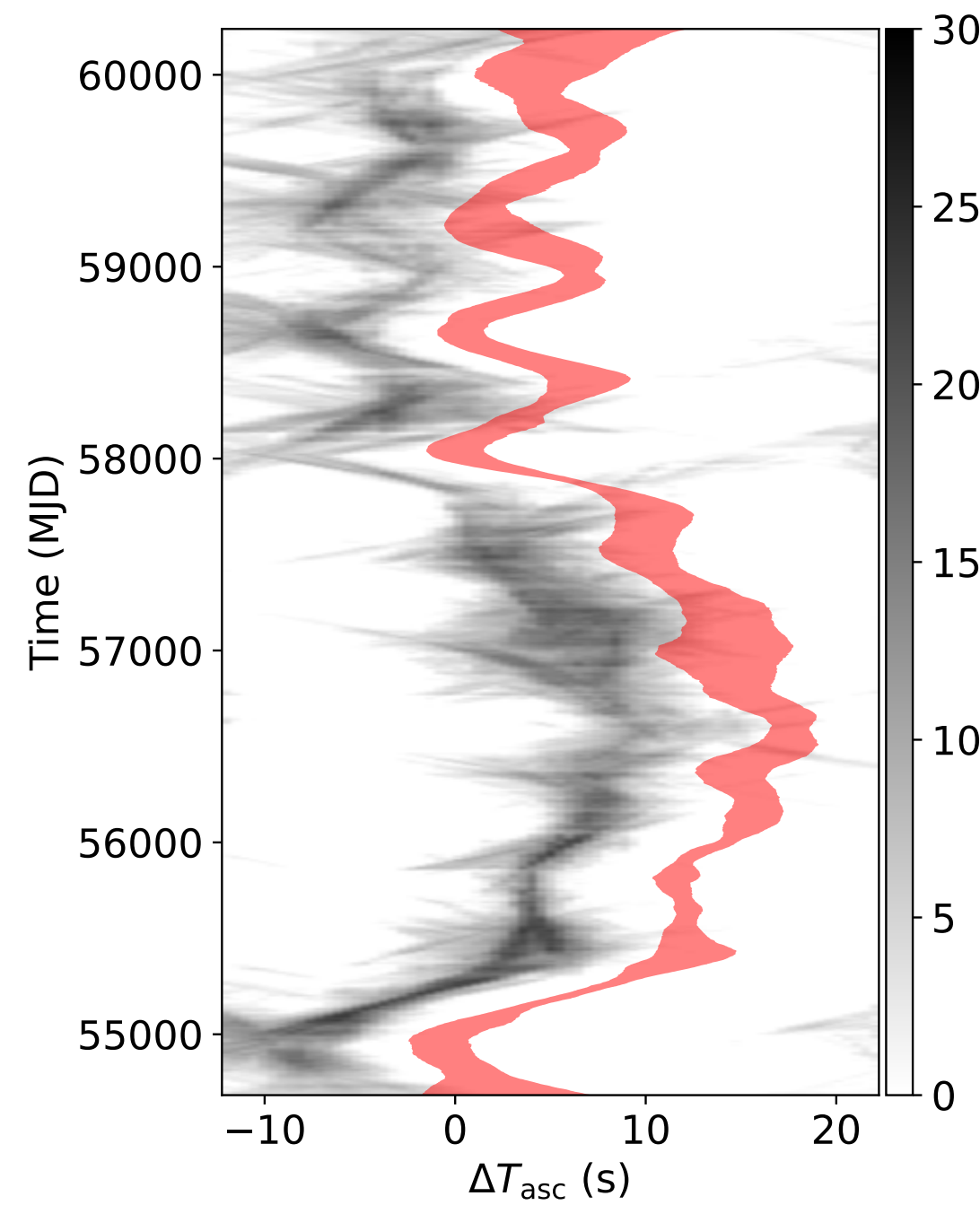
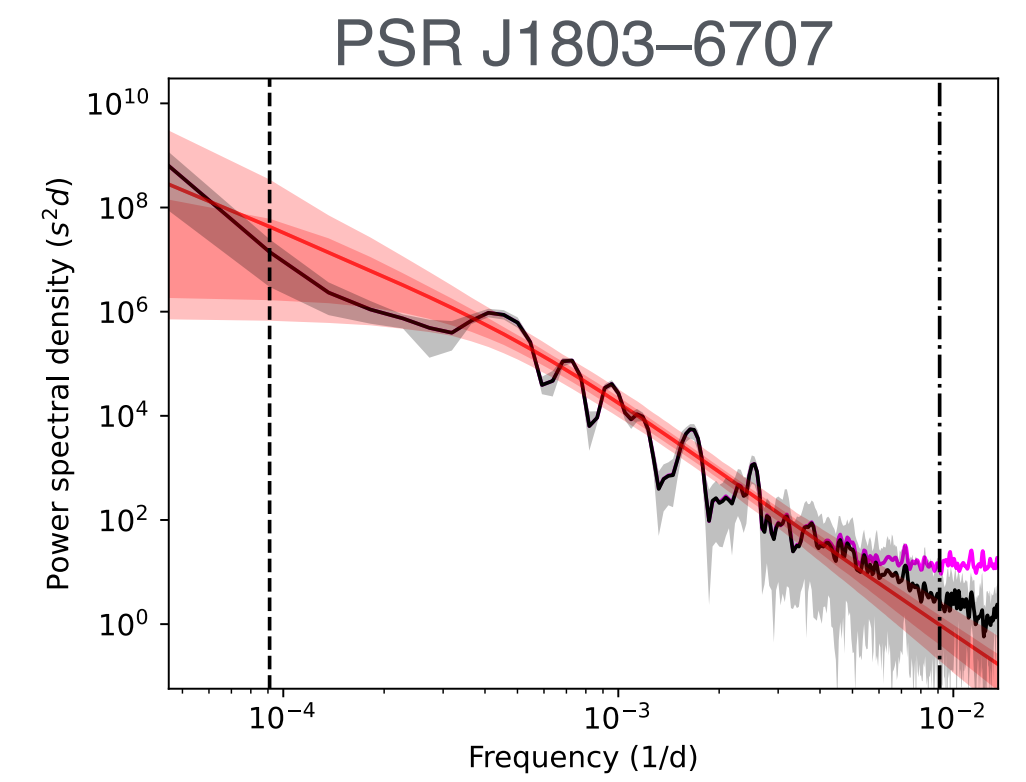
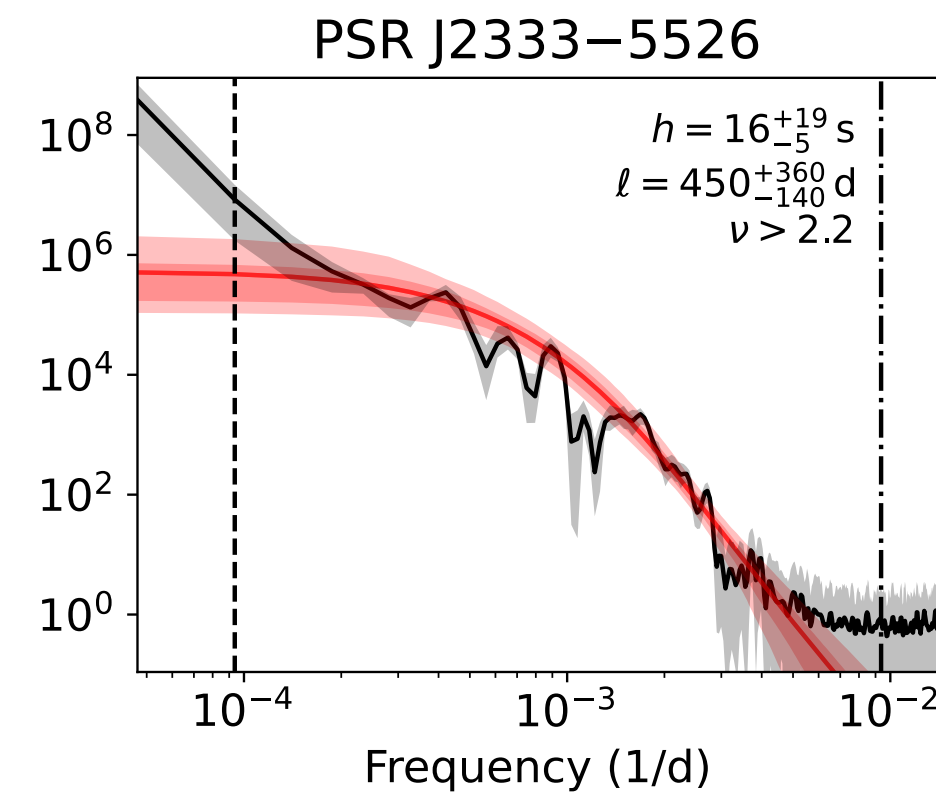
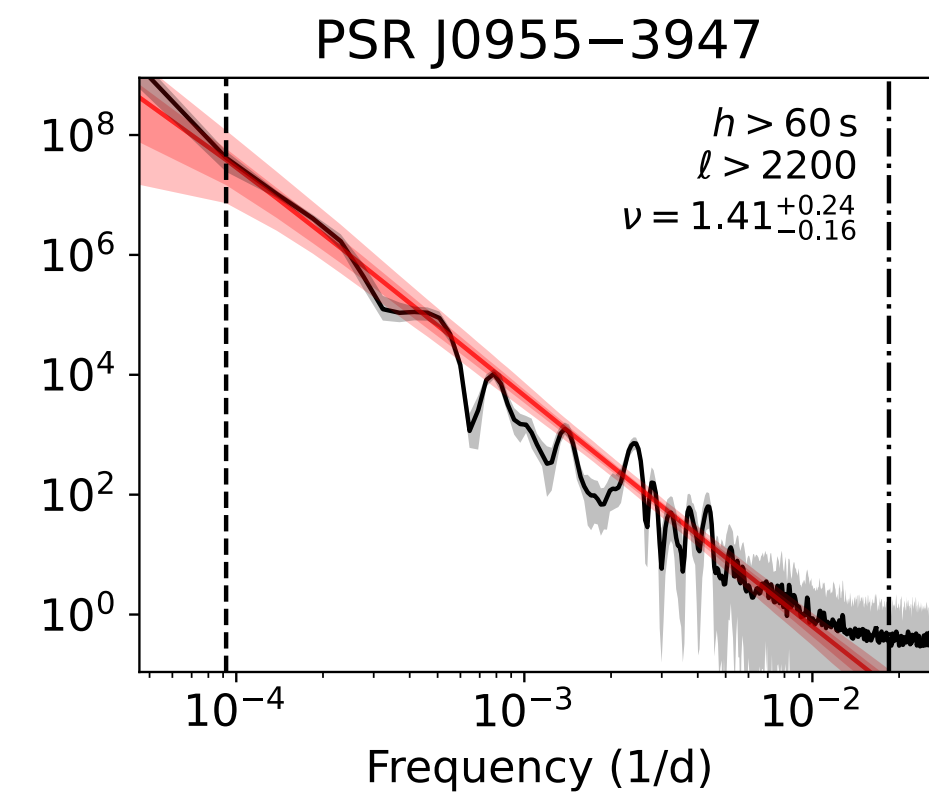
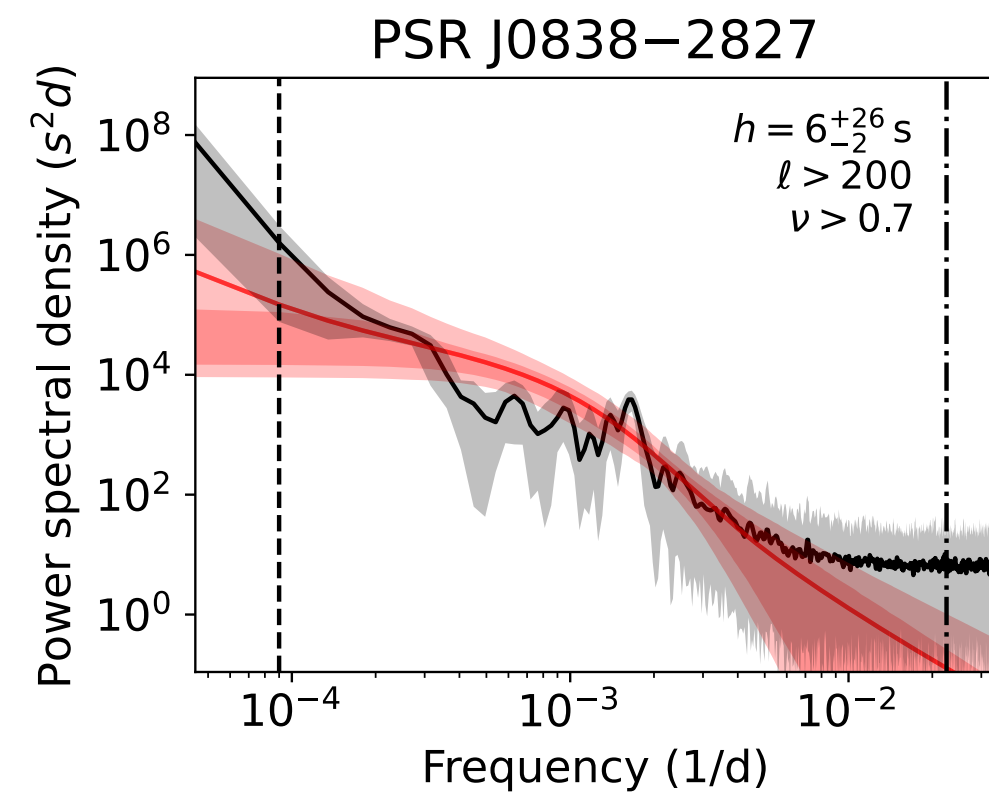
$$h_0^{95\%} \approx 9.2 \times h_0^{\text{sd}}$$

Ashok, A. et al. 2024, Phys. Rev. D 109, 104002

Redback Orbital Period Variations

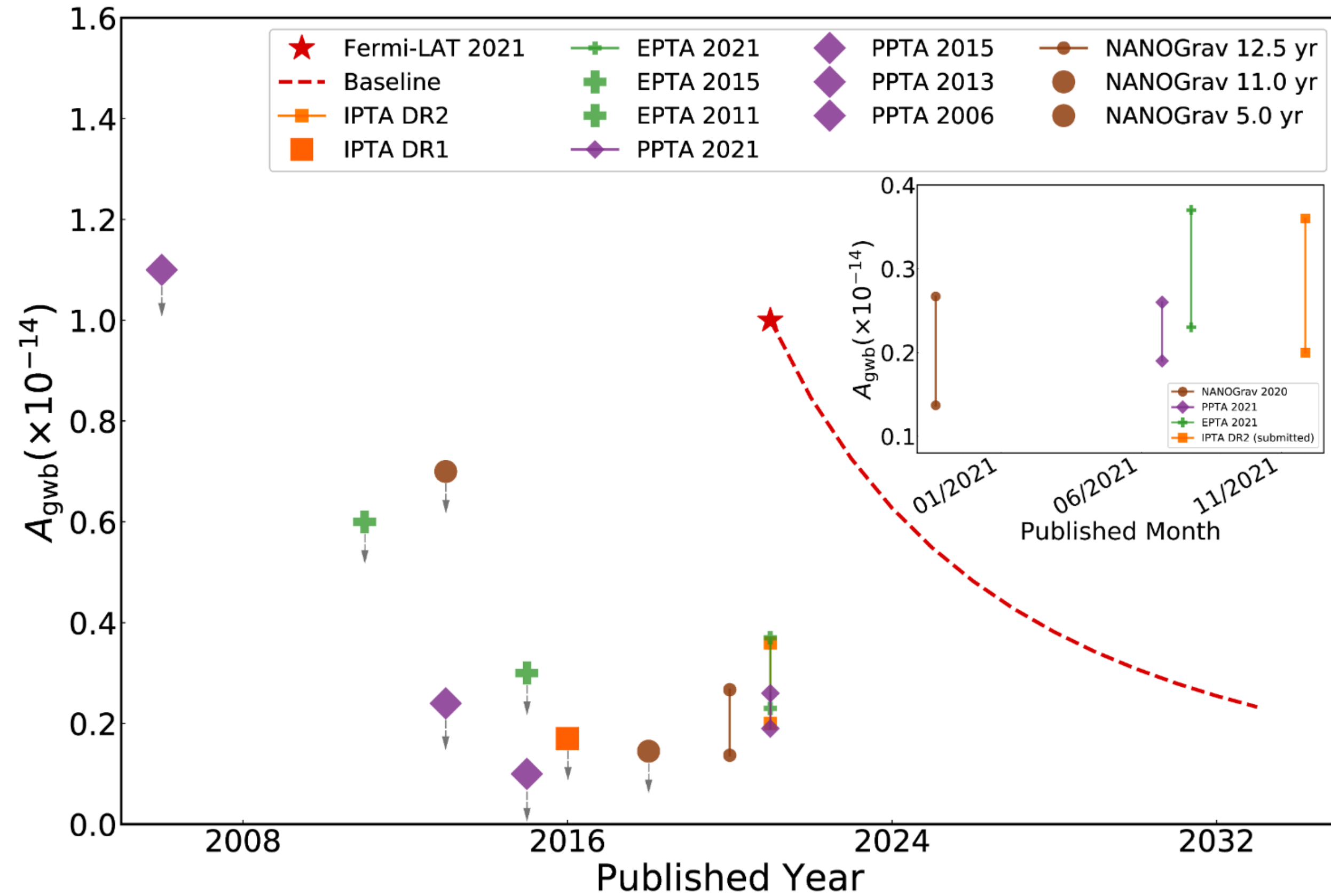
Thongmeearkom, T. et al., 2024, MNRAS, 530, 4676

Burgay, M. et al., A&A, submitted

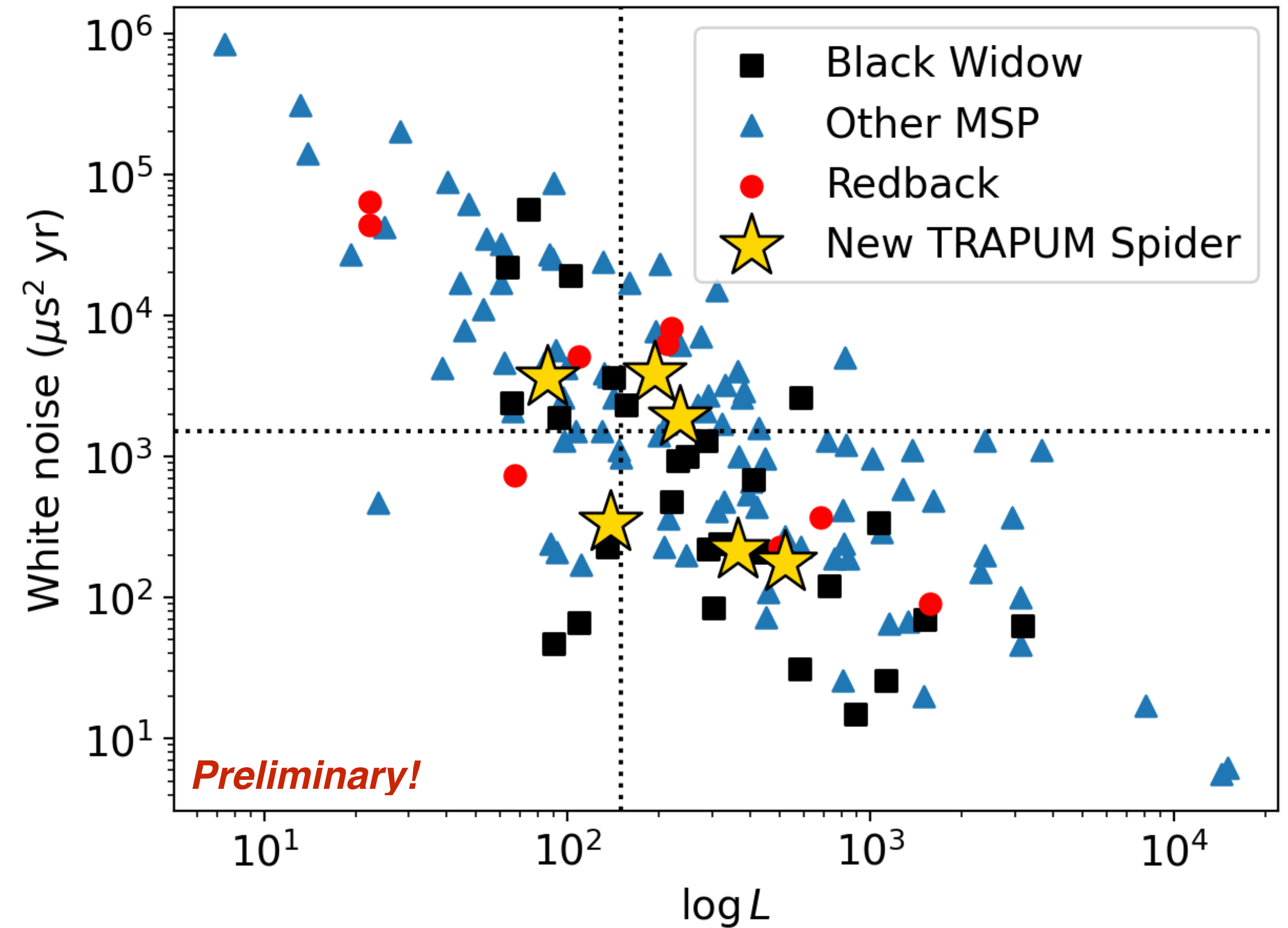


The Gamma-ray Pulsar Timing Array

The Fermi-LAT collaboration, 2022, Science, 376, 521



GPTA DR2, in prep!



Summary

- Fermi-LAT gamma-ray data provides a treasure map for new MSPs
- 34 new MSPs from TRAPUM searches of Fermi-LAT sources
 - Many new spiders, mass estimates for 6 new redbacks
 - [C.J.C. et al. 2023, MNRAS, 519, 5590](#)
 - [Dodge, O. et al., 2024, MNRAS, 528, 4337](#)
 - [Thongmearkom, T., et al. 2024, MNRAS, 530, 4676](#)
- Radio + gamma-ray MSP discoveries enable lots of NS science:
 - Targeted CW searches, even in old LIGO data, for new MSPs
 - Equation-of-state constraints from new pulsar mass measurements
 - Pulsar timing array projects to constrain the nano-Hertz GWB

Thank you for listening!