# The <sup>12</sup>C( $\alpha$ , $\gamma$ )<sup>16</sup>O reaction

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# Motivation



- Together with the 3 $\alpha$  process, the <sup>12</sup>C( $\alpha,\gamma$ )<sup>16</sup>O reaction determines the <sup>12</sup>C/<sup>16</sup>O ratio in the universe.
- For stellar evolution, the <sup>12</sup>C/<sup>16</sup>O ratio determines the evolution of massive stars, which in turn effects all later stages of nucleosynthesis.

TABLE I. Astrophysical environments and burning stages where the  ${}^{12}C(\alpha, \gamma){}^{16}O$  reaction plays an important role. The temperatures of these environments dictate the energy ranges where the  ${}^{12}C(\alpha, \gamma){}^{16}O$  cross section must be well known for an accurate calculation of the reaction rate.

Burning stages	Astrophysical sites	Temperature range (GK)	Gamow energy range (MeV)
Core helium burning	AGB stars and massive stars	0.1-0.4	0.15-0.65
Core carbon and oxygen burning	Massive stars	0.6–2.7	0.44–2.5
Core silicon burning	Massive stars	2.8-4.1	1.1–3.4
Explosive helium burning	Supernovae and x-ray bursts	$\approx 1$	0.6-1.25
Explosive oxygen and silicon burning	Supernovae	> 5	> 1.45

# Motivation Highlight: Black Hole Mass Gap Link to LIGO

#### • Farmer *et al.* (2020), Mehta *et al*. (2022)

THE ASTROPHYSICAL JOURNAL, 924:39 (21pp), 2022 January 1



**Figure 9.** Relative uncertainties in the  ${}^{12}C(\alpha, \gamma){}^{16}O$  reaction rate of this work, expanded from those presented in deBoer et al. (2017). The uncertainties are normalized to the central value for clearer presentation. The regions of fading blue color represent  $0.5\sigma$  steps in the Gaussian uncertainty distribution.

Following Gialanella et al. (2001)



# Motivation Highlight: White Dwarf Seismology

• Chidester *et al*. (2022,2023)



**Figure 9.** Relative uncertainties in the  ${}^{12}C(\alpha, \gamma){}^{16}O$  reaction rate of this work, expanded from those presented in deBoer et al. (2017). The uncertainties are normalized to the central value for clearer presentation. The regions of fading blue color represent  $0.5\sigma$  steps in the Gaussian uncertainty distribution.

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<sup>12</sup>C( $\alpha,\gamma$ )<sup>16</sup>O S-factor

 $S(E) = \sigma(E) E e^{2\pi\eta}$ 



### "World Data Set" for the ground state transition

In hindsight, it was not the best idea to only report "E1 and E2" cross sections, since this makes it harder to understand discrepancies between the different measurements





# Major experimental challenges

- Very low cross section
- Background reactions
  - <sup>13</sup>C(α,n)<sup>16</sup>O
  - High Q-value (n,γ) on nearby material
- Inverse kinematics has issues
- Solutions
  - Very clean target / beam lines
  - Time-of-flight method to separate prompt γ-ray signals from (n,γ) delayed ones
  - Recoil separators
  - More exotic solutions



### **Transfer reaction studies**

• SubCoulomb measurements of Asymptotic Normalization Coefficients ( $C_{\lambda c}$ )



$$C_{\lambda c} = \frac{(2\mu_{\alpha}a_c)^{1/2}}{\hbar W_c(a_c)} \frac{\tilde{\gamma}_{\lambda c}}{\left[1 + \sum_{c'} \tilde{\gamma}^2_{\lambda c'} (dS_{c'}/dE)(\tilde{E}_{\lambda})\right]^{1/2}},$$



Reference	$ANC_{\alpha}$ (fm <sup>-1/2</sup> )		
	6.92 MeV, 2 <sup>+</sup>	7.12 MeV, 1 <sup>-</sup>	
Transfer Brune et al. (1999) Belhout et al. (2007) Oulebsir et al. (2012) Avila et al. (2015)	$\begin{array}{c} 1.14(10)\times 10^{5}\\ 1.40(50)\times 10^{5\mathrm{c}}\\ 1.44(28)\times 10^{5}\\ 1.22(7)\times 10^{5} \end{array}$	$\begin{array}{c} 2.08(20)\times 10^{14}\\ 1.87(54)\times 10^{14}\\ 2.00(35)\times 10^{14}\\ 2.09(14)\times 10^{14} \end{array}$	

#### Other compound nucleus reactions



### Phenomenological R-matrix



### Data inconsistencies

- "World data sets"
- Normalized here, but still show a lot of scatter
- Larger discrepancies in ground state E2 data
- However, more modern measurements have, usually, produced more consistent data





# Measurement Highlights: ERNA

- Schürmann *et al*. (2005, 2011)
- High precision total cross section measurement over a wide energy range
- Cascade transition measurements over a difficult range to access using standard methods
- Is extremely helpful in constraining background contributions in the R-matrix fit
- New measurements under development at Caserta.







# Measurement Highlights: TAMU ANCs

- Avila et al. (2015)
- Built on the ground breaking sub-Coulomb transfer measurements by Brune et al. (1999)
- <sup>12</sup>C(<sup>6</sup>Li,d)<sup>16</sup>O
- Improved measurements of the 2<sup>+</sup> and 1<sup>-</sup> ANCs, first measurements of the o<sup>+</sup> and 3<sup>-</sup> ANCs
  - Around 10% uncertainties achieved



# Measurement Highlight: High Precision ${}^{12}C(\alpha,\gamma){}^{16}O$ low energy measurements

- Plag et al. (2012)
- Used time-of-flight to separate out background from (n,γ) reactions produced by neutrons from the <sup>13</sup>C(α,n)<sup>16</sup>O reaction
- BaF<sub>2</sub> array for detailed angular distribution measurements
- All angular distribution data is given, not just E1 and E2 data



## Measurement Highlight : ${}^{16}O(\gamma, \alpha)$ ${}^{12}C$



Long campaign (hundreds of hours) led by Moshe Gai

## Measurement Highlight : Alternative transfer reactions

- Shen et al. (2019, 2020)
- <sup>12</sup>C(<sup>11</sup>B,<sup>7</sup>Li)<sup>16</sup>O
- Probing the model and reaction sensitivity of ANC determination
- 2<sup>+</sup> ANC is remeasured in (2019)
- GS ANC is measured in (2020)



TABLE I.Present ANC of the <sup>16</sup>O GS and other availableresults in the literature.Shen et al. (2020)

Reference	ANC $(fm^{-1/2})$	Method
Adhikari (2009) [14] Morais (2011) [16]	$13.9 \pm 2.4$ 3390 (WS1) 1230 (WS2) 750 (FP)	<sup>16</sup> O + Pb breakup <sup>12</sup> C( <sup>16</sup> O, <sup>12</sup> C) <sup>16</sup> O
Sayre (2012) [11] Adhikari (2017) [15] Present	730 (PP) 709 $637 \pm 86$ $337 \pm 45$	<i>R</i> matrix <sup>12</sup> C( <sup>7</sup> Li, <i>t</i> ) <sup>16</sup> O <sup>12</sup> C( <sup>11</sup> B, <sup>7</sup> Li) <sup>16</sup> O

### On the horizon: ${}^{16}O(e,e'\alpha){}^{12}C$

- MIT group
  - See Friščić et al. (2019)



FIG. 3. First-order Feynman diagram for the electrodisintegration of <sup>16</sup>O involving one virtual photon  $\gamma^*$  exchange to be compared with Fig. 2. Again, the kinematic variables here will be discussed in more detail in Sec. III.



# On the horizon: ${}^{16}O(\gamma, \alpha)$ ${}^{12}C$

- Jefferson Laboratory
- Ernst Rehm and Claudio Ugalde
- Bubble Chamber + Bremsstrahlung beam
- Previous tests at HIγS
- Not sure on current status?



# On the horizon: Coulomb dissociation of <sup>16</sup>O on lead

#### • FAIR at GSI

- Aims to get 10% uncertainty at 1 MeV
- Will cover a wide energy range
- Some measurements made, but still under analysis



#### measure

- charge
- track points
- time
- photon from <sup>12</sup>C\*

#### to determine

excitation energy

Figure courtesy of Rene Reifarth

# On the Horizon: Felsenkeller planned Gas Jet Target based study of ${}^{12}C(\alpha,\gamma){}^{16}O$

- <sup>12</sup>C(α,γ)<sup>16</sup>O potential for Felsenkeller with <sup>12</sup>C+ beam, extended windowless <sup>4</sup>He gas target, γ-calorimeter, and 4π detector
- At Felsenkeller, planned to cover energy range from 0.6 2.0 MeV
- Gas target is complete, detector procurement is underway





80

60

40

20

0.3

S(E) [keV barn]



#### Recent questions about $\alpha$ particle ANCs

- Revised <sup>6</sup>Li ANC effects the ANCs of any (<sup>6</sup>Li,d) type alpha transfer measurement
- In the case of <sup>12</sup>C(α,γ)<sup>16</sup>O, it would imply smaller ANCs and therefore a smaller cross section by about 30%



Hebborn *et al*. (arXiv, July 2023)

# Multiple GS and 2<sup>+</sup> ANC solutions

Shen *et al*. (2020)

- <sup>12</sup>C(<sup>11</sup>B,<sup>7</sup>Li)<sup>16</sup>O a transfer was used to measure the ground state ANC
- E2 external capture is small, but at very low energies its interference with the 2<sup>+</sup> subthreshold state is substantial
- However, effect is lower in energy than we can measure directly
- Implies a larger 2<sup>+</sup> ANC is needed



# Revised method of extracting ANCs from scattering phase shifts

deBoer et al. (2017)

- Blokhintsev et al. (2023), Mukhamedzhanov et al. (arXiv, 2024)
- Extraction of ANCs using Tisshauser *et al*. (2009) phase shifts
- Question about uncertainties
- Larger 2+ ANC (1.14 → 1.42)10<sup>5</sup>
   fm<sup>-1/2</sup>

Transition to the ground state via	$S(300 \mathrm{keV})$	$S(300{\rm keV})$
resonance $+$ direct capture	Present	Ref. [2]
E1	98	85
E2	70	45
E1 + E2	168	130
cascade	Present	Ref. [2]
$0^+_2 + 3^- + 2^+ + 1^-$	6	7
total	Present	Ref. [2]
E1 + E2 +cascade	174	137

#### Improvement in reaction rate tabulation!

- Reaction rates are tabulated over "standard" temperature grids.
- Astrophysics codes then interpolate the values given in tables
- Have to watch out for too few temperature steps!
- Frank Timmes and Ebraheem Farag



MESA interpolation is off by more than  $1\sigma$  for deBoer et al. (2017) tabulated rates!!!

# The R-matrix fit ends up being a big multidimensional fit problem

#### • Fit parameters

- Level energies
- Level widths (or reduced width amplitudes)
- Asymptotic Normalization Coefficients (or reduced width amplitudes)
- Data normalization factors
- For the case described in the deBoer *et al.* (2017), there were 64 level parameters and over 100 data normalization fit parameters
- >10,000 experimental data points
- Computationally challenging



# Improving uncertainty estimation for R-matrix fits

- A more general problem
- Bayesian methods provide a way to improve and gain more detailed information
- See de Souza et al. (2020) for an application to <sup>3</sup>H(d,n)<sup>4</sup>He
- Computationally intensive, but probably doable
- Daniel Odell at Ohio University has developed the Bayesian R-matrix Inference Code Kit (BRICK) for use with the AZURE2 R-matrix code



Figures courtesy of Daniel Odell

# Summary

- Improvements in the uncertainty of the low energy S-factor are hampered by inconsistent data
  - Newer data are much more consistent! We're on the right track, but measurements are very challenging
  - Ground state E2 data is in the worst shape, because you usually need to measure a more complete angular distribution to obtain it
  - ERNA recoil separator measurements provided a LOT more constraint for the R-matrix fit because they covered a wide energy range. More of these type measurements would be very useful!
- Make new measurements using new techniques (reduce systematic uncertainties)
  - ${}^{16}O(\gamma,\alpha){}^{12}C$  (HI $\gamma$ S, Jefferson Lab, ELI-NP)
  - <sup>16</sup>O(e,e'α)<sup>12</sup>C
  - Additional types of transfer measurements
  - Improved traditional measurements, but in low background environments with very high beam intensities (reduce statistical / outlier uncertainties)
    - JUNA, LUNA, Dresden
- Include more detailed experimental uncertainties
  - Energy uncertainty, experimental resolution
  - Bayesian uncertainty estimation
  - Improved computational resources

#### Questions?



My 2017 estimate of S(300 keV):

140 ± 21 (MC) +18/-11 (model)

But see Shen *et al*. (2020)

Assumes ANC uncertainties are accurate

## Strategy in 2017 (and earlier)

Used a broader distribution

function for data point

than a Gaussian)

uncertainties (larger tails

 Use χ<sup>2</sup> minimization to find a best description of the data within the R-matrix framework (used MINUIT<sub>2</sub>)

 $\chi^2 = \sum_{i} \left( \frac{\sum_{j} R_{ij}^2}{N_i} \right) + \frac{(n_i - 1)^2}{n_i^2 \sigma_{\text{syst},i}^2}$ 

$$\chi^2 = \sum_i \left(\frac{\sum_j R_{ij}^2}{N_i - \nu}\right) + \frac{(n_i - 1)^2}{n_i^2 \sigma_{\text{syst},i}^2}$$

$$R_{ij} = \frac{f(x_{i,j}) - n_i y_{i,j}}{n_i \sigma_{i,j}}$$

Additional data normalization fit parameter, with common systematic uncertainty. Also fitting to reduced  $\chi^2$ .

0 /

$$\frac{(n_i - 1)^2}{n_i^2 \sigma_{\text{syst},i}^2}$$

$$+ \sum \frac{(P_{\text{fit},k} - P_{\text{exp},k})^2}{\sigma^2}$$
Add additional terms for subthreshold ANC to include uncertainty constraints from transfer reactions.



$$L = \sum_{j} \log \left[ \frac{1 - e^{-R_{ij}^2/2}}{R_{ij}^2} \right]$$

Sivia and Skilling (2006)

Kind of a Frequentist / Bayesian hybrid method

## R-matrix framework

$$R_{cc'} = \sum_{\lambda} \frac{\gamma_{\lambda c} \gamma_{\lambda c'}}{E_{\lambda} - E},$$

$$\mathbf{U} = \rho^{\frac{1}{2}} \mathbf{O}^{-1} (1 - \mathbf{R} \mathbf{L}_0)^{-1} (1 - \mathbf{R} \mathbf{L}_0^*) \mathbf{I} \rho^{-\frac{1}{2}}$$

$$T_{cc'} = e^{2i\omega_c}\delta_{cc'} - U_{cc'},$$

$$\sigma_{\alpha\alpha'} = \frac{\pi}{k_{\alpha}^2} \sum_{Jll'ss'} g_J \left| T_{cc'}^J \right|^2,$$



Bardayan et al. (2015)

# Simple Monte Carlo

- Created thousands of "synthetic" data sets by assuming that the error bars on the data represented an underlying Gaussian (probably should have used lognormal) Probability Density Function.
- Refit
- Histogrammed S-factor calculated at many energies to get uncertainty.
- Calculated many different variations on assumptions about the *R*-matrix fit and included those as well.
- Even more computationally expensive



Monte Carlo of experimental data Model uncertainties

### Model assumptions

- External capture contribution for GS?
  - Often neglected completely
  - May play a significant role in the E2 cross section (given recent improvements in the uncertainty of this cross section)
- Inconsistent measurements of the ground state  $\alpha$ -particle ANC in <sup>16</sup>O
- Additional measurements needed!

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Reference	ANC $(fm^{-1/2})$	Method
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Morais (2011) [16]	3390 (WS1)	$^{12}C(^{16}O, ^{12}C)^{16}O$
	1230 (WS2)	
	750 (FP)	
Sayre (2012) [11]	709	R matrix
Adhikari (2017) [15]	$637 \pm 86$	${}^{12}C(^{7}Li, t){}^{16}O$
Present	$337\pm45$	$^{12}C(^{11}B, ^{7}Li)^{16}O$



# **Background contributions**



Best solution: Fit experimental data up in energy until the cross section gets smaller or there is a natural gap in the level structure.

# Fit to higher energy

- Challenges
  - Multiple particle decay channels
  - Multiple entrance channel data
- Additional advantages
  - Combine <sup>12</sup>C(α,γ), <sup>15</sup>N(p,γ), and
     <sup>15</sup>N(p,α) R-matrix analyses into single consistent analysis
  - Check on systematic uncertainties like effective energy determination and partial width consistency



# *R*-matrix LEGO blocks (reaction components)

Resonance
Subthreshold state
HS External Capture



### "E1 and E2" data

- Can in principle measure the "E1 cross section" by measuring the differential cross section at 90 degrees.
  - Problems when E2 >> E1
- To get E<sub>2</sub> one needs to measure the angular distribution.
- Main issue: often only the E1 and E2 derived cross sections have been reported in the literature, and the original angular distributions are omitted.
- Problems with extracting E2 data
  - Number of parameters to use for fit.
  - Systematic problems with data
  - Can often be corrected/understood if original angular distribution data are given
  - See Brune and Sayre (2013)



# Measurement Highlights: High Precision ${}^{12}C(\alpha,\gamma){}^{16}O$ low energy measurements

- Makii et al. (2009)
- Used time-of-flight to separate out background from (n,γ) reactions produced by neutrons from the <sup>13</sup>C(α,n)<sup>16</sup>O reaction
- Only two energies, but highly accurate and precise
- Just wish they could have done more measurements



