



# A Neutrino Oscillation Global Fit Using GAMBIT

**Wilf Shorrock**

**On behalf of the GAMBIT Collaboration**

3<sup>rd</sup> September 2024

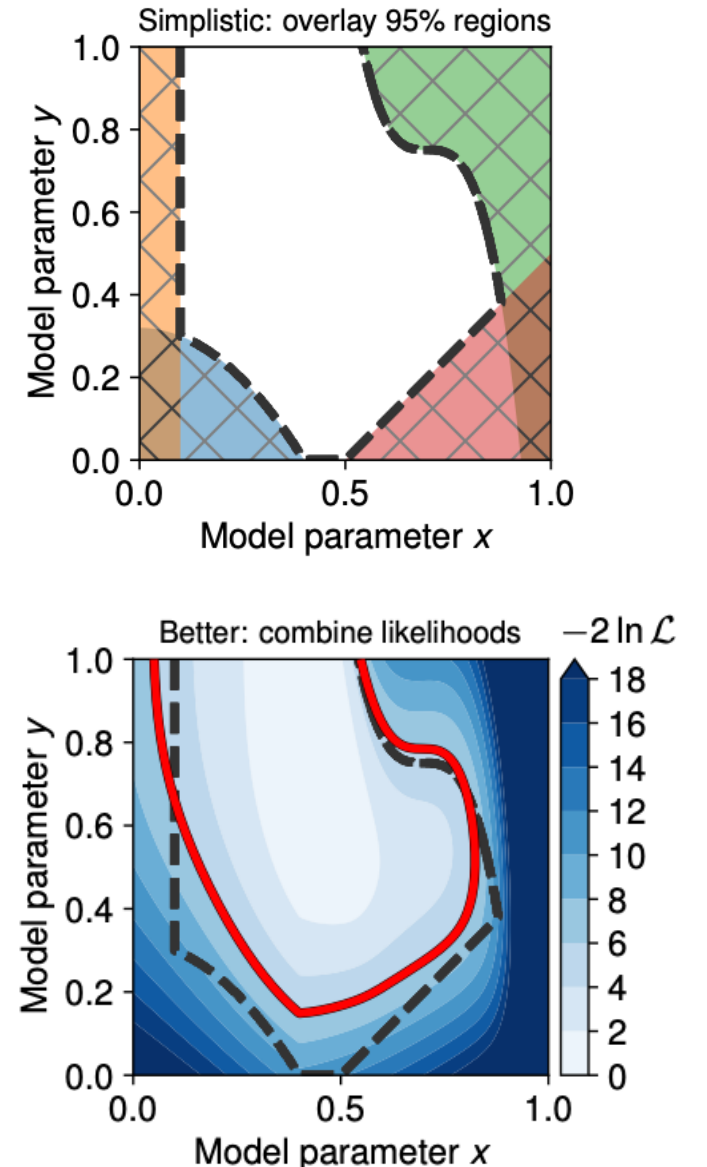
*NOW 2024*



# Global Fits

- Utilise data from multiple experiments
- Combine the unique strengths of each experiment
- Return the most accurate confidence/credible regions for model variables
- General approach: combine likelihoods from each experiment
- Past and present global fits (non-exhaustive):
  - NuFIT<sup>1,2</sup>
  - Valencia<sup>3</sup>
  - Bari<sup>4</sup>

- 1) JHEP 09 (2020) 178
- 2) NuFIT 5.3 (2024) [www.nu-fit.org](http://www.nu-fit.org)
- 3) JHEP 02 (2021) 071
- 4) <https://doi.org/10.1016/j.pnpnp.2005.08.002>



# GAMBIT: The Global And Modular BSM Inference Tool

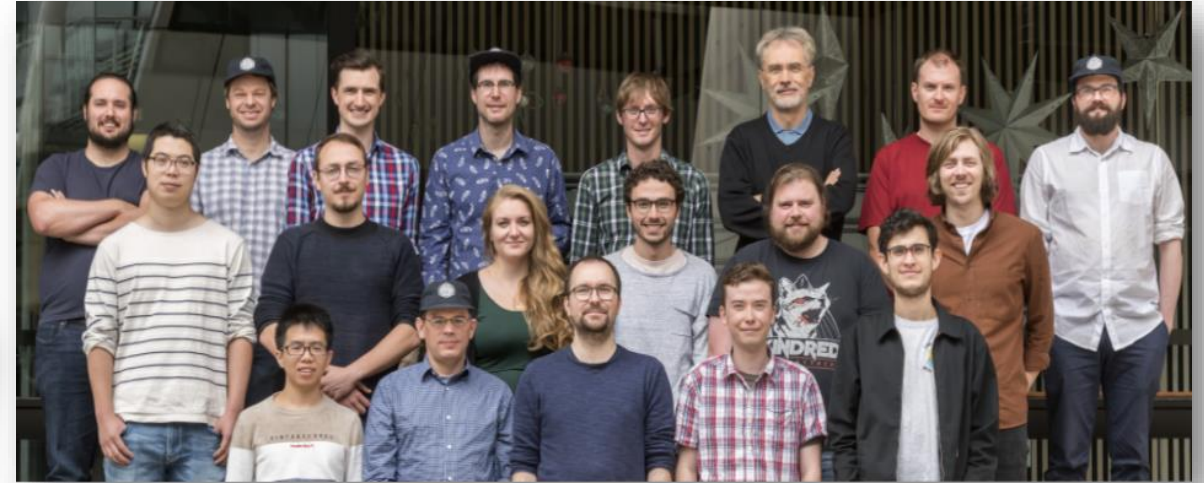
[gambitbsm.org](http://gambitbsm.org)

[github.com/GambitBSM](https://github.com/GambitBSM)

EPJC 77 (2017) 784

arXiv:1705.07908

- Extensive model database, beyond SUSY
- Fast definition of new datasets, theories
- Extensive observable/data libraries
- Plug&play scanning/physics/likelihood packages
- Various statistical options (frequentist /Bayesian)
- Fast LHC likelihood calculator
- Massively parallel
- Fully open-source



**Members of:** ATLAS, Belle-II, CLiC, CMS, CTA, Fermi-LAT, DARWIN, IceCube, LHCb, NOvA, SHiP, T2K, XENON

**Authors of:** BubbleProfiler, Capt'n General, Contur, DarkAges, DarkSUSY, DDCalc, DirectDM, Diver, EasyScanHEP, ExoCLASS, FlexibleSUSY, gamLike, GM2Calc, HEPLike, IsaTools, MARTY, nuLike, PhaseTracer, PolyChord, Rivet, SOFTSUSY, SuperIso, SUSY-AI, xsec, Vevacious, WIMPSim

**Recent collaborators:** V Ananyev, P Athron, N Avis-Kozar, C Balázs, A Beniwal, LL Braseth, T Bringmann, A Buckley, J Butterworth, JE Camargo-Molina, C Chang, J Cornell, M Danninger, A Fowlie, T Gonzalo, W Handley, S Hoof, A Jueid, F Kahlhoefer, A Kvellestad, M Lecroq, C Lin, M Lucente, FN Mahmoudi, DJE Marsh, G Martinez, H Pacey, MT Prim, T Procter, F Rajec, A Raklev, R Ruiz, A Scaffidi, P Scott, W Shorrock, C Sierra, P Stöcker, W Su, J Van den Abeele, A Vincent, M White, A Woodcock, Y Zhang ++

**70+ participants in many experiments and numerous major theory codes**

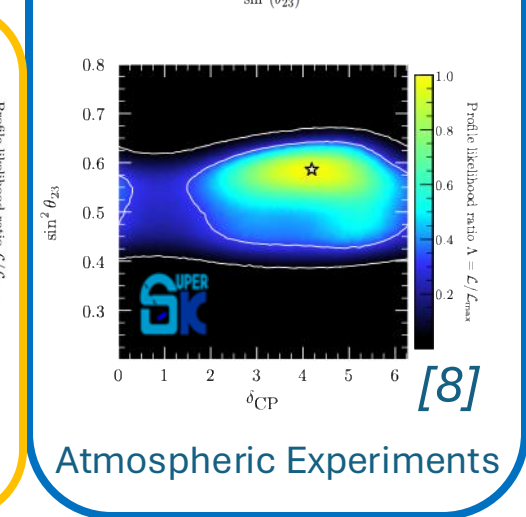
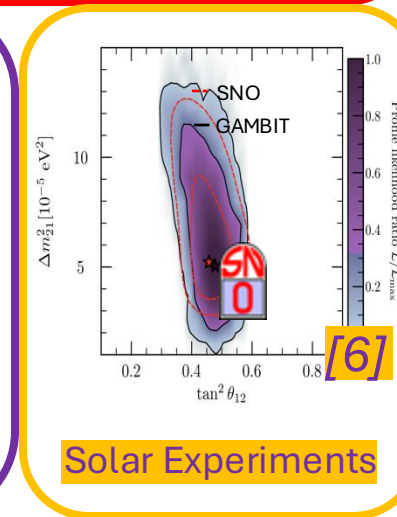
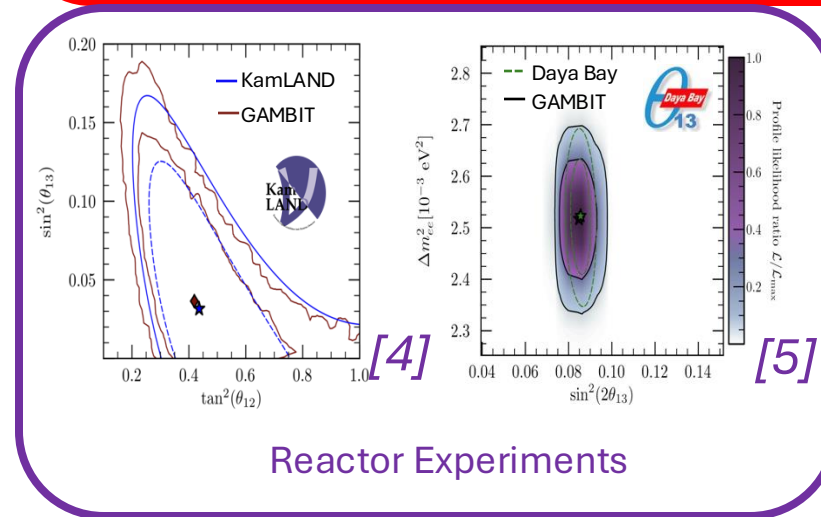
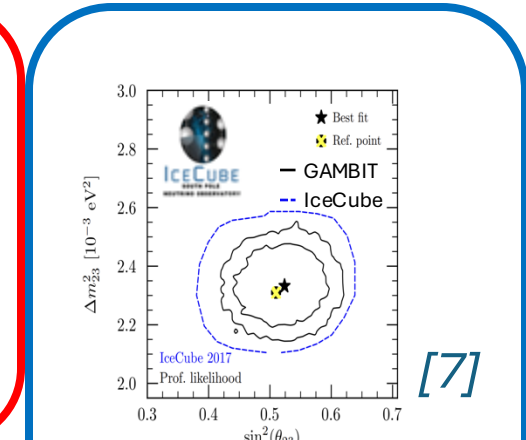
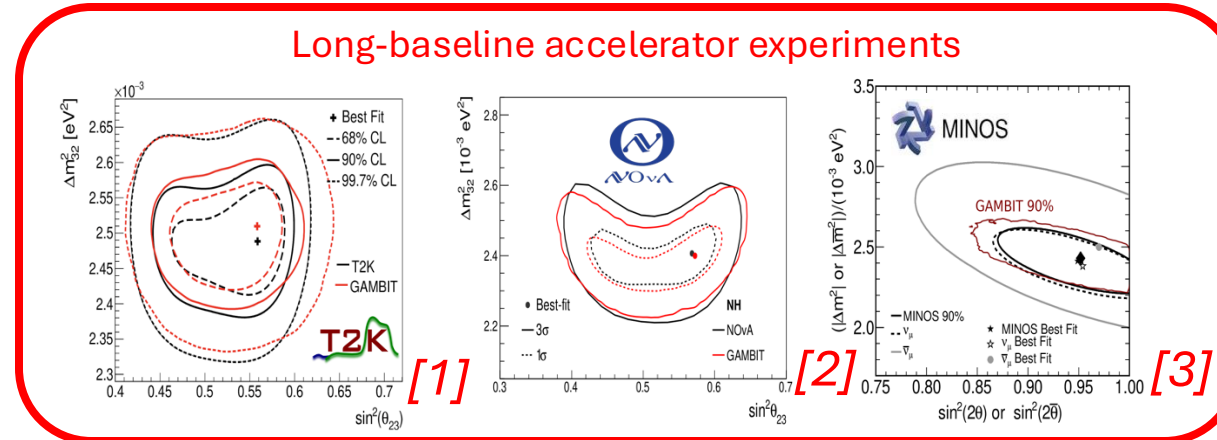
# GAMBIT's Approach to Neutrino Oscillations

- Keep things **open source**
- Provide a tool where anyone can utilise neutrino oscillation data from **all major experiments**
- Use **published information** wherever possible
- Use **advanced sampling techniques** for efficient exploration of parameter space
- Publish global fits of the standard three-neutrino oscillation results with **clear procedures**



# Included Experiments

- Daya Bay
- IceCube
- KamLAND
- MINOS
- NOvA (2020)
- SNO
- T2K (2020)
- More in future

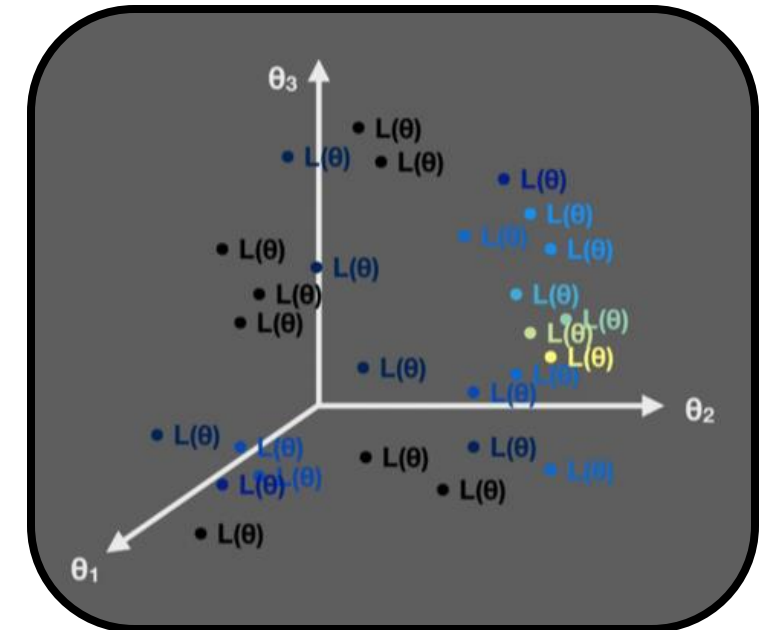


[1] T2K 2020 result. <https://doi.org/10.5281/zenodo.3959558>  
 [2] NOvA 2020 result. <https://doi.org/10.5281/zenodo.4142045>  
 [3] MINOS 2013 paper. <https://doi.org/10.1103/PhysRevLett.110.251801>  
 [4] KamLAND 2011 paper. <https://doi.org/10.1103/PhysRevD.83.052002>

[5] Daya Bay 2018 paper. <https://doi.org/10.1103/PhysRevLett.121.241805>  
 [6] SNO paper. <https://arxiv.org/abs/1109.0763>  
 [7] IceCube 2018 paper <http://link.aps.org/doi/10.1103/PhysRevLett.120.071801>  
 [8] Super-K paper. <https://arxiv.org/abs/1710.09126>

# Method

- Calculate likelihood ( $L$ ) of each experiment by comparing data to simulation
- Scan through parameter space, calculating  $L$  for each experiment
- Converge to best fit
- Below is the calculation for the number of detected events for neutrino flavour  $\alpha$  in energy bin  $i$



$$N_i^\alpha = N_{\text{bkg},i} + \int_{E_i}^{E_{i+1}} dE_{\text{rec}} \int_0^\infty dE_\nu \overset{\text{Reconstruction function}}{\downarrow} R(E_{\text{rec}}|E_\nu) \overset{\text{Neutrino beam flux}}{\downarrow} \frac{d\Phi}{dE_\nu} \overset{\text{Efficiency}}{\downarrow} \sigma_\alpha(E_\nu) \epsilon(E_\nu) P_{\nu_\beta \rightarrow \nu_\alpha}(E_\nu)$$

$\uparrow$   
 Total cross section

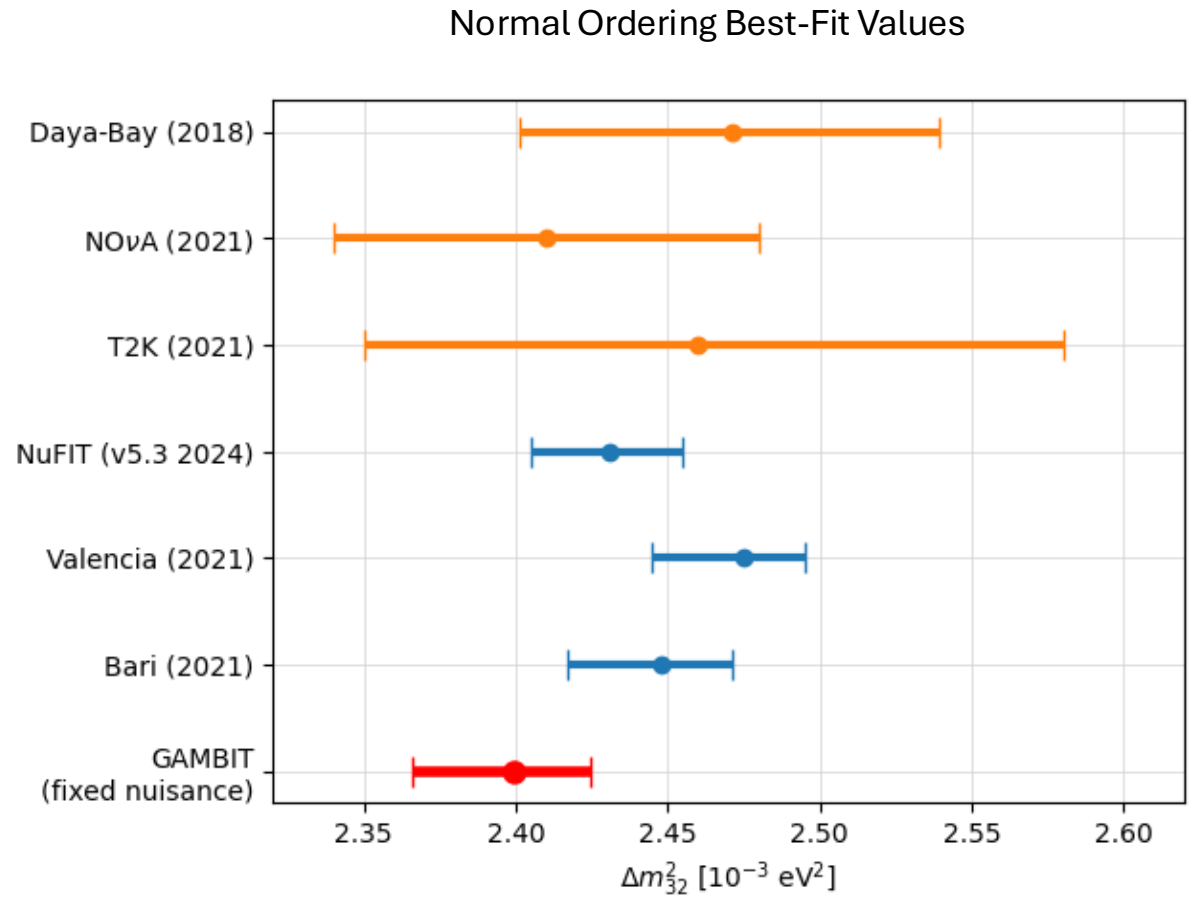
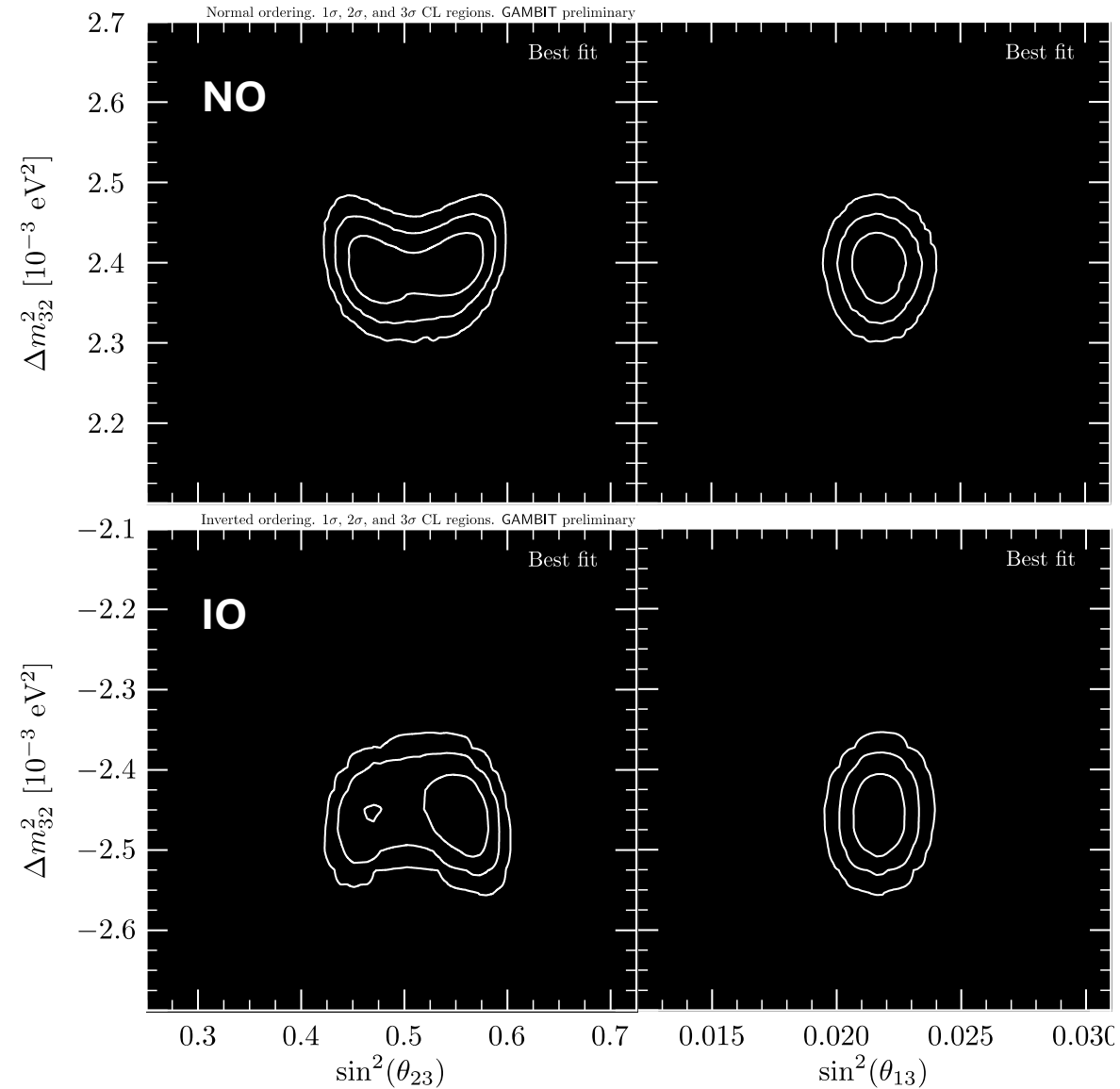
# Results

## **Please note:**

These results are preliminary, with no inclusion of nuisance parameters in the scans.

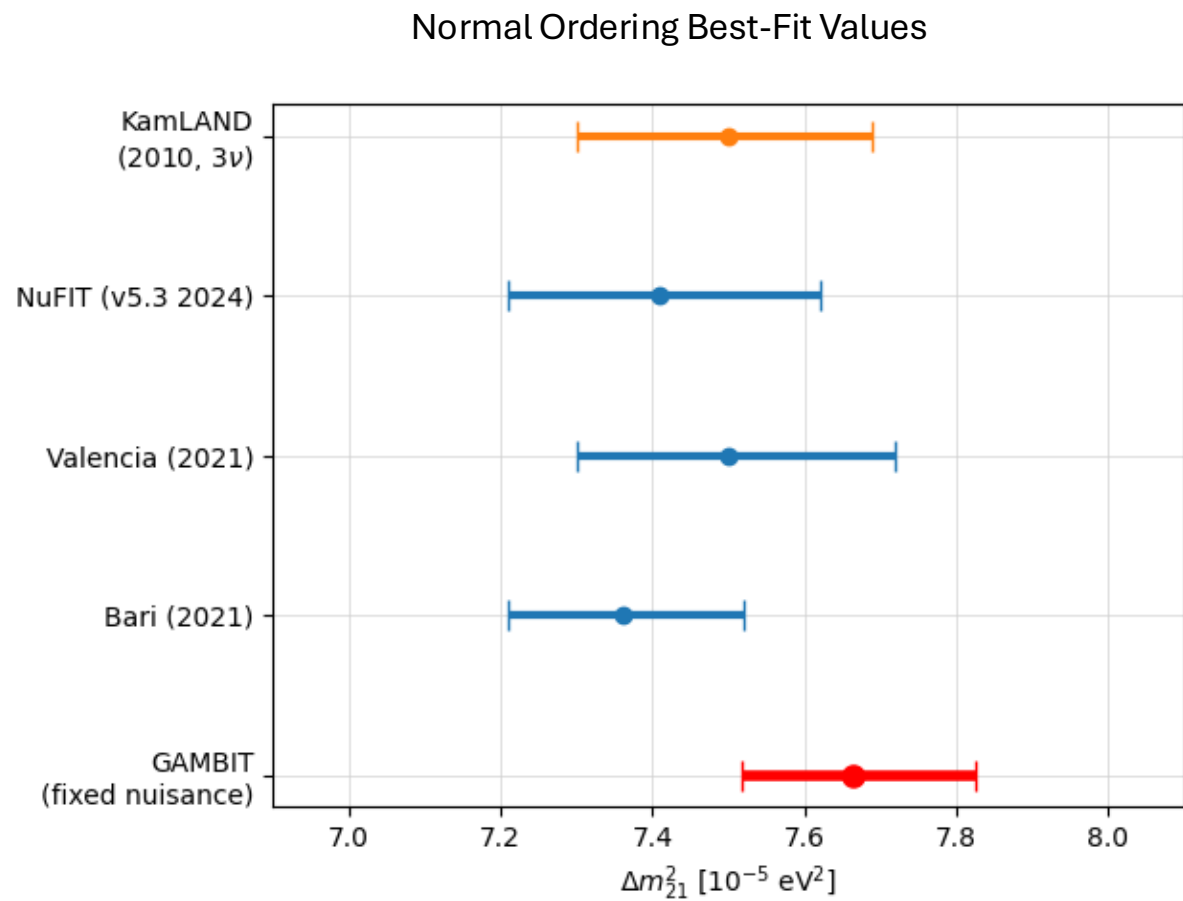
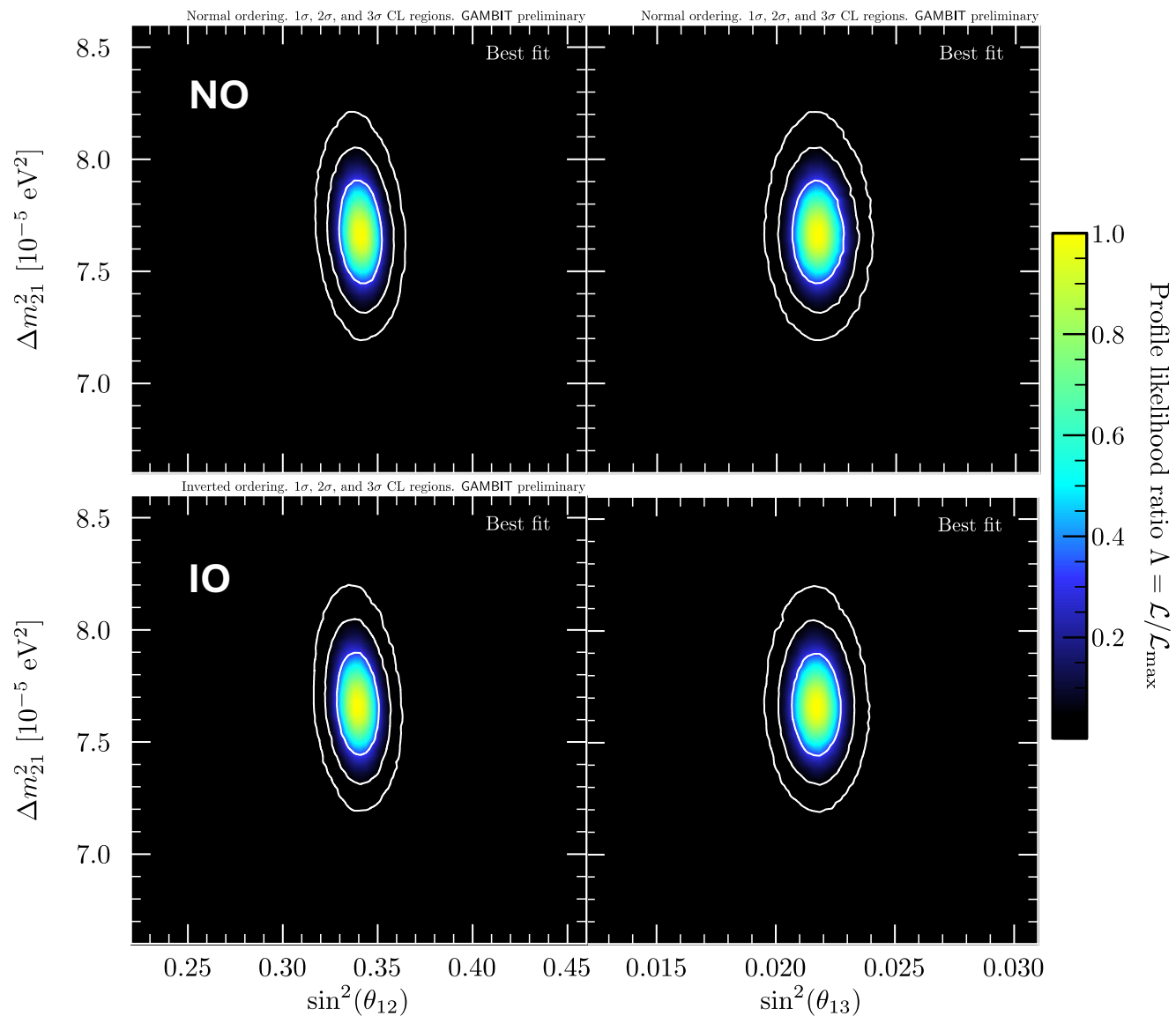
We plan to publish full results soon.

# Results ( $\Delta m_{32}^2$ )

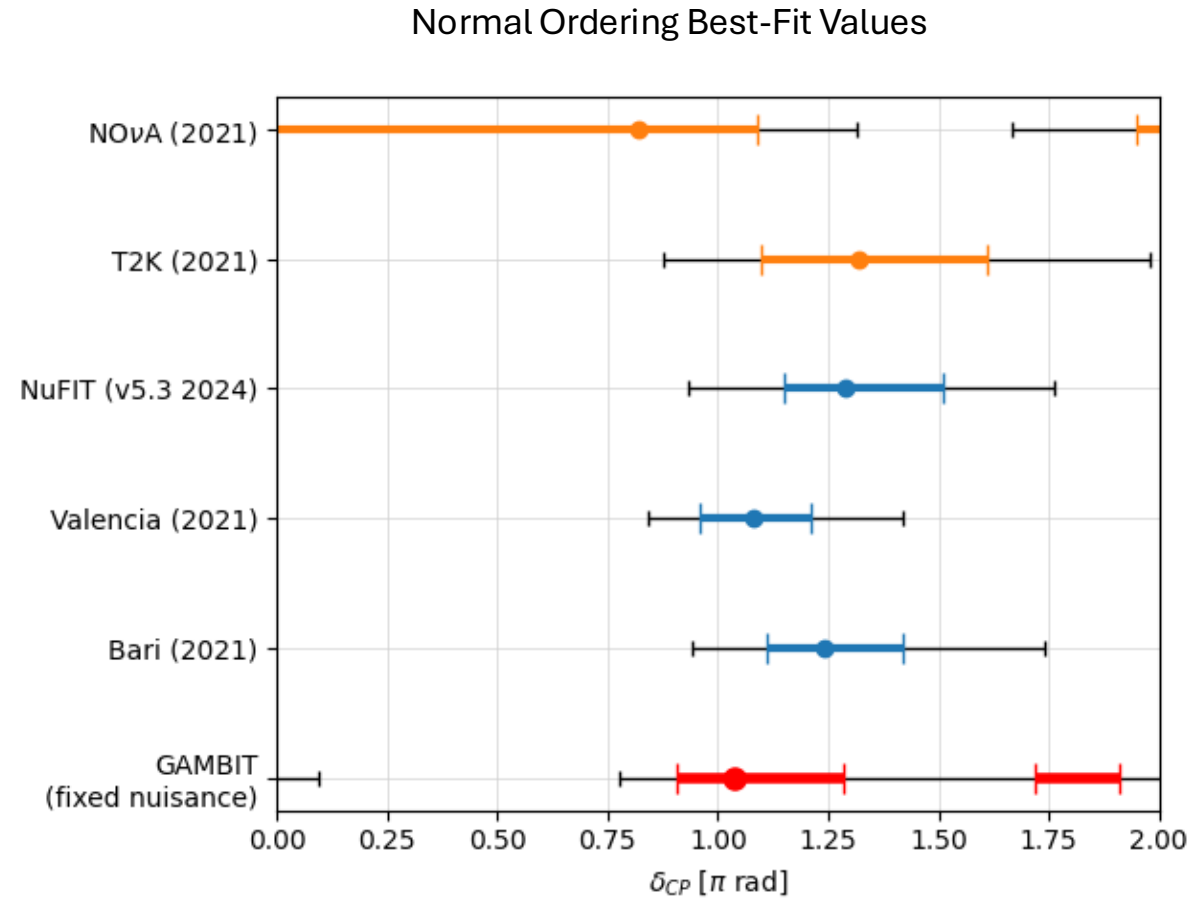
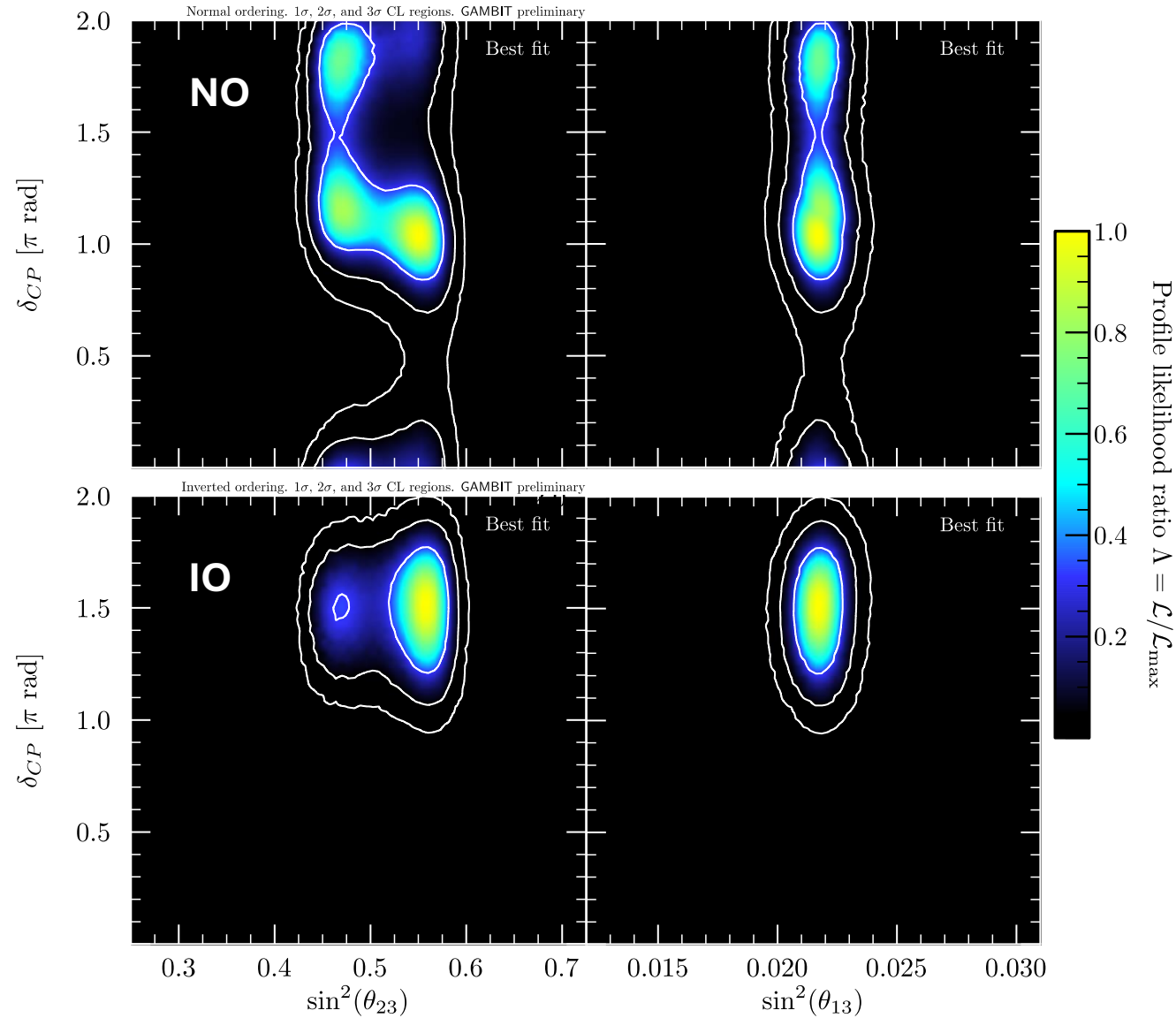




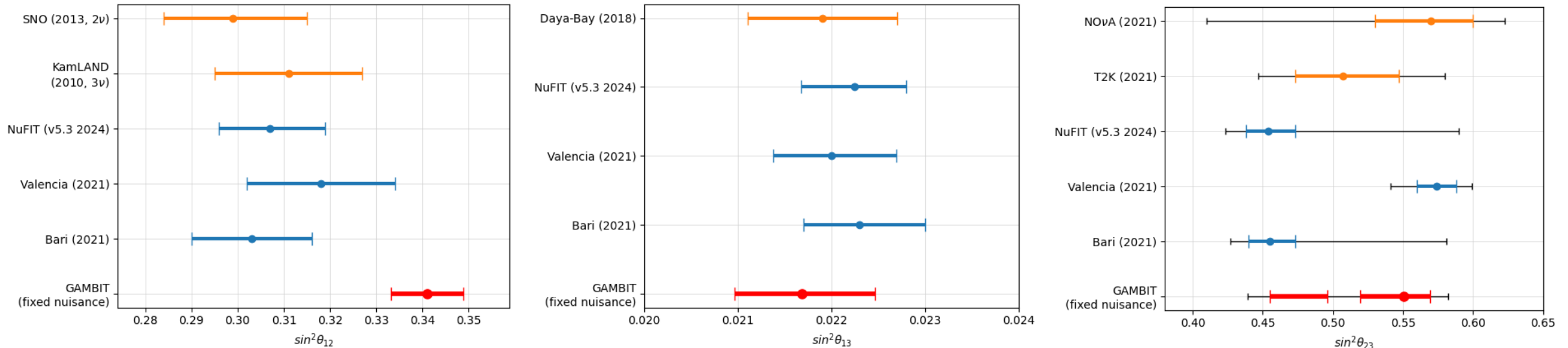
# Results ( $\Delta m_{21}^2$ )



# Results ( $\delta_{CP}$ )



# Results (Mixing Angles)



- All best-fit values of parameters agree with previous global fits, except  $\theta_{12}$
- We used a new (2016) solar model<sup>1</sup> in our fit, with higher flux. A higher value of  $\theta_{12}$  is expected
- In our upcoming paper we will show results with both solar models
- This highlights one of GAMBIT's uses, testing new models independent of large collaborations

1) Solar Model B16(AGSS09): arXiv: 1611.09867

# Summary

- Preliminary (fixed nuisance) results for GAMBIT neutrino oscillation global fit
- Initial results are consistent with previous global fits where expected
- Full results coming...
- Will provide an open-source tool for testing new models with all available data

	Normal ordering ( $\pm 1\sigma$ )
$\theta_{12}$ [ $^\circ$ ]	$35.7 \pm 0.5$
$\theta_{13}$ [ $^\circ$ ]	$8.34 \pm 0.25$
$\theta_{23}$ [ $^\circ$ ]	$48.3^{+1.0}_{-1.4}$
$\sin^2 \theta_{12}$	$0.340 \pm 0.009$
$\sin^2 \theta_{13}$	$0.0211^{+0.0013}_{-0.0012}$
$\sin^2 \theta_{23}$	$0.557^{+0.018}_{-0.024}$
$\Delta m_{21}^2$ [ $10^{-5}$ eV $^2$ ]	$7.66^{+0.15}_{-0.14}$
$\Delta m_{32}^2$ [ $10^{-3}$ eV $^2$ ]	$2.41 \pm 0.03$
$\delta_{\text{CP}}$ [ $^\circ$ / $\pi$ rad]	$188^{+27}_{-23} / 1.05^{+0.15}_{-0.13}$

NO results from GAMBIT's fixed nuisance scan

# Backups

# Diver Scanner

- Open source **differential evolution** parameter sampler
- Population-based self-adaptive  $\lambda$ jDE algorithm (rand-to-best/1/bin)
- Mutation -> crossover -> selection