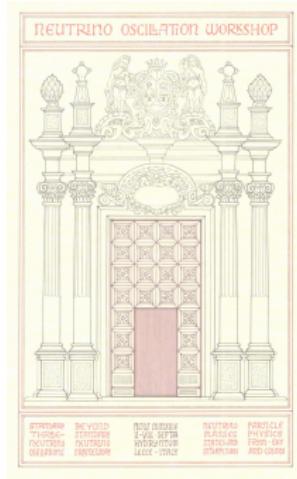


Gravitational Waves: Towards the Einstein Telescope



NOW 2024
Neutrino Oscillation Workshop

Otranto, 2024 Sep 05



Archisman Ghosh

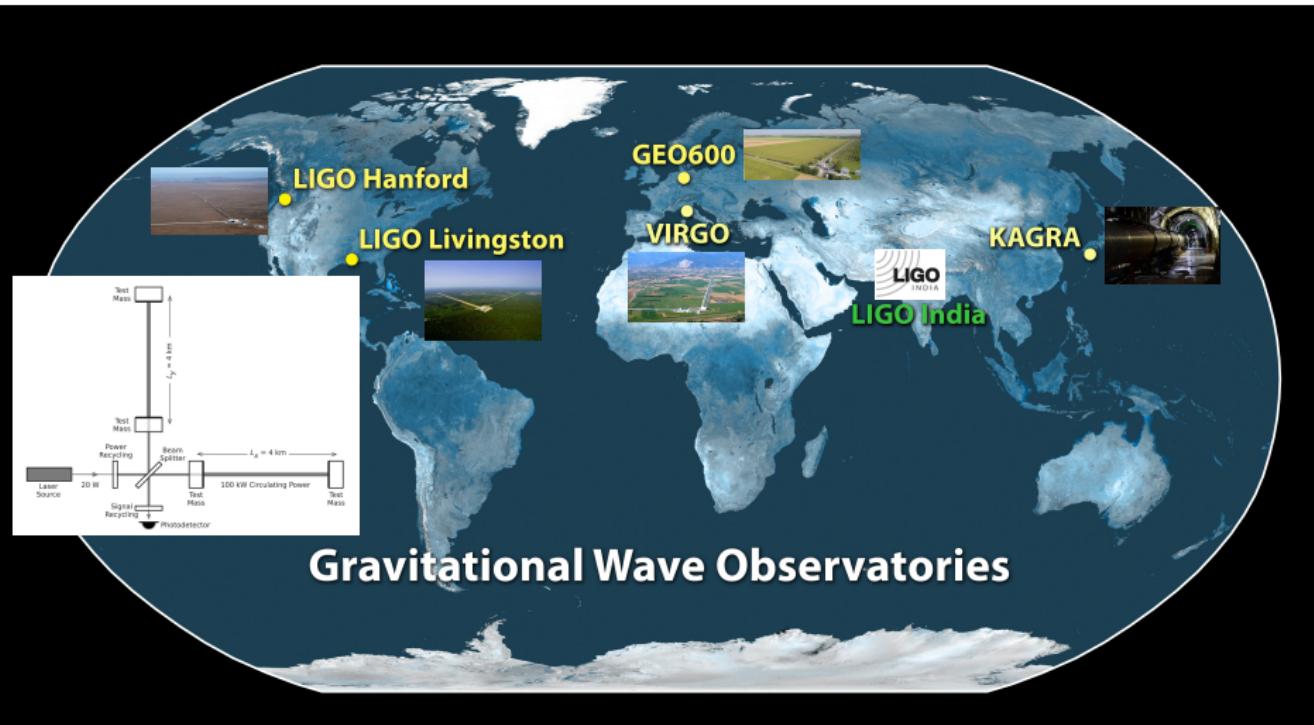
Ghent University

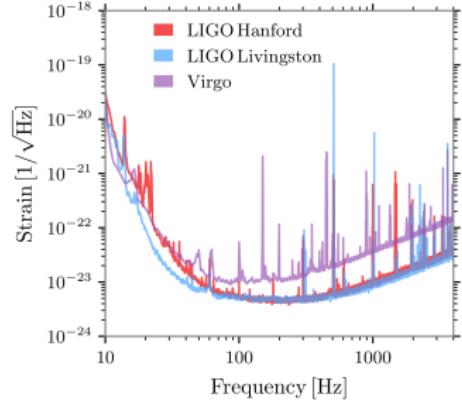
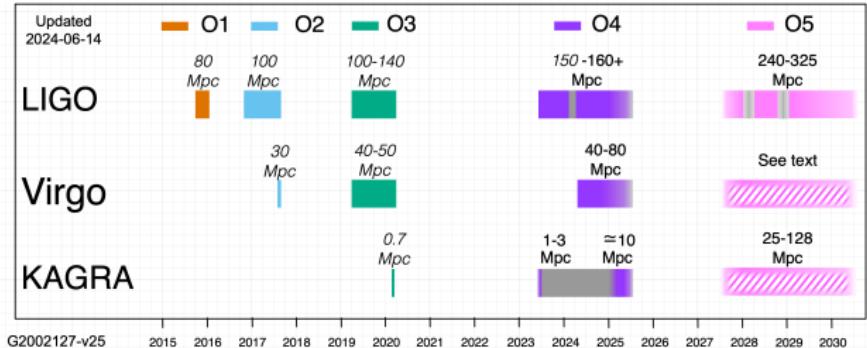
for the Einstein Telescope Collaboration



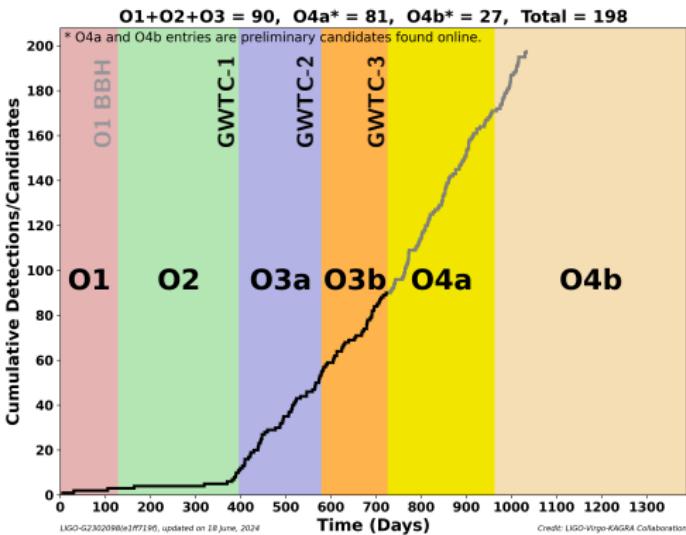
LIGO-Virgo-KAGRA basics

A global network of interferometric GW detectors

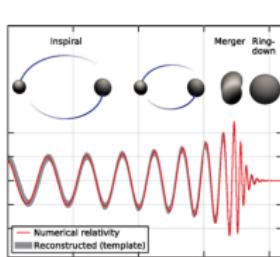




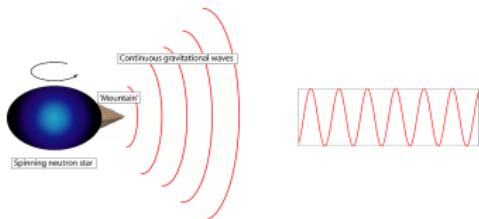
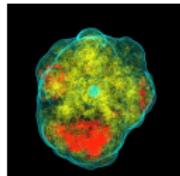
LVK: Abbott+ arXiv:2111.03606



Gravitational-wave sources



	Modelled	Unmodelled
Transient	Compact binary coalescences NS-NS, NS-BH, BBH	Bursts Supernova explosions
Persistent	Continuous waves Spinning deformed NS	Stochastic background Astrophysical + Cosmological



Current observational results

GW150914: the discovery of GWs



first direct detection of gravitational waves

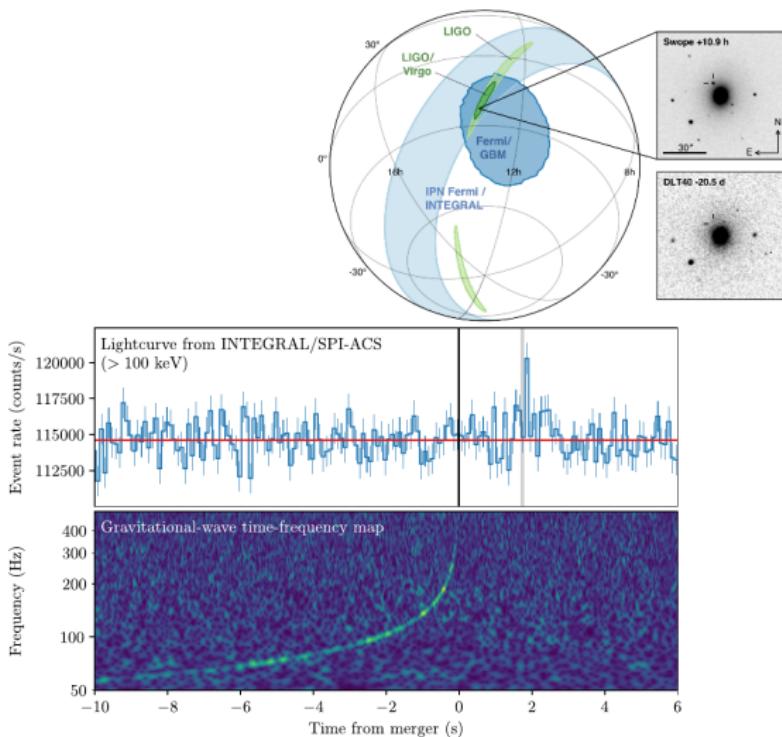
first direct observation of a black hole

discovery of “heavy” black holes $M > 30M_{\odot}$

first observation of a black hole binary

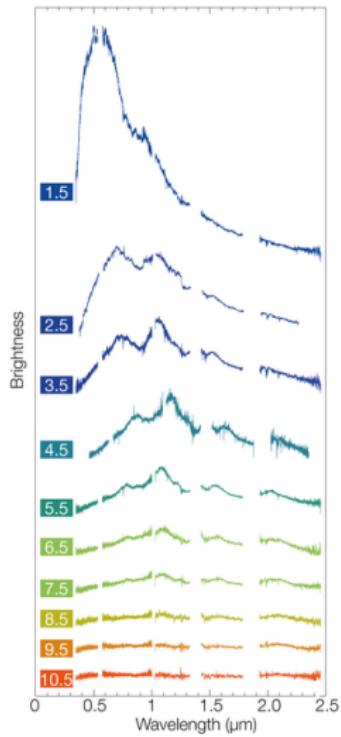
first observation of black hole “formation”

GW170817



LVC: Abbott+ *Astrophys. J.* **848** #2, L13 (2017)
6 of 25

Kilonova Pian+ 2017



birth of **multimessenger science** with GW and EM

confirming BNS coalescence as **progenitors of short GRBs**

demonstrating „ as prolific sites of **heavy element formation**

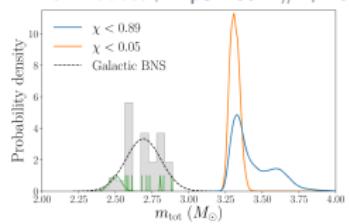
constraining **speed of gravity** w.r.t. speed of light

access to elusive **neutron star equation-of-state**

providing a novel measurement of the **Hubble constant**

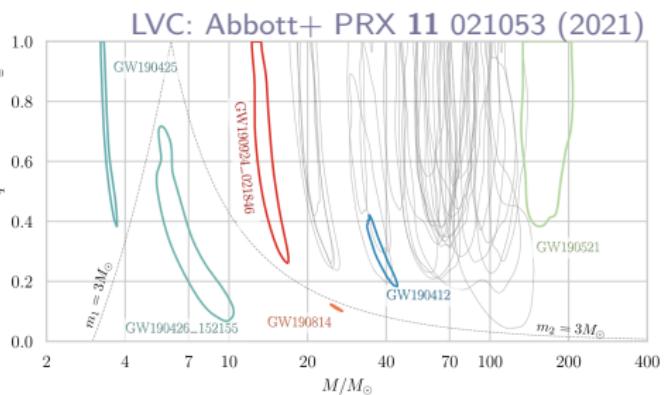
O3a highlights

LVC: Abbott+ ApJL 892 #1, L3 (2020)

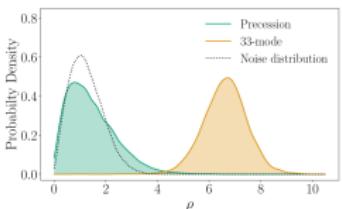
**GW190425**

potential BNS

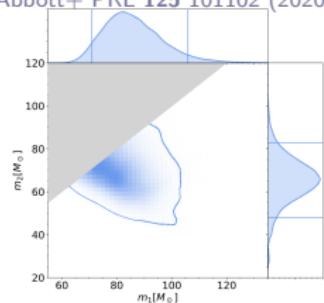
heavy, no counterpart

**GW190814**

lighter object in “mass gap” between NS and BH
mass ratio > 9 | “higher harmonics” in GW signal



LVC: Abbott+ ApJL 896 #2, L44 (2020)

**GW190521**

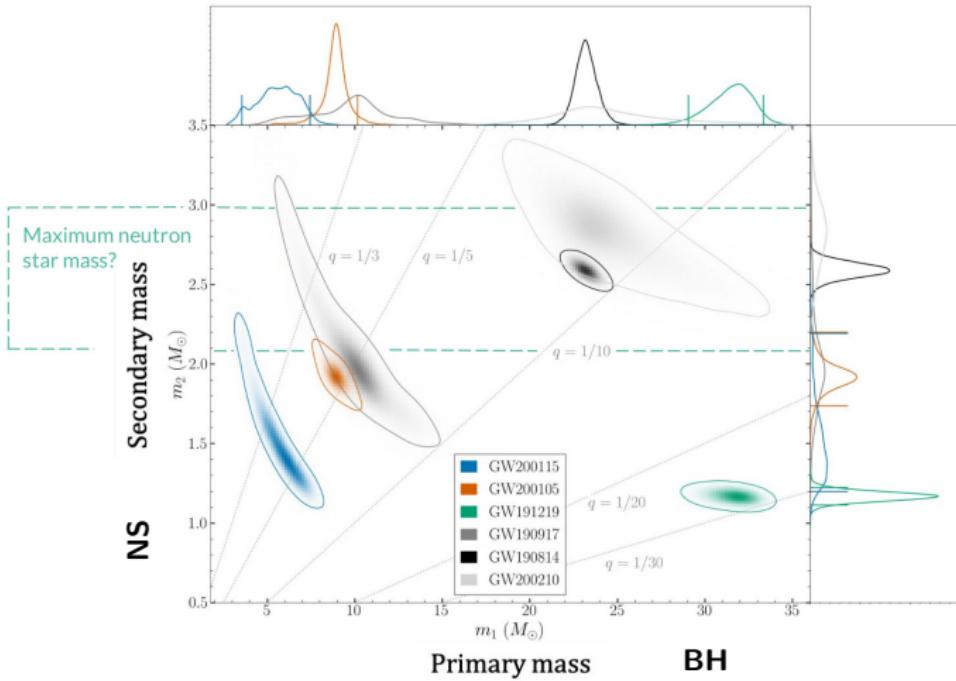
$$m_1 = 85^{+21}_{-14} M_{\odot}$$

$$m_2 = 66^{+17}_{-8} M_{\odot}$$

$$M = 142^{+28}_{-16} M_{\odot} \text{ (IMBH)}$$

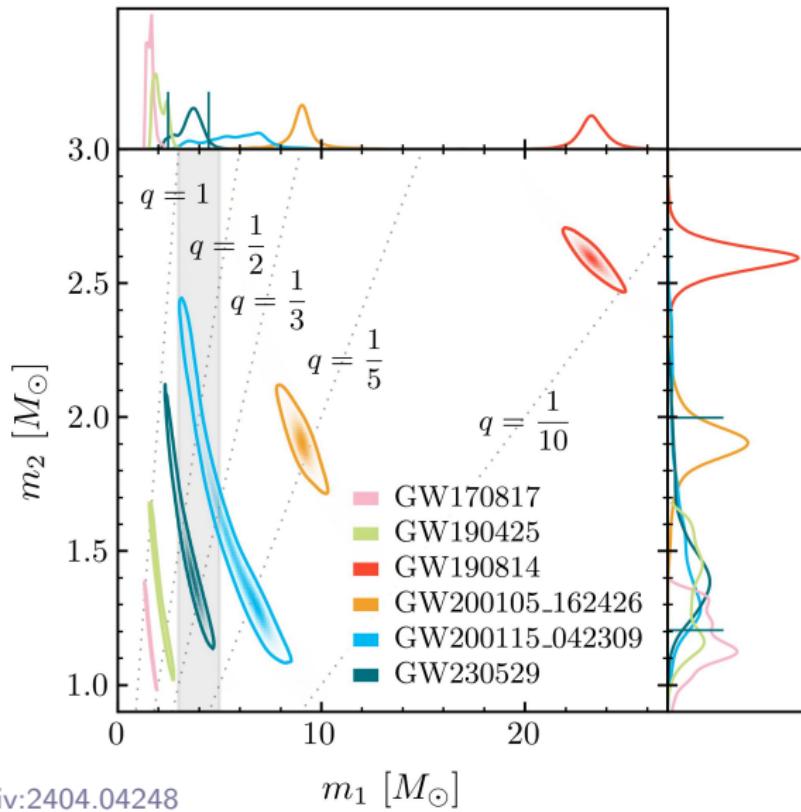
O3b highlights

neutron star – black hole mergers!



LVK: Abbott+ ApJL **915** #1 L5 (2021); LVK: Abbott+ arXiv:2111.03606; LIGO-G2102416

O4a: GW230529



FILLING THE MASS GAP

with observations of compact binaries from gravitational waves



Mass of compact object (M_\odot) 1

10

2

3

4

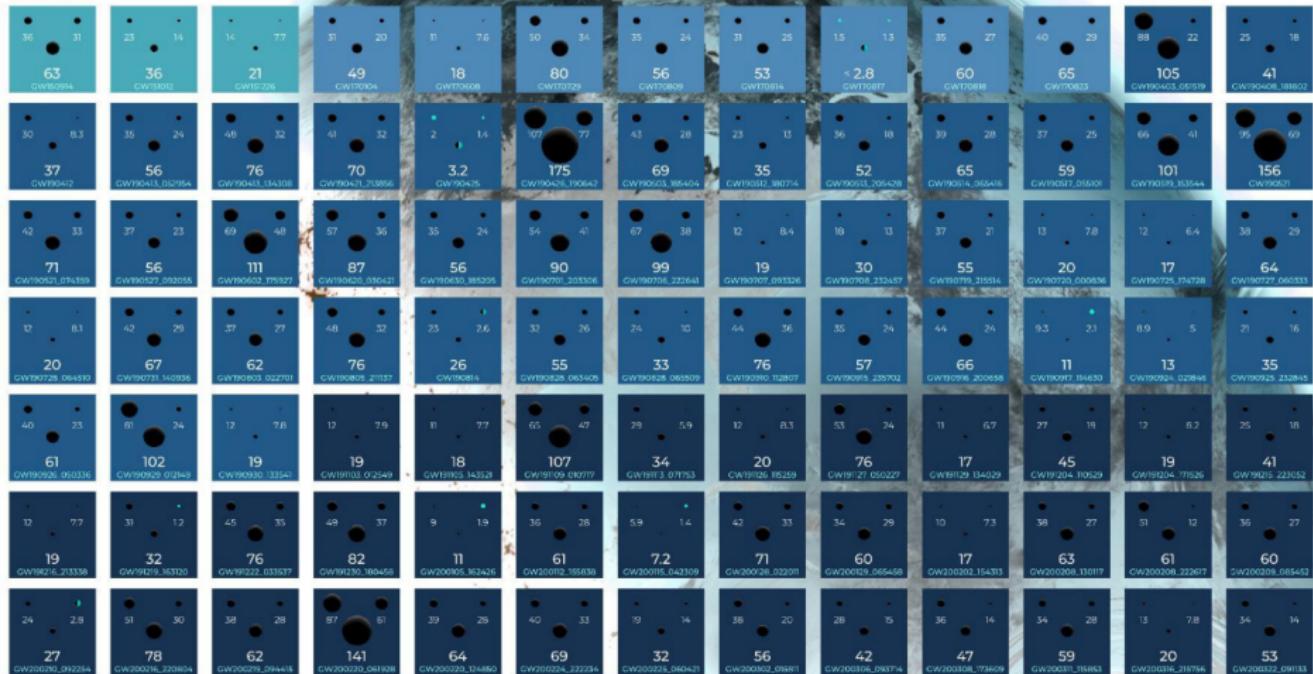
5

6

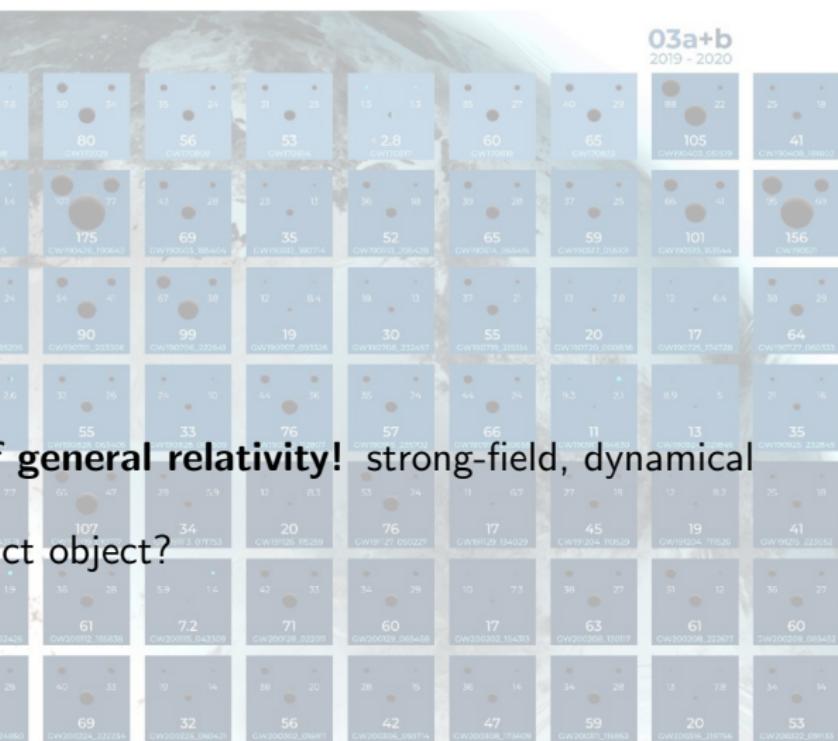
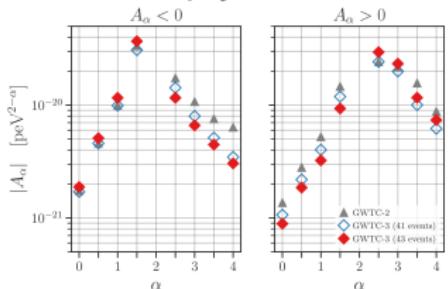
Includes components of compact binary mergers detected with a False Alarm Rate (FAR) of less than 0.25 per year

OBSERVING 01 — RUN

2015 - 2016



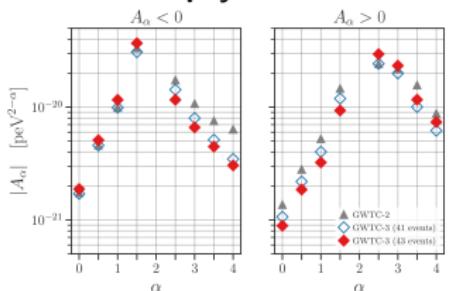
Fundamental physics



Several tests of general relativity! strong-field, dynamical

Nature of compact object?

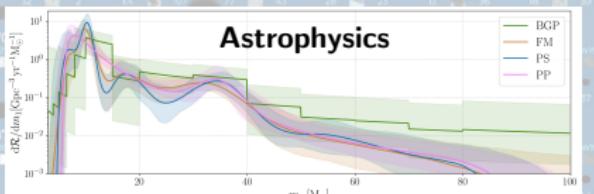
Fundamental physics



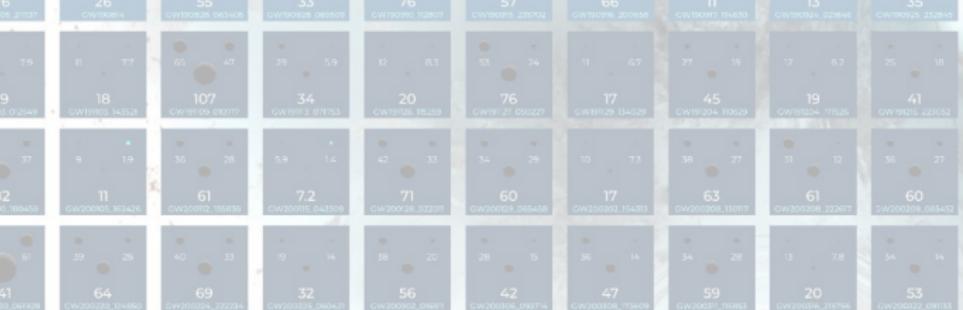
mass of graviton

$$m_g^{37} < 1.27 \times 10^{-23} \text{ eV}/c^2$$

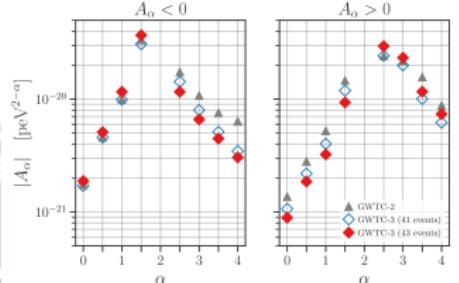
LVC: Abbott+ arXiv:2112.06861



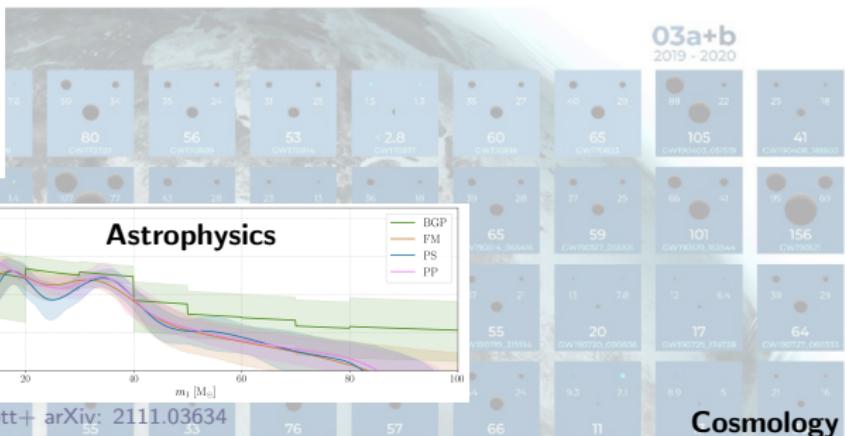
LVC: Abbott+ arXiv: 2111.03634



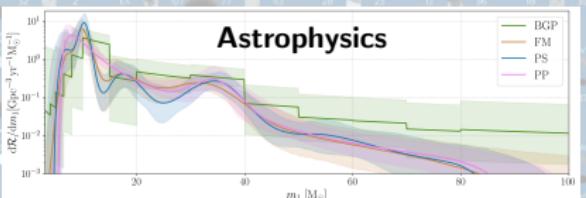
Fundamental physics



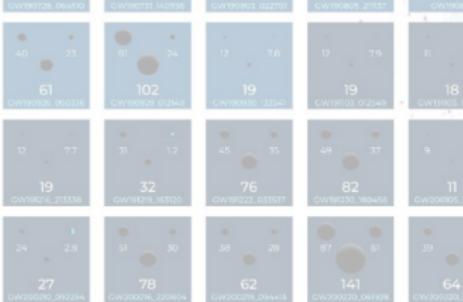
$$m_g < 1.27 \times 10^{-23} \text{ eV}/c^2$$



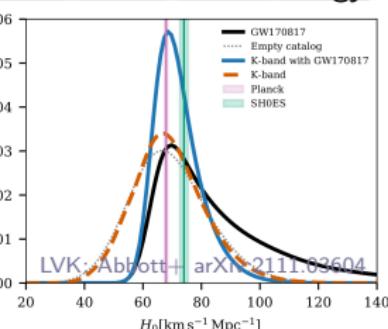
Astrophysics



LVC: Abbott+ arXiv: 2111.03634



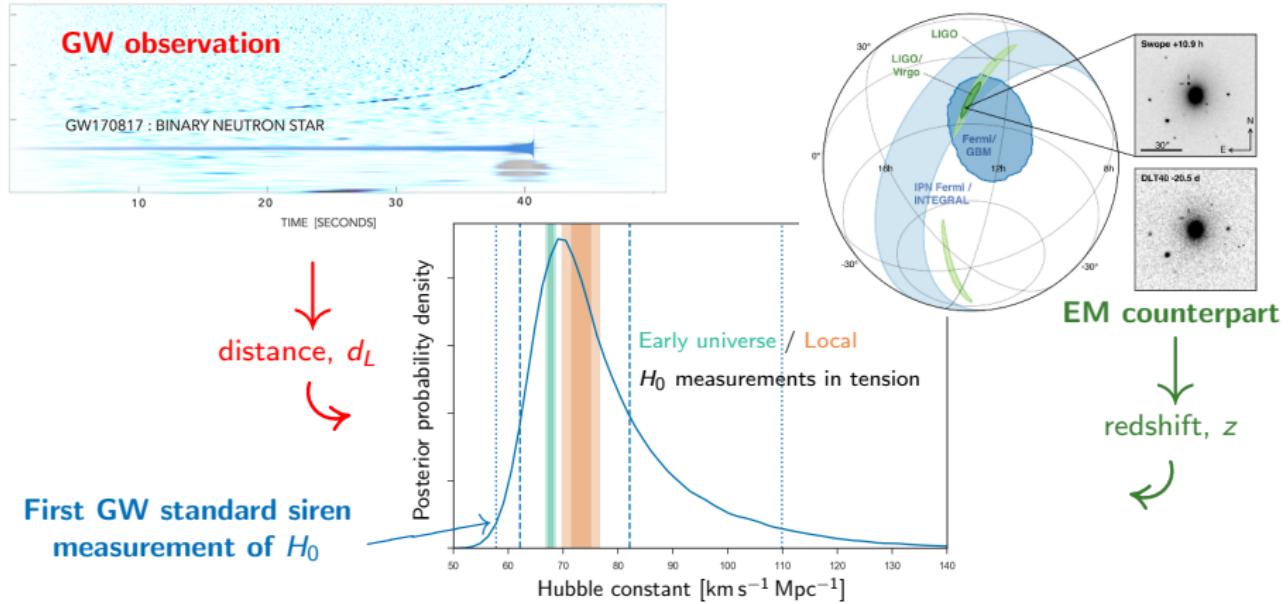
3a+b
9-2020



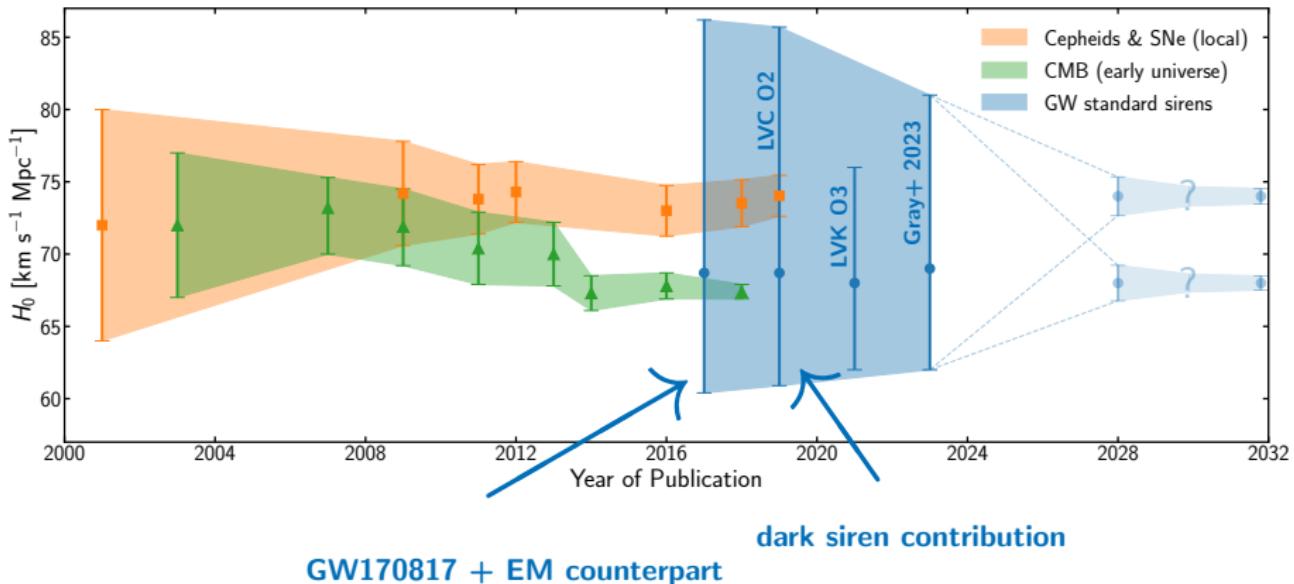
A gravitational-wave standard siren measurement of the Hubble constant

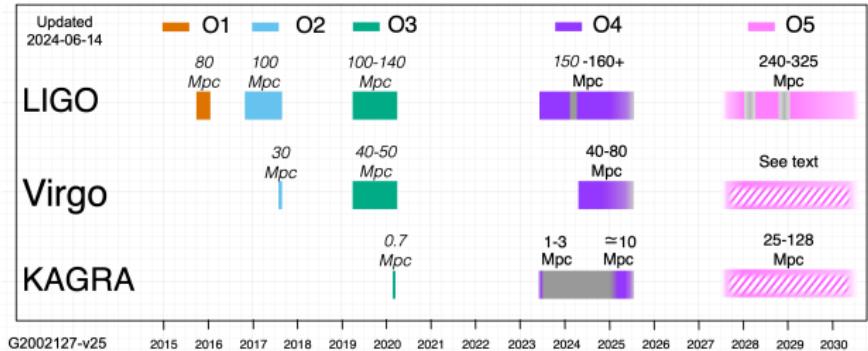
self-calibrated distance indicator

The LIGO Scientific Collaboration and The Virgo Collaboration*, The 1M2H Collaboration*, The Dark Energy Camera GW-EM Collaboration and the DES Collaboration*, The DLT40 Collaboration*, The Las Cumbres Observatory Collaboration*, The VINROUGE Collaboration* & The MASTER Collaboration*

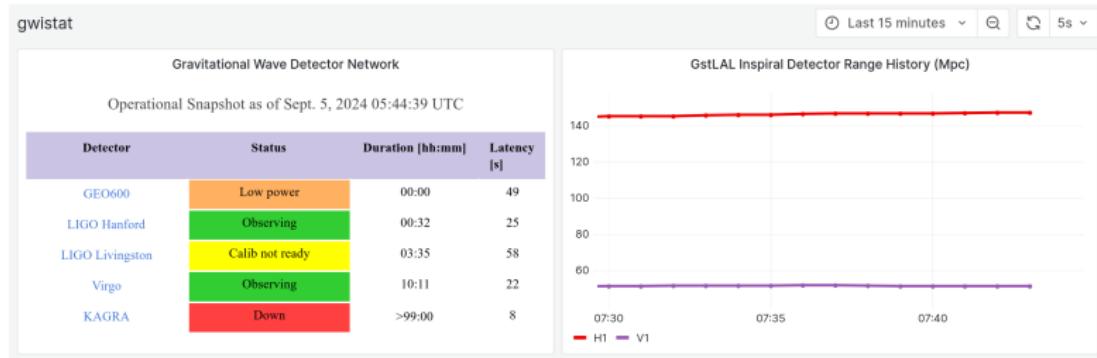


Gravitational-waves to resolve the H_0 tension?

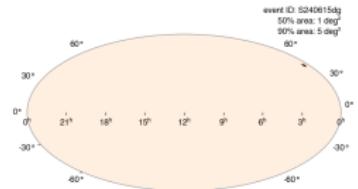
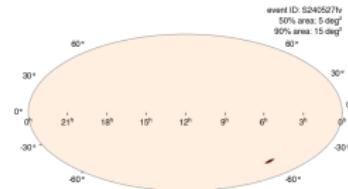
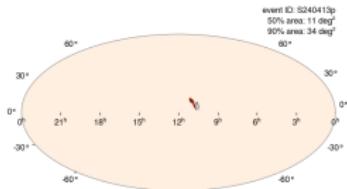




O4a: 2023 May 24 – 2024 Jan 16
O4b: 2024 Apr 10 – 2025 Jun 09



<https://gracedb.ligo.org/>: 124 significant detection candidates in O4



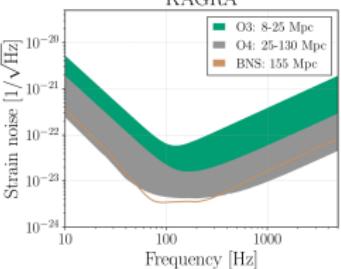
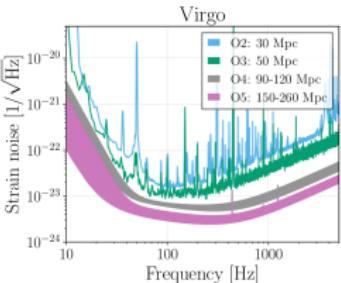
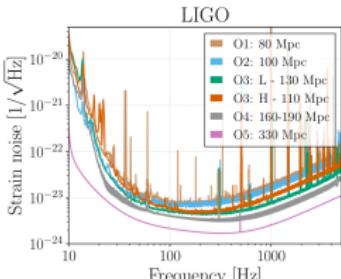
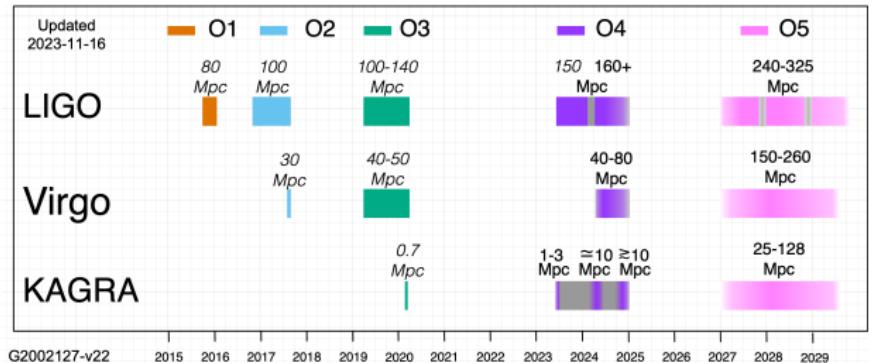
Towards 3G and the Einstein Telescope

Updated
2023-11-16

LIGO

Virgo

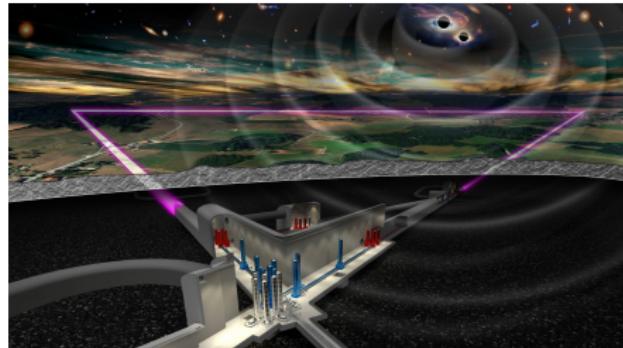
KAGRA



O5 / A+ → A[#] / V_nEXT upgrades

Einstein Telescope, Cosmic Explorer, LISA

3G ground-based detectors: ET and CE



Sos-Enattos, Sardinia | Euregio Meuse-Rhine (EMR)

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Strategic Plan on Research Infrastructures
ROADMAP 2021 Part I STRATEGY REPORT Part II LANDSCAPE ANALYSIS Part III PROJECTS & LANDMARKS Annex PEOPLE

INTERCONNECTIONS

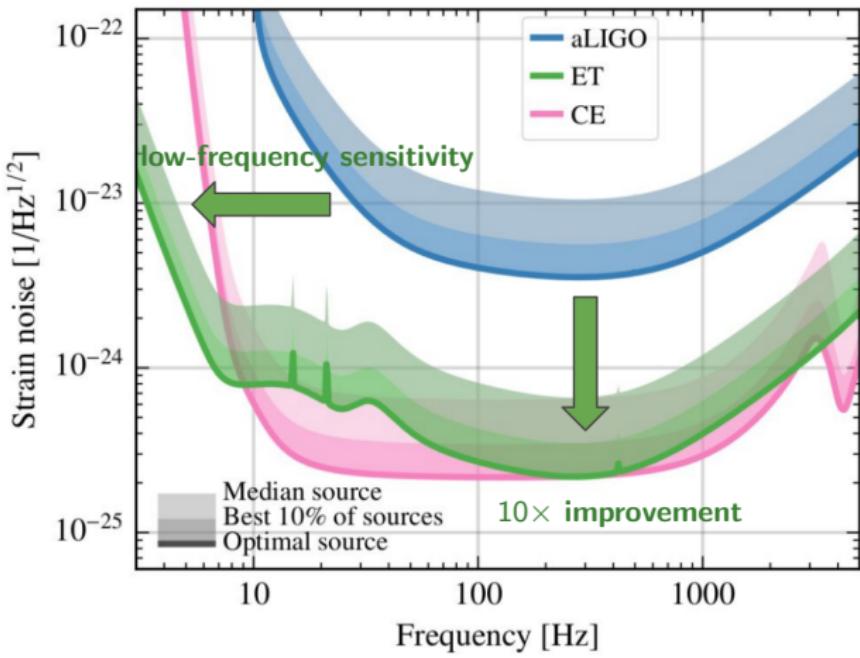


POLITICAL SUPPORT

Lead
IT
Prospective member
BE, ES, NL, PL



3G ground-based detectors: ET and CE



Particularly for ET:

longer arms ($\gtrsim 10$ km)
underground
active subtraction
“xylophone” (HF-LF)
cryogenic
glass \longrightarrow silicon
laser wavelength?
coating material?
quantum technology

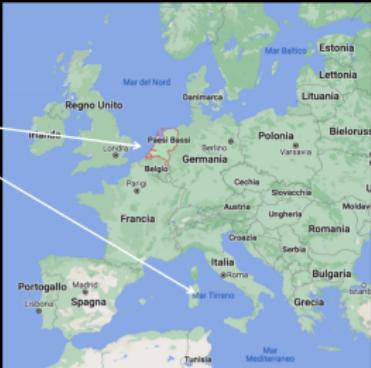
Science with the Einstein Telescope: a comparison of different designs

Marica Branchesi,^{1,2} Michele Maggiore,³ David Alonso,³ Charles Badger,⁴ Biswajit Banerjee,^{1,2} Freja Beirnaert,¹ Elsin Belgacem,⁵ Svetlana Bhagat,⁹ Guillame Bouleau,¹⁰ Sohrab Borhanian,¹² Darius Brown,¹³ Leon Meun Chan,¹⁴ Giulia Cusin,¹⁵ Stefan L. Danilishin,¹⁰ Jerome Degallaix,¹⁶ Valerio De Luca,¹⁷ Arnaud Demarteau,¹⁸ Tim Dietrich,^{1,2} Fabrizio Di Giacomo,¹⁹ Daniel Dittmar,²⁰ Alessandro Drago,²¹ Andrea Fossi,²² Gianfranco Gammie,²³ Boris Goncharov,^{2,24} Archisman Ghosh,²⁵ Francesca Golimini,²⁶ Ish Gupta,²⁷ Pawan Kumar Gupta,²⁸ Nandini Haera,^{1,2,27} Steven Hild,^{2,27} Tanya Hunter Hindler,² Kiong Hung,²³ Francesco Iaccovi,²⁹ Justin Januart,¹⁸ Kamiel Janssens,^{10,11} Alexander C. Jenkins,³⁰ Chinmay Kalaghatgi,^{18,31} Xishua Khorasheh,^{22,32} Tjonne G.F.C. Li,³³ Yufeng Li,³⁴ Eleonora Loffredo,¹ Elisa Maggio,³⁵ Michele Mancarella,^{3,36,37} Alessandro Manzoni,³⁸ Katerina Maroulis,³⁹ Andrea Bassetti,^{1,2} Robert Meyer,⁴⁰ Andrew L. Miller,⁴¹ Michael Mousset,⁴² Niccolò Mattioli,⁴³ Raula Narro,²⁵ Miriam Orentet,⁴⁴ Gor Oganessian,^{1,2} Costantino Paciolla,³⁷ Cristiano Palombo,⁴⁵ Paola Pani,¹ Antonio Pasquetti,⁴⁶ Albino Perego,^{17,48} Carla Pérezig,⁴⁹ 39,40,51 Mauro Pieroni,³⁹ 50,51 Omella Julianne Piccini,⁵² Anna Pucher,^{18,46} Paola Puppo,⁴⁵ Angelo Ricciardo,⁵³ Antonio Riotta,^{4,46} Samuele Ronchini,^{1,2} Mai Sakellariou,⁵⁴ Anindita Samajdar,⁵⁵ Giacomo Sartori,^{56,57,58} James Steinberg,^{1,2} Sebastian Steinberg,^{59,60} Aude Utien,^{16,17,51} Christophe Vuilleumier,⁵² and Tong Zhou,^{61,62}

JCAP07(2023)068

- 
 - Two scenarios
 - D of 10 km
 - D of 15 km

- Two scenarios
 - 2L of 15 km
 - 2L of 20 km

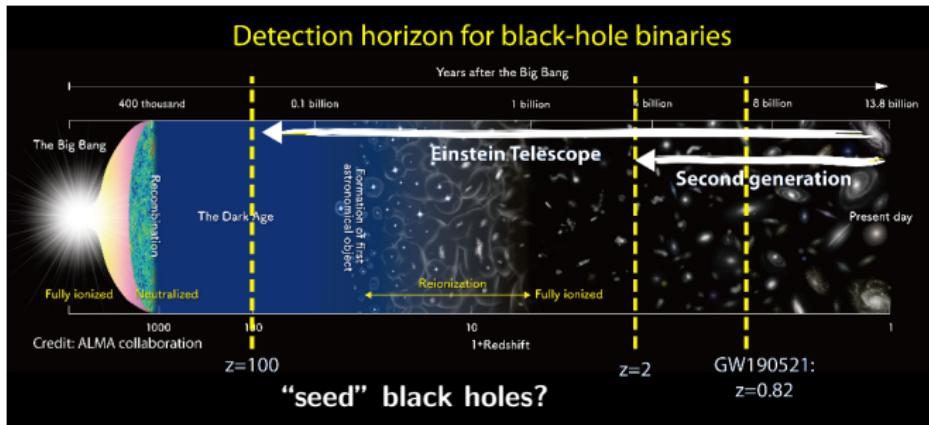
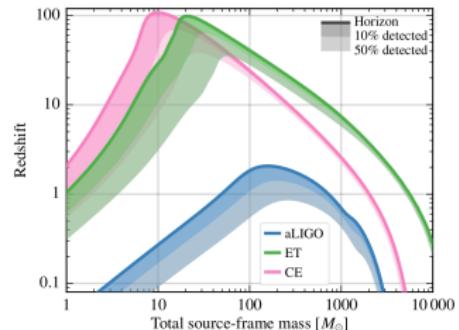


Branchesi, Maggiore et al. 2023,

Contents		
1	Introduction	6
2	Detector geometries and sensitivity curves	6.1
3	Cohesiveness of compact binaries	6.1.1
3.1	Binary Black Holes	Physical near the BH horizon
3.1.1	Comparison between geometries	Testing the GR predictions for space-time dynamics near the horizon
3.1.2	Effects of a change in the ASD	Constraining the effects for echoes and near-horizon structures
3.1.3	Golden events	Constraining tidal effects and multipolar structure
3.2	BH population studies	Nuclear physics
3.2.1	Comparison between geometries	Radius estimation from Fisher-matrix computation
3.2.2	Effects of a change in the ASD	Full parameter estimation results
3.2.3	Golden events	Connected uncertainty of nuclear-physics parameters
3.2.4	Dependence on the population model	Conclusions
3.3	ET in a network of 3G detectors	Population studies
3.3.1	Multimessenger astrophysics	Merge rate reconstruction
4	BNS sky-localisation and pre-m merger alerts	Constraints on PBHs from high-redshift merges
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4.2.2	Aftershow: survey and pointing modes	Hubble parameter and dark energy from joint GW/EM detections
4.3	Kilonovae: joint GW and optical detections	Hubble parameter and dark energy from BNS tidal deformability
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5.2	Angular sensitivity	Cosmic Strings
5.3	Astrophysical backgrounds	Gravitational phase transition
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5.6		Search for dark matter with CWS
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6.1.3	Constraining tidal effects and multipolar structure	
6.2	Nuclear physics	
6.2.1	Radius estimation from Fisher-matrix computation	
6.2.2	Full parameter estimation results	
6.2.3	Connected uncertainty of nuclear-physics parameters	
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6.3	Population studies	
6.3.1	Merge rate reconstruction	
6.3.2	Constraints on PBHs from high-redshift merges	
6.3.3	Other PBH signatures	
6.4	GW detection	
6.4.1	Hubble parameter and dark energy from joint GW/EM detections	
6.4.2	Hubble parameter and dark energy from BNS tidal deformability	
6.4.3	Hubble parameter from high-sn ratio events	
6.5	Cosmological stochastic backgrounds	
6.5.1	Cosmic Strings	
6.5.2	Gravitational phase transition	
6.5.3	Source separation	
6.6	Continuous waves	
6.6.1	CWs from spinning neutron stars	
6.6.2	Transient CWS	
6.6.3	Search for dark matter with CWS	
6.6.4	Conclusions	
7	+Summary and Appendix	

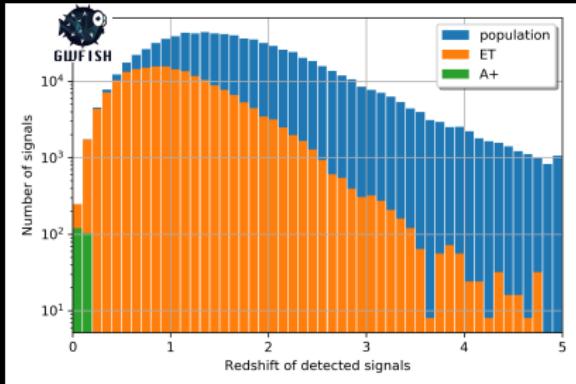
3G ground-based detectors: ET and CE

arXiv:1903.09260; courtesy: Evan Hall, Salvatore Vitale

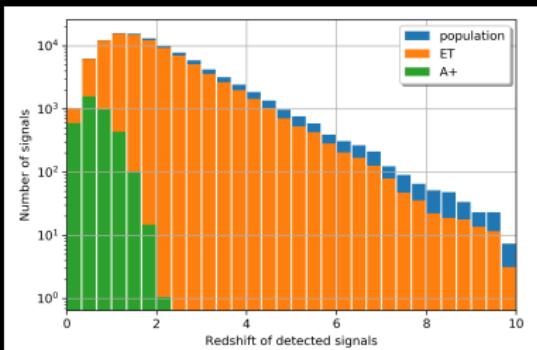


COMPACT OBJECT BINARY POPULATIONS

BINARY NEUTRON-STAR MERGERS



BINARY BLACK-HOLE MERGERS



Sampling **astrophysical populations**
of binary system of compact objects
along the cosmic history of the
Universe

10^5 BNS detections per year
 10^5 BBH detections per year

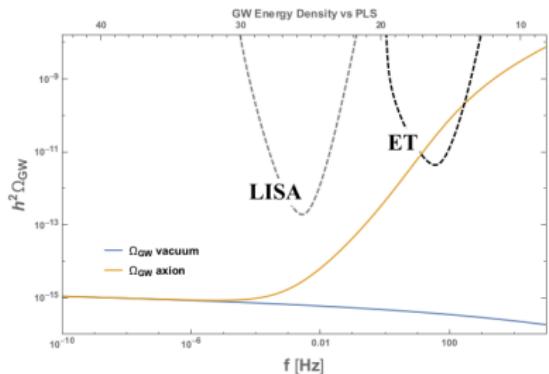
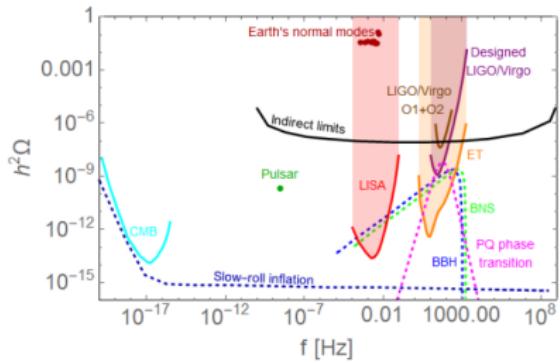
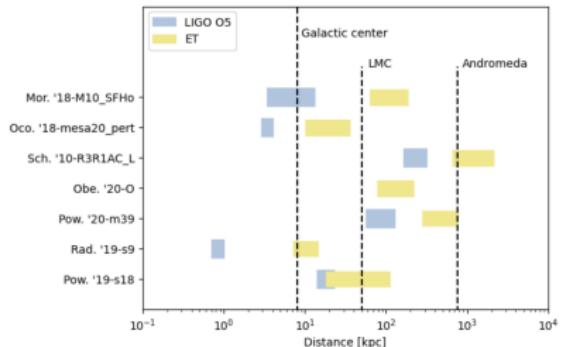
Other sources

Supernovae?

Isolated NS?

GW background?

Something exotic??



Data analysis challenges

- $\mathcal{O}(10^6)$ compact binary detections \Rightarrow big data!
- Very long signals \Rightarrow need new data analysis methods!
- Overlapping signals.



ET Collaboration

Site Characterization

electronic Infrastructure

Waveforms

Common Tools Stellar Collapse | NS

Data Analysis Platform Nuclear Physics

Instrument Science

Observational Science

Fundamental Physics Synergies w/ GW

Cosmology Multimessenger Obs

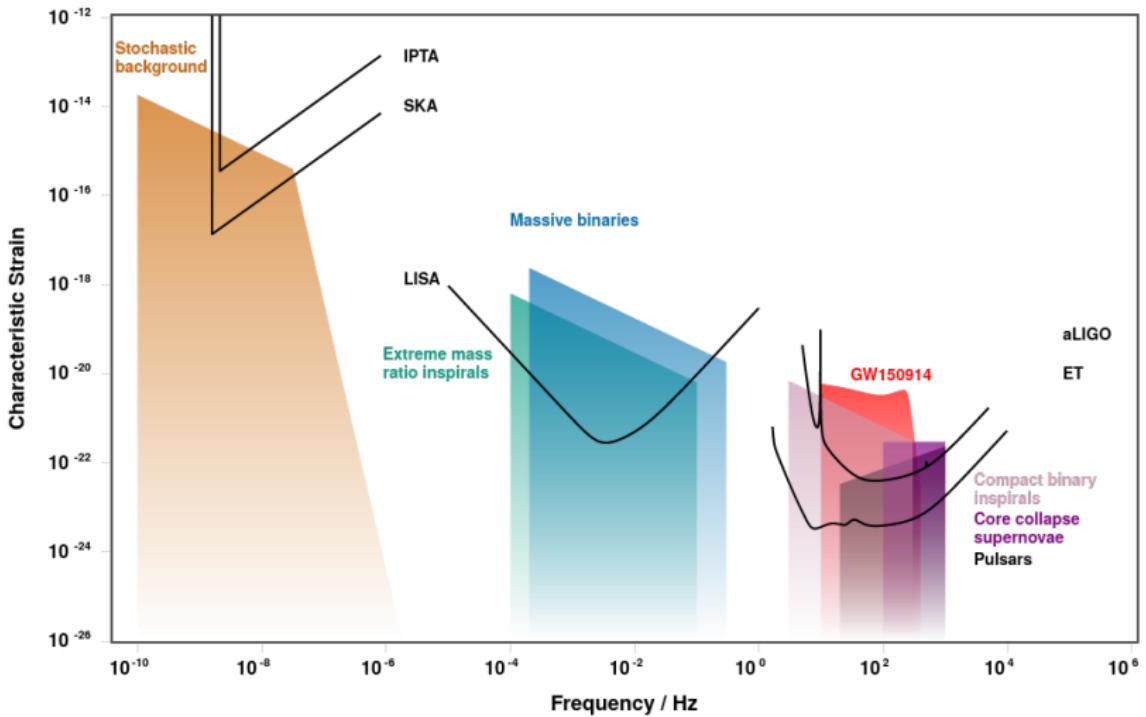
Population Studies

Forum on future EM and neutrino experiments?

EXTRA SLIDES

GWs in other bands

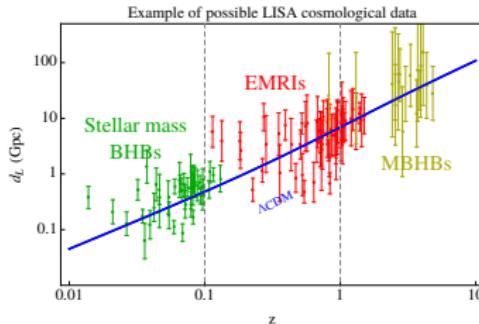
Moore, Cole, & Berry, <http://gwplotter.com/>



LISA sources

Standard sirens for LISA

Nicola Tamanini

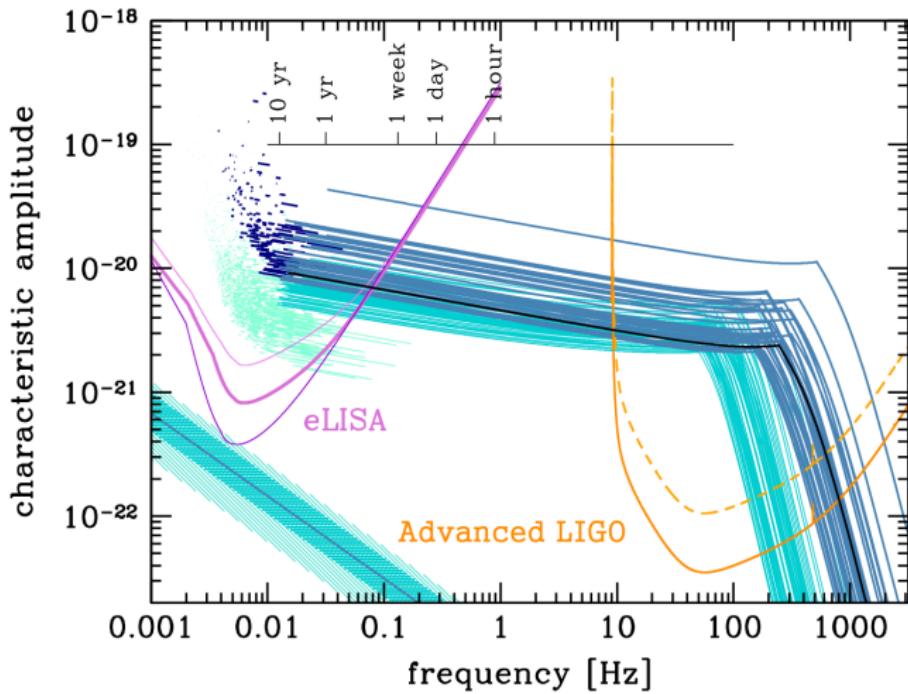


EM counterparts!

- ▶ StMBHBs: [Del Pozzo *et al.*, 1703.01300; Kyutoku & Seto, 1609.07142]
- ▶ EMRIs: [MacLeod & Hogan, 0712.0618]
- ▶ MBHBs: [Tamanini *et al.*, 1601.07112; Petiteau *et al.*, 1102.0769]

- **StMBHBs:** Del Pozzo *et al.* (2017); Kyutoku & Seto (2016)
- **EMRIs:** MacLeod & Hogan (2007)
- **MBHBs:** Tamanini *et al.* (2016); Petiteau *et al.* (2011)

Multiband GW astronomy



Sesana 2016