# Material Assay Campaign for the DarkSide-20k experiment



Roberto Santorelli

on behalf of the DarkSide-20k collaboration











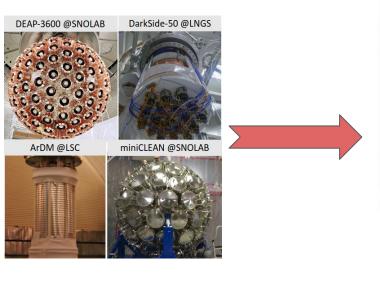
# **Outline**

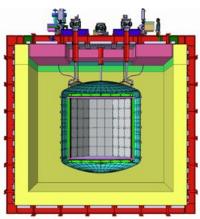
- DarkSide-20k
- Radiopurity requirements for WIMP searches with LAr
- The assay campaign:
  - ♦ Tools & organization
  - Results so far
- Conclusions and outlook

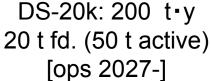
# **The Global Argon Dark Matter Collaboration**

Joint effort of **all** former DM experiments using Liquid Ar targets: Combined expertise from four LAr experiments to explore dark matter to the neutrino floor and beyond

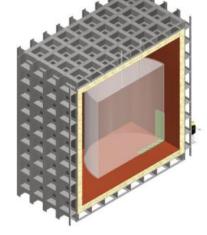
- > 500 scientists
- > 100 institutions
- 14 counties





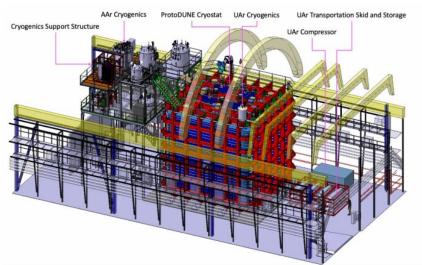


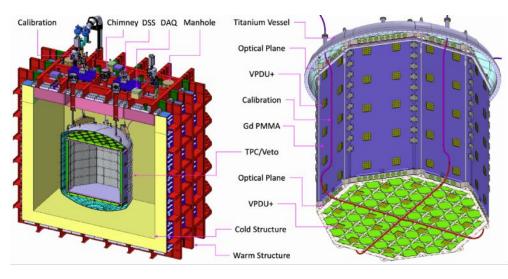




ARGO: ~ 3000 t·y 300 t fd. (400 t active) [early 2030's]

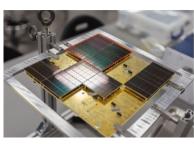
# **The DarkSide-20k experiment**

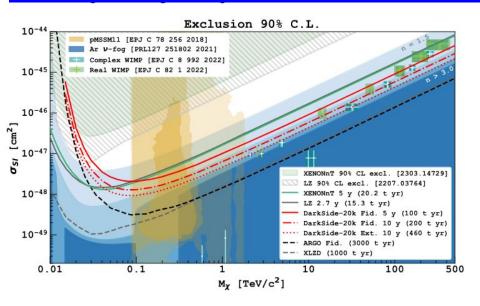




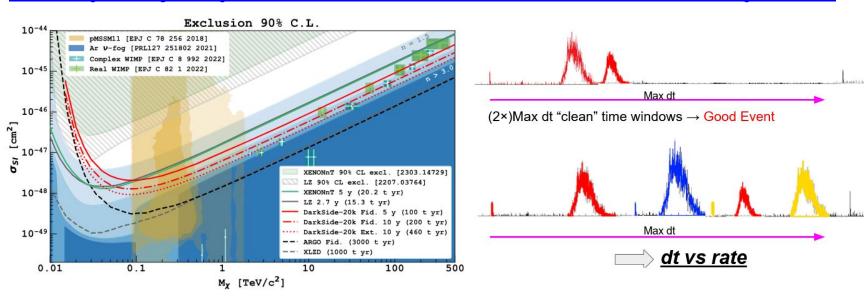


- Nested detector with
  - ProtoDUNE-like cryostat: 8 x 8 x 8 m<sup>3</sup> hosting 650 t of atm Ar
  - SS vessel+structure hosting 99 t underground Ar.
  - UAr + acrylic Veto for neutrons and γ.
  - Central Dual-phase Ar TPC:
    - **■** 50 t UAr (**20 t fiducial**)
    - 348 cm drift
    - Cathode @ -**73.4 kV**
- **21 m<sup>2</sup> of SiPM + 5 m<sup>2</sup> in veto.** 
  - Coverage in TPC Optical plane > 90%

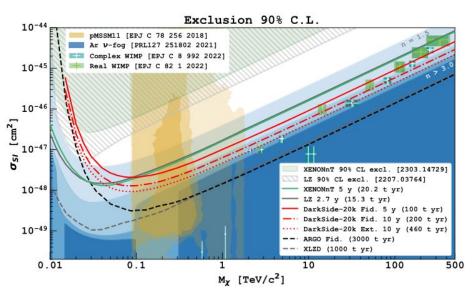


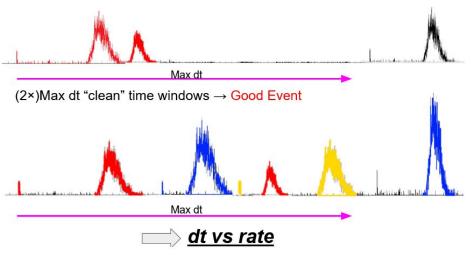


- neutrons after cuts < 0.1 in 10 y</p>
- < 0.05 from  $\beta$  and  $\gamma$

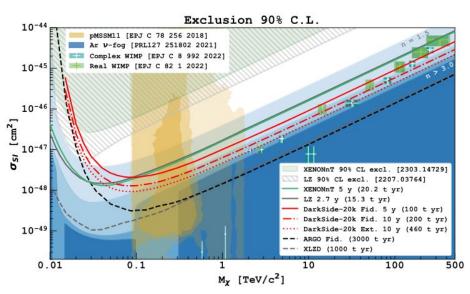


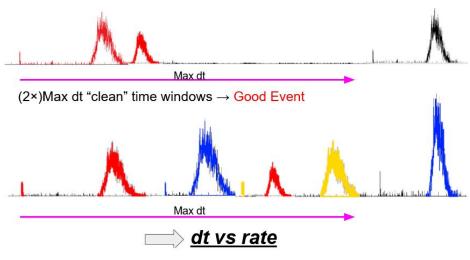
- neutrons after cuts < 0.1 in 10 y</p>
- $\blacksquare$  < 0.05 from  $\beta$  and  $\gamma$





- GOAL: Complete control over every component that goes into the detector (Down to the resistor level)
- To investigate the secular equilibrium of the U-Th decay chains
- Knowledge on the chemical composition of the components in order to calculate the n-yield via X(α,xn)Y



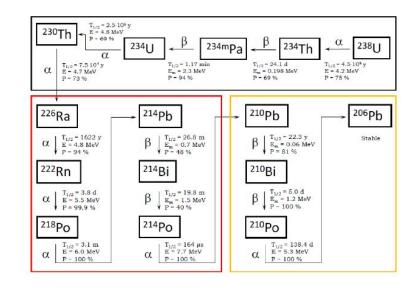


- GOAL: Complete control over every component that goes into the detector (Down to the resistor level)
- To investigate the secular equilibrium of the U-Th decay chains
- Knowledge on the chemical composition of the components in order to calculate the n-yield via X(α,xn)Y

- Full Monte Carlo simulations for the (α,xn) calculation with the detailed detector geometry
- Control on the surface exposure to air for the <sup>210</sup>Pb/<sup>210</sup>Po surface contamination, Cosmogenics...
- Handling, transportation, recontamination QA/QC...

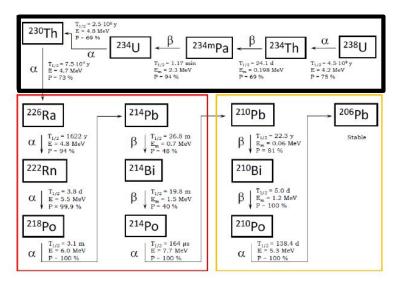
BREAKING OF THE
SECULAR EQUILIBRIUM AT

226Ra



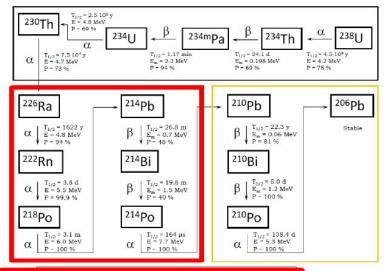
SECOND BREAKING OF THE EQUILIBRIUM IN <sup>210</sup>Pb

# **ICPMS**



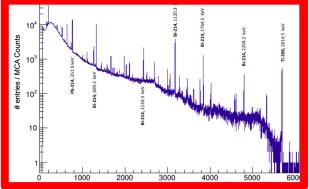


- Very little mass needed
- Relatively fast
- Digestion
- Destructive
- Only U-238 (Th-232)

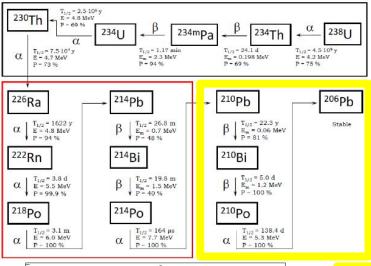






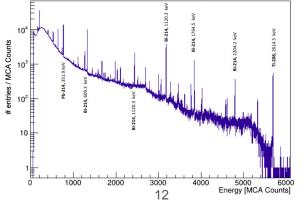


- Larger mass needed
- Slow
- Non-destructive
- Gamma activity (40K+...)

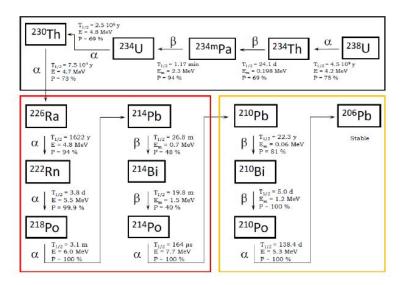


## **Radiochemical**







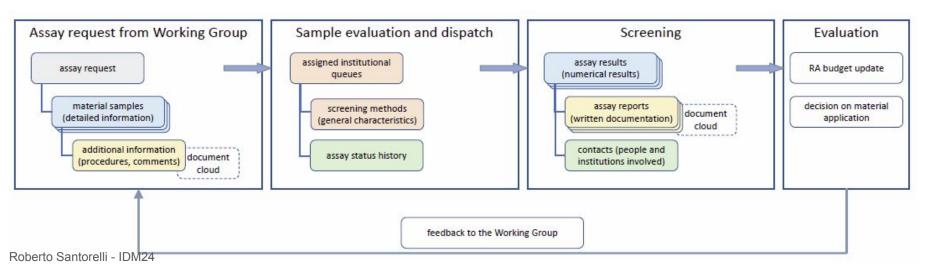


- XIA surface (and not so surface) screening
- Monitor the history of exposure of materials to cosmogenic radiation.
- Worldwide effort involving
  - o underground facilities: LSC (Sp), Gran Sasso (It), Boulby (UK), SNOLAB (Ca)
  - + CIEMAT (Sp), Jagiellonian Univ. (PI), Aix-Marseille U. (Fr), Mendeleev U. (Ru), Temple (US)...

# **DarkSide materials DB structure**

Online database that centralizes the full assay process

- New material or component? Assay request!
- Sample allocation depending on available mass and needs
- Information on sample/assay status
- Storage and organization of results.



14

# **DarkSide materials Web interface**

#### Online database that centralizes the full assay process

- New material or component? Assay request!
- Sample allocation depending on available mass and needs
- Information on sample/assay status
- Storage and organization of results.

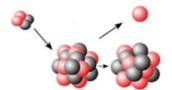
#### DB: Web-interface



# **Neutron yield calculation**

$$Y_i(E_{\alpha}) = \frac{\eta_i}{\eta} \int_0^{E_{\alpha}} \frac{\sigma_{(\alpha, X_n)}^i(E)}{\varepsilon(E)} dE$$

 $E_{\alpha}$  is the initial energy of the  $\alpha$  particle;  $\eta_i$  is the number density of nuclide i;  $\eta$  is the number density of the material;  $\sigma^i(\alpha,Xn)(E)$  is the neutron production x-sec for the nuclide i $\varepsilon(E) = -\frac{dE}{dx}$  is the stopping power of the material

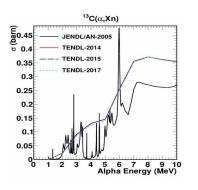


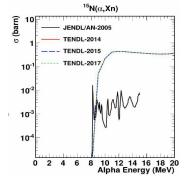
DarkSide is the first experiment with the (α,n) neutron background fully calculated with Geant4



(http://win.ciemat.es/SaG4n/)

"Neutron production induced by α-decay with Geant4", Nucl. Instrum. Methods A 960, 163659 (2020)





- Exploiting evaluated libraries (JENDL)
- Detailed geometries (actual geometry and border effects)
- Neutron transport, precise tracking
- Biasing techniques

# (α,n) white paper

### 2405.07952

#### White paper on (a, n) neutron yield calculations

D. Cano-Ott, <sup>1</sup> S. Cebrián, <sup>2</sup> M. Gromov, <sup>3,4</sup> M. Harańczyk, <sup>5</sup> A. Kish, <sup>6</sup> H. Kluck, <sup>7</sup> V. A. Kudryavtsev, <sup>8</sup> I. Lazanu, <sup>9</sup> V. Lozza, <sup>1</sup>, <sup>1</sup>, <sup>1</sup>, <sup>6</sup>, Luzón, <sup>2</sup> E. Mendoza, <sup>1</sup> M. Parvu, <sup>9</sup> V. Pesudo, <sup>1</sup> A. Pocar, <sup>12</sup> R. Santorelli, <sup>1,a</sup> M. Selvi, <sup>13</sup> S. Westerdale, <sup>14</sup> and G. Zuzel<sup>5</sup>

<sup>1</sup>CIEMAT, Centro de Investigaciones Energéticas, Medioambientales y Tecnológicas, Madrid 28040, Spain

<sup>2</sup>CAPA, Centro de Astropartículas y Física de Altas Energías, Universidad de Zaragoza, Zaragoza 50009, Spain

<sup>3</sup>Skobeltsyn Institute of Nuclear Physics, Lomonosov Moscow State University, Moscow 119234, Russia
<sup>4</sup>Joint Institute for Nuclear Research. Dubna 141980. Russia

<sup>5</sup>M. Smoluchowski Institute of Physics, Jagiellonian University, 30-348 Krakow, Poland <sup>6</sup>Fermi National Accelerator Laboratory, Batavia, IL 60510, U.S.A

<sup>7</sup>Institut für Hochenergiephysik der Österreichischen Akademie der Wissenschaften, 1050 Wien, Austria

<sup>8</sup>Department of Physics and Astronomy, University of Sheffield, Sheffield S3 7RH, UK

9 Faculty of Physics, University of Bucharest, POBox 11, 077125, Magurele, Romania

Laboratório de Instrumentação e Física Experimental de Partículas (LIP), 1649-003, Lisboa, Portugal

<sup>11</sup>Universidade de Lisboa, Faculdade de Ciências (FCUL), Departamento de Física, 1749-016 Lisboa, Portugal

<sup>12</sup>Amherst Center for Fundamental Interactions and Physics Department, University of Massachusetts, Amherst, MA 01003, USA

13 INFN - Sezione di Bologna, Bologna 40126, Italy

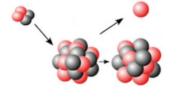
<sup>14</sup>Department of Physics and Astronomy, University of California, Riverside, CA 92507, USA (Dated: Tuesday 28th May, 2024-00:28, Version: F1.0)

Understanding the radiogenic neutron production rate through the  $(\alpha, n)$  reaction is essential in many fields of physics like dark matter searches, neutrino studies, nuclear astrophysics and medical physics. This white paper provides a review of the current landscape of  $(\alpha, n)$  yields, neutron spectra and correlated y-rays calculations, and describes the existing tools and the available cross sections. The uncertainties that contribute to  $(\alpha, n)$  yield calculations are also discussed with plans for a program to improve the accuracy of these estimates. Novel ideas to measure  $(\alpha, n)$  cross sections for a variety of materials of interest are presented. The goal of this study is to reduce the uncertainty in the expected sensitivity of next-generation physics experiments in the keV–MeV regime.

#### CONTENTS

I. Introduction

alphan@ciemat.es



# Multidisciplinar WG on (α,n) neutron yield studies alphan@ciemat.es

Members of several DM experiments + Neutrino + Nuclear + IAEA:

- DarkSide-20k
- XENON
- LZ
- CRESST ...

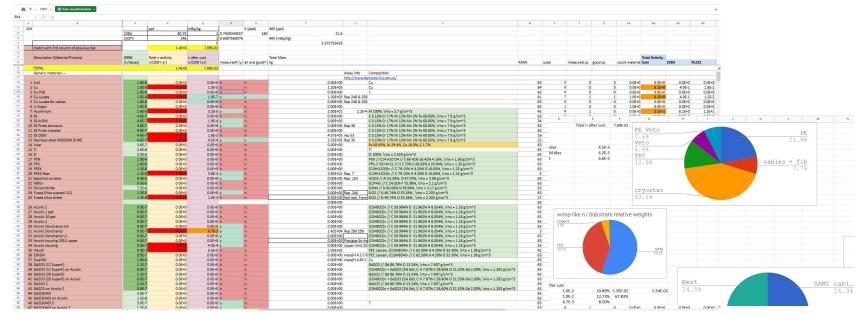
Understanding the radiogenic neutron production rate through the  $(\alpha, n)$  reaction is essential in many fields of physics like dark matter searches, neutrino studies, nuclear astrophysics and medical physics. This white paper provides a review of the current landscape of  $(\alpha, n)$  yields, neutron spectra and correlated  $\gamma$ -rays calculations, and describes the existing tools and the available cross sections. The uncertainties that contribute to  $(\alpha, n)$  yield calculations are also discussed with plans for a program to improve the accuracy of these estimates. Novel ideas to measure  $(\alpha, n)$  cross sections for a variety of materials of interest are presented. The goal of this study is to reduce the uncertainty in the expected sensitivity of next-generation physics experiments in the keV–MeV regime.

# **Background budget spreadsheet**

# R.A. budget spreadsheet

Storing the results of the samples, materials composition, contribution to the bkg ...

(α,n) neutron background γ event rate(VETO+TPC)



# **Operation protocols**

Protocols that account for all steps in the manipulation of every component that makes it into the detector:

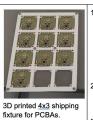
- Storage upon arrival
- Storage during any intermediate step
- Every time that things need to be unbagged and rebagged
- Manufacturing
- Assembly
- Cleaning
- Testing
- Final storing
- Final assembly

# **Operation protocols**

Written procedures for each step, including radio-purity control Storage of wafers

#	Step	Details
1	Receipt of wafers	Wafer shipments arrive at Liverpool containing wafers for both STFC and Liverpool. Typically half the quantity received will be forwarded to STFC in the original NOA packaging. Half will be retained at Liverpool. Upon receipt of wafers they (all wafers, including ones forwarded to STFC) are tagged in the database as being located at Liverpool using ds20k_veto_location_gui. For the wafers retained, these are typically vacuum-bagged in batches of 5 in Reber food bags. These are then transferred to the cleanroom and moved to nitrogen storage for picking later.

PCB shipping



Preparing

for shipping

nuts.

used.

Secure lid of shipping fixture using m4 bolts and

Place the fixture into an ESD bag.

Place PCBs into a 3D

printed PLA shipping

ultrasonically cleaned in

fixture. Shipping fixtures are

ethanol every time they are

Shipping fixture

Vacuum holds

bag.

secure and stal

Vacuum seal the bag.

Place the vacuum sealed bag into another ESD bag and seal the outer bag at atmospheric pressure. In this way the inner bag is surrounded by low Rn cleanroom air to avoid contamination in the case of vacuum loss. \*

**vPDU** testing, Initial steps

No.	Step	Description	Remarks
1	vPDU received	Move the vPDU in the clean room, Remove the vPDU from PeliCase in the clean room lobby	
2.	Environment check	Particle counter measurement before un-bagged the vDPU	ISO7: 352,000 particles per cubic meter or less ≥0.5 μm sized particles, 83,200 particles per cubic meter or less ≥1 μm sized particles, and 2,930 particles per cubic meter or less ≥5 μm sized particles
3.	Un-bagged vDPU for testing	Remove vPDU from the bags and update the status on the database	

# **Operation protocols**



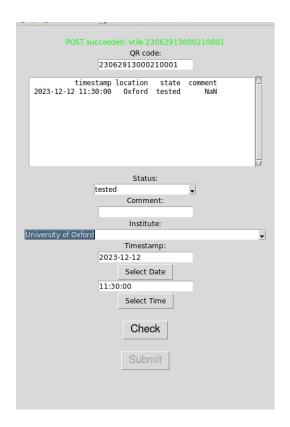


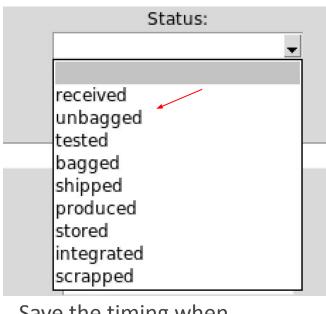


- ESD bag
- Reber bag for radon protection
- Desiccant and humidity control inside Reber bags
- Second Reber bag for radon protection

Protective cases

# **Exposure and tracking Database**





Save the timing when the status is changed on DB

Every relevant action on a vPDU is registered in a DB with an identifying code and a time stamp.

Each facility has an appointed person responsible for the interaction with the Materials WG.

# vTile populated vs sum of the components

	Summing components [mBq]	vTile fully assembled after testing [mBq]
U-up	1.3	1.1
U-mid	0.8	0.8
U-low	31.7	58.6
Th-232	0.7	0.34
U-235	0.1	0.05
K40	7.8	4.7

Passive (PCB, Resistors, Capacitors) and active el. components.
Copper pillars, connectors, soldering paste...

"Populated" tile radioactivity in very good agreement with the sum of the components

Relatively small difference in <sup>210</sup>Pb under study (not a show stopper)

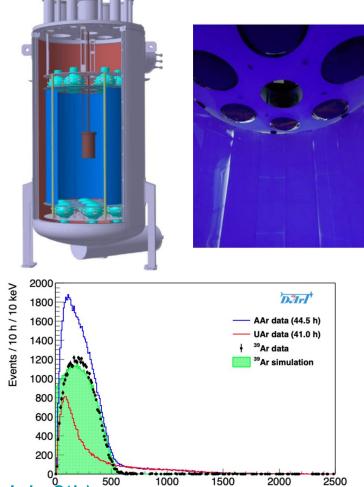
# **Other relevant topics**

- Cleaning protocols
- Control on the surface exposure to air:
  - <sup>222</sup>Rn decay daughters on surfaces (<sup>210</sup>Pb 22 y)
- Monitor the history of exposure of materials to cosmogenic radiation
  - "Study of cosmogenic activation above ground for the DarkSide-20k experiment"
     Astropart. Phys. 152 (2023) 102878
- Evaluation of the systematic uncertainty from the material composition and cross-sections
- Knowledge on the chemical composition of the components in order to calculate the n yield via  $X(\alpha,xn)Y$  with Geant4
- MANY (Measurement of Alpha Neutron Yields) collaboration
  - ->  $(\alpha,n)$  cross-section measurement with a thick target and  $\alpha$  beam

# **UAr assay @ LSC**







Energy [keV]

25

Talk by V. Pesudo (Monday, July 8th)

# **Conclusions and outlook:**

- DarkSide-20k has developed a very detailed material assay campaign for the full control / mitigation of the backgrounds
- Three assays per sample for the full control of the U-238 decay chain
- New tools for neutron yield calculation
- Control of cosmogenics, Rn, surface exposure...
- Cleaning and surface protocols

. . . .

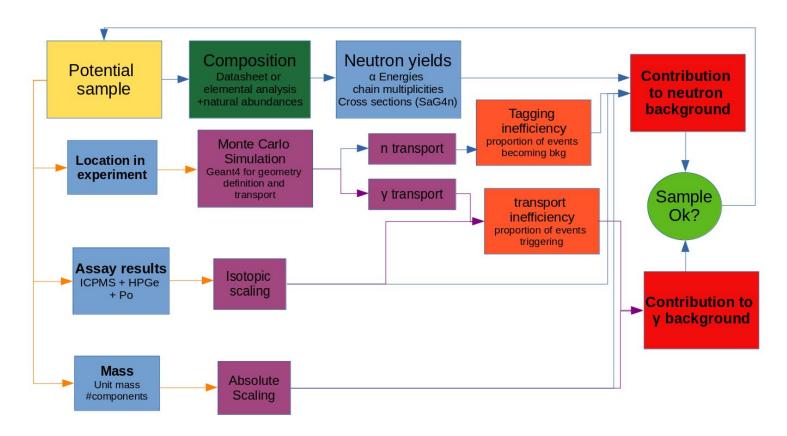
Pushing the state-of-the-art in many fronts in this field

Thank you!

Credits to: Chiara Chiarelli (L'Aquila, Italy)

# **Backup**

# **The decision making process**



# Th-232 equilibrium

