

Universität
Münster



XENON

WIMP Dark Matter searches with the XENONnT experiment

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on behalf of the XENON collaboration

living.knowledge



Bundesministerium
für Bildung
und Forschung

Contract number 05A20PM1

XENON Collaboration

~180 Scientists
29 Institutions

AMERICA

UC San Diego

San Diego



THE UNIVERSITY OF CHICAGO

Chicago

COLUMBIA UNIVERSITY
IN THE CITY OF NEW YORK

New York City

PURDUE
UNIVERSITY

Lafayette



Zurich



Karlsruhe Institute of Technology



Münster



Freiburg



Mainz



Heidelberg



Amsterdam



Stockholm



Rehovot



Nagoya



Kobe

ASIA



Tsinghua University

Beijing



Westlake University

Hangzhou



The Chinese University of Hong Kong

Shenzhen



THE UNIVERSITY OF TOKYO

Tokyo



Nagoya

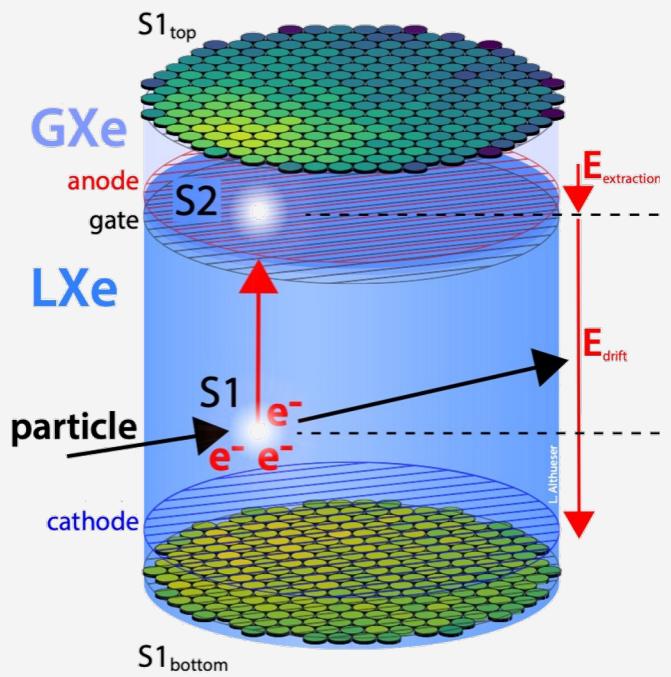


Kobe



The XENONnT Experiment

- Dark matter direct detection experiment
- At the INFN Laboratori Nazionali del Gran Sasso (LNGS) in Italy
- Underground ultra-low background experiment



Dual Phase Time Projection Chamber

- Scintillation and ionization
- Photosensor arrays at top and bottom
- Detection of direct scintillation light by PMTs (**S1**)
- Drift and extraction of electrons into the gas phase
- Secondary proportional scintillation signal (**S2**)



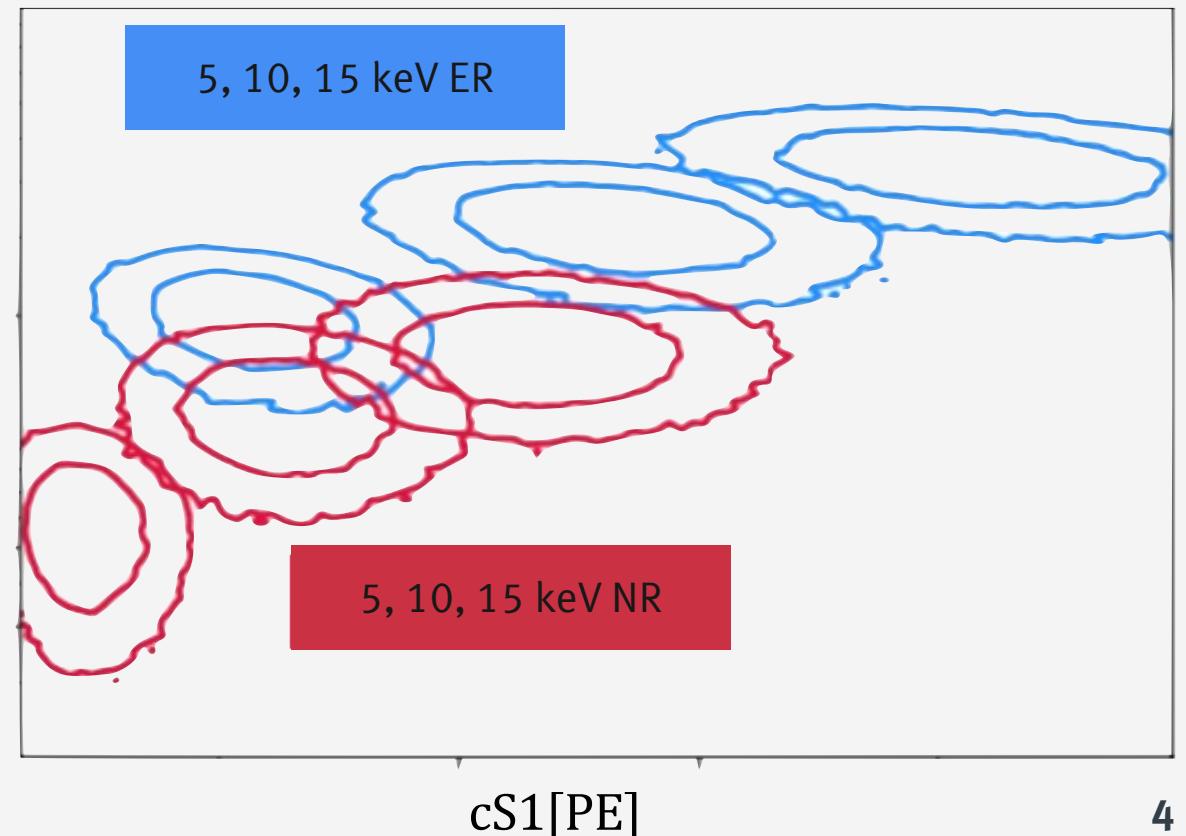
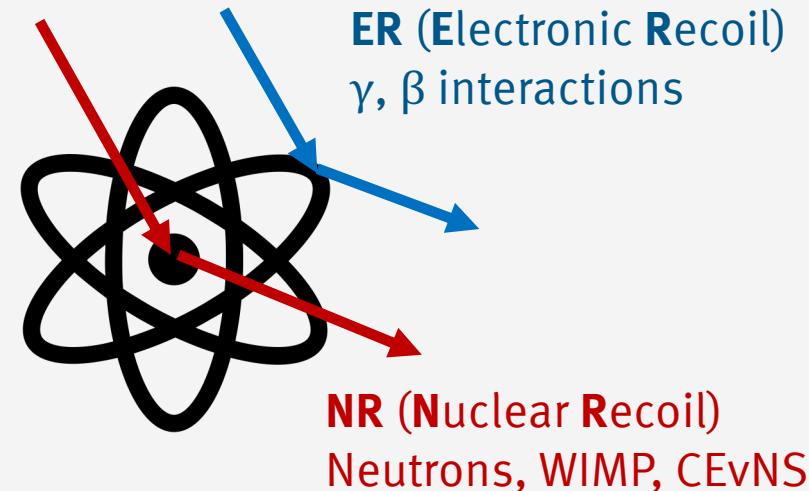
Signals in Liquid Xenon

Signal Type Discrimination

- Separation of ER and NR interactions by ratio of light to charge signals
- XENON detectors conceived and designed to search for NR signals created by **WIMPs**
- ~ 1keV ER recoil energy deposited in the liquid xenon target is enough to produce a characteristic scintillation and charge signal

Additional physics cases

- Solar axions, axion-like particles
- Solar neutrinos, SN neutrinos, CEvNS
- (neutrinoless) Double β -Decay, double EC
- Other BSM processes

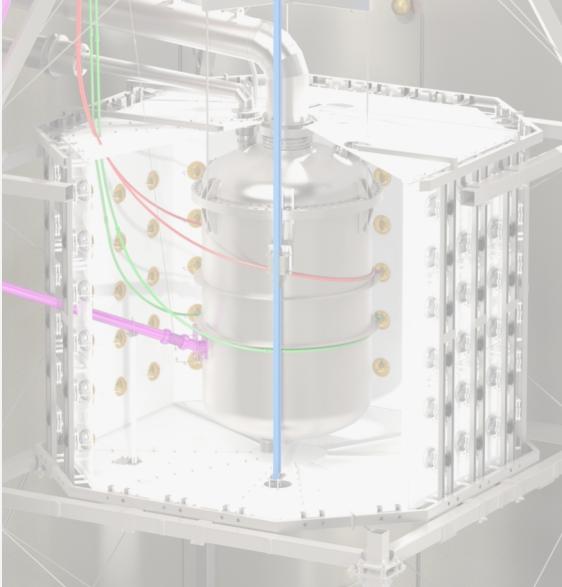


Three Nested Detectors



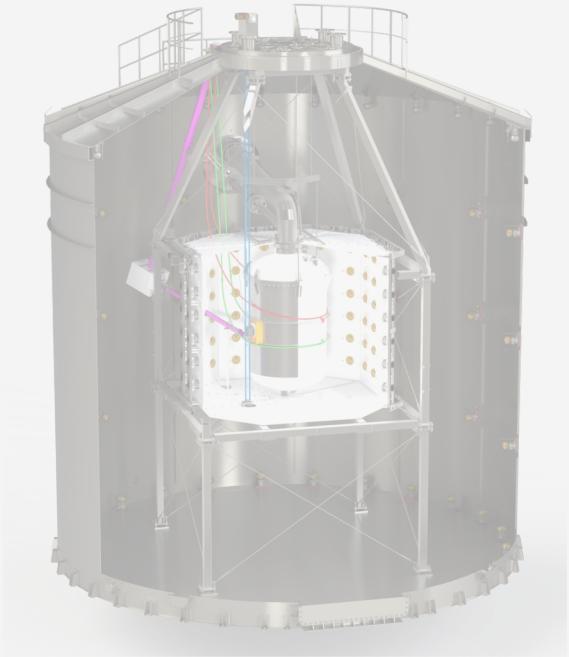
Time Projection Chamber (TPC)

- 8.5 t ultra pure xenon target
- 5.9 t Xe observed by 494 PMTs
- Active target diameter/height:
1.3m/1.5m



Neutron Veto (NV)

- (Gd-salted) Water Cherenkov detector
- 120 8" PMTs inside enclosure of reflective panels
- Neutron tagging efficiency of 53% in pure water



Muon Veto (MV)

- 700 t Water Cherenkov detector
- 84 8" PMTs
- Active veto against muons
- Passive shielding

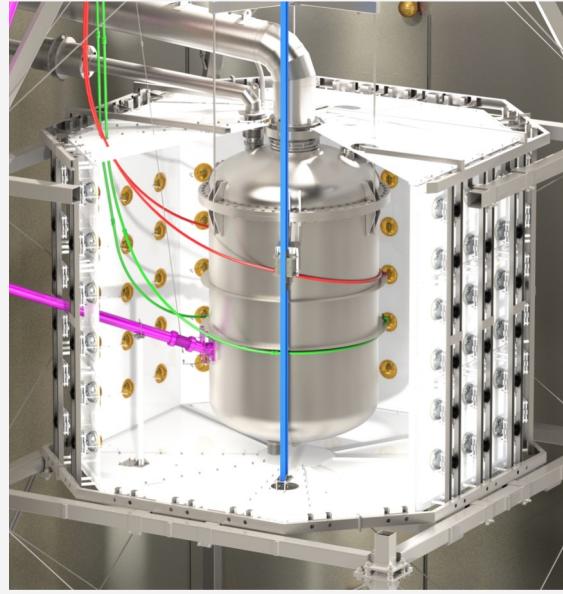
Three Nested Detectors

The XENONnT Neutron Veto: performances
without and with Gd-doping
Marco Selvi – Today 17:10



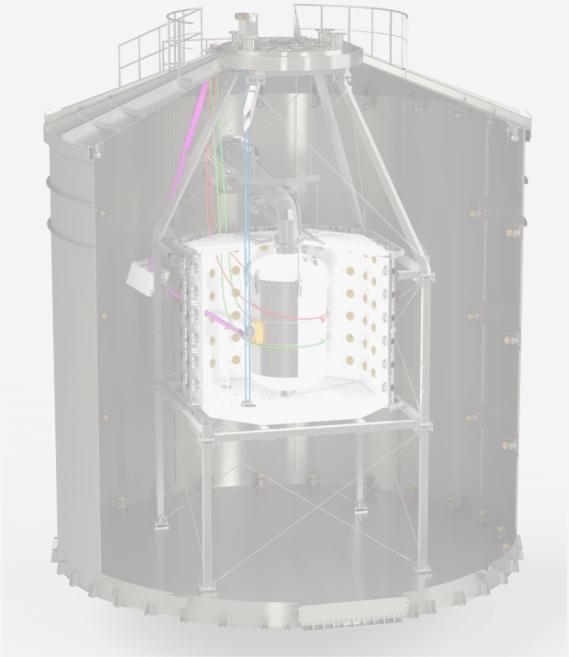
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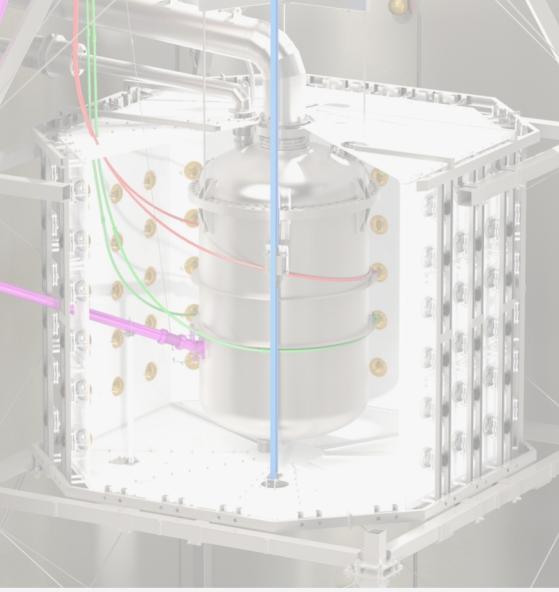
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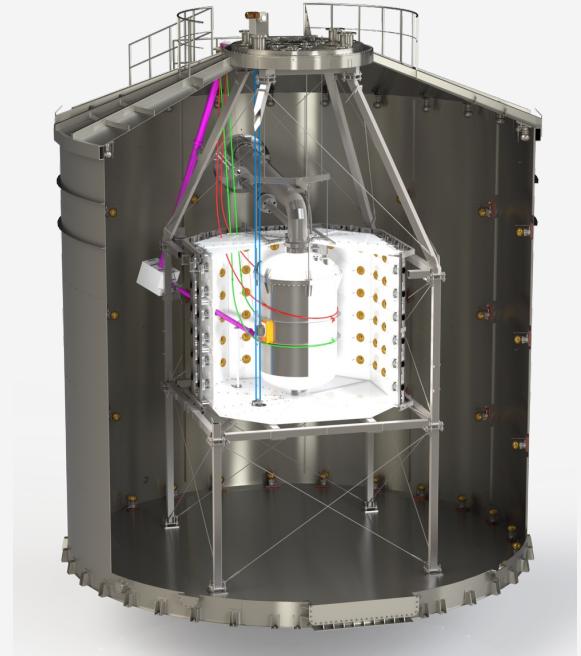
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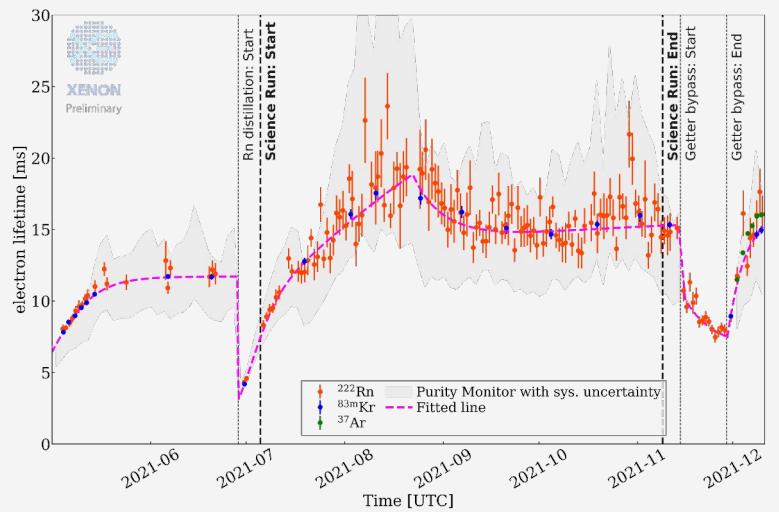
Selected Upgrades

	Max TPC drift time	Electron Lifetime	e ⁻ survival @ max. drift length
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XENON1T	0.67 ms	0.65 ms	30%
XENONnT	2.2 ms	> 15 ms	> 90%

Liquid xenon purification

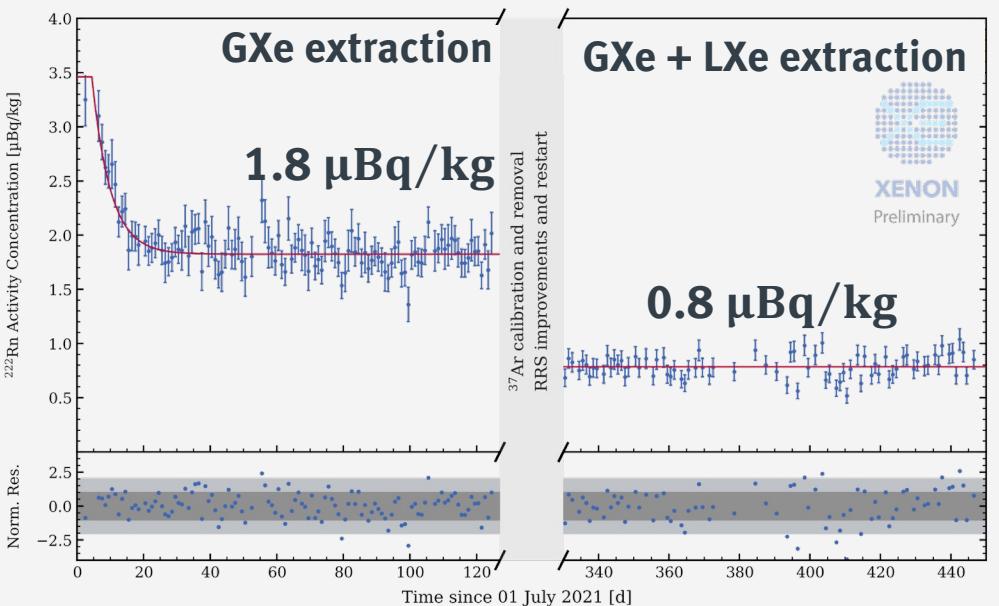
- LXe purity is crucial to drift electrons along the entire detector length
- Novel liquid-phase purification system with replaceable filters and extremely low radon emanation
- High flow of 2 liters LXe per minute, recirculate entire inventory in 18h



Selected Upgrades

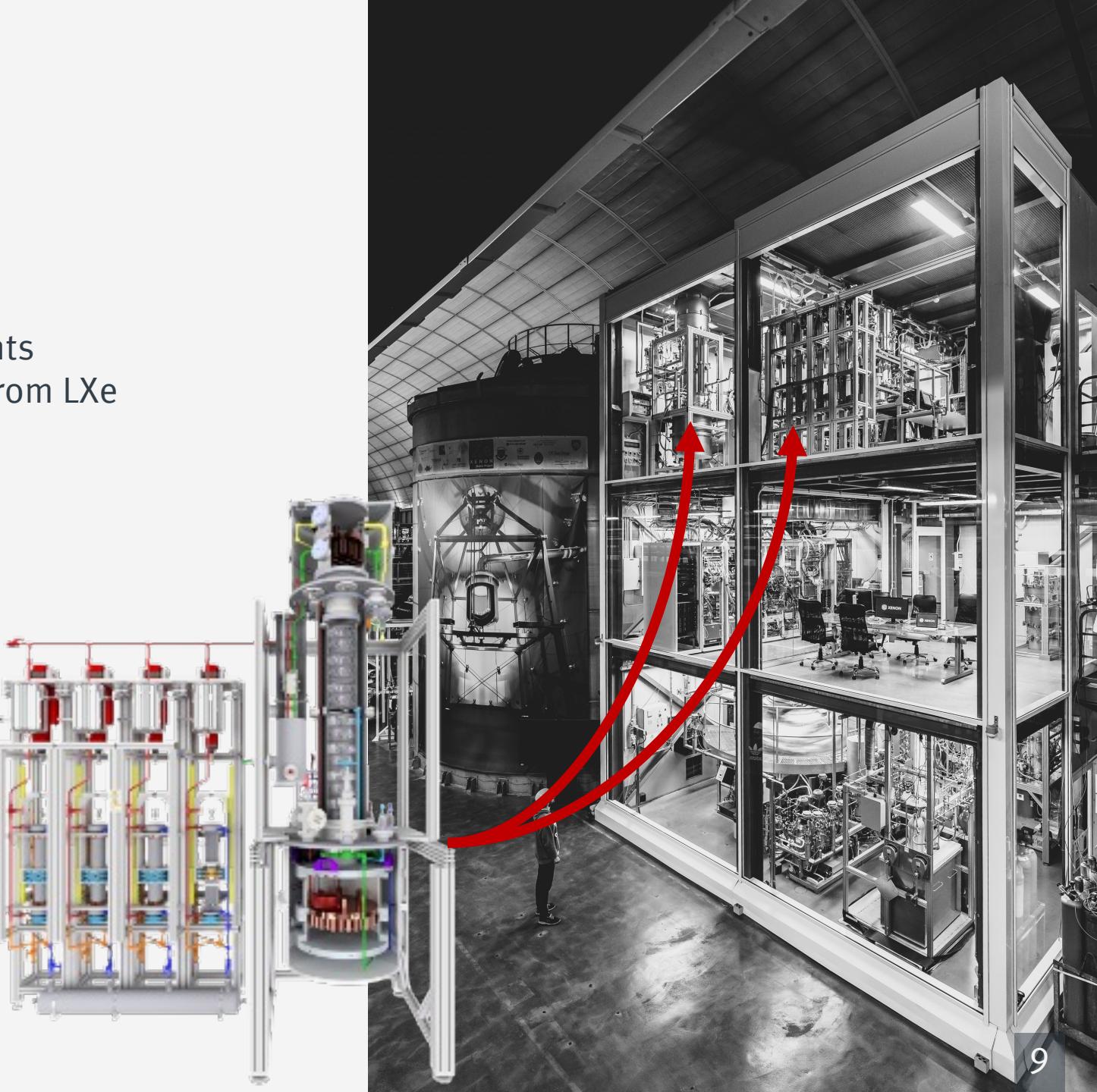
Radon Distillation Column

- ^{222}Rn daughters are the main background
- Continuous emanation from detector components
- Continuous cryogenic distillation removing Rn from LXe
- High flow: 71 kg/h (200 slpm)
- $< 1 \mu\text{Bq}/\text{kg}$ ^{222}Rn activity

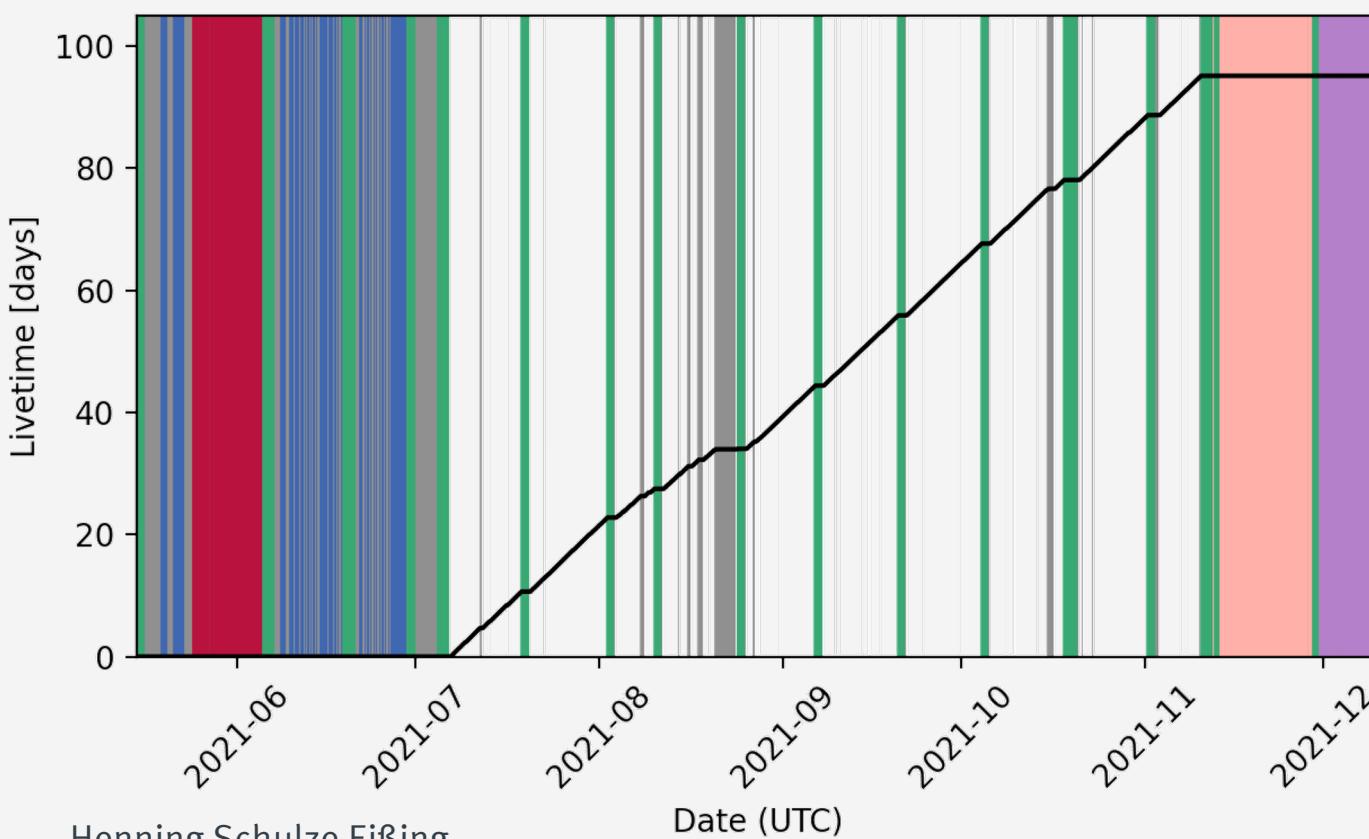


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Eur. Phys. J. C 82, 1104 (2022)
J. Inst. 17, P05037 (2022)



XENONnT SR0



Data Taking

- 97.1 days exposure from July 6th to November 11th 2021
- 95.1 days lifetime corrected
- Rn column in gas-only mode

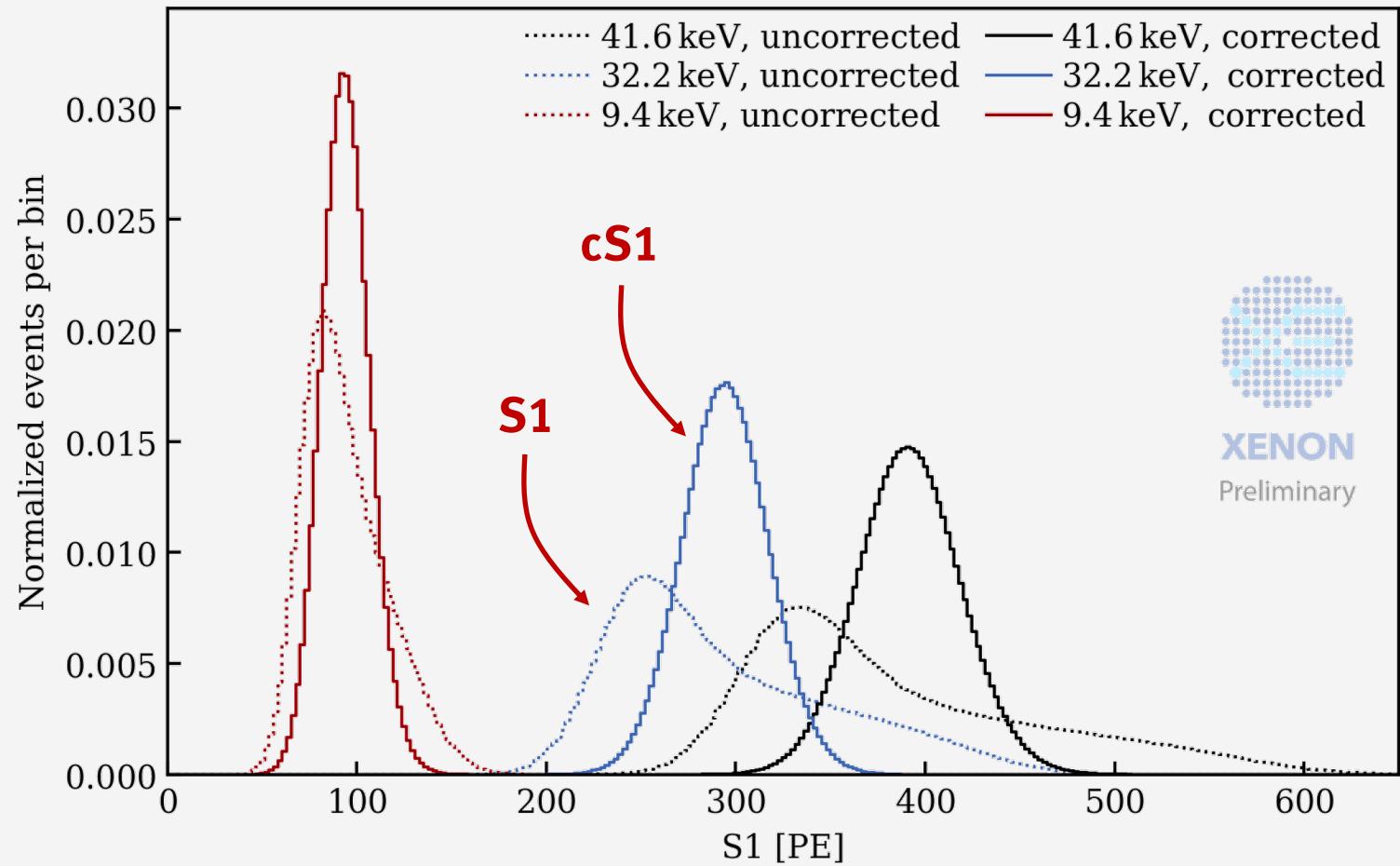
Detector Conditions

- >96% PMTs working, gain stable at 3%
- 23 V/cm drift field, extraction field in LXe 2.9 kV/cm
- Localized high single-electron emission, occurring seemingly at random: cured with short anode ramp-down

Calibration and Analysis

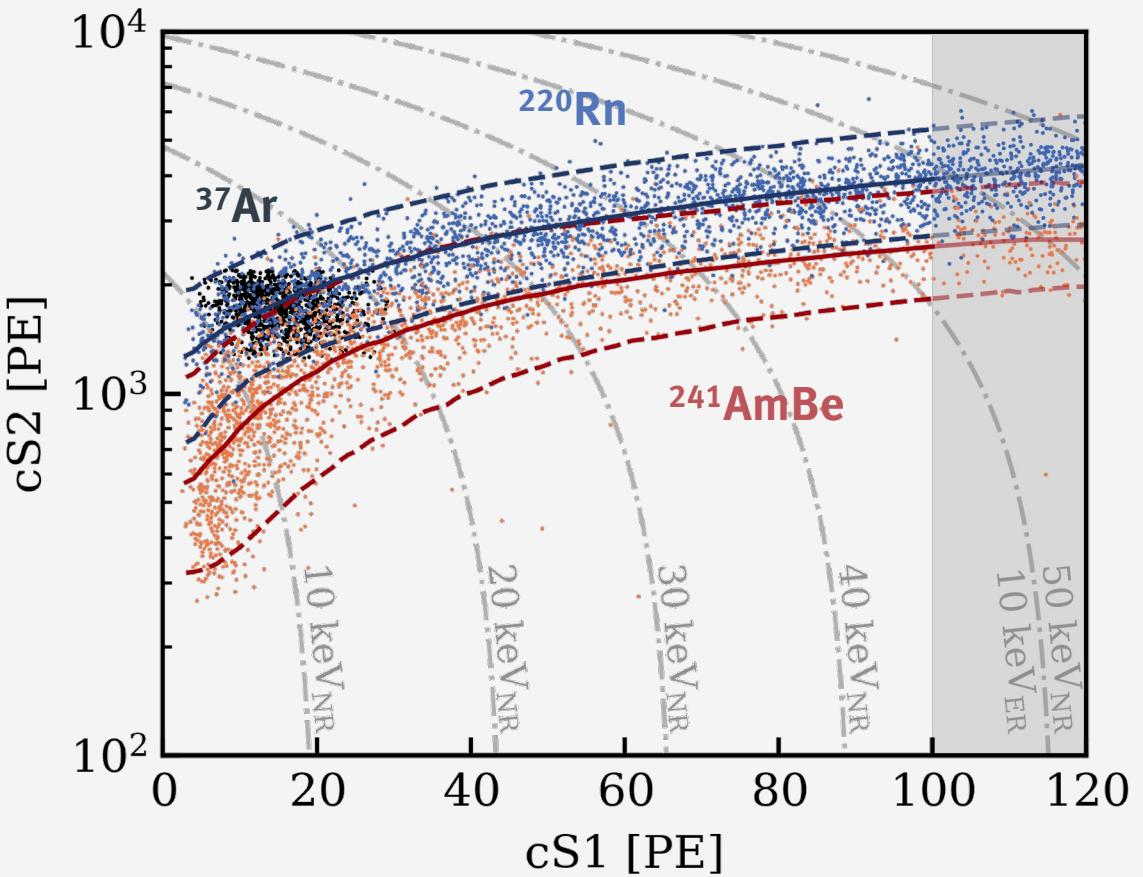
Multi-step analysis effort

- Peak and event reconstruction
- Signal Corrections – Compensate inhomogeneous (spatial, temporal) detector response
- Data quality validation
- Cuts against backgrounds
- Background modeling
- Detector response modeling
- Inference



arXiv:2406.13638

ER and NR Response

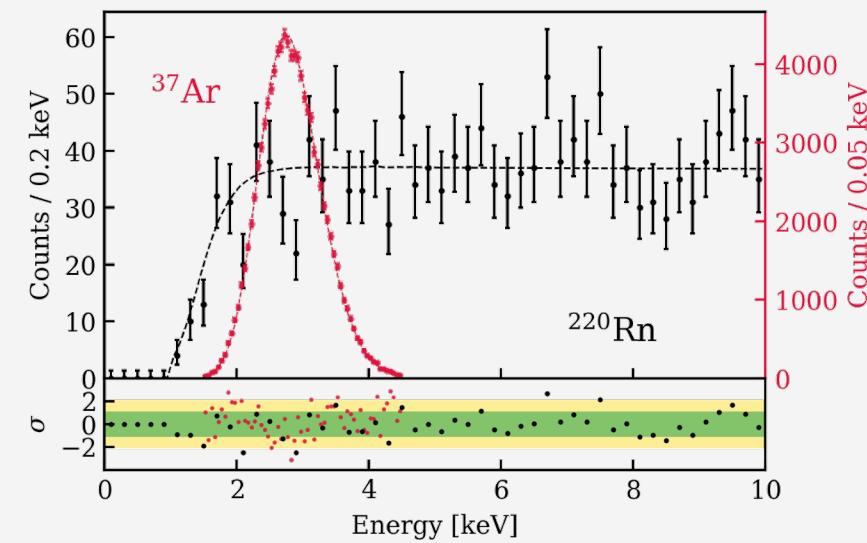


Phys. Rev. Lett. 131, 041003 (2023)

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Electronic Recoil Calibration

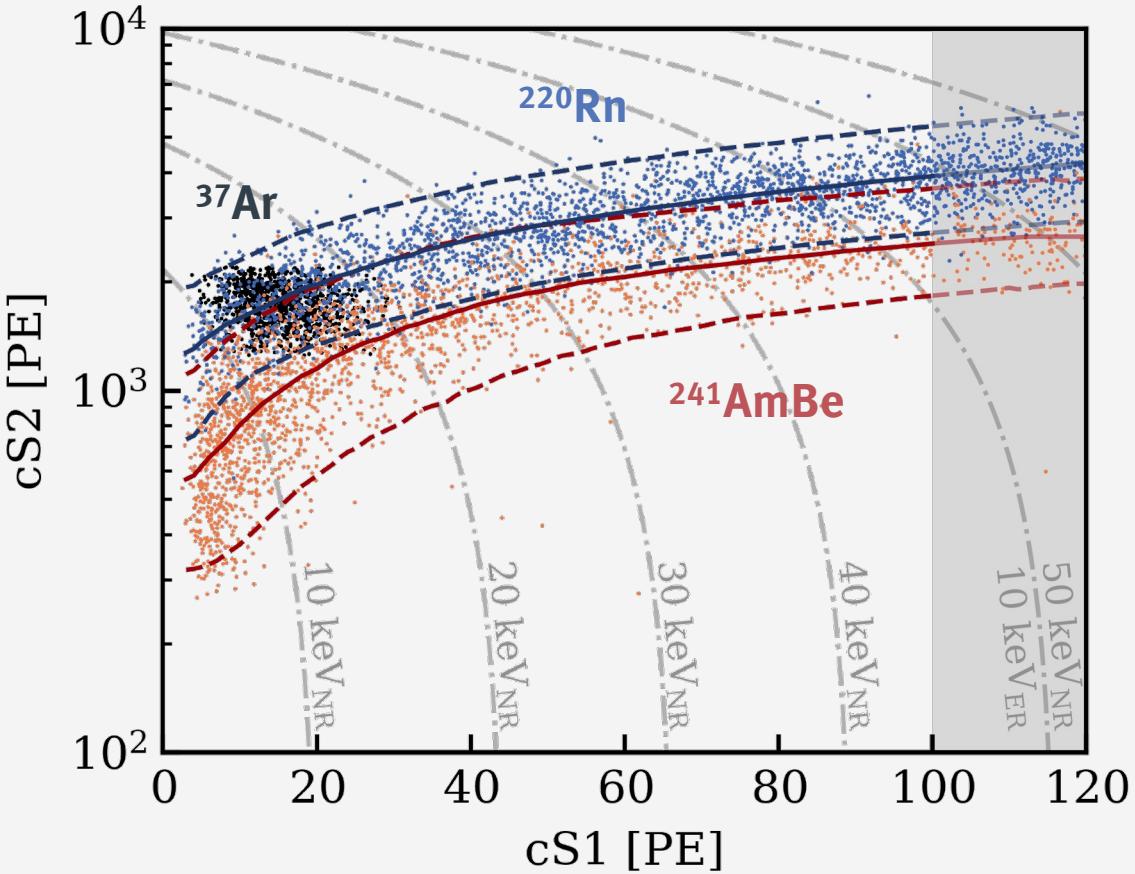
- ^{212}Pb from ^{220}Rn gives a roughly flat β -spectrum to estimate cut acceptances and energy threshold
- ^{37}Ar gives a mono-energetic 2.82 keV peak to model the low energy response and resolution near the detector energy threshold



Nuclear Recoil Calibration

- External $^{241}\text{AmBe}$ neutron source with clean NR selection via coincident 4.4 MeV γ -ray observed in the neutron veto

Physics Searches



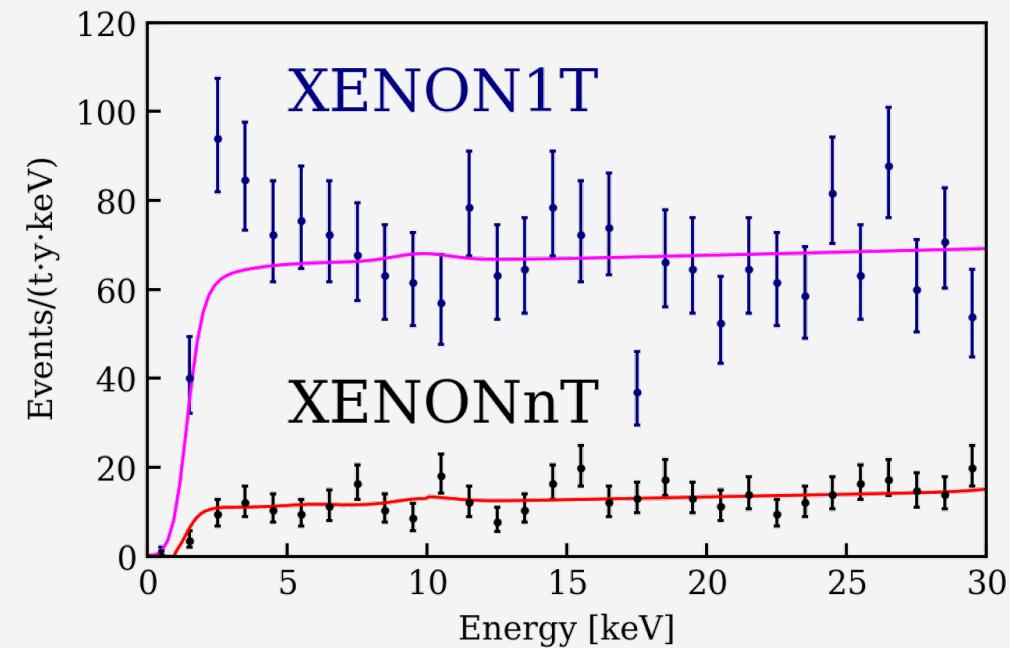
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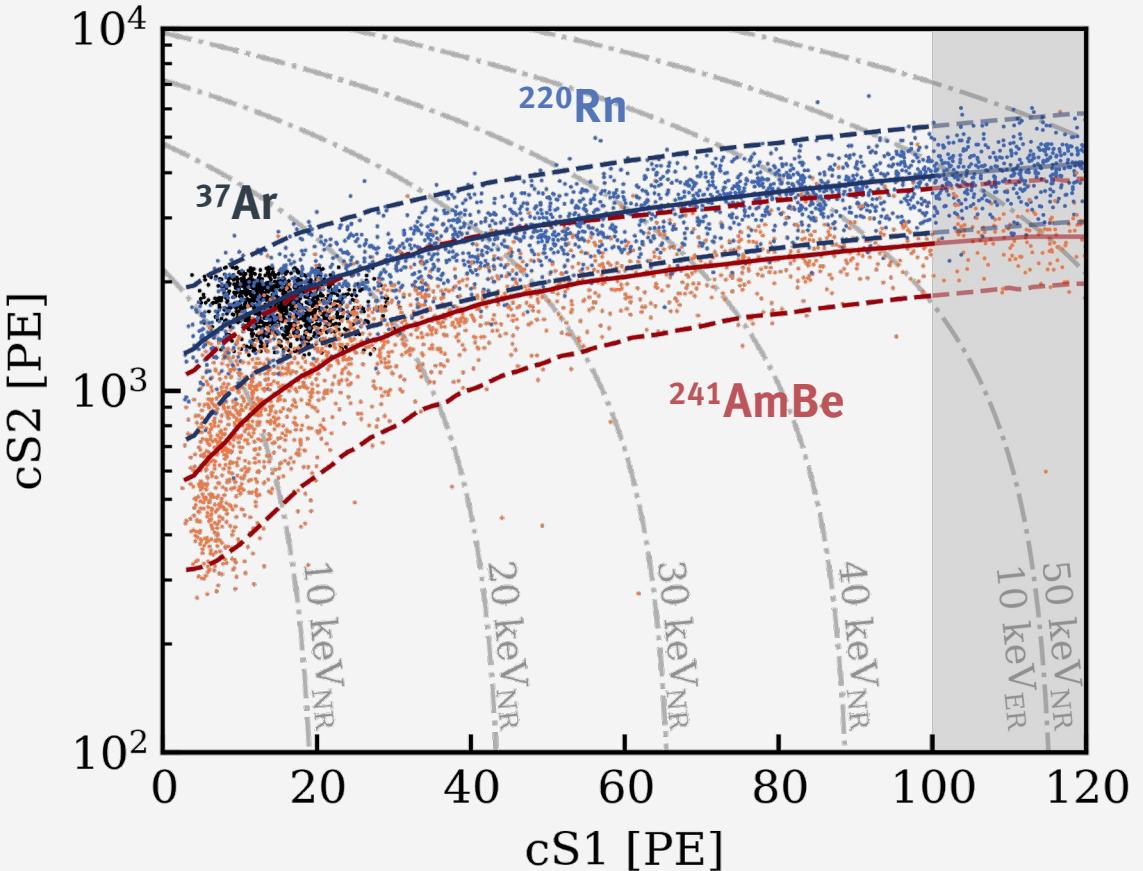
Electronic Recoils

- Search in energy reconstructed from S1 and S2
- Unprecedented low background: 15.8 ± 1.3 events / $(t \cdot y \cdot \text{keV})$
- No low-energy ER excess found excluding BSM explanation of XENON1T excess

Phys. Rev. Lett. 129, 161805 (2022)



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Phys. Rev. Lett. 129, 161805 (2022)

Nuclear Recoils

- Search in cS1, cS2 and r (radius) for WIMP dark matter
- Even lower NR background boosted by ER/NR discrimination
- Blind analysis leading to stringent limits on WIMP interactions

Phys. Rev. Lett. 131, 041003 (2023)

Backgrounds in WIMP search

ER Background

- Dominated by radon background (^{214}Pb beta decay, GXe extraction only)
- Sub-dominant ^{85}Kr background (cryo. Distillation)

Surface Background

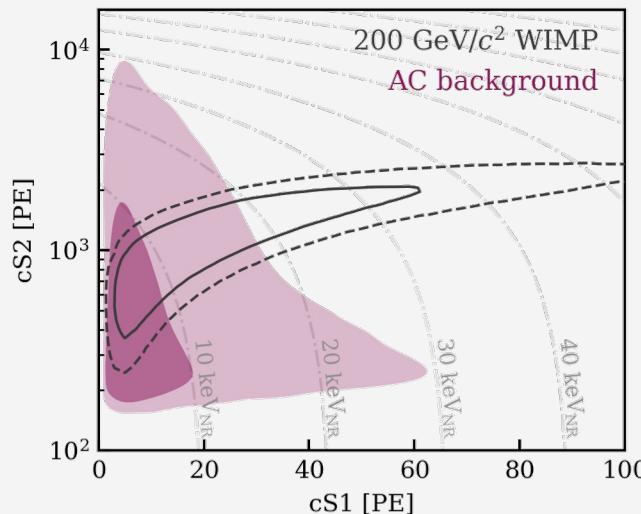
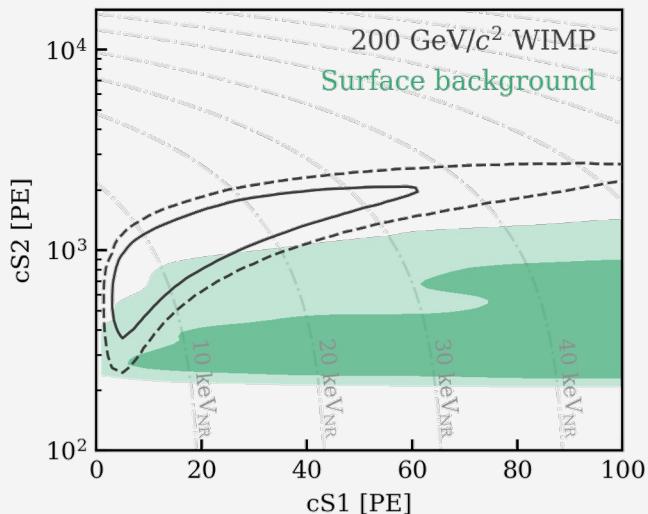
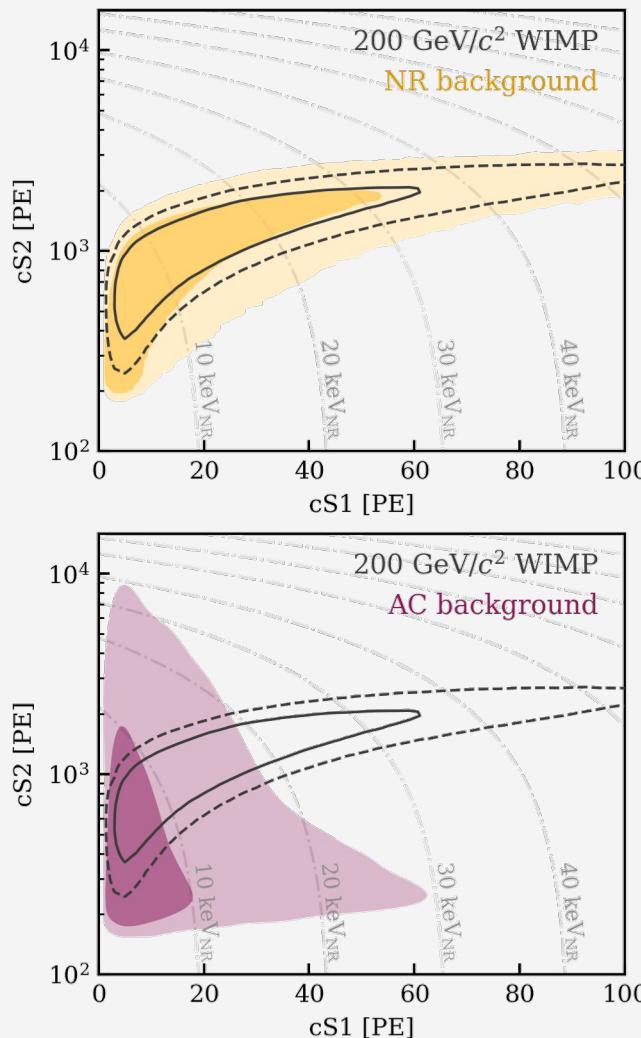
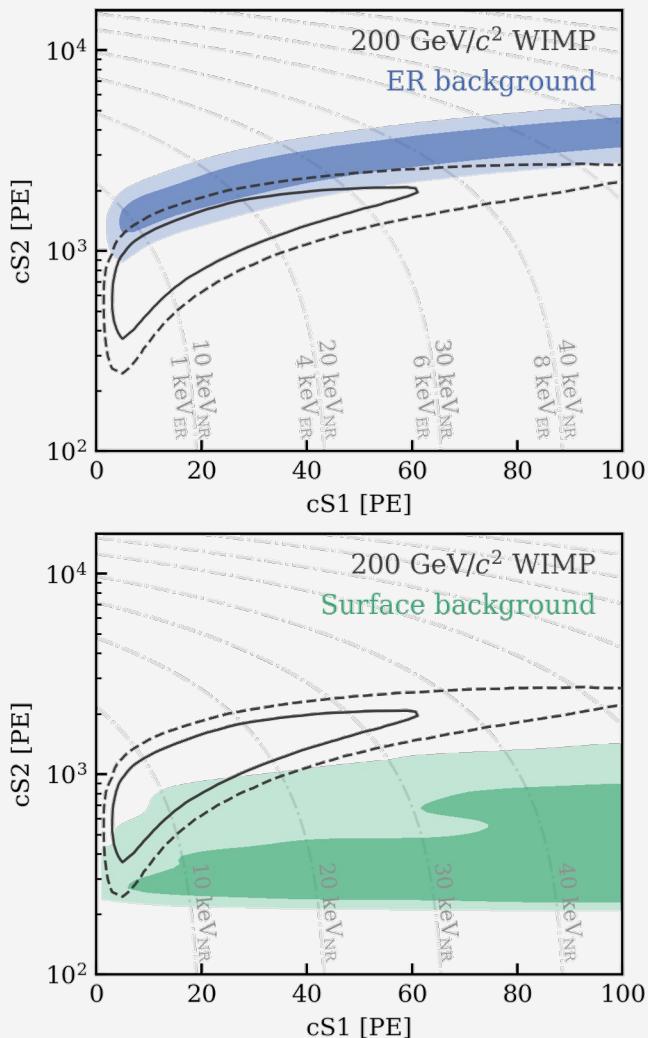
- ^{210}Pb plate-out at the PTFE walls leading to ^{210}Po α -decays with electron loss
- Suppressed by fiducial volume cut

Accidental Coincidences

- Random pairing of isolated S1 and S2 signals
- Suppression using a gradient BDT cut based on S2 shape, R and Z information

NR Background

- Radiogenic neutron rate constrained by NV tagging
- CEvNS constrained from solar ^8B neutrino flux

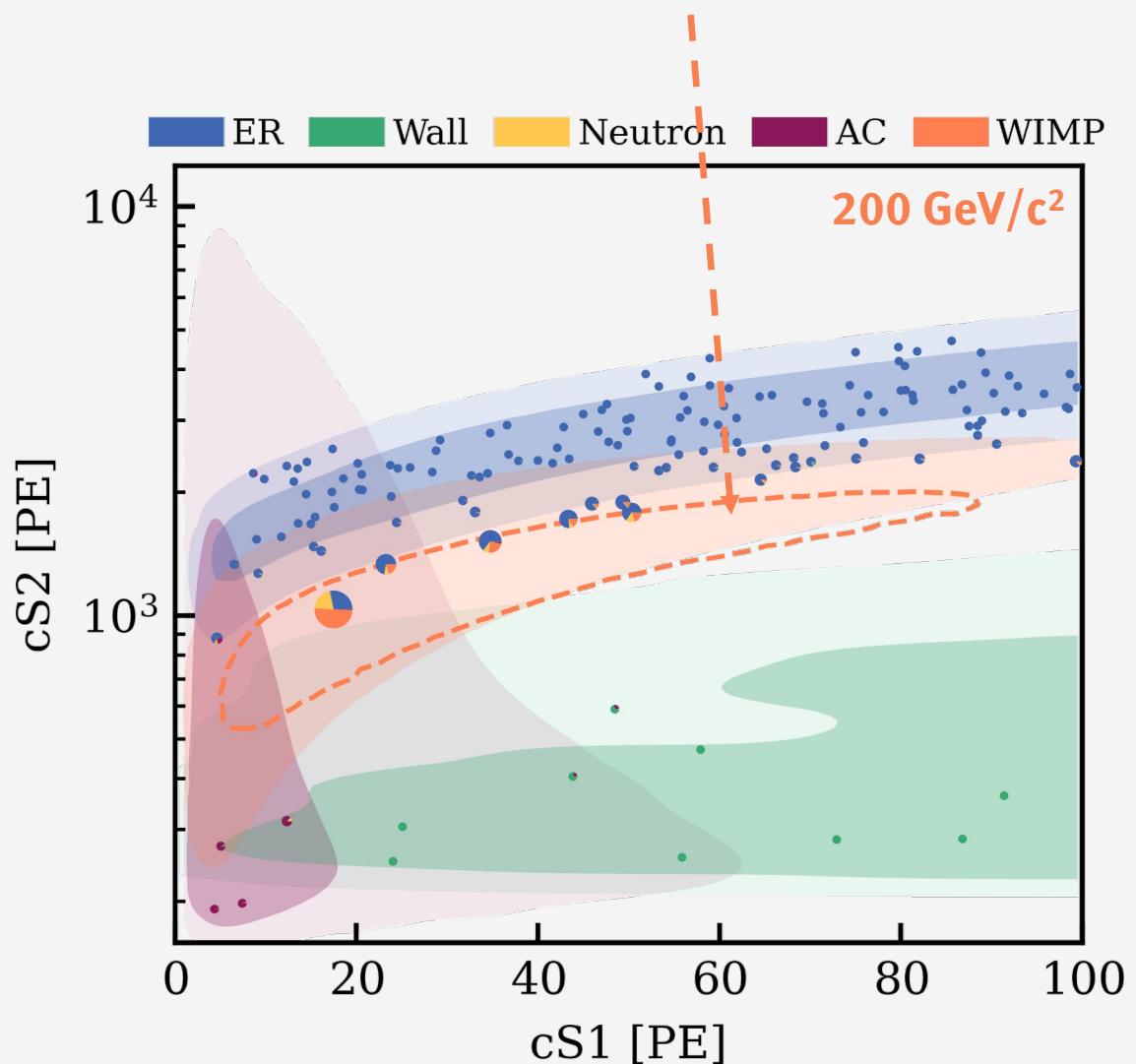


WIMP search results

	Nominal	Best Fit (200 GeV/c ²)
ER	134	135(+12)(-11)
Neutrons	1.1(+0.6)(-0.5)	1.1 ± 0.4
CEvNS	0.23 ± 0.06	0.23 ± 0.06
AC	4.3 ± 0.2	4.32 ± 0.16
Surface	14 ± 3	12(+0)(-4)
Total Background	154	152 ± 12
WIMP	-	2.6
Observed	-	152

No significant excess found!

Signal like region, containing 50% of WIMP signal with highest signal-to-noise



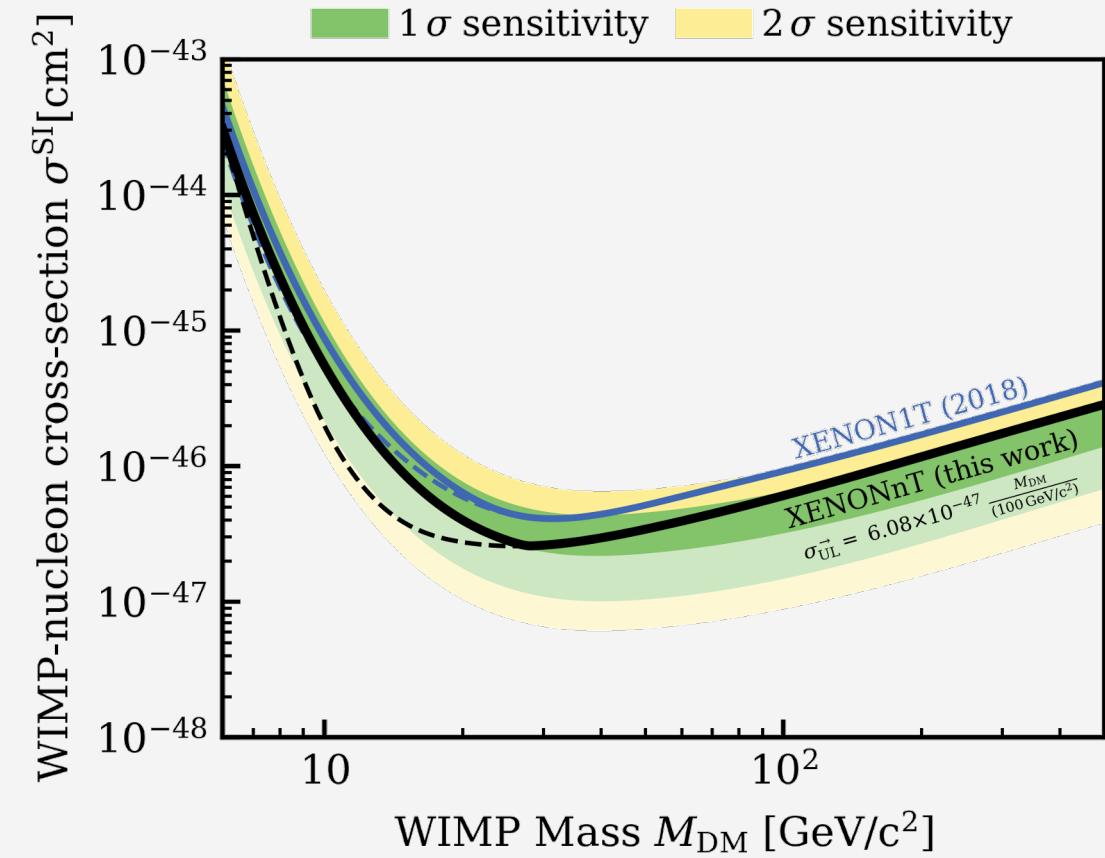
Limit on SI WIMP-nucleon Cross Section

Limit Setting

- Unbinned maximum likelihood
- Power constraint limits (PCL) to avoid spurious exclusion limits
- Only exclude the parameter space that the detector is sensitive to.
- Conservative choice at median of sensitivity band
- This needs to be discussed within the community

Results

- Strongest limit: $2.6 \cdot 10^{-47} \text{ cm}^2$ at WIMP mass of $28 \text{ GeV}/c^2$
- Factor 1.6 improvement w.r.t. XENON1T (with considerably shorter lifetime)
- Similar improvements in spin-dependent limits



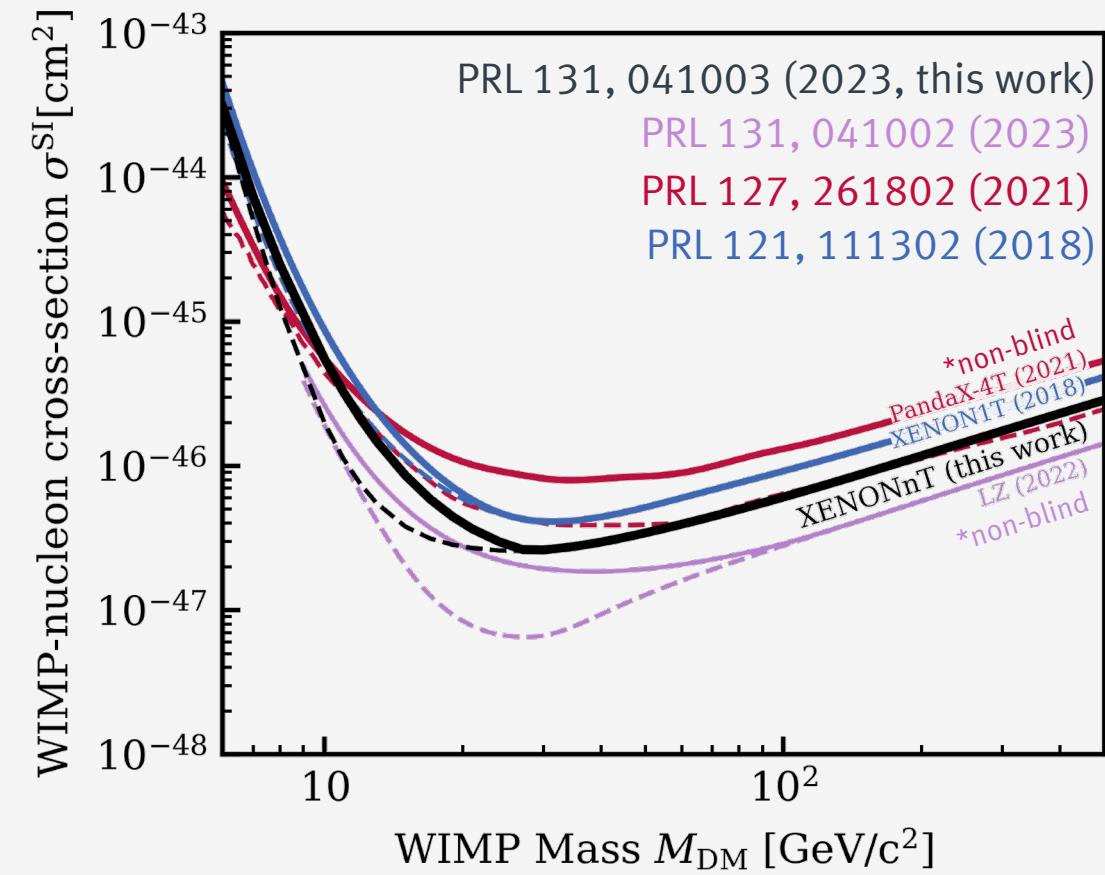
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Summary and Outlook

XENONnT SR0

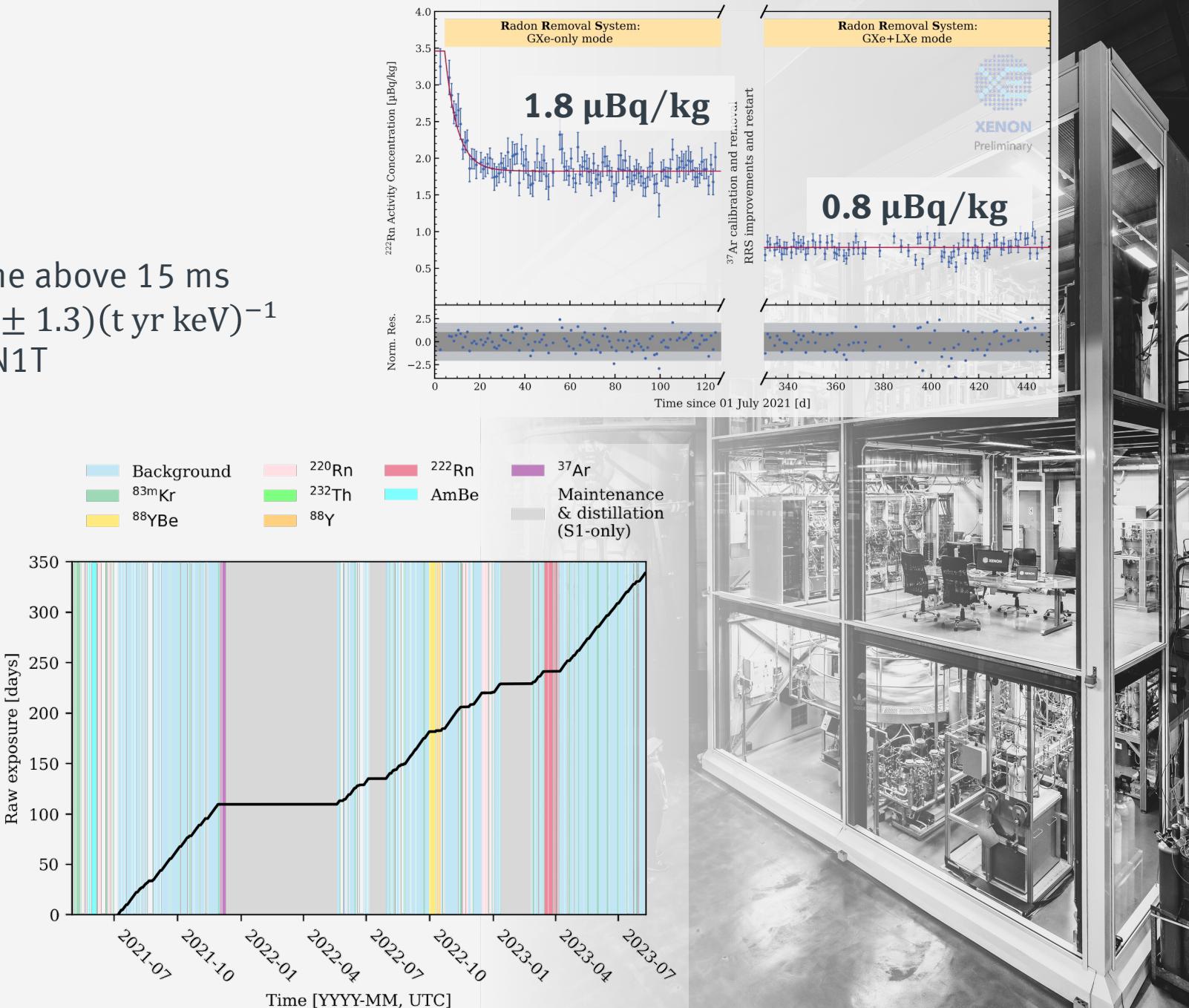
- Ultrapure target with an electron lifetime above 15 ms
- Lowest ER background in the field: $(15 \pm 1.3)(t \text{ yr keV})^{-1}$
~5x background reduction w.r.t. XENON1T

First Results

- Blinded electronic recoil (ER) and nuclear recoil (NR) searches
- No significant excess over background found

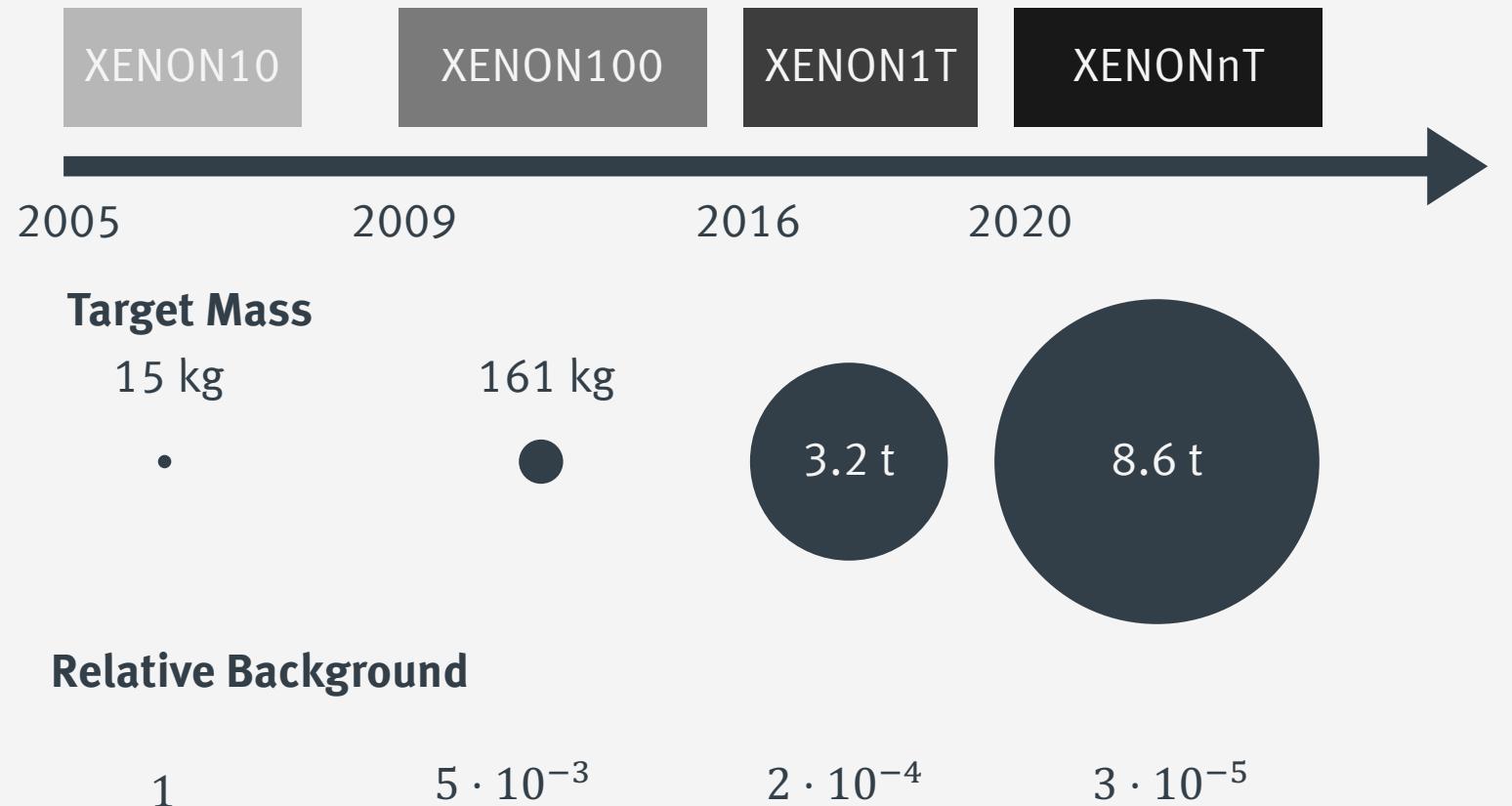
Prospects

- ~2x improved ^{222}Rn level
- Improved neutron tagging by Gd-loaded neutron veto
- WIMP search with increased exposure in preparation
- Continue to accumulate low BG data



Backup Slides

XENON Dark Matter Project



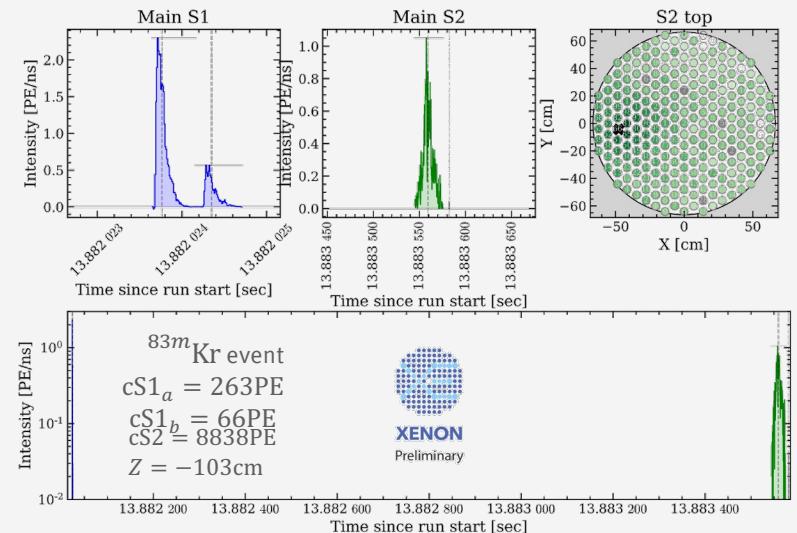
Selected Upgrades

Triggerless Data Acquisition

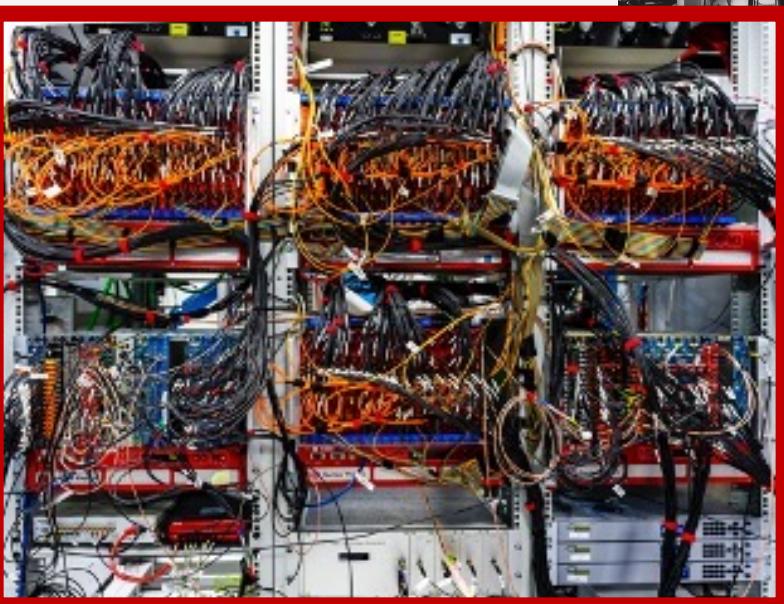
- All data above per-channel threshold stored long-term
- Fully live processing

Open Source Software

- Available on GitHub
- Processing: strax, straxen
- Simulation: fuse, appletree, WFSim, epix
- Inference: alea



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222Rn Background Reduction

Radon in XENONnT

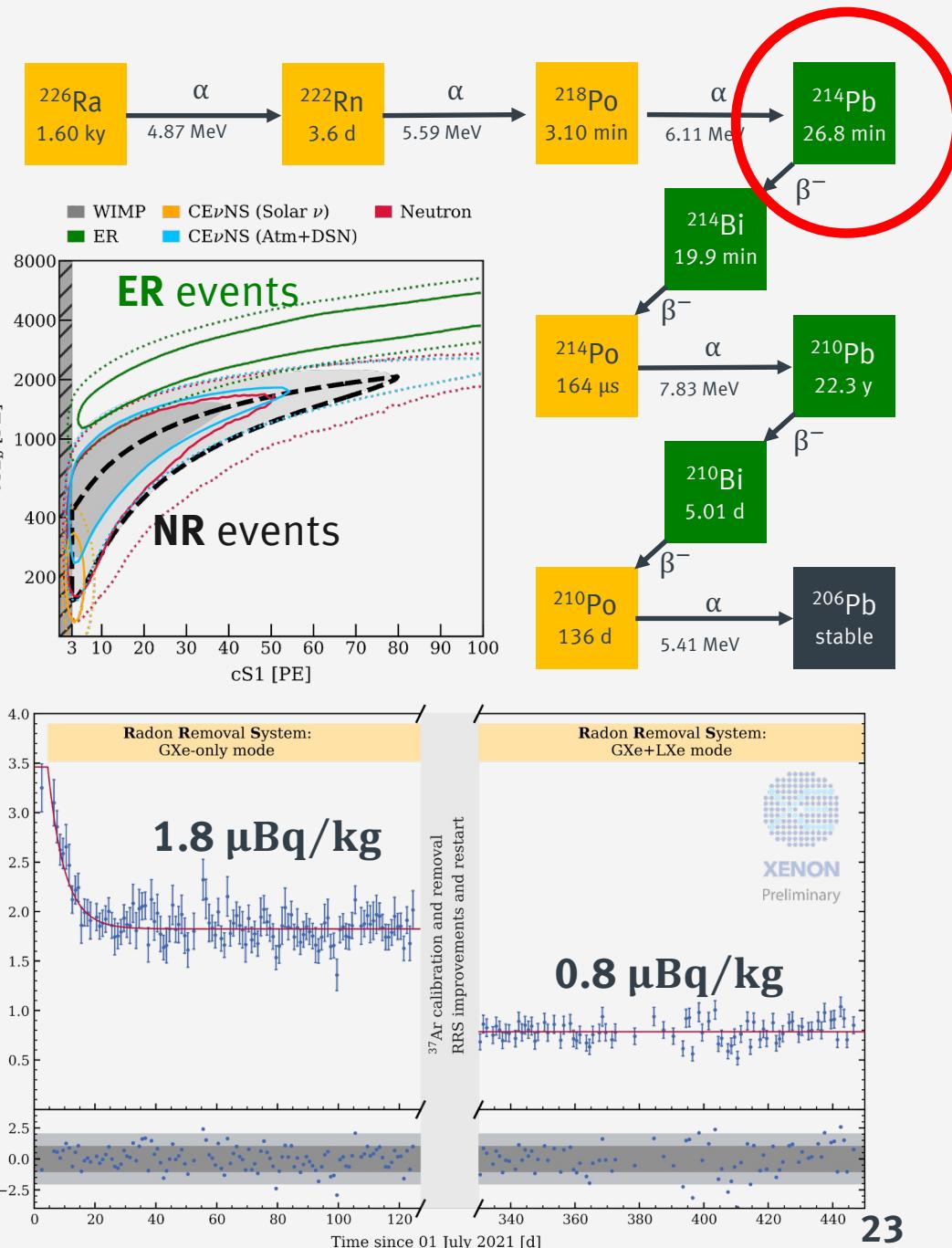
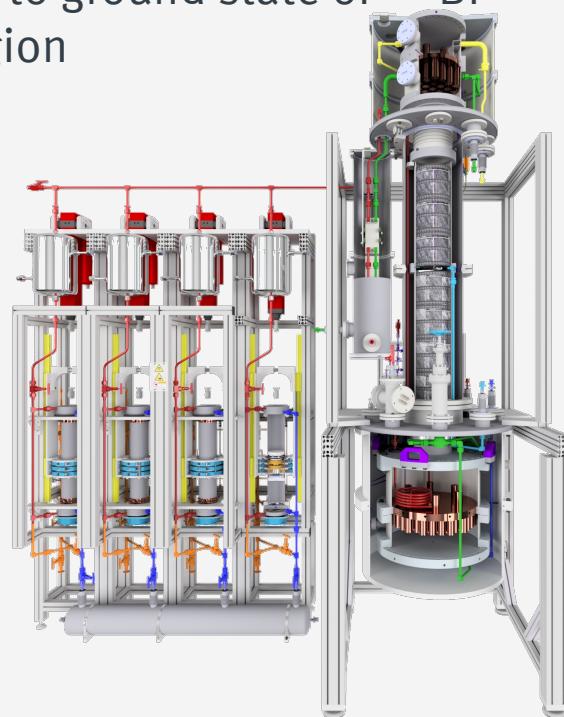
- 222Rn daughters are the main background
- Continuous emanation from detector components
- Homogeneous distribution in detector

214Pb

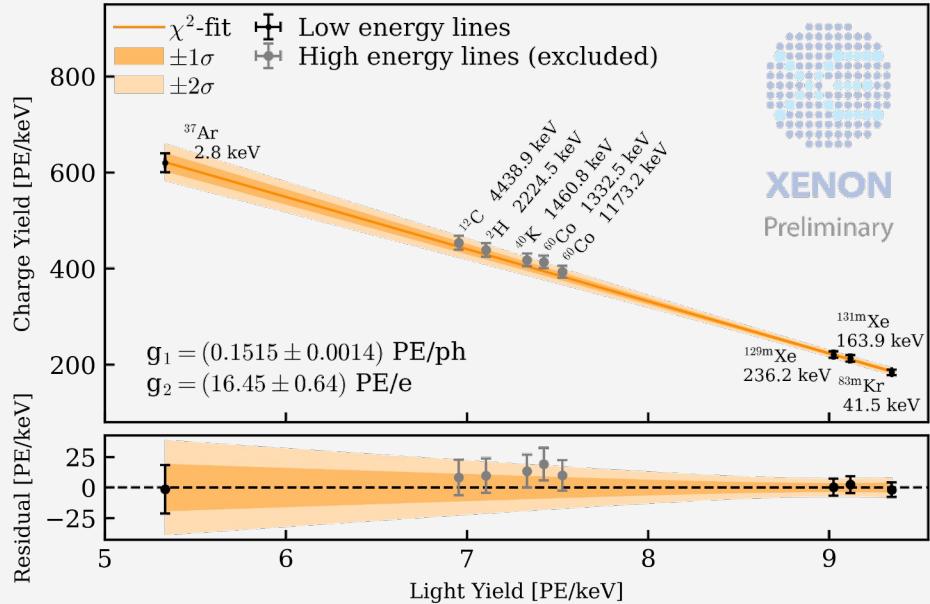
- Beta decay with high probability to ground state of ^{214}Bi
- Leakage of ER events into NR region

Radon reduction

- Cryogenic distillation removing Rn from LXe
- Newly developed Distillation column
- High flow: 71 kg/h (200 slpm)
- $< 1 \mu\text{Bq}/\text{kg}$ 222Rn activity



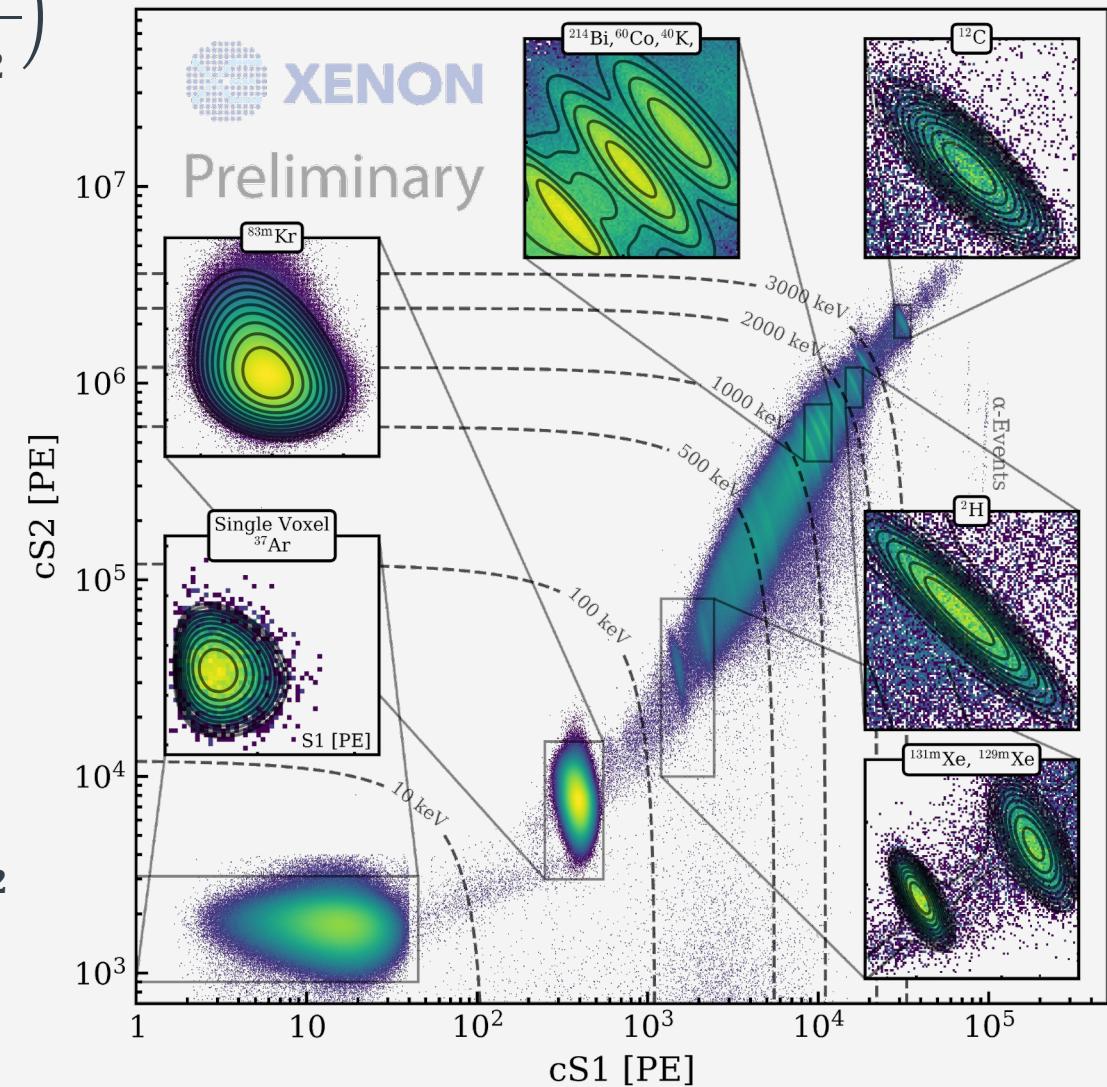
Energy Reconstruction



$$E = W \cdot \left(\frac{cS1}{g_1} + \frac{cS2}{g_2} \right)$$

Combined Energy Scale

- Use anti-correlation of charge and light signals
- Photon detection efficiency g_1 and charge amplification g_2
- Low energy calibration sources
 ^{37}Ar , $^{83\text{m}}\text{Kr}$, $^{129\text{m}}\text{Xe}$ and $^{131\text{m}}\text{Xe}$
- Reconstruction bias correction of O(1-2)%
- Calibration from 2.8 keV up to MeV region



Detection and Selection Efficiencies

Detection efficiency

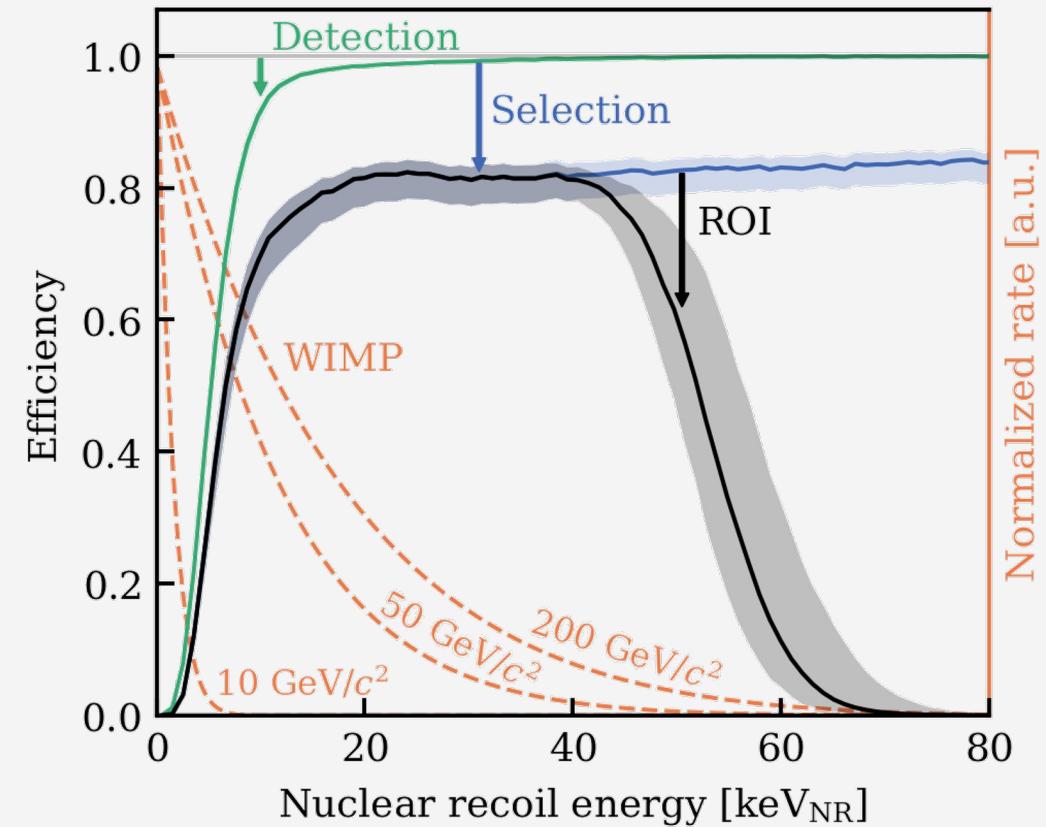
- Driven by 3-fold PMT coincidence requirement to identify peaks as S1
- Data- and simulations driven with good agreement

Selection efficiency (~80% acceptance)

- Data quality selections to remove unphysical events
- S1 and S2 peaks must have PMT hit patterns, top/bottom area ratios, etc. consistent with real events
- S2 width consistent with the expected diffusion along the drift path
- Single-site events

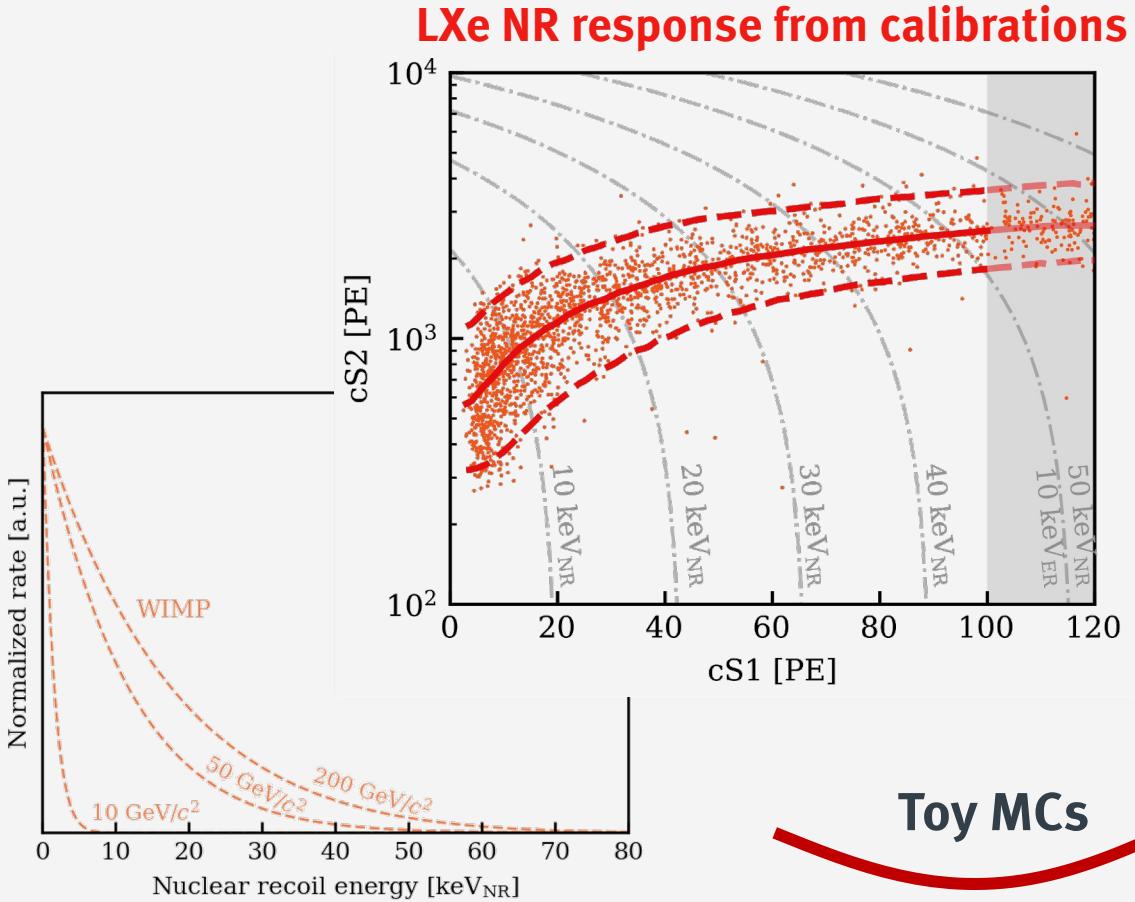
Region of interest (ROI)

- Constructed to fully contain the WIMP recoil spectra
- cS1: 0-100 PE | cS2: $10^{2.1} - 10^{4.1}$ PE



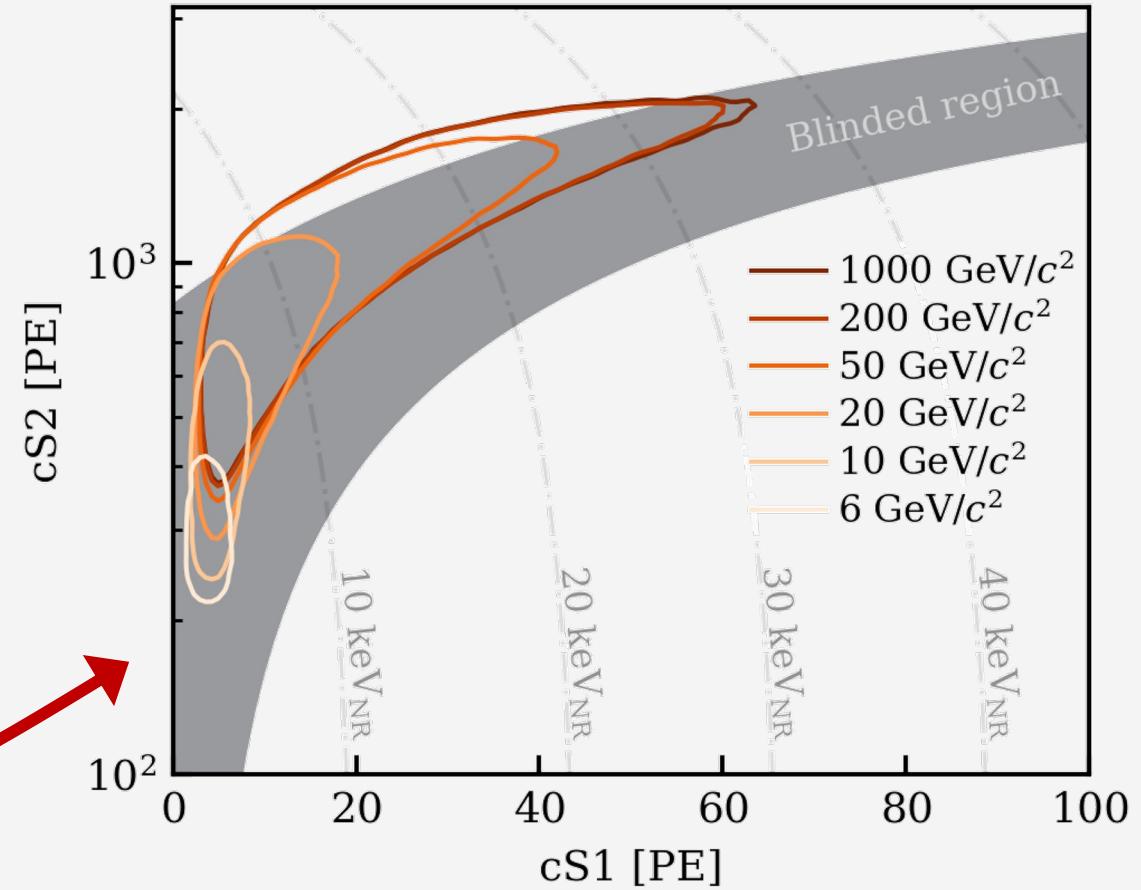
Phys. Rev. Lett. 131, 041003 (2023)

WIMP Signal Model



Theoretical WIMP NR energy spectra

Henning Schulze Eißing



arXiv:2406.13638

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