

# Status of the Quantum Sensors for the Hidden Sector (QSHS) Experiment

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Quantum Sensors for the Hidden Sector Collaboration





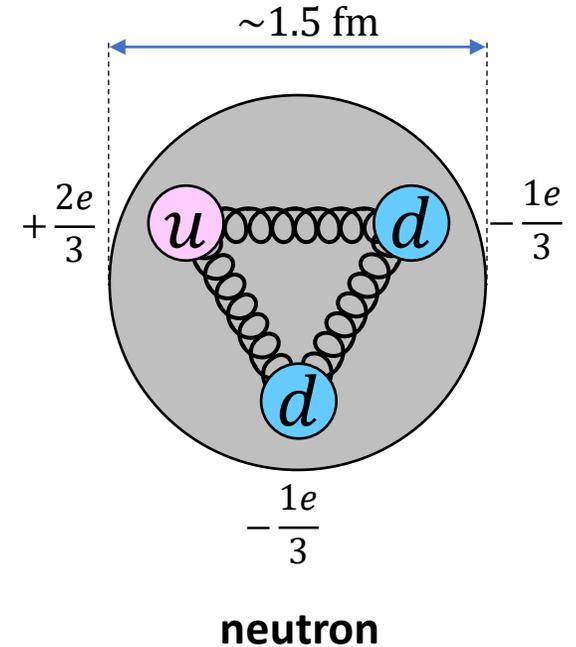
# The Strong CP Problem

Conservation of Charge-Parity (CP) in the strong interaction is unexpected.

$$\mathcal{L}_{CPV(QCD)} = \frac{(\Phi + \arg \det |M|)}{32\pi^2} \vec{E}_{QCD} \cdot \vec{B}_{QCD} \approx \bar{\theta} \cdot (10^{-16} e \text{ cm})_{\text{neutron}}$$

A well-known example of conservation of strong CP is the neutron Electric Dipole Moment (EDM):

- The current experimental limits on  $|d|$  give  $10^{-26} e \text{ cm}$  therefore  $\bar{\theta} < 10^{-10}$
- Either it is coincidentally small, or the CP violating effects must add up to zero, driving the minimisation of  $\bar{\theta} = (\Phi + \arg \det |M|)$





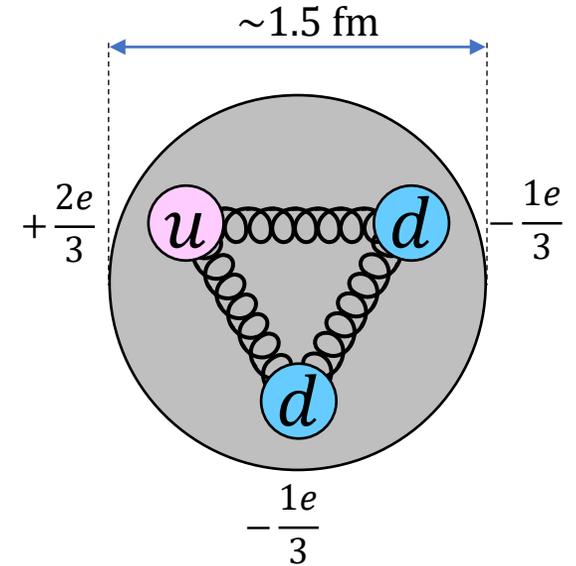
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- The Peccei-Quinn (PQ) mechanism adds an additional QCD Lagrangian term that turns  $\bar{\theta}$  into a dynamic variable.
- This naturally minimises  $\bar{\theta}$  and solves the ‘Strong CP’ problem.
- The PQ mechanism produces Axions – are these DM candidates?



**neutron**

$$m_a f_a \sim m_\pi f_\pi$$

$$m_a = 5.70 \mu\text{eV} \left( \frac{10^{12} \text{ GeV}}{f_a} \right)$$



# Why is Axion DM Wave Like

If the local halo density  $\rho_H = \sim 0.45 \text{ GeV cm}^{-3}$

And the virial velocity in the local halo is  $230 \text{ kms}^{-1}$

Assume: 1) Axions have a mass that provides a good fraction of the closure density, i.e. all DM is axions

2)  $m_a c^2 = 4 \mu\text{eV}$

- High number density of  $10^{14}$  axions per  $\text{cm}^3$
- Long Wavelength - This gives us de-Broglie wavelength of about 400 m.

Axion DM can be described to a very good approximation by a classical pseudoscalar field in the same way that photons can be approximated as an EM field.



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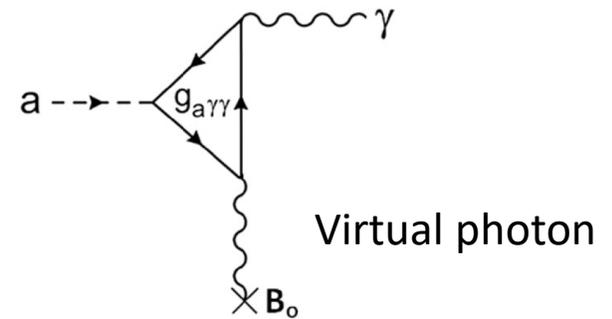
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- This could be a useful characterisation when it comes to their detection.
- Method of detection is the reverse Primakoff effect.
- Axions are very light, may be detectable by conversion to RF/UHF photons in a cryogenically cooled resonant cavity.
- **Photon signal at the yocto-watt ( $10^{-24} \text{ W}$ ) level**

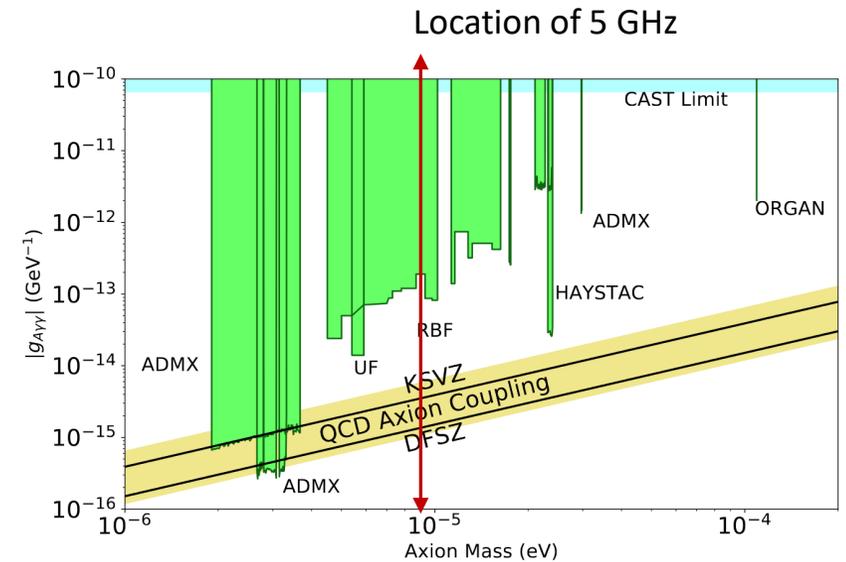
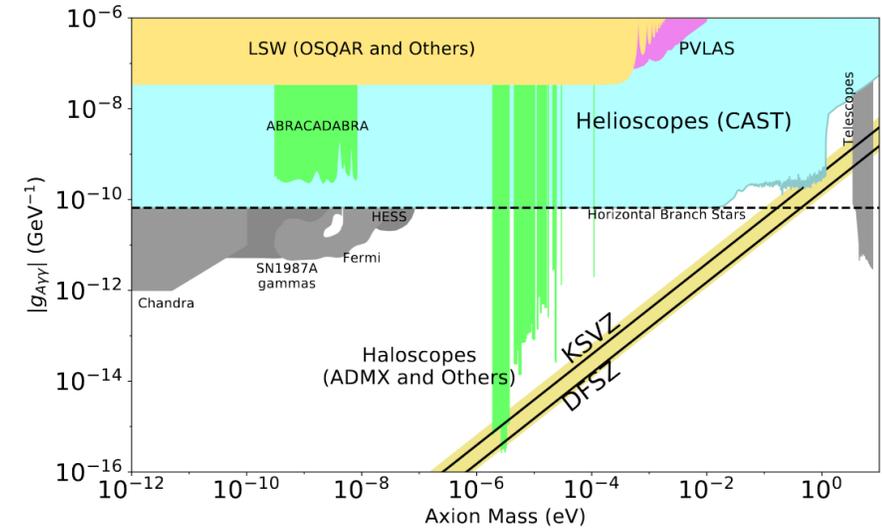


Inverse Primakoff effect in a static Magnetic field  $g_{a\gamma\gamma} \sim 10^{-15} \text{ GeV}^{-1}$



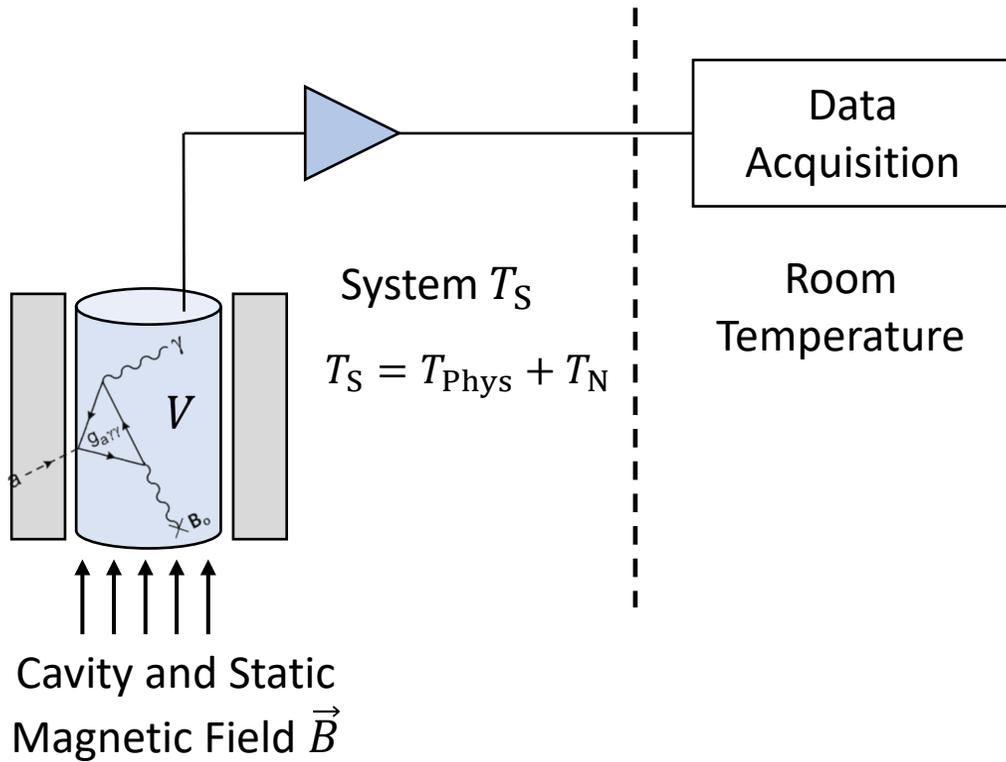
# Axion Resonant Detectors

<https://pdg.lbl.gov/2020/review/s/rpp2020-rev-axions.pdf>

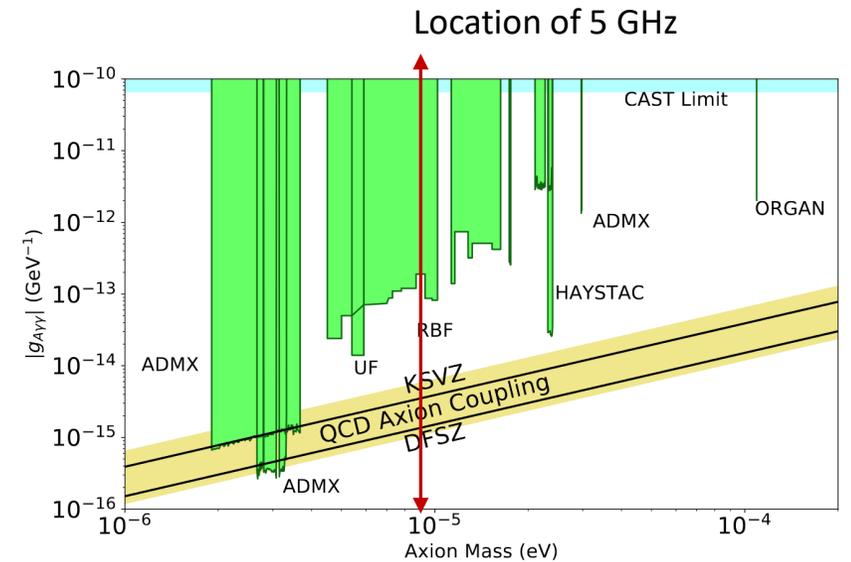
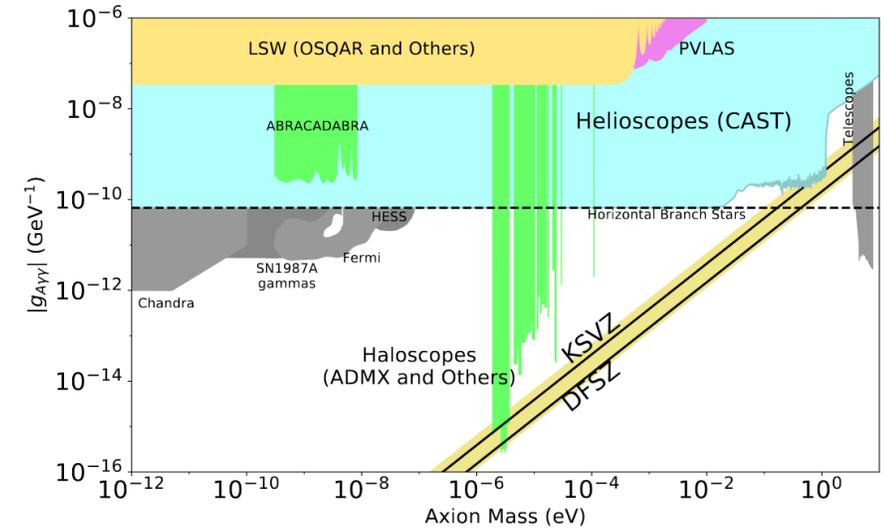




# Axion Resonant Detectors



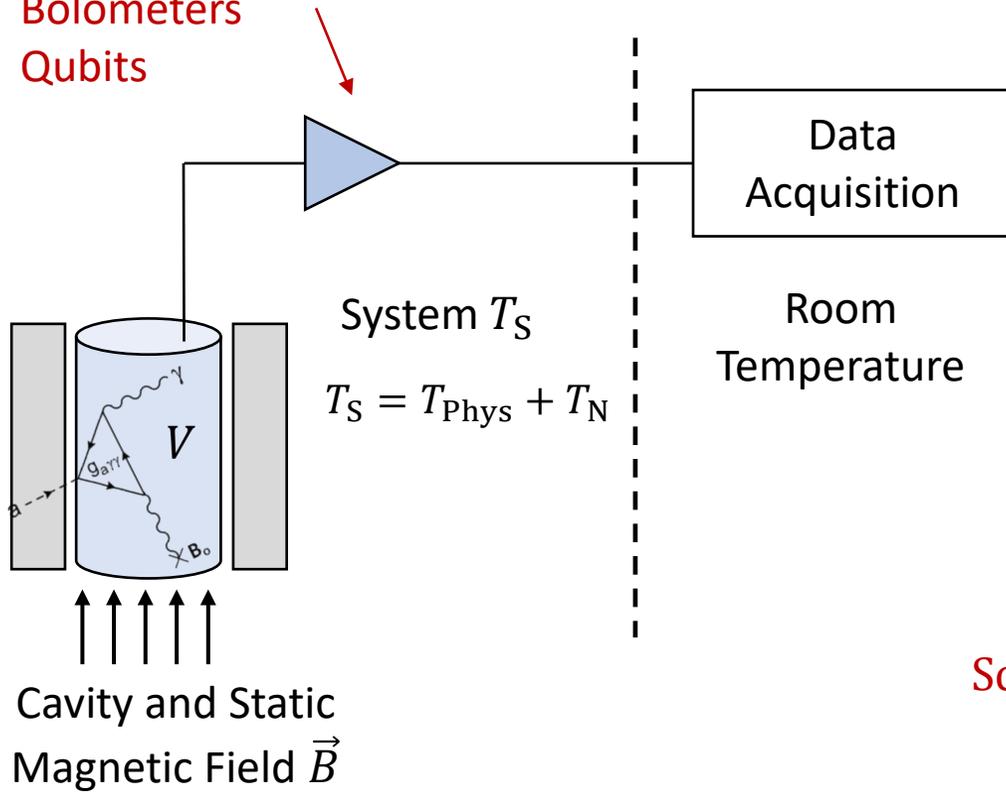
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# Axion Resonant Detectors

- SQUIDS
- Josephson Parametric Amplifiers & Travelling Wave Parametric Amplifiers
- Bolometers
- Qubits

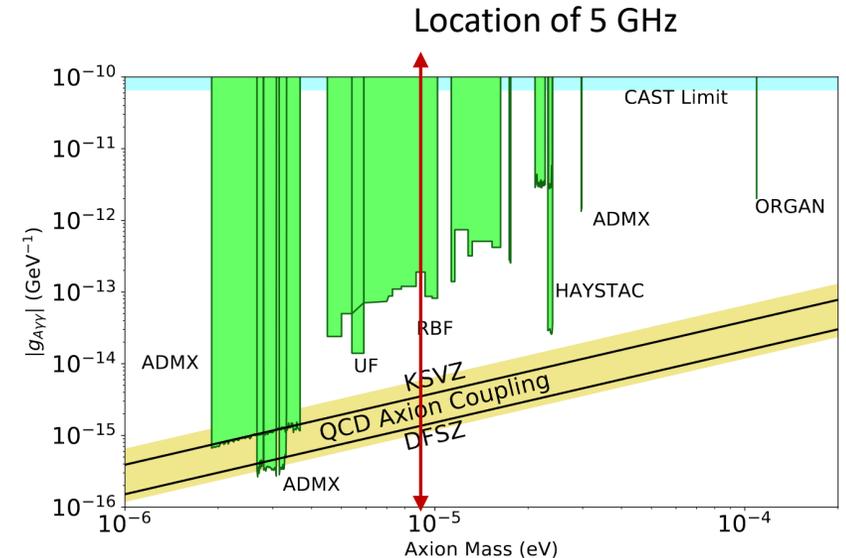
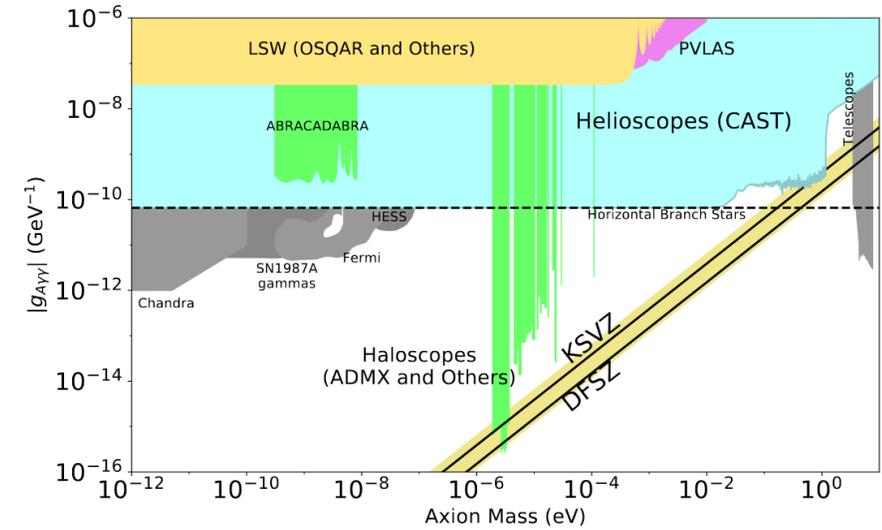


$$\text{SNR} \propto \frac{B^2 V}{T_S}$$

$$\text{Scan Rate} \propto \frac{B^4 V^2 Q_L}{T_S^2}$$

- The huge improvement in SNR and scan-rate over the last few decades is due to the improvement in system noise  $T_S$ .
- Significant further improvements in  $T_S$  needs to be made to help create a tractable scan through the frequencies available to resonant detectors.

<https://pdg.lbl.gov/2020/review/s/rpp2020-rev-axions.pdf>



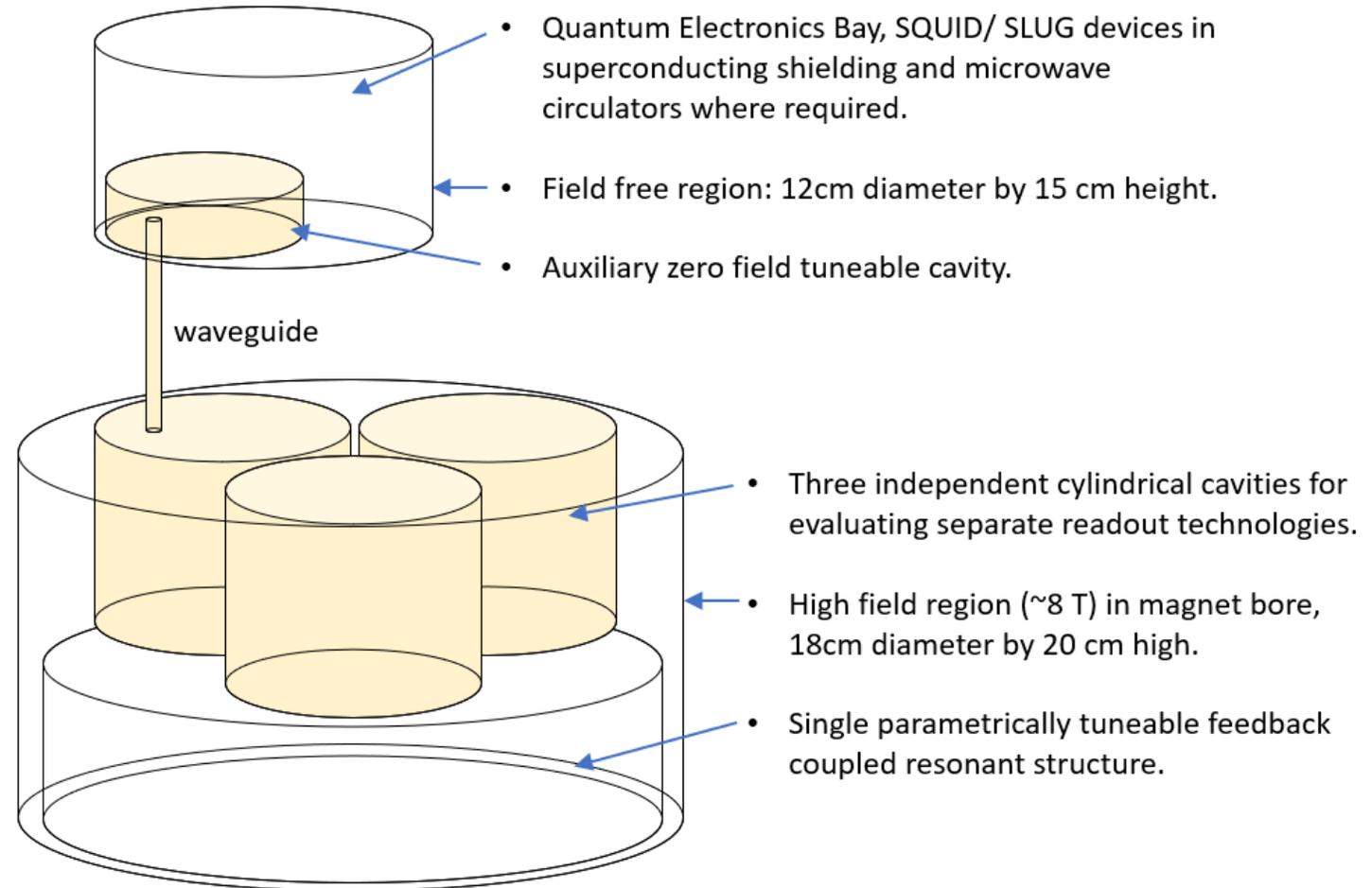


# QSHS Test Facility – Science

## Science Goals

- Test of ultra-low-noise electronics developed by the collaboration
- Tests of tuneable resonator hardware
- Science from a search of QCD axions ( $\sim 5 \text{ GHz} \approx 20 \text{ } \mu\text{eV}$ )
- Test of active resonant feedback.

I won't discuss this here but see: E.J. Daw, Resonant feedback for axion and hidden sector dark matter searches, Nucl. Instrum. Meth. A 921 (2019) 50 [arXiv:1805.11523].



An STFC funded facility to be located in Sheffield, UK

We have a 100 DU with ADMX

- 10 mK target temperature
- At least a 8 T magnet field
- 20 cm bore by 20 cm high



DF Lab



DF Aux Lab – Compressors, pumps, benches, etc

- A Dilution Fridge (DF) and 8 T Superconducting magnet are on order – delivery expected in 12-14 months
- Refurbishment of lab space about to commence.



# QSHS Test Facility

An STFC funded facility located in Sheffield, UK

- We have a MOU with ADMX
- 8.5 mK base temperature
- 8 T magnetic Field (6 T currently)
- 180 mm bore by 200 mm high



DF Lab



Plant Room – Compressors, pumps, etc.

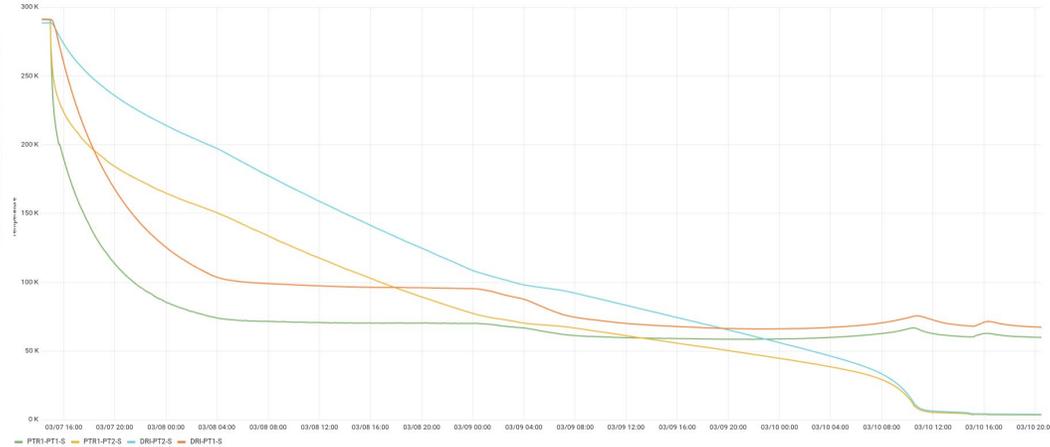
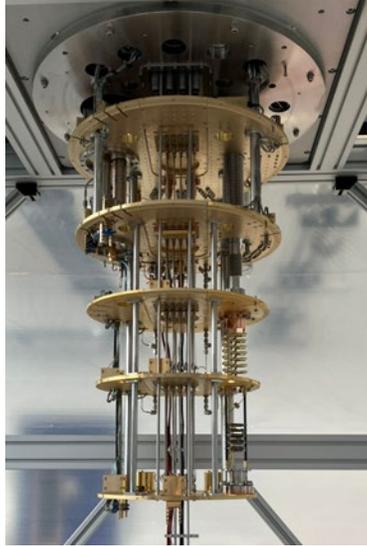
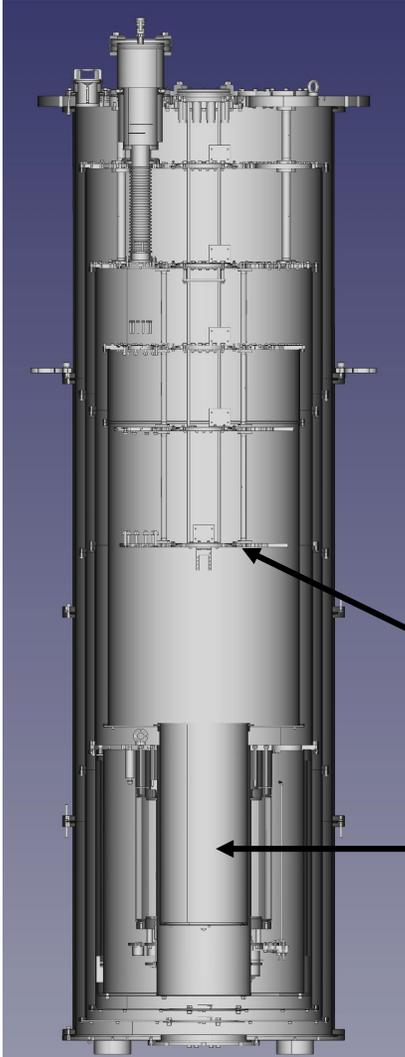


Auxiliary Lab

- Dilution Fridge now installed, thermometry calibration and cavity support installation underway.

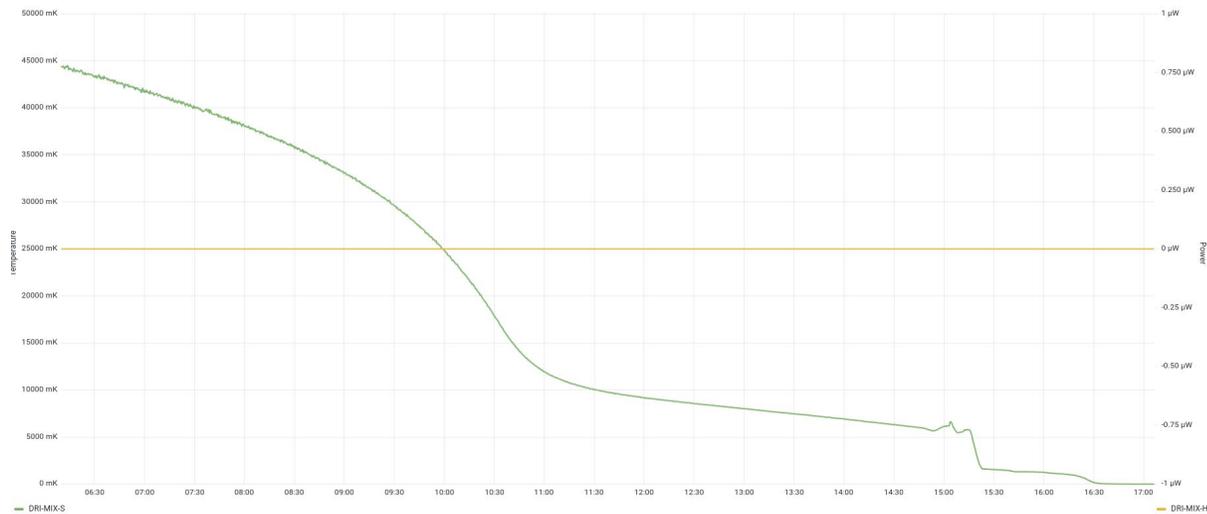


# QSHS Test Facility



1<sup>st</sup> stage amp here.

Cavities go here

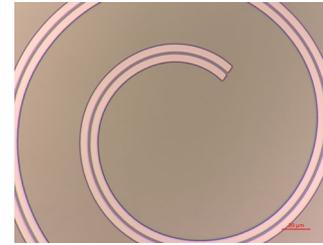
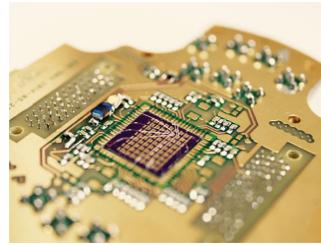
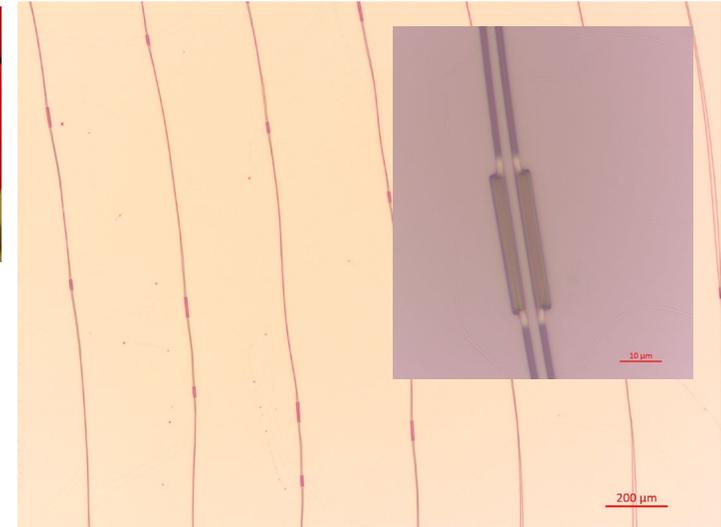
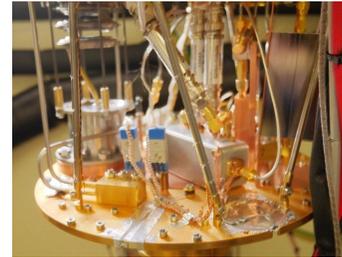




# Quantum Electronics for QSHS 1

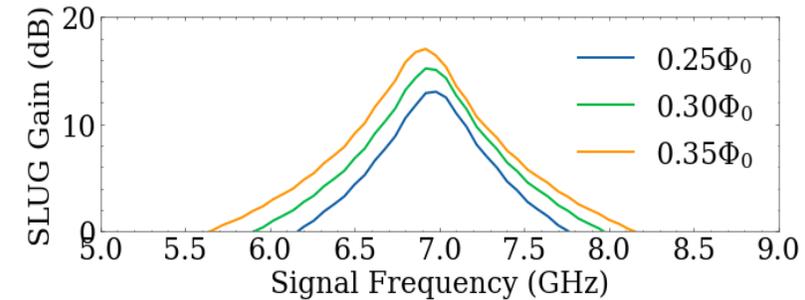
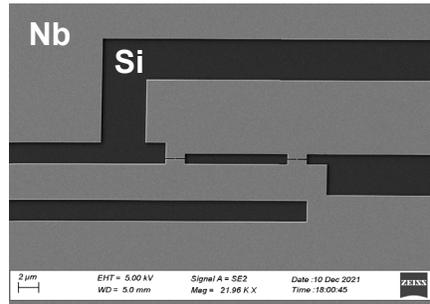
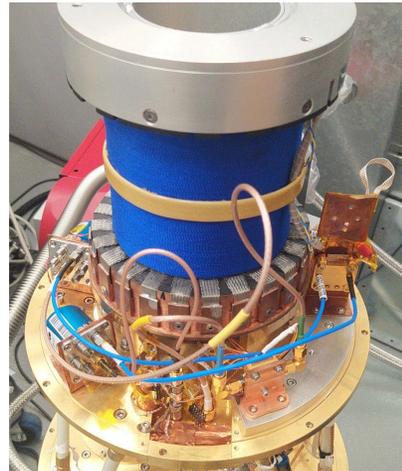
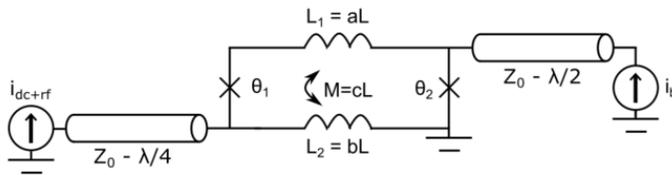
## Parametric Amplifiers

Josephson Parametric Amplifiers  
Travelling Wave Parametric Amplifiers



## SLUG loaded (SQUID) Amplifiers

High frequency RF amplifiers



Calculated frequency-dependent gain of a SLUG amplifier. The signal is input to the SLUG via a  $\lambda/4$  transmission line resonator with characteristic impedance  $Z_0 = 2 \Omega$  and bare resonant frequency  $f_{res} = 8$  GHz.

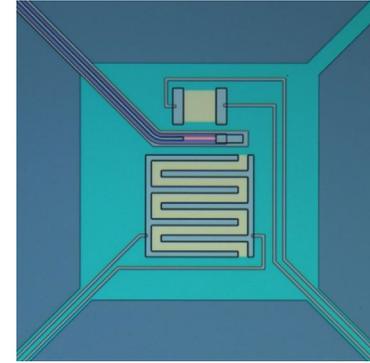
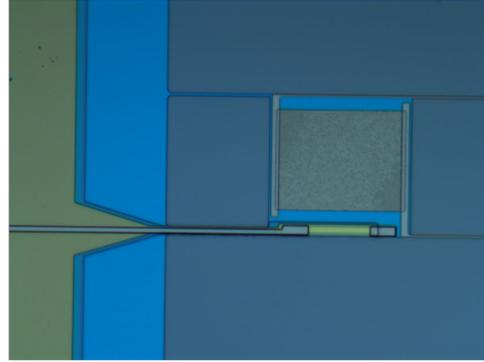


CMD29 (Manchester, Aug. 2022) – Developing a SLUG Microwave Amplifier for Axion Detection



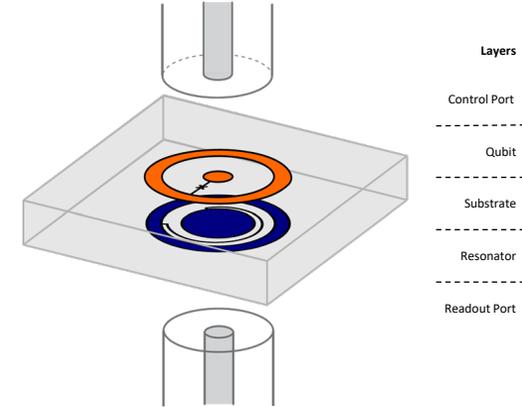
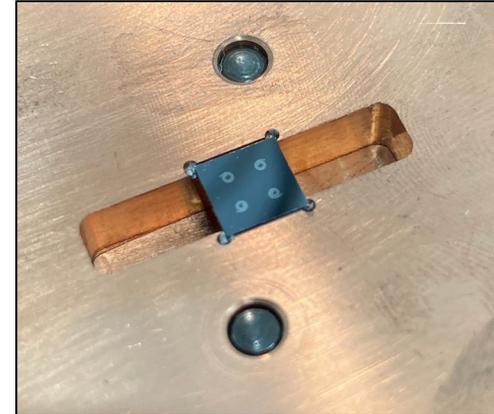
# Quantum Electronics for QSHS 2

## Bolometers



- At  $\sim 10$  mK, noise equivalent powers (NEP's) of  $< 10^{-21}$   $\text{WHz}^{-1/2}$  should be possible.
- Will permit a broad-range search over the cavities resonant bandwidths

## Qubits



- Fabricated a test device with qubits and resonators
- Built and measured a waveguide sample holder
- Demonstrated feasibility of multiplexed readout with waveguide architecture



# Summary

- The QSHS collaboration is building a world-class programme in this area over eight institutions
- A new UK based facility furthering in development that will search for axions via the resonant cavity method.
- A dilution fridge has been installed.
- Development of quantum devices is taking place to improve scan rate and detection of very faint signals:
  - Amplifiers
  - Qubits
  - Bolometers
- Collaboration with US colleagues (ADMX) on resonators.
- The long-term goal is a large-scale UK facility.