

# Tracing Cosmic Evolution through Dark Matter Phenomena: The Origin of Supermassive Black Holes in JWST Era

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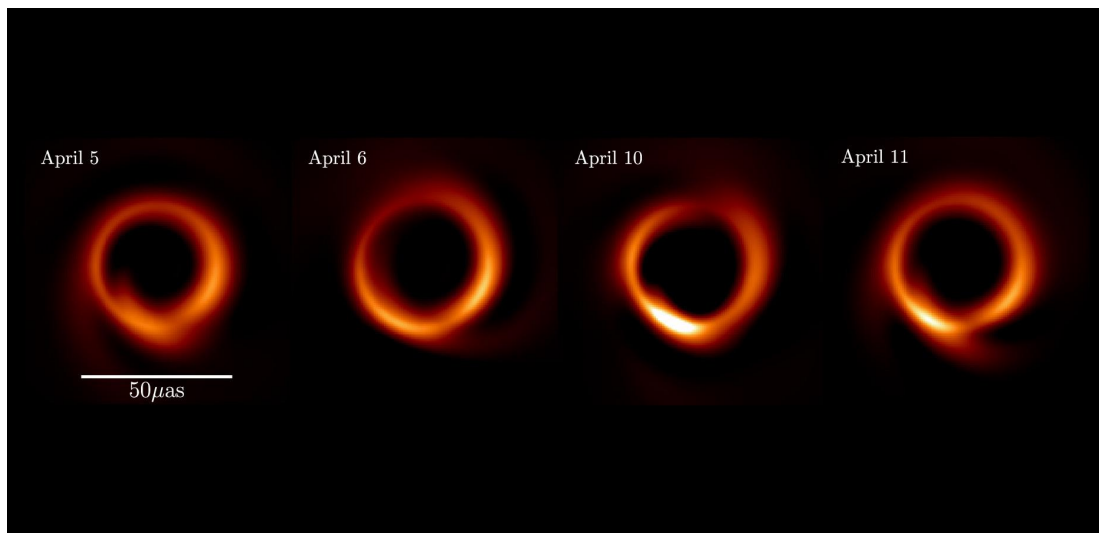
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# Content

- Supermassive Black Holes (SMBHs): an overview
- Results from the latest JWST observations
- Beyond Standard Model physics on SMBH seeding

# Supermassive Black Hole (SMBH)

- Mass  $>10^6$  solar mass black holes
- Accrete surrounding material through gravity and form accretion disk
- Produce copious amount of X-ray emission, also a radio source
- Could be observed through emission line broadening (fast moving clouds)



M87, Medeiros et al. (2023) [2304.06079]

Balancing the two forces gives:

$$\frac{GMm}{R^2} = \frac{L\kappa m}{4\pi R^2 c}$$

And solving for this luminosity we get:

$$L = \frac{4\pi GMc}{\kappa}$$

This has a profound implication. If our accreting object radiates at more than the Eddington luminosity, even a glut of “fuel” will be blown away by radiation pressure: we get a natural feedback process with a limiting accretion rate. We derive this by setting the accretion luminosity equal to the Eddington luminosity:

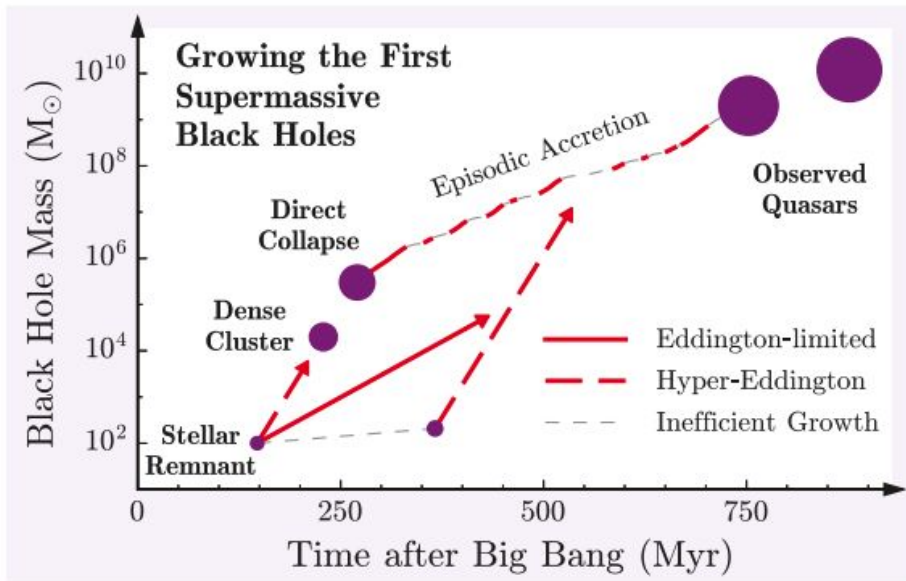
$$\epsilon \dot{M} c^2 = \frac{4\pi GMc m_p}{\sigma_T}$$

From which the limiting *Eddington accretion rate* is:

$$\dot{M}_{\text{Edd}} = \frac{4\pi GM m_p}{\epsilon c \sigma_T}$$

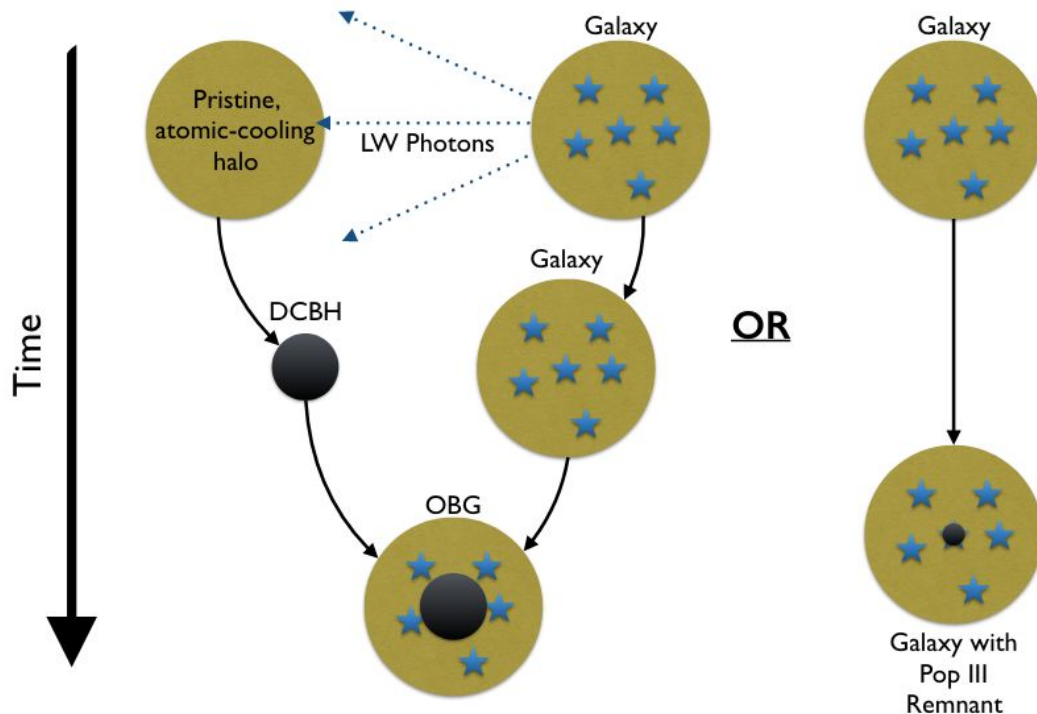
<https://www-astro.physics.ox.ac.uk/~garret/teaching/lecture7-2012.pdf>

# Black Hole Growth and Eddington Limit

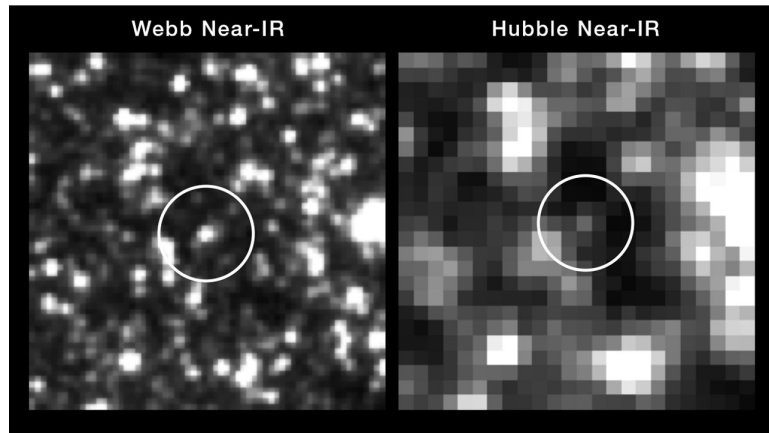
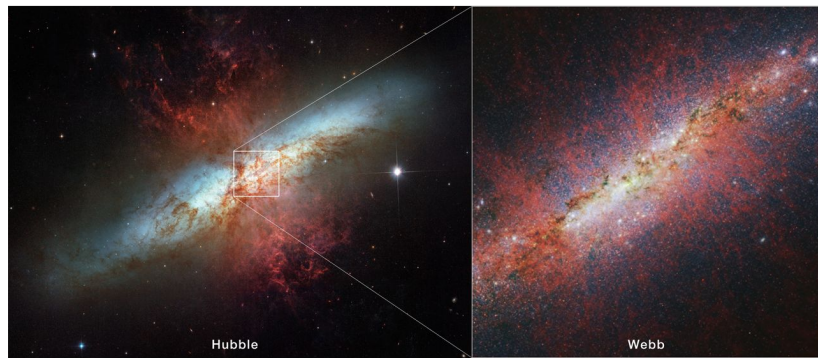
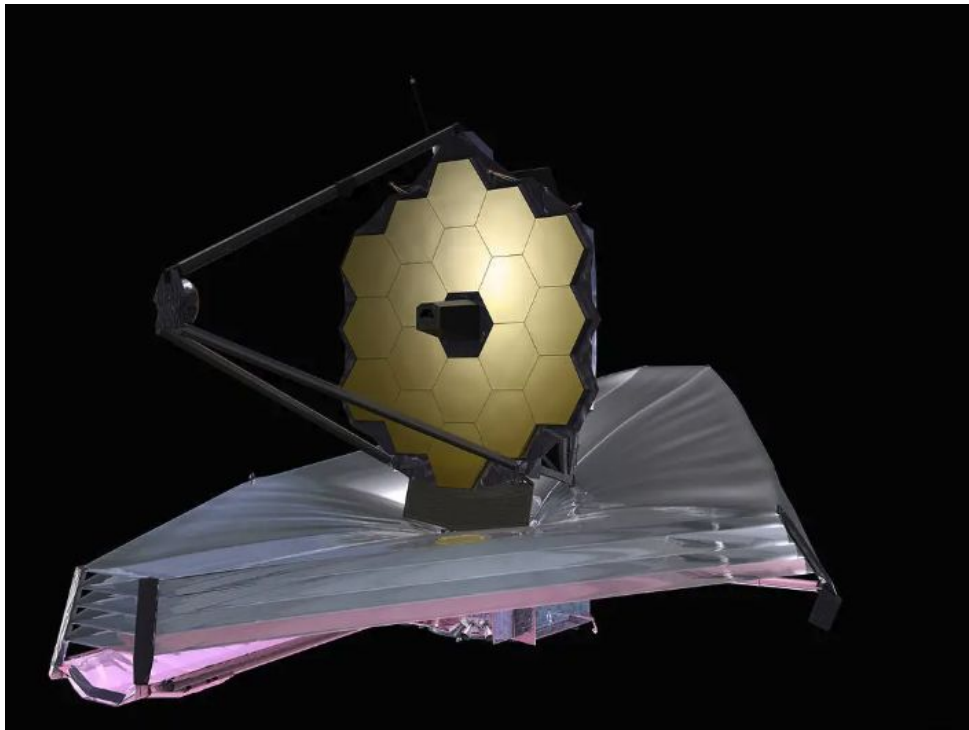


# Light Seed vs. Heavy Seed in $\Lambda$ CDM Model

- Pop III star remnants
- Dense Star Cluster
- Direct Collapse Black Hole (DCBH)



# James Webb Space Telescope (JWST)



<https://www.nasa.gov/image-article/james-webb-space-telescope-jwst/>

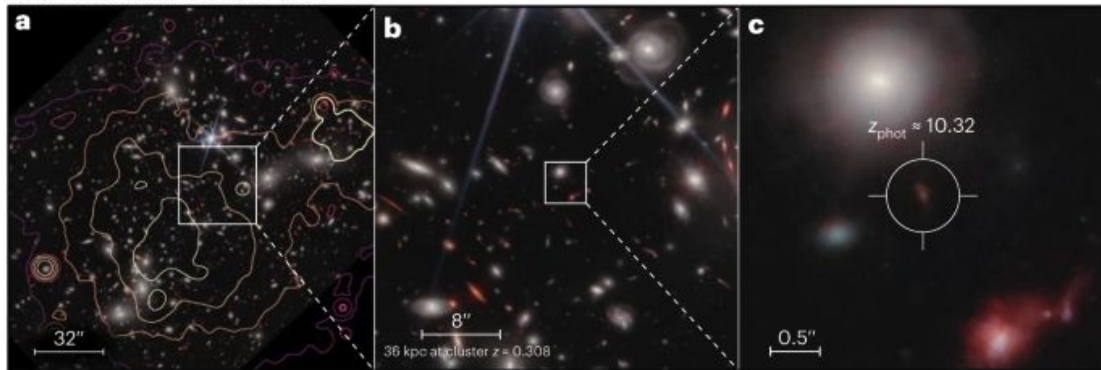
# Questions on Cosmology from JWST Discovery

- Stress testing standard universe model (Boylan-Kolchin (2023) [2208.01611])
  - Discovery of too massive galaxies with stellar mass  $>10^9$  solar mass at high redshift  $z > 10$
  - Challenging current cosmic structure/galaxy formation theory and simulations
  - Need new theory/models to explain, or over-prediction of stellar mass (Wang et al. (2024) [arXiv:2403.02399])
- Early Supermassive Black Hole formation
  - See next page

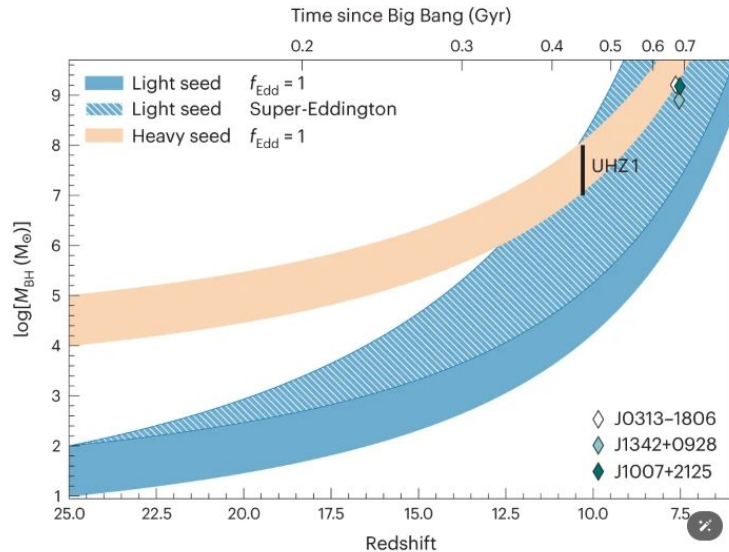
# Some Recent Discoveries from JWST

**Fig. 2: JWST and Chandra images of UHZ1.**

JWST NIRCам zoom-in on UHZ1



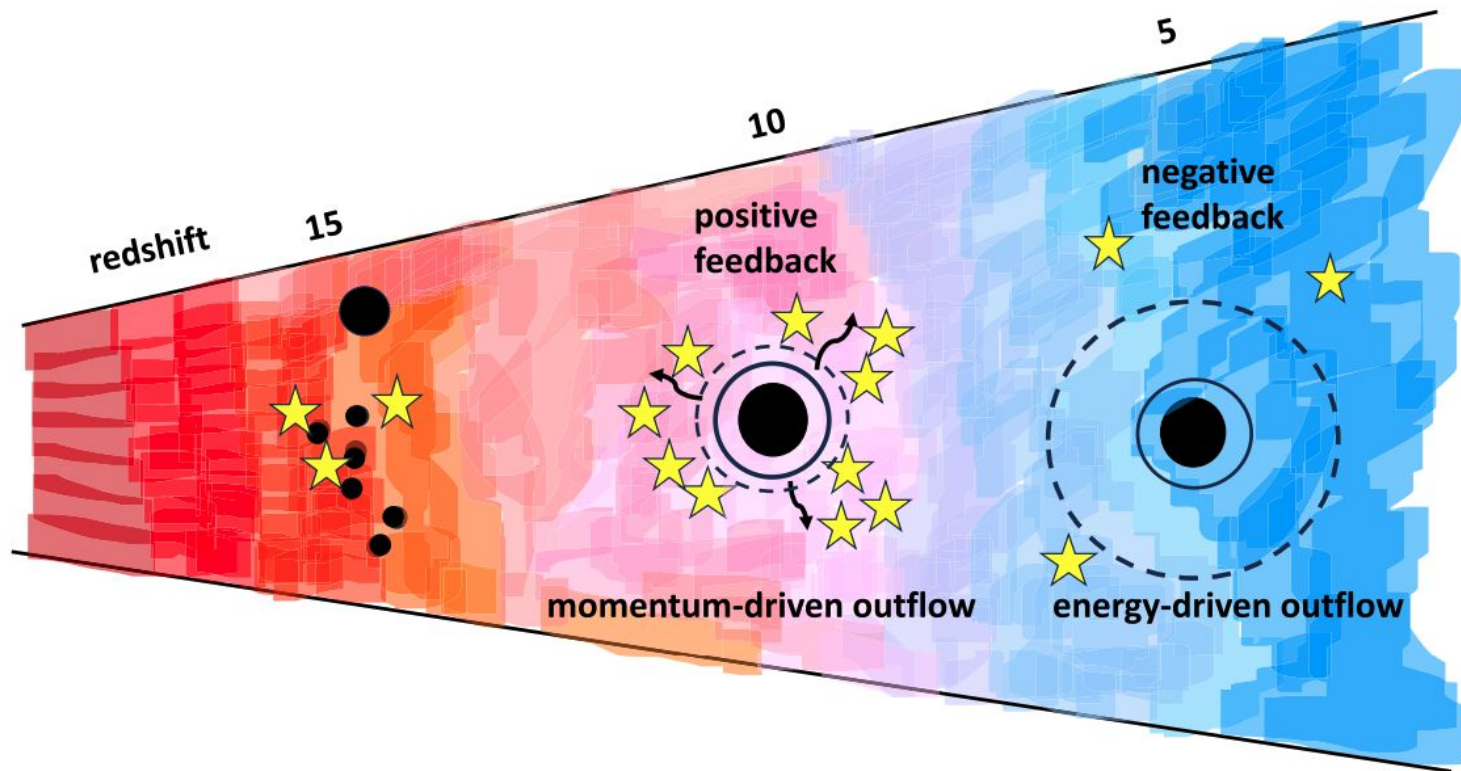
JWST NIRCам UHZ1 images



$z=10$  AGN: UHZ-1



# Co-evolution of SMBH with Galaxy



# Beyond Standard Model Path

- Supermassive Dark Star (WIMP)
- (Massive) Primordial Black Holes (PBH)

# About Dark Stars



<https://www.sci.news/astronomy/webb-supermassive-dark-stars-12096.html>

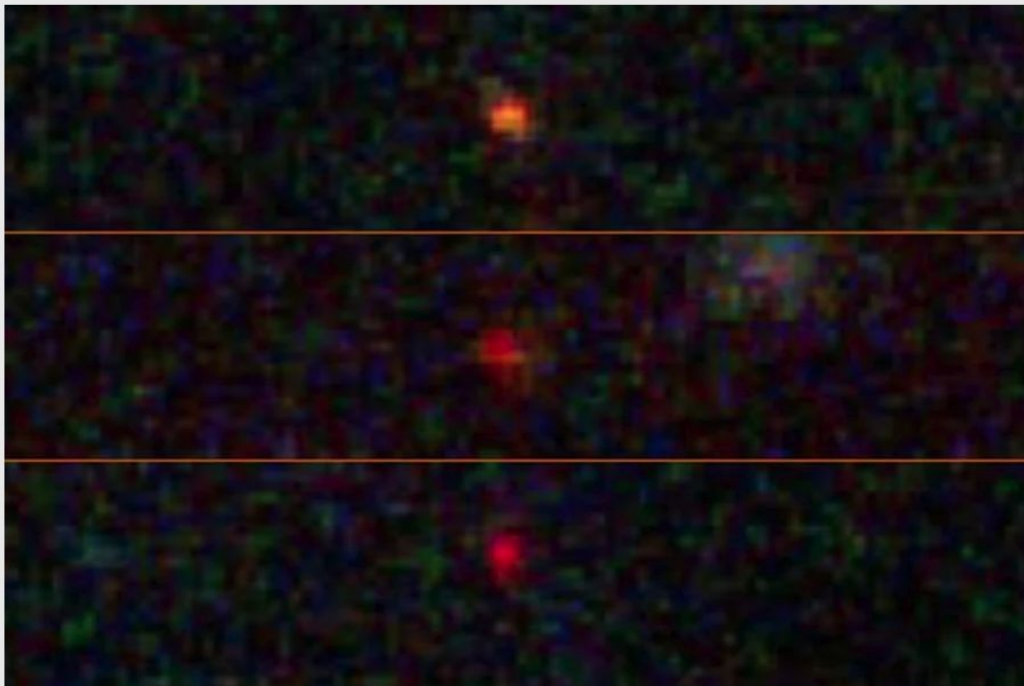
- Formed in the center of mini-halo
- Powered by DM(WIMPs), no fusion reaction
- Capture DM to refuel
- Can grow supermassive to about  $\sim 10^7$  solar mass
- Collapse into massive black hole seed after fuel depleted

Freese et al. (2008) [0802.1724]

JULY 20, 2023 | 5 MIN READ

## JWST Might Have Spotted the First Dark Matter Stars

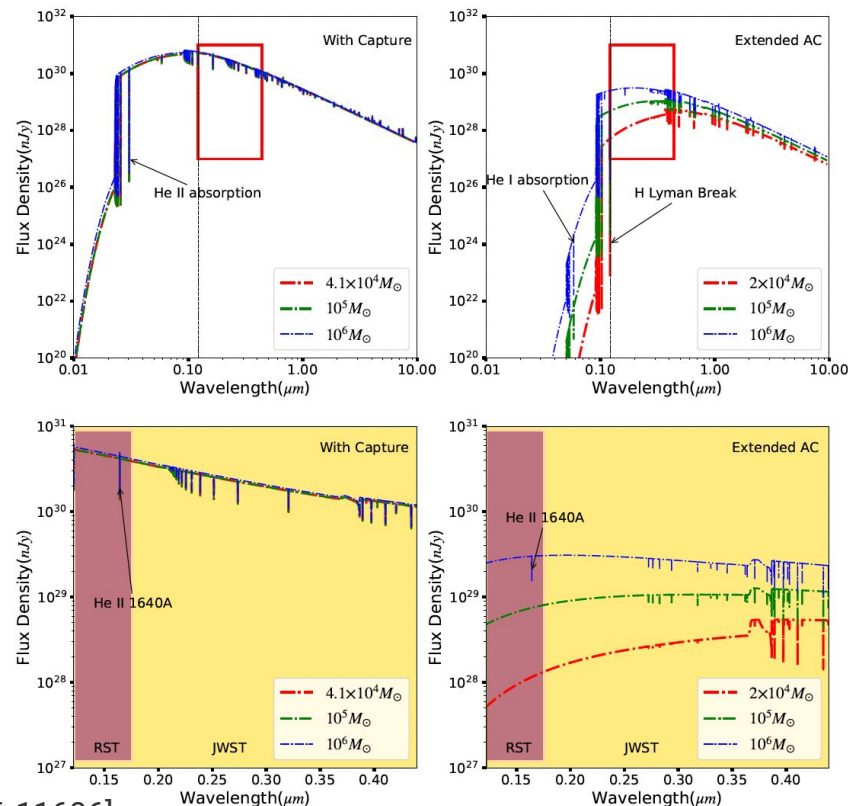
Stars fueled by the self-annihilation of dark matter might have been spotted for the first time by JWST

BY [STEPHANIE PAPPAS](#)

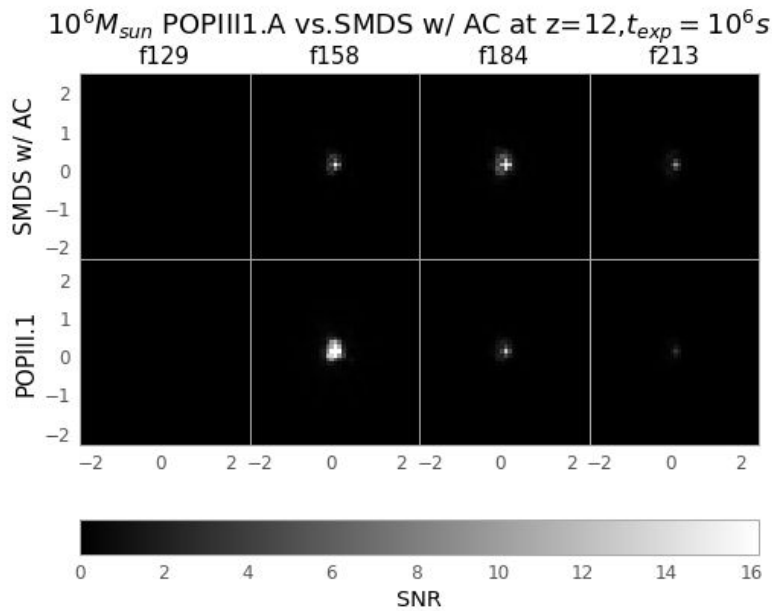
<https://www.scientificamerican.com/article/jwst-might-have-spotted-the-first-dark-matter-stars/>

# Dark Stars Detectable by Roman Space Telescope (RST)

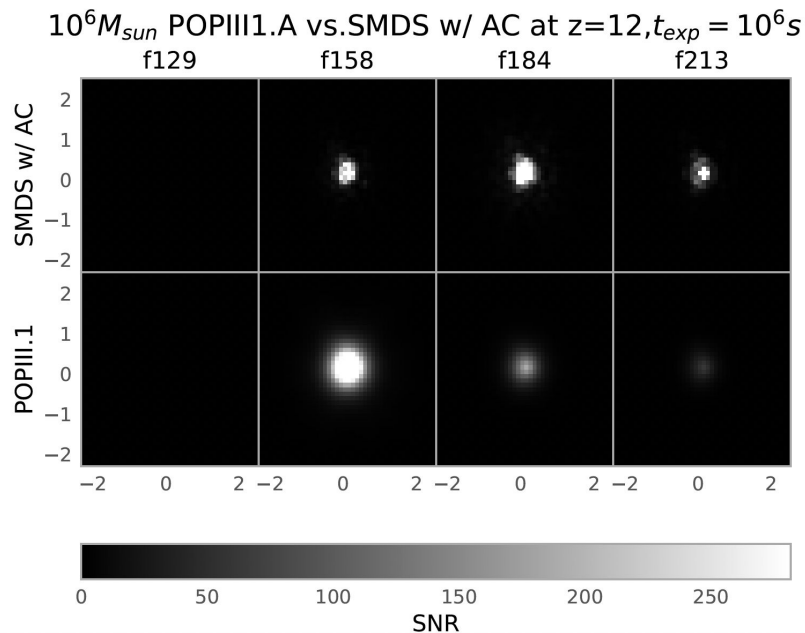
- RST have similar detection limit as JWST but have much wider field of view
- RST could detect supermassive dark stars up to  $z \sim 14$ , and capable of distinguishing them from early galaxies by unique spectral features
- Lensing by foreground object can help us find less massive dark stars



# Dark Stars Detectable by Roman Space Telescope (RST)

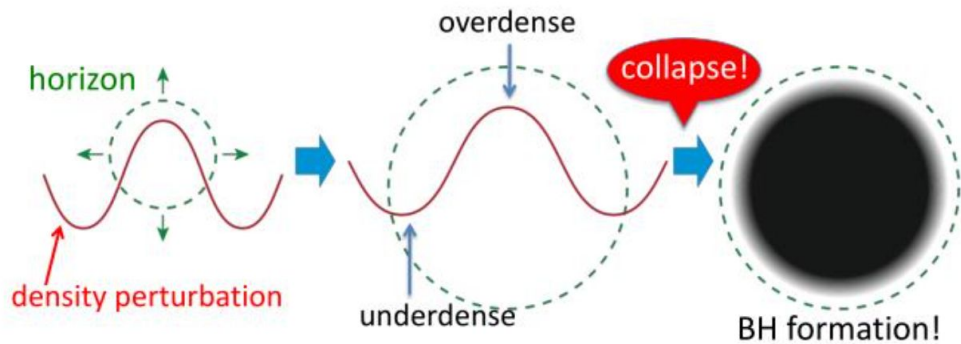


No Lensing



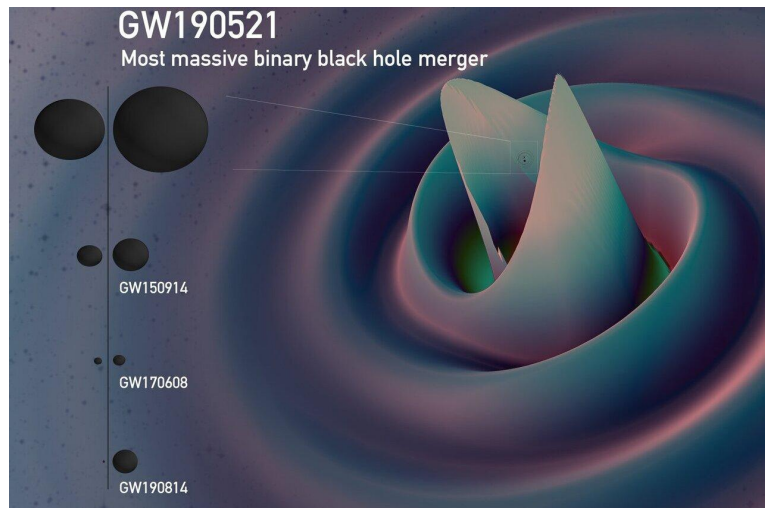
100x Lensing

# Primordial Black Holes



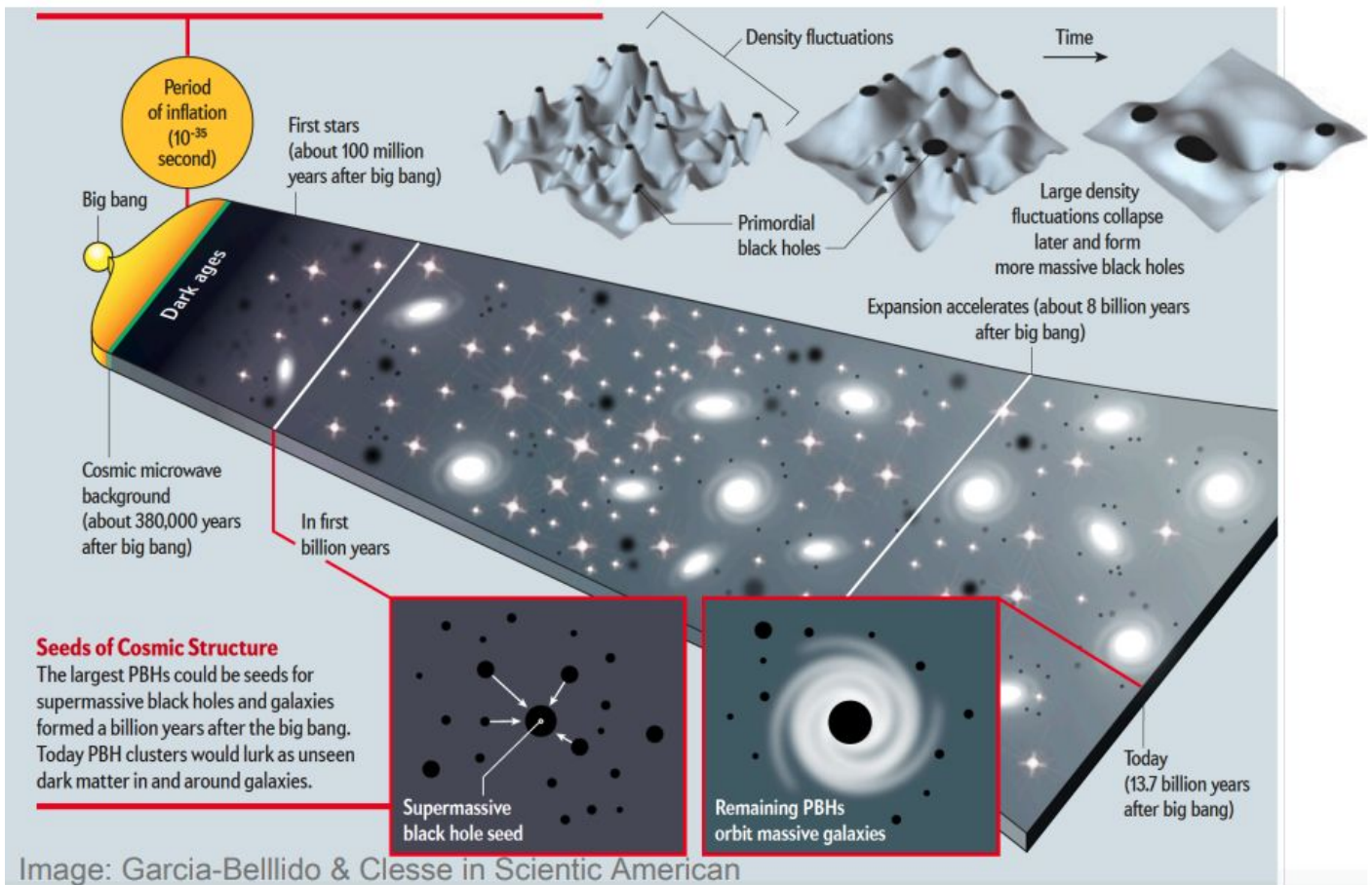
## Formation

[https://en.wikipedia.org/wiki/Primordial\\_black\\_hole](https://en.wikipedia.org/wiki/Primordial_black_hole)



## Gravitational Wave as a Probe

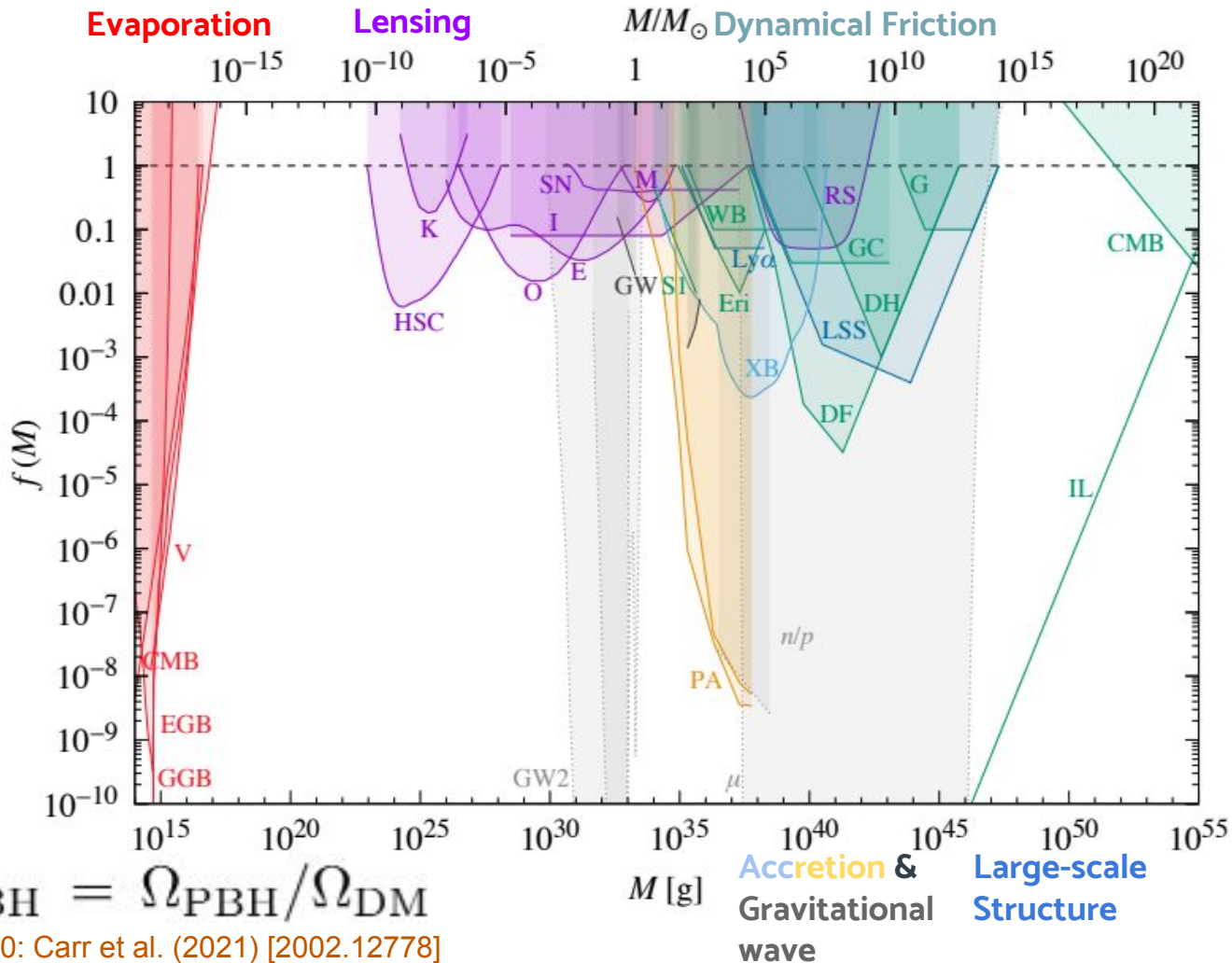
<https://phys.org/news/2021-03-gw190521-event-primordial-black-holes.html>



**Credit:**  
[https://kspa.soe.ucsc.edu/sites/default/files/Lecture1\\_PN.pdf](https://kspa.soe.ucsc.edu/sites/default/files/Lecture1_PN.pdf)

Image: Garcia-Bellido & Clesse in Scientific American



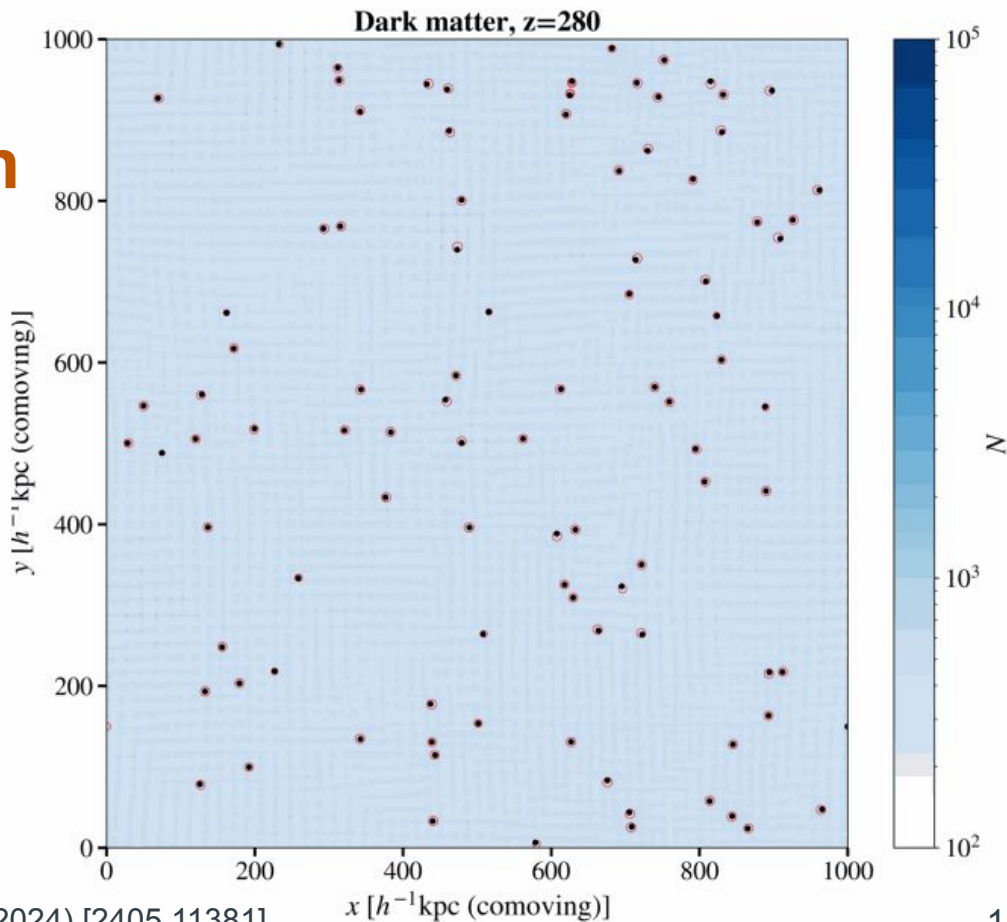


$$f_{\text{PBH}} = \Omega_{\text{PBH}} / \Omega_{\text{DM}}$$

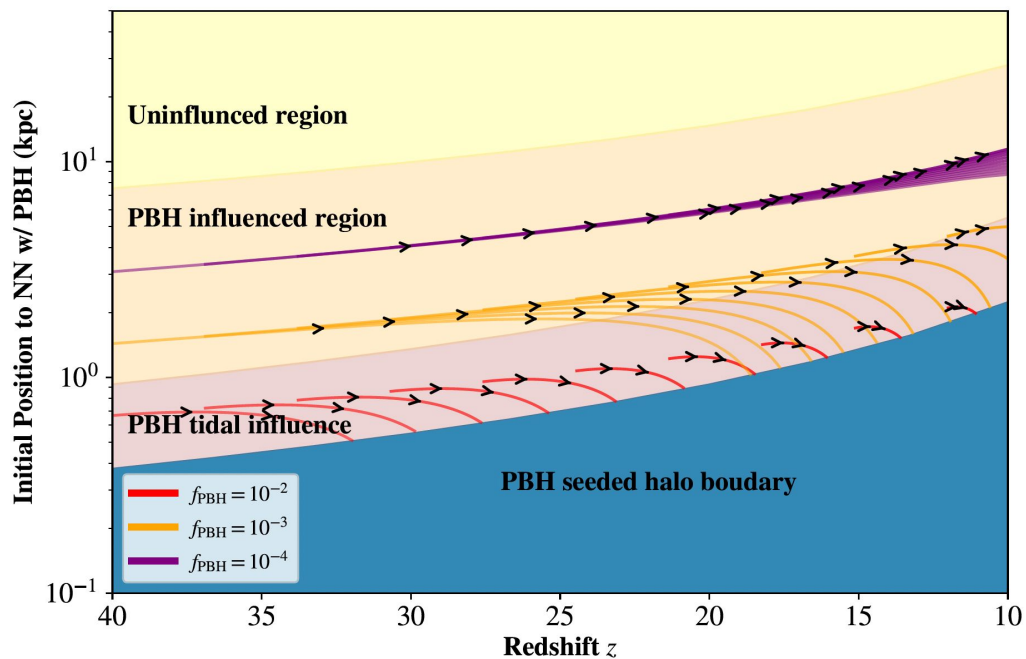
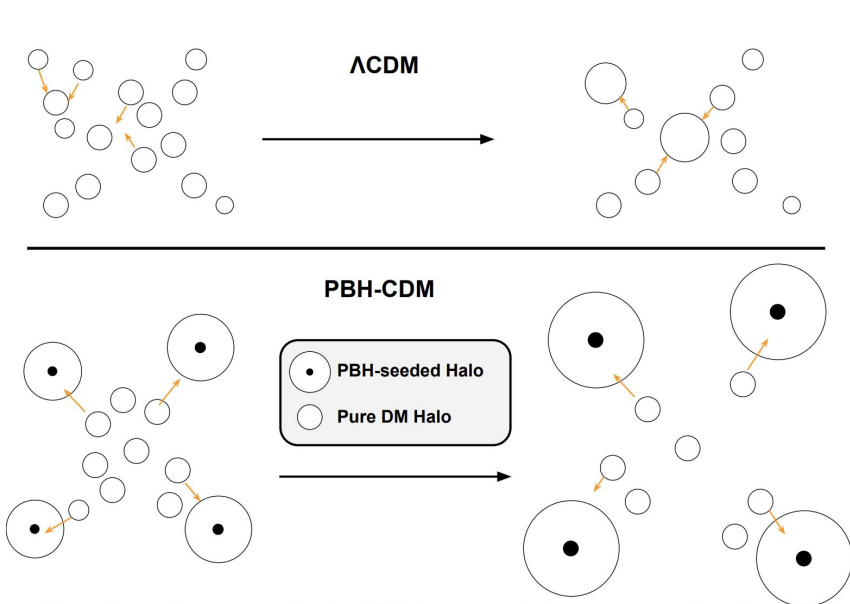
Fig 10: Carr et al. (2021) [2002.12778]

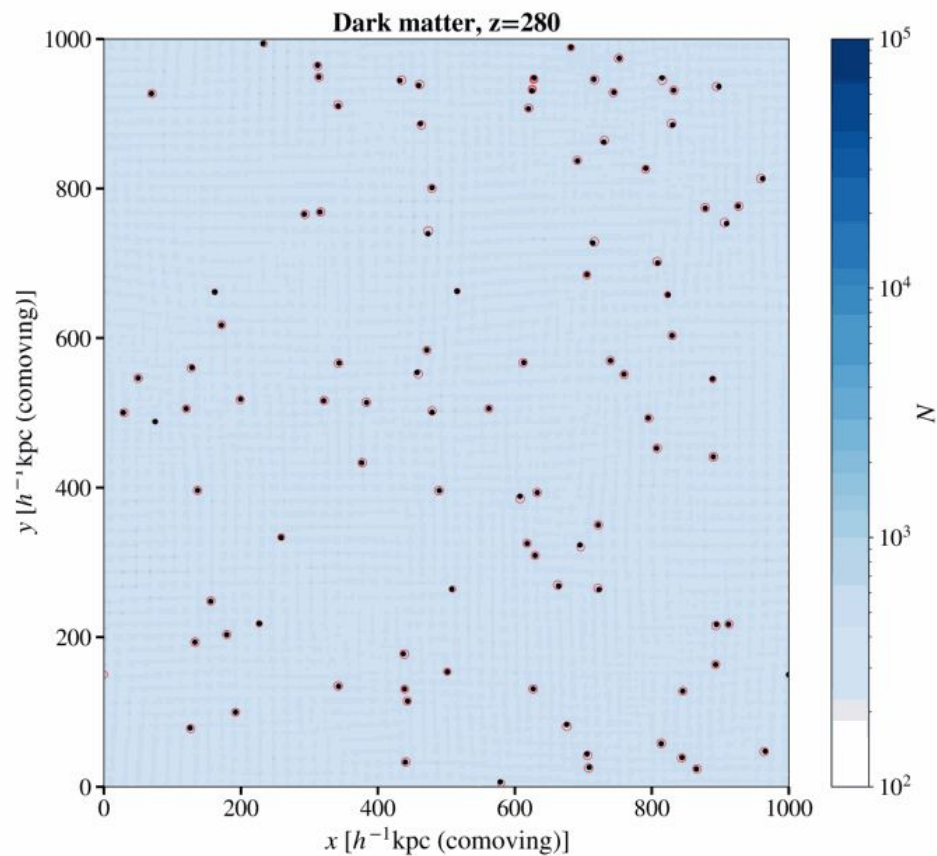
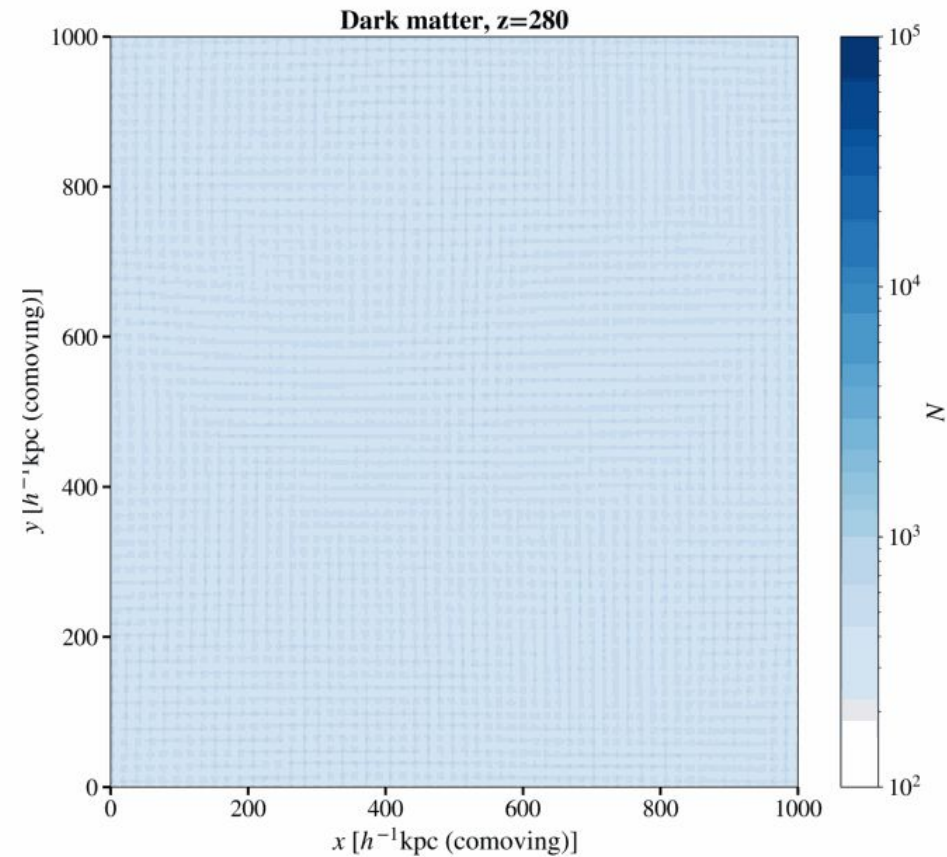
# Primordial Black Holes and Structure Formation

- $10^6$  solar mass **Primordial Black Hole** candidate used in this work
- **Primordial Black Holes** will seed the formation of halos surrounding them
- **Primordial Black Hole** seeded halos will engulf newly formed neighboring halos



# Halo Dynamics in PBH Universe





# In the Future

- ❑ Run more cosmological simulations with different initial conditions/scales for Primordial Black Holes
- ❑ Study how would different Primordial Black Hole candidates influence the formation of the first stars and galaxies
- ❑ Find more possible observational features from Dark Stars and Primordial Black Holes
- ❑ Data mining from available JWST dataset

# Summary

- ❑ In the era of deep sky surveys, we are closer to understand the formation of SMBH in early universe
- ❑ Early JWST results favor the heavy seed scenario
- ❑ Different astrophysical phenomenologies from beyond standard model DM could also form massive seeds to solve the problem involving supermassive black hole formation

Thank you!  
Grazie!