

EXPERIMENTAL DARK MATTER SEARCH AT MASS > 1 GeV



MARIA ELENA MONZANI, SLAC/KIPAC
IDM - L'AQUILA, JULY 10, 2024



**YOUR
PRESENTER
AT GRAN
SASSO LAB,
CIRCA
2002 (?)**

**...
THANKS
FOR
HAVING ME
BACK!!!**

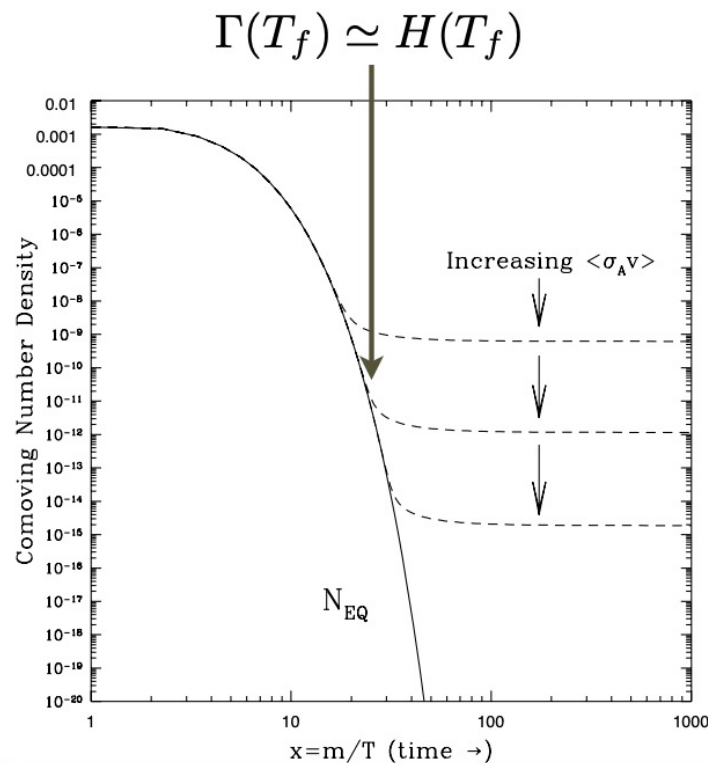


**RECAP: DIRECT DETECTION
OF WIMP DARK MATTER**

WIMPs AS THERMAL RELICS

A trigger from naturalness versus the hierarchy problem, and thermal relic **WIMPs** as natural dark matter candidates.

Thermal relics directly coupled to the baryon/photon primordial bath: $\chi \bar{\chi} \leftrightarrow \text{SM} \bar{\text{SM}}$ (with SM is some lighter Standard Model state)



$$\Omega_\chi h^2 \simeq \frac{3 \cdot 10^{-27} \text{cm}^{-3} \text{s}^{-1}}{\langle \sigma_A v \rangle_{T=T_f}}$$

WIMP miracle: “fixed” DM pair annihilation cross section into “visible” particles.

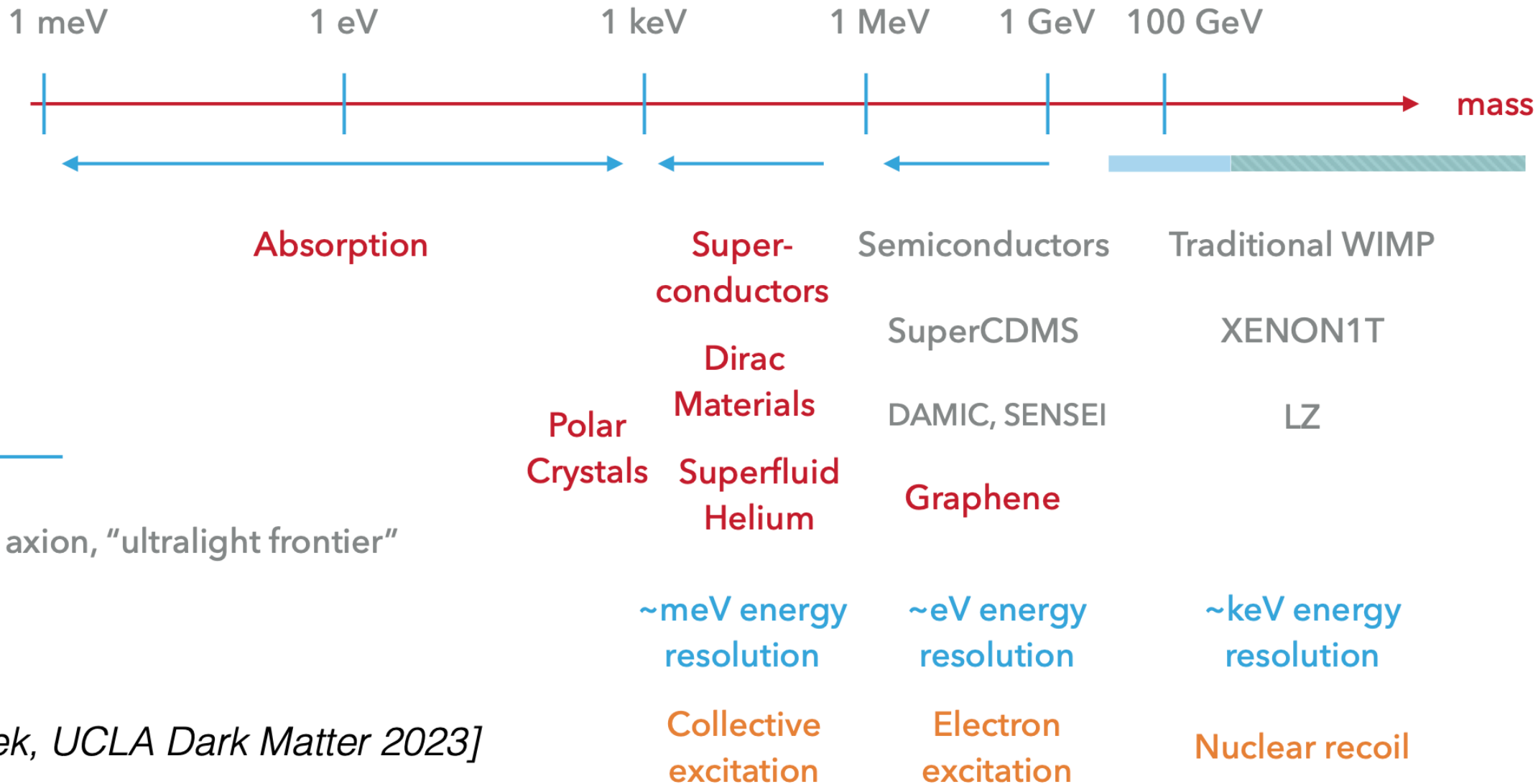
A recipe that can work below about **100 TeV** (unitarity limit [Griest & Kamionkowski 1990]; in realistic models up to about 15 TeV) and gets inefficient below about **1 GeV**.

Experimental Panorama

L. CANONICA

D. BAXTER

THIS TALK

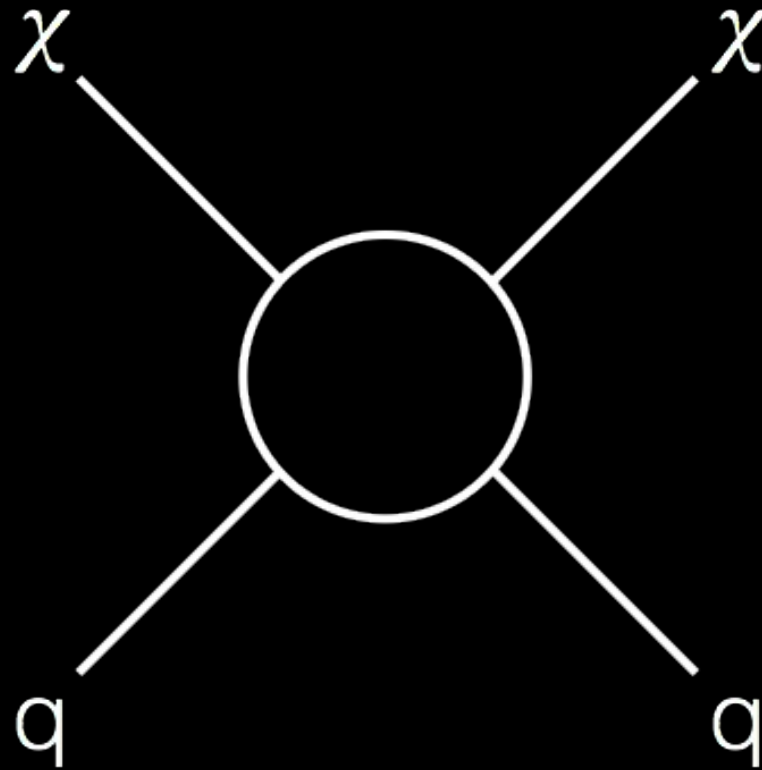


[Zurek, UCLA Dark Matter 2023]

P. ULLIO, JULY 8, 2024

THREE WAYS TO LOOK FOR WIMPS

Annihilation



Production



Indirect Detection



Colliders

Scattering



Direct
Detection

THE PRINCIPLE OF DIRECT DETECTION

The background of the slide is a dark space filled with numerous stylized atoms. Each atom consists of a central nucleus made of red and green spheres, surrounded by a cloud of blue spheres representing electrons. A large, white, three-dimensional arrow points from the top left towards the center of the slide, passing through several atoms.

WIMPs and Neutrons
scatter from the
Atomic Nucleus

**BUILD A MASSIVE
TANK (OR TOWER)
OF NUCLEI**

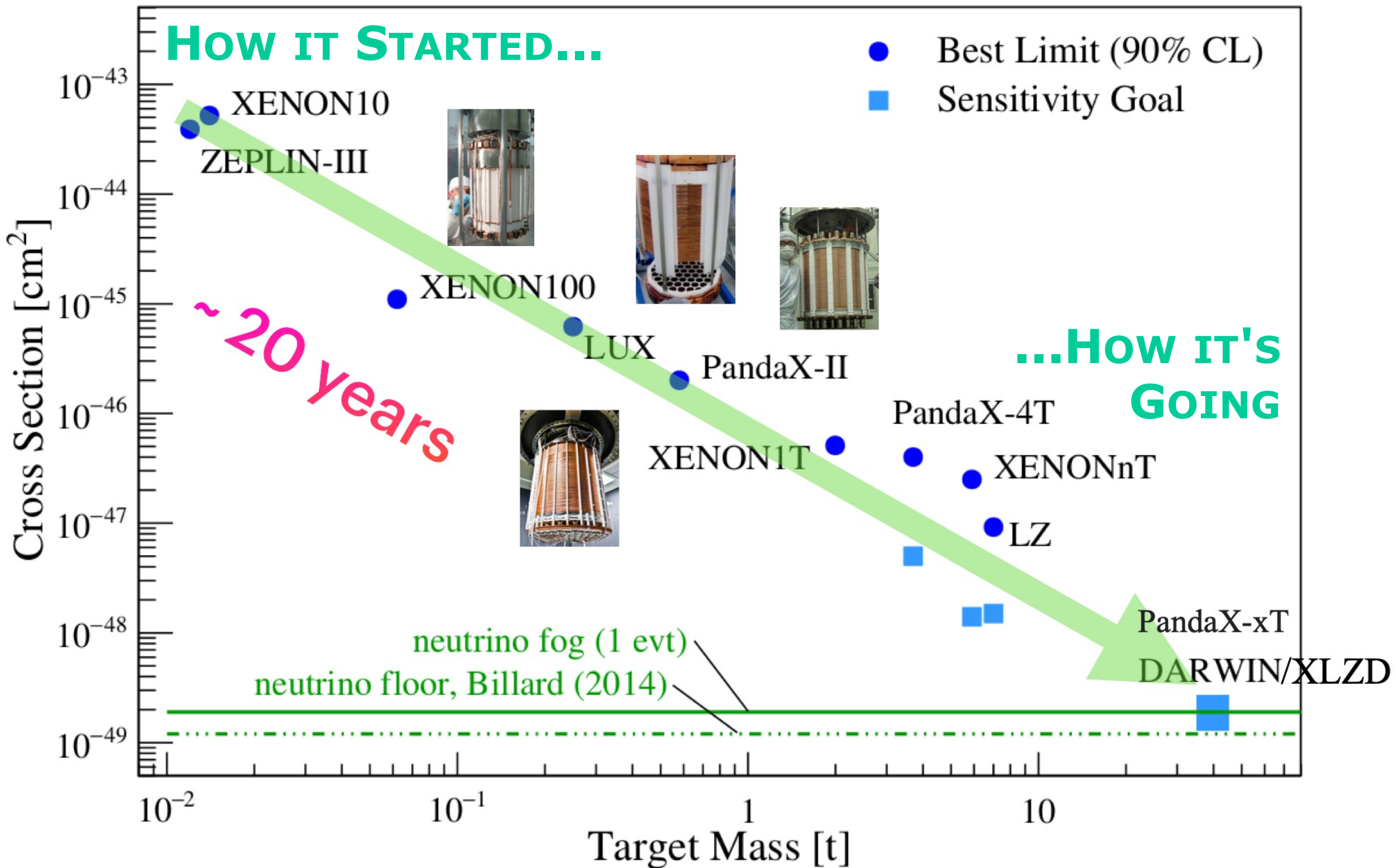
**“HIDE” IT DEEP
UNDERGROUND**

**WAIT FOR WIMPS
TO HIT THE NUCLEI**

Photons and Electrons
scatter from the
Atomic Electrons

**LOOK FOR RECOILS
FROM NUCLEI THAT
HAVE BEEN BUMPED
BY A WIMP**

$$m_\chi = 50 \text{ GeV}$$



P. ULLIO, JULY 8, 2024

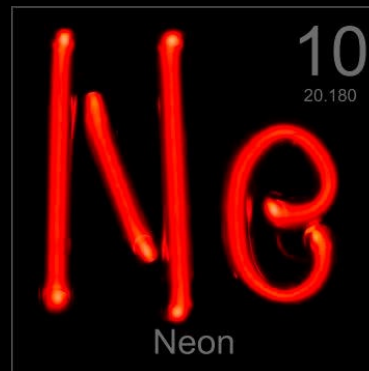
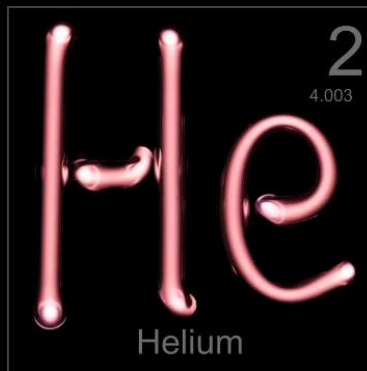
[Snowmass 2021, arXiv:2203.08084, adapted by Baudis 2024]

A deep-field astronomical image showing a vast number of galaxies, including spiral, elliptical, and irregular shapes, scattered across a dark cosmic background. A grid of thin blue lines is overlaid on the image, with several bright star-like points at the intersections. The text is positioned in the lower right quadrant.

**THE SEARCH FOR WIMP DM
IN THE NOBLE LIQUIDS ERA**

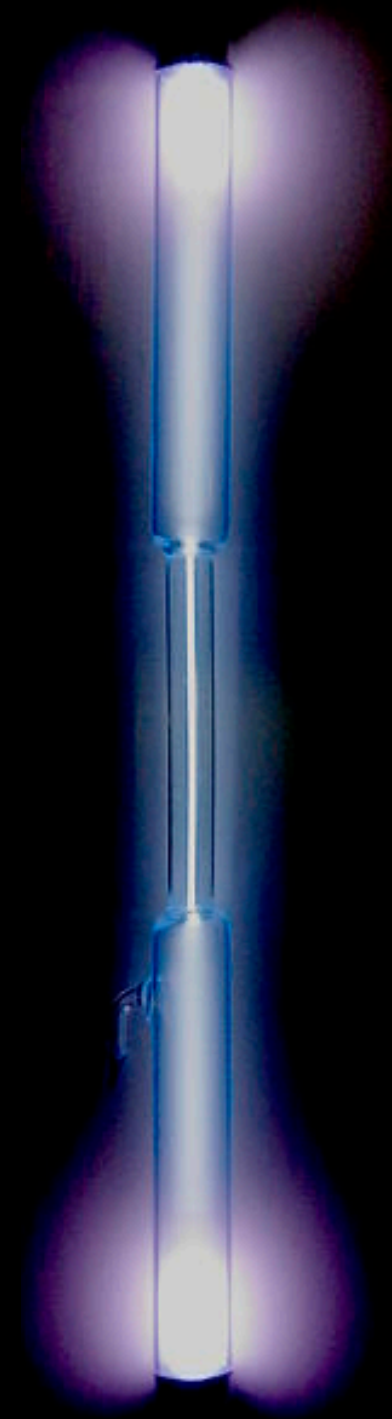
WHY NOBLE LIQUIDS?

- **KINEMATIC MATCHING TO WIMPS**
- **LOTS OF NUCLEONS PER ATOM**
- **TRANSPARENT TO THEIR OWN LIGHT**

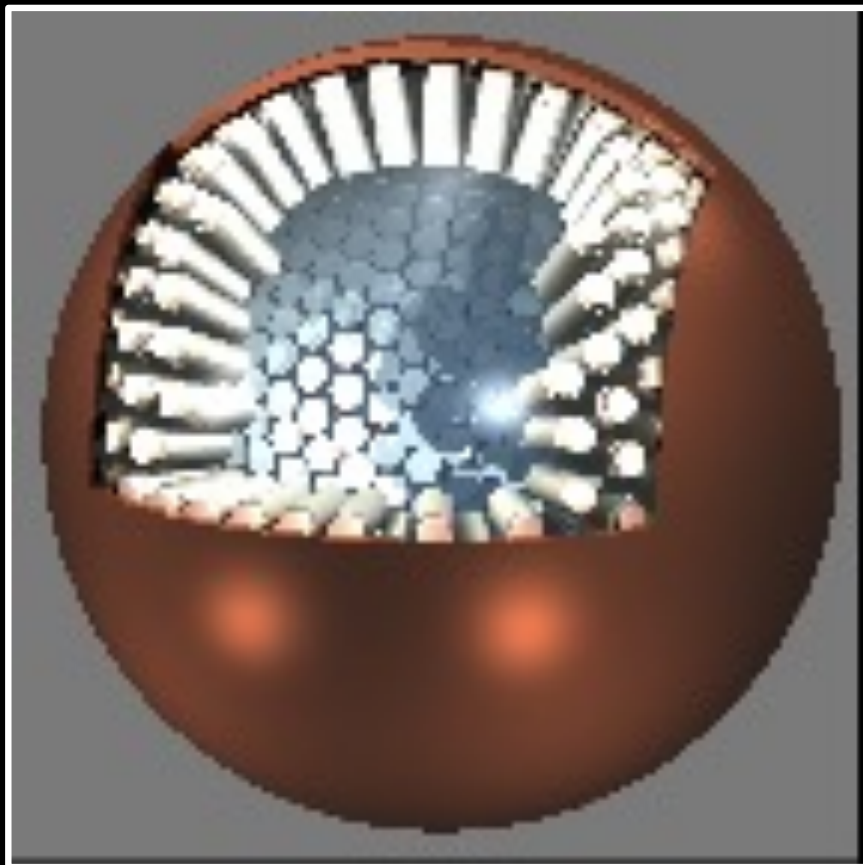


WHY NOBLE LIQUIDS?

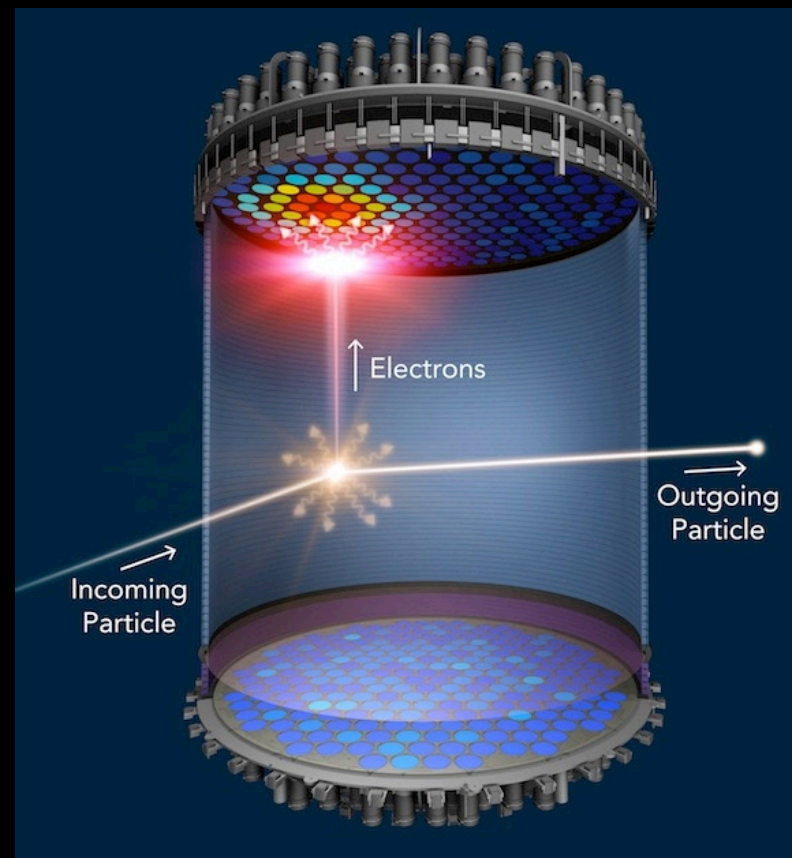
- **KINEMATIC MATCHING TO WIMPS**
- **LOTS OF NUCLEONS PER ATOM**
- **TRANSPARENT TO THEIR OWN LIGHT**
- **VERY DENSE (SELF-SHIELDING)**
- **BACKGROUND REJECTION (CHARGE/LIGHT)**
- **LIQUID: CAN BE PURIFIED IN A LOOP**
- **“EASY” TO MAKE A LARGER DETECTOR**



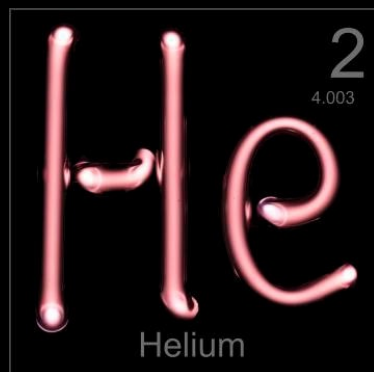
SINGLE PHASE VS DUAL PHASE



4π Scintillation

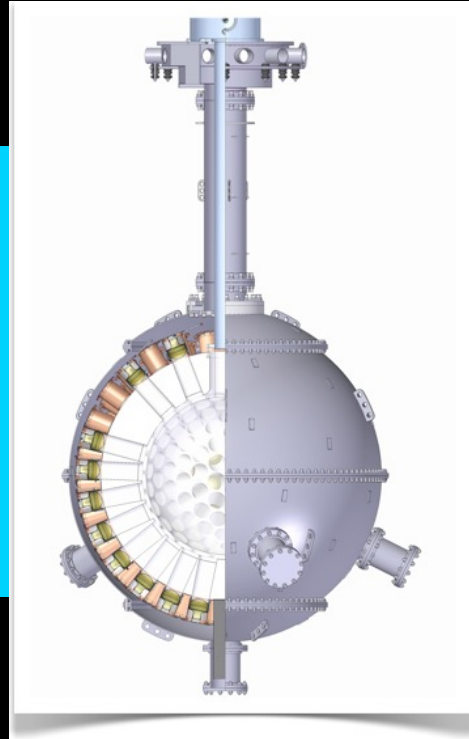


Time Projection Chamber (TPC)

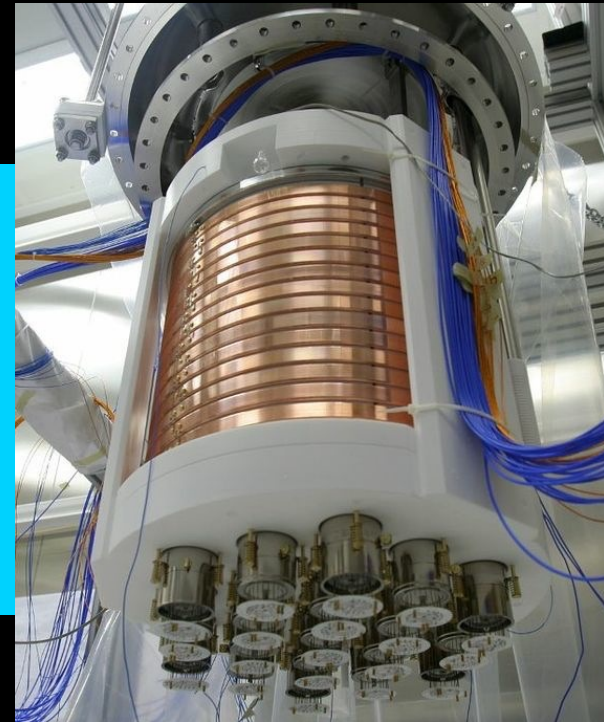


SINGLE PHASE

DUAL PHASE



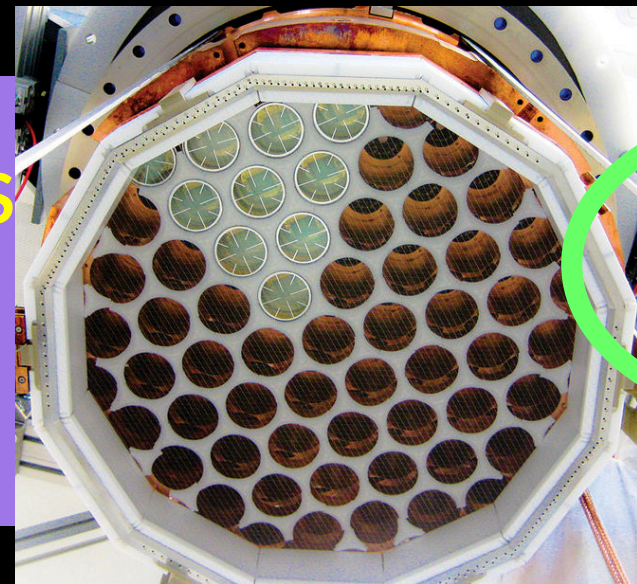
DEAP
MiniClean



DarkSide
ARGO

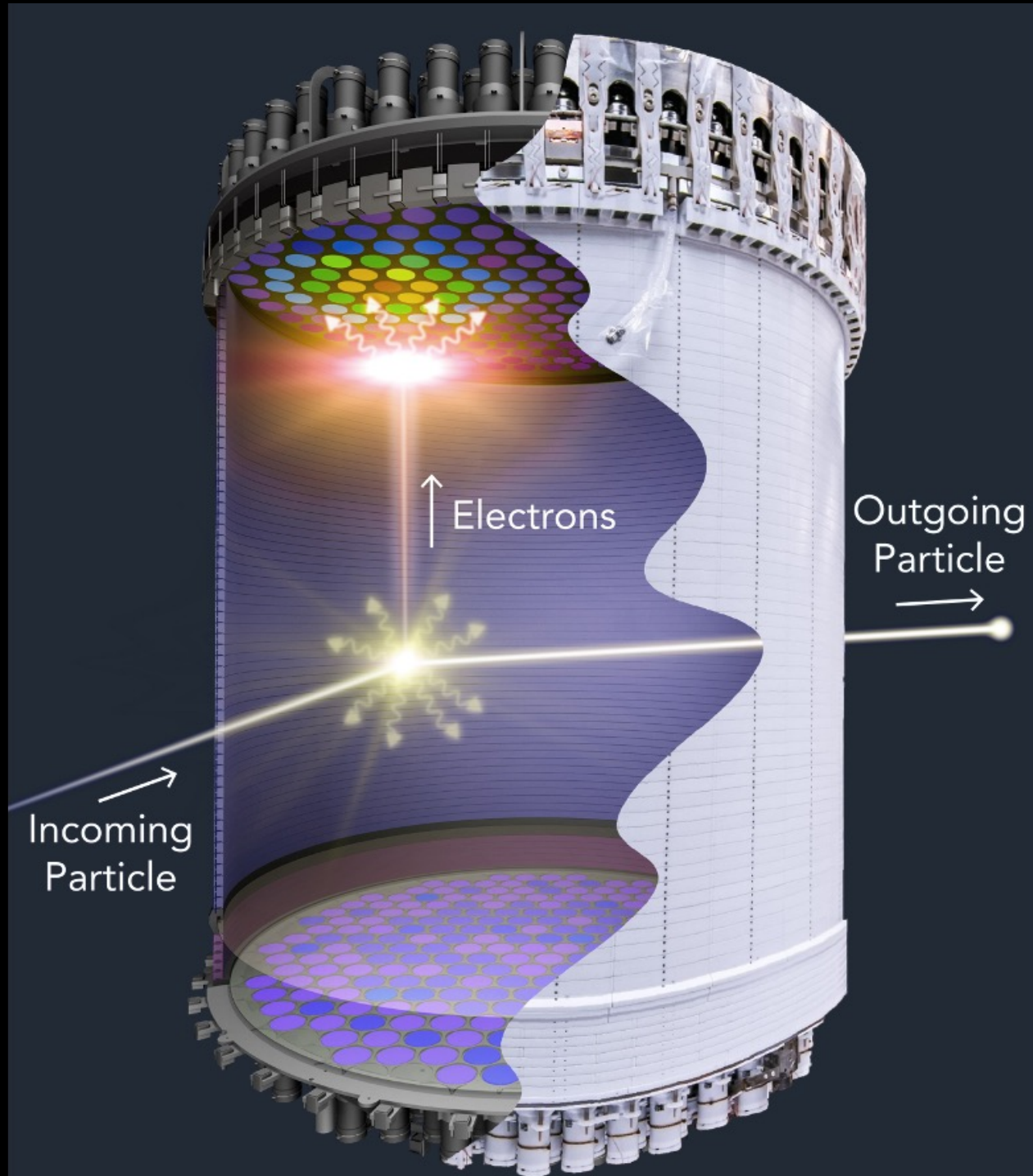


XMASS

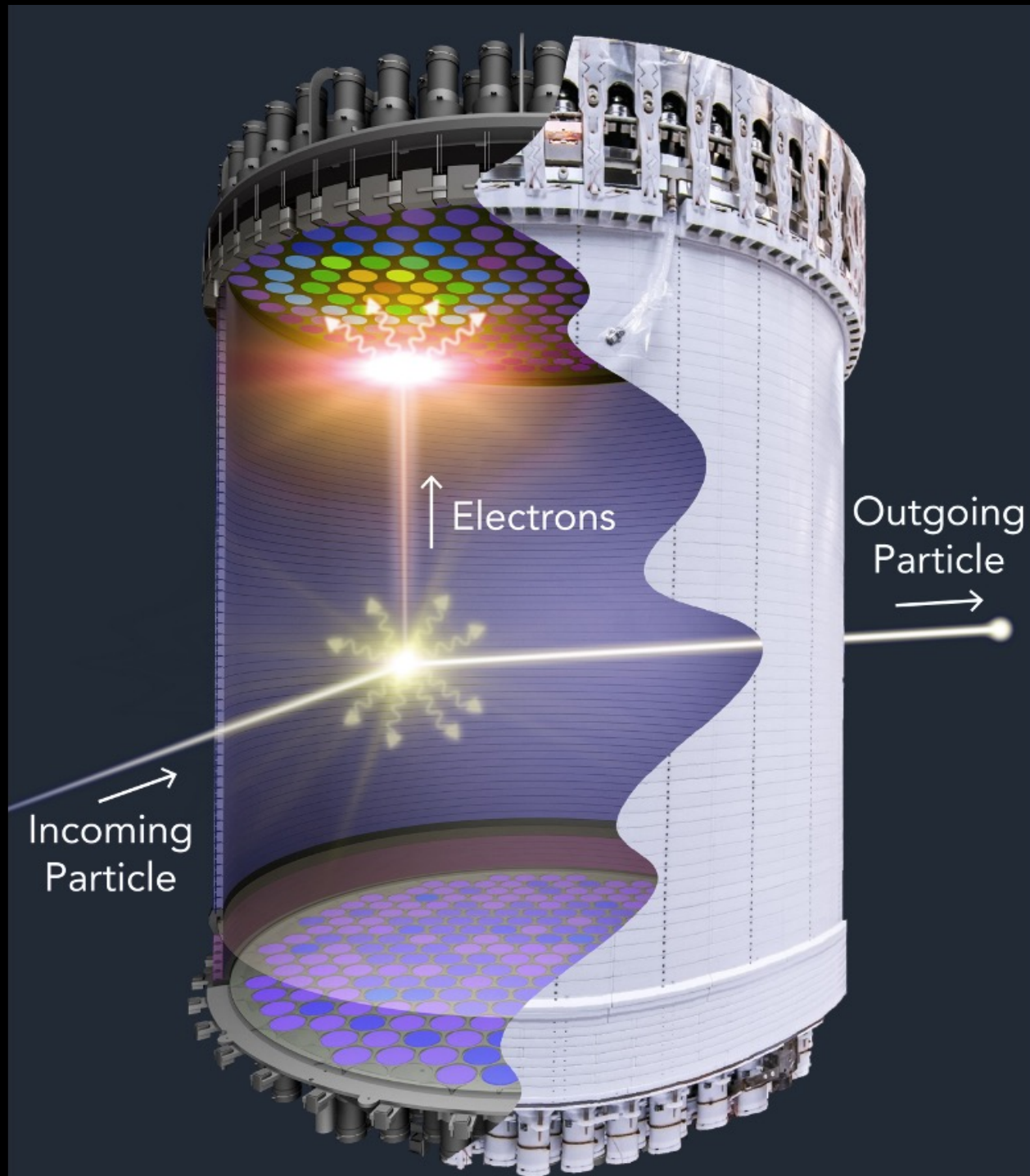


LUX/LZ
XENONnT
Panda-4X

(DUAL PHASE) NOBLE LIQUID TPC



(DUAL PHASE) NOBLE LIQUID TPC



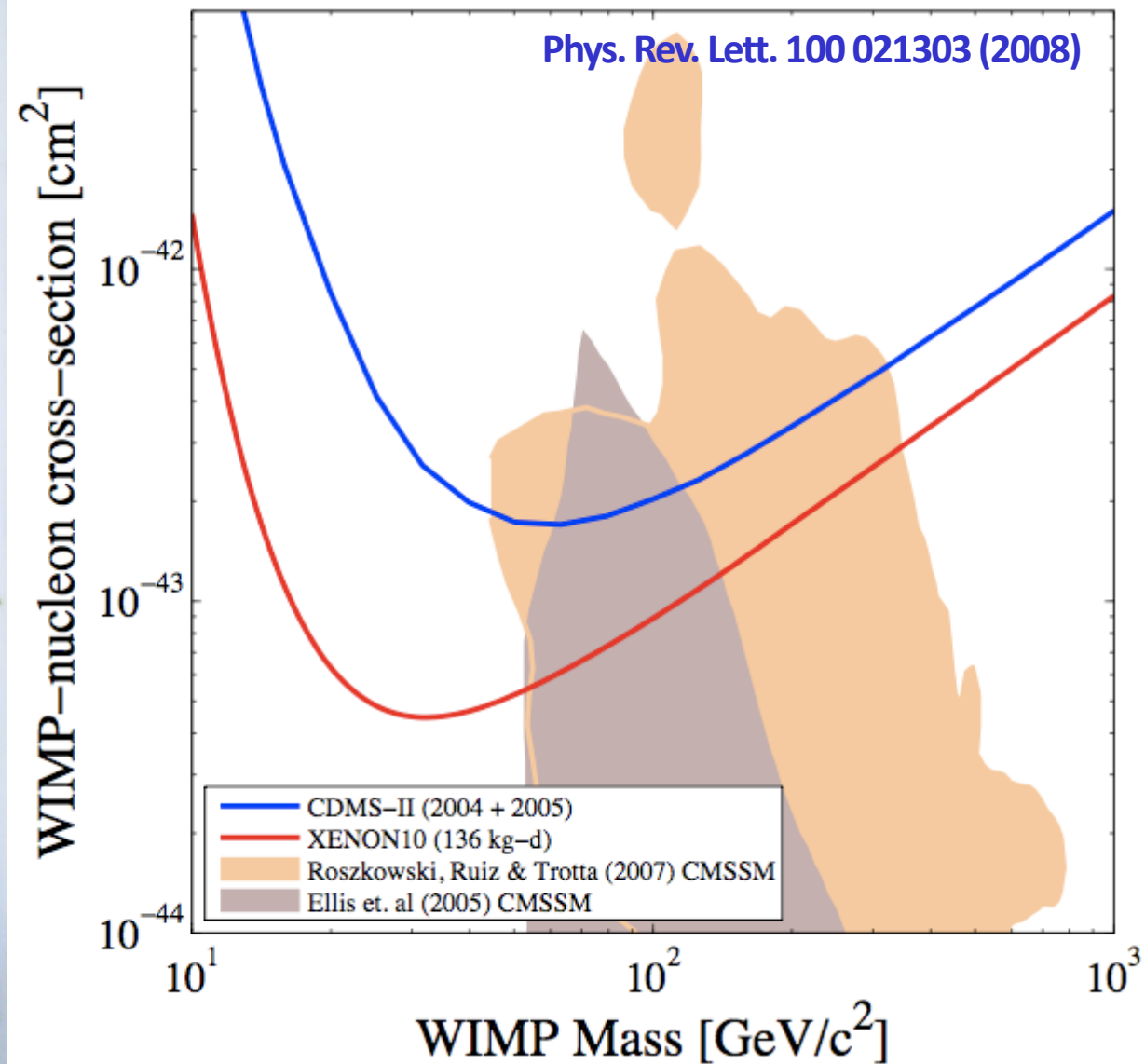
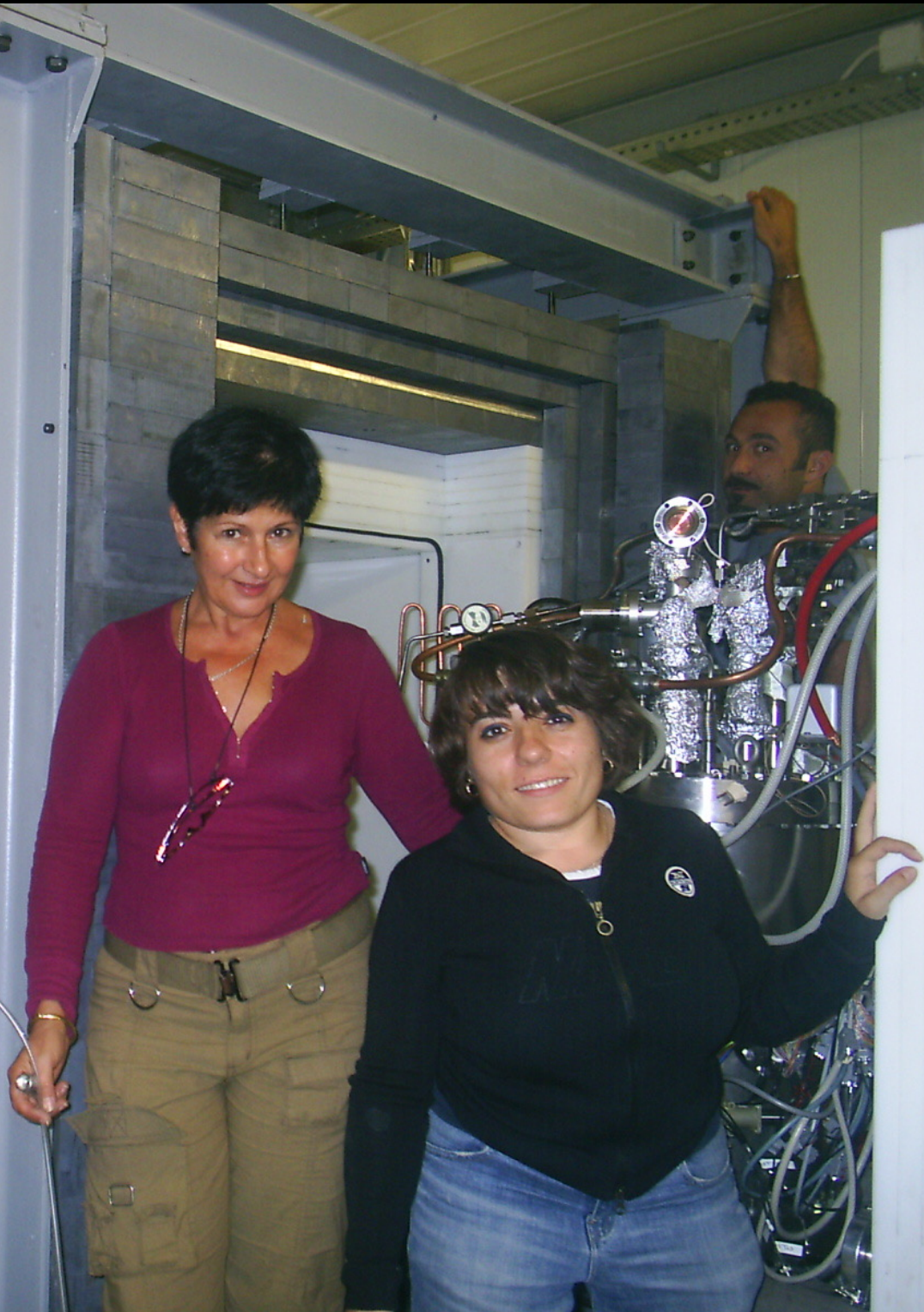
- WIMP-induced nuclear recoils: \sim few keV energy
 - S1, S2 \rightarrow event energy
 - S2 image \rightarrow xy coordinate
 - S1-S2 timing \rightarrow z coord.
 - S2/S1 (Xe) \rightarrow recoil type
 - S1 PSD (Ar) \rightarrow recoil type
- No long-lived isotopes (Xe)
- Self-shielding
- Recoil discrimination

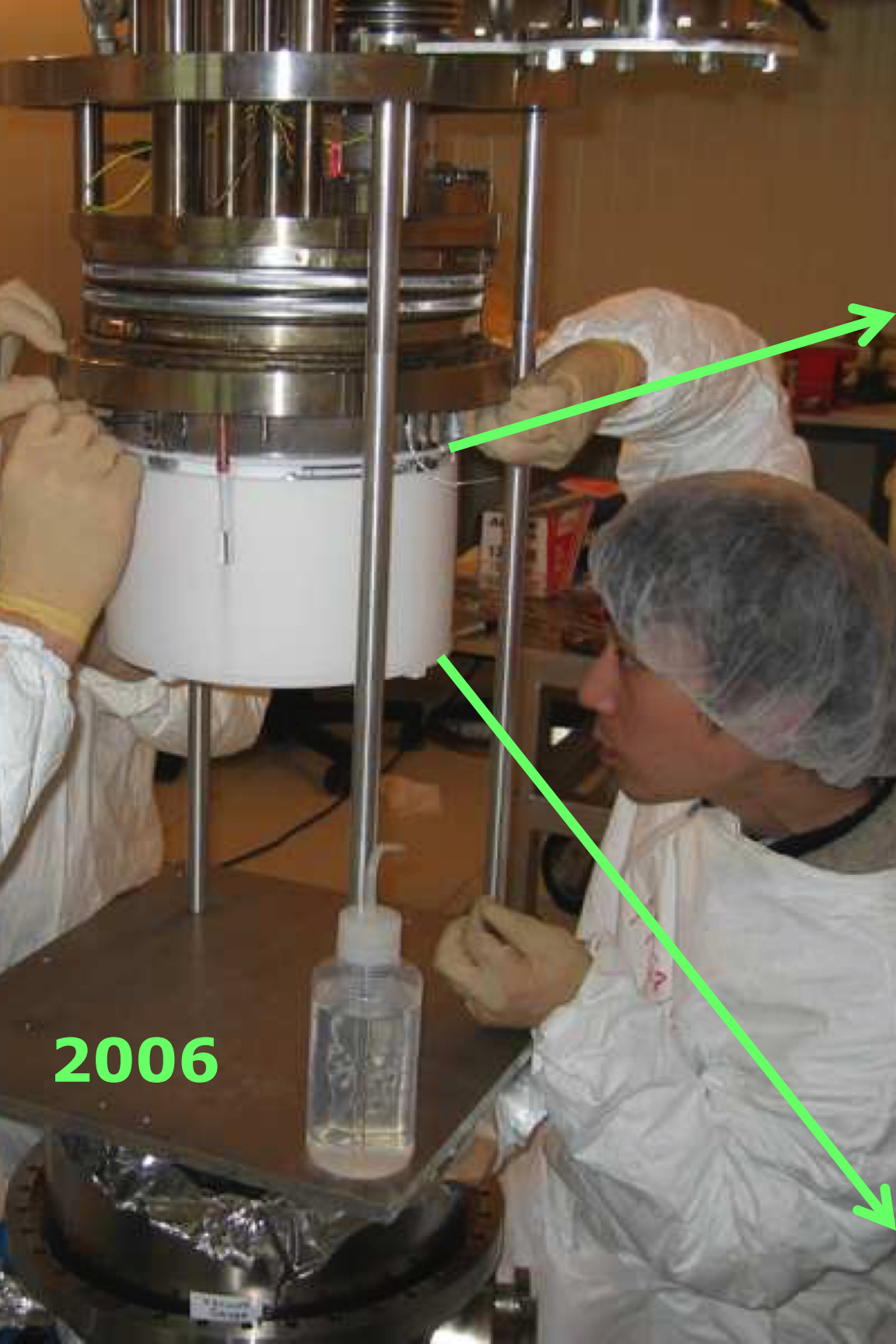
THE XENON10 EXPERIMENT AT GRAN SASSO



THE XENON10 EXPERIMENT AT GRAN SASSO

- Operated in 2006-2007
- 15 kg Liquid Xenon target
- $\sigma < 4.5 \cdot 10^{-44} \text{ cm}^2 @ 30 \text{ GeV}$



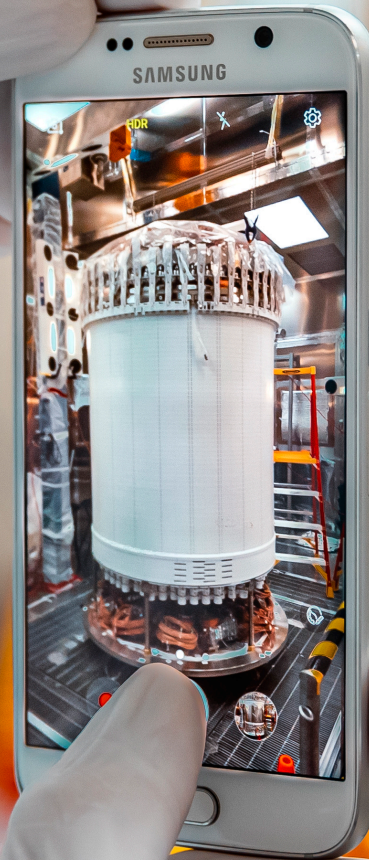


2006



2019

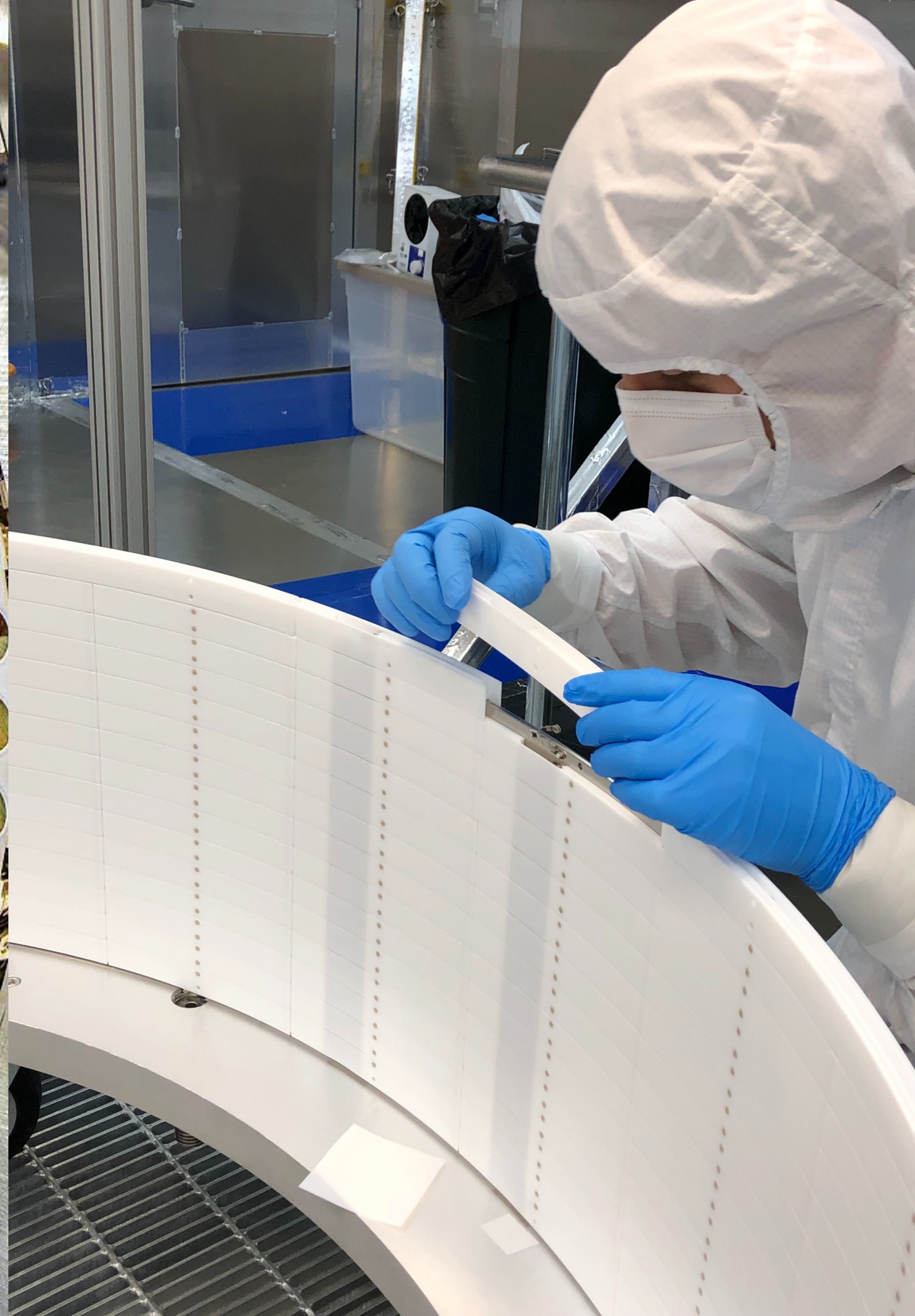
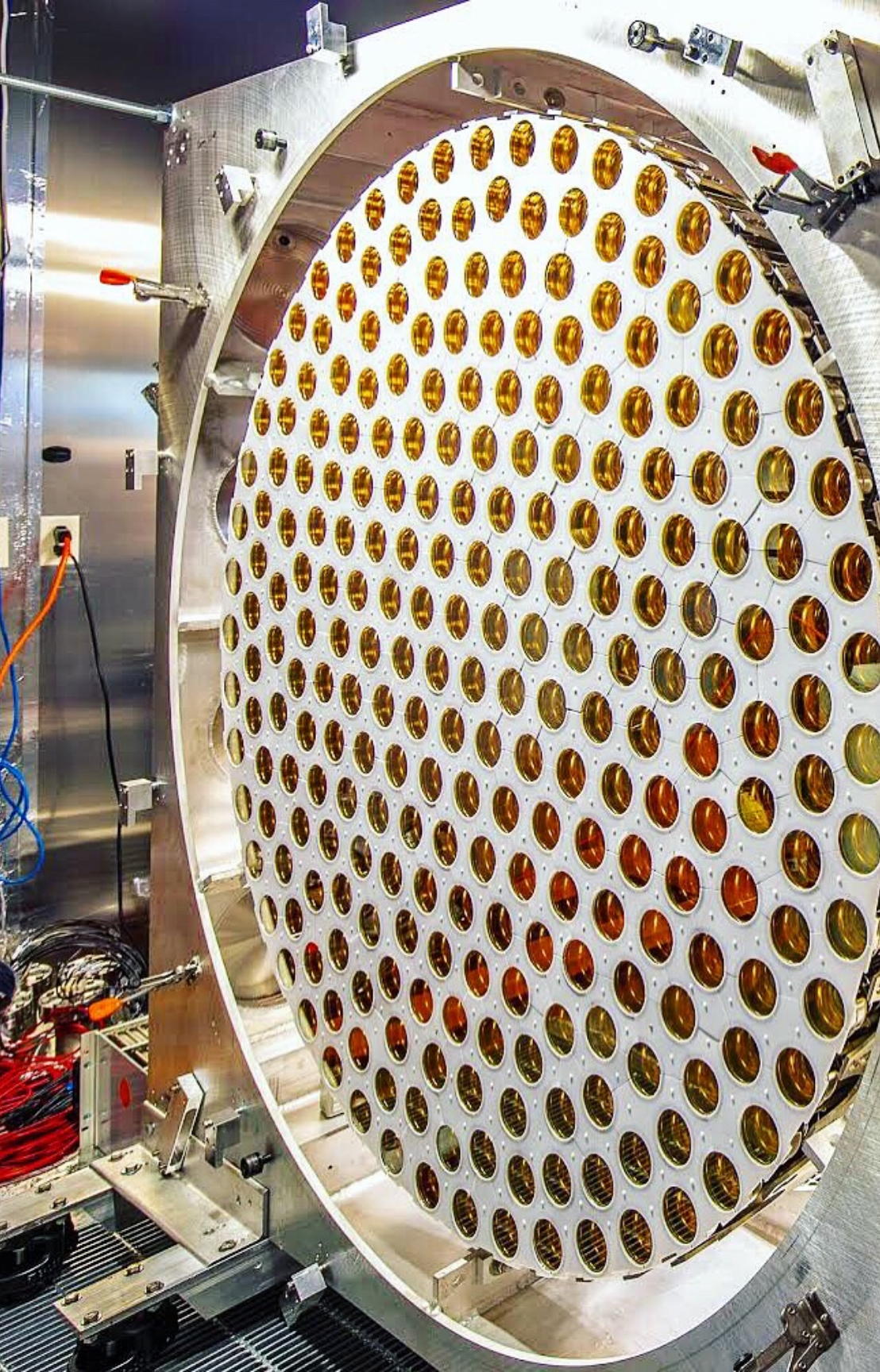
HOW TO BUILD A 10-TON LXE TPC

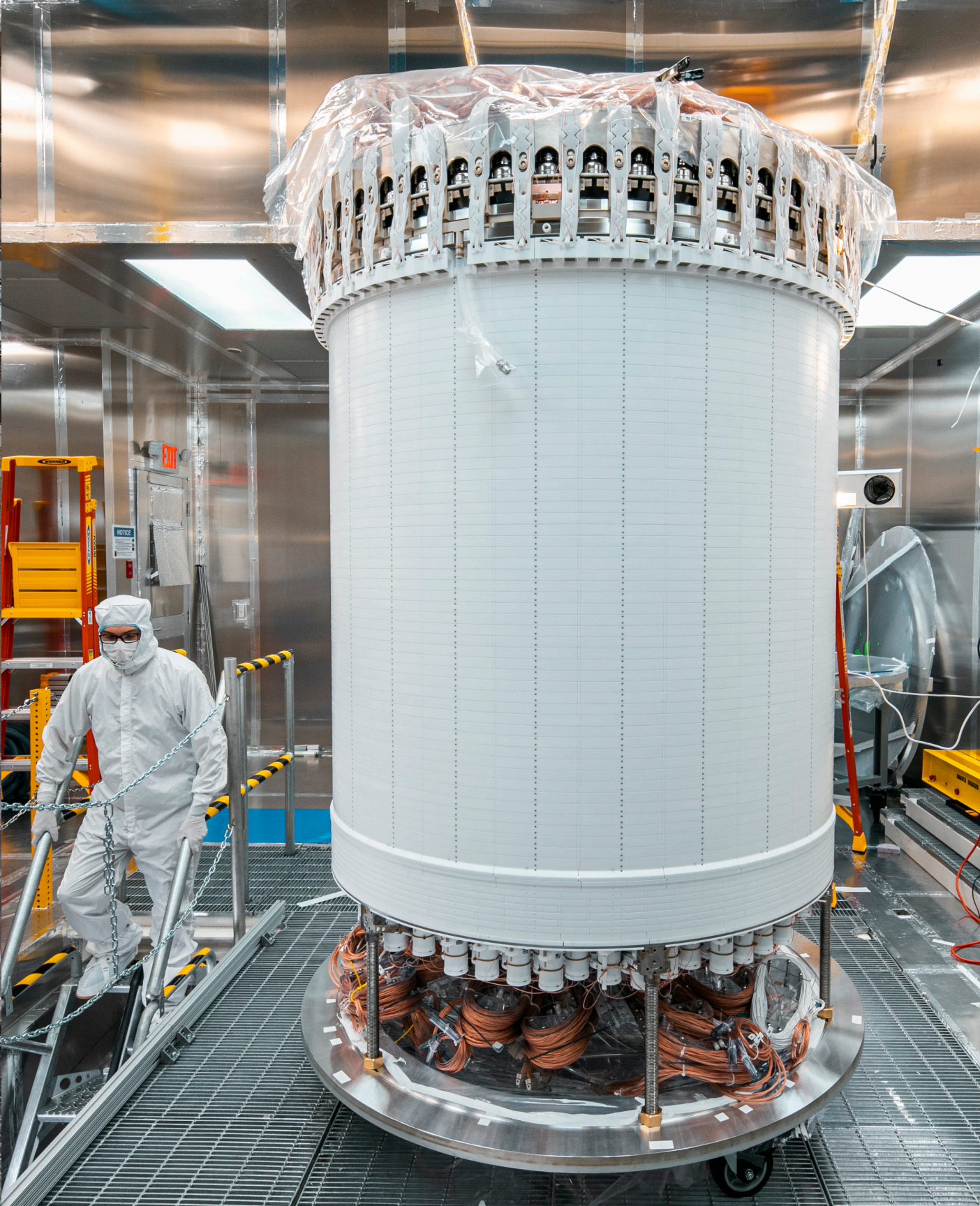




PHOTOSENSORS

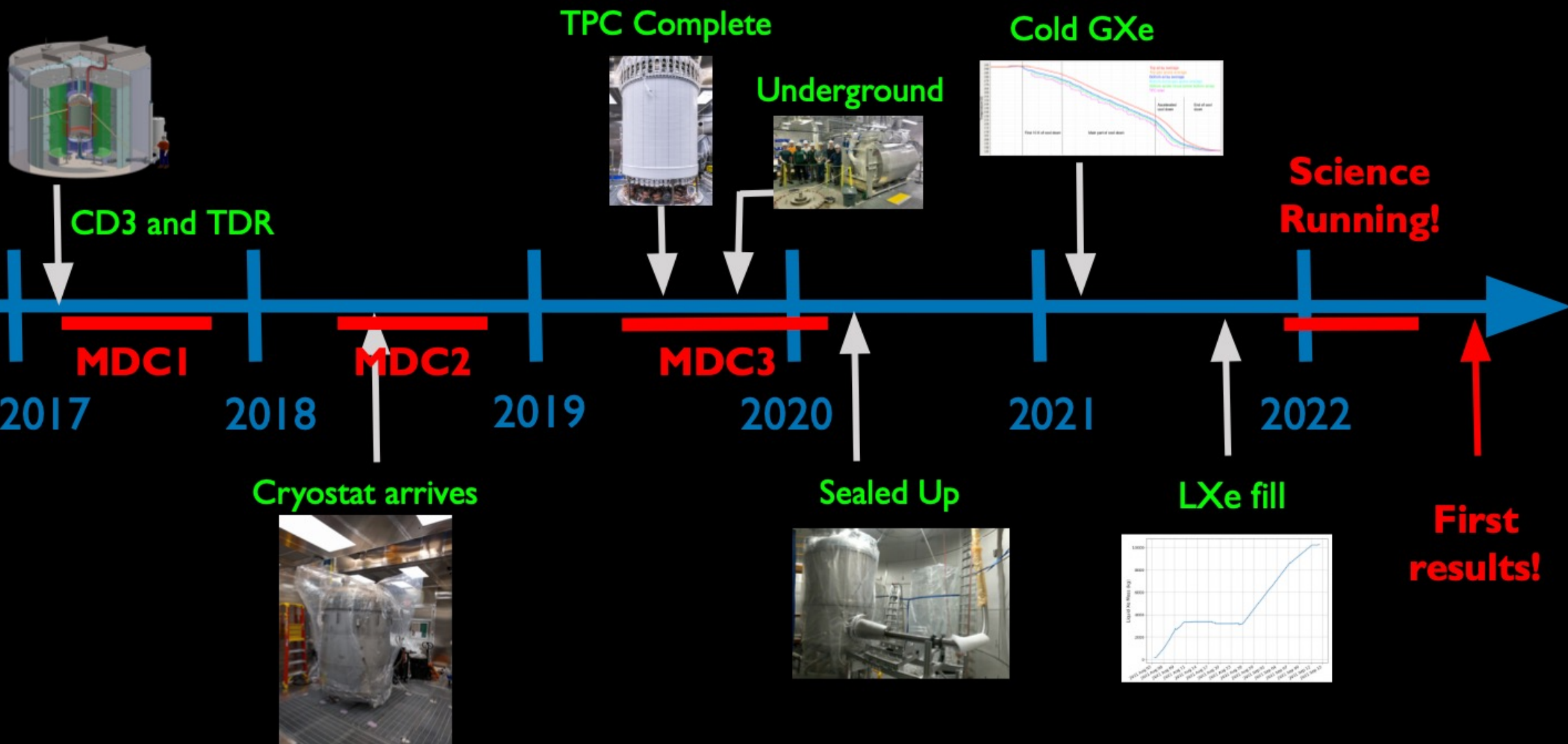








LZ CONSTRUCTION & DATA TAKING TIMELINE



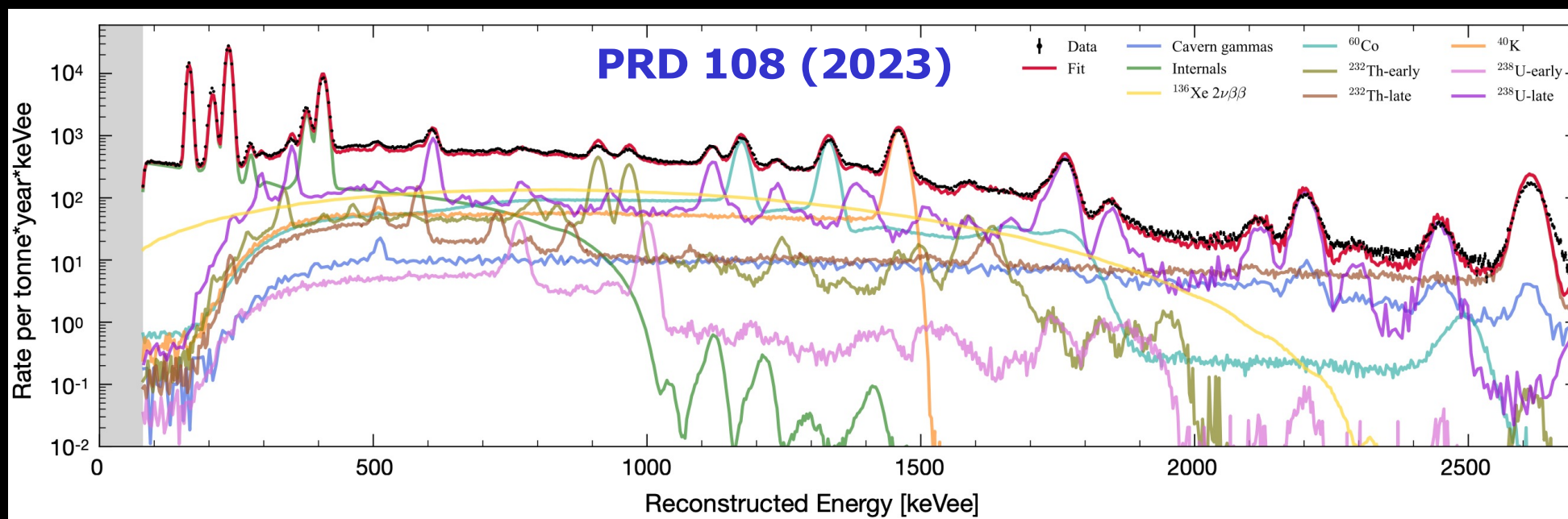
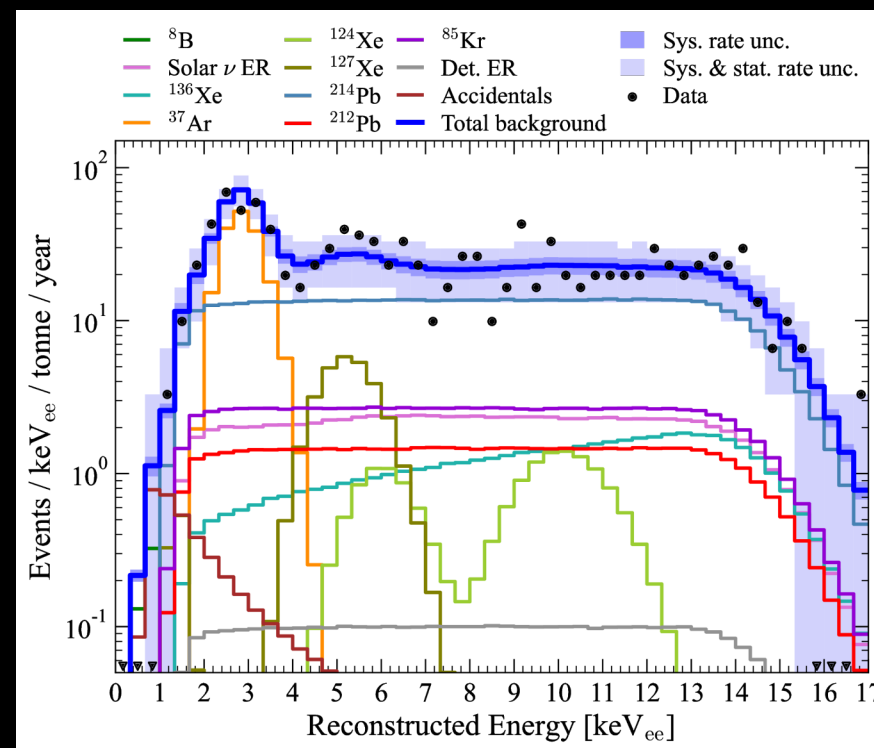
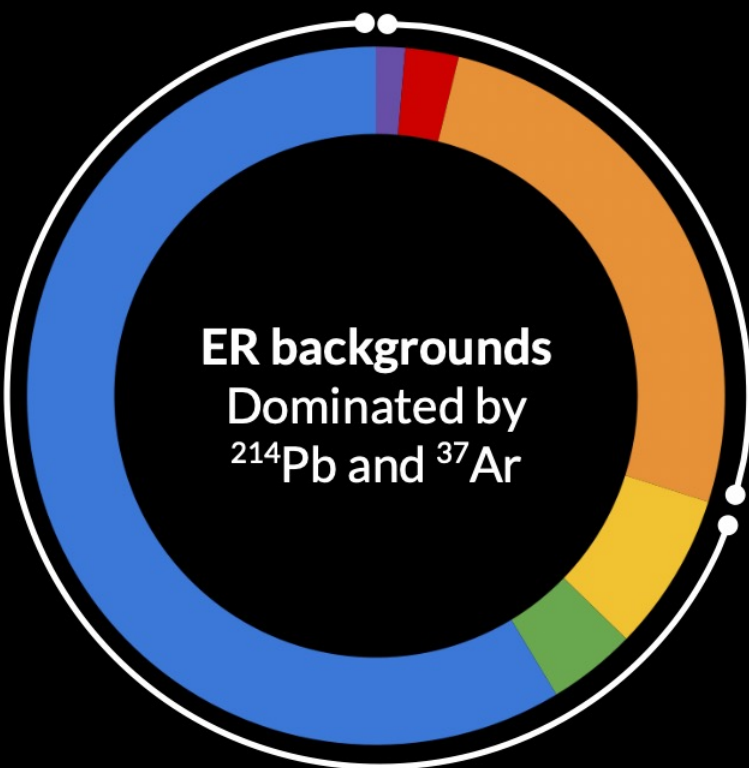
**ENABLED BY AN EXTENSIVE CAMPAIGN
OF MOCK DATA CHALLENGES (MDC)**

SEARCHING FOR WIMPs WITH SUPERCOMPUTERS

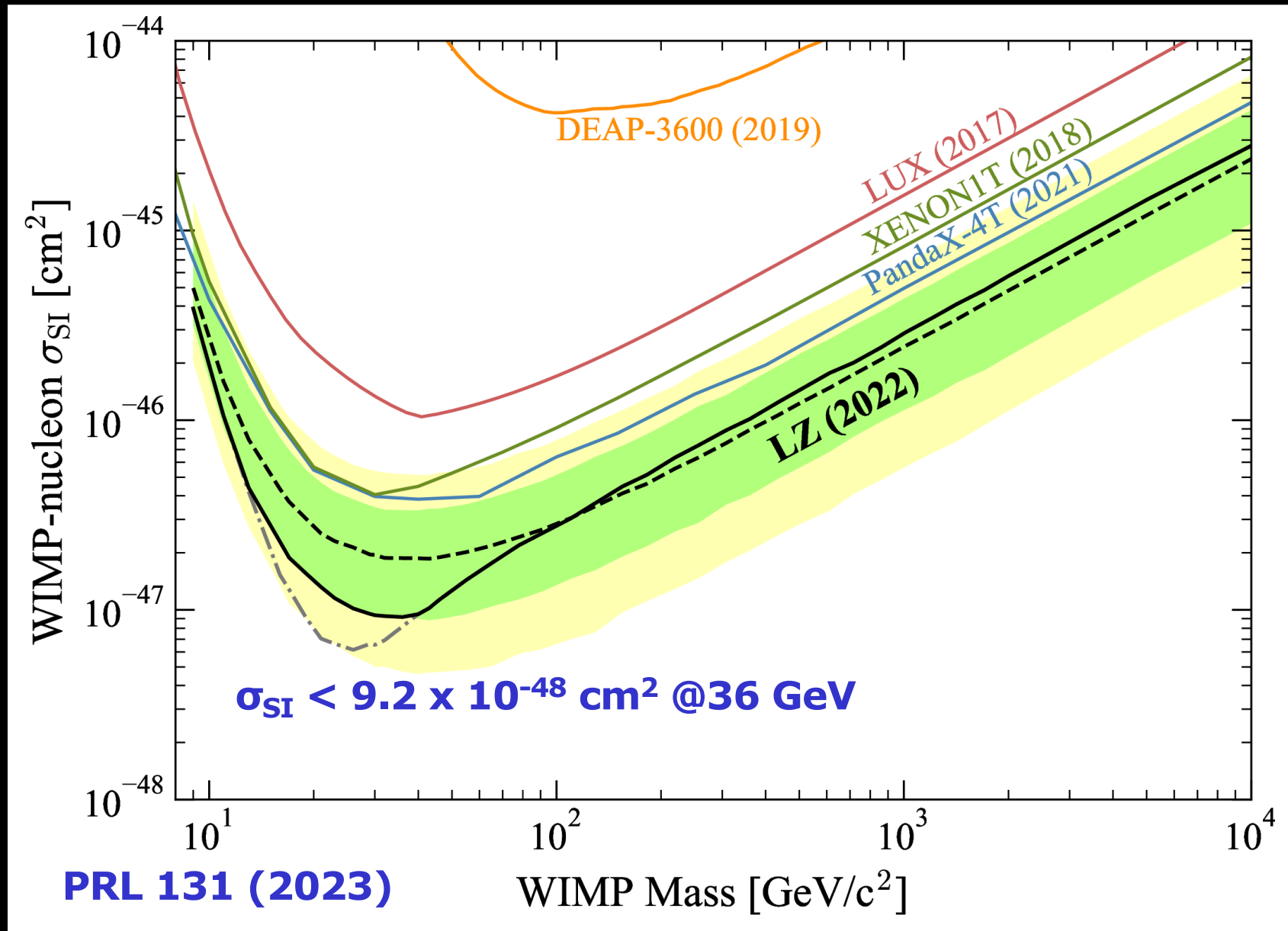
- **AUTOMATIC 24/7 PROCESSING OF ALL DETECTOR DATA**
- **LARGE-SCALE SIMULATIONS WITH DETAILED MODELING**
- **EXTENSIVE CAMPAIGN(x3) OF MOCK DATA CHALLENGES**



EXTREMELY DETAILED BACKGROUND MODEL

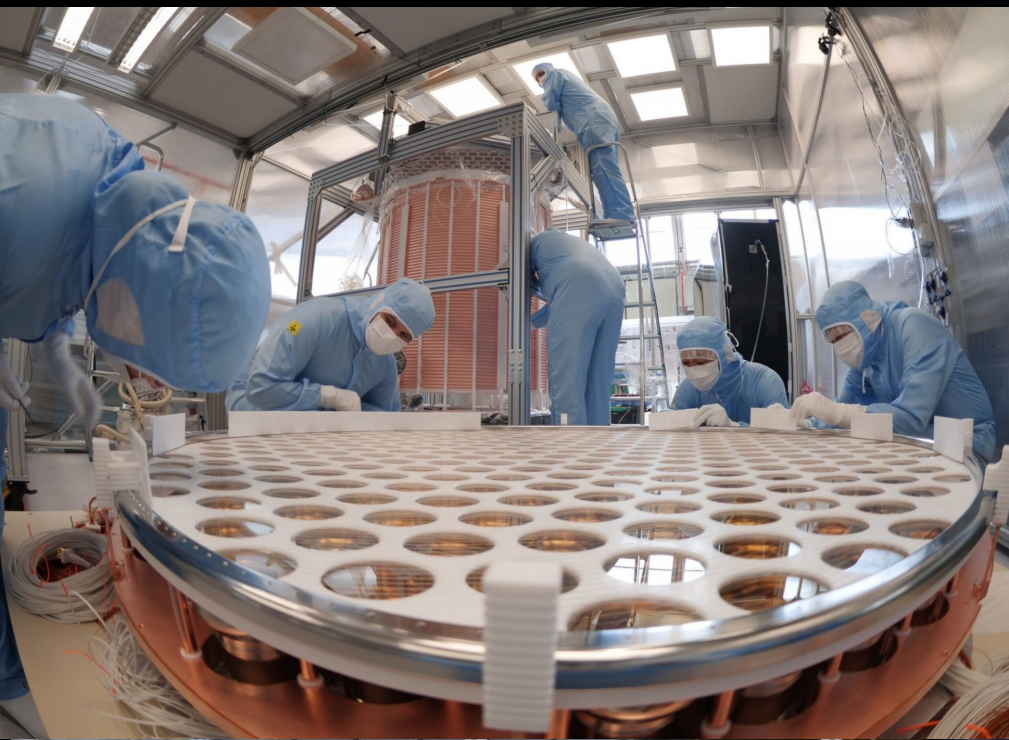


WORLD RECORD WIMP SENSITIVITY (SI, 2022)



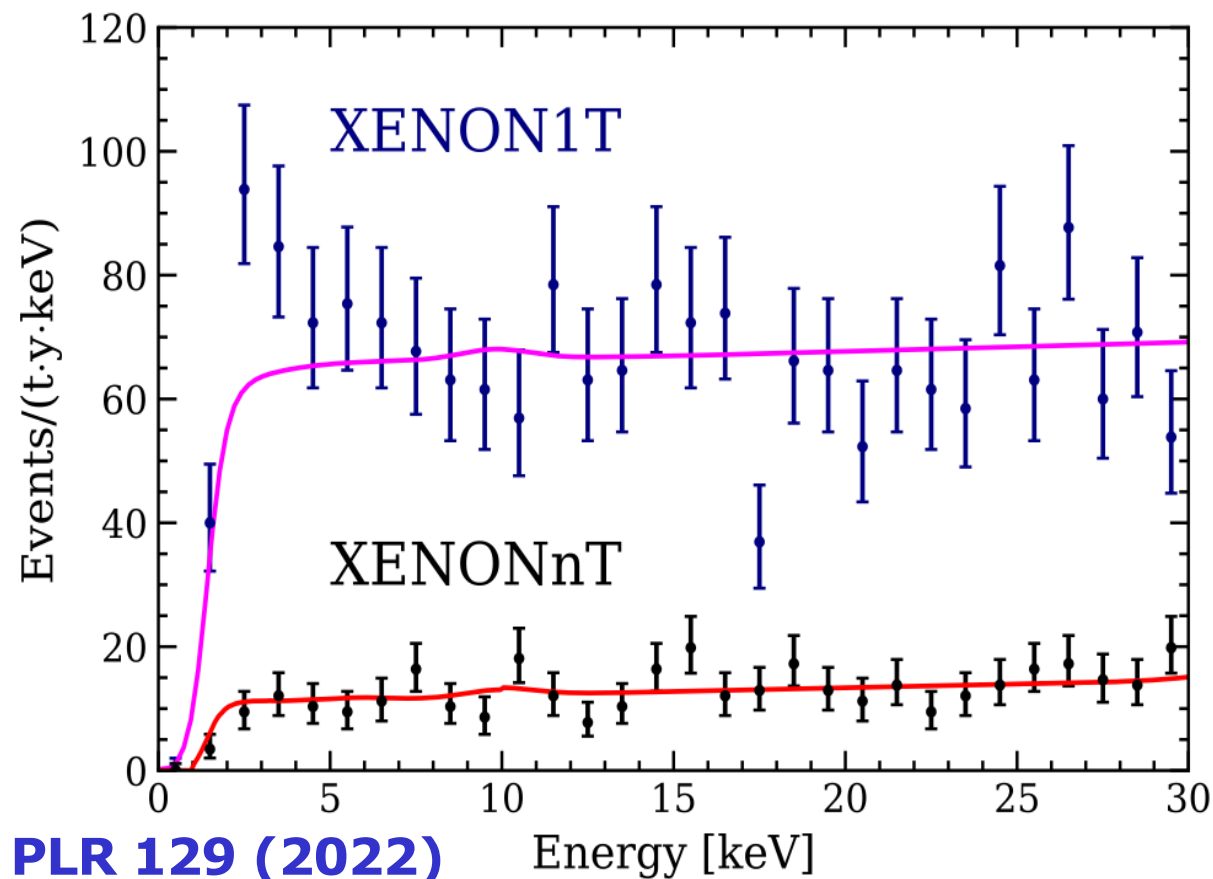
See also: Olcina, Uson, Dey, Green, Swain

XENONnT AT GRAN SASSO (2020+)

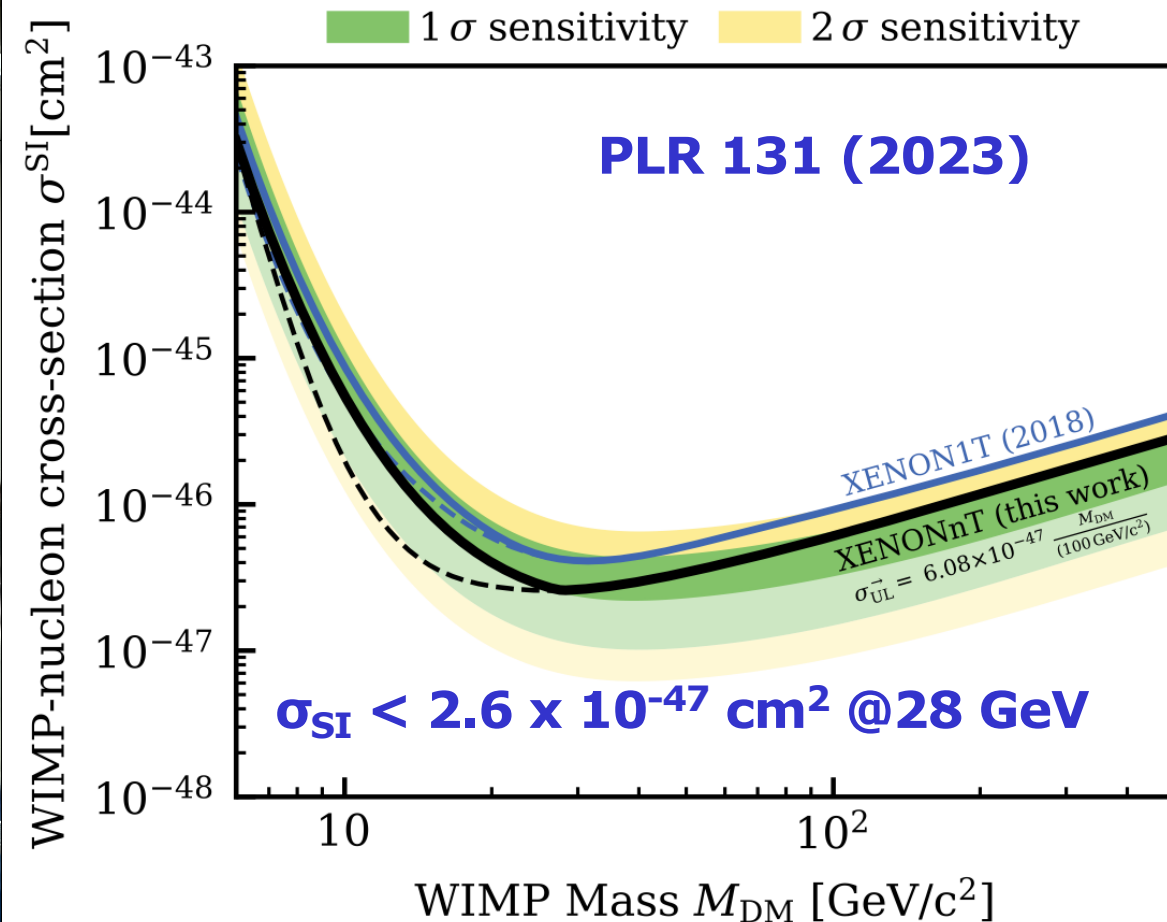
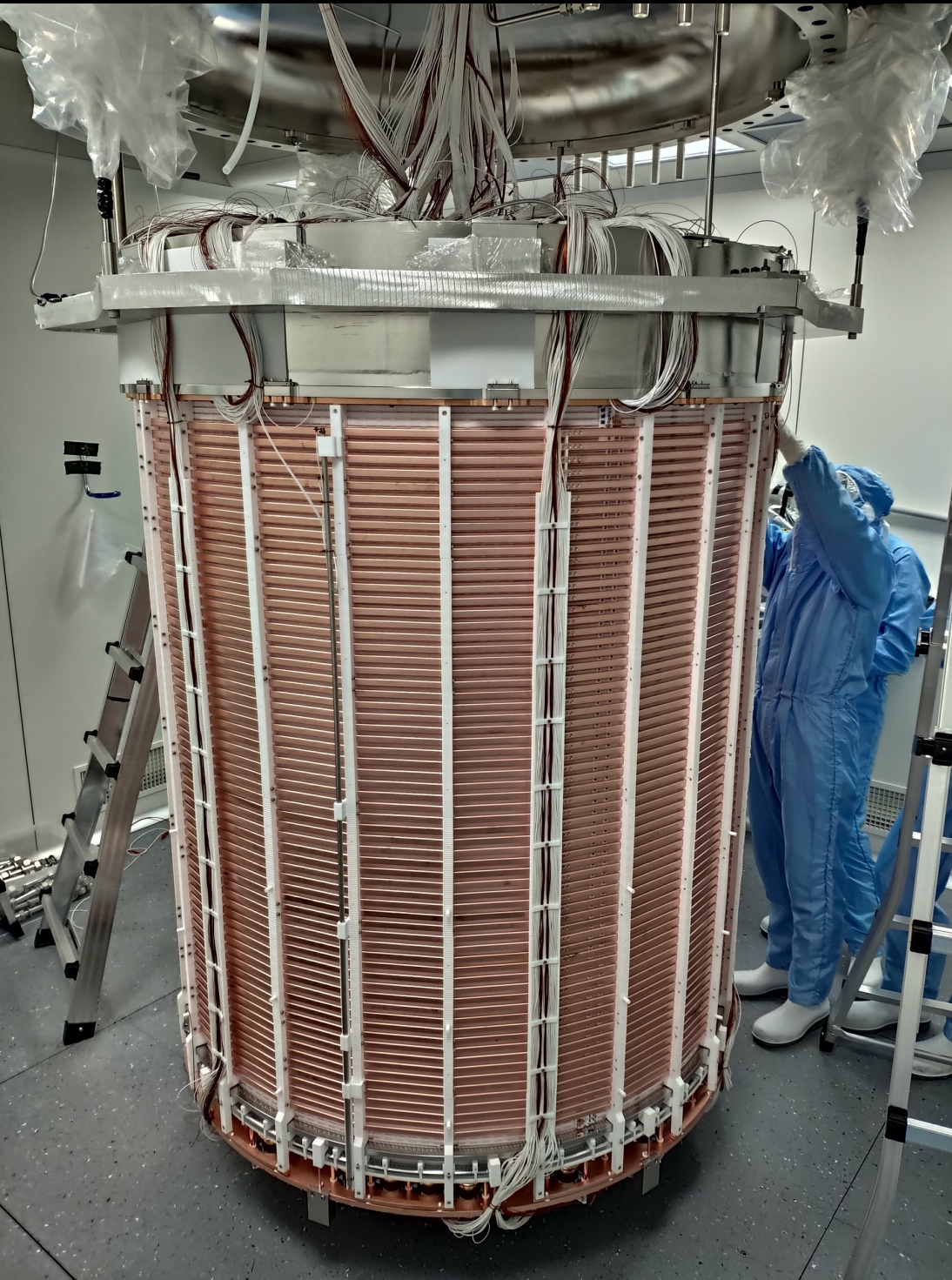


Initial results with 97.1 live days (2022)

- 6.9 tons active LXe volume
- 5x bckg reduction wrt Xenon-1T
- Unfortunately ER excess ruled out
- No axions this time around 🙄



XENONnT AT GRAN SASSO (2020+)

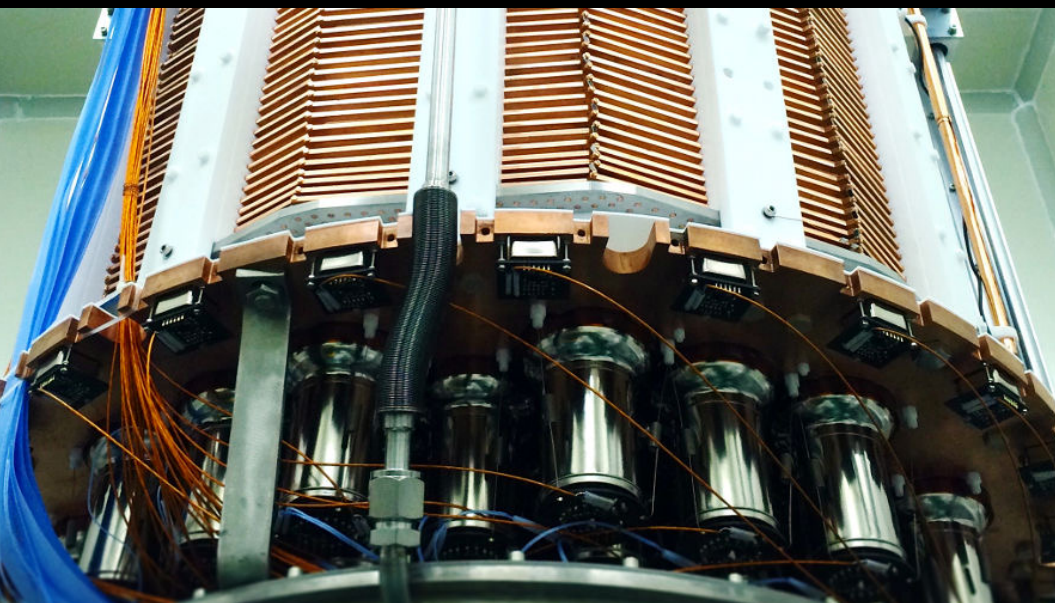



**See also: Eißing, Liu,
Selvi, Gao, Ferrari**

New results from XENONnT

Palazzo dell'Emiciclo, Sala Ipogea

PANDAX-4T AT JINPING LAB (2020+)



- Jinping Lab (China), 6800'
- Deepest underground lab on 
- 3.7 tons active LXe volume




2020/11
–
2021/04

Commissioning (Run 0)
95 days

2021/07
–
2021/10

Tritium removal
xenon distillation, gas flushing, etc

2021/11
–
2022/05

Physics run (Run 1) **NEW!!!** 
164 days

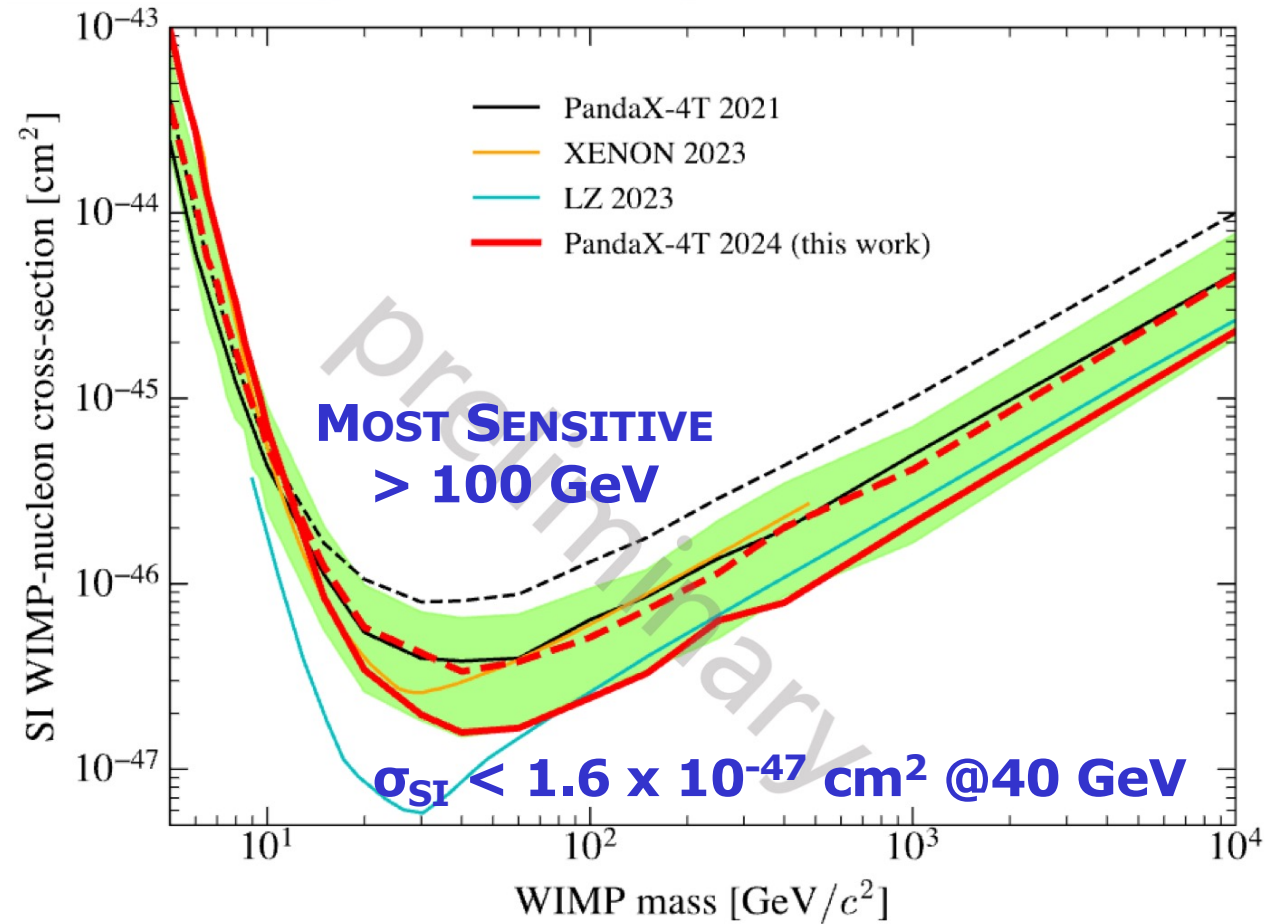
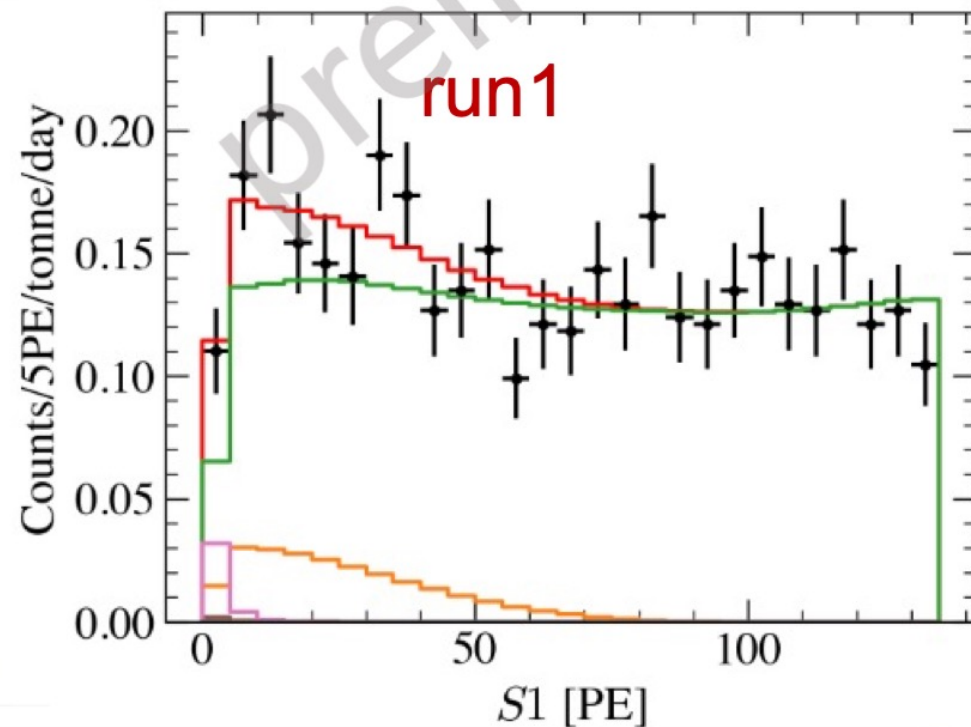
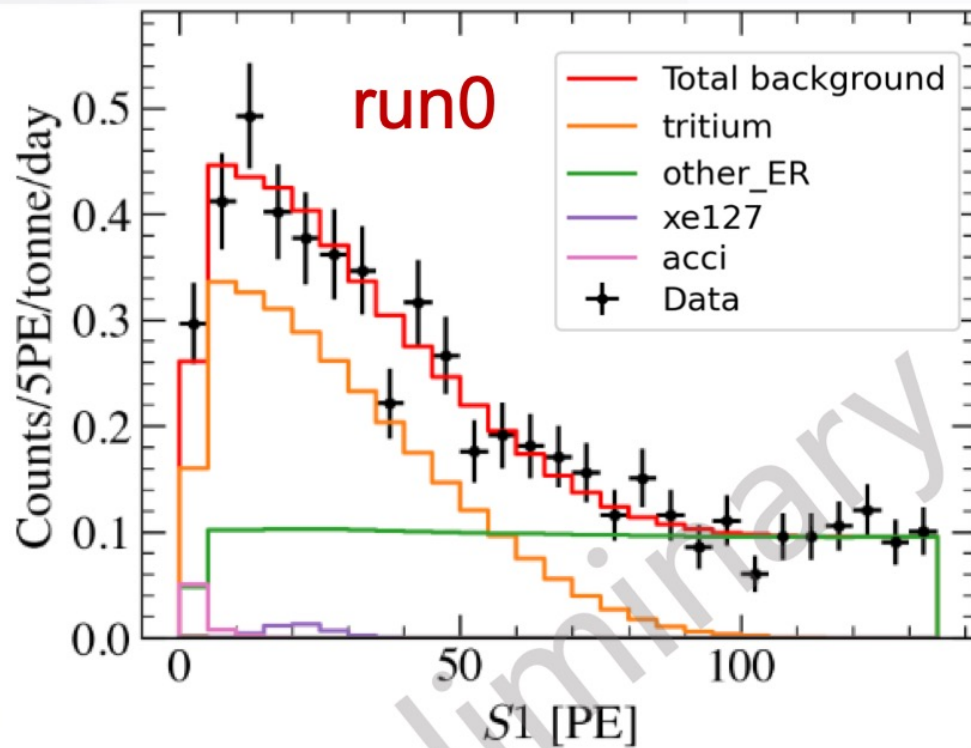
2022/09
–
2023/12

CJPL B2 hall renovation
xenon recuperation, detector upgrade

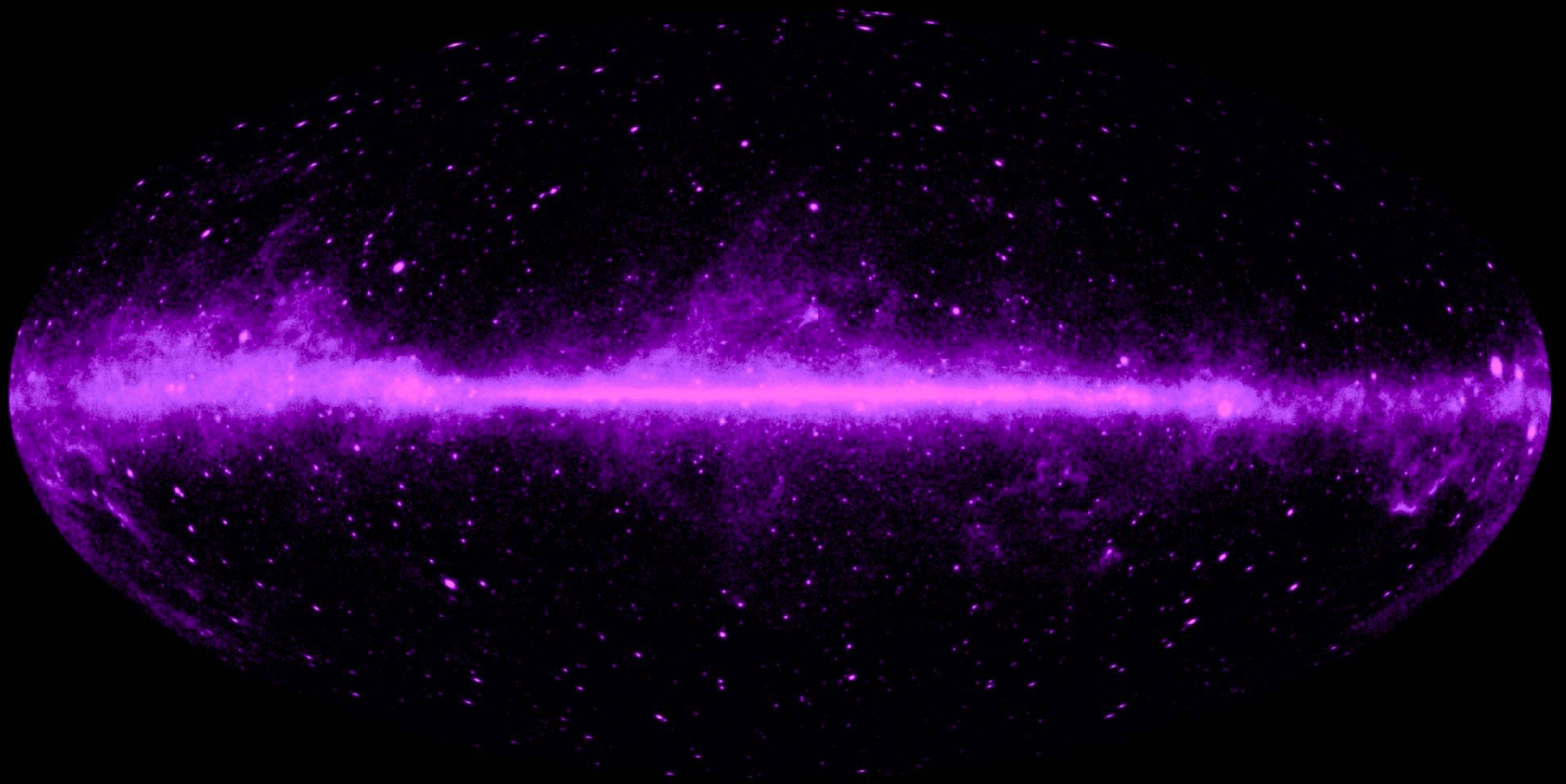
**Current
Status**

Resuming physics data-taking

PANDAX-4T AT JINPING LAB (2020+)



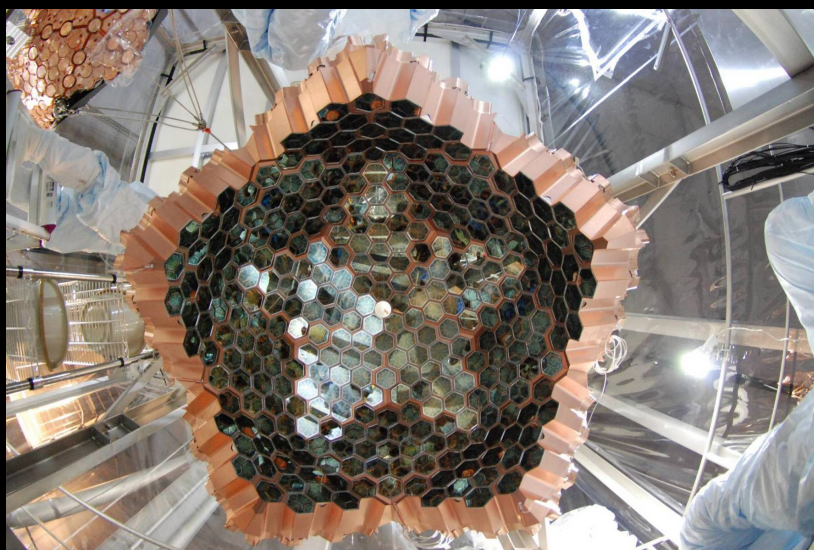
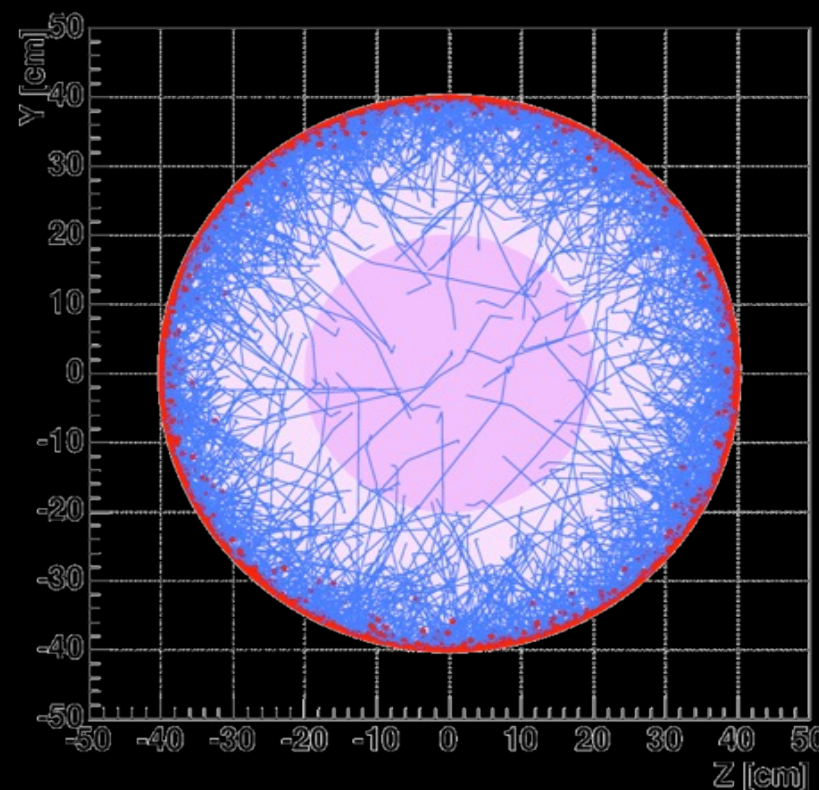
See also: Lin and Li



"OTHER" TECHNIQUES: SINGLE PHASE & LAR

SINGLE PHASE Xe: XMASS @KAMIOKA

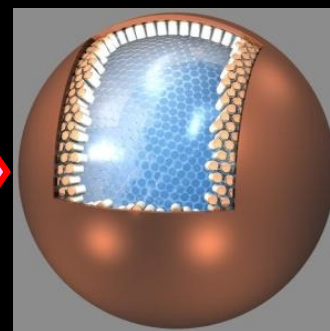
- XMASS-I: 835 kg / 642 PMTs
- Simple concept, good light collection
 - 14 pe/keVee
- Position reconstruction/self-shielding
- Surface Rn-backgrounds crucial
 - 10x lower than LUX, XENON-100
- 40 kg fiducial at 40 keV threshold



XMASS-I

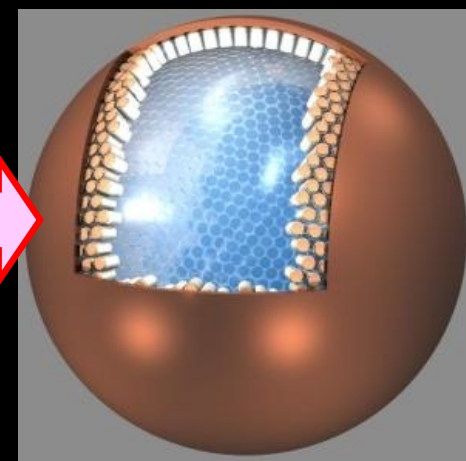


XMASS-I.5

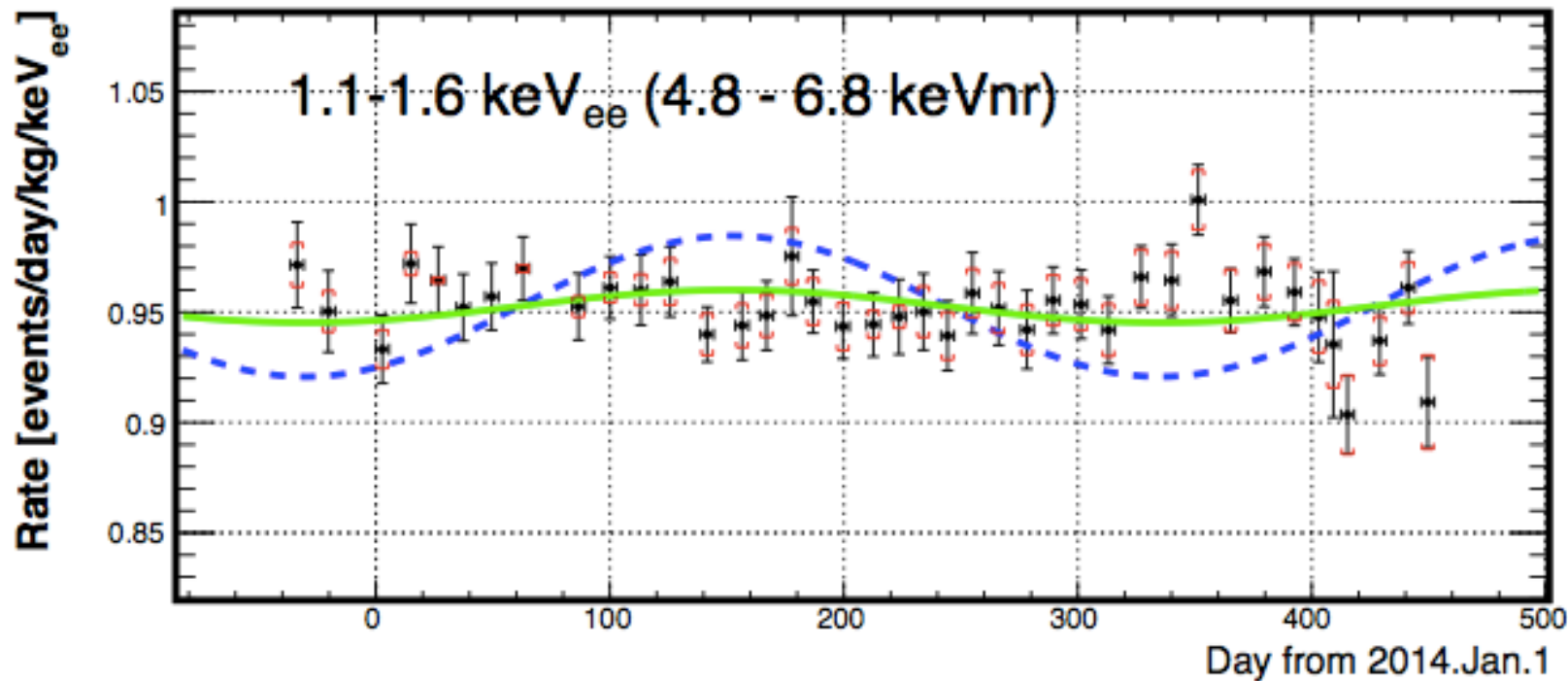


5 ton, 1 ton FV
 $\phi 1.5$ m, ~1k PMTs

XMASS-II



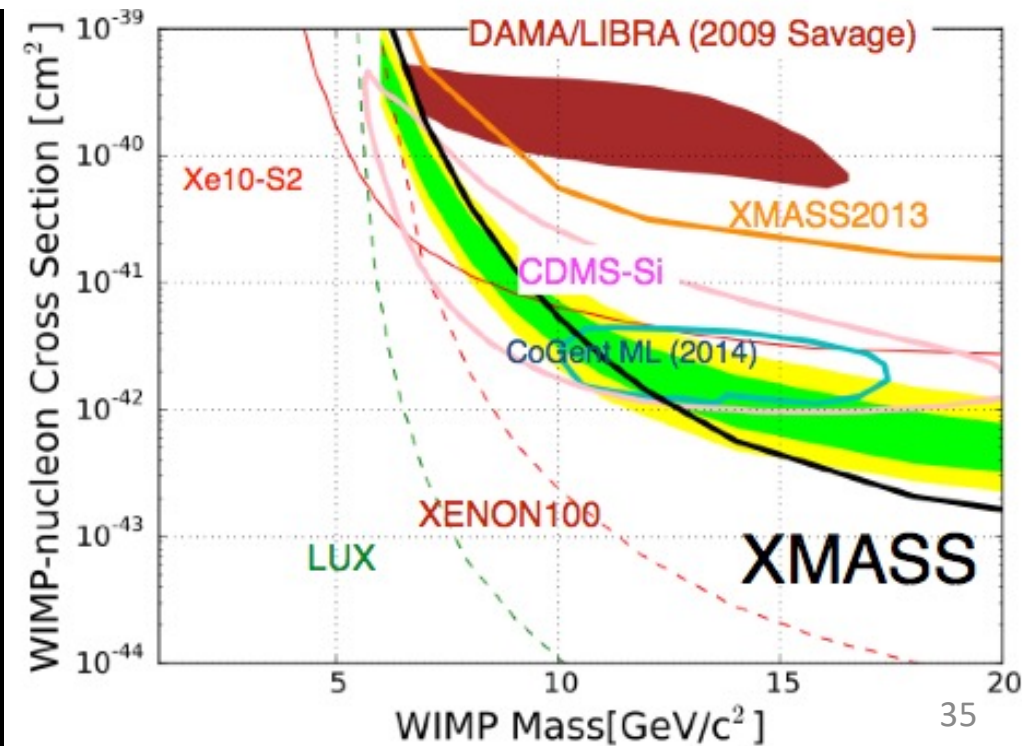
XMASS: ANNUAL MODULATION SEARCHES



- 7 GeV/c² x 10⁻⁴⁰ cm²
- 8 GeV/c² x 10⁻⁴⁰ cm²

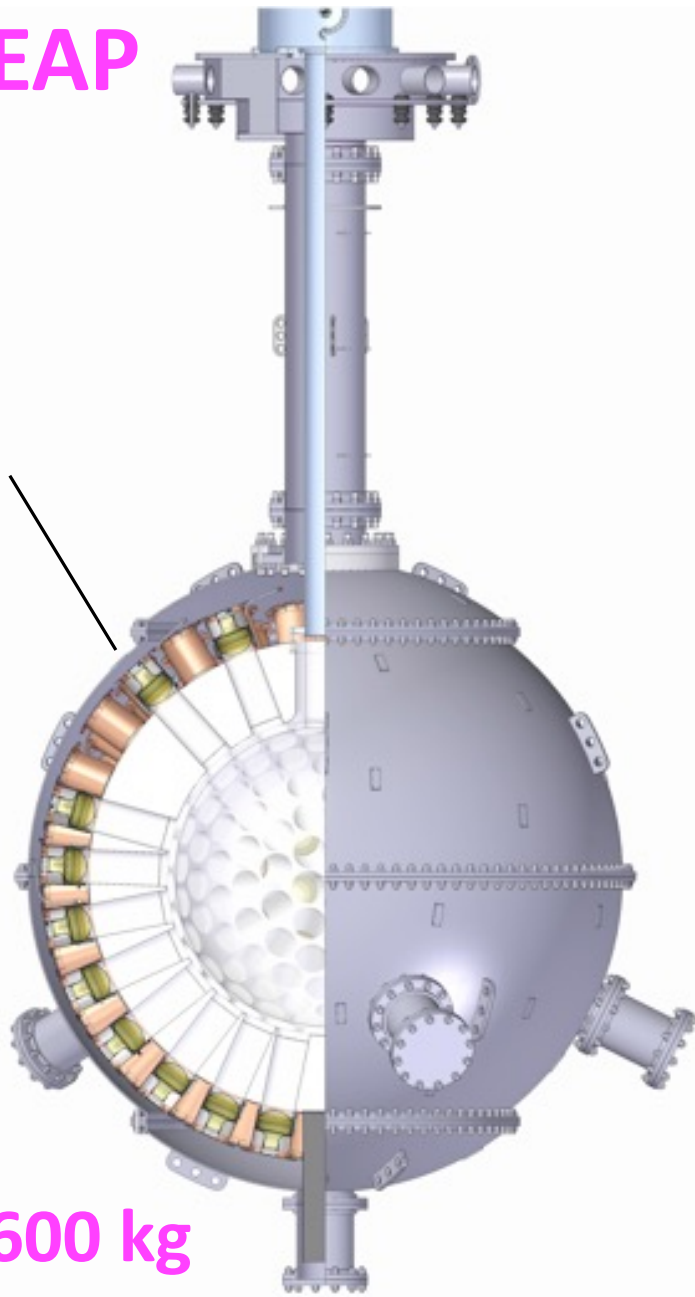
PhysLett B 759 (2016)

- Annual modulation, 2 WIMP cases
 - A² nuclear recoil signal model
- Full modulation analysis
 - fixed phase
 - limits consistent with previous Xenon-100 results



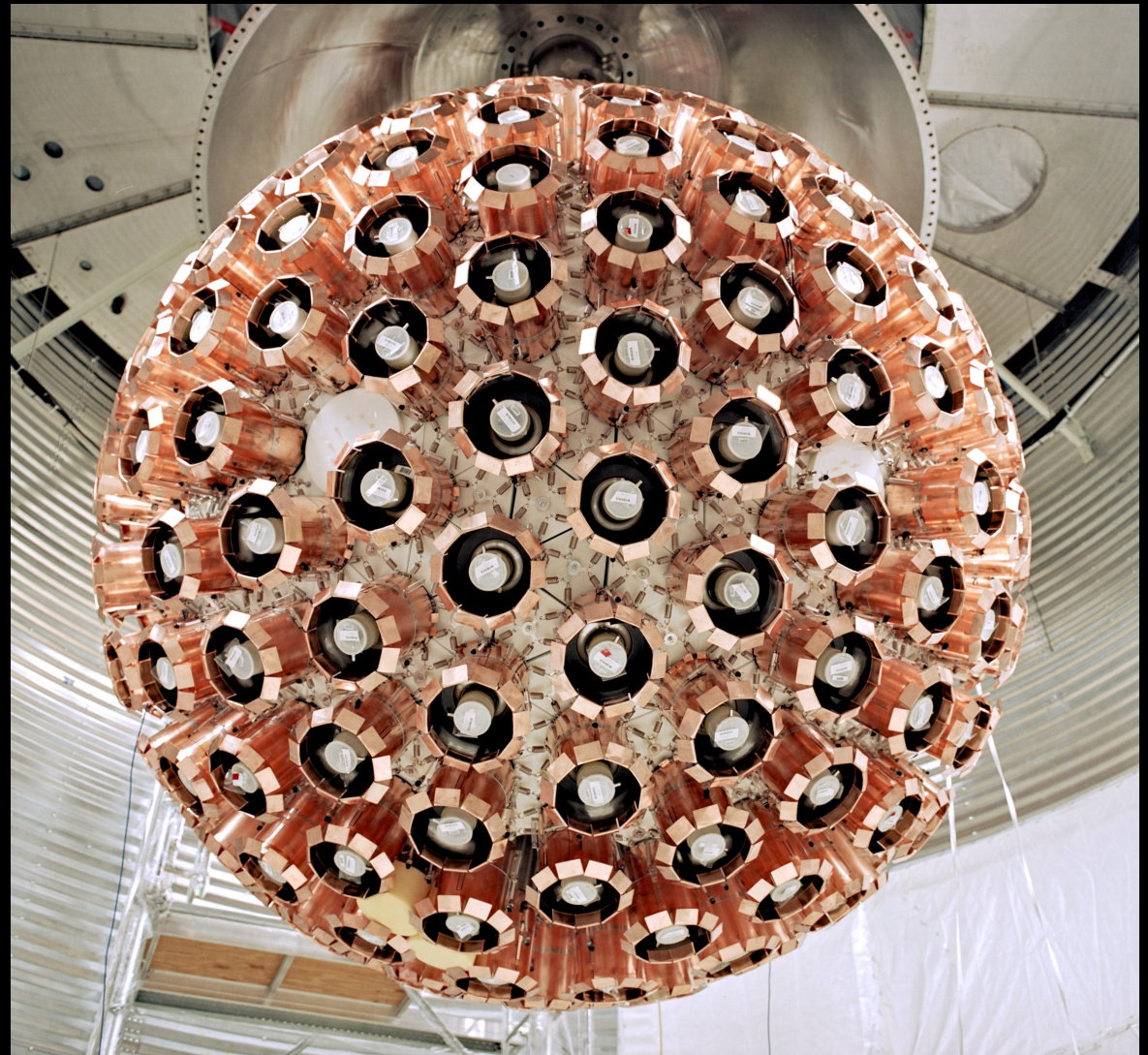
SINGLE PHASE LAR: DEAP @ SNOLAB

DEAP

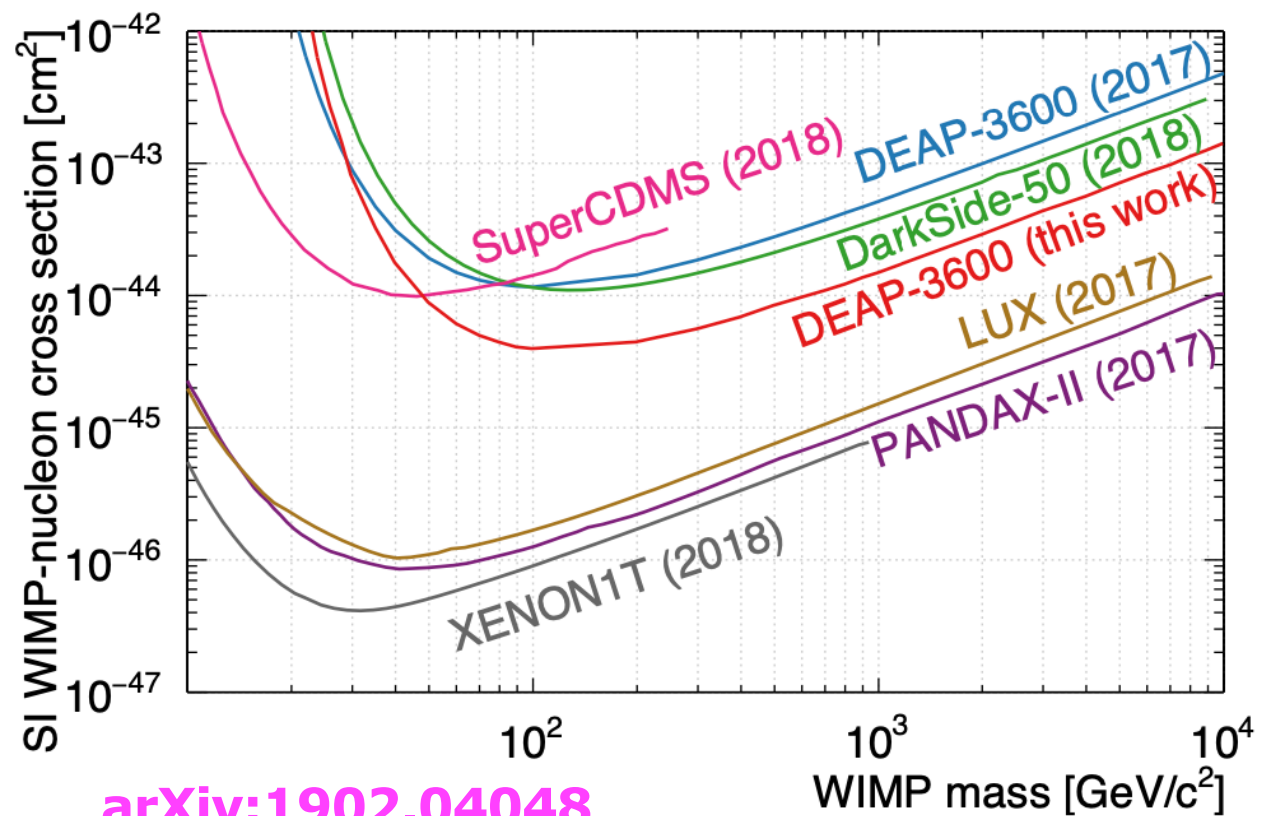
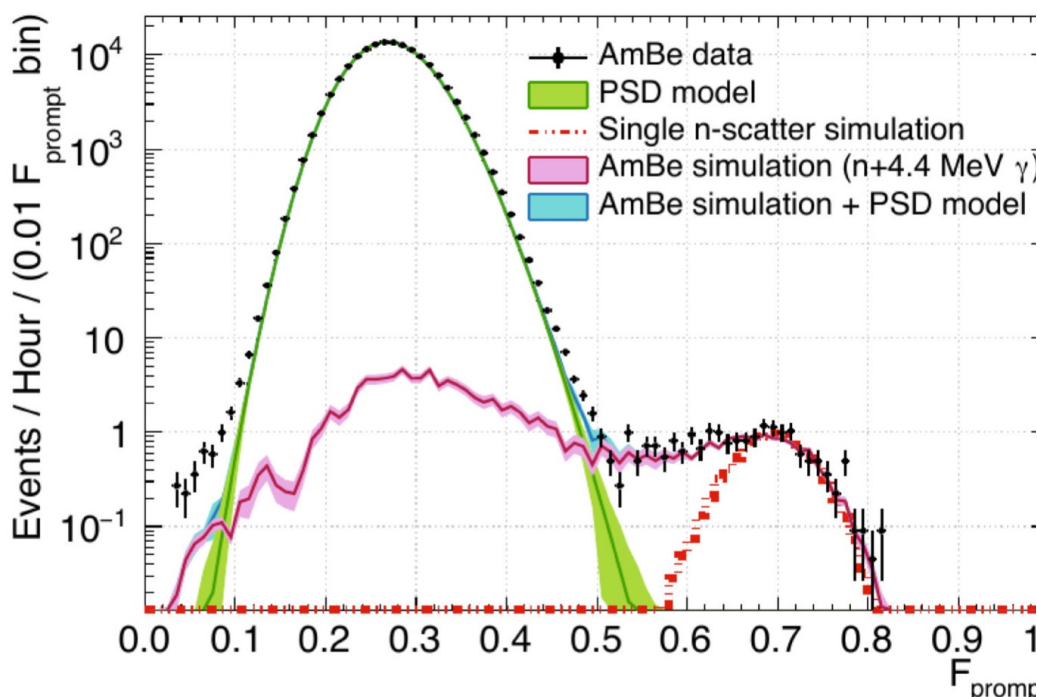
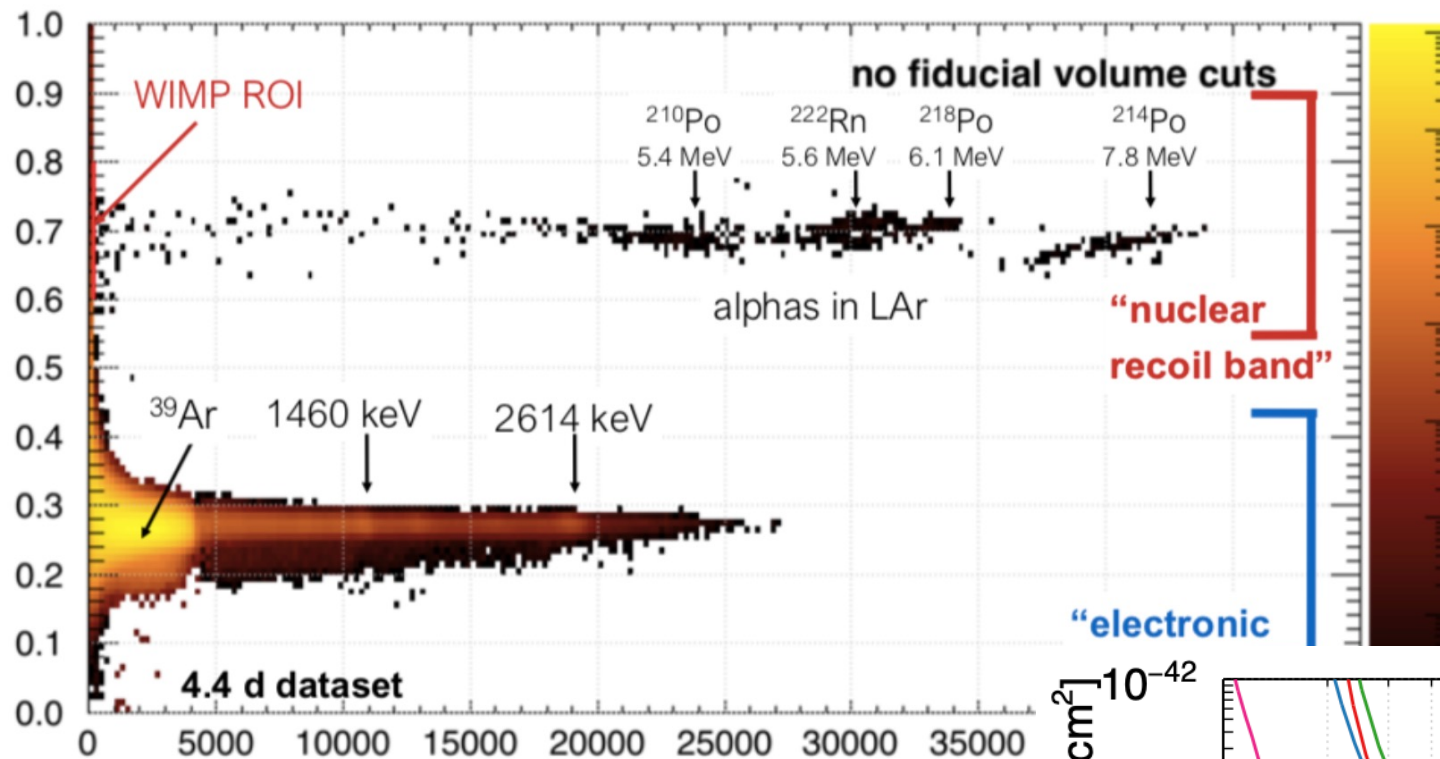


3600 kg
instrumented
1000 kg fiducial

- Significant pulse shape discrimination
- Ar is abundant (1% of Earth's atm.)
 - Intrinsic background: ^{39}Ar , ~600 keV endpoint beta (~1 Bq/kg atmosph. Ar)



SINGLE PHASE LAR: DEAP @ SNOLAB



Exceptional Pulse-Shape discrimination ($>10^8$)

(see P. Adhikari's talk)

- WIMP search: 231 live days, 3.3 t target
- $\sigma < 3.9 \times 10^{-45} \text{ cm}^2$ @100 GeV (2019)

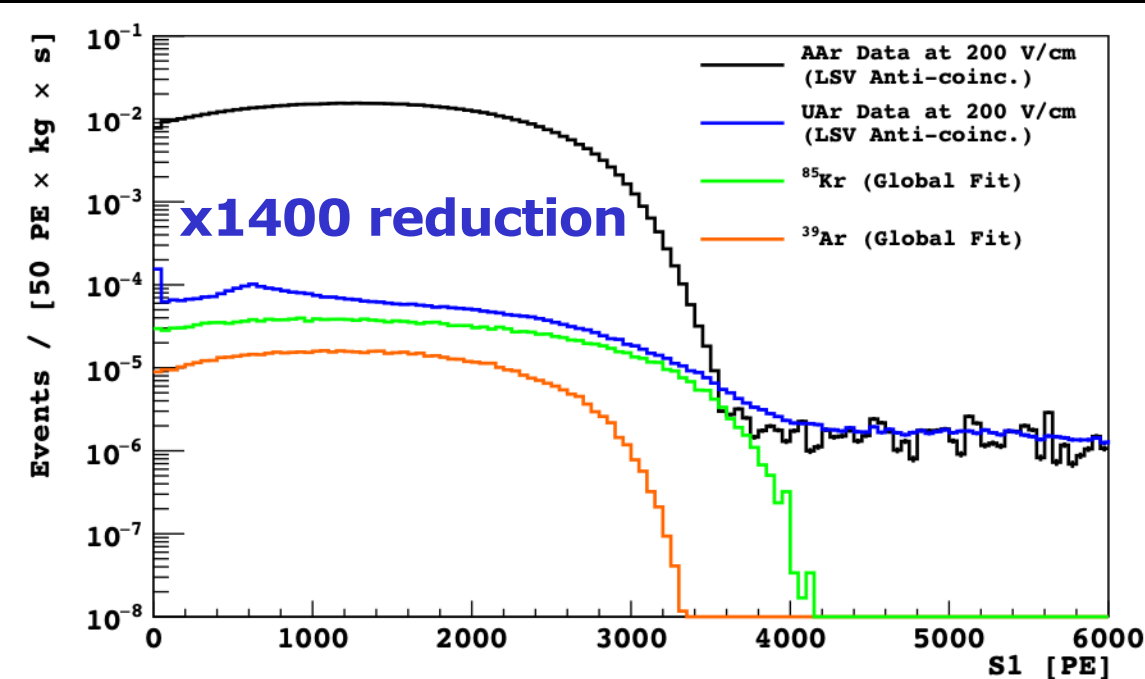
LIQUID ARGON TPC: DARKSIDE-50 @ LNGS



- 3-fold discrimination (S1 pulse shape, S2/S1, 3D-reconstruction)
- Underground Ar, depleted in ^{39}Ar
- x1400 ^{39}Ar depletion observed
- UAr WIMP search in 2015-2107



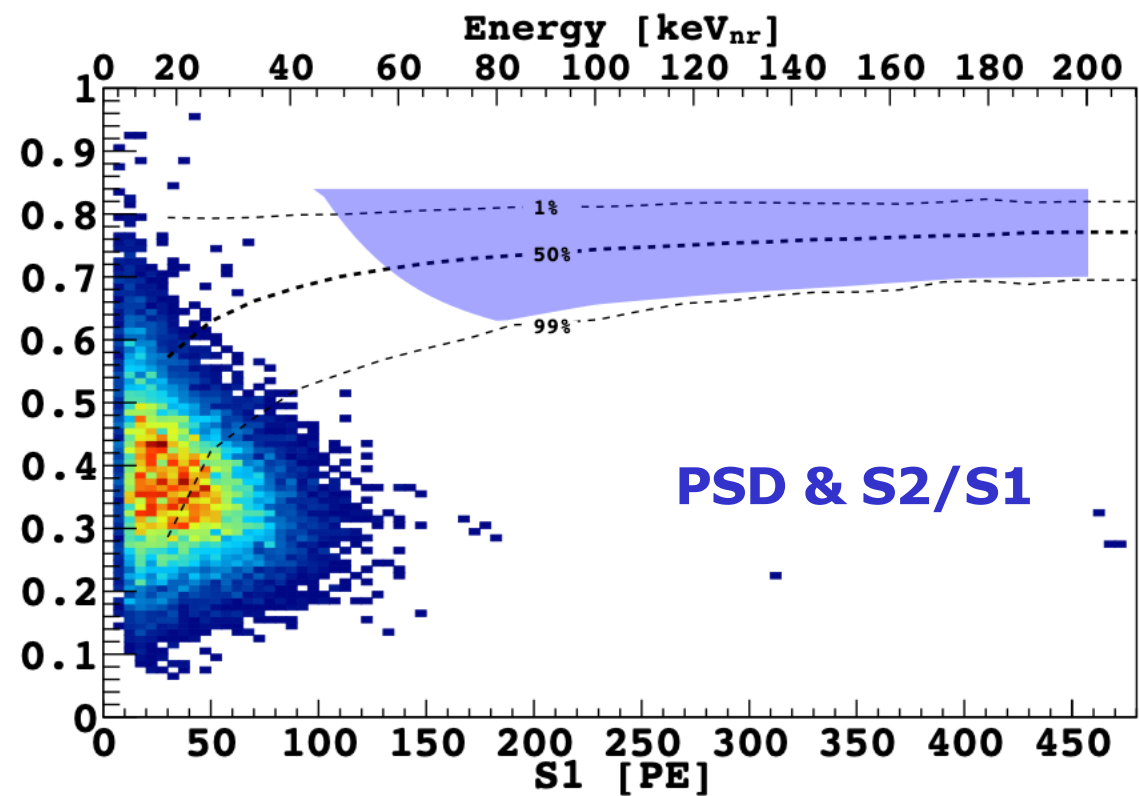
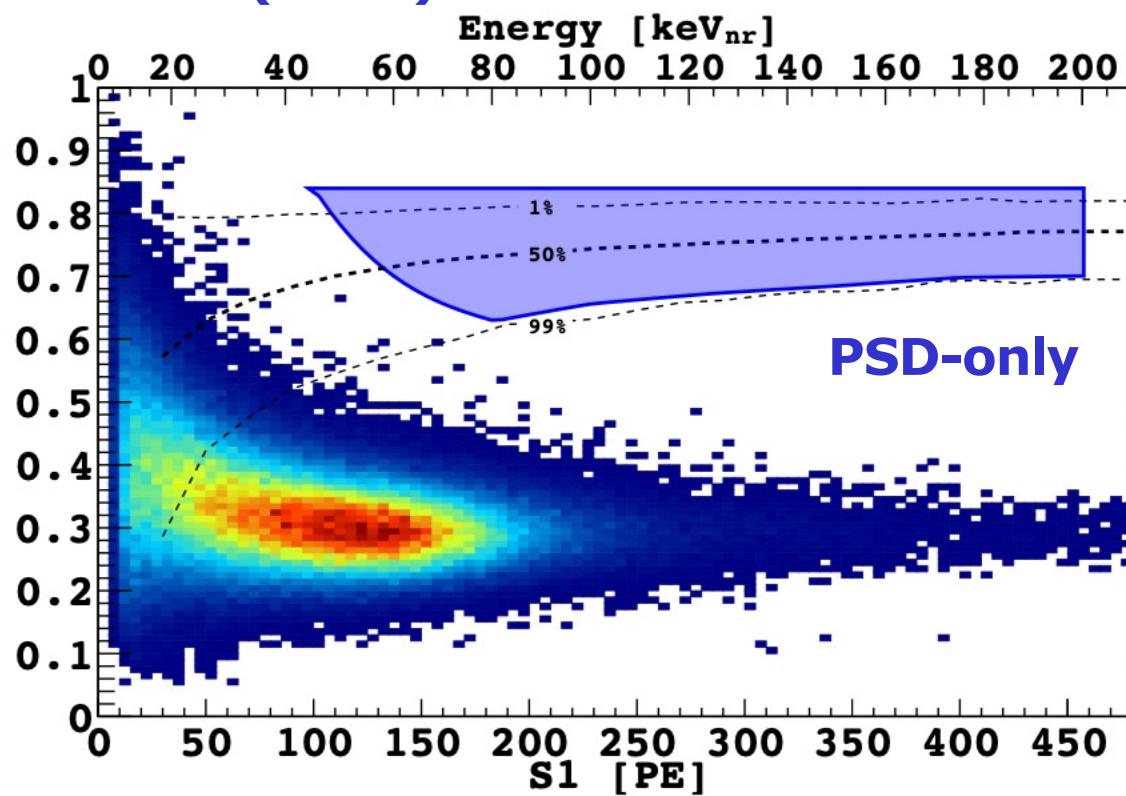
LIQUID ARGON TPC: DARKSIDE-50 @ LNGS

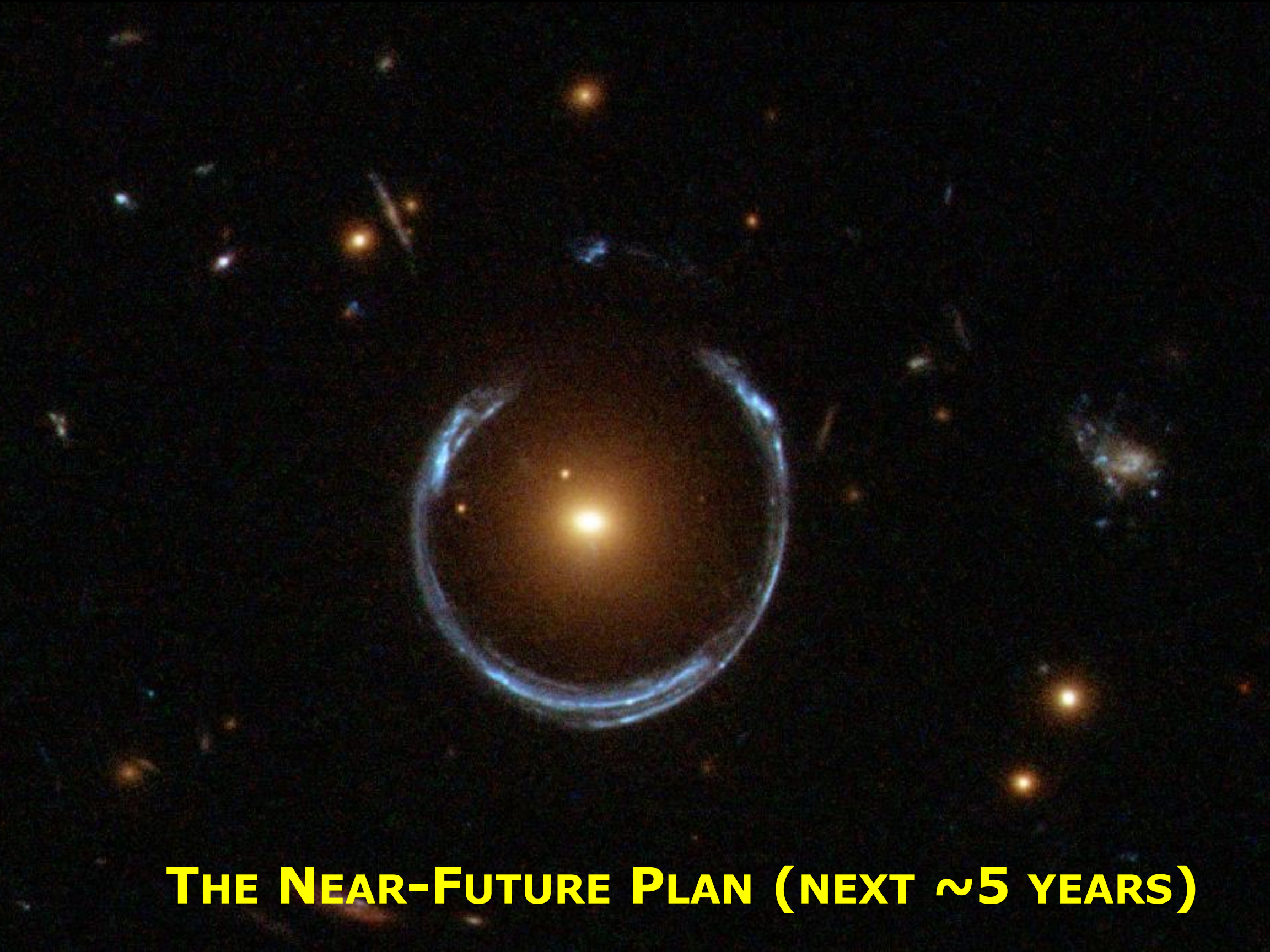


UAr WIMP search in 2015-2108

- 532 live days, 50 kg target
- x1400 ^{39}Ar depletion observed
- active neutron & muon vetoes
- entirely background-free!
- $\sigma < 1.14 \times 10^{-44} \text{ cm}^2 @ 100 \text{ GeV}$

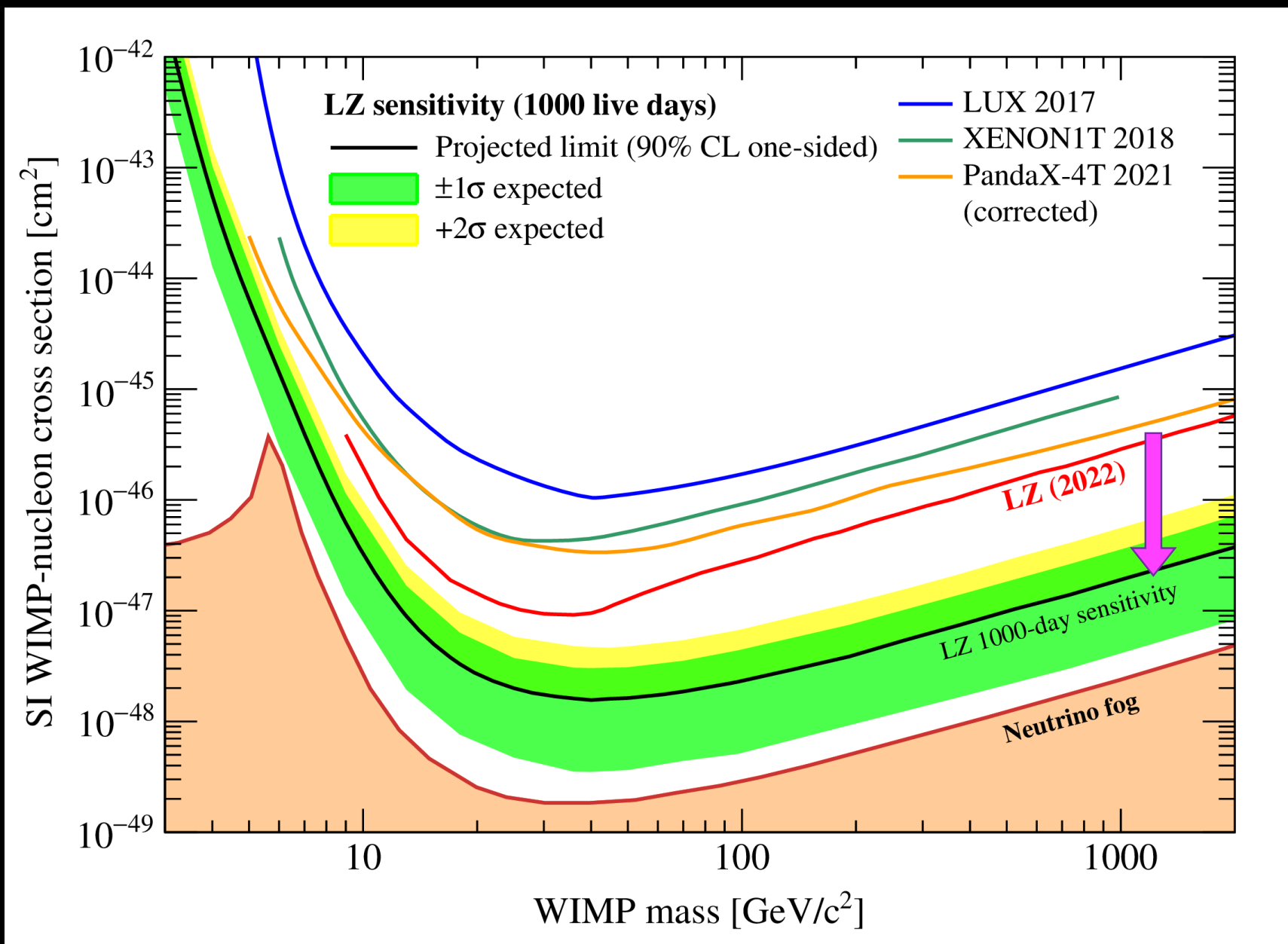
PRD 98 (2018)





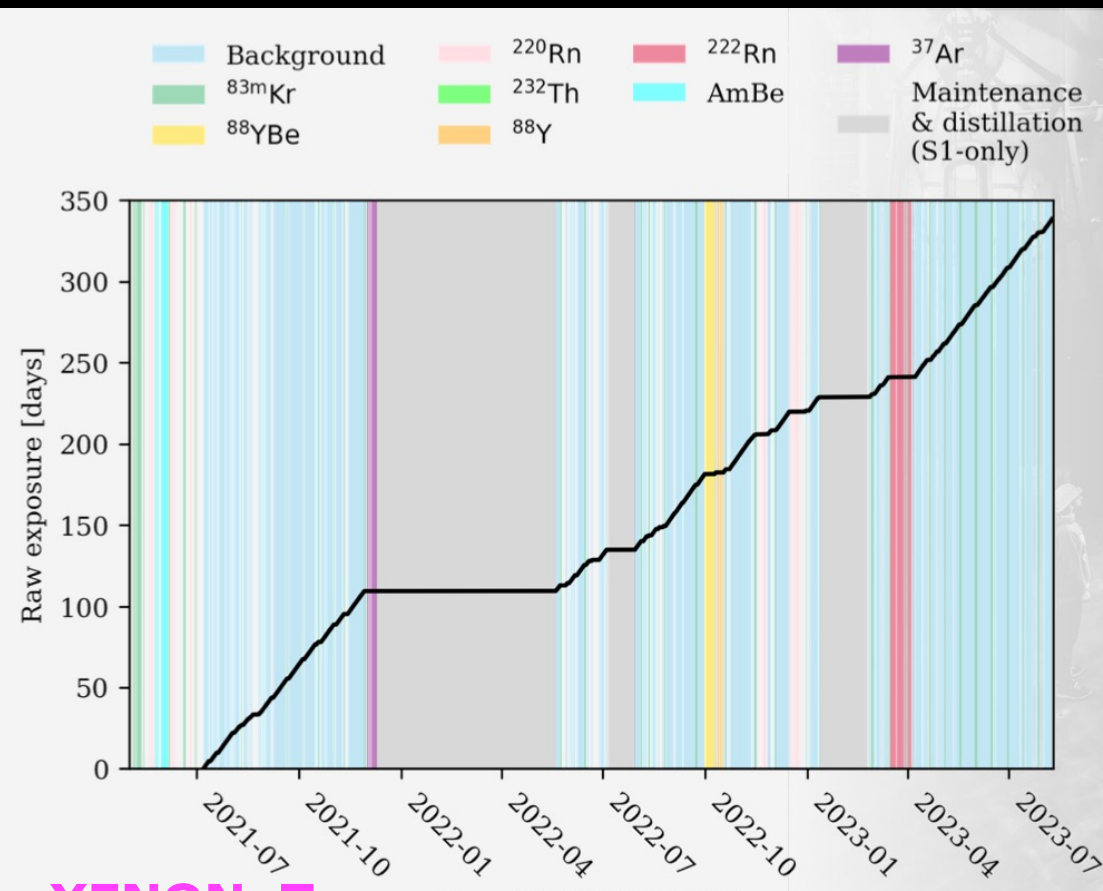
THE NEAR-FUTURE PLAN (NEXT ~ 5 YEARS)

LZ OPERATIONS EXTENDED THROUGH 2028



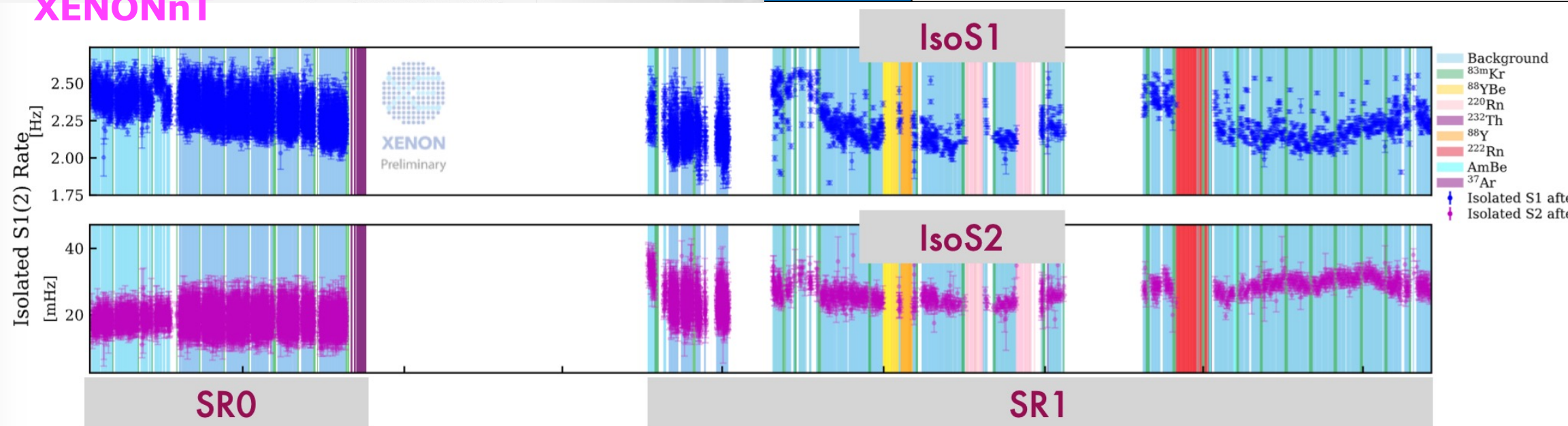
WHAT'S NEXT FOR LZ? LOTS. MORE. DATA.

XENONnT & PANDAX-4T TAKING DATA TOO



2020/11 – 2021/04	Commissioning (Run 0) 95 days
2021/07 – 2021/10	Tritium removal xenon distillation, gas flushing, etc
2021/11 – 2022/05	Physics run (Run 1) 164 days
2022/09 – 2023/12	CJPL B2 hall renovation xenon recuperation, detector upgrade
Current Status	Resuming physics data-taking

XENONnT



THE GRAND UNIFICATION OF LAR EXPERIMENTS

Global Argon Dark
Matter Collaboration

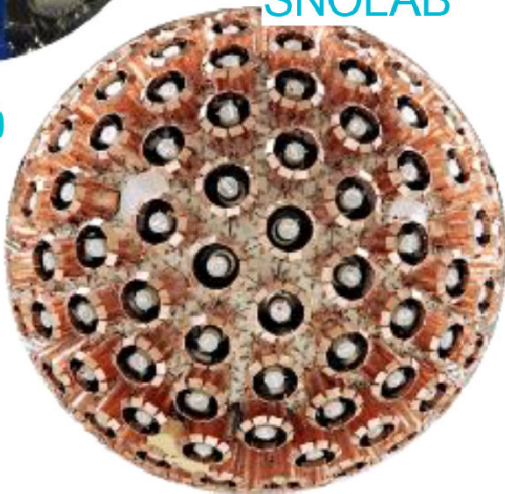
MiniCLEAN
SNOLAB



ArDM
Canfranc

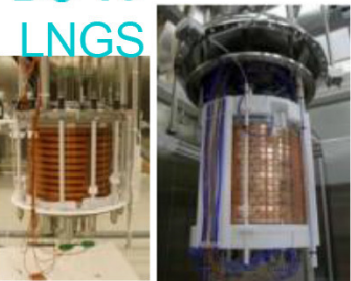


DEAP-3600
SNOLAB

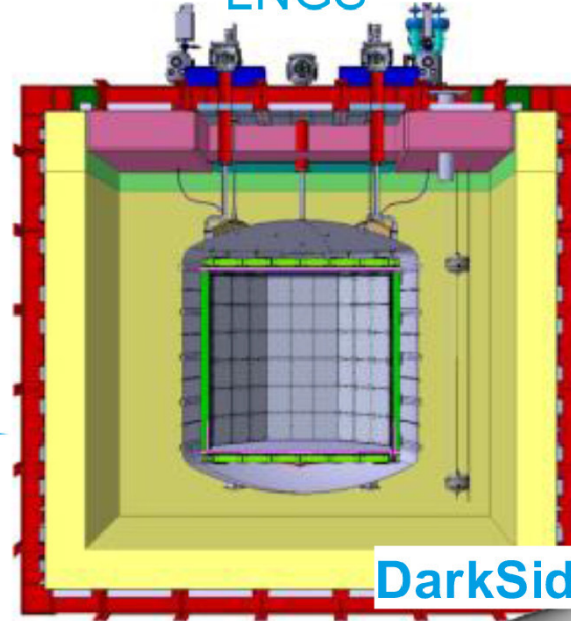


DarkSide50

DS-10
LNGS



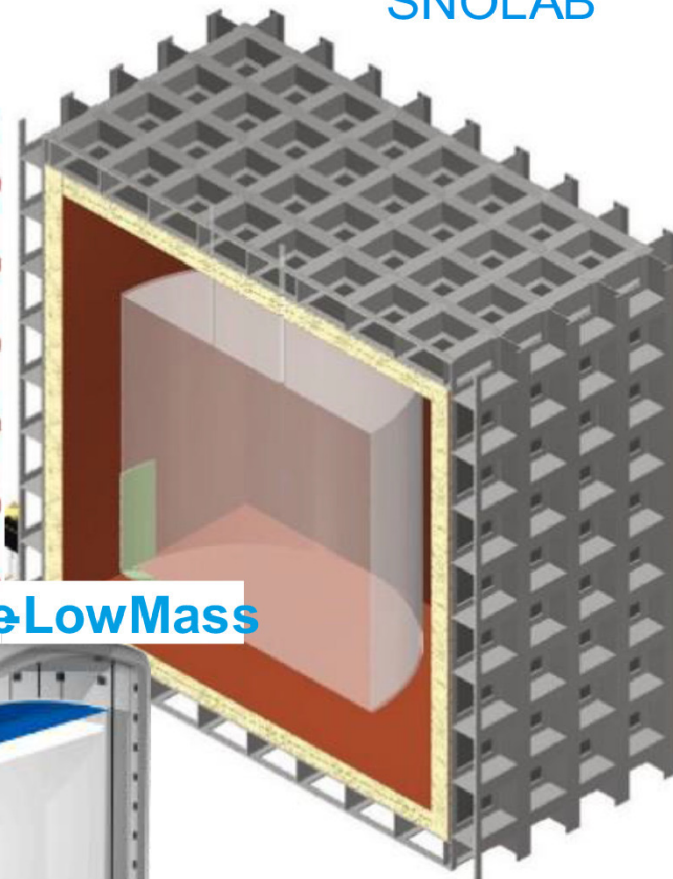
DarkSide20k
LNGS



DarkSideLowMass



Argo
SNOLAB



2011

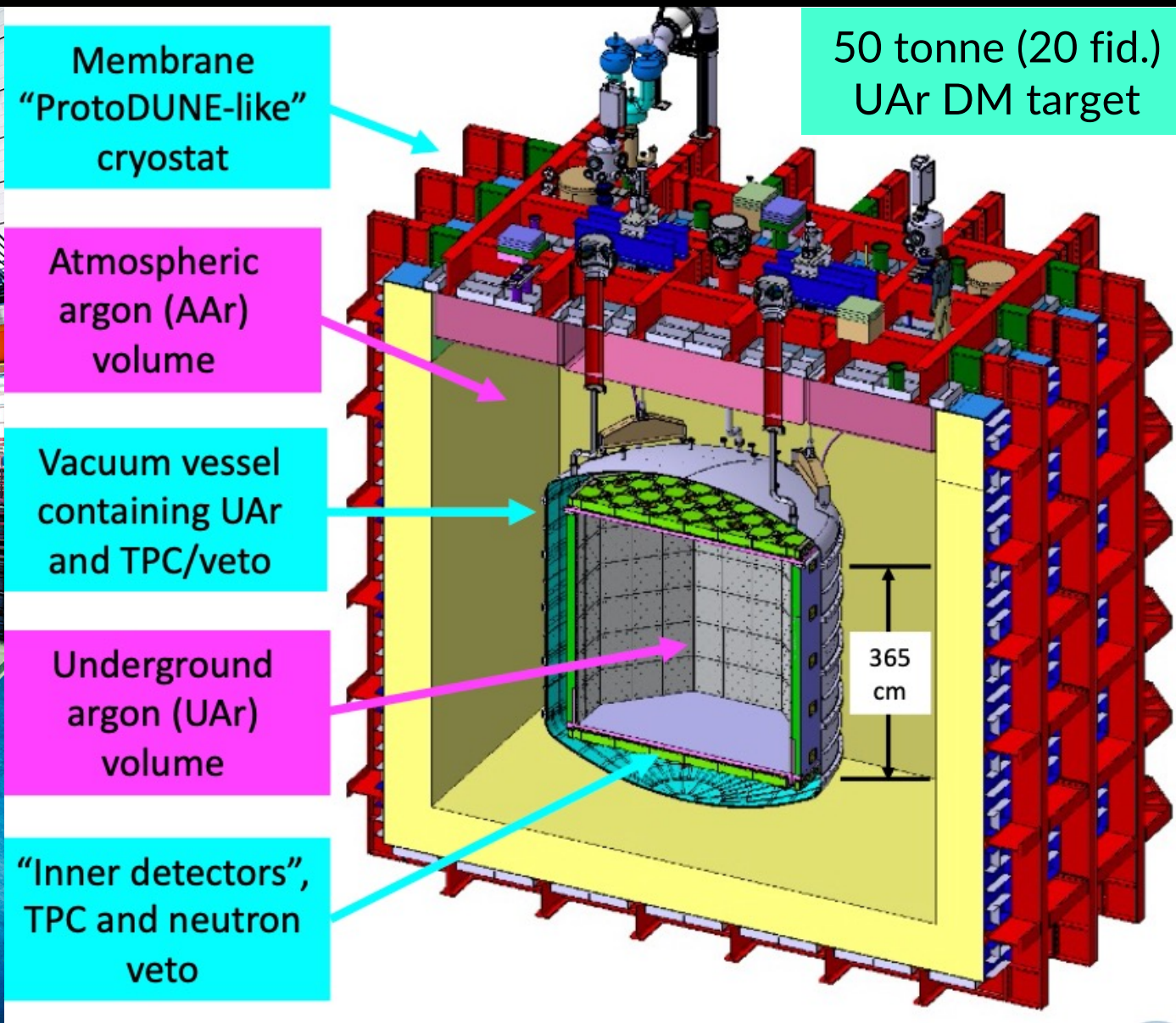
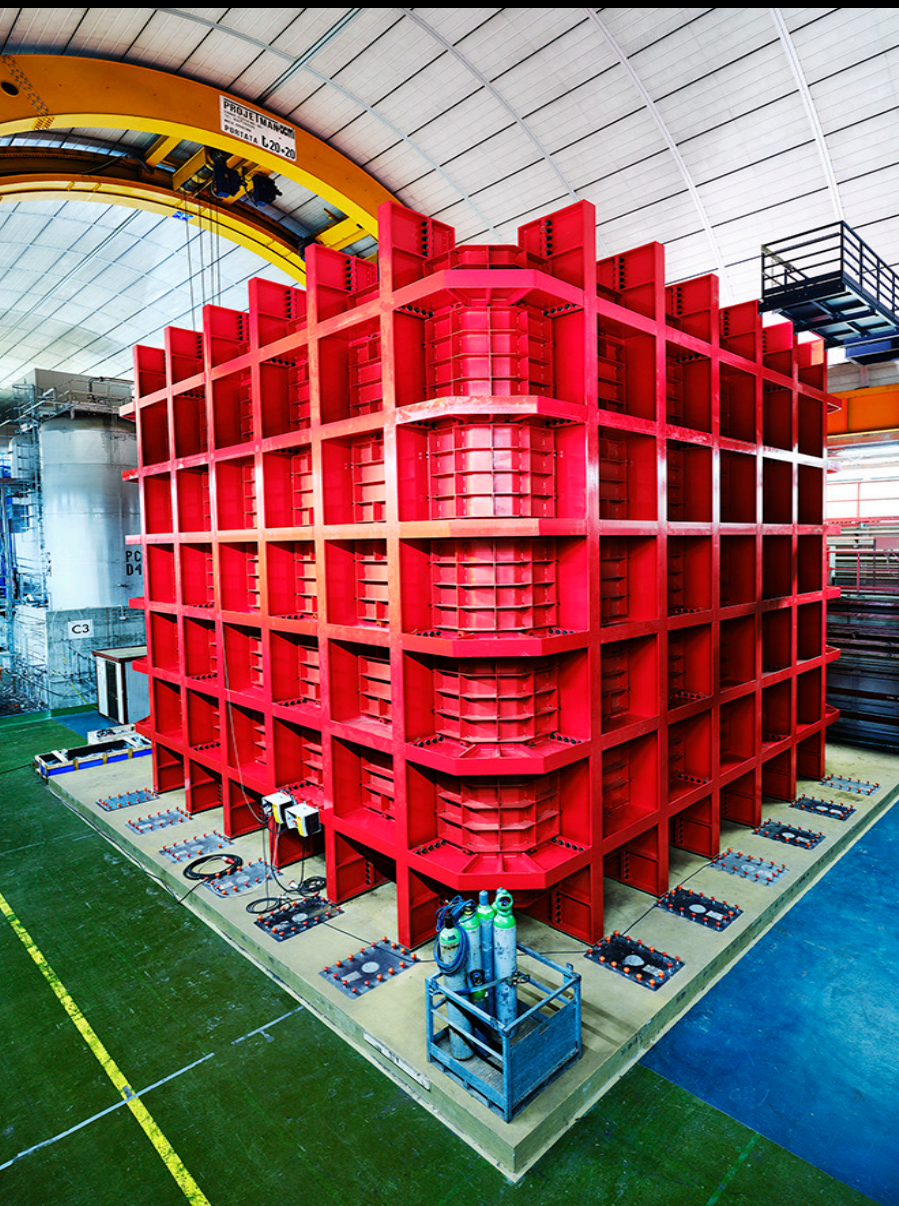
2013

2016

2026

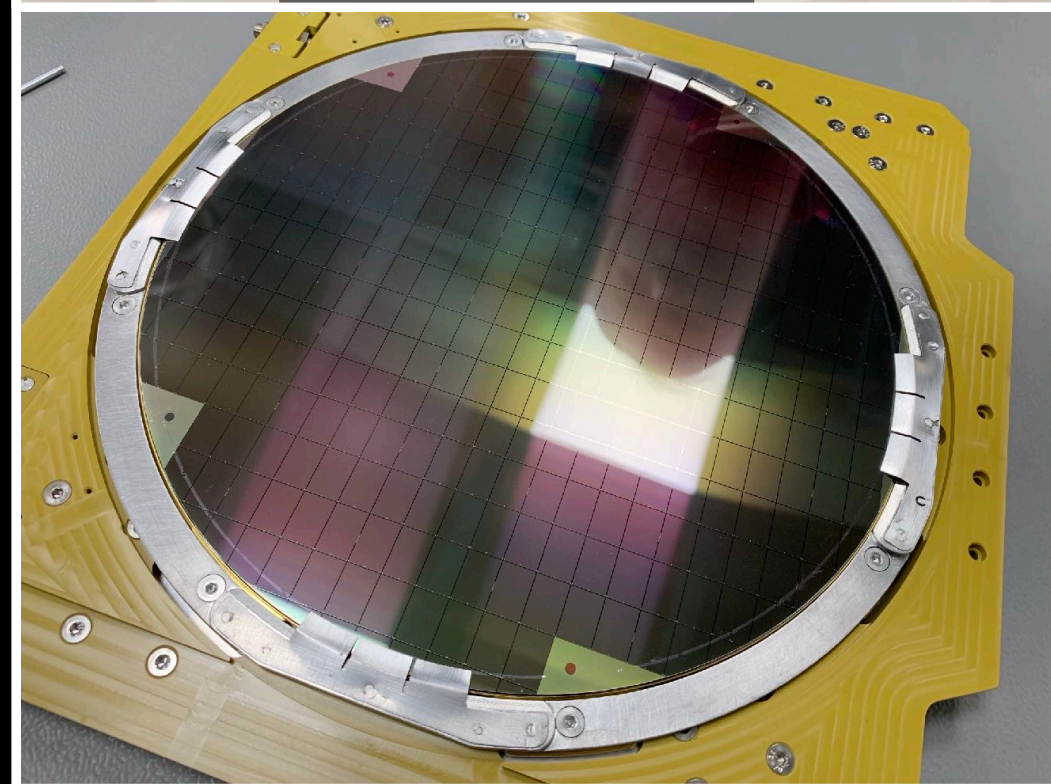
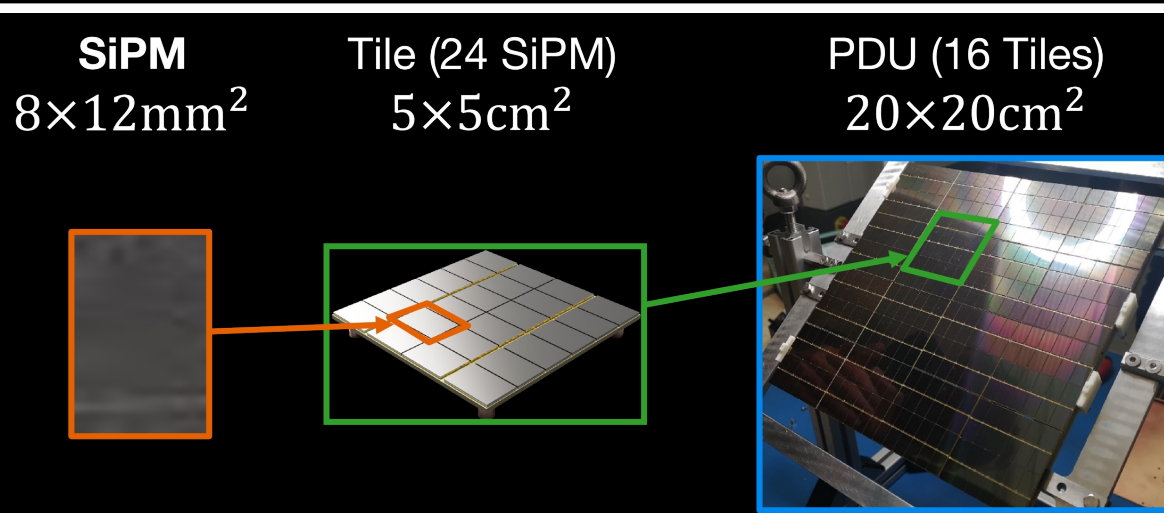
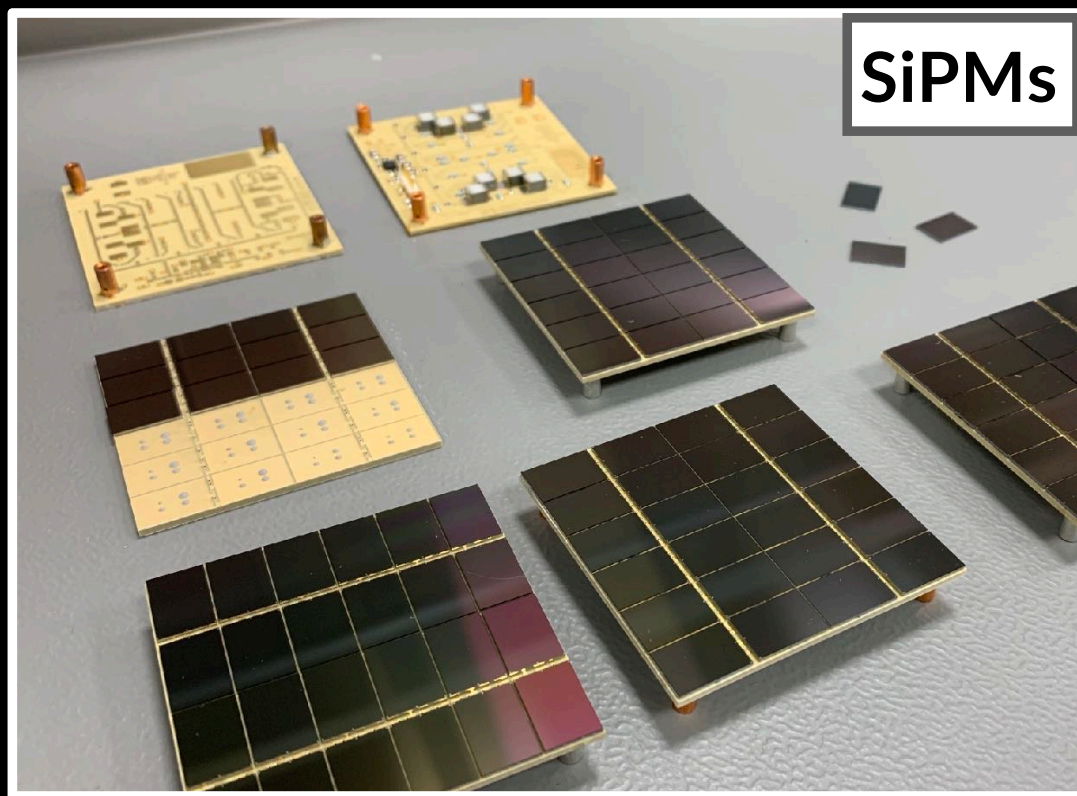
2030+

DARKSIDE-20K AT GRAN SASSO

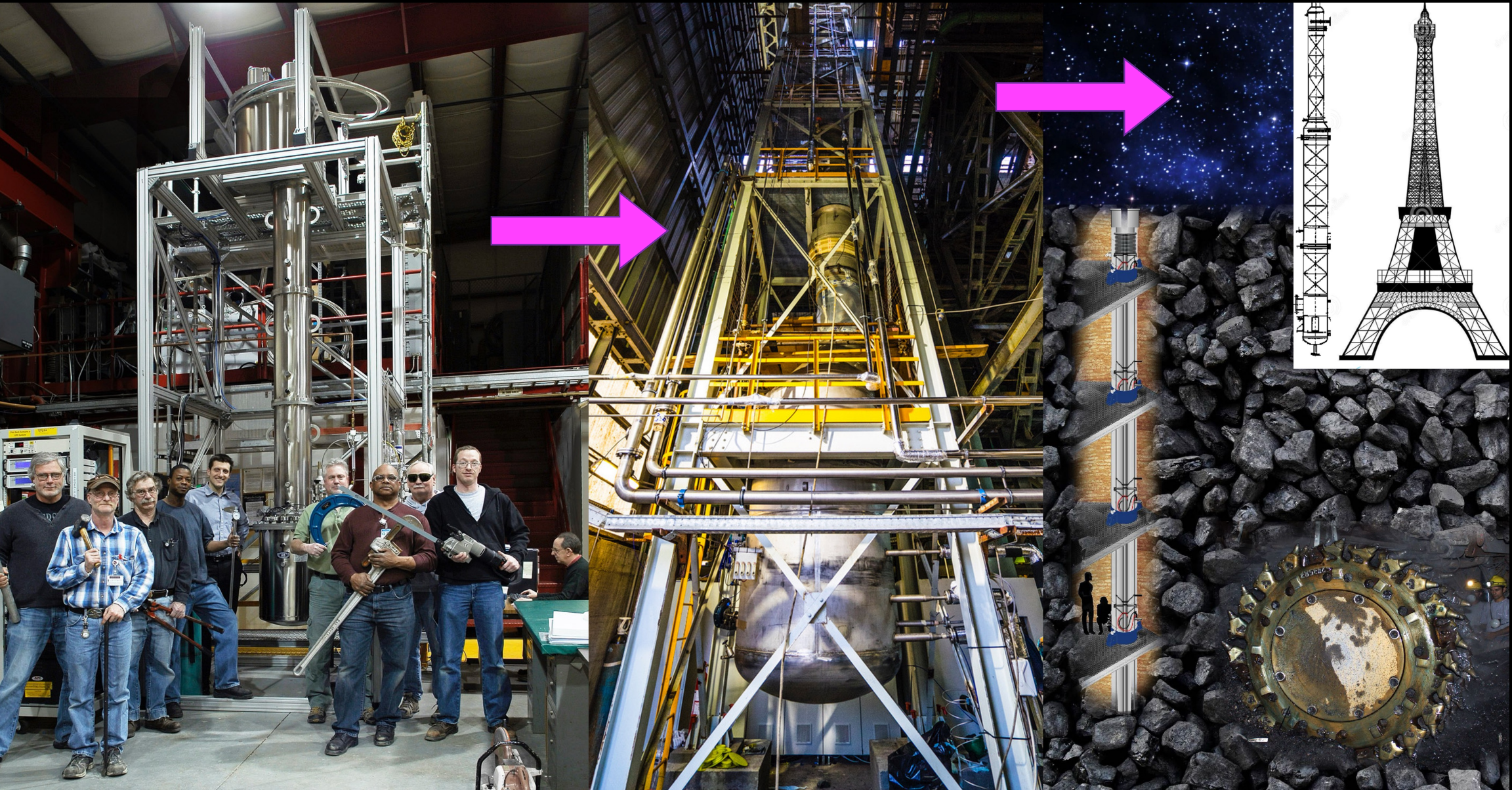


See also: Ahmad, Stefanizzi, Pesudo, Santorelli, Calabrese, Pandola, Van Uffelen, Marasciulli, Pino, Salomone, Thieme

DARKSIDE-20K: PHOTODETECTOR TECHNOLOGY



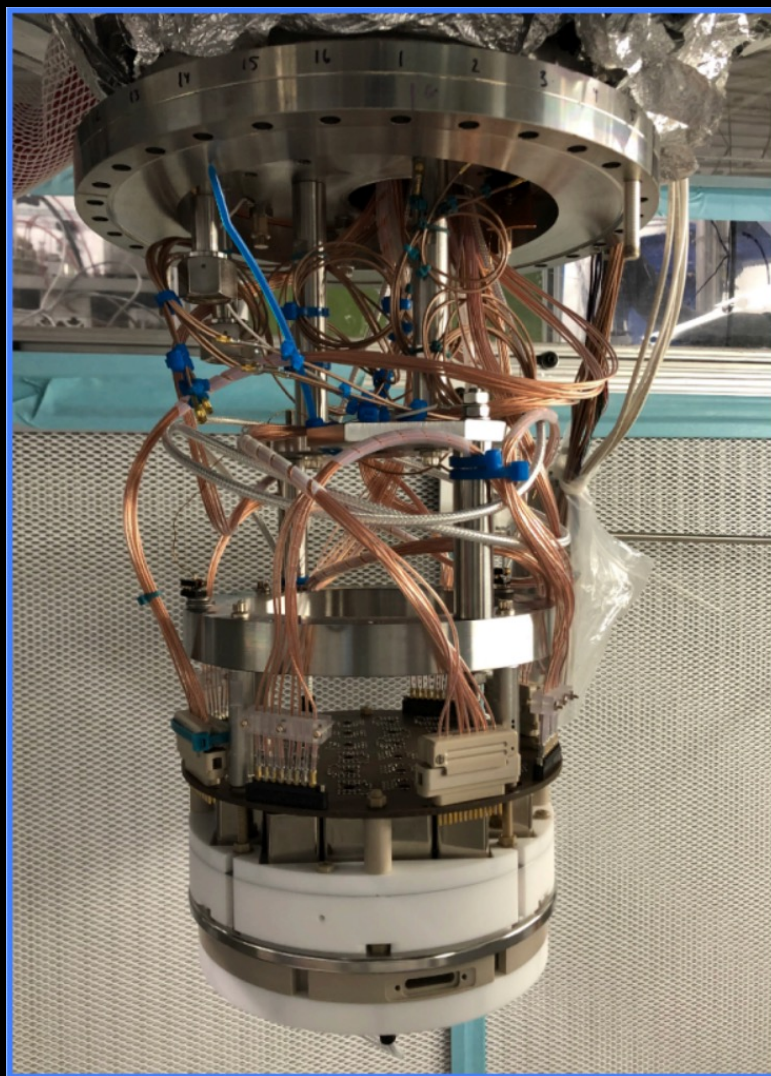
DARKSIDE-20K: UAr PURIFICATION



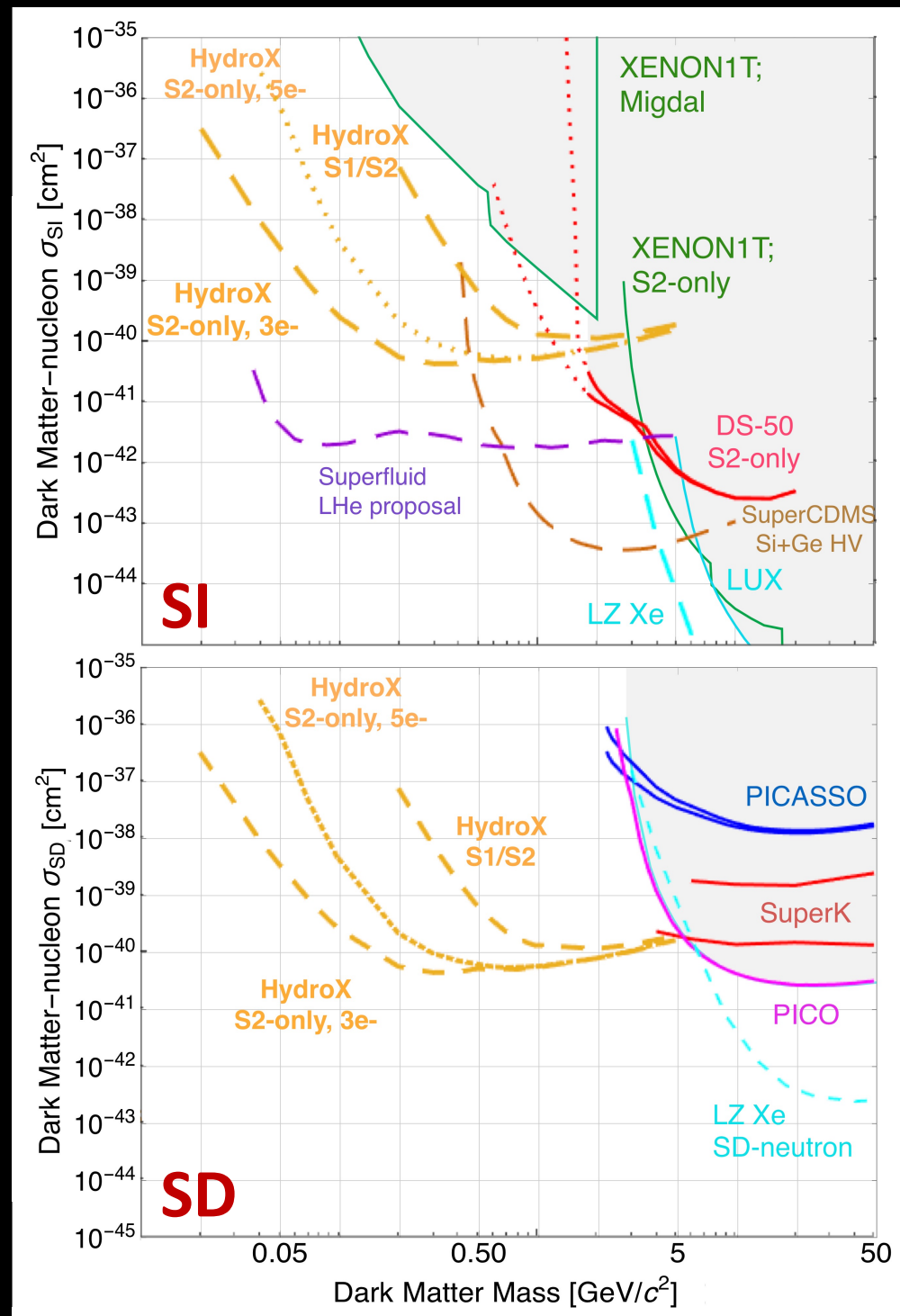
GOAL: CHEMICAL + ISOTOPIC DISTILLATION

WHAT ELSE IS FUN ON THIS TIME SCALE?

HYDROX (HYDROGEN IN XE): POTENTIAL LZ UPGRADE

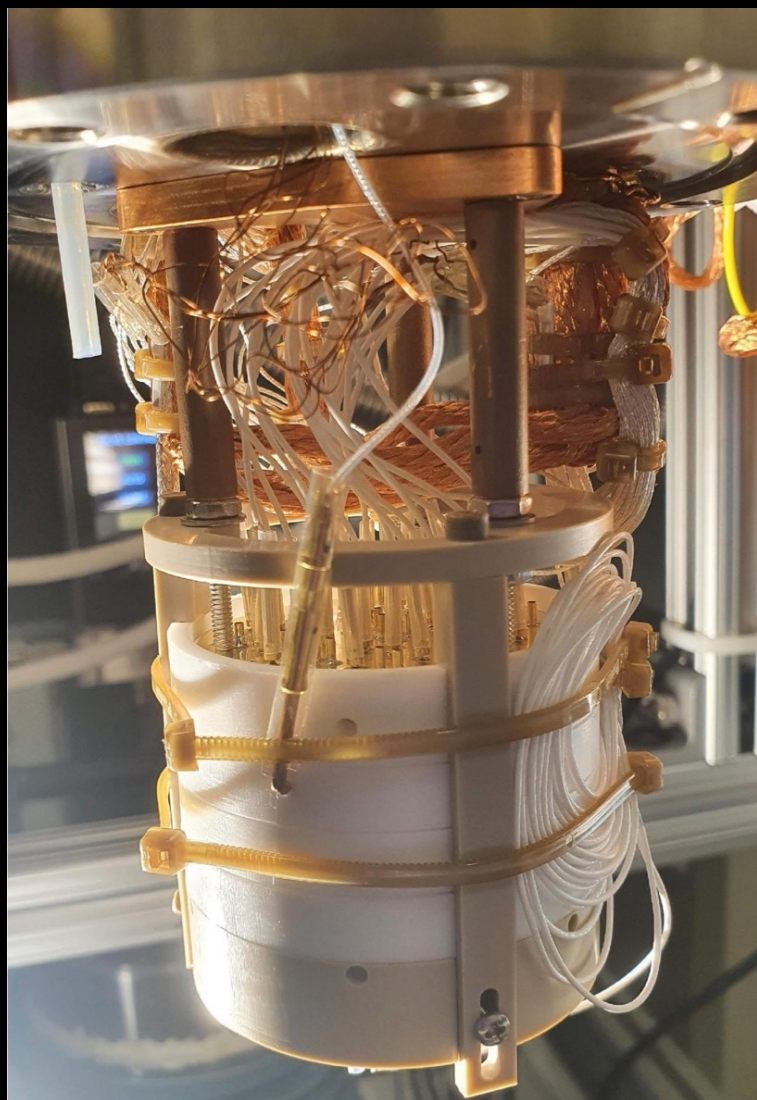


SEE A. WANG'S TALK

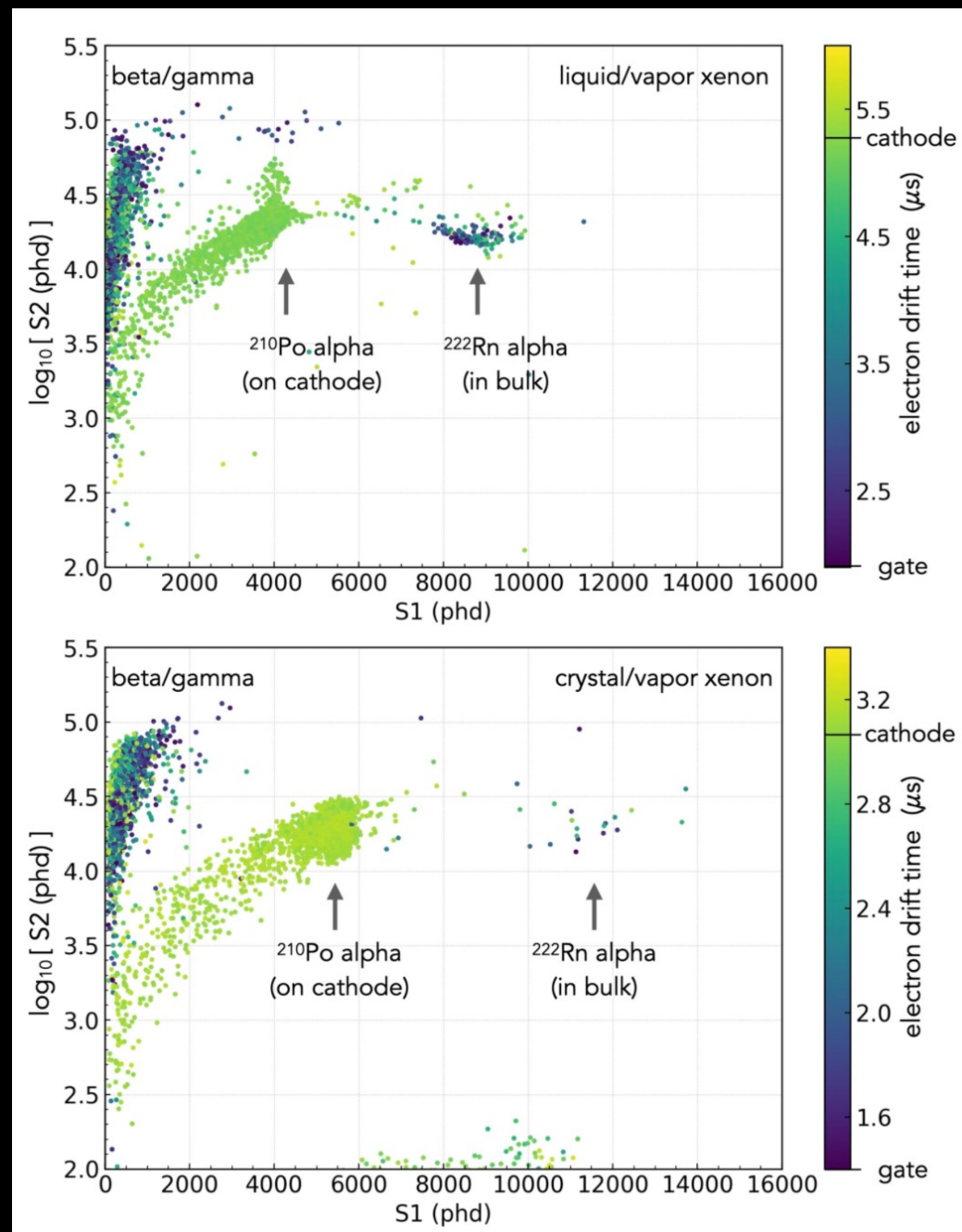


WHAT ELSE IS FUN ON THIS TIME SCALE?

CRYSTALiXe (CRYSTAL/VAPOR Xe): POTENTIAL LZ UPGRADE



SEE P. SORENSEN'S TALK

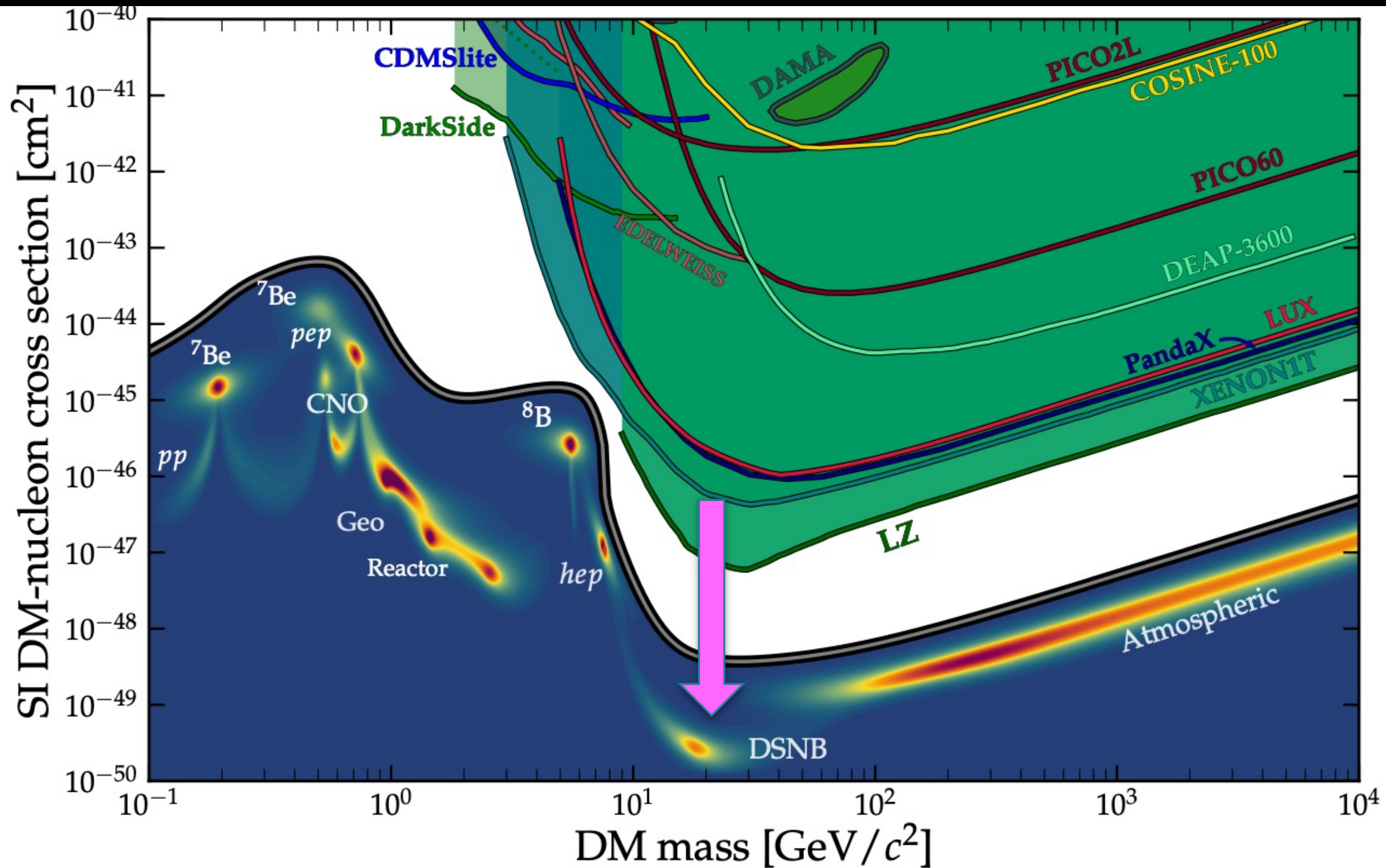


arXiv:2312.15082 [hep-ex] 2023

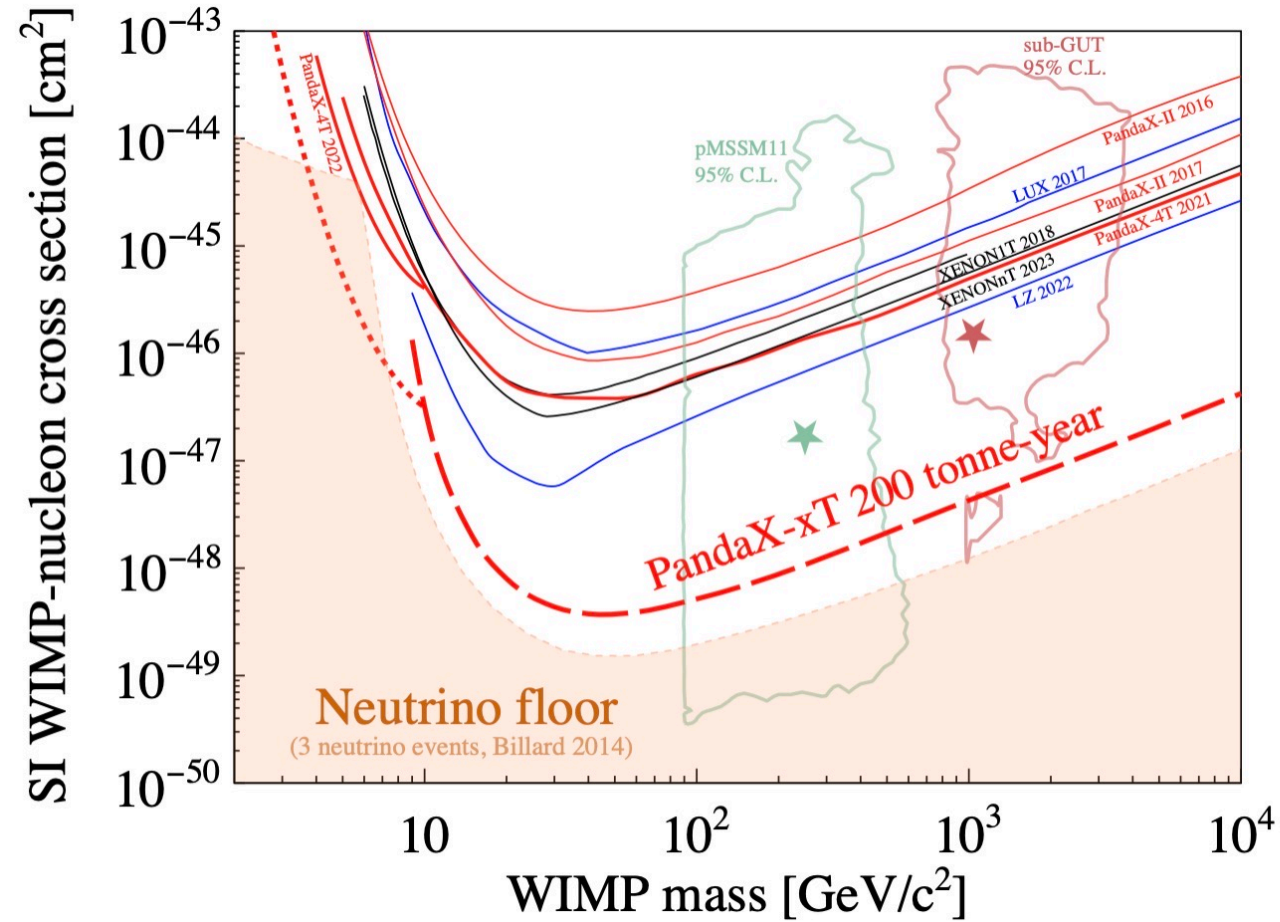
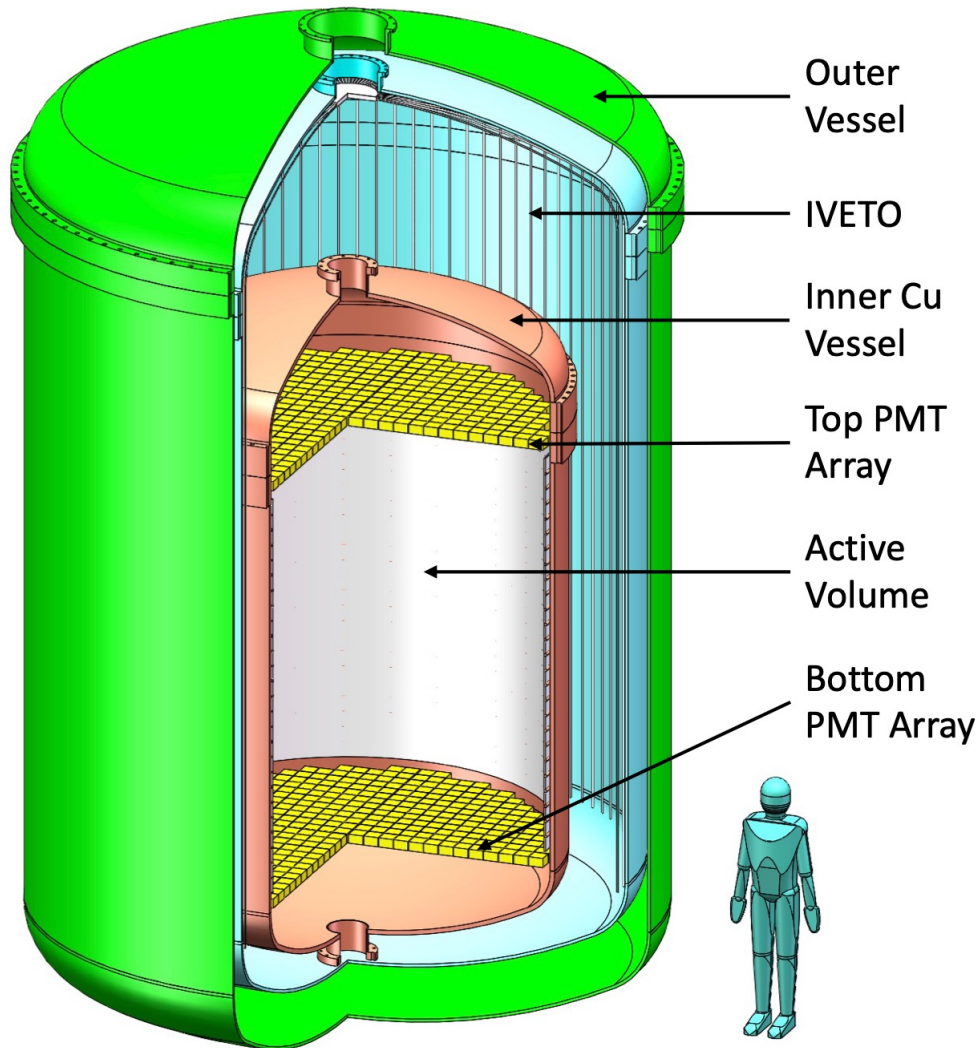


THE FUTURE-FUTURE PLAN (NEXT DECADE?)

RACE TO THE BOTTOM? OR WIMP DISCOVERY?



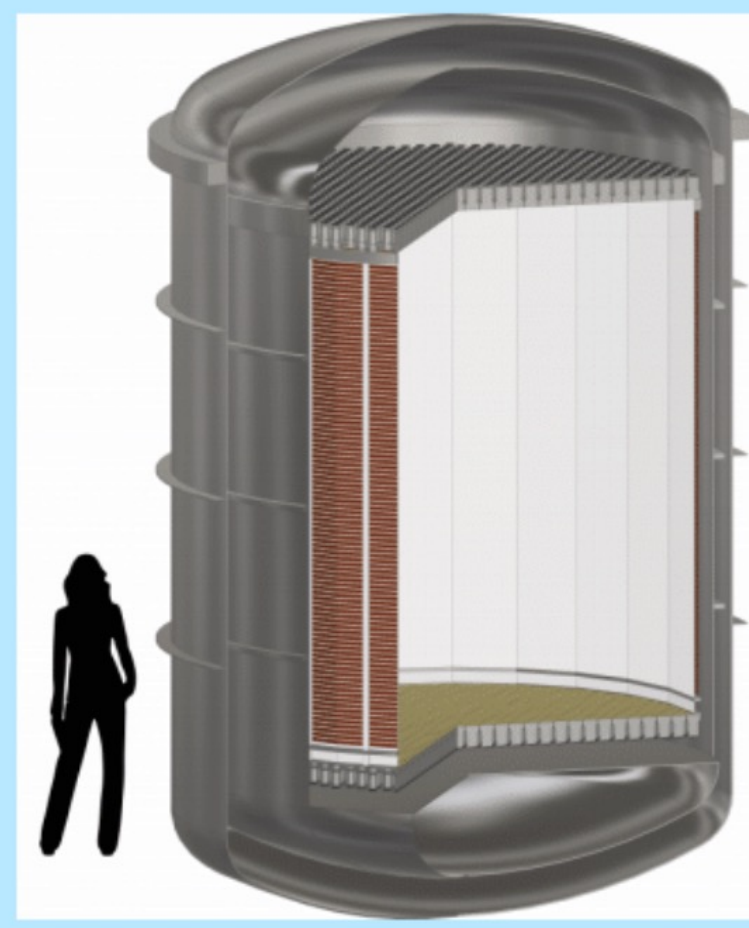
PANDAX-xT: MULTI-TEN-TONNES LXe



43 TONNES, 10-YEAR RUN

ARXIV: 2402.03596 [HEP-EX] 2024

THE GRAND UNIFICATION OF LXe EXPERIMENTS



XLZD = XENON + LUX-ZEPLIN (LZ) + DARWIN

**MOU (2001) → CONSORTIUM (NOW) → COLLABORATION
TO BUILD THE ULTIMATE LXe DARK MATTER DETECTOR**

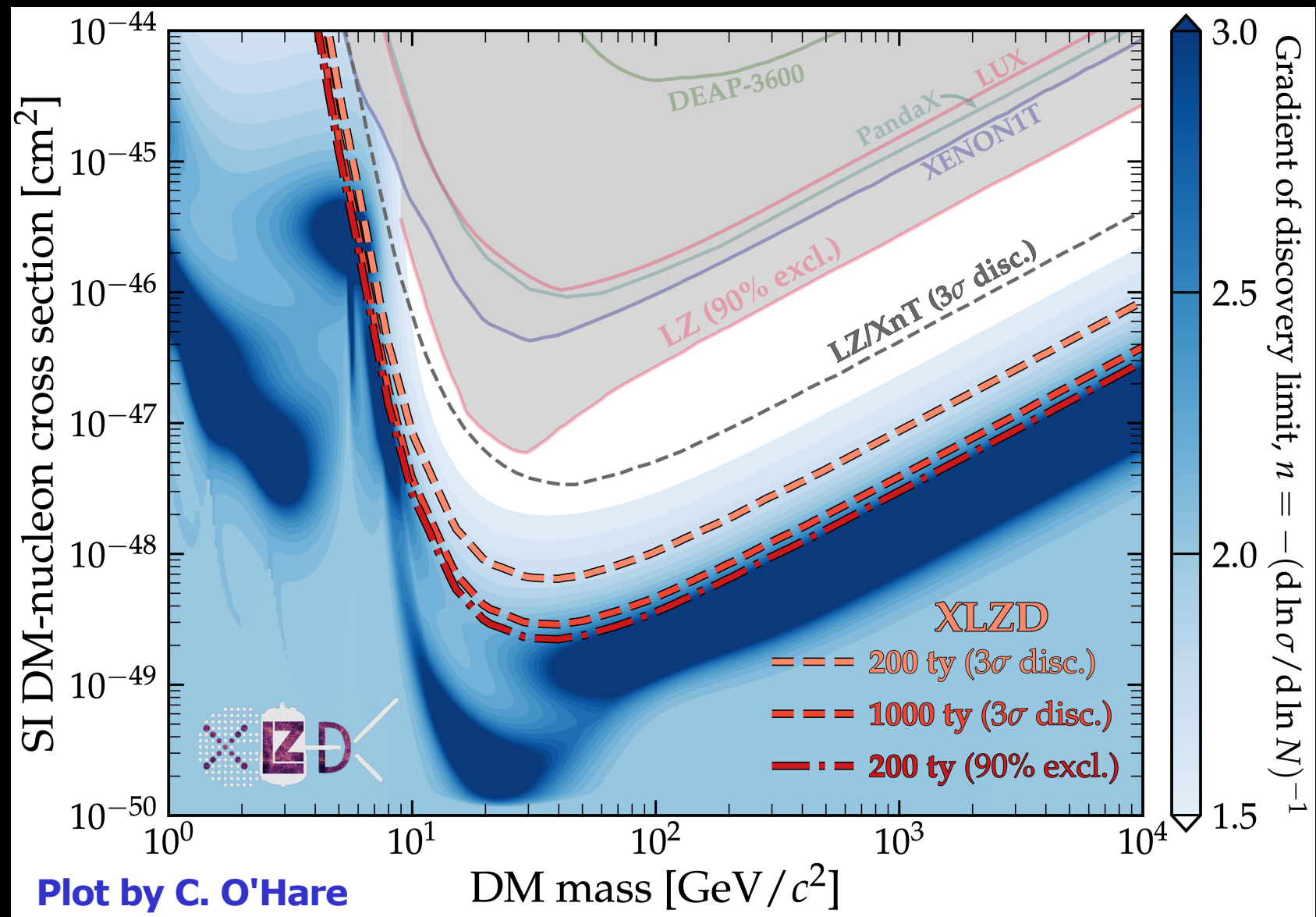
THE GRAND UNIFICATION OF LXe EXPERIMENTS

MOU (2001) → CONSORTIUM (NOW) → COLLABORATION
TO BUILD THE ULTIMATE LXe DARK MATTER DETECTOR

60-TONNES
SCALE LXe
DETECTOR
OBSERVATORY

REACH DOWN
AND INTO THE
NEUTRINO FOG

SEE TALKS BY
L. BAUDIS &
C. CAPELLI



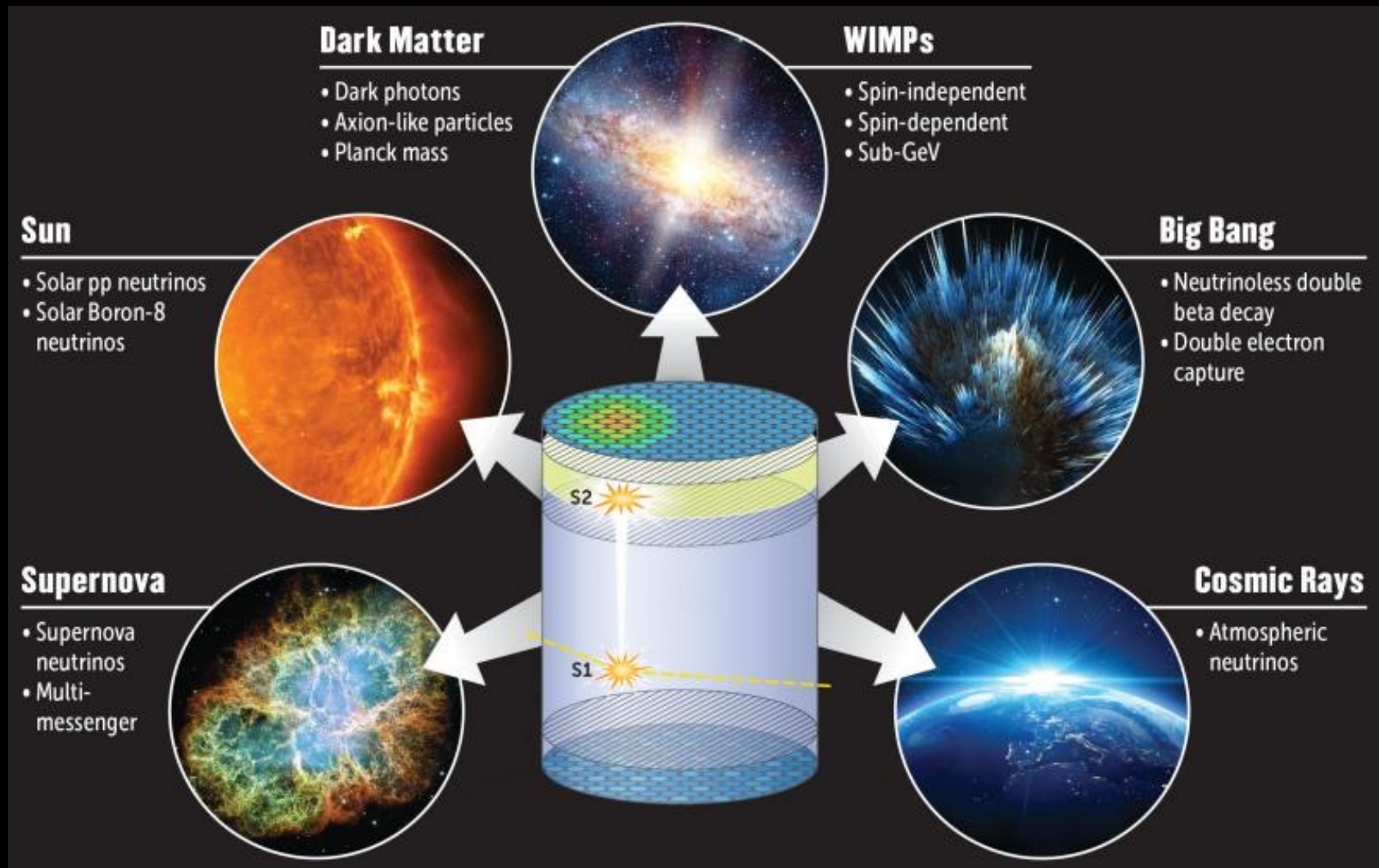
THE GRAND UNIFICATION OF LXe EXPERIMENTS

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TO BUILD THE ULTIMATE LXe DARK MATTER DETECTOR

60-TONNES
SCALE LXe
DETECTOR
OBSERVATORY

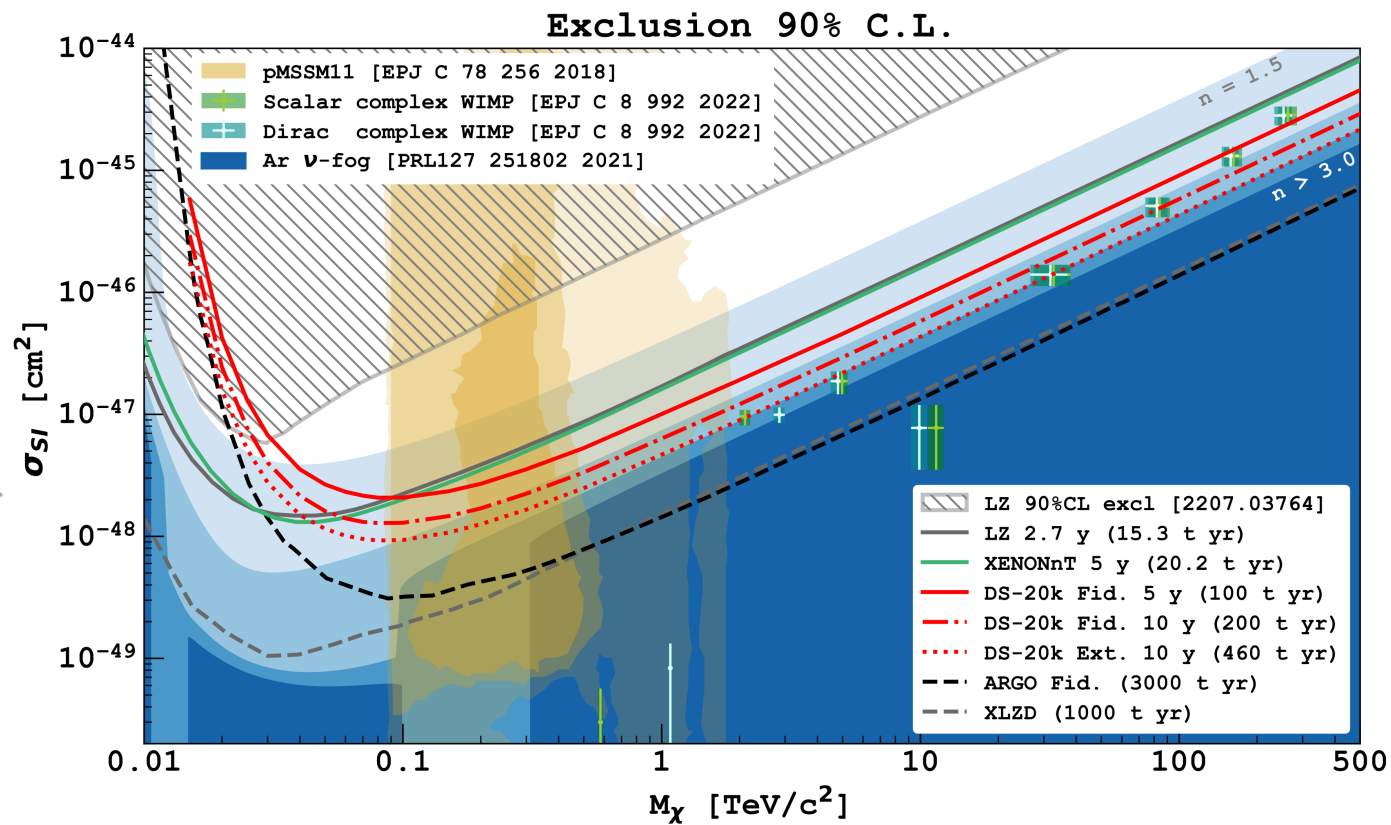
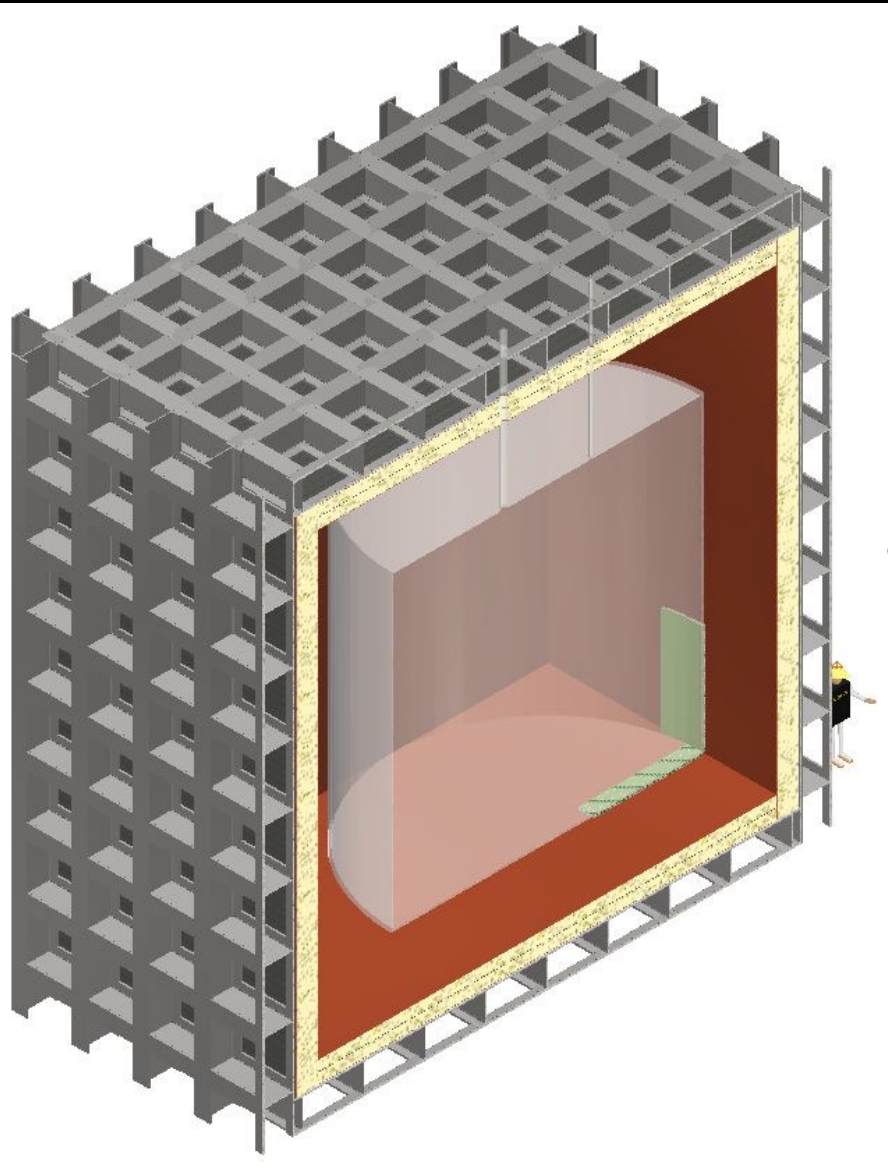
REACH DOWN
AND INTO THE
NEUTRINO FOG

SEE TALKS BY
L. BAUDIS &
C. CAPELLI



THE ULTIMATE DETECTOR: ARGON VERSION

ARGO: 400 TONS UAr (300 FIDUCIAL), 250 m² PDCs
COVERING FULL ACRYLIC VESSEL SURFACE



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What are the new XnT results?



0

Migdal effect

0

CNS Boron-8

0

EFT models

0

2nBB, 0nBB

0

ER channels

0

Other Results

<https://www.menti.com/altuoh1wbtg9>