

Development of Position Reconstruction in the LUX-ZEPLIN Outer Detector

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LUX-ZEPLIN

- LUX-ZEPLIN (LZ) is the world's most sensitive direct dark matter detection experiment, aiming to discover Weakly Interacting Massive Particles (WIMPs).

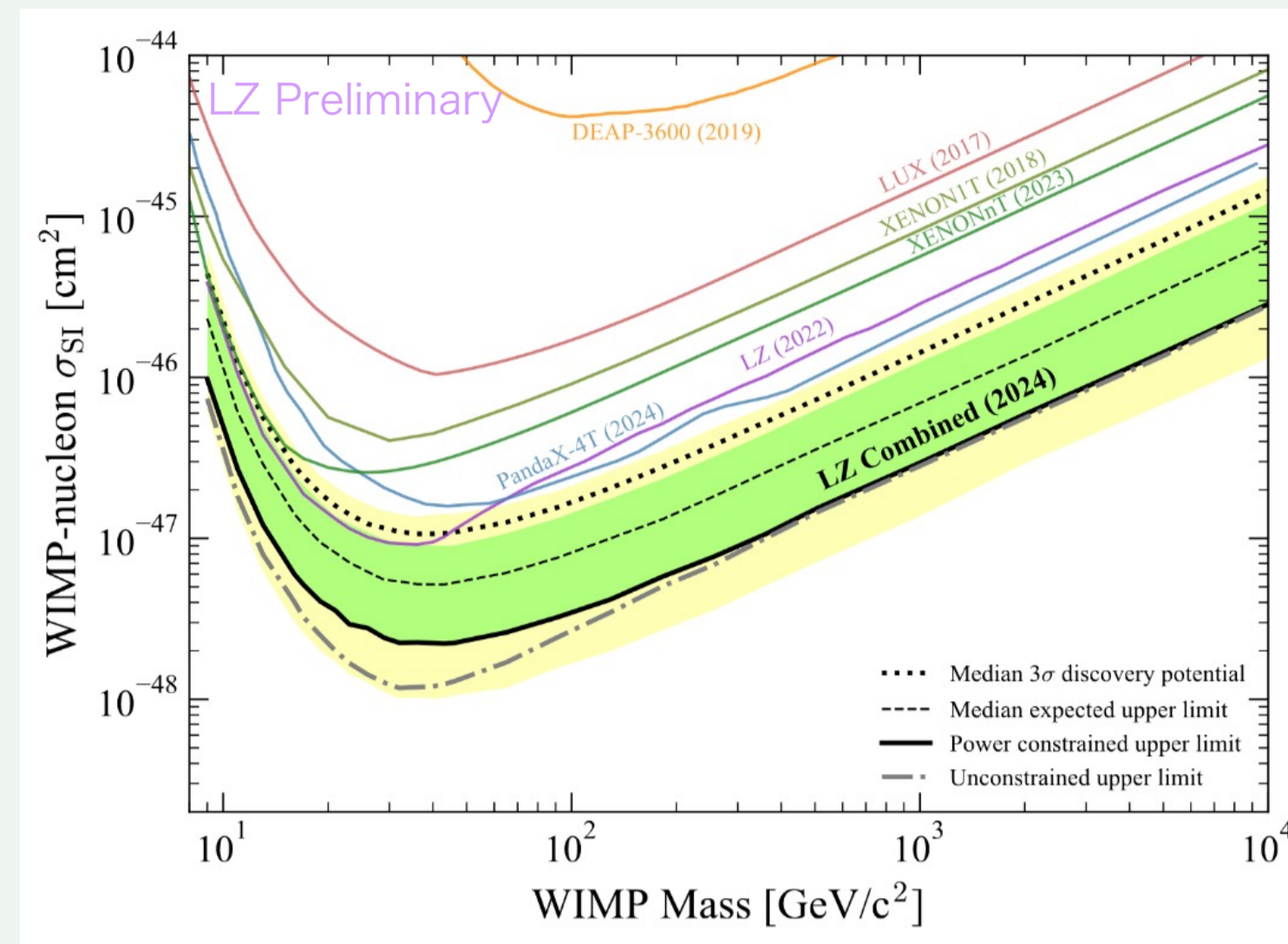


Figure 1: Combined WS2022 + WS2024 power constrained sensitivity limit (black line) for spin-independent WIMP cross-sections vs WIMP mass. The 1σ and 2σ sensitivity bands are shown in green and yellow. [1, 2]

- LZ utilises:
 - Dual-phase liquid Xe (LXe) time projection chamber (TPC).
 - LXe skin veto detector.
 - Gd-loaded liquid scintillator (GdLS) outer detector (OD).

Time Projection Chamber

- 7t of liquid xenon (LXe), gas phase with 8mm of gaseous xenon (GXe).
- Primary scintillation (S1) and ionisation e⁻s produced by interactions with LXe.
- Secondary scintillation (S2) signal from electroluminescence.

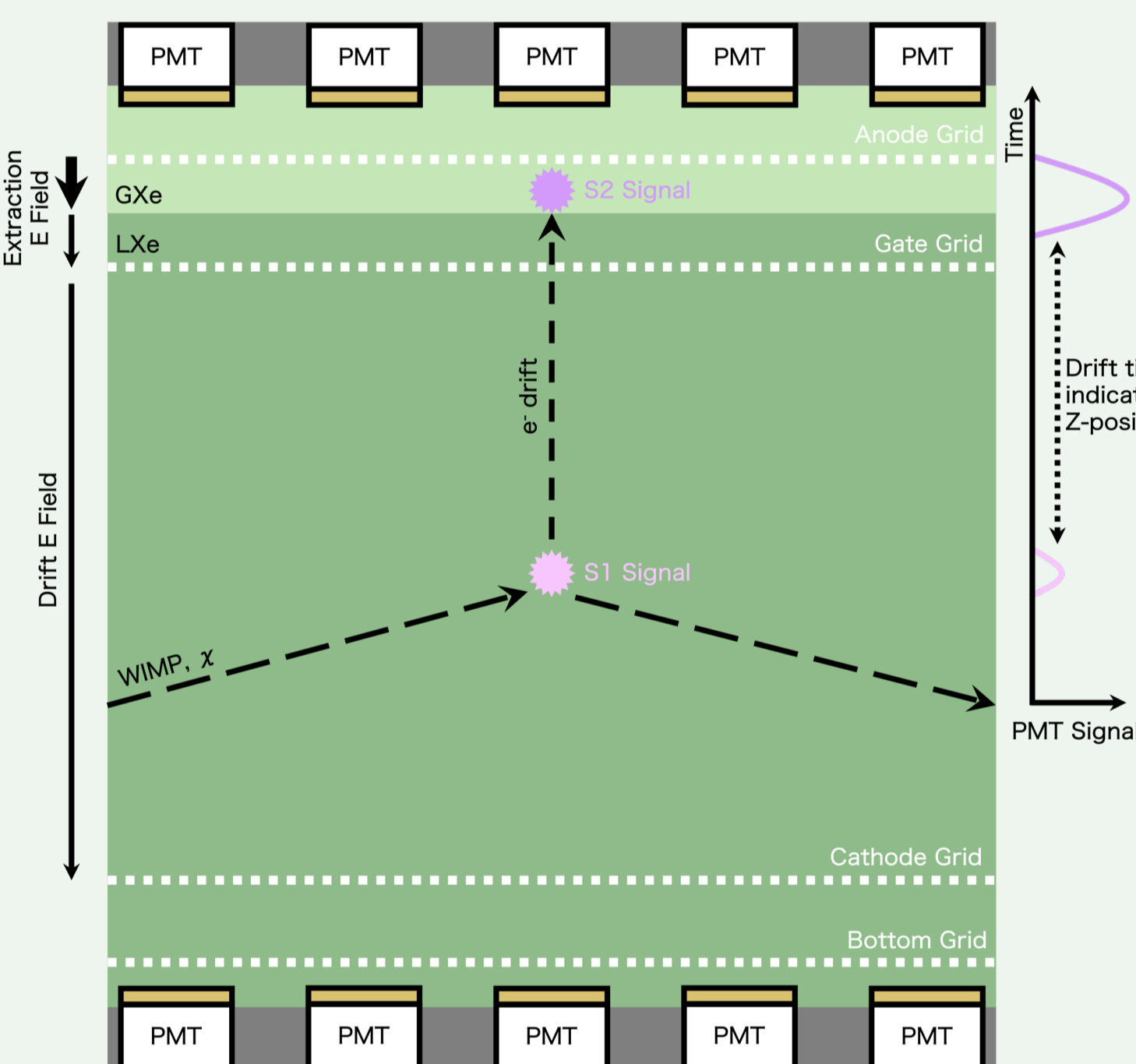


Figure 2: Schematic Diagram of the LZ TPC.

- Particle discrimination – S1:log(S2).
- 3D position reconstruction:
 - (X,Y) from S2 hit pattern.
 - Z from drift time.

Outer Detector

- 17t of GdLS across 10 tanks of varying geometry.
- Cylindrical array of 120 PMTs. [3]
- Primary focus:
 - Veto single scatter neutrons with an efficiency of $89\pm 3\%$.
 - Characterise the neutron background.

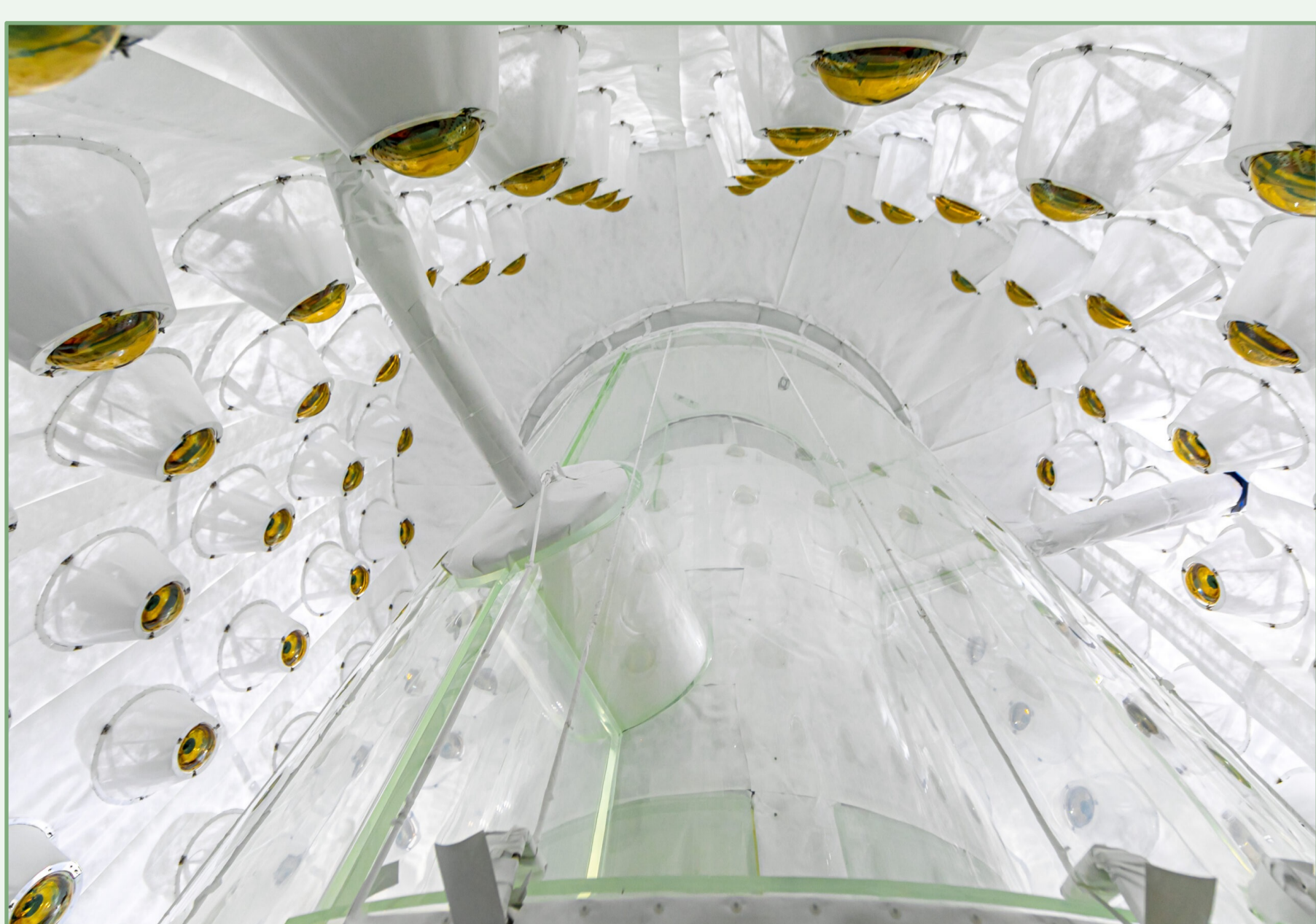


Figure 3: Fully assembled LUX-ZEPLIN Outer Detector

Outer Detector Position Reconstruction

- Current OD position reconstruction uses a centre of gravity (CoG) method.
 - Positions are being largely mis-reconstructed.
- Motivation:
 - Additional quantities for background discrimination.
 - Light collection efficiency mapping.
 - Equivalent energy thresholds.
 - Reduced neutron veto dead-time.
- A machine learning approach has been investigated, utilising convolutional neural network (CNN) algorithms.

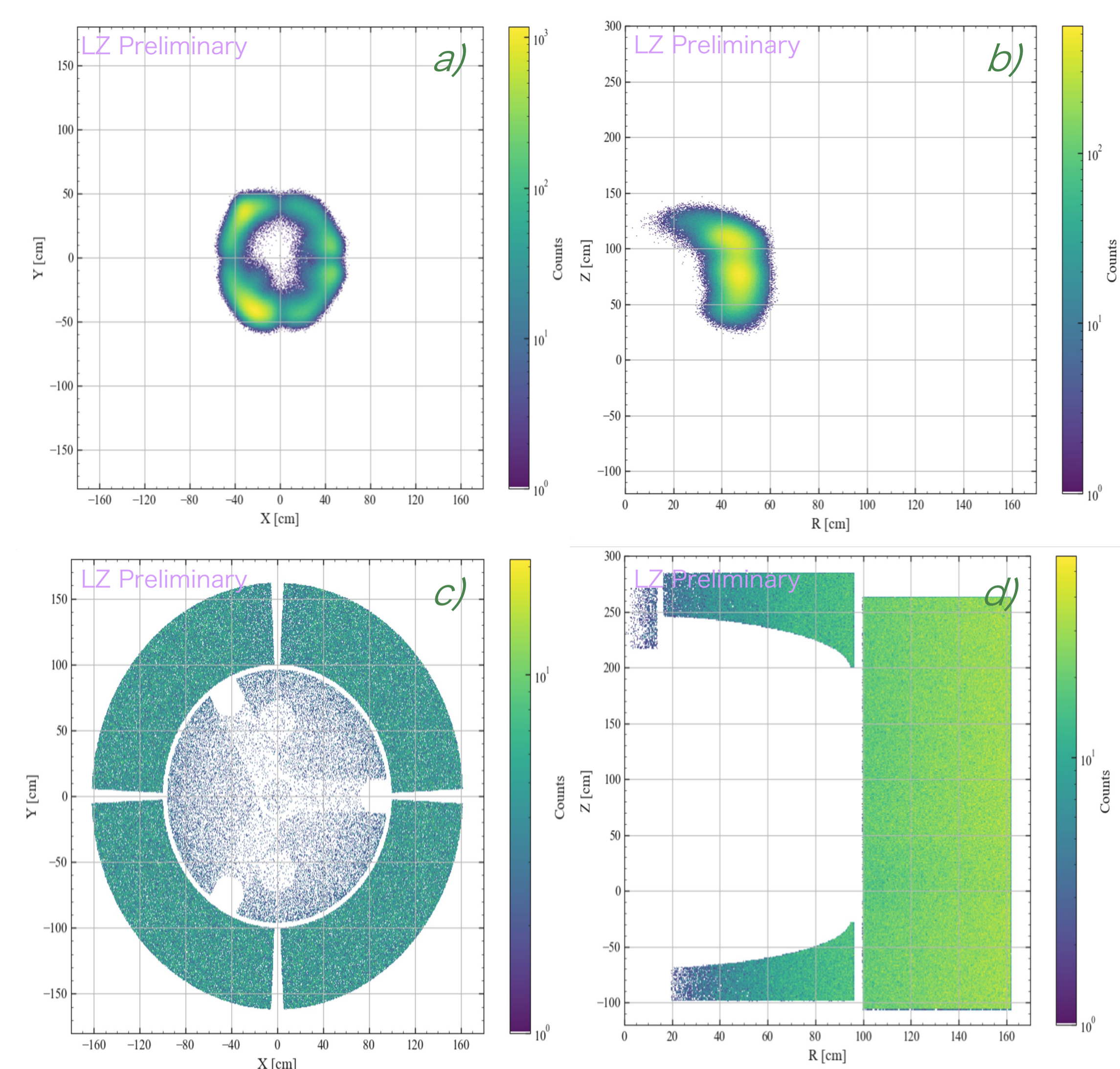


Figure 4: Event positions in XY (left) and RZ (right) from a,b) preWS2024 ²²⁸Th calibration data using CoG method, c,d) photon bomb simulations using truth positions.

Convolutional Neural Networks

- CNN designed to make predictions of event positions in (R, θ , Z) within the OD scintillator tanks.
- Two channel input.
 - OD PMT pulse areas.
 - OD PMT pulse timings.

$$\text{Timing} = \text{Time}_{\text{Channel Peak}} - \text{Time}_{e5\% \text{ Area Fraction}}$$

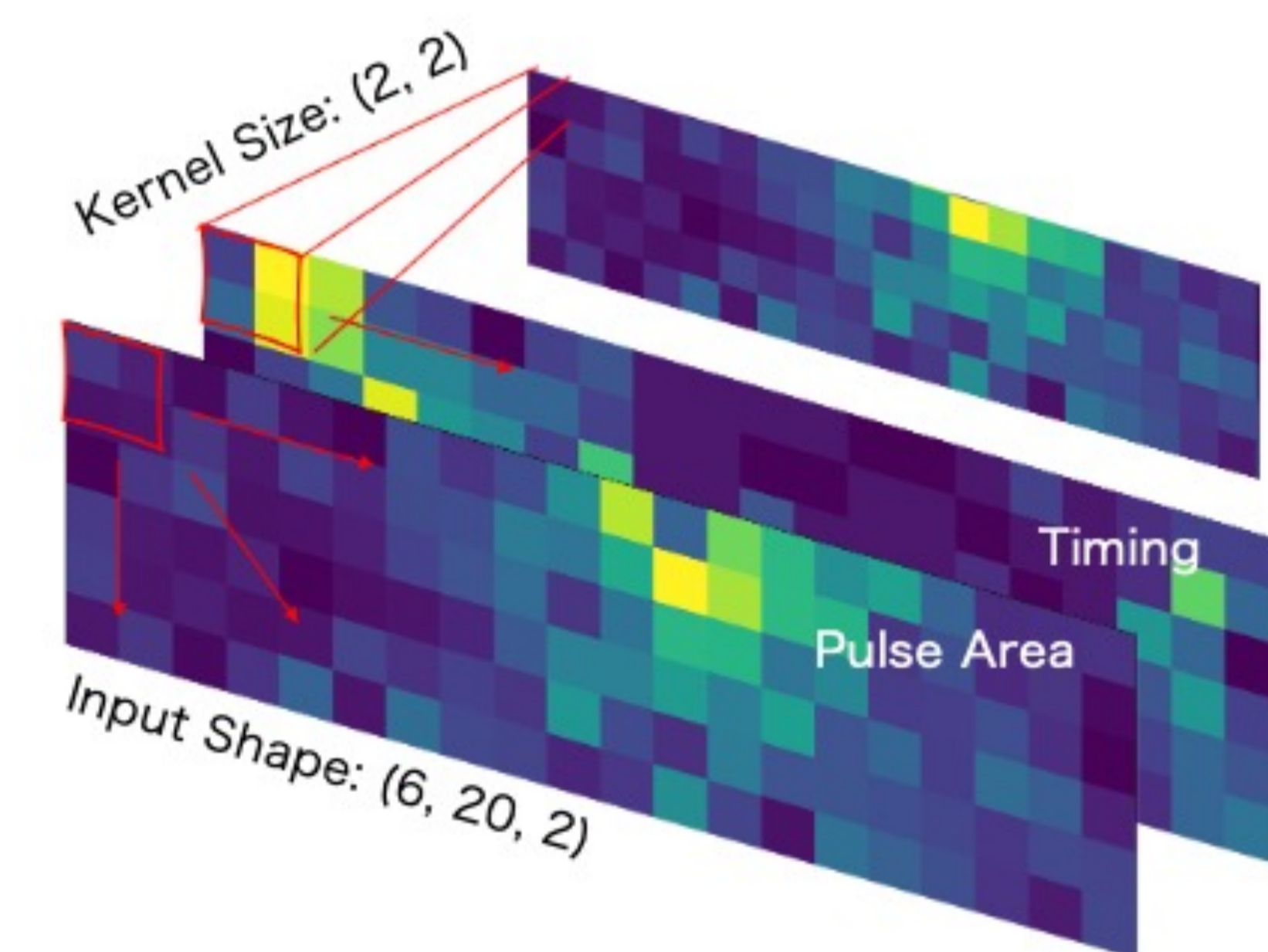


Figure 5: Schematic diagram of CNN. Channel 1 is the PMT pulse area, channel 2 is the respective PMT pulse timing information.

- Custom loss function:

$$\text{Loss} = \frac{1}{N} \sum_{i=1}^N \left(\frac{|\Delta R_i|}{\sigma_R} + \frac{|\Delta \theta_i|}{\sigma_\theta} P_{\theta,i} + \frac{|\Delta Z_i|}{\sigma_Z} \right)$$

$$P_{\theta,i} = \begin{cases} 1 & \text{if } \theta_{pred,i} \in [-\pi, +\pi] \\ \frac{\theta_{pred} - \pi}{\pi} + 1 & \text{if } \theta_{pred,i} \notin [-\pi, +\pi] \end{cases}$$

- N is the number of samples in a batch.
- $\Delta y_i = y_{true,i} - y_{pred,i}$.
- $P_{\theta,i}$ is the angular penalty.
- σ_x is a pre-defined position resolution.

- Obtained an overall OD scintillator tank (R, θ , Z) position resolution of (8.35 cm, 0.16 rad, 12.89cm)

Future Work

- Expand training data provided to the CNN.
 - Flat OD gamma simulations.
 - Calibration data.
- Add a fourth output, (R, θ , Z) \rightarrow (R, θ , Z, E).
 - Energy dependent component loss denominators.

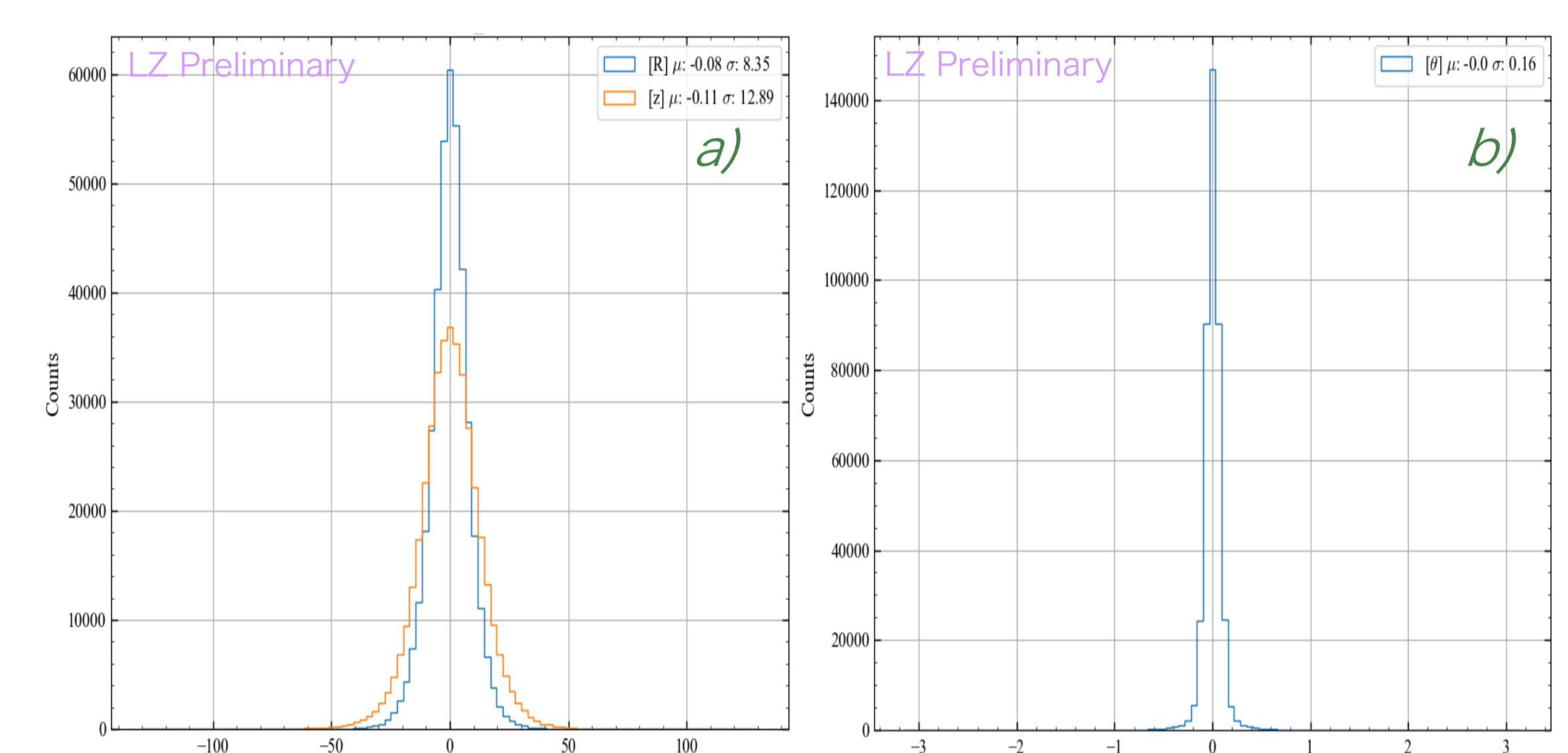


Figure 6: Plotted differences between the truth and model predicted event positions in a) R (blue) and Z (orange), and in b) Theta θ (blue). μ is the mean, and σ is the standard deviation describing the position resolution across the LZ outer detector.

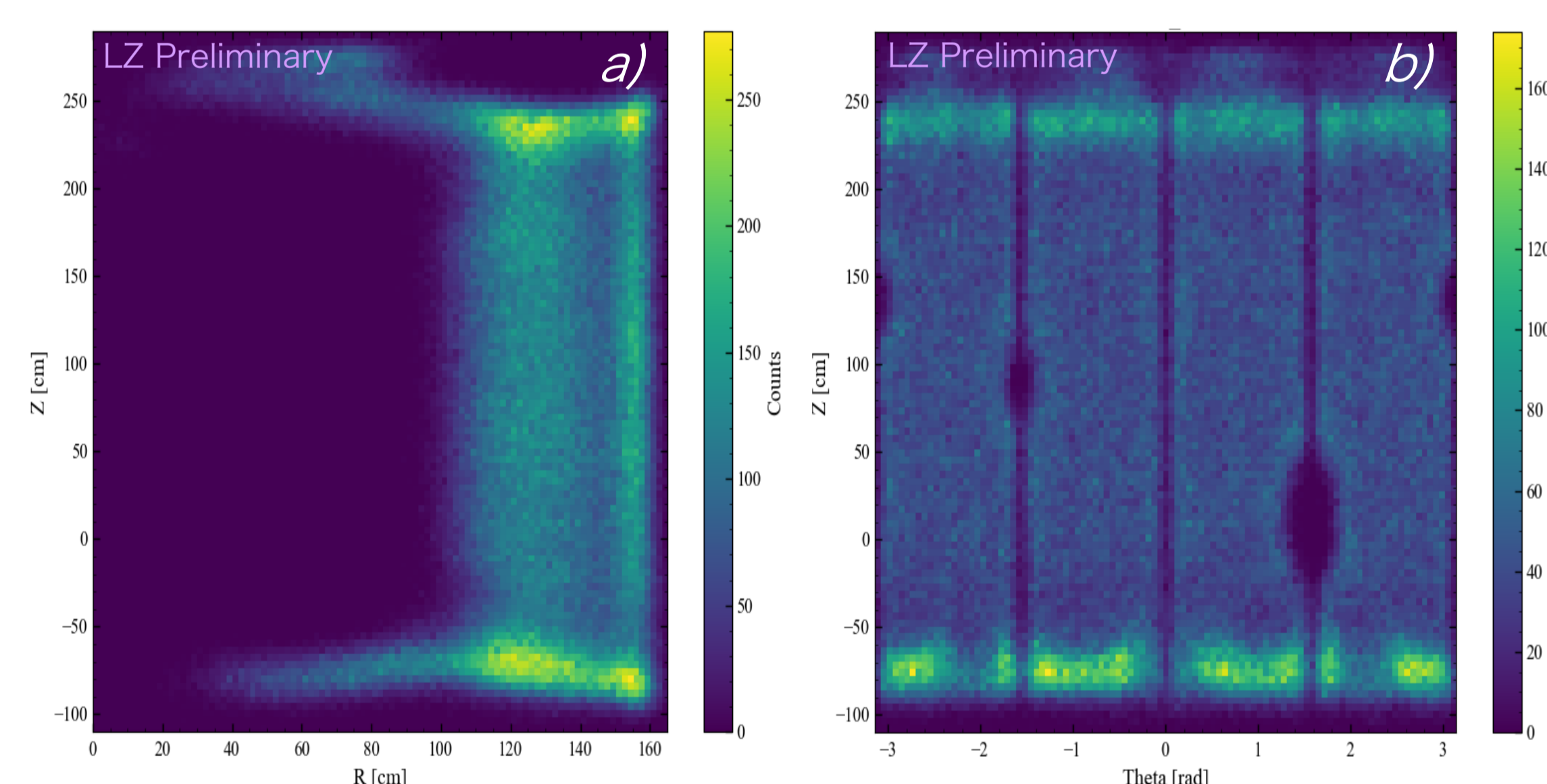


Figure 7: CNN model predicted positions in a) RZ space and b) θ Z space.

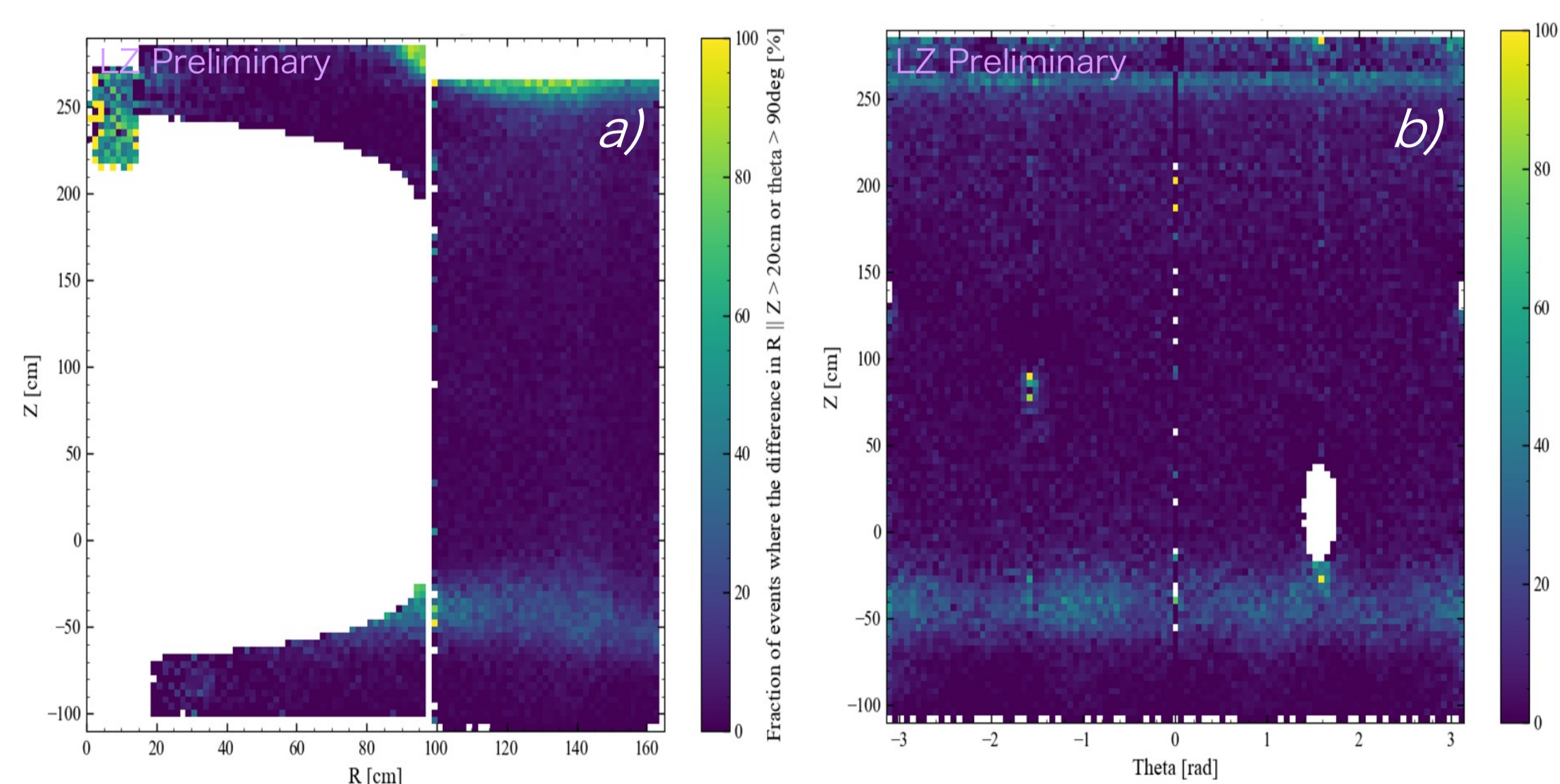


Figure 8: Fraction of events at their truth position which differ from their corresponding model predicted position by greater than 20cm or 90 degrees ($\pi/2$ rad) in a) RZ space and in b) θ Z space. These indicate regions of poor model predictions within the OD.

References

1. A. Cottle. 2024. "New Dark Matter Results from the LUX-ZEPLIN Experiment". LIDINE 2024. Sao Paulo, Brazil.
2. S. Haselschwardt. 2024. "New Dark Matter Results from the LUX-ZEPLIN Experiment". TeVPA 2024. Chicago, USA.
3. B. J. Mount. et al. 2017. "LUX-ZEPLIN Technical Design Report". LUX-ZEPLIN Collaboration. arXiv:1703.09144