# Electromagnetic counterparts of GW sources

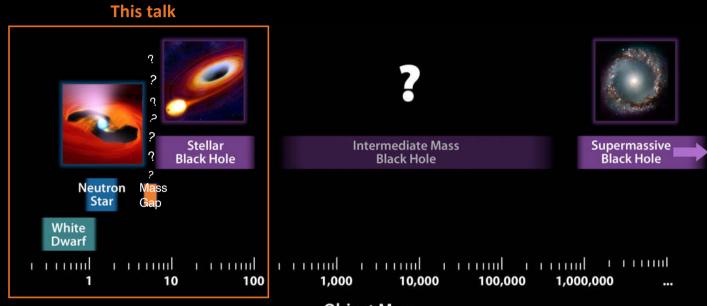
Eleonora Troja University of Tor Vergata





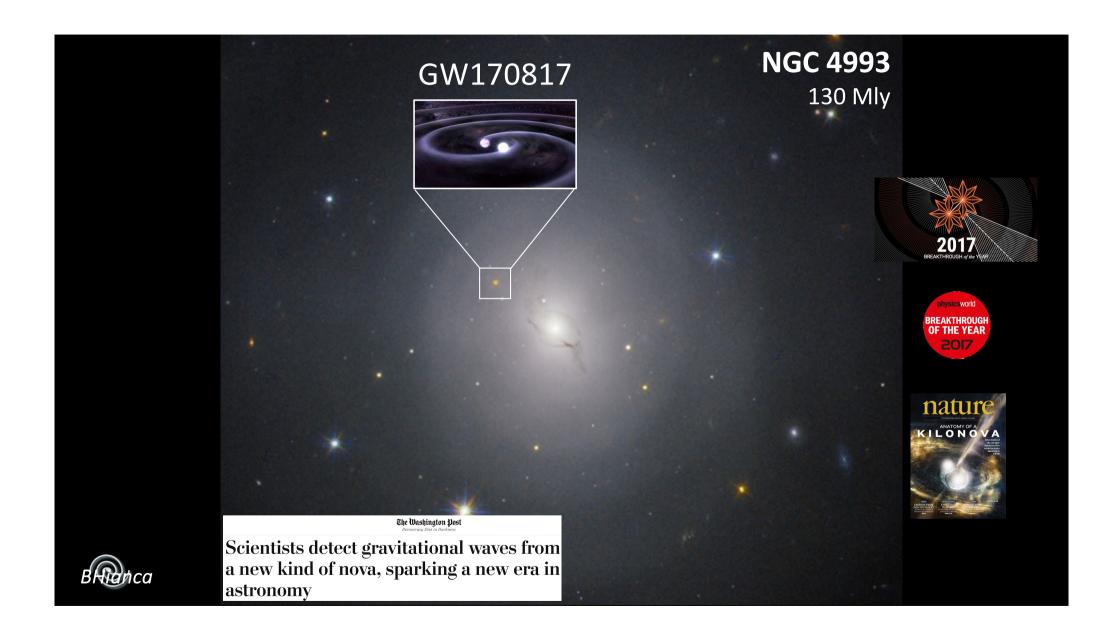


# **Observed Mass Ranges of Compact Objects**

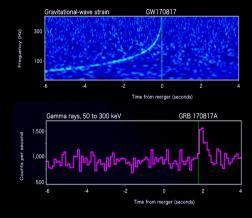


**Object Mass** (Relative to the Sun)

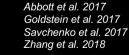




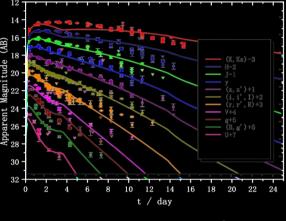
# A GRB followed by a kilonova



#### GRB170817A



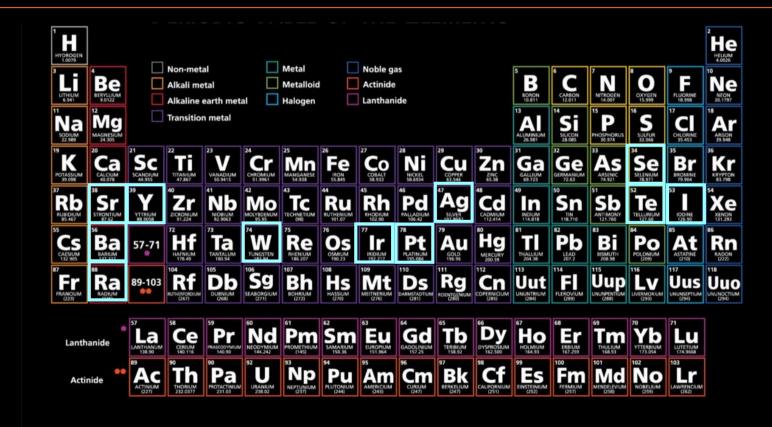




The kilonova AT2017gfo

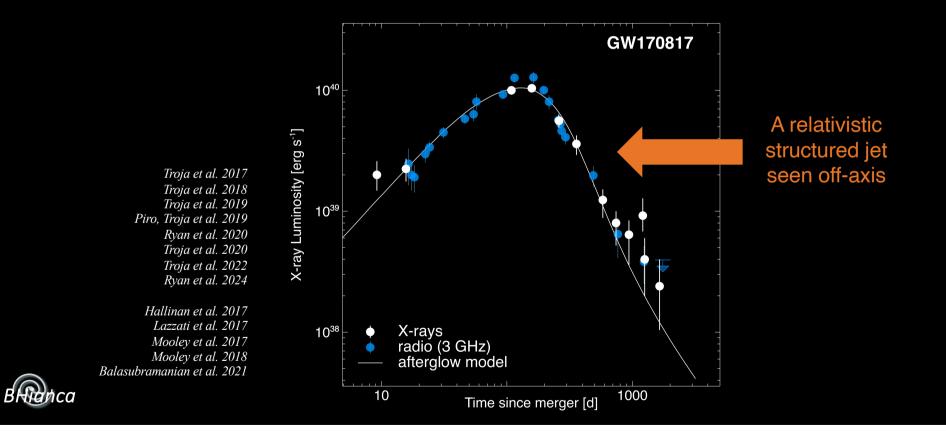
Coulter et al. 2017, Valenti et al. 2017, Drout et al. 2017, Arcavi et al. 2017, Lipunov et al. 2017, Soares-Santos et al. 2017, Tanvir et al. 2017, Andreoni et al. 2017, Smartt etal. 2017, Chornock et al. 2017, Kasliwal et al. 2017, Pian et al. 2017, Troja et al. 2017, Cowpertwaite et al. 2017,...

#### Cauldron of heavy elements



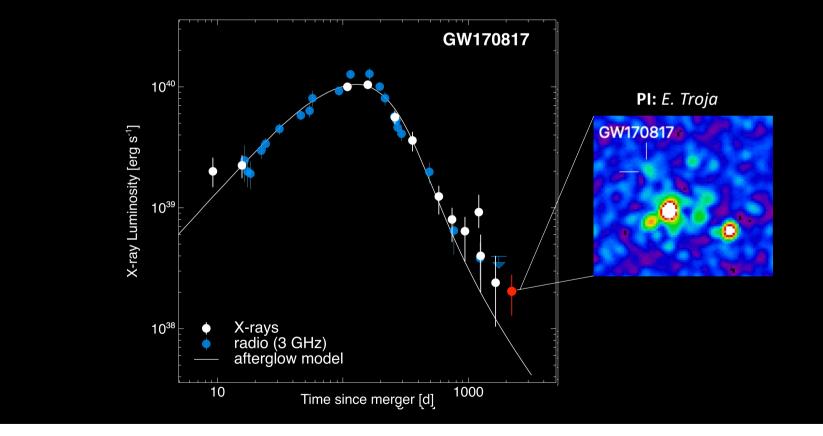


# Last but not least, the broadband afterglow

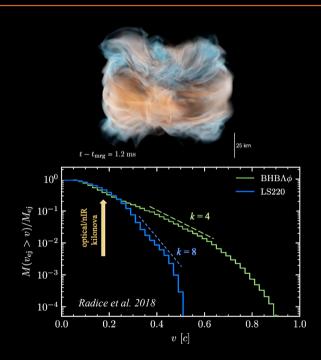


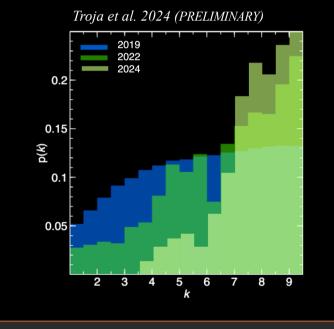
# GW170817, lately

BRanca



# Beyond the GRB jet: novel constraints on the ejecta



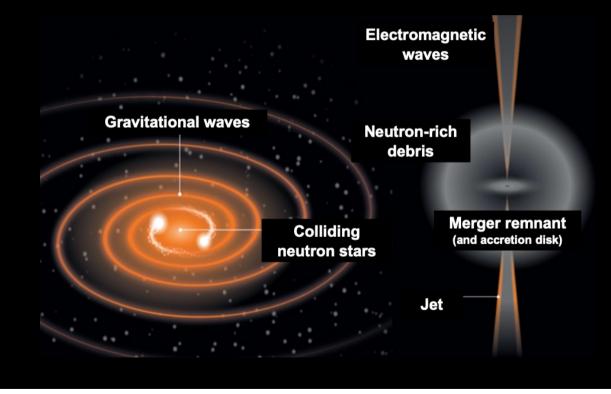


Exploring a phase of dense matter otherwise not accessible



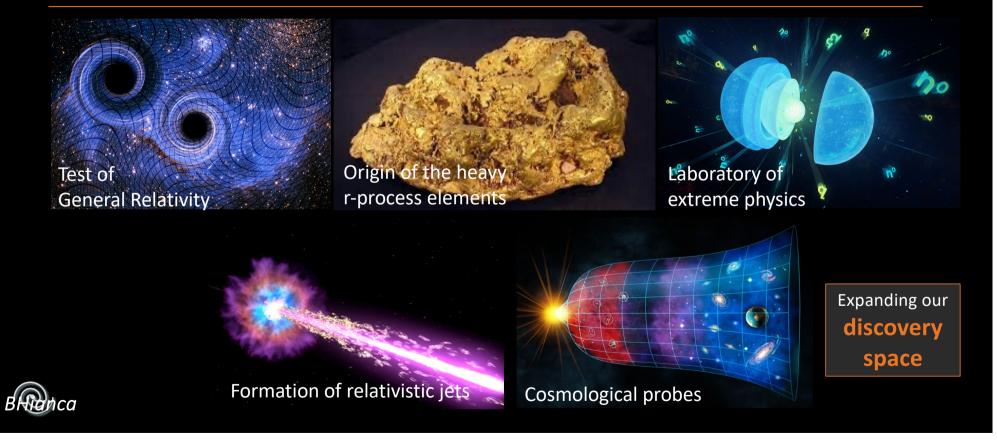
# Why multi-messenger astronomy?

All messengers are needed to complete the big picture of cosmic explosions





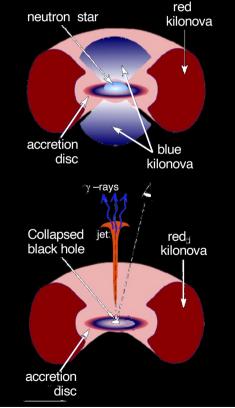
# Key questions of modern astrophysics



# Outstanding questions

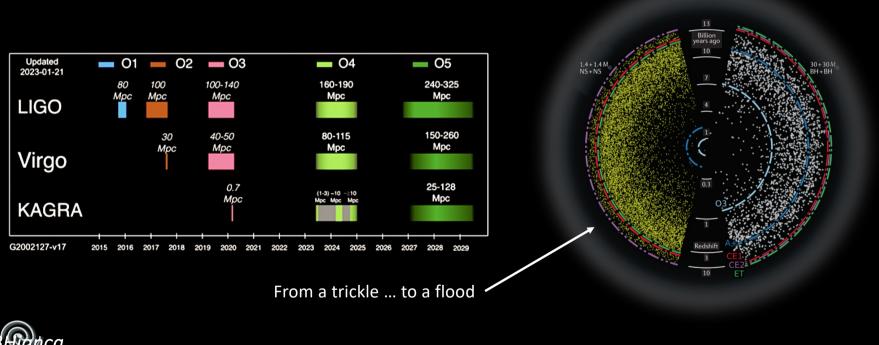
- Do all NS mergers produce relativistic jets and GRBs like GRB170817A?
- Do all NS mergers forge large amounts of heavy elements? How robust is the r-process production?
- What is the remnant of a NS merger, a massive NS or a prompt BH? Does it produce FRBs?
- How do NS-BH mergers look like?
- In which conditions binary BHs emit light?

We need MORE EVENTS to map and understand the diversity of initial conditions and outcomes.



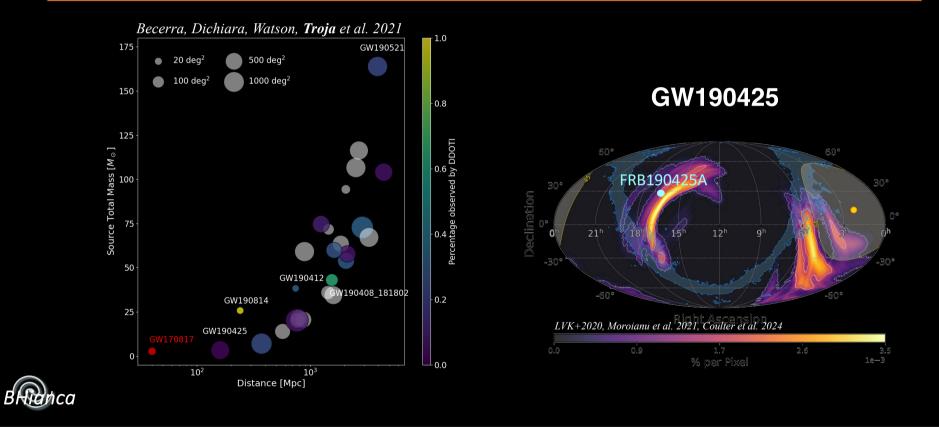


#### Finding new GW sources

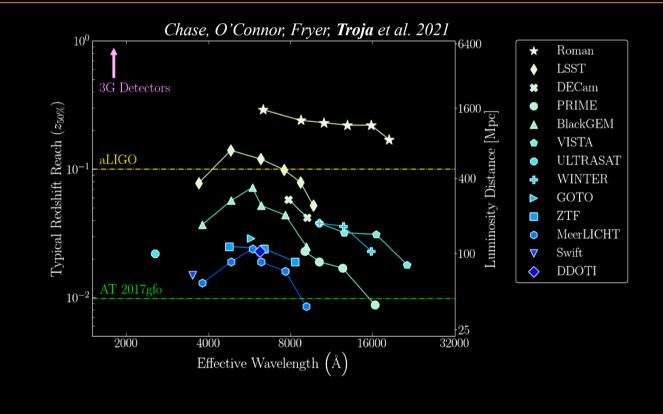




# Localizations, localizations, localizations

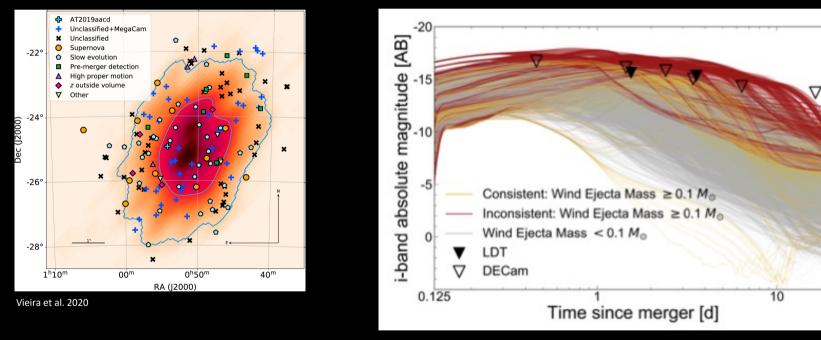


#### Kilonova detectability





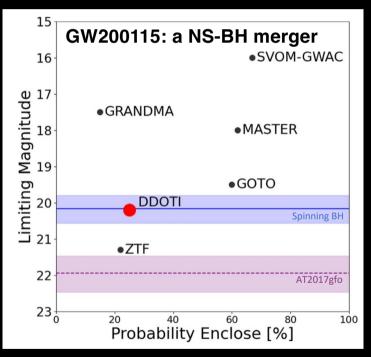
# Constraints from GW follow-up: GW190814

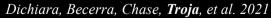


Thakur, Dichiara, **Troja**, et al. 2020



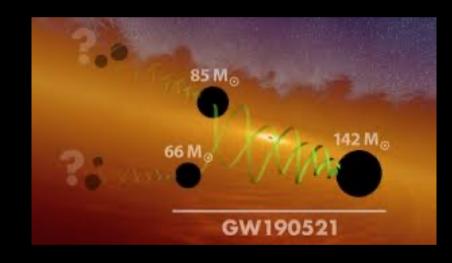
# Constraints from GW follow-up: GW200115

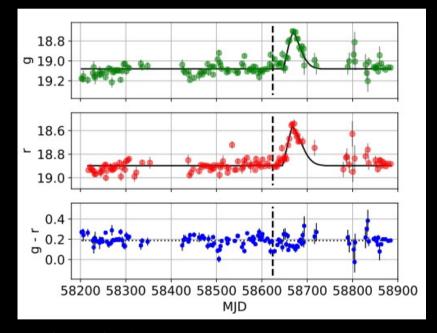






# Constraints from GW follow-up: GW190521

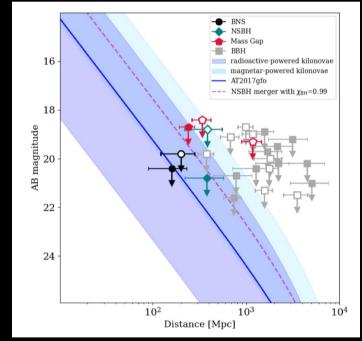




Graham et al. 2021



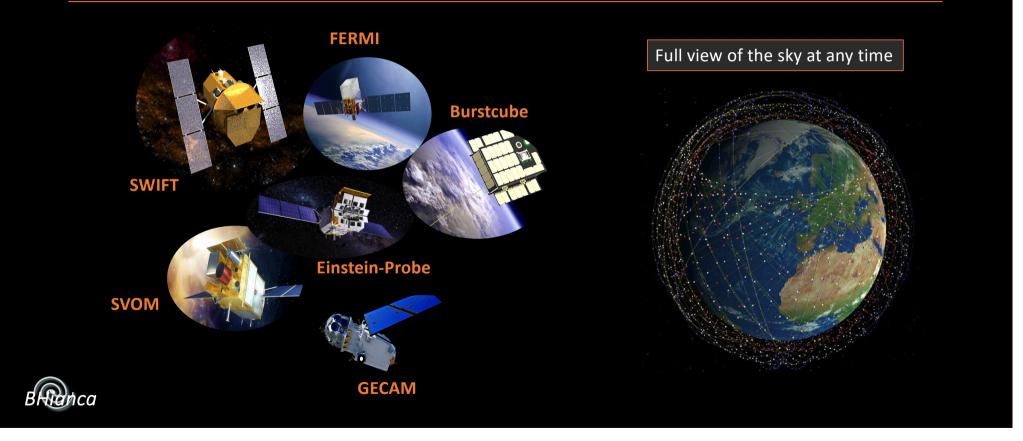
# Constraints from GW follow-up: summary



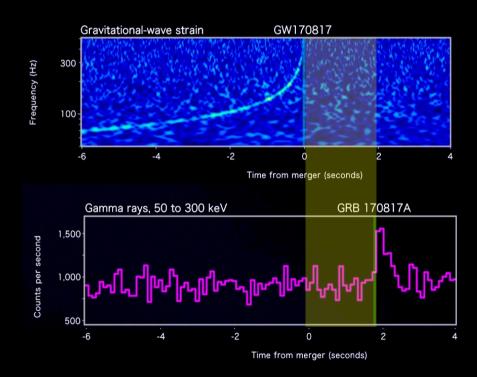
Becerra, Dichiara, Watson, Troja, et al. 2021



# GW sources at high-energies



#### The GRB/GW connection

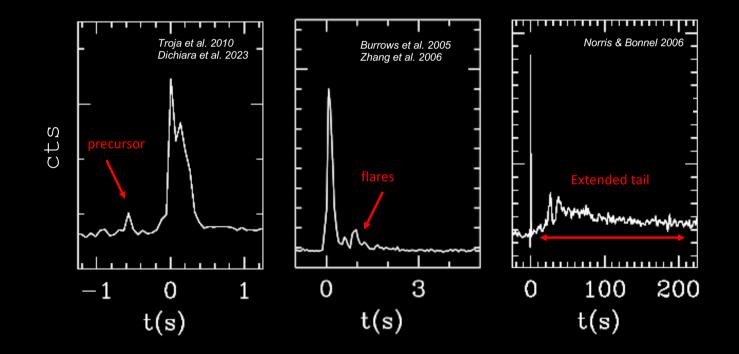


#### Why a 1.7 s delay?

- Viewing angle?
- Jet-launching?
- NS collapse?
- Modified gravity?

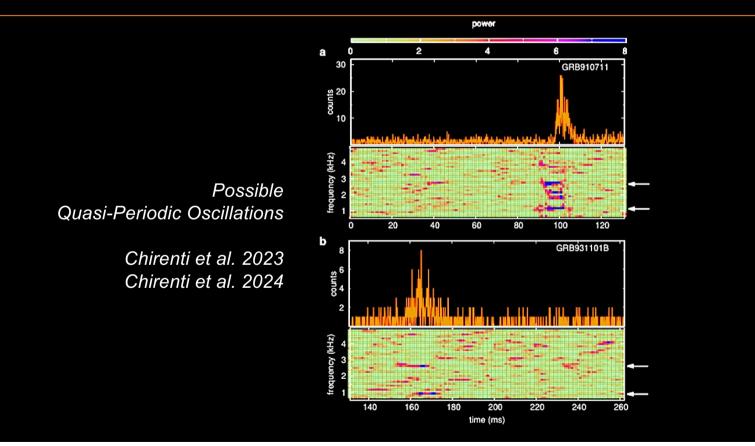


# A new look at the gamma-ray emission



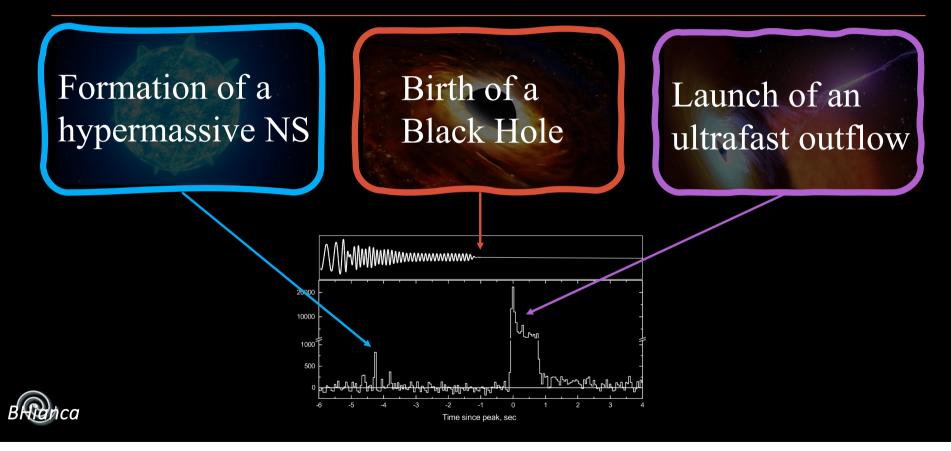


# A new look at the gamma-ray emission

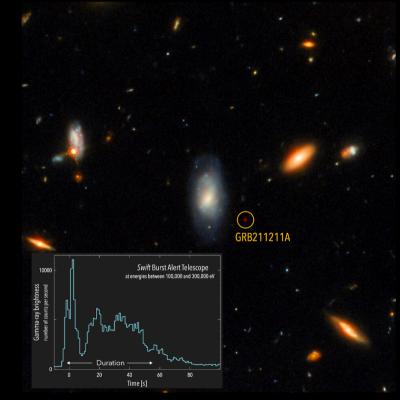




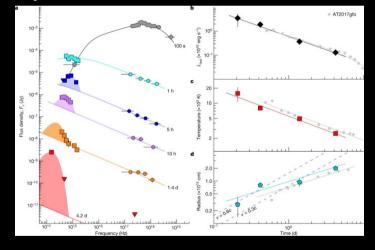
# Stellar forensics: a multi-messenger approach



# New Enigma: Long GRBs from NS mergers



Troja et al., Nature, 2022



See also Yang et al. 2022, Rastinejad et al. 2022, Yang et al. 2024, Levan et al. 2024



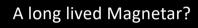
# Who will solve the enigma?

#### LIGO and Virgo

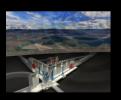




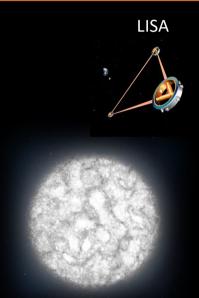
Neutron Star - Black Hole merger?







Einstein Telescope



White Dwarf – Neutron Star merger?



# Summary and final thoughts

The combination of multiple messengers is the **most powerful tool** to explore our Universe. EM observations can break the degeneracy of some GW parameters (distance, inclination, merger remnant). To fully realize the promise of multi-messenger astronomy, electromagnetic facilities must keep up with the advancement of GW detectors

Our knowledge will be built upon a few key events in the next few years, but the **rate of events** might drammatically increase in the next decade

Expect the unexpected: GW170817 was one NS-NS merger, a wide range of diverse outcomes is likely





# Thanks!