"Who Ordered That?" The Dance of the Muon

Tao Han Pitt PACC Pitt PACC
University of Pittsburgh
June 6, 2024 June 6, 2024

University of Bologna

CENTER

WHO ORDERED THAT?!

• A surprising discovery

In 1936, Anderson & Neddermeyer used a cloud chamber and found a (\pm) charged particle in the cosmic ray: m_e < $M < m_{p+1}$

- Pair production from γ^*
- Highly ionizing: a slow-motion $M \sim 200$ me-
- Decaying to electron with a lifetime

$\tau \sim 2 \times 10^{-6}$ s

• Highly penetrating: no strong interactions, thus NOT Yukawa's "meson" to mediate p*+, n* strong nuclear force.

"Who ordered THAT ?!" -- I. I. Rabi (1944 Nobel Laureate)

• The "usher" for particle physics: -- there are two "mesons"!

In 1948, Lattes et al. used a photographic emulsion detector, observing two charged particles: $\pi^{\pm} \rightarrow \mu^{\pm} \rightarrow e^{\pm}$

The discovery of μ^{\pm} led to another discovery of π^{\pm} ! -- both particles named by Lattes.

> The $\mathbf{\pi}^{\pm,0}$ are the Yukawa mesons mediating the strong force!

(Yukawa, 1949) (Powell, 1950)

Neutrons and μ^{\pm} decays involve the weak nuclear force:

Lederman et al. & Friedman et al. made it with polarized muons: dN_e $\frac{dN_e}{d\cos\theta}\propto 1+\vec{p_e}\cdot\vec{n}_{\mu}\sim 1-\frac{1}{3}$ 3 $\cos\theta_e$ Phys. Rev. 105, 1415 (1957); Phys. Rev. 106, 1290 (1957)

• Heavy Quark Onia:

"Lederman's shoulder": 1968-1969@BNL di-muon expt: ushered the $J/\psi(c\bar{c})$ Discovery

Fig. 1 - Dimuon yield from the 1968 BNL experiment. Source: J. H. Christenson, et al. (1970). Observation of massive muon pairs in hadron collisions. Physical Review Letters, 25(21), 1523.

"Oops-leon" (Leon Lederman, 1976) the $\mathbf{Y}(b\bar{b}) \rightarrow \mathbf{\mu}^+ \mathbf{\mu}^-$ Discovery

(1976)

• Two flavors of neutrinos:

• Neutrino Deeply Inelastic Scattering In addition to $e^{\pm} + N \rightarrow^{(\gamma^*)} e^{\pm} + \text{hadrons}$: Phys. Rev. Lett. 9, 36 (1962) Lederman-Schwartz-Steinberger made the v_μ beam! And there " ν " is NOT ν_e ! $\pi^{\pm} \rightarrow \mu^{\pm} + \nu \Rightarrow \nu + N \rightarrow \mu^{\pm} + N'$ (1). Muon flavor identified \rightarrow flavor physics, before Cabibbo ! --- Two-neutrino mixing by S. Sakata: Prog. Theor. Phys. 28 (1962) 870. Discovery of "weak neutral current" $F_2^e(x) = 2xF_1^e(x) = x$ $\sum e_q^2 [q(x) + \bar{q}(x)]$ *q* Precisions: $F_2^{\nu}(x) = 2xF_1^{\nu}(x) = x\sum [d(x) + \bar{u}(x)],$ $F_3^{\nu}(x) = 2[d(x) - \bar{u}(x)]$ $s(x)-\bar{s}(x),\;V_{cd},\;V_{cs},\; \sin\theta_W$ (1988) (DIS, 1990) $\nu_{\mu} + N \rightarrow^{(Z)} \nu_{\mu} + \text{hadrons}$ (2). Neutrino beam and scattering opens a new avenue

• Who ordered That? A Tri-peat

THE SEARCH FOR MUON-ELECTRON DIFFERENCES

AND MUON-PROTON DEEP INELASTIC SCATTERING*

Martin L. Perl

Stanford Linear Accelerator Center Stanford University, Stanford, California !

THE DISCOVERY OF THE TAU LEPTON*

 (1975)
Martin L. Perl Stanford Linear Accelerator Center Stanford University, Stanford, California 94309

Consequently, that led to yet another discovery: DONuT experiment @ FNAL (2000) detected the v_{τ}

 \rightarrow three generations of leptons completed! → 2022 Panofsky Prize!

Byron Lundberg Kimio Niwa Regina Rameika Vittorio Paolone

9

FNAL: "the muon campus" Continuing Explorations

The muon anomalous magnetic moment

Fermilab Muon Department

• **Paul Sherrer Institute (PSI), Laboratory for Particle Physics**

- AICap: µ-e conversion
- FAST: muon lifetime
- Lambshift in muonic hydrogen
- MEG: $\mu \rightarrow e\gamma \omega$ 10⁻¹⁴ level
- µ3e: flavor violating decay
- μ -Mass: muonium transition
- MUSE: **µ**-proton scattering
- PIONEER: pion decays

muon physics leader in Switzerland

11

test neutrino Majorana nature

$$
\mu^- + N(A,Z) \to e^+ + N'(A,Z-2)
$$

(referred to as $\mu^- \rightarrow e^+$)

• Lepton Flavor non-universality

Meson's 2-body decays scale with m_l^2

Meson's 3-body decays, like the β -decay, are flavor-universal in the SM:

 \rightarrow Universality in *K,D,B* meson decays has been a sensitive test to the SM, and possible new physics BSM.

For reference on flavor anomalies, e.g., e-Print: [1704.05435](https://arxiv.org/abs/1704.05435)

• Nu-Storm @ FNAL / PIP-II

Schematic layout of the nuSTORM facility

nuSTORM's physics program: Three themes

The physics program for the nuSTORM facility encompasses three central themes.

- 1. The neutrino beams produced at the nuSTORM facility will enable short-baseline (SBL) oscillation searches for light-sterile neutrinos with unprecedented sensitivity over a wide parameter space and, if sterile neutrinos are discovered, offers the opportunity to carry out an extremely comprehensive study of their properties.
- 2. These same beams may be exploited to make detailed studies of neutrino-nucleus scattering over the neutrino-energy range of interest to present and future long-baseline (LBL) neutrino oscillation experiments such as $T2HK$ (7), LBNE (8) and LBNO (9).
- 3. The storage ring itself, and the muon beam it contains, can be used to carry out a R&D program that can facilitate the implementation of the next step in the incremental development of muon accelerators for particle physics.

• SBNE: Short-Baseline Neutrino Experiments

- Pioneer the Liquid Argon TPC technology
- Resolve LSND / MiniBooNE anomalies
- Search for (light) new physics

• LBNE: Long-Baseline Neutrino Experiments DUNE: Deep Underground Neutrino Experiment, the "ultimate" neutrino experiment

Neutrinos can travel long distances through rock and other matter without a scratch. The LBNE neutrino beam will travel 800 miles straight through the earth from Batavia, Illinois, to the Sanford Lab in Lead, South Dakota-no tunnel necessary. The trip will take less than one-hundredth of a second, enough time for some of the muon neutrinos to transform into electron neutrinos and tau neutrinos. Scientists call this process neutrino oscillation.

DUNE will pursue major science goals:

- Leptonic CP violation, precision measurements of θ_{13} , Δm^2_{13}
- Dark matter searches
- Proton decay
- Supernova, formation of neutron star/black holes

MUONS BEYOND HEP

• Muonic atom for precision physics

- $r_B = (\alpha \text{ m})^{-1}$: $r_B(\mu) = r_B(e^-)/207 = 2.2 \times 10^{-5} \text{ nm}$
- Wavefunction overlap: $(m_{\mu}/m_e)^{3}$ ~10⁷ stronger
- Lamb shift: 10⁵ larger

Very sensitive to probe the **proton size**/properties: PSI, CREMA Best precision for the helium nucleus size: $r_{\alpha} = 1.67824(13)_{exp}$ (82)_{theo} fm.

• Muons for material science: PSI

SuS: Swiss Muon Source

µSR - Muon Spin Rotation, Relaxation or Resonance: A research tool using muons as sensitive local magnetic probes in matter.

Research at the LMU focuses mainly on magnetic properties of materials and on positive muons or muonium (bound state of a positive muon and an electron) as light protons or hydrogen substitutes in matter.

Worldwide unique: The Low-Energy Muon Beam and µSR Spectrometer for the study of thin films, layers and surfaces, the possibility to perform high-field µSR with a field up to 9.5 Tesla, and the Extraction of Muons On Request for high frequency resolution and slow relaxation measurements.

• Muons for material science: KEK

KEK, Japan:

Muon Science Laboratory

Institute of Materials Structure Science High Energy Accelerator Research Organization, KEK

One important experimental technique which the team uses is called Muon Spin Rotation / Relaxation / Resonance (μ SR). μ SR is used to map magnetic fields inside matter on a nanometer scale by means of muons shot into samples. Using this technique, scientists can examine the magnetic properties of materials. For example, they can examine the magnetic flux through type-Il superconductors, and [determine][simulate][?] the location of the trace amounts of hydrogen atoms contained in some materials. Other examples include studies of muon-catalyzed fusion and the non-destructive analysis of the interior of solids, which takes advantage of the fact that negatively charged muons behave as heavy electrons.

Transmission Muon Microscope TMM @ KEK Advantages of TMM versus an electron microscope

• **Muon Tomography: Cosmic muons** Atmospheric muon flux ~ $200/m^2/s$ Muons are penetrating!

nature

(November 2, 2017, and update Jan. 2023)

Cosmic-ray particles reveal secret chamber in Egypt's Great Pyramid

A previously unknown chamber has been found in the largest of the pyramids in Giza, Egypt. Credit: Tomasz Tomaszewski/VISUM creativ/eyevine

A Muon Collider Who Ordered That Collider? $\overline{}$ **Ordered That Col** "Who or derected that is not that is not the contract of the c A MUON COLLIDER $\overline{1}$ and $\overline{1}$ in understanding particle physics.

aring the sam
isn't another ISIT L ANULIER EIECLIUM. Although sharing the same EW interactions, it isn't another electron:

 $m_\mu \approx 207 \, m_e$ $\tau(\mu \to e \bar{\nu}_e \nu_\mu) \approx 2.2 \ \mu s$ $c\tau \approx 660$ m.

It is these features: heavy mass, short lifetime that dictate the physics.

• **Advantages of a muon collider** Advantages of a Muon Collider \mathbf{A} Muon Collider of a Muo

• Much less synchrotron radiation energy loss than e's:

 $\Delta E \sim \gamma^4 = (\frac{E}{m})$

 $\Delta E \sim \gamma^4 = (\frac{E}{m_t})$

 m_{μ}

 $)^4$

 $)^4$

- Unlike the proton as a composite particle, E_{CM} efficient in $\mu^+ \mu^-$ annihilation E_{CM} efficient in $\mu^+\mu^-$ annihilation
	- Much smaller beam-energy spread: $\Delta E/E \sim 0.01\% - 0.001\%$

• **Disadvantages of a muon collider**

• Production: Protons on target \rightarrow pions \rightarrow muons: Require sophisticated scheme for **µ** capture & transport

"Never play with an unstable thing!"

• Very short lifetime: in micro-second, Muons cooling in (x,p) 6-dimensions

 \rightarrow Difficult to make quality beams and a high luminosity

[Note: $E_{\mu} \sim 1 \text{ TeV} \rightarrow \gamma \sim 10^4 \rightarrow \gamma \tau = 0.02 \text{ s} \rightarrow \text{d} = 6,000 \text{ km}$]

• Beam Induced Backgrounds (BIB) from the decays in the ring at the interacting point, [Note: $\sigma_{\text{pp}}(\text{total})$ ~100 mb; $\sigma_{\mu\mu}(\text{total})$ ~100 nb]

• Neutrino beam dump (environmental hazard) $\sigma_{\nu} \sim G_{F}^2 E^2 \rightarrow$ Shielding?

The recent excitement: the "Muon Shot" Muon Accelerator Project (MAP)

[https://arxiv.org/abs/1907.08562,](https://arxiv.org/abs/1907.08562) J.P. Delahauge et al., arXiv:1901.06150/

U.S. P5 (Particle Physics Project Prioritization Panel) The path to 10 TeV pCM (partonic c.m. energy):

collider. At the end of the path is an unparalleled global facility on US soil This is our Muon Shot. **Multiple 19th and Media in the Mundeev of Although We do not know it a muon collider is ultimately teasible, the roa
Later Fermilab strengths and capabilities to a series of proton beam impro beam facilities**, each producing world-class science while performing between the solution of the part is an unparalleled alobal facility on LIS

Collider benchmark points: Collider benchmark points:

- $L \sim 1 \text{ fb}^{-1}\text{/yr}$ Beam Energy Spread $\%$ 0.004 $\Delta E_{cm} \sim 5 \text{ MeV}$ At **few TeV energy** one can still exploit high partonic energy for a striking • The Higgs factory: Parameter Units Higgs **CoM Energy TeV** 0.126 Avg. Luminosity $10^{34} \text{cm}^{-2}\text{s}^{-1}$ 0.008 Beam Energy Spread $\%$ 0*.*004 Higgs Production $/10^7$ sec 13'500 Circumference km 0.3 $E_{cm} = m_H$
- Multi-TeV colliders: Repetition Rate Hz 15 15 12 6

We can borrow **CLIC** physics case (see below) Lumi-scaling scheme: $\sigma L \sim \text{const.}$ ⇤ *x,y* cm 1*.*71 0*.*5 0*.*25 ng scheme: O $L \sim$ const.

$$
L \gtrsim \frac{5 \text{ years}}{\text{time}} \left(\frac{\sqrt{s}_{\mu}}{10 \text{ TeV}}\right)^2 2 \cdot 10^{35} \text{cm}^{-2} \text{s}^{-1}
$$

The conceivable choices: $E_{\text{max}} = Z T_{\text{c}} V I \sqrt{3} \text{GeV}/10 \text{TeV}^2 \text{6} \cdot 10^{35}$ If much less, we could only bet on Direct Discoveries ! $(3 \text{TeV}/10 \text{TeV})^2$ 6 $\cdot 10^{35}$ $Ecm = 3 TeV - 14³ TeV^{/10} TeV^{/0}$ parameters of the enabled facilities are summarized in Table 1.

European Strategy, arXiv:1910.11/75; arXiv:1901.06150; arXiv: European Strategy, arXiv:1910.11775; arXiv:1901.06150; arXiv:2007.15684.

Just like in hadronic collisions: $\mu^+\mu^ \rightarrow$ exclusive particles + remnants • **"Semi-inclusive" processes**

Costantini, F. Maltoni, et al., arXiv:2005.10289; Y. Ma et al., arXiv:2007.14300

• Underlying sub-processes:

Precision Higgs physics 10M H, 500K HH @ 10 TeV

Muon Collider Forum Report:<https://arxiv.org/abs/2209.01318>

Muon Collider Higgs Precision Projections (SMEFT)

Higgs pair production & triple coupling:

 \rightarrow dictate EW phase transition & impact $\begin{bmatrix} 1 & 1 \\ 1 & 1 \end{bmatrix}$ on early universe cosmology!

Table 7: Summary table of the expected accuracies at 95% C.L. for the Higgs couplings at a variety of muon collider collider energies and luminosities.

TH, D. Liu³⁰I. Low, X. Wang, arXiv:2008.12204 D. Buttazzo, D. Redogolo, F. Sala, arXiv:1807.04743;

Pushing the "Naturalness" limit The searches for top quark partners (most wanted in "naturalness"); & gluinos, gauginos …

 \rightarrow Higgs mass fine-tune: $\delta m_H/m_H \sim 1\%$ (1 TeV/ Λ)² Thus, $m_{stop} > 8$ TeV \rightarrow 10⁻⁴ fine-tune!

• **WIMP Dark Matter (a conservative SUSY scenario)**

Consider the "minimal EW dark matter": an EW multi-plet

- The lightest neutral component as DM
- Interactions well defined \rightarrow pure gauge
- Mass upper limit predicted \rightarrow thermal relic abundance

Covering the thermal target **WIMP Dark Matter**

TH, Z. Liu, L.T. Wang, X. Wang: arXiv:2009.11287; arXiv:2203.07351 $\frac{33}{6}$

Summary: "Who ordered That?" **The muon is such a pleasant surprise Nature offers us!**

- **Leads to many discoveries**
- **Provides deeper understanding of Nature**
- **Continues to play a key role in going forward**

Far reaching scope: all involves with muon physics \rightarrow

m. Accel.
Unknowr

k Matter

trinos

Muon physics has taken many spot-lights in contemporary HEP! Look forward to more surprises with muons!

Lots of recent works! -- my apologies not to cover properly D. Buttazzo, D. Redogolo, F. Sala, arXiv:1807.04743 (VBF to Higgs) A. Costantini, F. Maltoni, et al., arXiv:2005.10289 (VBF to NP) M. Chiesa, F. Maltoni, L. Mantani, B. Mele, F. Piccinini, and X. Zhao, arXiv:2005.10289 (SM Higgs) R. Capdevilla, D. Curtin, Y. Kahn, G. Krnjaic, arXiv:2006.16277; arXiv:2101.10334 (g-2, flavor) P. Bandyopadhyay, A. Costantini et al., arXiv:2010.02597 (Higgs) D. Buttazzo, P. Paradisi, arXiv:2012.02769 (g-2) W. Yin, M. Yamaguchi, arXiv:2012.03928 (g-2) R. Capdevilla, F. Meloni, R. Simoniello, and J. Zurita, arXiv:2012.11292 (MD) D. Buttazzo, F. Franceschini, A. Wulzer, arXiv:2012.11555 (general) G.-Y. Huang, F. Queiroz, W. Rodejohann, arXiv:2101.04956; arXiv:2103.01617 (flavor) W. Liu, K.-P. Xie, arXiv:2101.10469 (EWPT) H. Ali, N. Arkani-Hamed, et al, arXiv:2103.14043 (Muon Smasher's Guide) Richard Ruiz et al., arXiv:2111.02442 (MadGraph5) … …