



Cogenesis of baryon and dark matter with PBH and QCD axion

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Introduction and The Framework

Motivation

- Apart from solving DM and neutrino mass problem, we aim to bring solution of strong CP problem and matter anti-matter asymmetry within **same framework**.
- We aim give cosmological observations of **high-scale leptogenesis**.
- We consider a type-I seesaw framework extended by Peccei-Quinn symmetry.
 1. Complex PQ scalar field : $\sigma \equiv \frac{v_{PQ} + \rho}{\sqrt{2}} e^{ia/f_a}$ ($\sigma \sim (1, 1, 0)$).
 2. A heavy quark : $Q \sim (3, 1, 0)$.
 3. 3 right-handed neutrinos : $N_R \sim (1, 1, 0)$.

Lagrangian

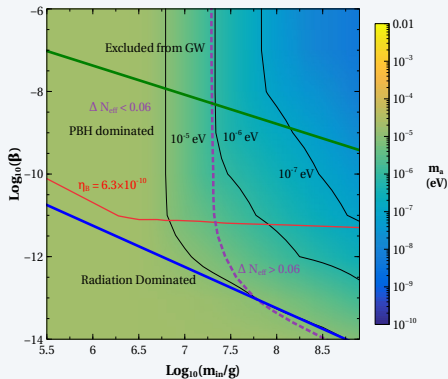
$$\mathcal{L}_Y = - \left[y \bar{Q}_L \sigma Q_R + G_{ij} \bar{L}_i H_{jR} + F_{ij} \bar{L}_i \tilde{H} N_{jR} + \frac{1}{2} y_{ij} \bar{N}_{iR}^c \sigma N_{jR} \right] + \text{h.c.}$$

- We take : $v_{PQ} = f_a \sim M_i$.

Interplay between Leptogenesis and Axion with PBH

- PBH open up new parameter space to probe.
- Form $m_{\text{in}} = \frac{4\pi}{3} \gamma \frac{\rho_R(T_{\text{in}})}{H^3(T_{\text{in}})}$ with initial fraction $\beta = \frac{\rho_{\text{BH}}(T_{\text{in}})}{\rho_R(T_{\text{in}})}$.
- Presence of PBH
 1. Changes T_{OSC} as $3M_{\text{p}}\mathcal{H}^2 = \rho_r + \rho_{\text{BH}}$.
 2. Dilutes existing axions abundance. (2209.14307).
- Baryon-to-photon ratio

$$\eta_B \approx 10^{-2} \kappa_1 \frac{\epsilon}{\xi}.$$



Hierarchical; axion 100% DM

DM : Either Axion or RHN

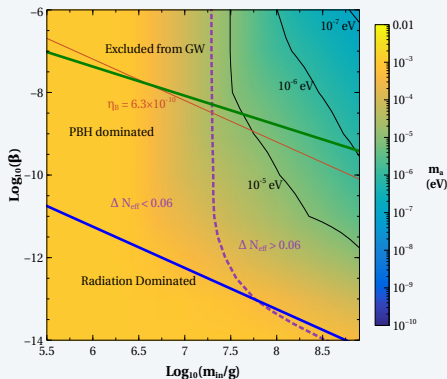
For axion : Vacuum Misalignment.

For RHN : i) PBH evaporation; **X** ii) Axion portal; **X** iii) Decay of W^\pm, Z, h ; **✓**.

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Resonant; axion DM 1%

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For axion : Vacuum Misalignment.

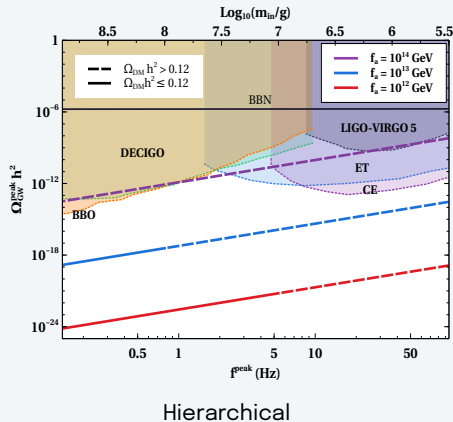
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Probes from Gravitational Waves

- PBH \implies GWs. (2012.08151)

Imposing Leptogenesis

- $f^{\text{peak}} \simeq 1.7 \times 10^3 \text{ Hz} \left(\frac{m_{\text{in}}}{10^4 \text{ g}} \right)^{-5/6}$.
- Hierarchical : $\Omega_{\text{gw}}^{\text{peak}} \simeq 6.29 \times 10^{-22} \left(\frac{f_{\sigma}}{10^{12} \text{ GeV}} \right)^{16/3} \left(\frac{10^7 \text{ g}}{m_{\text{in}}} \right)^{14/9}$.
- Resonant : $\Omega_{\text{gw}}^{\text{peak}} \simeq 1.71 \times 10^{-7} \left(\frac{\epsilon_1}{0.1} \right)^{16/3} \left(\frac{10^7 \text{ g}}{m_{\text{in}}} \right)^{14/9}$.

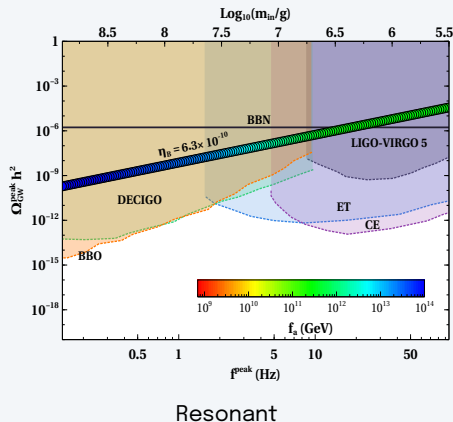


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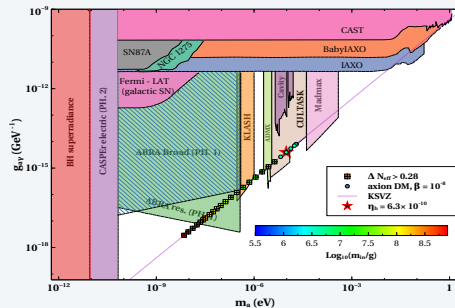
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Probes from Axion detection

- Axion-photon coupling is

$$g_{a\gamma} = -\frac{\alpha}{2\pi f_a} \left(\frac{2}{3} \frac{4m_d+m_u}{m_u+m_d} \right) = -1.92 \frac{\alpha}{2\pi f_a}.$$
- ★ indicates correct observed baryon asymmetry.



Conclusion

- This unified set up address strong CP problem, neutrino mass, DM and baryon asymmetry.
- **Resonant** enhancement of CP parameter may help in probing high-scale **leptogenesis** via future GWs experiments.
- **Axion** detection experiment can also be connected to high-scale **leptogenesis**.
- **Hot axions** from PBH (ΔN_{eff}) can get detected in future CMB S4, CMB-HD.



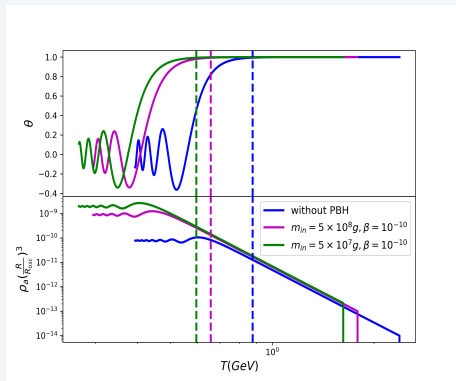
**Thank you for your attention
See you during the poster session.**

Nayan Das

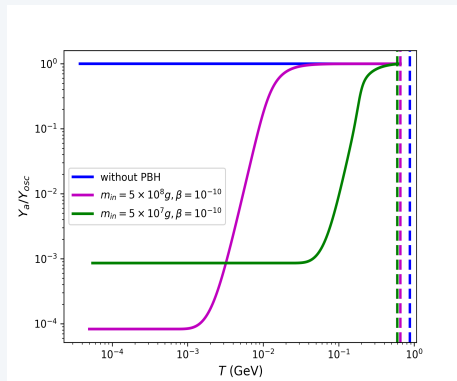
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Backup Slide : Axion evolution in presence of PBH



Effect on T_{osc}



Entropy Dilution