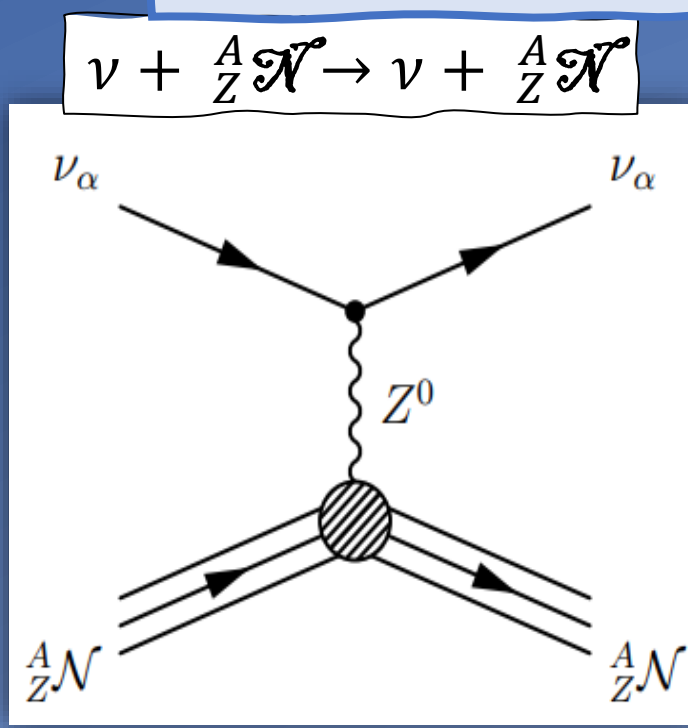
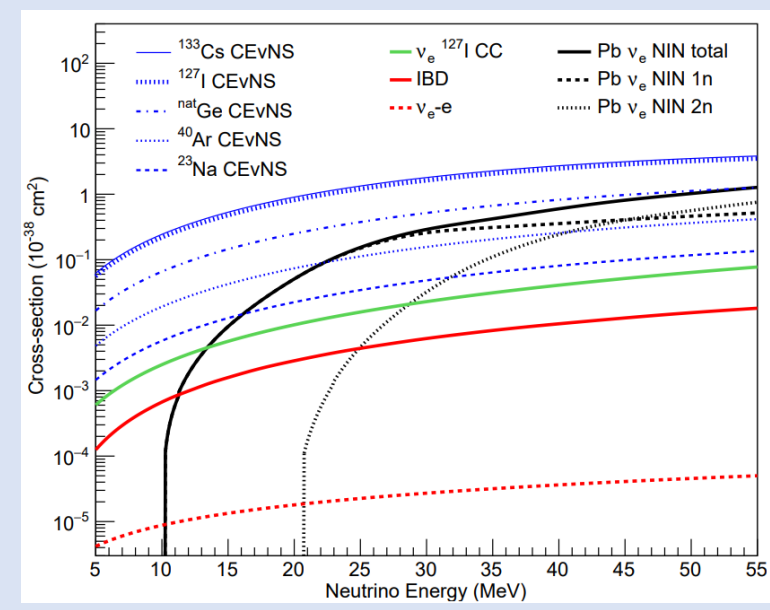
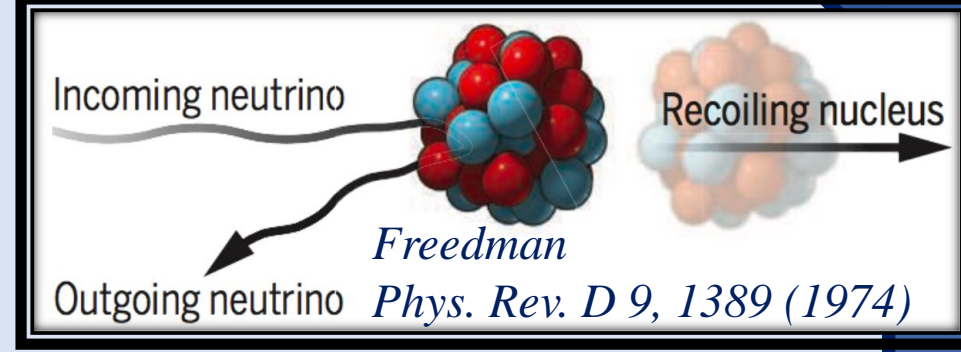


## CEvNS

### Coherent Elastic Neutrino-Nucleus Scattering Weak-neutral-current process

The de Broglie wavelength of the exchanged  $Z^0$  boson mediator is of the order of the radius of the nucleus.

- All the nucleons respond coherently (recoils as a whole)
- “Large” cross section ( $\propto N^2$ ) compared to other low energy neutrino interactions



Room to study **electroweak** and **nuclear physics**

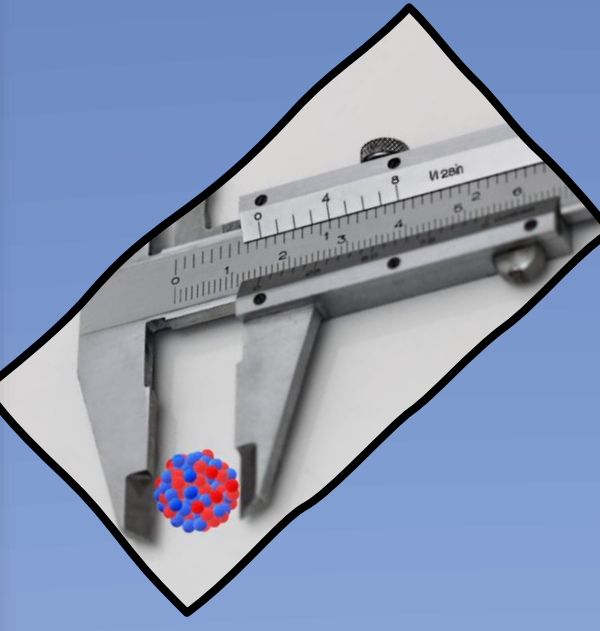
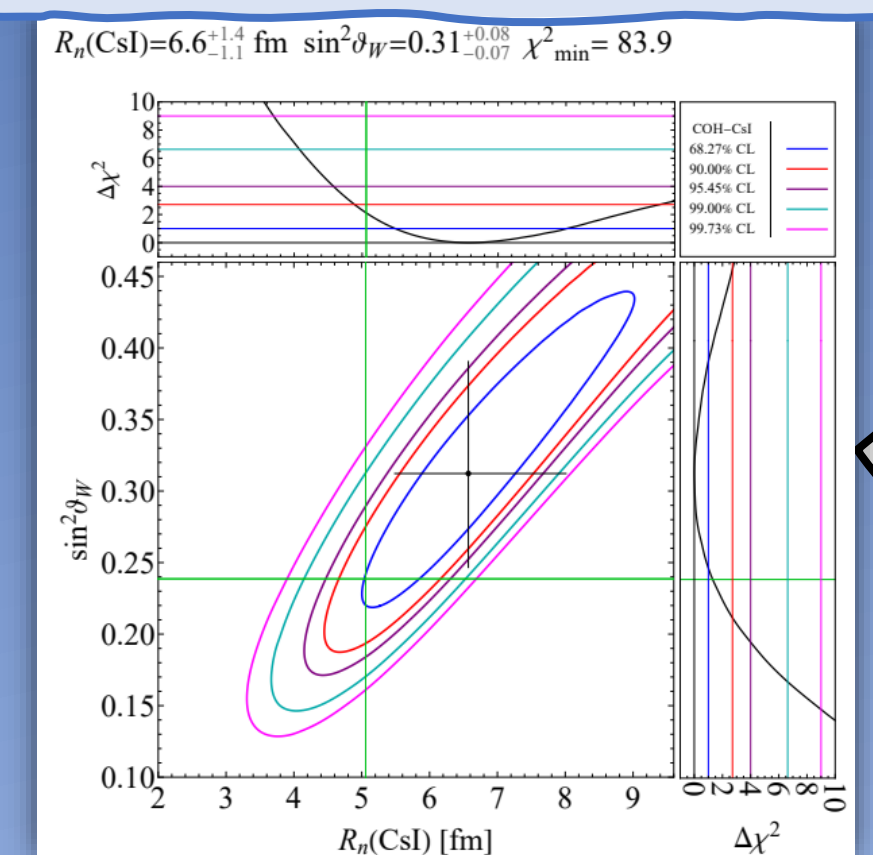
MAC. et al. Eur.Phys.J.C 83 (2023) 7, 683

## CEvNS Cross Section

$$\frac{d\sigma}{dT_{nr}} \cong \frac{G_F^2 m_N}{\pi} \left(1 - \frac{m_N T_{nr}}{2E_\nu^2}\right) [g_V^p(\sin^2\theta_W)ZF_Z(q^2) + g_V^n N F_N(q^2)]^2$$

$\sim -0.0227$

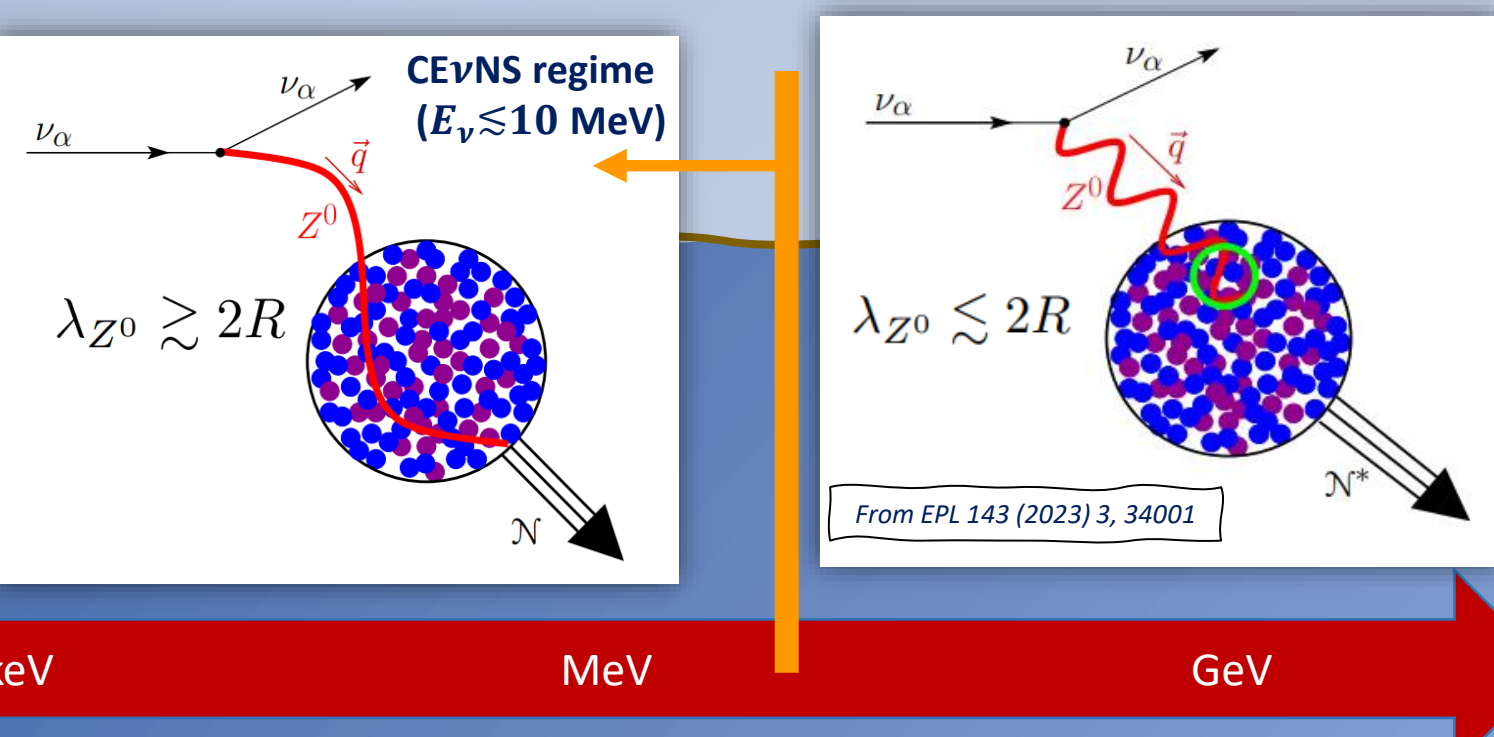
$\sim -0.5117$



## Coherency condition

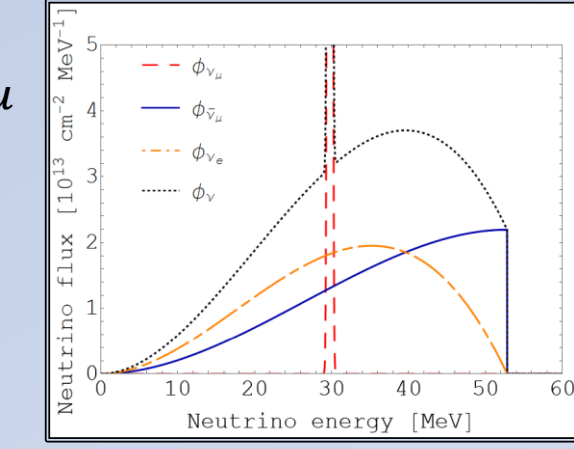
$$qR_N \lesssim 1$$

- $E_\nu \sim \frac{200 \text{ MeV fm}}{5 \text{ fm}} \sim 40 \text{ MeV}$
- Neutrinos with energies of the order of few tens of MeV see the nucleus as a whole (not its constituents)
- Low energy neutrinos produce small nuclear recoils ( $\sim \text{keV}_{nr}$ ): low experimental threshold needed

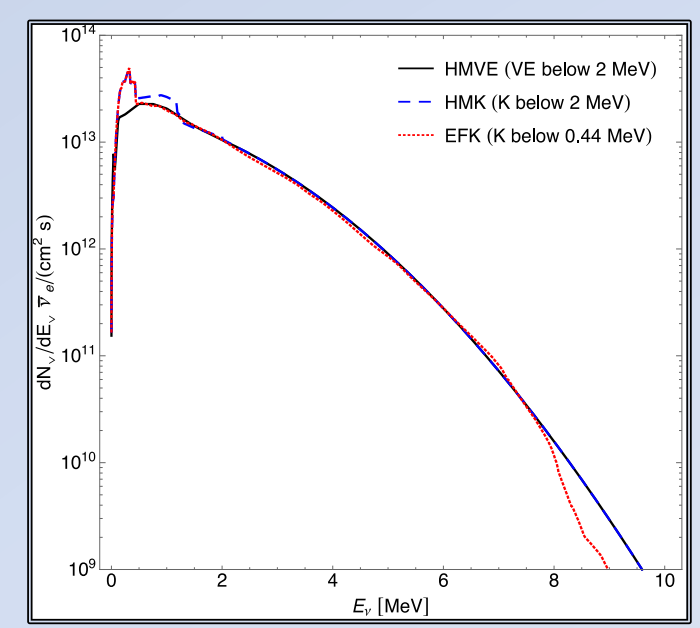


## Neutrino Sources

Spallation Neutron Source (SNS)  
 $\pi$  decay-at-rest ( $\sim 30 \text{ MeV}$ )  
Pulsed neutrino source  $\nu_e, \nu_\mu, \bar{\nu}_\mu$



Reactor Neutrinos  
Produced at nuclear power plant ( $\sim 1 - 10 \text{ MeV}$ )  
Only  $\bar{\nu}_e$



## Experimental Measurements

### COHERENT EXPERIMENT

**CsI detector:** first measurement in 2017, updated results in 2021

- Data reject no-CEvNS hypothesis at  $11.6 \sigma$
- 14.6 kg;  $\sim 7 \text{ keV}_{nr}$  threshold
- About 320 events observed consistent with the SM at  $1 \sigma$ .

**Liquid Argon (LAR) detector** observed CEvNS in 2020

- Single Phase detector (24.4 kg)
- $\sim 20 \text{ keV}_{nr}$  threshold
- $3.5 \sigma$  significance with about 120 events consistent with the SM.

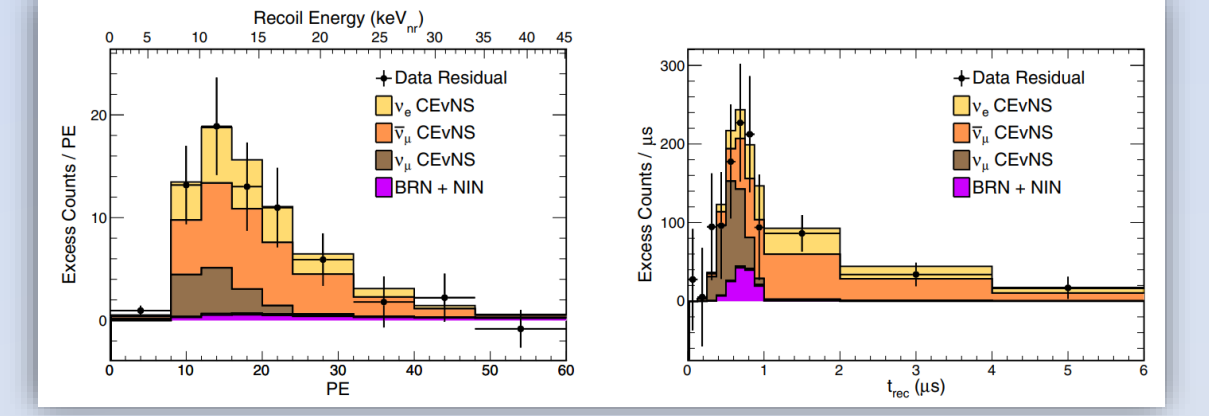
In both cases, strong constraints from both energy and timing information

### Reactor Data

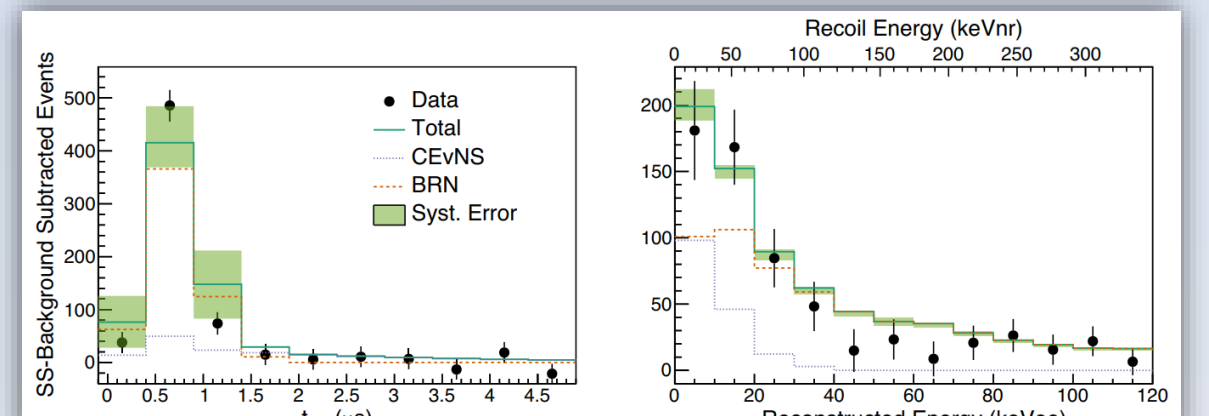
- Observation of CEvNS with a 3 kg Ge detector, 10.39 meters from the 2.96 GW<sub>th</sub> Dresden-II nuclear power plant.
- 3 kg ultra low-noise Ge detector, 0.2 keV<sub>ee</sub>

Non-standard behavior of the quenching factor at very low recoils (discussions ongoing in the community).  
See MAC. et al. Phys.Lett.B 852 (2024) 138627

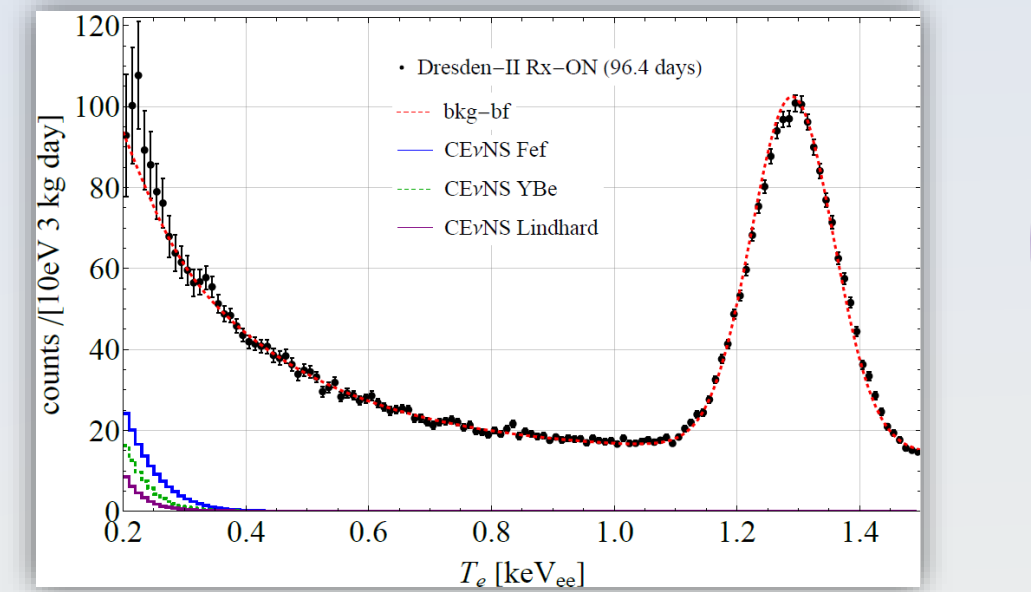
Phys.Rev.Lett. 129 (2022) 8, 081801



Phys.Rev.Lett. 126 (2021) 1, 01200213



Phys.Rev.Lett. 129 (2022) 21802



## Neutrino Electromagnetic Properties

$\nu$ -electromagnetic property:  $\nu$  interaction mediated by a photon (electromagnetic interaction).

In the most general form, the neutrino electromagnetic vertex function  $\Lambda_\mu^{ij}$  is

$$\Lambda_\mu(q) = f_Q(q^2)\gamma_\mu - f_M(q^2)i\sigma_{\mu\nu}q^\nu + f_E(q^2)\sigma_{\mu\nu}q^\nu\gamma_5 + f_A(q^2)(q^2\gamma_\mu - q_\mu\not{q})\gamma_5$$

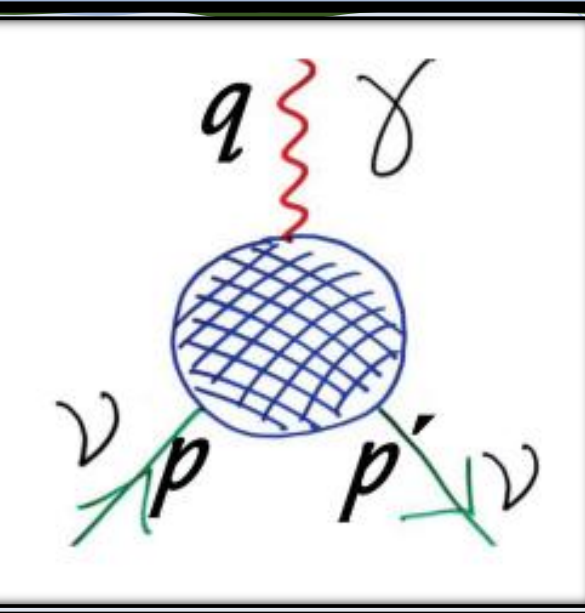
Charge Form Factor

Magnetic Moment

Electric Form Factor

Anapole Moment

See the review Rev.Mod.Phys. 87 (2015) 531



## Neutrino charge radius and CEvNS

The NCR modifies the neutrino-proton coupling through the radiative correction  $\phi_{\nu_e-W}$ , that can be expressed in terms of  $\langle r_{\nu_e}^2 \rangle$ :

$$g_V^p \approx \frac{1}{2} - 2 \sin^2 \theta_W - 2\phi_{\nu_e-W} + \text{flavor independent radiative correction (See Eur.Phys.J.C 83)}$$

$$g_V^p \approx \frac{1}{2} - 2 \sin^2 \theta_W - \frac{\sqrt{2}\pi\alpha_{EM}}{3G_F} \langle r_{\nu_e}^2 \rangle + \text{flavor independent radiative correction}$$

We can let it free to vary in the fit to obtain a constraint.

To deal with experiments performed at  $q^2 \neq 0$ , we define a radiative correction associated to the NCR dependent the momentum transfer (recall  $q^2 = -2T_{nr}$ )

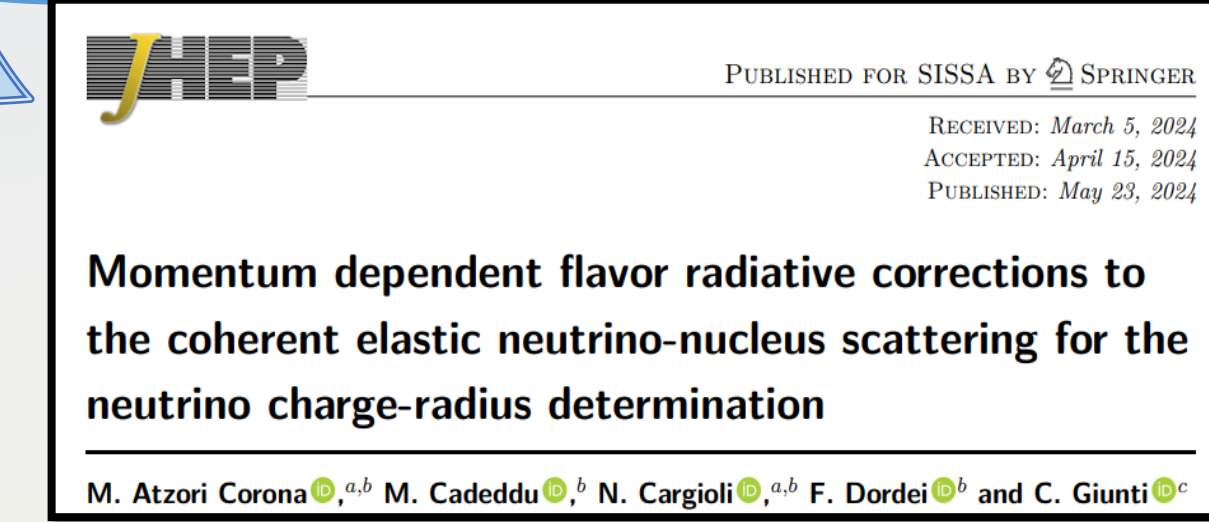
$$\langle r_{\nu_e}^2 \rangle^{\text{eff}} = -\frac{G_F}{2\sqrt{2}\pi^2} [3 - 12R_\ell(q^2)]$$

$$R_\ell(q^2) = \int_0^1 dx x(1-x) \ln \left[ \frac{m_\ell^2 - q^2 x(1-x)}{M_W^2} \right]$$

### Warning:

CEvNS experiments allow the extraction of  $\langle r_{\nu_e}^2 \rangle^{\text{eff}}$ , but the physical NCR is defined at zero momentum transfer.

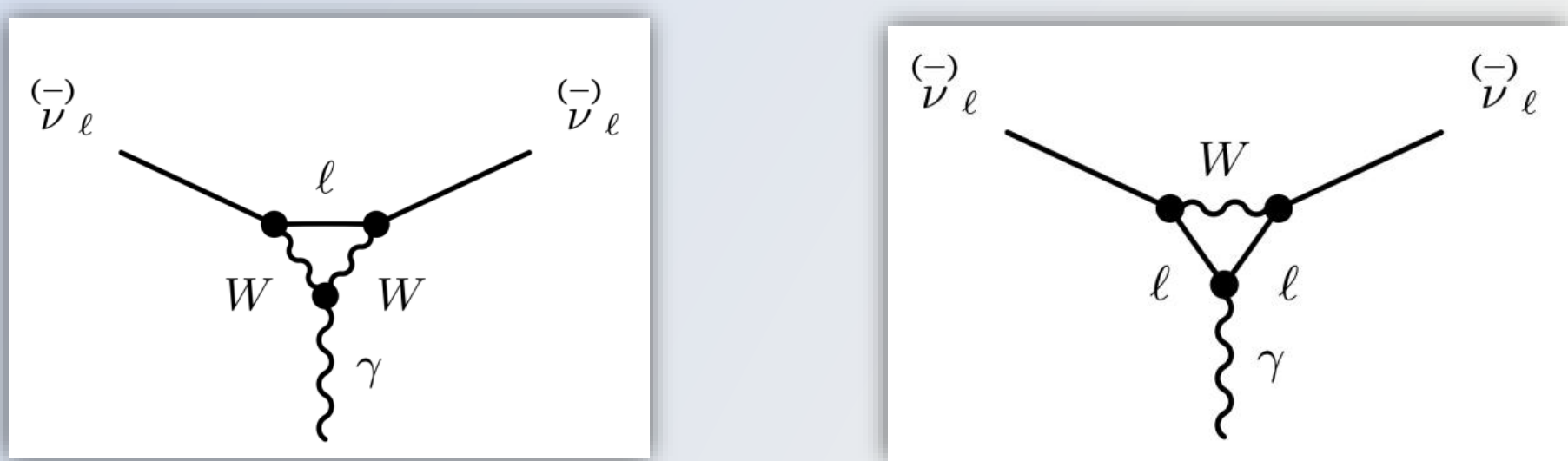
$$\mathcal{F}_{\nu_e}(T_{nr}) = \frac{\langle r_{\nu_e}^2 \rangle^{\text{eff}}(T_{nr})}{\langle r_{\nu_e}^2 \rangle^{\text{eff}}(0)} \equiv \frac{\langle r_{\nu_e}^2 \rangle^{\text{eff}}(T_{nr})}{\langle r_{\nu_e}^2 \rangle^{\text{SM}}}$$



Momentum dependent flavor radiative corrections to the coherent elastic neutrino-nucleus scattering for the neutrino charge-radius determination  
M. Atzori Corona, M. Cadeddu, N. Cargioli, F. Dordei and C. Giunti

## Neutrino Charge Radius

$\nu$ EM in the standard model: do we have such diagrams?



YES!

Neutrinos are neutral particles, therefore, they can not couple directly with photons  $\rightarrow$  **1 loop diagrams** are needed!

Expanding the charge form factor in a series of powers of  $q^2$  we get

$$f_Q(q^2) = f_Q(0) + q^2 \frac{df_Q(q^2)}{dq^2} \Big|_{q^2=0} + \dots$$

and we define the **neutrino charge radius (NCR)** as

$$\langle r_\nu^2 \rangle \equiv 6 \frac{df_Q(q^2)}{dq^2} \Big|_{q^2=0}$$

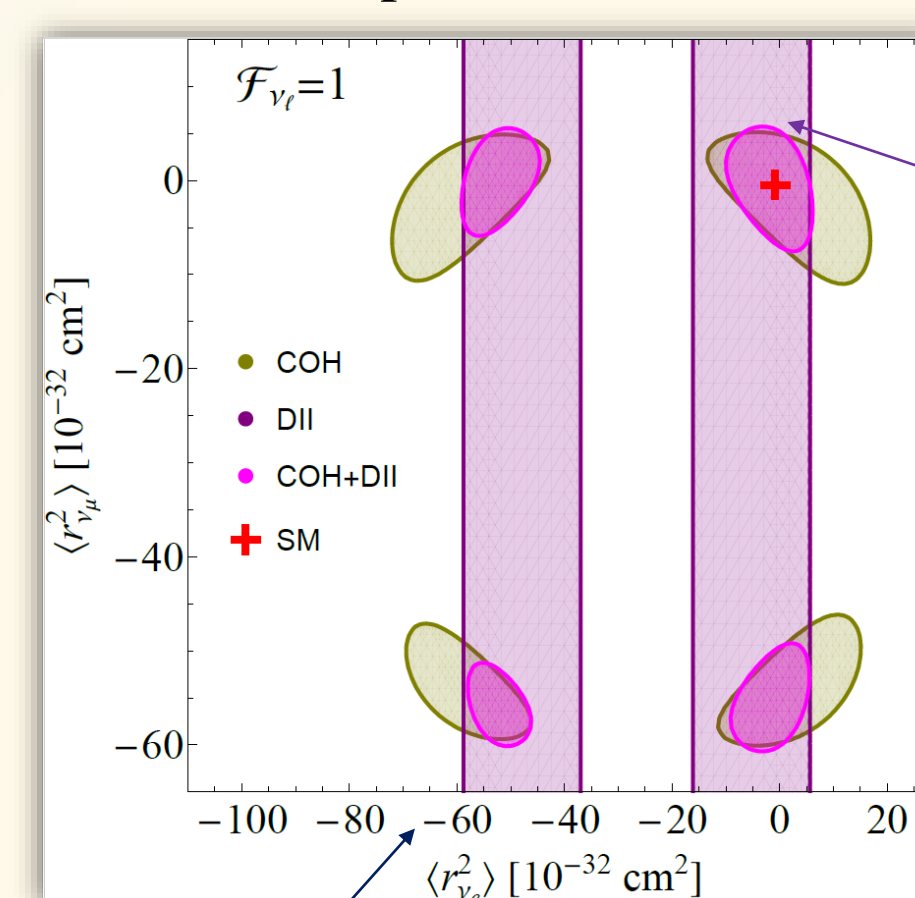
Note that it is defined at zero momentum transfer!

This quantity can be well evaluated within the standard model and gives

$$\langle r_{\nu_e}^2 \rangle_{SM} = -\frac{G_F}{2\sqrt{2}\pi^2} \left[ 3 - 2 \log \left( \frac{m_p^2}{m_W^2} \right) \right] \rightarrow \begin{cases} \langle r_{\nu_e}^2 \rangle_{SM} = -8.2 \times 10^{-33} \text{ cm}^2 \\ \langle r_{\nu_\mu}^2 \rangle_{SM} = -4.8 \times 10^{-33} \text{ cm}^2 \\ \langle r_{\nu_\tau}^2 \rangle_{SM} = -3.0 \times 10^{-33} \text{ cm}^2 \end{cases}$$

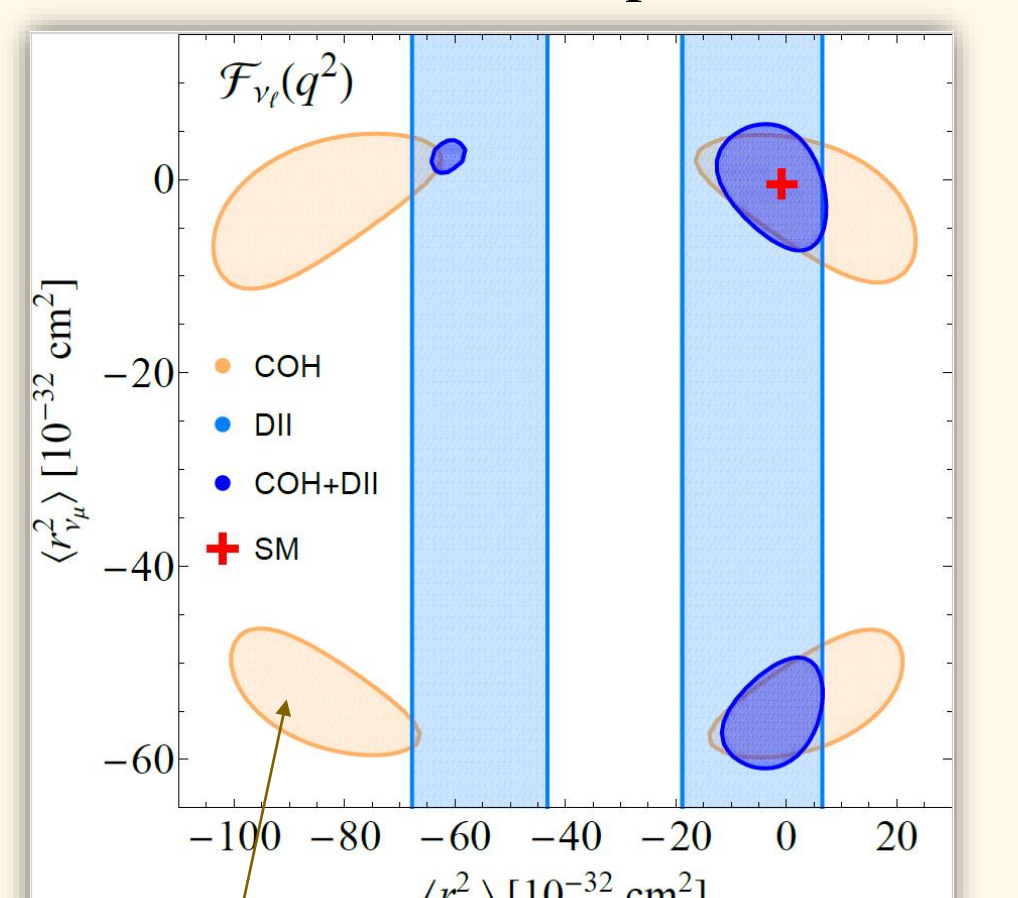
## State-Of-The-Art Constraints

No momentum dependence (JHEP 09 (2022) 164)

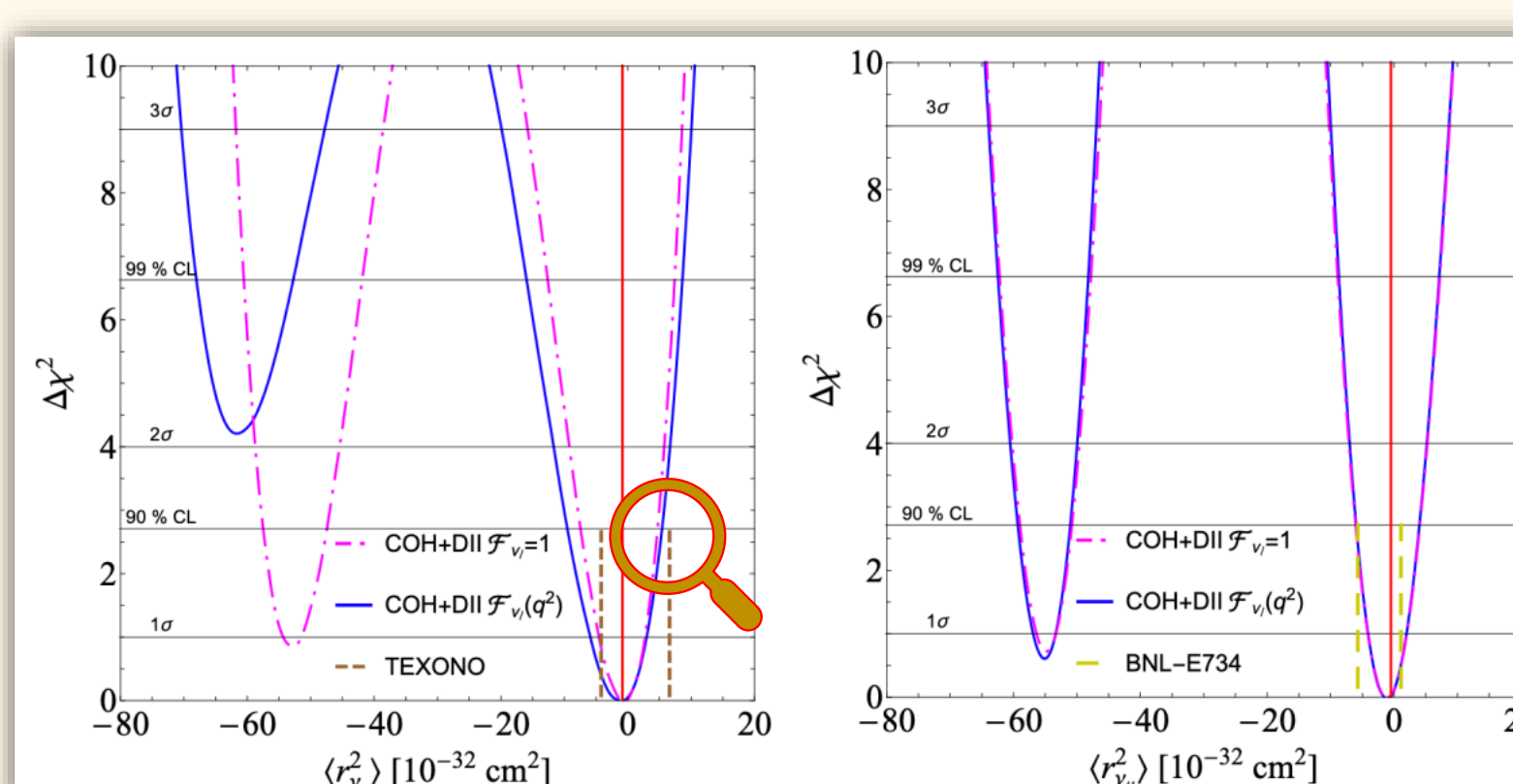


Negative large values produce a degeneracy in the cross section

With momentum dependence



COHERENT results are more affected by reactors due to the larger momentum transfer



- By combining measurements at different momentum transfer, due to the impact of the NCR form factor, the allowed regions in the parameter space are reduced.

At 90% C.L.

$$-9.5 < \langle r_e^2 \rangle [10^{-32} \text{ cm}^2] < 5.5 \quad \text{Improved best limit!}$$

$$-59.2 < \langle r_\mu^2 \rangle [10^{-32} \text{ cm}^2] < -51.0$$

$$-5.9 < \langle r_\tau^2 \rangle [10^{-32} \text{ cm}^2] < 4.1$$

World leading limits:

$$2 < \langle r_e^2 \rangle [10^{-32} \text{ cm}^2] < 6.6 \quad \text{TEXONO}$$

$$-5.7 < \langle r_\mu^2 \rangle [10^{-32} \text{ cm}^2] < 1.1 \quad \text{BNL}$$