DOUBLE-COPY SUPERTRANSLATIONS

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Based on: P. FERRERO, D.F., C. HEISSENBERG & M. ROMOLI [2402,11595]

THE DOUBLE COPY (DC) $H_{\mu\nu} = A_{\mu} * \tilde{A}_{\nu}$

 $GR = (YM)^2$ [KLT 1986, BCJ 2008-2010]

THE DOUBLE COPY (DC)

$$H_{\mu\nu} = A_{\mu} * \widetilde{A}_{\nu}$$
 [KLT 1986, BCJ 2008-2010]

$$\mathcal{A}_{m}^{\text{TREE}} = \oint_{i \in \Gamma_{3}^{n}}^{m-2} \sum_{i \in \Gamma_{3}^{i}}^{i} \frac{m_{i} c_{i}}{\Pi_{\alpha_{i}} S_{\alpha_{i}}} \rightarrow M_{m}^{\text{TREE}} = i(2\kappa)^{m-2} \sum_{i \in \Gamma_{3}^{i}}^{i} \frac{m_{i} \widetilde{m}_{i}}{\Pi_{\alpha_{i}} S_{\alpha_{i}}} - GAUGE INVARIANT- ON SHELLYM$$

THE DOUBLE COPY (DC)

$$-I_{\mu\nu} = A_{\mu} * \widetilde{A}_{\nu}$$
 [KLT 1986, BCJ 2008-2010]

$$\mathcal{A}_{m}^{\text{TREE}} = \mathcal{A}_{i \in \overline{\Gamma}_{3}^{4}}^{m-2} \underbrace{\sum_{i \in \overline{\Gamma}_{3}^{4}}^{m_{i} \subset i}}_{\text{Track}, S_{\alpha_{i}}} \longrightarrow M_{m}^{\text{TREE}} = i(2\kappa)^{m-2} \underbrace{\sum_{i \in \overline{\Gamma}_{3}^{4}}^{m_{i} \in \widetilde{M}_{i}}}_{\text{Track}, S_{\alpha_{i}}} - GAUGE INVARIANT- ON SHELLYM$$





PRODUCT OF TWO COULONB POTENTIALS



SCHWAR ? SCHILD (KERR - SCHILD FORT)

- GAUGE FIXED ON SHELL

2

THE DOUBLE COPY (DC)

$$-I_{\mu\nu} = A_{\mu} * \widetilde{A}_{\nu}$$
 [KLT 1986, BCJ 2008-2010]

$$A_{m}^{TREE} = \int_{i \in \Gamma_{3}^{*}}^{m-2} \frac{m_{i} C_{i}}{\Pi_{\alpha_{i}} S_{\alpha_{i}}} \longrightarrow M_{m}^{TREE} = i(2\kappa)^{m-2} \sum_{i \in \Gamma_{3}^{*}}^{+} \frac{m_{i} \widetilde{m}_{i}}{\Pi_{\alpha_{i}} S_{\alpha_{i}}} - GAUGE INVARIANT- ON SHELLYM$$





PRODUCT OF TWO COULONB POTENTIALS



SCHWARZSCHILD (KERR- SCHILD FORT) - GAUGE FIXED - ON SHELL

HOW MUCH DOES THE DC KNOW ABOUT SPACETIME GEOMETRY ?

THE DOUBLE COPY (DC)

$$-I_{\mu\nu} = A_{\mu} * \widetilde{A}_{\nu}$$
 [KLT 1986, BCJ 2008-2010]

$$A_{m}^{TREE} = g^{m-2} \sum_{i \in \Gamma_{3}^{i}} \frac{m_{i} C_{i}}{TT_{\alpha_{i}} S_{\alpha_{i}}} \longrightarrow M_{m}^{TREE} = i(2\kappa)^{m-2} \sum_{i \in \Gamma_{3}^{i}} \frac{m_{i} \widetilde{m}_{i}}{TT_{\alpha_{i}} S_{\alpha_{i}}} - GAUGE INVARIANT- ON SHELLYM$$





PRODUCT OF TWO COULONB POTENTIALS



SCHWARZSCHILD (KERR- SCHILD FORT) _ GAUGE FIXED _ ON SHELL

HOW MUCH DOES THE DC KNOW ABOUT SPACETIME GEOMETRY ?

HARD TO TELL, IN GENERAL

ASYMPTOTIC SYMMETRIES (AS)

AS MAY PROVIDE A NON-TRIVIAL, INTERNEDIATE SETUP

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RESIDUAL DIFFEORDAPHISONS AT J FOR ASYMPTOTICALLY - FLAT SPACES ENCODE DATA ON SOFT THEOREMS AND MEMORY EFFECTS



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AS MAY PROVIDE A NON-TRIVIAL, INTERNEDIATE SETUP

CELESTIAL M->m+a(s, +) SCHERE RESIDUAL DIFFEONDRPHISMS AT J g t FOR ASYMPTOTICALLY - FLAT SPACES H ENCODE DATA ON SOFT THEOREMS AND MEMORY EFFECTS PARTICLE CONTENT OF THE DC: J $\mathcal{E}_{i}, \mathcal{E}_{j} = \Box \otimes \Box = \Box \Box \oplus \Box$ • [0(0-2]]

AS MAY PROVIDE A NON-TRIVIAL, INTERNEDIATE SETUP

CELESTIAL it M->m+a(s,+) SCHERE RESIDUAL DIFFEONDRPHISMS AT J g t FOR ASYMPTOTICALLY - FLAT SPACES H ENCODE DATA ON SOFT THEOREMS 0 AND MEMORY EFFECTS PARTICLE CONTENT OF THE DC: J $\mathcal{E}_i \mathcal{E}_j =$ \oplus \otimes • = [0(0-2]] AS LESS KNOWN AS WELL KNOWN

AS MAY PROVIDE A NON-TRIVIAL, INTERNEDIATE SETUP



HOW MUCH DOES THE DC KNOW OF THE FULL RESULTING AS?

* A STEP TOWARDS UNDERSTANDING THE DC AT THE LEVEL OF SYMMETRIES

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* MAY (HOPEFULLY) HELP SIMPLIFYING COMPUTATIONS

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* UNIFIED FRAMEWORK ENCOMPASSING GAUGE SYMMETRIES OF DIFFERENT THEORIES



3 SETUP : THE FIELD THEORETICAL DC

J TECHNICALIA: SQUARING ASYMPTOTICS

3 DC SUPERTRANSLATIONS

3 SETUP : THE FIELD THEORETICAL DC

| THE DC FIELD HAN

FOCUS ON THE SYMMETRY ASPECTS

THREE BUILDING BLOCKS :

| THE DC FIELD HAN

Focus on THE SYMMETRY ASPECTS

THREE BUILDING BLOCKS :

| THE DC FIELD HAN

FOCUS ON THE SYMMETRY ASPECTS THREE BUILDING BLOCKS:

*

*

*

*

$$\begin{array}{c} \textcircled{1} \\ H_{\mu\nu} := A_{\mu}^{a} \circ \textcircled{1}_{aa'}^{-1} \circ \widetilde{A}_{\nu}^{a'} := A_{\mu} * \widetilde{A}_{\nu} = \blacksquare \textcircled{1} \\ \hline \\ A_{\mu}, \widetilde{A}_{\nu} & YM \text{ fields of groups } G \text{ and } \widetilde{G} \\ \hline \\ [f \circ g](x) = \int d^{p} f(y) g(x - y) \\ \hline \\ \hline \\ \\ \varPhi_{aa'} & BIADJOINT "SPECTATOR" iscala field } \end{array}$$

The *- product is s.t. $\partial_{\mu} (A_{\rho} * \widetilde{A}_{\varsigma}) = (\partial_{\mu} A_{\rho}) * \widetilde{A}_{\varsigma} = A_{\rho} * (\partial_{\mu} \widetilde{A}_{\varsigma})$

[ANASTASION, BORSTEN, DUFF, HUGHES & NAGY 2014, 2918]

| THE DC FIELD HANN

(2) $S_0 H_{\mu\nu} = \partial_{\mu} \alpha_{\nu} + \partial_{\nu} \hat{\alpha}_{\mu}$ So = LINEARISED DIFF * * So = TWO-FORN TRANSF

 $\alpha_{\mu} = \in * \widetilde{A}_{\mu}, \widetilde{\alpha}_{\mu} = A_{\mu} * \widetilde{\epsilon}$ [E, E GAUGE PARAMETERS FOR A, A,]

[ANASTASION, BORSTEN, DUFF, HUGHES & NAGY 2014, 2918]

| THE DC FIELD HANN

(2) $S_0 H_{\mu\nu} = \partial_{\mu} \alpha_{\nu} + \partial_{\nu} \hat{\alpha}_{\mu}$ So = LINEARISED DIFF * * So = TWO-FORM TRANSF

 $\alpha_{\mu} = \in * \widetilde{A}_{\mu}, \widetilde{\alpha}_{\mu} = A_{\mu} * \widetilde{\epsilon}$ [E, E GAUGE PARAMETERS FOR A, A,]

(3)
$$R_{\mu\nu\rho\sigma} \sim F_{\mu\nu} * \tilde{F}_{\rho\sigma}$$
 $[F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}]$
 $\delta_{\sigma}R_{\mu\nu\rho\sigma} = 0$
 $R_{\mu\nu\rho\sigma} = \square \otimes \square \Rightarrow GEOMETRY OF H_{\mu\nu}$ INVOLVES A TORSION
FROM $R_{\mu\nu\rho\sigma}$ LINEAR EOM FOR $H_{\mu\nu}$, WITH NON-LOCALLY COUPLED SOURCES

EXPLOITING HAN OFF-SHELL MEETS A SUBTLETY:

- /

EXPLOITING HAN OFF-SHELL MEETS A SUBTLETY:

 $\sim DC FIELD \\ H_{\mu\nu}^{s} = \Box \otimes \Box = \Box \oplus \Box \\ [GL(D)] \qquad \qquad H_{\mu\nu}^{s}$



EXPLOITING HAD OFF-SHELL MEETS A SUBTLETY: \sim DC FIELD Hru STARETRIES OF AN OFF-SHELL TWO-FORM $H_{\mu\nu} = \square \otimes \square = \square \oplus \square$ [GL(D)]SYMMETRIES OF AN OFF-SHELL GRAVITON ~ DC MULTIPLET NEEDS A LAGRANGIAN FOR HAN ALSO PROVIDING A PROPAGATING SCALAR ____ ⊕ $\mathcal{E}_i \mathcal{E}_j = \square \otimes \square$ [0(0-2)]



WHAT DO WE NEED? A LAGRANGIAN FOR THE REDUCIBLE GL(D)-TENSOR HAND

WHOSE EOT PROPAGATE GRAVITON + KALB-RAMOND + SCALAR:

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WHOSE EOT PROPAGATE GRAVITON + KALB-RAMOND + SCALAR:

$$\mathcal{L}_{HL} = \frac{1}{2} H^{\alpha\beta} \left\{ M_{\alpha\mu} M_{\beta\nu} \Pi - M_{\alpha\mu} \partial_{\beta} \partial_{\nu} - M_{\beta\nu} \partial_{\alpha} \partial_{\mu} \int H^{\mu\nu} \right\}$$

[A. CAMPOLEONI & D.F. 2013, P. FERRERO & D.F. 2020]

LAGRANGIAN CONSTRUCTION EXTENDED TO CUBIC LEVEL
 FIRST-ORDER DEFORMATION OF GAUGE SYMMETRY COMPUTED

WHAT DO WE NEED? A LAGRANGIAN FOR THE REDUCIBLE GL(D)- TENSOR HAND

WHOSE EOT PROPAGATE GRAVITON + KALB-RAMOND + SCALAR:

$$\mathcal{L}_{HL} = \frac{1}{2} H^{\alpha\beta} \left\{ M_{\alpha\mu} M_{\beta\nu} \Box - M_{\alpha\mu} \partial_{\beta} \partial_{\nu} - M_{\beta\nu} \partial_{\alpha} \partial_{\mu} \int H^{\mu\nu} \right\}$$

[A. CAMPOLEONI & D.F. 2013, P. FERRERO & D.F. 2020]

LAGRANGIAN CONSTRUCTION EXTENDED TO CUBIC LEVEL
 FIRST-ORDER DEFORMATION OF GAUGE SYMMETRY COMPUTED

HERE FOCUS ON THE LINEAR THEORY:

$$\Box H_{\mu\nu} - \partial_{\mu}\partial^{\alpha} H_{\alpha\nu} - \partial_{\nu}\partial^{\alpha} H_{\mu\alpha} = 0$$

$$\Box H_{\mu\nu} - \partial_{\mu}\partial^{\alpha} H_{\alpha\nu} - \partial_{\nu}\partial^{\alpha} H_{\mu\alpha} = 0$$

$$\Box H_{\mu\nu} - \partial_{\mu}\partial^{\alpha} H_{\alpha\nu} - \partial_{\nu}\partial^{\alpha} H_{\mu\alpha} = 0$$

$$A_{\mu} * A_{\nu}$$





WE WORK IN SINGLE- COPY LORENZ GAUGE:

 $\partial^{\alpha}A_{\alpha} = 0 = \partial^{\beta}A_{\beta}$



WE WORK IN SINGLE- COPY LORENZ GAUGE: ALL FIELDS (SINGLE- AND DOUBLE- COPY) AND ALL PARAMETERS (IDEM) SATISFY :

$$\partial^{\alpha}A_{\alpha} = 0 = \partial^{\beta}\tilde{A}_{\beta}$$

$$\Box \begin{cases} FIELD \\ PARAMETER \\ = 0 \end{cases}$$
$$O \cdot \begin{cases} FIELD \\ PARAMETER \\ = 0 \end{cases}$$
3 TECHNICALIA: SQUARING ASYMPTOTICS

ALTERNATIVE PERSPECTIVE : [CAMPIGLIA AND NAGY 2021]

EXPLORING Y->+00

TECHNICALIA : BASICS

1. (RETARDED) BONDI CORDS: (u:t-r, r, [Z;])

- 3. D = 4
- 4 \forall FIELD G(x) : $\Box G = 0$

EXPLORING Y->+00

TECHNICALIA : BASICS

1. (RETARDED) BONDI CORDS: (u:t-r, r, [Z;])

3. D = 4

4
$$\forall$$
 FIELD $G(x)$: $\Box G = O$

$$G_{+}(x) = \int_{0}^{+\infty} d\omega \frac{\omega}{2} \int_{0}^{2} d\tilde{w} e^{-i\omega \tilde{w} - \frac{i\omega r |\tilde{w} - \tilde{z}|^{2}}{1 + |\tilde{z}|^{2}}} Q(\omega q(\tilde{w}))$$

EXPLORING Y-7+00

* TECHNICALIA : BASICS

1. (RETARDED) BONDI COORDS: $(u = t - r, r, \{z;\})$ 2. $A_{\mu} \sim \frac{1}{r}$ (LEADING ORDER = RADIATION)

3. D = 4

4 \forall FIELD G(x) : $\Box G = O$ $G_{+}(x) = \int_{0}^{+\infty} d\omega \frac{\omega}{2} \int_{0}^{1/2} \frac{1}{\omega} e^{-i\omega \hat{y} \cdot n} - \frac{i\omega r |\vec{w} \cdot \vec{z}|^{2}}{(1 + |\vec{z}|^{2})} Q(\omega q(\vec{w}))$

 $\omega r | \vec{w} \cdot \vec{z} |^2 \sim O(1)$

LEADING BEHANDUR AT LARGE Y

EXPLORING Y-7+00

* TECHNICALIA : BASICS

1. (RETARDED) BONDI CORDS: (u:t-r, r, [Z;]) 2. An ~ 1 (LEADING ORDER = RADIATION) 3. D = 44 \forall FIELD G(x) : $\Box G = 0$ $G_{+}(x) = \int_{-\infty}^{+\infty} d\omega \frac{\omega}{2} \left[\frac{1}{2} \vec{w} e^{-i\omega \vec{v} - \frac{i\omega r |\vec{w} - \vec{z}|^2}{4 + i\vec{z} i^2} Q(\omega q(\vec{w})) \right]$

 $\omega_{r}|\vec{w}\cdot\vec{z}|^{2}\sim O(1)$

LEADING BEHANDUR AT LARGE Y RELEVANT FOR REGIONS INVOLVING BEHAVIOUR IN W

 $\omega \left| \vec{w} \cdot \vec{z} \right|^2 \sim O(1/r)$

$$\omega \left| \vec{w} \cdot \vec{z} \right|^2 \sim O(1/r)$$

COLLINEAR REGION

always brings about leading RADIATION behaviour $G \sim \frac{1}{r} \mathcal{N}^{(u, \vec{z})}$





COLLINEAR REGION

always brings about leading RADIATION behaviour $G \sim \frac{1}{r} \mathcal{N}(u, \vec{z})$



$$\left| \vec{w} \cdot \vec{z} \right|^2 \sim \frac{1}{\gamma}$$

COLLINEAR REGION

always brings about leading RADIATION behaviour $G \sim \frac{1}{r} \mathcal{N}(u, \overline{z})$



ALTOGETHER, Y FIELD OR PARAMETER

$$G \sim \frac{1}{\gamma} \gamma^{c}(u, \vec{z}) + \frac{1}{\gamma^{r}} \gamma^{R}(\vec{z})$$





COLLINEAR REGION

always brings about leading RADIATION behaviour $G \sim \frac{1}{r} \mathcal{N}(u, \overline{z})$



ALTOGETHER, Y FIELD OR PARAMETER

$$G \sim \frac{1}{\gamma} \mathcal{N}(u, \vec{z}) + \frac{1}{\gamma \beta} \mathcal{N}^{R}(\vec{z})$$

IN PARTICULAR : THE LEADING BEHAVIOUR DEPENDS ON B (>0)

 \sim Asymptotics of Single-COPY FIELDS $A_{\mu}, \tilde{A}_{\mu}, \Phi$

THUS:
$$G_{sc} \sim \frac{1}{r}$$
 \forall SINGLE - COPY FIELDS

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ASYMPTOTICS OF DOUBLE-COPY FIELD HAN = AM * A,

* SOFT THEOREMS
$$\Rightarrow \beta_A = \beta_{\bar{A}} = \beta_{\bar{A}} = 1$$

THUS:
$$G_{sc} \sim \frac{1}{r}$$
 \forall SINGLE - COPY FIELDS

ASYMPTOTICS OF DOUBLE-COPY FIELD HAN = AN * AN

*
$$H_{\mu\nu} = \int \frac{d^{q}u}{(2\pi)^{q}} \delta(n^{2}) e^{i\kappa x} Q^{(H)} \mathcal{E}_{\mu\nu}$$

DUE TO THE PROPERTIES OF THE CONVOLUTION: $Q^{(H)} = \frac{Q^{(A)}Q^{(A)}}{Q^{(A)}}$

* SOFT THEOREMS
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THUS:
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 \forall SINGLE - COPY FIELDS

ASYMPTOTICS OF DOUBLE-COPY FIELD HAN = AM * AN

*
$$H_{\mu\nu} = \int \frac{d^4 u}{(2\pi)^4} S(u^2) e^{iux} Q^{(H)} \mathcal{E}_{\mu\nu}$$

DUE TO THE PROPERTIES OF THE CONVOLUTION: $Q^{(H)} = \frac{Q^{(A)} Q^{(\bar{A})}}{Q^{(\bar{B})}}$
 $\sim Q^{(H)}$ SAVES LIKE ITS SC COUNTERPARTS : $\beta_H = 1$

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DUE TO THE PROPERTIES OF THE CONVOLUTION: $Q^{(H)} = \frac{Q^{(A)} Q^{(\bar{A})}}{Q^{(\bar{\Phi})}}$
 $\sim Q^{(H)}$ saws like ITS SC COUNTERPARTS : $\beta_{H} = 1$
 $\sim H_{\mu\nu} \sim \frac{3}{2}$

~ ASYMPTOTICS OF SINGLE - COPY FIELDS A, A, P

* SOFT THEOREMS
$$\Rightarrow \beta_A = \beta_{\bar{A}} = \beta_{\bar{A}} = 1$$

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$$G_{sc} \sim \frac{1}{r}$$
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ASYMPTOTICS OF DOUBLE-COPY FIELD HAN = AM * A,

*
$$H_{\mu\nu} = \int \frac{d^{4}u}{(2\pi)^{4}} \, \delta(u^{2}) \, e^{iux} \, Q^{(H)} \, \mathcal{E}_{\mu\nu}$$

DUE TO THE PROPERTIES OF THE CONVOLUTION: $Q^{(H)} = \frac{Q^{(A)} \, Q^{(\tilde{A})}}{Q^{(\tilde{a})}}$
 $\sim Q^{(H)}$ SOLUS LIKE ITS SC COUNTERPARTS : $\beta_{H} = 1$
 $\sim H_{\mu\nu} \sim \frac{3}{4}$
 $\sim GRAVITON, KALB- RAMOND FIELD AND DILATON $\sim \frac{3}{4}$$

12

* SOFT THEOREMS
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ASYMPTOTIC SYMMETRIES (2.K.2. LARGE GAUGE TRANSFORMATIONS)

ASYMPTOTIC SYMMETRIES (2.K.2. LARGE GAUGE TRANSFORMATIONS)

 \sim NEED PARAMETERS THAT GET TO $\mathcal{J}: \mathcal{E} \sim O(r^{\circ})$ $r \rightarrow +\infty$ A SYMPTOTIC SYMMETRIES (J.K.J. LARGE GAUGE TRANSFORMATIONS) NEED PARAMETERS THAT GET TO \mathcal{F} : $\mathcal{E} \sim O(r^{\circ})$ $r \rightarrow +\infty$

2 LGT -> PARAMETERS ~ E⁽³⁾(Z) (i.e. NOT JUST CONSTANT)

 \sim Asymptotic sympetries (2.k.2. Large Gauge Transformations) \sim NEED PARAMETERS THAT GET TO $\mathcal{F}: \mathcal{E} \sim O(Y^{\circ})$ $Y \rightarrow +\infty$

2 LGT → PARAMETERS ~ E⁽¹⁾(Z) (i.e. NOT JUST CONSTANT)

C ANALYSIS SIMILAR IN SPIRIT TO THAT INVOLVING FIELDS, BUT TECHNICALLY MORE INVOLVED (AND NORE INTERESTING). NEED PARAMETERS THAT GET TO \mathcal{G} : $\mathcal{E} \sim O(r^{\circ})$ $r \to +\infty$

2 LGT → PARAMETERS ~ E⁽³⁾(Z) (i.e. NOT JUST CONSTANT)

NALYSIS SIMILAR IN SPIRIT TO THAT INVOLVING FIELDS, BUT TECHNICALLY MORE INVOLVED (AND MORE INTERESTING). OUTCOME :

*
$$\alpha_{\mu} = \mathcal{E} * \widetilde{A}_{\mu} \sim \alpha^{(3)}(\overline{z}) + \widehat{\alpha}^{(1)}(\overline{z}) \frac{\mu_{\mu}}{r} + O(\frac{1}{r}) \frac{POLYHOMOGENEOUS}{EXPANSION}$$

* $\alpha^{(3)} = \alpha^{(3)}(\overline{z}) \qquad DC \qquad ASYMPTOTIC SYMMETRIES$

ASYMPTOTIC SYMMETRIES (2.4.2. LARGE GAUGE TRANSFORMATIONS) NEED PARAMETERS THAT GET TO \mathcal{F} : $\mathcal{E} \sim O(r^{\circ})$ $r_{2+\infty}$

2 LGT → PARAMETERS ~ E⁽³⁾(Z) (i.e. NOT JUST CONSTANT)

NALYSIS SIMILAR IN SPIRIT TO THAT INVOLVING FIELDS, BUT TECHNICALLY MORE INVOLVED (AND MORE INTERESTING). OUTCOME :

*
$$\alpha_{\mu} = \mathcal{E} * \widetilde{A}_{\mu} \sim \alpha^{(3)}(\overline{\tilde{e}}) + \widehat{\alpha}^{(1)}(\overline{\tilde{e}}) \frac{\ell_{\mu}r}{r} + O(\frac{t}{r}) \frac{\ell_{\nu}r}{\epsilon \times \ell_{\mu}} + O(\frac{t}{r})$$

* $\alpha^{(3)} = \alpha^{(3)}(\overline{\tilde{e}}) \quad DC \quad ASYMPTOTIC \quad SYMMETRIES$

CAN COMPUTE THE PREDICTION OF THE DC FOR THE ASYMPTOTIC SYMMETRIES AND CHARGES OF GRAVITON, TWO-FORM AND SCALAR

3 DC SUPERTRANSLATIONS

* PHYSICAL QUANTITIES ARE ENCODED IN CHARGES

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* SCHEMATIC STRUCTURE OF AN ASYMPTOTIC CHARGE:

$$Q \sim \int d^2 z \gamma_{z\overline{z}} \mathcal{E}^{(m)}(z,\overline{z}) R^{(m)}_{xxxx}$$

* PHYSICAL QUANTITIES ARE ENCODED IN CHARGES * SCHEMATIC STRUCTURE OF AN ASYMPTOTIC CHARGE:

$$Q \sim \int d^{2}z Y_{z\overline{z}} \stackrel{(m)}{\mathcal{E}(\overline{z},\overline{z})} R_{xxxx}^{(m)}$$

$$o R_{xxxx}^{(m)} \quad \text{conforment} \sim O(r^{-m}) \text{ of the FIELD STRENGTH}$$

$$o M + M \quad S.T. \quad Q = \lim_{r \to r\infty} Q(r)$$

* PHYSICAL QUANTITIES ARE ENCODED IN CHARGES * SCHEMATIC STRUCTURE OF AN ASYMPTOTIC CHARGE:

$$Q \sim \int_{2\overline{z}}^{(m)} \mathcal{E}(\overline{z},\overline{z}) R_{XXXX}^{(m)}$$

$$O R_{XXXX}^{(m)} = O(r^{-m}) \text{ of the FIELD ST}$$

$$O M + M \qquad S.T. \quad Q = \lim_{r \to r\infty} Q(r)$$

* FOR THE DC

$$R_{\mu\nu\rho\sigma} = -\frac{1}{2} F_{\mu\nu} * \tilde{F}_{\rho\sigma} = R_{\mu\nu\rho\sigma}^{s} + R_{\mu\nu\rho\sigma}^{A}$$

$$H_{\mu\nu}^{s} = \frac{1}{2} H_{(\mu\nu)} + H_{\mu\nu}^{A} \sim \frac{1}{2} H_{(\mu\nu)}$$

How much does the DC KNOW ABOUT SPACETINE GEONETRY?

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 $SH_{\mu\nu}^{S} = \partial_{\mu}S_{\nu} + \partial_{\nu}S_{\mu}$ $S= \frac{1}{2}(\alpha^{m} + \overline{\alpha}^{m})$

HOW MUCH DOES THE DC KNOW ABOUT SPACETIME GEOMETRY?



HOW MUCH DOES THE DC KNOW ABOUT SPACETIME GEOMETRY?



$$\xi^{U}_{LEADING} := T(\vec{z})$$

$$\hat{\boldsymbol{S}}^{A} = - \hat{\boldsymbol{D}}^{A} T(\hat{\boldsymbol{z}})$$
$$\hat{\boldsymbol{S}}^{T} = \frac{1}{2} \Delta T(\hat{\boldsymbol{z}})$$

HOW MUCH DOES THE DC KNOW ABOUT SPACETINE GEONETRY?

AT LEAST A BIT, ABOUT ASYMPTOTICS :



 $\xi_{LEADING}^{U} := T(\vec{z})$

 $\xi^{A} = -D^{A}T(\vec{z})$

 $\vec{\xi} = \frac{1}{2} \Delta T(\vec{z})$

DC - BHS SUPERTRANSLATIONS

$$Q^{S} = \frac{1}{8\pi G} \int d^{2} \mathcal{E} \mathcal{F}_{z\overline{z}} T(z,\overline{z}) R^{S}_{nrnr}$$
$$= \frac{-1}{32\pi G} \int d^{2} \mathcal{F}_{z\overline{z}} (\varepsilon * \widetilde{A} + A^{n} * \widetilde{\varepsilon})^{(0)} (F_{nr} * \widetilde{F}_{nr})^{(3)}$$

HOW MUCH DOES THE DC KNOW THE FULL UNDERLYING AS?
IT ACTUALLY INPROVED OUR KNOWLEDGE ON THE MATTER

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~ SCALAR SECTOR

H & = P PROVIDES THE SCALAR DOF OF THE DC MULTIPLET

IT ACTUALLY INPROVED OUR KNOWLEDGE ON THE MATTER

~ SCALAR SECTOR

 $H_{\alpha}^{S} = \Pr{PROVIDES THE SOLAR DOF OF THE DC MULTIPLET}$ $R_{\alpha}^{S} ALSO ENCODES \otimes -LY$ RWY ASYMPTOTIC CHARGES $R_{\alpha}^{P} \sim \left(d^{2}z \gamma_{z\overline{z}} (D^{2}+1)T(z,\overline{z}) \varphi^{(1)}\right)$ FOR THE SOLAR FIELD

IT ACTUALLY INPROVED OUR KNOWLEDGE ON THE MATTER

~ SCALAR SECTOR

 $H_{x}^{S} = \Psi PROVIDES THE SALAR DOF OF THE DC MULTIPLET$ \Longrightarrow $Q^{S} ALSO ENCODES \otimes -LY$ TWY ASYMPTOTIC CHARGES FOR THE SALAR FIELD $Q^{P} \sim \left(\int_{a}^{2} Z Y_{z\overline{z}} \left(D^{2} + 1 \right) T(z,\overline{z}) \Psi^{(1)} \right)$

* FIRST CONSTRUCTED BY CAMPIGLIA, COITO AND MIZERA 2018, WITH NO SYMMETRY EXPLANATION

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~ SCALAR SECTOR

 $H_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR THE SOLAR FIELD$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR FIELD$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR FIELD$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR THE SOLAR FIELD$ $<math display="block">R_{x}^{S} = \Psi PROVIDES THE SOLAR FIELD$

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* LATER INTERPRETED AS DUAL TO TWO-FORD ASYMPTOTIC CHARGES

IT ACTUALLY INPROVED OUR KNOWLEDGE ON THE MATTER

~ SCALAR SECTOR

 $H_{\alpha}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">R_{\alpha}^{S} = \Psi PROVIDES THE SOLAR DOF OF THE DC MULTIPLET$ $<math display="block">Q_{\alpha}^{S} = Q_{\alpha}^{S} = Q_{\alpha}^$

* FIRST CONSTRUCTED BY CAMPIGLIA, COITO AND MIZERA 2018, WITH NO SYMMETRY EXPLANATION

* LATER INTERPRETED AS DUAL TO TWO-FORM ASYMPTOTIC CHARGES * IN OUR CONTEXT NATURAL CONSEQUENCE OF BAS SUPERTRANSLATIONS, ONCE P IS IDENTIFIED AS THE DILITON OF THE DC MULTIPLET





$$\begin{split} \delta H^{A}_{\mu\nu} &= \partial_{\mu}\Lambda_{\nu} - \partial_{\nu}\Lambda_{\mu} \\ \Lambda^{\mu} &= \frac{1}{2} \left(\alpha^{\mu} - \tilde{\alpha}^{\mu} \right) \end{split}$$

~ TWO-FORN SECTOR

$$\begin{split} \delta H^{A}_{\mu\nu} &= \partial_{\mu}\Lambda_{\nu} - \partial_{\nu}\Lambda_{\mu} \\ \Lambda^{\mu} &= \frac{1}{2} \left(\alpha^{\mu} - \tilde{\alpha}^{\mu} \right) \end{split}$$

* OUT PUT OF THE DC : POLYHONDGENEOUS EXPANSION:

$$\Lambda \sim \Sigma_{t}^{t} \left\{ \frac{1}{\gamma^{m}} \varepsilon^{(m)} + \frac{l_{y} \gamma}{\gamma^{m}} \lambda^{(m)} \right\}$$

~ TWO-FORN SECTOR

$$\begin{split} \delta H^{A}_{\mu\nu} &= \partial_{\mu}\Lambda_{\nu} - \partial_{\nu}\Lambda_{\mu} \\ \Lambda^{\mu} &= \frac{1}{2} \left(\alpha^{\mu} - \tilde{\alpha}^{\mu} \right) \end{split}$$

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PREVIOUS LITERATURE ON THE SUBJECT DDES NOT INCLUDE THESE TERMS

~ TWO-FORT SECTOR

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PREVIOUS LITERATURE ON THE SUBJECT DOES NOT INCLUDE THESE TERMS

IS IT A PROBLEM ? DOES THE DC PROVIDE A MISLEADING/SUPERFLUOUS INDICATION ?

~ TWO-FORT SECTOR

$$\begin{split} \delta H_{\mu\nu}^{A} &= \partial_{\mu} \Lambda_{\nu} - \partial_{\nu} \Lambda_{\mu} \\ \Lambda^{\mu} &= \frac{1}{2} \left(\alpha^{\mu} - \tilde{\alpha}^{\mu} \right) \end{split}$$

* OUT PUT OF THE DC : POLYHONDGENEOUS EXPANSION:

$$\Lambda \sim \sum_{i}^{i} \left\{ \frac{1}{\gamma^{m}} \in {}^{(m)} + \frac{l_{ij} \gamma}{\gamma^{m}} \lambda^{(m)} \right\}$$

PREVIOUS LITERATURE ON THE SUBJECT DOES NOT INCLUDE THESE TERMS

IS IT A PROBLEM ? DOES THE DC PROVIDE A MISLEADING/SUPERFLUOUS INDICATION?

THE ANSWER IS CONTAINED IN THE CHARGE:

$$Q^{S} = \frac{z}{8\pi G} \int d^{2} z \gamma_{z\overline{z}} \Lambda^{(1)} R^{A(2)}_{irmr}$$

~ TWO-FORN SECTOR

$$\begin{split} & \int H_{\mu\nu}^{A} = \partial_{\mu}\Lambda_{\nu} - \partial_{\nu}\Lambda_{\mu} \\ & \Lambda \sim \sum_{i}^{i} \left\{ \frac{1}{Y^{m}} e^{(m)} + \frac{b_{i}Y}{Y^{m}} \lambda^{(m)} \right\} \\ & \Lambda \sim \sum_{i}^{i} \left\{ \frac{1}{Y^{m}} e^{(m)} + \frac{b_{i}Y}{Y^{m}} \lambda^{(m)} \right\} \\ & PREVIOUS LIFERATURE ON \\ & THE SUBJECT DOES NOT INCLUDE \\ & THESE TERMS \\ & IS IT A PROBLEM ? DOES THE DC PROVIDE A \\ & MISLEADING/SUPERFLUOUS INDICATION? \\ & THE ANSWER IS CONTAINED IN THE CHARGE: \\ & WITHOUT THE GIF TERNS \\ & Q^{S} = \frac{1}{3\pi G} \int d^{2} Y_{2\overline{2}} \left(\Lambda^{i} A^{(2)} R^{A(2)} \right) \\ & R^{i} A^{(2)} R^{i} C R \\ & IS IT A PROBLEM ? DOES THE DC PROVIDE A \\ & MISLEADING = 0 \\ \hline \\ & MISLEADING =$$

US EXPANSION:

A

~ TWO-FORT SECTOR

$$\begin{split} S H_{\mu\nu}^{A} &= \partial_{\mu} \Lambda_{\nu} - \partial_{\nu} \Lambda_{\mu} \\ \Lambda^{\mu} &= \frac{1}{2} \left(\alpha^{\mu} - \overline{\alpha}^{\mu} \right) \\ & \Lambda \sim \Sigma_{1}^{i} \left\{ \frac{4}{\gamma^{\mu}} e^{\binom{(n)}{i}} + \frac{b_{3}\gamma}{\gamma^{\mu}} \lambda^{\binom{(n)}{i}} \right\} \\ & PREVIOUS LITERATURE OW \\ & THE SUBJECT DOES NOT INCLUDE \\ & THESE TERMS \\ & IS IT A PROBLEM ? DOES THE DC PROVIDE A \\ & HISLEADING/SUPERFLUOUS INDICATION? \\ & THE ANSWER IS CONTAINED IN THE CHARGE: \\ & WITHOUT THE GIFTERS \\ & LORENZ GAUGE = \lambda^{1} \int_{1}^{(1)} e^{2} \gamma_{z\overline{z}} \int_{1}^{1} \int_{1}^{(1)} R^{A(2)} \\ & THE DC INPLEMENTS THE \\ & CORRECT ASYMPTOTIC \\ \end{split}$$

EXPANSION OF THE PARAMETERS



~ D>4

2 MORE GENERAL GAUGE THEORIES, MULTIPLE COPIES, HIGHER SPINS.

