



## Search for a light sterile neutrino with the KATRIN experiment

#### NuMass 2024 Workshop, Genova

*Shailaja Mohanty* (shailaja.mohanty@kit.edu) for the KATRIN collaboration Institute for Astroparticle Physics | February 29, 2024



#### www.kit.edu



#### Non-standard or Sterile Neutrino

Sterile neutrino = SM neutral singlet fermion

 Existence could be revealed through effects of mass and mixing with active neutrinos (neutrino oscillations, β- decay, 0νββ-decay)



#### Non-standard or Sterile Neutrino

Sterile neutrino = SM neutral singlet fermion

- Existence could be revealed through effects of mass and mixing with active neutrinos (neutrino oscillations, β- decay, 0νββ-decay)
- Theoretical motivation:

- Singlet fermions naturally appear in the dark sector
- Members of dark sector could mix with active neutrinos via neutrino portal coupling
- Sterile neutrinos can live at any mass scales: GeV, keV,





#### Non-standard or Sterile Neutrino

Sterile neutrino = SM neutral singlet fermion

- Existence could be revealed through effects of mass and mixing with active neutrinos (neutrino oscillations, β- decay, 0νββ-decay)
- Theoretical motivation:

eV

DM e	exists	$\Longrightarrow$	uncharged	particles	under	SM	gauge	group
$\implies$	single	et fern	nions					

- Singlet fermions naturally appear in the dark sector
- Members of dark sector could mix with active neutrinos via neutrino portal coupling
- Sterile neutrinos can live at any mass scales: GeV, keV,

- Experimental hints for eV scale :
  - Appearance LSND (3 $\sigma$ ) and MiniBooNE (4.8 $\sigma$ ) excess observations Explained by ( $\nu_{\mu} \rightarrow \nu_{s} \rightarrow \nu_{e}$ )
  - Disappearance SAGE and GALLEX: Gallium anomaly (2.9 $\sigma$  deficit) Explained by  $\nu_e \rightarrow \nu_s$
  - The Gallium anomaly reaffirmed by BEST experiment



#### Interpretation



- SBL anomalies could be explained by an additional neutrino flavor (v<sub>s</sub>)
- There must be at least one additional mass squared difference,  $3\nu + 1$  framework  $\Delta m_{SBI}^2 \approx (1-2) \text{ eV}^2$
- Allowed by solar, atmospheric and long baseline experiments, achieved with  $|U_{e4}|^2 \ll 1$



#### Differential decay rate:

Single  $\beta$ -decay

$$R_{\beta}(E, m_{\nu}^{2}, m_{4}^{2}, |U_{e4}|^{2})$$

$$= \underbrace{(1 - |U_{e4}|^{2}) \cdot R_{\beta}(E, m_{\nu}^{2})}_{\text{Active branch}} + \underbrace{|U_{e4}|^{2} \cdot R_{\beta}(E, m_{4}^{2})}_{\text{Sterile branch}}$$

$$= \cos^{2} \theta \cdot R_{\beta}(E, m_{\nu}^{2}) + \sin^{2} \theta \cdot R_{\beta}(E, m_{4}^{2})$$
• A kink at  $E_{0} - m_{4}$ 

$$m_{\nu}^{2} = \sum_{k=1}^{3} |U_{ek}|^{2} m_{k}^{2} \xrightarrow{3+1} \sum_{k=1}^{3}$$

$$m_{\nu}^{2} = \sum_{k=1}^{3} |U_{ek}|^{2} m_{k}^{2} \xrightarrow{3+1} \sum_{k=1}^{3}$$

$$m_{\mu}^{2} = 0.6$$

$$m_{4}^{2} = 0.4$$

0







2

 $|U_{ek}|^2$ 



Talk by V. Hannen

#### **KATRIN Experiment**

- Kinematics-based neutrino mass experiment (expected sensitivity: 0.3 eV (90% CL) after 1000 days of measurement time)
- Current result: m<sub>β</sub>< 0.8 eV (90%) CL, (Nature Phys. 18 (2022) 2, 160-166)</p>



#### Sterile Signal in $\beta$ -decay Spectrum



Measured integral spectrum N<sub>exp</sub>(qU) is fitted to the model N<sub>model</sub>(qU, Θ):

$$N_{ ext{model}}(qU,\Theta) = A \cdot \int R_{eta}(E,\Theta) \cdot f(E,qU) + Bg$$

- 6 model parameters:
  - A Signal amplitude
  - E<sub>0</sub> effective endpoint energy
  - m<sup>2</sup> effective mass of electron anti-neutrino
  - Bg Background rate
  - m<sub>4</sub><sup>2</sup> sterile neutrino mass
  - $|U_{e4}|^2$  sterile neutrino mixing





#### **Dataset and Analysis Strategy**

- Data selection and combination, active neutrino model configuration are the same as for the active neutrino mass analysis. Talk by W. Xu
- Unblinding procedure<sup>a</sup>
  - Code validation on Monte Carlo twins
  - Tritium spectrum, model, systematics treatment and budget (pull term approach) same as active neutrino mass analysis
  - Two independent analysis teams with independent codes:
    - KaFit (exact model evaluation)
    - Netrium (use neural nets for swift model interpolation)

<sup>a</sup>M. Aker et al. "Improved eV-scale sterile-neutrino constraints from the second KATRIN measurement campaign". In: *Phys. Rev. D* 105 (7 2022), p. 072004. DOI: 10.1103/PhysRevD.105.072004.



#### Analysis method

- Extend Tritium β- spectrum model to 3+1 framework
- Grid Scan: 50  $\times$  50  $\left[\log(|U_{e4}|^2), \log(m_4^2)\right]$  plane
- Contours are drawn at  $\Delta \chi^2 = \chi^2 \chi^2_{BF}$  = 5.99 (95% CL, 2 dof)
- Energy range: [*E*<sub>0</sub> − 40, *E*<sub>0</sub> + 135] eV
- Sensitive to  $m_4^2 \leq$  1600 eV<sup>2</sup> and  $|U_{e4}|^2 \leq$  0.5
- Two complementing analyses
  - Case-I Fixed neutrino mass:  $m_{\nu}^2$  = 0 ( $m_{1,2,3} \ll m_4$ )
  - Case-II Free neutrino mass: m<sup>2</sup><sub>ν</sub> as nuisance parameter





#### **Data collection status**



Table: KATRIN Neutrino Mass Measurements (KNM)

Campaign	Time (hrs)	$ ho { m d} \sigma$ (m $^{-2}$ )	Bg (mcps)	
KNM1	522	$1.11  imes 10^{21}$	370	
KNM2	294	$4.23 imes10^{21}$	278	
KNM3a	220	$2.08  imes 10^{21}$	137	
KNM3b	224	$3.75 imes10^{21}$	258	
KNM4	1267	$3.77 imes10^{21}$	150	
KNM5	1232	$3.78 imes10^{21}$	160	

#### KNM1, KNM2, KNM3b operated in Nominal Analyzing Plane (NAP) mode



#### **Results from First Two Science Runs**

• 5.24  $\times$  10<sup>6</sup> electrons for 40 eV below E<sub>0</sub>,  $10^{3}$ 1265 hours of data Best fit:  $-m_4^2 = 59.9 \text{ eV}^2$ ,  $|U_{e4}|^2 = 0.011$ ,  $10^{2}$  $m_{4}^{2}$  (eV<sup>2</sup>)  $m_{\nu}^2 = 0.0 \text{ eV}^2$  $-\Delta\chi^2_{null}=0.66$ Active neutrino mass set free  $m_{\nu}^2 = 0 \text{ eV}^2$   $m_{\nu}^2 \text{ free}$ .....KNM1 .....KNM1 Best fit:  $- m_4^2 = 87.4 \text{ eV}^2, |U_{e4}|^2 = 0.019,$ ---KNM2 ---KNM2  $m_{\nu}^2 = 0.57 \text{ eV}^2$ -KNM1+2 --- KNM1+2  $-\Delta \chi^2_{null} = 1.69$ Signal-to-background ratio of up to 235  $10^{-2}$  $|U_{\mathcal{A}}|^2$ 

 $10^{-1}$ 



#### **Data collection status**

- Significant experimental development: Shifted Analyzing Plane (SAP) background reduction method Lokhov et al., EPJ C 82 (2022) 3, 258
- KNM1 to KNM5: 20 % of expected KATRIN data





#### **Sensitivity Results From Five Science Runs**

- **Case-I**:  $m_{\nu}^2$  = 0 eV<sup>2</sup>
- 40 eV fit range,  $|U_{e4}|^2 \in [0, 0.5]$
- Stat. only + all systematics 95% CL
- Gain in overall sensitivity with increased statistics
   S. Mohanty, PoS EPS- HEP2023 (2024)



Campaign	KNM1	KNM2	KNM3a	KNM3b	KNM4	KNM5	KNM1-5
No. of signal electrons ( $\times 10^6$ )	2.0	4.3	1.1	1.4	10.2	16.8	35.8



#### Impact of Systematics

Calculating 68% CL uncertainty on 
$$|U_{e4}|^2$$
:  $\sigma_{syst} = \sqrt{\sigma_{Stat+Syst}^2 - \sigma_{Stat}^2}$ 

- Statistically dominated uncertainties
- Largest systematic contribution: Penning Bg (low m<sup>2</sup><sub>4</sub>), Column Density (high m<sup>2</sup><sub>4</sub>)





## Sensitivity comparison to other experimental results

Translation of parameters:

 $\sin^2(2\theta) = 4|U_{e4}|^2(1-|U_{e4}|^2)$ 

- Large Δm<sup>2</sup><sub>41</sub> solutions of RAA and BEST+GA anomalies excluded
- Current KATRIN data extends exclusion bounds from SBL oscillation experiments for  $\Delta m_{41}^2 \ge 10 \text{ eV}^2$
- Probing large parameter space for light sterile neutrino anomalies
- Expected KNM1-5 sensitivity yields improved constraints in the sterile parameter space





#### Impact of active neutrino on sterile neutrino search

Possible treatments for  $m_{\nu}^2$ : Extension of Case-II

• Free  $m_{\nu}^2$ 

Correlation between  $m_4^2$  and  $m_{\nu}^2$ .

Pull term using **0**±**1** eV<sup>2</sup>

Intermediate sensitivity between two extremes (fixed and free)

■ m<sup>2</sup><sub>4</sub> > m<sup>2</sup><sub>ν</sub> ≥ 0: Limit m<sup>2</sup><sub>ν</sub> by mass of right-handed neutrino

Reasonable option of optimizing sensitivity in addition to free  $m_{\nu}^2$  case





- New physics beyond the SM can include sterile neutrinos at all mass scales
- KATRIN uniquely addresses SBL anomalies via spectral shape analysis
- Results from first two science runs (KNM1 + KNM2):
  - No significant sterile-neutrino signal observed
  - Improved exclusion limits w.r.t. complementary experiments



- New physics beyond the SM can include sterile neutrinos at all mass scales
- KATRIN uniquely addresses SBL anomalies via spectral shape analysis
- Results from first two science runs (KNM1 + KNM2):
  - No significant sterile-neutrino signal observed
  - Improved exclusion limits w.r.t. complementary experiments
- Sensitivity projection for five science runs (KNM1...5):
  - Increased dataset boosts sensitivity, potential to probe large parameter space of SBL anomalies and complementary to oscillation experiments
  - Sensitivity dominated by statistical uncertainties



- New physics beyond the SM can include sterile neutrinos at all mass scales
- KATRIN uniquely addresses SBL anomalies via spectral shape analysis
- Results from first two science runs (KNM1 + KNM2):
  - No significant sterile-neutrino signal observed
  - Improved exclusion limits w.r.t. complementary experiments
- Sensitivity projection for five science runs (KNM1...5):
  - Increased dataset boosts sensitivity, potential to probe large parameter space of SBL anomalies and complementary to oscillation experiments
  - Sensitivity dominated by statistical uncertainties
- $m_{\nu}^2 = 0 \text{ eV}^2$  gives better constraints than  $m_{\nu}^2$  set free.
- $-\,$  Reasonable option for optimized sensitivity  $m_4^2 > m_\nu^2 \ge 0\,$



- New physics beyond the SM can include sterile neutrinos at all mass scales
- KATRIN uniquely addresses SBL anomalies via spectral shape analysis
- Results from first two science runs (KNM1 + KNM2):
  - No significant sterile-neutrino signal observed
  - Improved exclusion limits w.r.t. complementary experiments
- Sensitivity projection for five science runs (KNM1...5):
  - Increased dataset boosts sensitivity, potential to probe large parameter space of SBL anomalies and complementary to oscillation experiments
  - Sensitivity dominated by statistical uncertainties
- $-m_{\nu}^2 = 0 \text{ eV}^2$  gives better constraints than  $m_{\nu}^2$  set free.
- $-\,$  Reasonable option for optimized sensitivity  $m_4^2 > m_\nu^2 \ge 0\,$

Outlook:

- Analysis on data for first five science runs ongoing
- Stay tuned for upcoming release!























# Backups

Institute for Astroparticle Physics



#### **Experimental hints**

- Appearance LSND (3σ) and MiniBooNE (4.8σ) excess observations. Explained by (ν<sub>μ</sub> → ν<sub>s</sub> → ν<sub>e</sub>)
- Disappearance SAGE and GALLEX: Gallium anomaly (2.9 $\sigma$  deficit). Explained by  $\nu_e \rightarrow \nu_s$
- The Gallium anomaly reaffirmed by BEST experiment







#### Measurement time distribution - Standard vs Flat



Institute for Astroparticle Physics



#### Raster scan on different measured time distributions



Institute for Astroparticle Physics

# Karlsruher Institut für Technologie

#### Monte Carlo breakdown



KNM4 systematic breakdown – Monte Carlo



#### Active neutrino correlation with sterile neutrino



FIG. 4. The correlation between active and sterile neutrino mass is approximately a linear slope  $m_b^2 = \alpha_{\rm slope} \cdot m_A^2 + {\rm const}$  for various values of  $m_a^2$  and  $|\mathcal{L}_{cd}|^2$  by analyzing simulated spectra. The gradient indicates the magnitude of  $\alpha_{\rm slope}$ . For small mixing  $|\mathcal{L}_{cd}|^2 < 0.01$ , we observe small slope values  $|\alpha_{\rm slope}| < 0.01$ . For larger mixing, we find a strong negative correlation for larger  $m_A^2 \lesssim 30 \ {\rm eV}^2$  and a weaker positive correlation for larger  $m_A^2 \lesssim 30 \ {\rm eV}^2$ 

Institute for Astroparticle Physics



# $m_{ u}^2 = 0 \; {f vs} \; m_4^2 > m_{ u}^2 \ge 0$



Institute for Astroparticle Physics



#### Impact of Measured Time Distribution

Objective: To investigate spikes in the raster contours.





#### Schematic overview of KaFit

- C++ based fitting framework used to analyse measured KATRIN data and simulated data
- Applicable for Frequentist (based on MINUIT class of ROOT) and Bayesian analysis
- Minimisation is performed with MINUIT by minimising the -2log(L)





# Testing applicability of Wilks' Theorem

Previously done

- Generate O(10<sup>3</sup>) twins with statistical fluctuations for particular choice of MC truth
- Perform fitting for sterile parameter values on a grid and for MC truth for each sample (m<sup>2</sup><sub>ν</sub> = 0)
- Evaluate  $\Delta \chi^2 = \chi^2_{\rm MC \ truth} \chi^2_{\rm best \ fit}$  for each sample
- Compare distribution of  $\Delta \chi^2$  values to  $\chi^2$ -distribution with 2 dof



Taken from H.9 of Schlüter, L. (2022). Neutrino-Mass Analysis with subeV Sensitivity and Search for Light Sterile Neutrinos with the KATRIN Experiment. PhD Thesis, TU München, Garching bei München.