



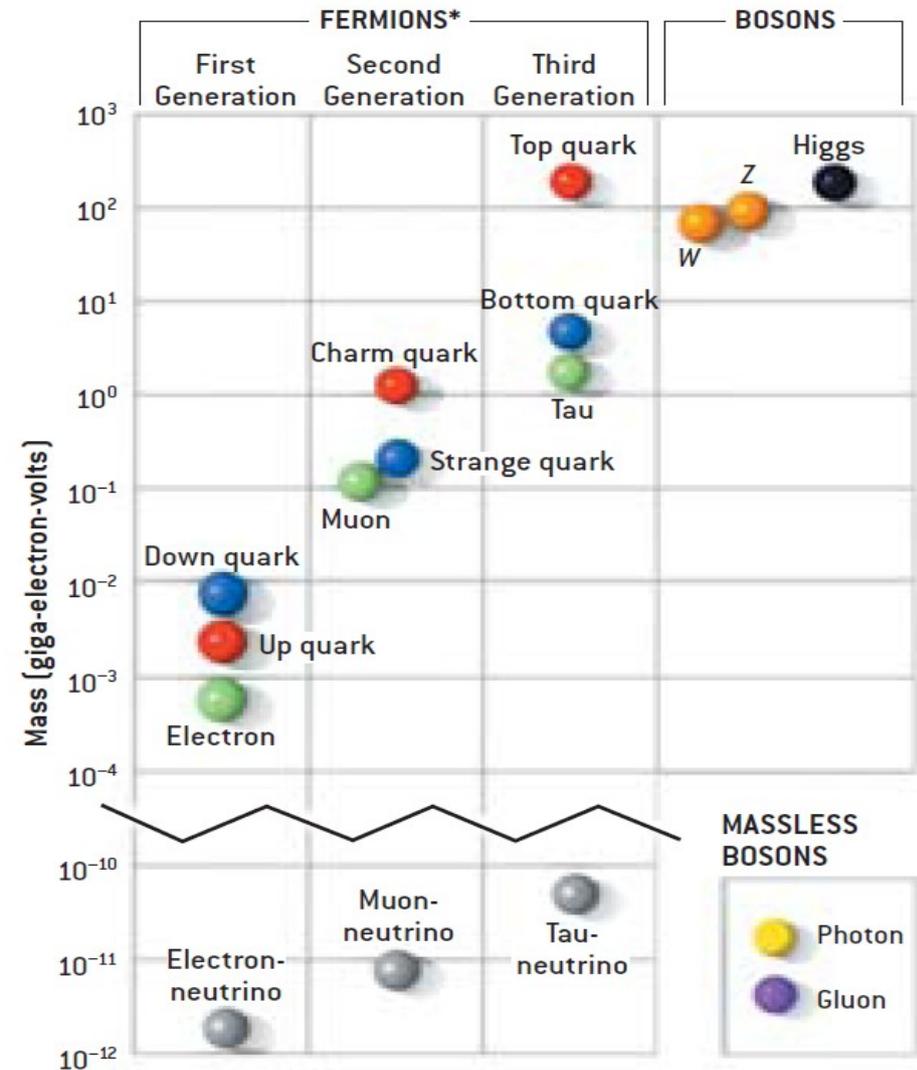
Universität  
Münster

# The Karlsruhe Tritium Neutrino Experiment

V.M. Hannen for the KATRIN collaboration

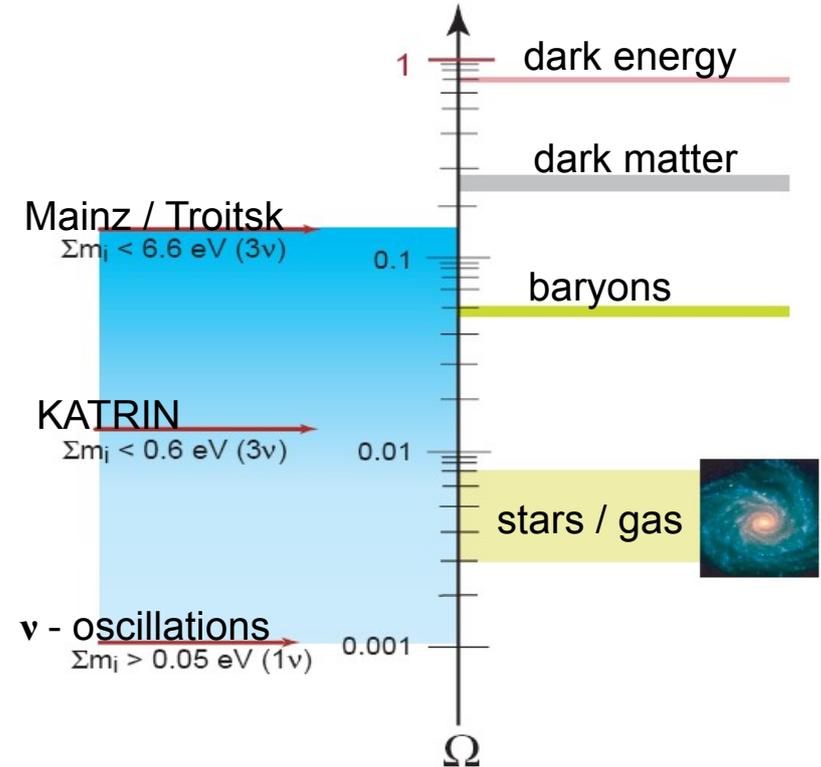
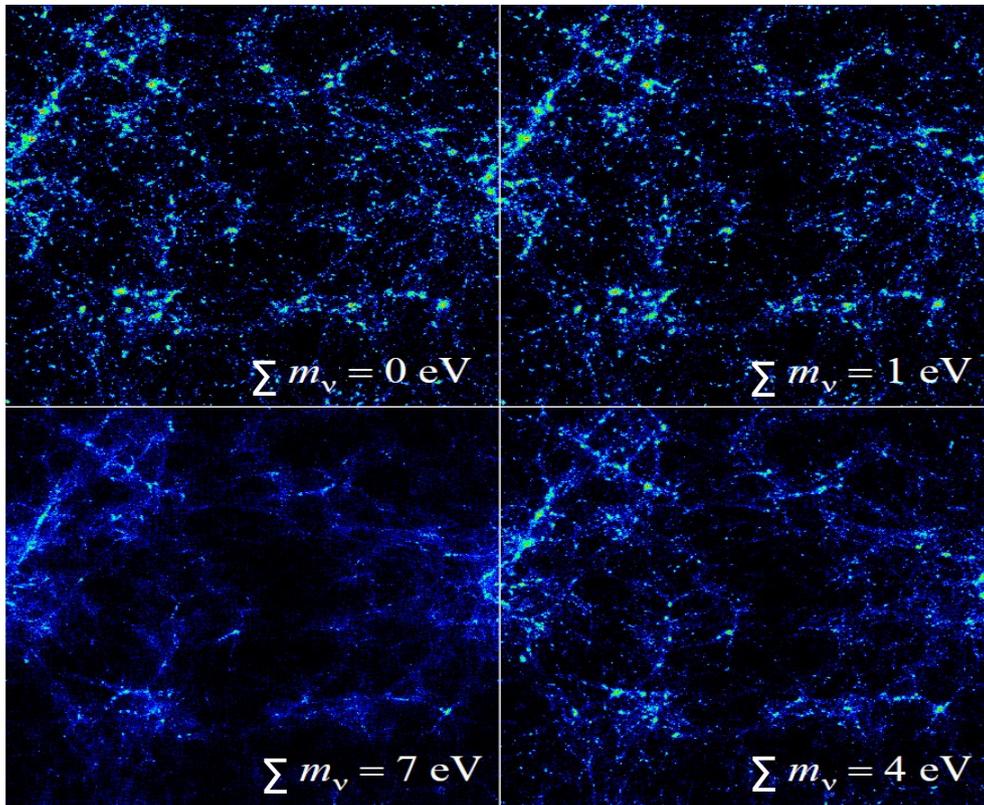
Institut für Kernphysik,  
Universität Münster

- Nature of the neutrino: Majorana or Dirac particle, i.e. is the neutrino its own anti-particle ?
- How to explain the many orders of magnitude difference between neutrino mass limits and masses of the charged fermions of the standard model  
→ sea-saw type I and type II mechanisms
- Possible connection to the generation of the observed matter - antimatter asymmetry in the universe  
→ leptogenesis



# Neutrino mass in cosmology

- Neutrinos are (after  $\gamma$ 's) the second most abundant particle species in the universe
- As part of the hot dark matter, neutrinos have a significant influence on structure formation



- For large  $\Sigma m_\nu$  values fine grained structures are washed out by the free streaming neutrinos

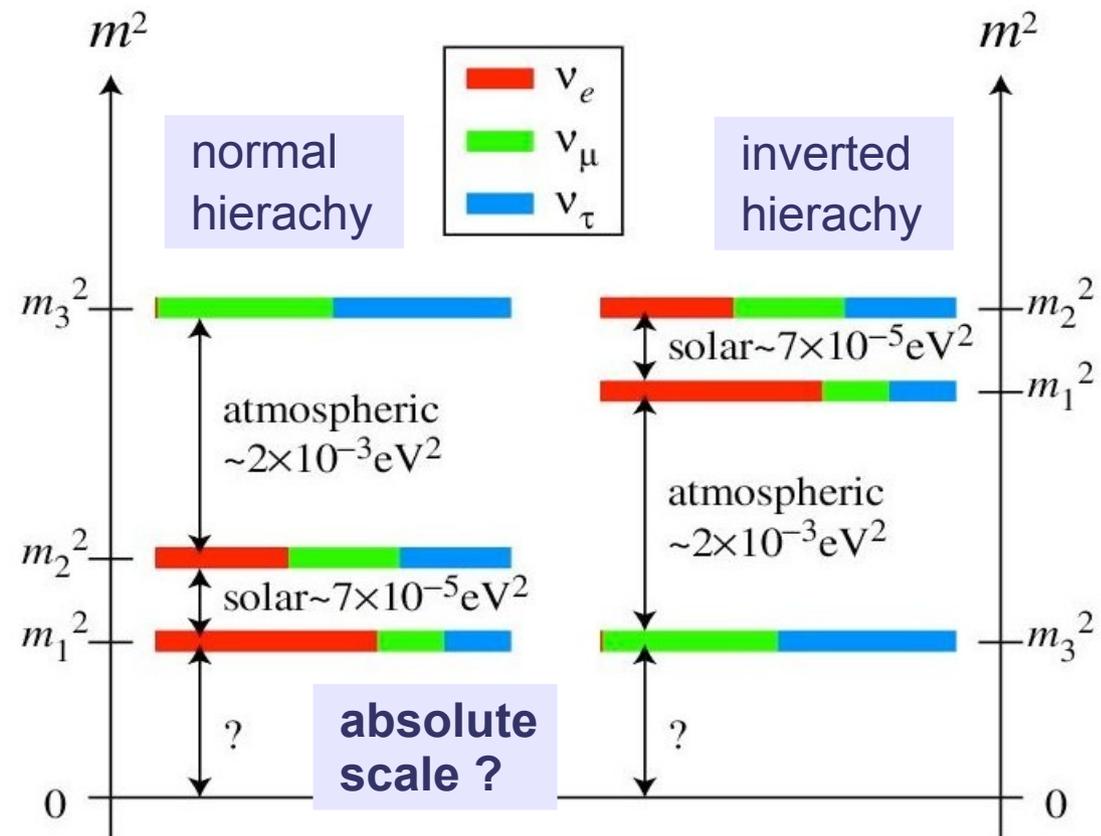
Chung-Pei Ma 1996

## What we know (from $\nu$ oscillations):

- Neutrino flavour eigenstates differ from their mass eigenstates
- Neutrinos oscillate, hence they must have mass
- Mixing angles and  $\Delta m^2$  values known (with varying accuracies)

## What we don't know :

- Normal or inverted hierachy ?
- Dirac or Majorana particle ?
- CP violating phases in mixing matrix ?
- **No information about absolute mass scale ! (only upper limits)**
- Existence of sterile neutrinos ?



## $\beta$ -decay: absolute $\nu$ -mass

model independent, kinematics

status:  $m_\nu < 0.8$  eV

potential:  $m_\nu \approx 0.2$  eV

e.g.: KATRIN, Project-8, ECHO  
HOLMES, NuMECS

## $0\nu\beta\beta$ -decay: eff. Majorana mass

model-dependent (CP-phases)

status:  $m_{\beta\beta} < 36$  meV to 156 meV

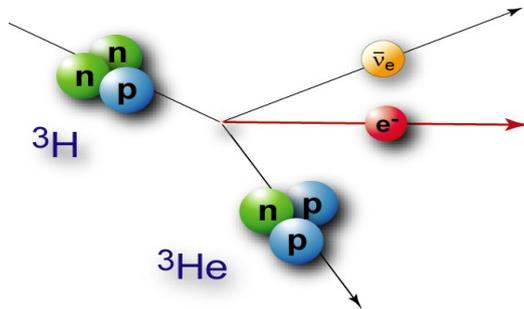
[<sup>136</sup>Xe, *Phys. Rev. Lett.* 130, 051801 (2023)]

potential:  $m_{\beta\beta} \approx 20$ -50 meV

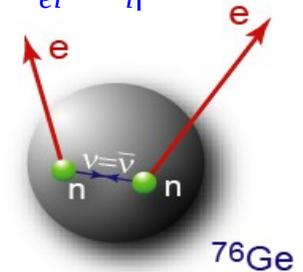
e.g.: KamLAND-Zen, GERDA, CUORE, EXO,  
SNO+, Majorana, Nemo 3, COBRA

## neutrino mass measurements

$$m_\nu^2 = \sum |U_{ei}|^2 m_i^2$$



$$m_{\beta\beta} = \left| \sum U_{ei}^2 m_i \right|$$



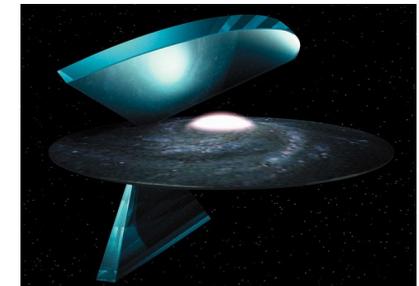
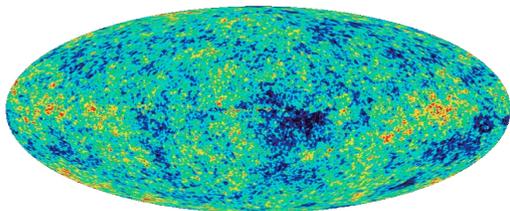
$$\sum m_i$$

## cosmology: $\nu$ hot dark matter $\Omega_\nu$

model dependent, analysis of CMB and  
structure formation data

status:  $\sum m_i < 0.11$  eV

[Planck PR4 release, *A&A*, 682 (2024) A37]



$$\frac{d\Gamma}{dE} = C p(E+m_e)(E_0-E)\sqrt{(E_0-E)^2 - m_\nu^2} F(Z+1, E) \Theta(E_0-E-m_\nu) S(E)$$

$$C = \frac{G_F^2}{2\pi^3} \cos^2 \theta_C |M|^2$$

(modified by final state distribution, recoil corrections, radiative corrections, ...)

$$m_\nu^2 = \sum_{i=1}^3 |U_{ei}|^2 m_i^2$$

Suitable Isotopes:

## Tritium

- $E_0 = 18.6$  keV,  $T_{1/2} = 12.3$  a
- $S(E) = 1$  (super-allowed)

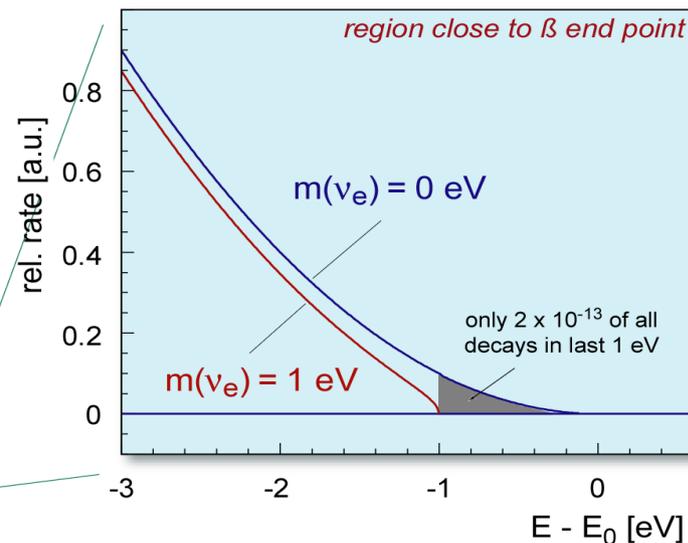
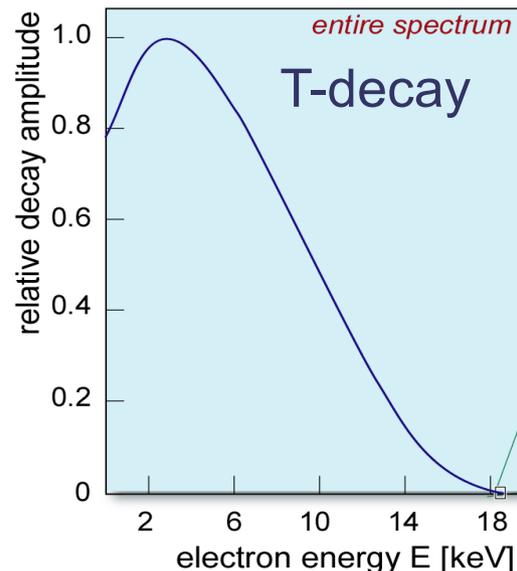
## Rhenium

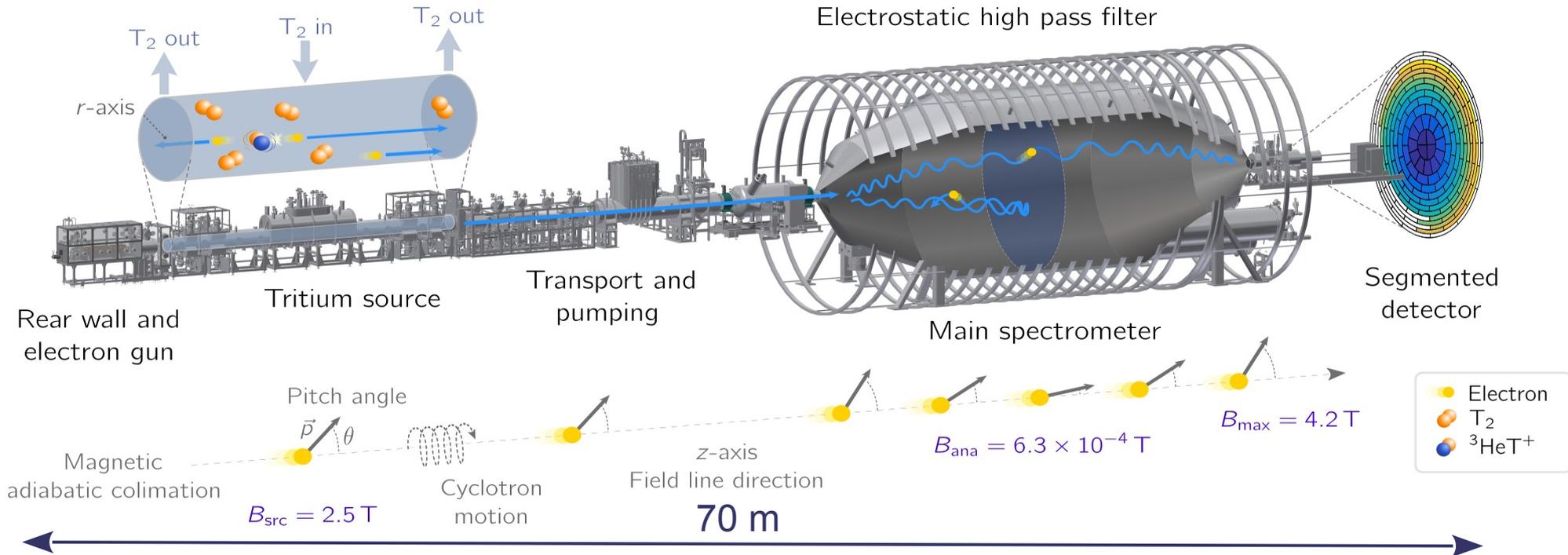
- $E_0 = 2.47$  keV,  $T_{1/2} = 43.2$  Gy

alternative approach:

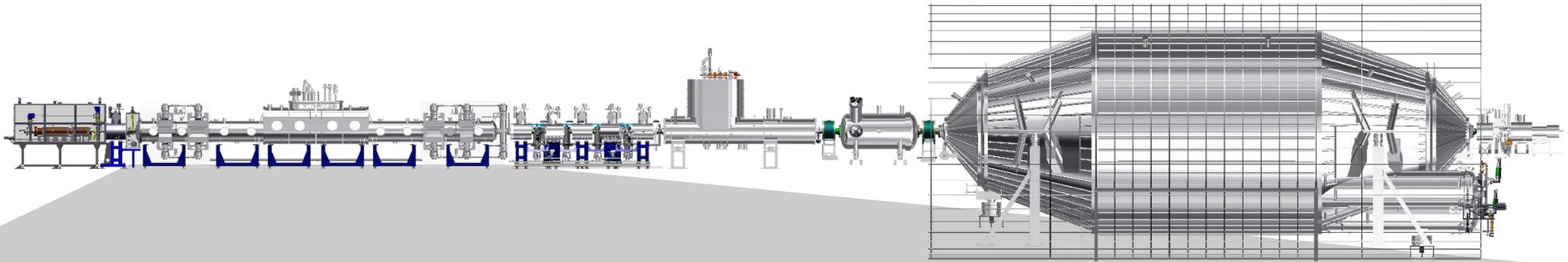
## Holmium (EC decay)

- $Q_{EC} \approx 2.5$  keV,  $T_{1/2} = 4570$  y

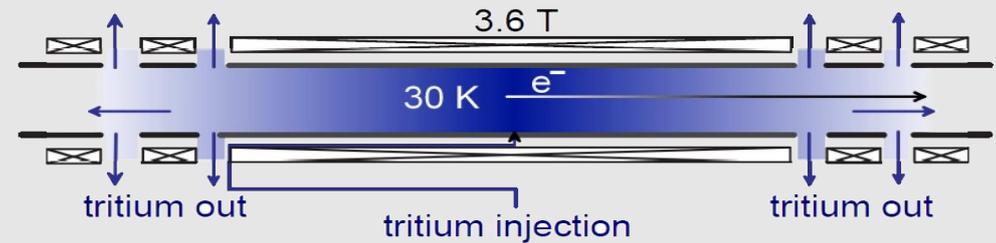




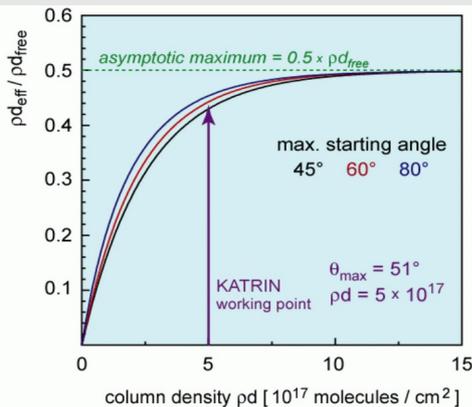
# Windowless Gaseous Tritium Source



- beam tube  $\varnothing = 9 \text{ cm}$  ,  $L = 10 \text{ m}$
- guiding field  $2.5 \text{ T}$
- Temperature  $80 \text{ K} \pm 0.01 \text{ K}$
- $T_2$  purity  $95\% \pm 0.1 \%$
- column density  $5 \cdot 10^{17} T_2 / \text{cm}^2$
- luminosity  $1.7 \cdot 10^{11} \text{ Bq}$



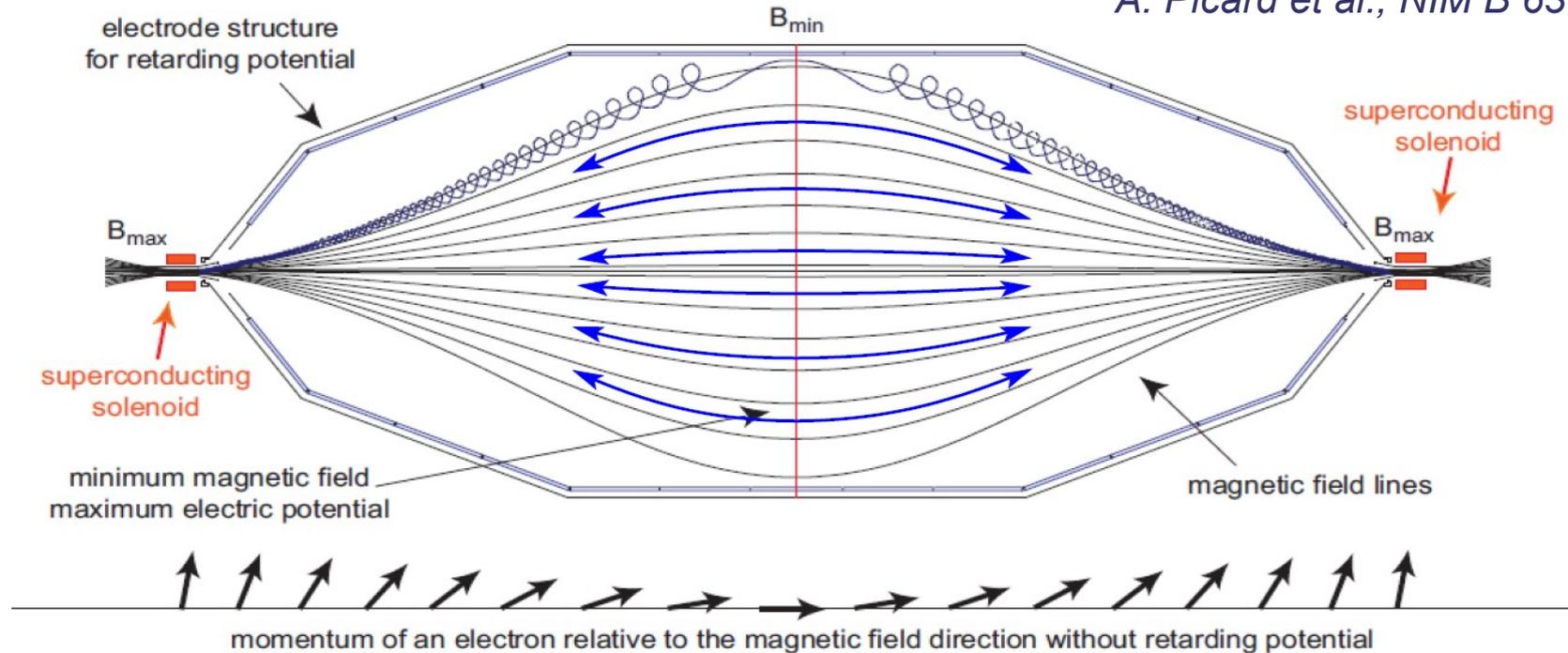
- $T_2$  flow rate  $5 \cdot 10^{19} \text{ molecules/s}$   
(40 g of  $T_2$  / day)



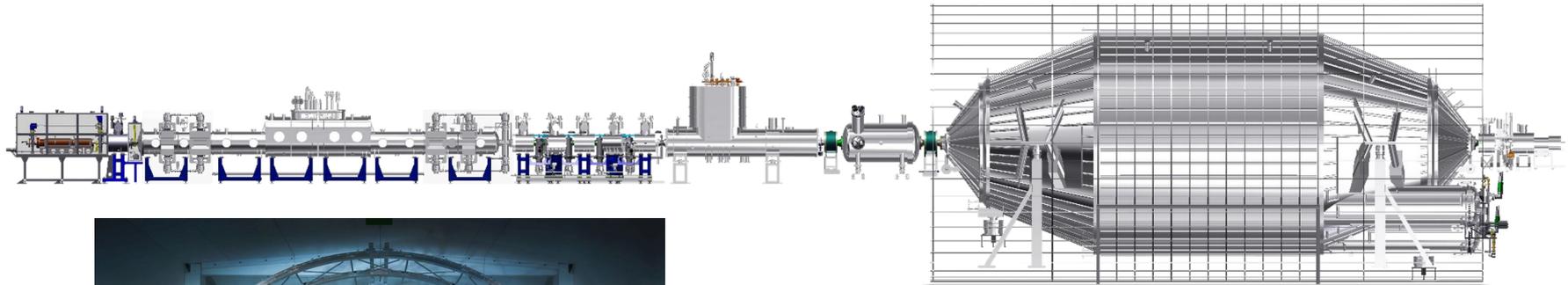
WGTS at Tritium Laboratory Karlsruhe

## Magnetic Adiabatic Collimation with Electrostatic Filter

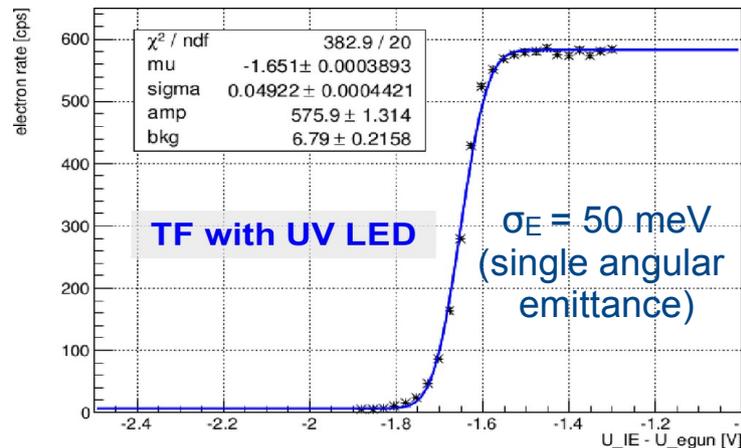
A. Picard et al., NIM B 63 (1992)



- adiabatic transport  $\rightarrow \mu = E_{\perp} / B = \text{const.}$
- $B$  drops by  $2 \cdot 10^4$  from solenoid to analyzing plane  $\rightarrow E_{\perp} \rightarrow E_{\parallel}$
- only electrons with  $E_{\parallel} > eU_0$  can pass the retardation potential
- Energy resolution  $\Delta E = E_{\perp, \text{max, start}} \cdot B_{\min} / B_{\max} < 1 \text{ eV}$

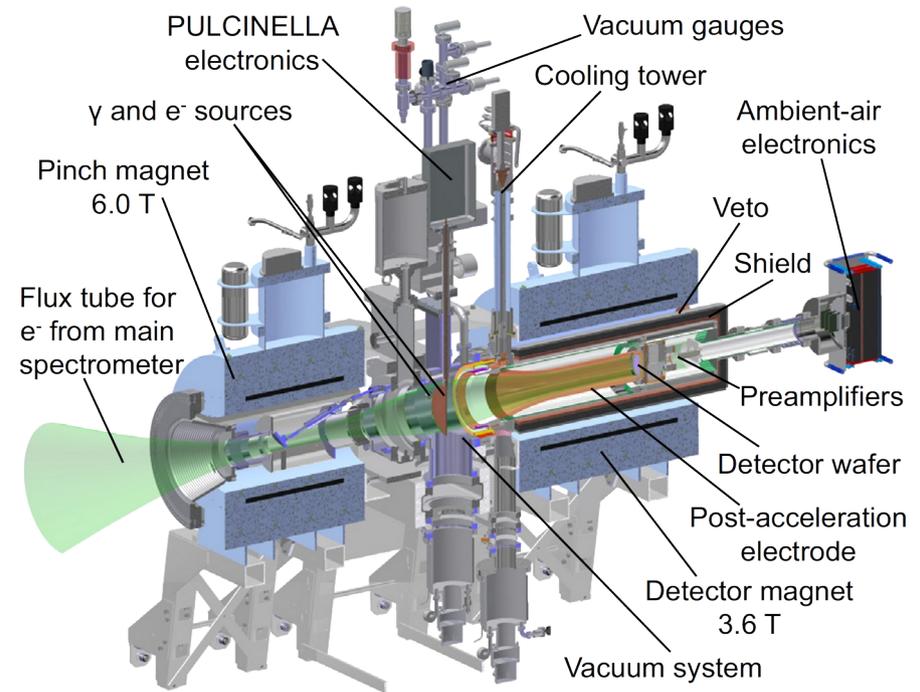


- 18.6 kV retardation voltage,  $\sigma < 60$  meV
- energy resolution 2.7 eV (SAP mode)
- pressure  $< 10^{-11}$  mbar
- Air coils for earth magnetic field compensation
- Double layer wire electrode for background reduction and field shaping



## Focal plane detection system

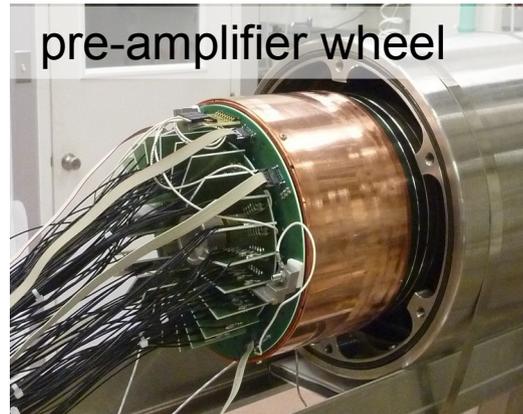
- segmented Si PIN diode:  
90 mm  $\varnothing$ , 148 pixels, 50 nm dead layer
- energy resolution  $\approx 1$  keV
- pinch and detector magnets up to 6 T
- 10 kV post acceleration
- active veto shield



detector magnets at KIT



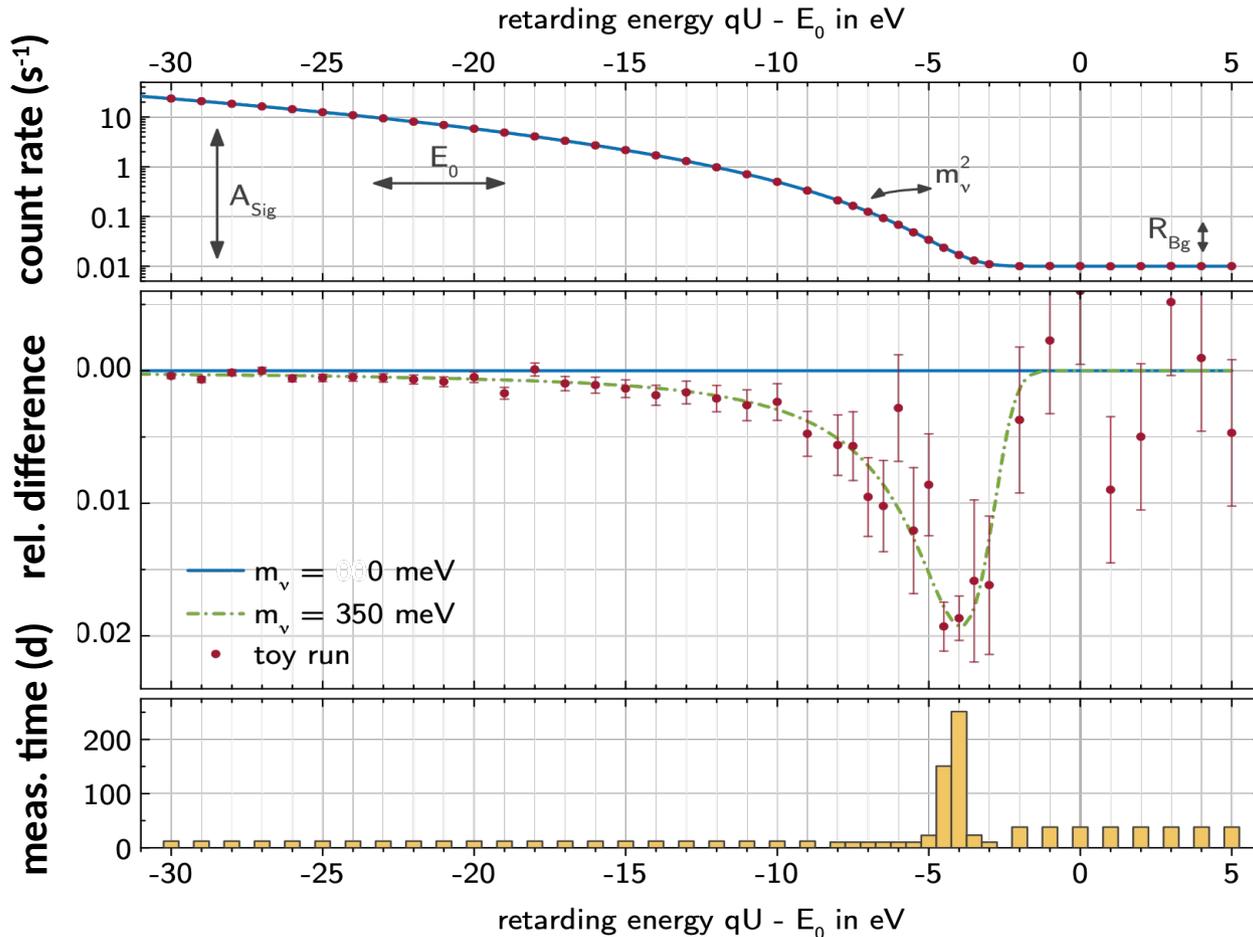
pre-amplifier wheel



segmented Si-PIN wafer



- Direct shape measurement of integrated  $\beta$  spectrum



Four fit parameters:

spectrum  
ampl.  $A_{sig}$

spectrum  
endpoint  $E_0$

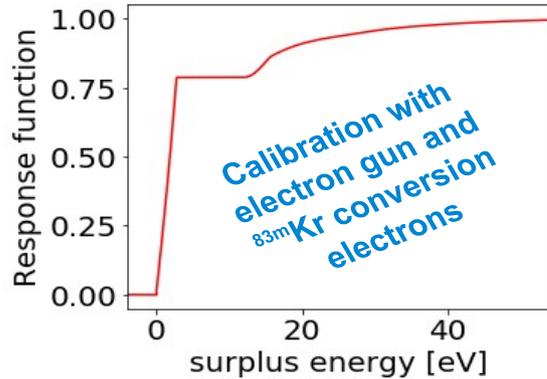
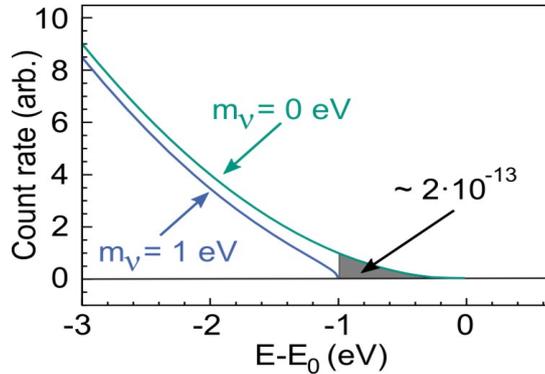
background  
rate  $R_{bg}$

squared  
mass  $m_\nu^2$

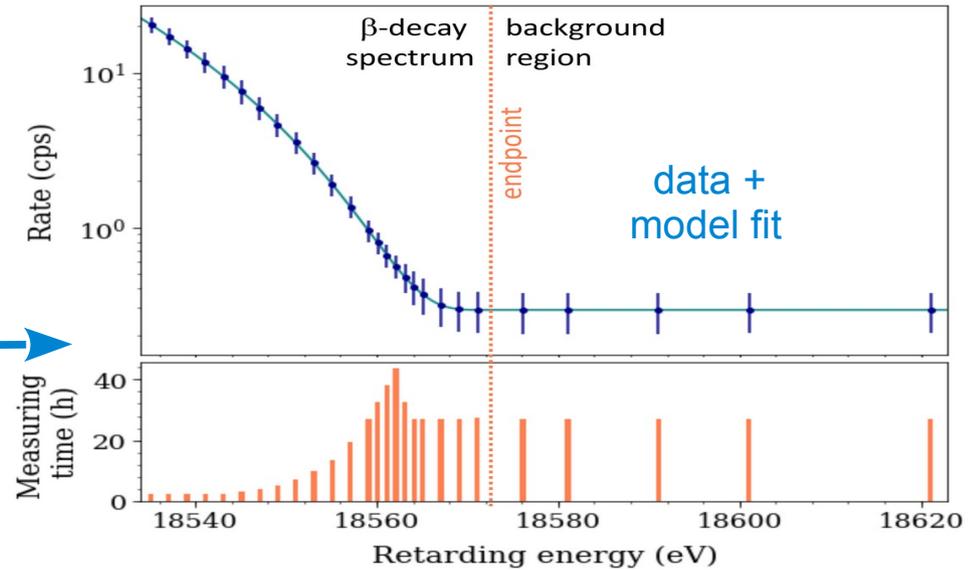
$\sim 10^{-8}$  of all  $\beta$ -decays in scan  
region  $\sim 40$  eV below endpoint

*Eur. Phys. J. C 79 (2019) 204*

Beta spectrum:  $R_\beta(E, m^2(\nu_e))$



KATRIN  $\beta$ -scan



$$R(qU) = A_s \cdot N_T \int_{qU}^{E_0} R_\beta(E, m^2(\nu_e)) \cdot f(E - qU) dE + R_{bg}$$

*PRL 123 (2019) 221802, EPJ C 79 (2019) 204*

*+ detailed analysis PRD 104 (2021) 012005*

*+ energy loss measurement EPJ C 81 (2021) 579*

Experimental response:  $f(E - qU)$

## 1st campaign (spring 2019)

- total statistics: 2 million events
- best fit result:  $m_\nu^2 = -1.0_{-1.1}^{+0.9} \text{ eV}^2$
- mass limit:  $m_\nu < 1.1 \text{ eV}$  (90% CL)

## 2nd campaign (autumn 2019)

- total statistics: 4.3 million events
- best fit result:  $m_\nu^2 = 0.26_{-0.34}^{+0.34} \text{ eV}^2$
- mass limit:  $m_\nu < 0.9 \text{ eV}$  (90% CL)

## Combine 1<sup>st</sup> and 2<sup>nd</sup> campaign:

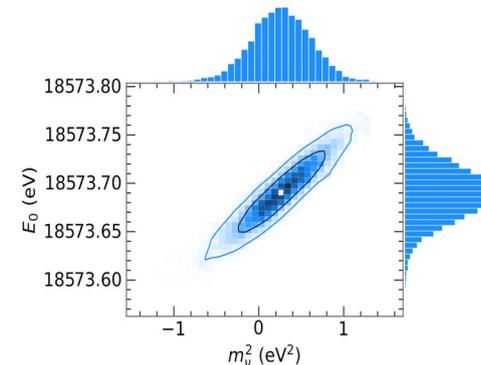
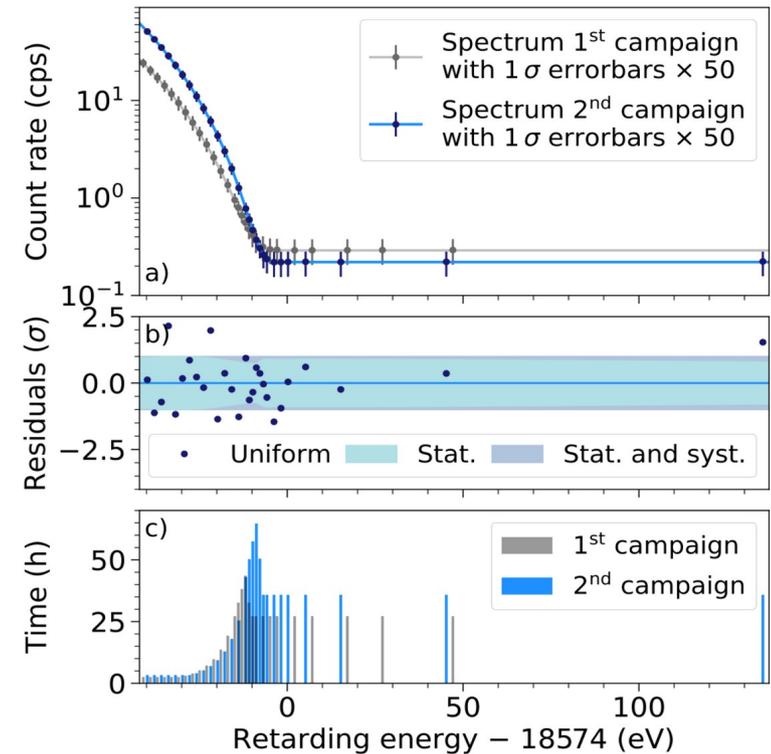
- mass limit:  $m_\nu < 0.8 \text{ eV}$  (90% CL)

## Cross-check: endpoint energy

$E_0 = 18573.69 \pm 0.03 \text{ eV} \rightarrow$  Q-value:  $18575.2 \pm 0.5 \text{ eV}$

$\rightarrow$  good agreement with Penning trap experiments:

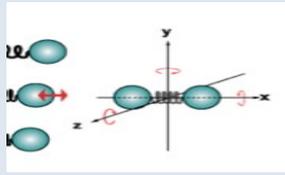
$Q = 18575.72 \pm 0.07 \text{ eV}$  PRL 114 (2015) 013003



*Nature Physics*  
18 (2022) 160

## Molecular final states

- Quantum-chemical computations



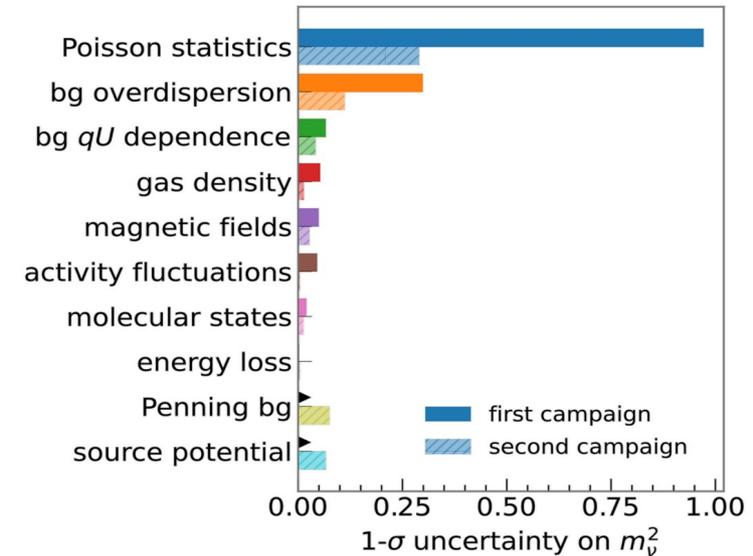
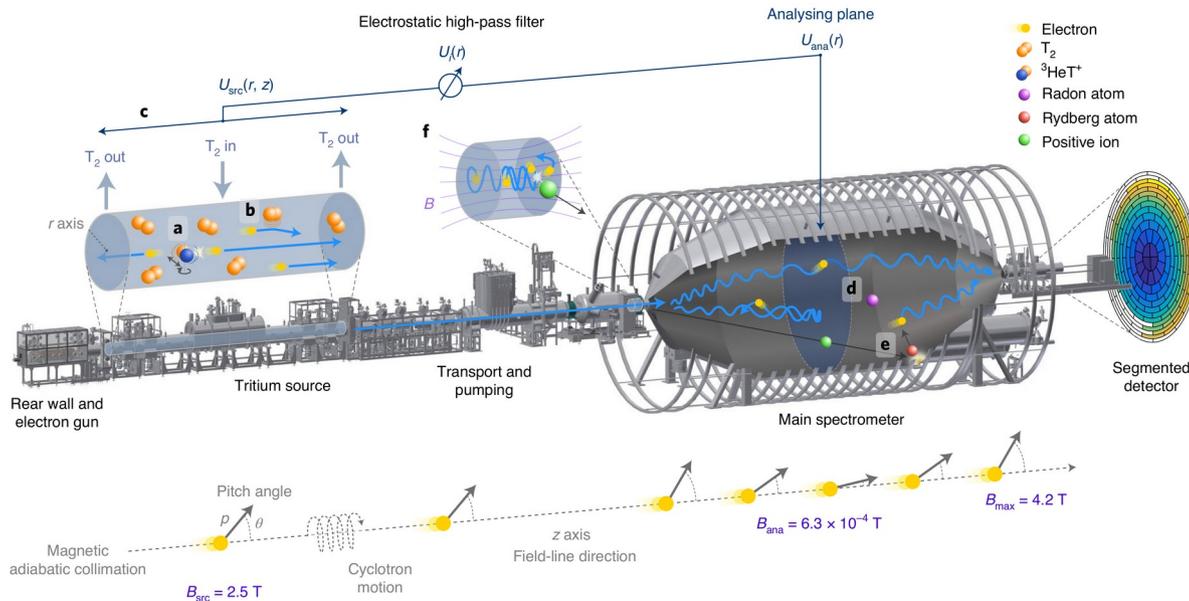
## Source electric potential

- plasma properties
- surface conditions

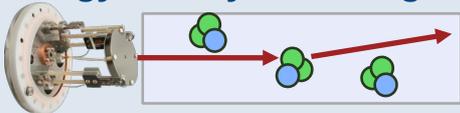


## Magnetic fields

- source
- spectrometer
- detector



## Energy loss by scattering

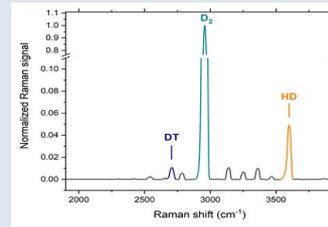


EPJ C 79 (2019) 204  
EPJ C 81 (2021) 579

## Activity fluctuations

- column density
- Tritium purity ( $T_2$ , DT, HT)

Sensors 20 (2020) 4827



## Background

- Volume dependent "Rydberg" background
- Voltage dependent background
- Time structure due to trapped electrons

JINST 13 (2018) T10004

Eur. Phys. J. A 44 (2010) 499

Astropart. Phys. 138 (2022) 102686

Main sources of background electrons originating inside or between the spectrometers:

- Time-dependent background due to inter spectrometer Penning trap
- Non-poissonian distributed background due to stored electrons from Radon decays
- Volume dependent background from ionization of Rydberg states created by radioactive decays on the inner vessel surface

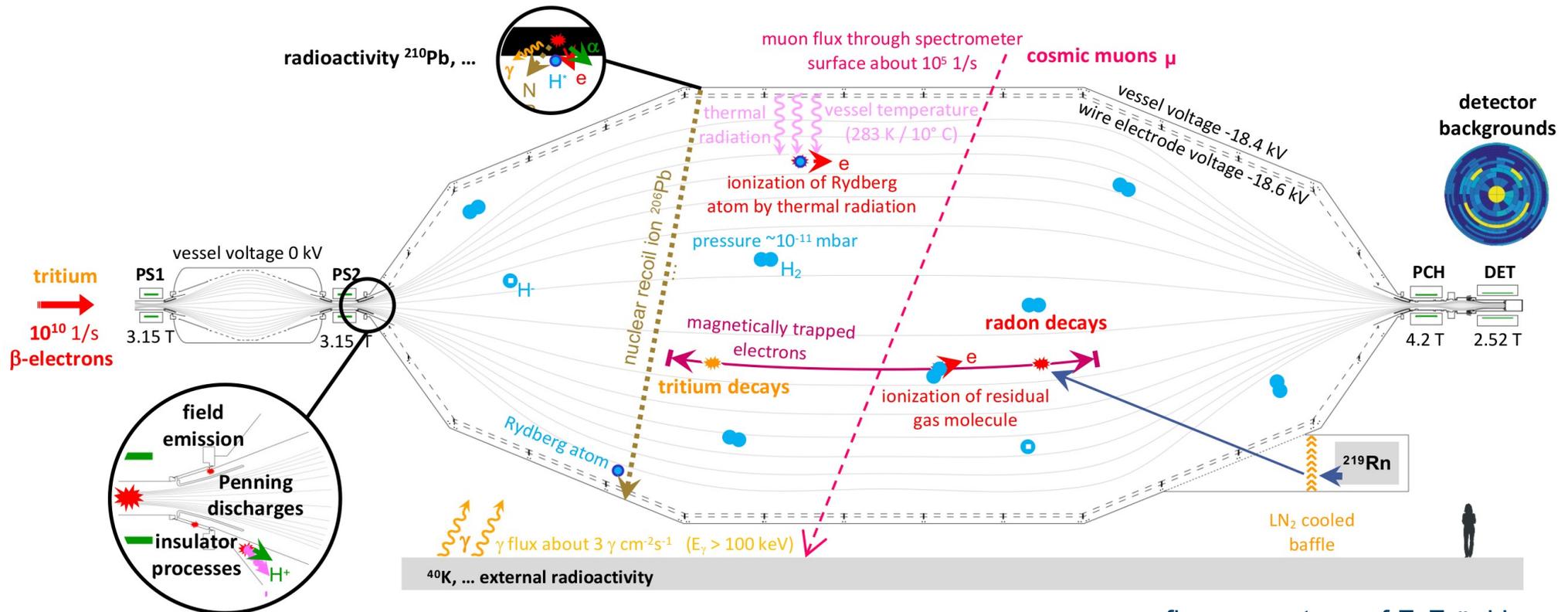
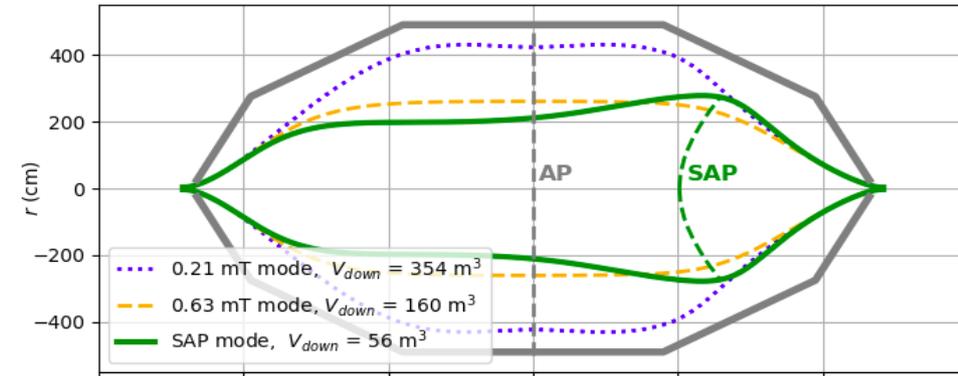


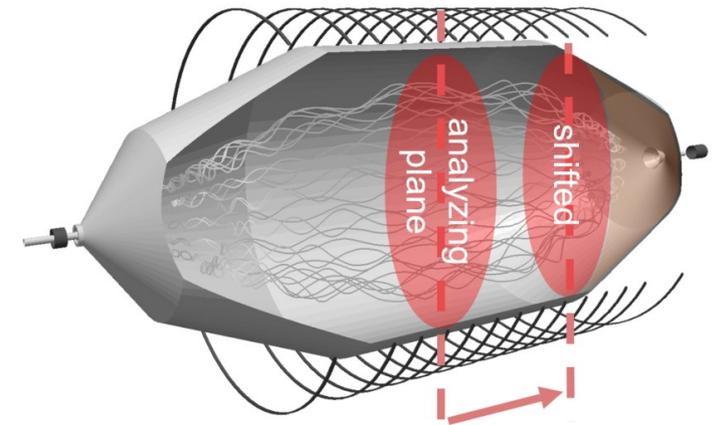
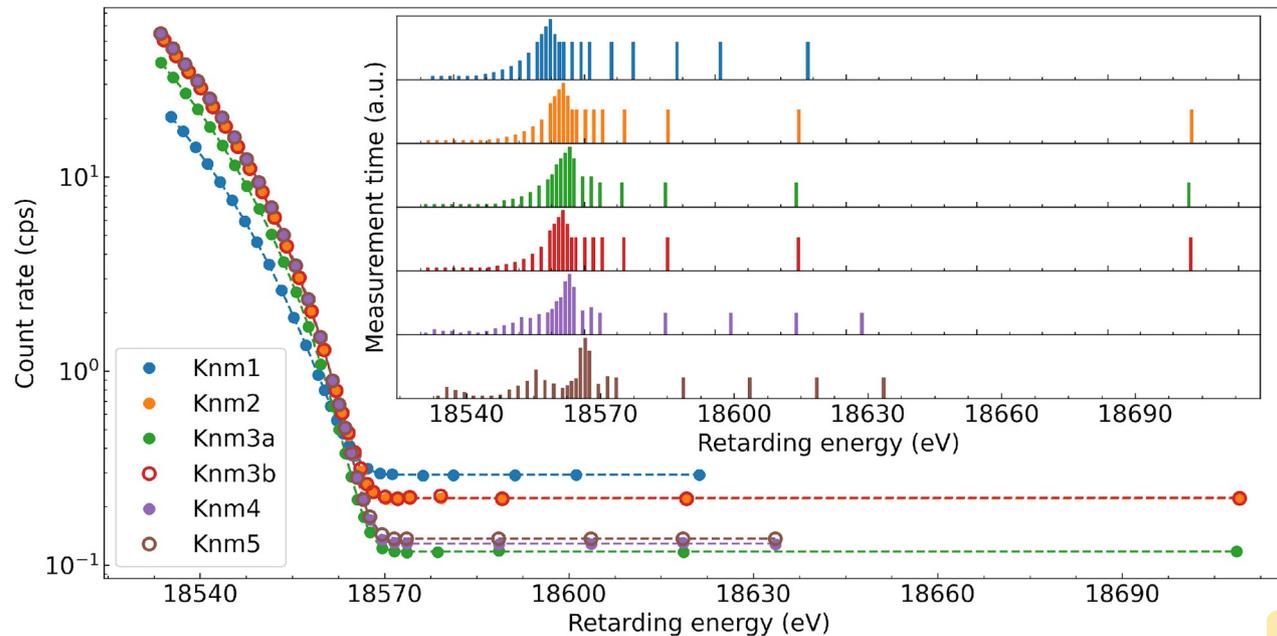
figure courtesy of F. Fränkle

# Reduction of experimental backgrounds

- Reduction of radon related backgrounds by LN<sub>2</sub> baffles
- Removal of inter-spectrometer Penning trap to eliminate background time dependence
- New spectrometer field configuration (shifted analysis plane) reduced background by factor 2 and removes non-poissonian backgrounds



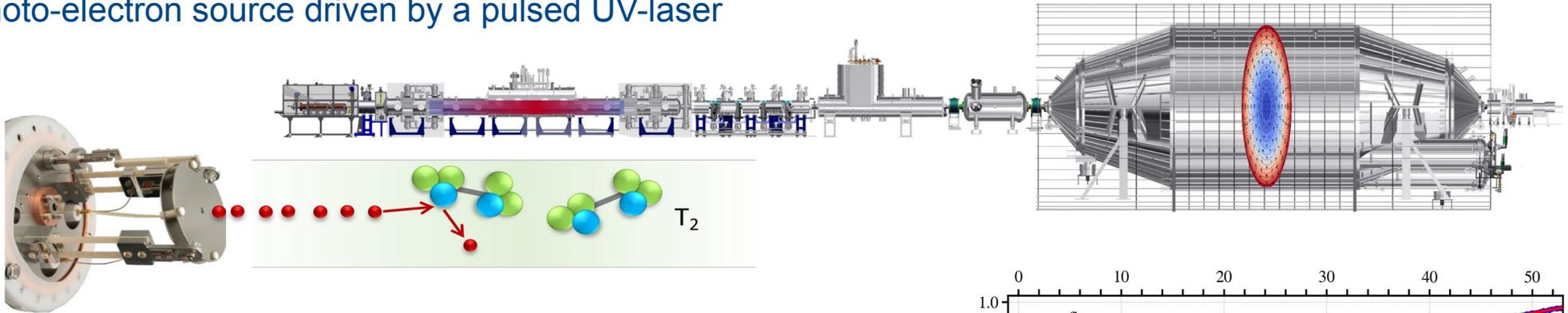
*Eur. Phys. J. C 82 (2022) 258*



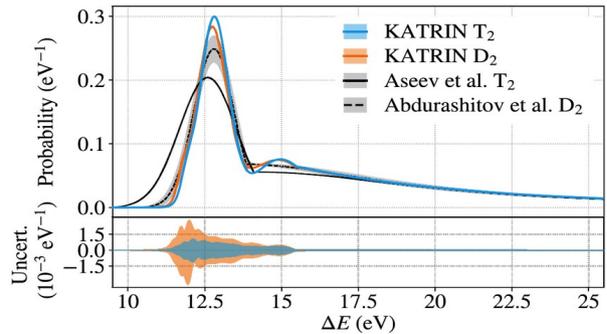
*see Talk by Joscha Lauer, Mo. 15:00*

# Response function measurement

- Determination of the experimental response function using a mono-energetic, angular selective photo-electron source driven by a pulsed UV-laser

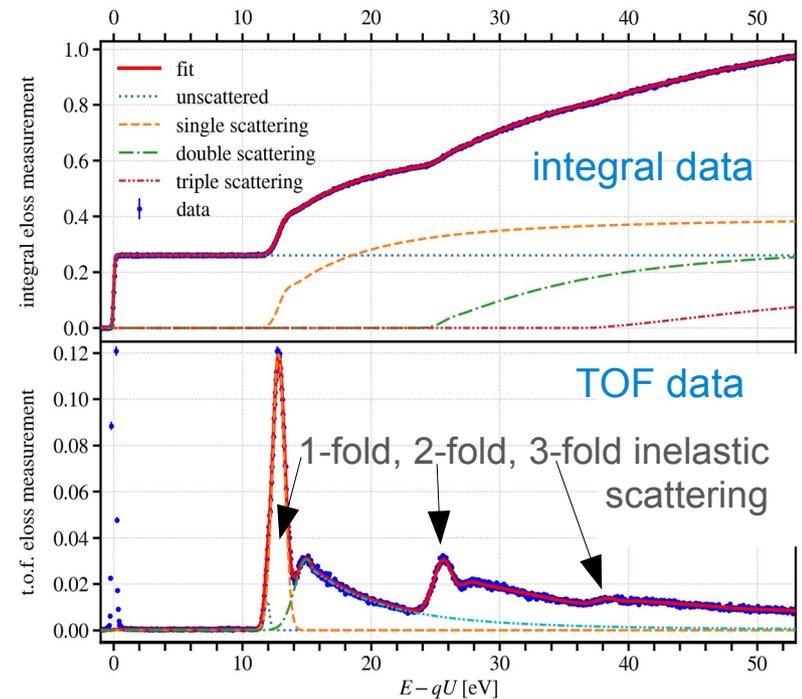


- Measurement of integral and differential (using TOF method) spectra at different column densities  
 → Extraction of spectrometer transmission function and energy losses due to scattering



EPJ C 81 (2021) 579

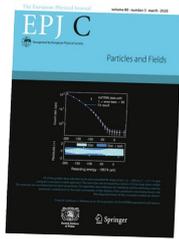
←  
Extract differential energy loss function





# KATRIN data taking

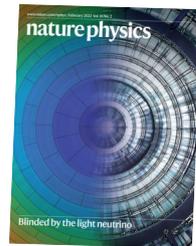
First tritium run



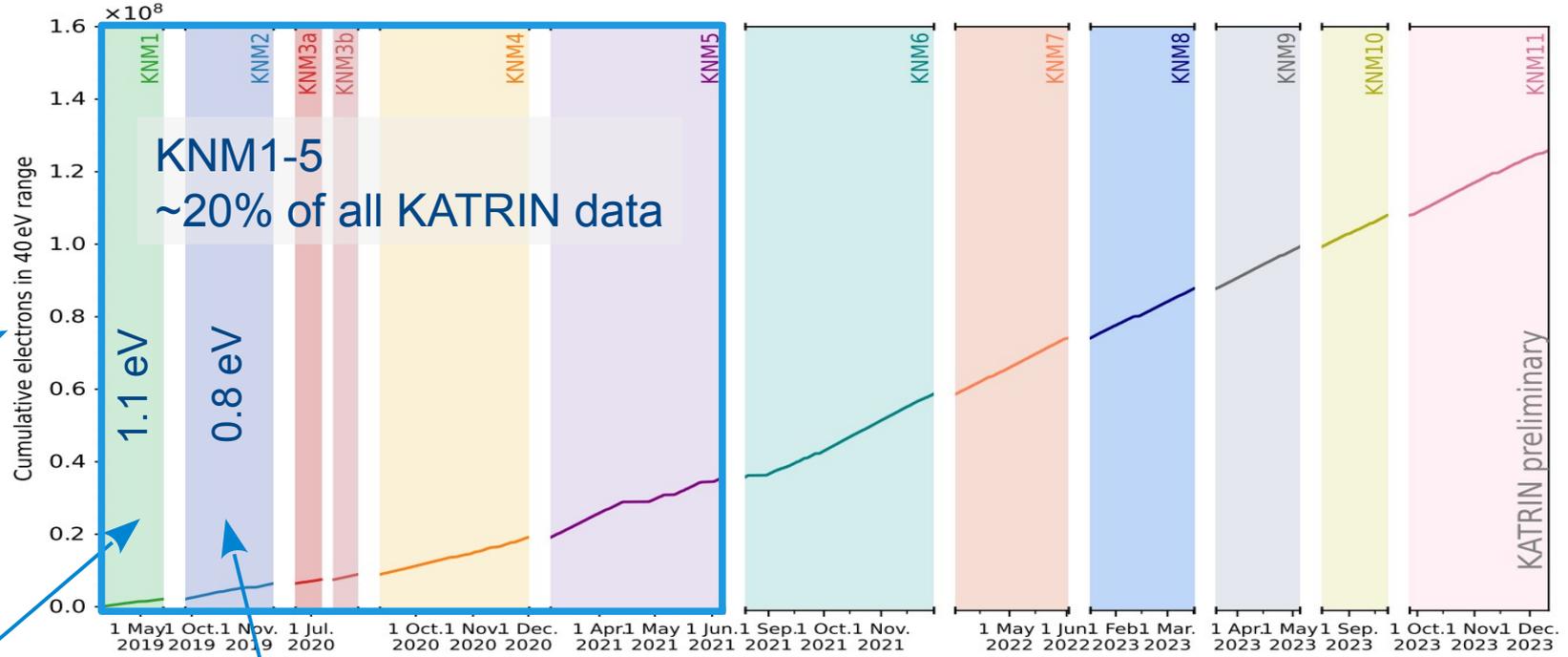
EPJC 80 (2020) 264



PRL 123 (2019) 221802



Nat. Phys. 18 (2022) 160



Light sterile neutrinos

Relic neutrinos

+

+



PR D 105 (2022) 072004



PRL 129 (2022) 011806

## Combined analysis of first five campaigns

- Currently in unblinding process
- Next data release in summer

## Sensitivity projection KNM1-5

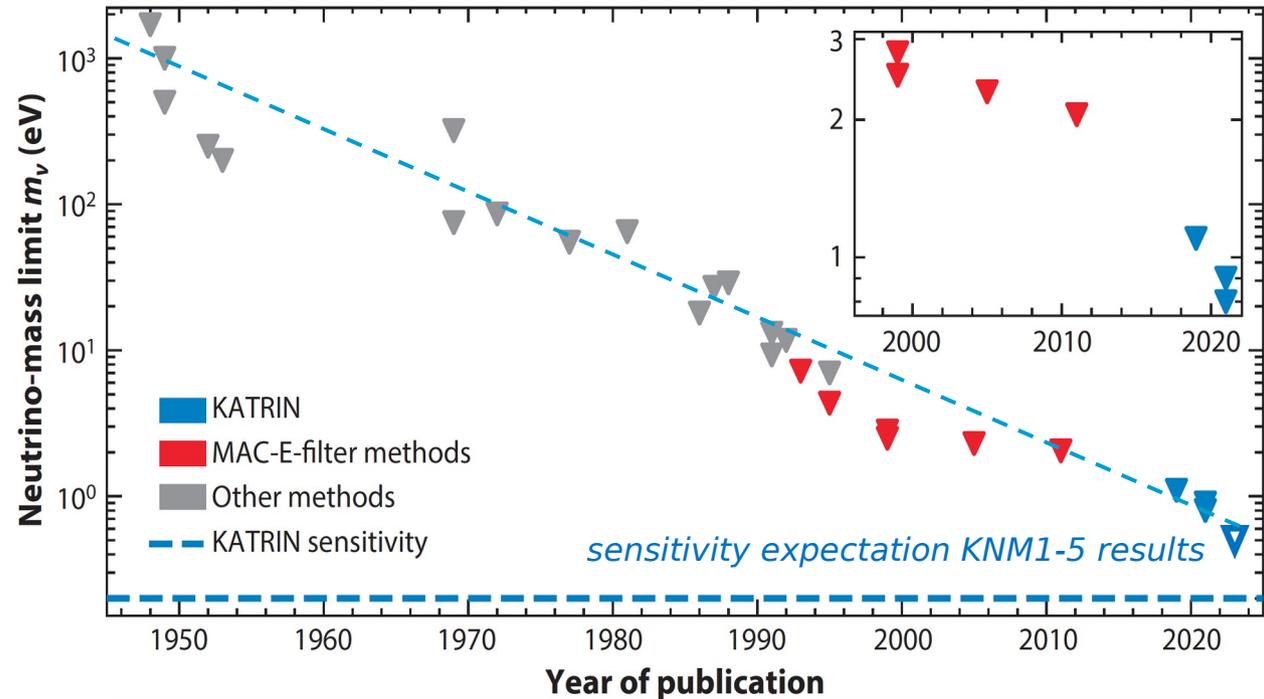
- $m_\nu \leq 0.5$  eV (90% CL)

## Improvements:

- Factor six in statistics
- Background reduction
- More accurate knowledge of systematics

## Continuation of current measurement program until end of 2025

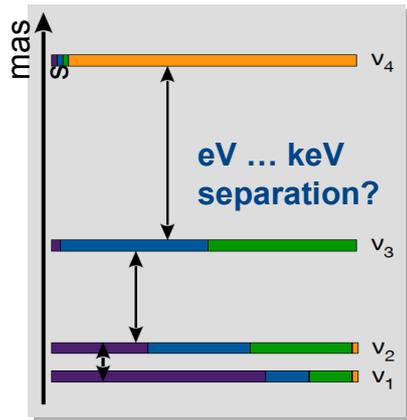
- collect 1000 days of beta scans
- expect to reach sensitivity  $m_\nu < 0.3$  eV (90% CL)



*Annu. Rev. Nucl. Part. Sci. 72 (2022) 259*

*More on analysis methods: talk  
by Weiran Xu, Thu. 16:00*

Is there a fourth  
(sterile) neutrino?

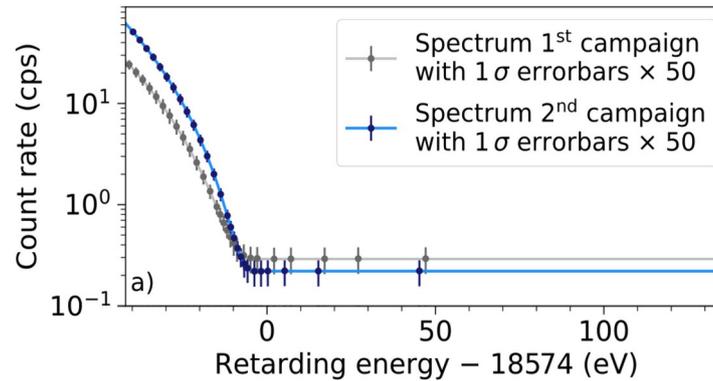


Neutrino mixing: “Kink” in  
normal  $\beta$ -spectrum (eV scale)  
or deep  $\beta$ -spectrum (keV scale)

*Phys. Rev. D* **105** (2022) 072004

*Phys. Rev. Lett.* **126** (2021) 091803

*JCAP* **02** (2015) 020



Search for exotic  
weak interactions  
(spectrum shape)

*JHEP* **01** (2019) 206

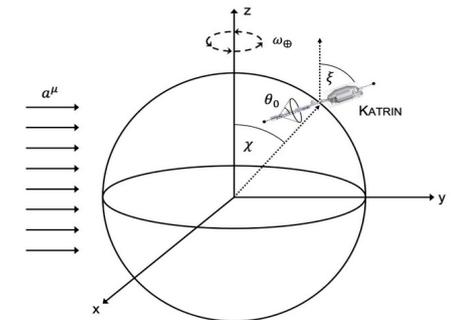
*PoS DISCRETE2022* (2024) 011

High statistics, high  
precision  $\beta$  spectrum

Search for Lorentz  
invariance violation  
(sidereal modulation)

Constrain local over-  
density of cosmic relic  
neutrinos (peak search)

*PRL* **129** (2022) 011806

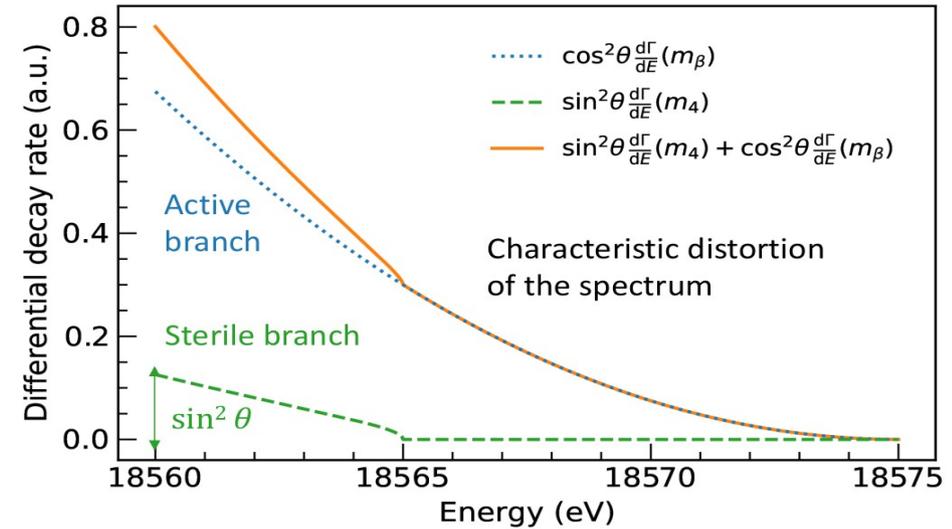
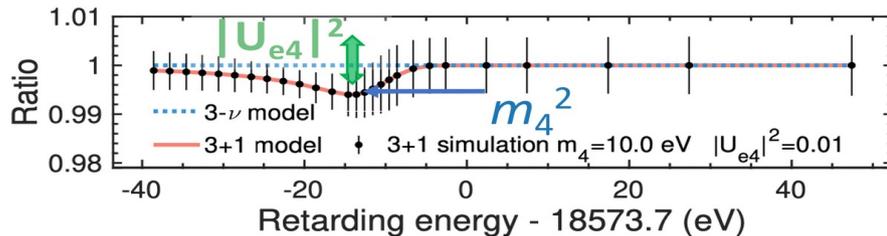
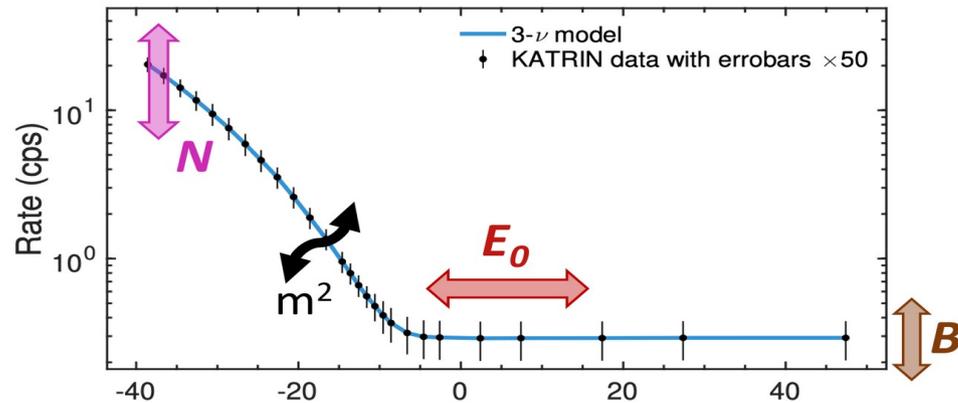


*Phys. Rev. D* **107** (2023) 082005

# (Light) sterile neutrino signature

- 3+1 sterile neutrino model
- Same data-set as for the neutrino mass
- Grid search in  $m_4, |U_{e4}|^2$  plane

$$\frac{d\Gamma}{dE} = \underbrace{(1 - |U_{e4}|^2) \frac{d\Gamma}{dE}(m_\beta^2)}_{\text{light neutrino}} + \underbrace{|U_{e4}|^2 \frac{d\Gamma}{dE}(m_4^2)}_{\text{heavy neutrino}}$$



## 6 Fit Parameters:

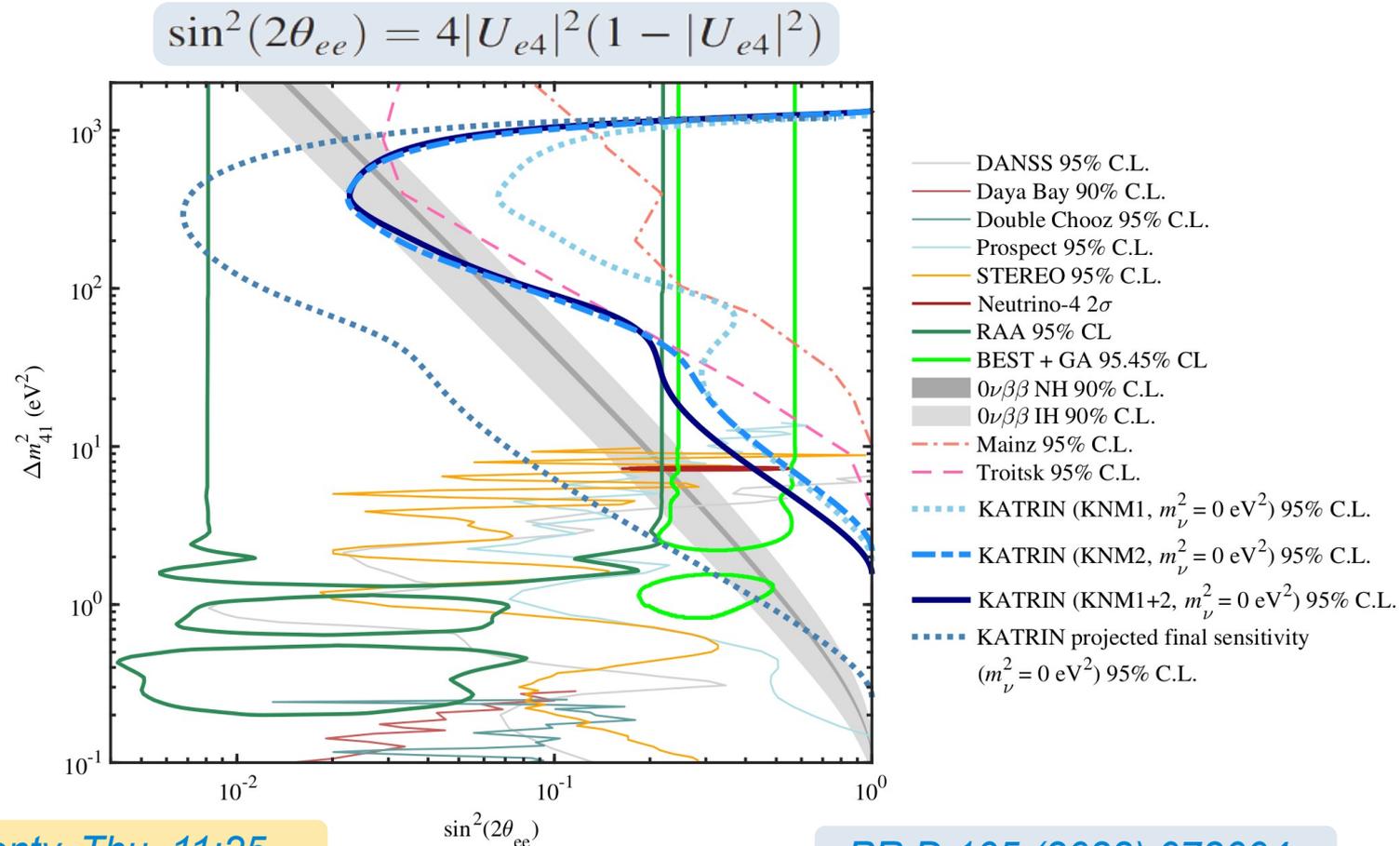
- $m^2$  - neutrino mass (free/constrained)
- $E_0$  - endpoint
- $N$  - amplitude
- $B$  - background rate
- $m_4^2$  - 4<sup>th</sup> neutrino mass
- $|U_{e4}|^2$  - 4<sup>th</sup> neutrino mixing

## Current dataset

- KATRIN starts to probe very interesting parameter space, complementary to oscillation searches
- approaching the BEST allowed regions with  $\Delta m_{41}^2 > 10 \text{ eV}^2$

## Final dataset

- Probing large portion of RAA, BEST and Neutrino-4 allowed regions
- comparable sensitivity to neutrinoless double  $\beta$ -decay



see Talk by Shailaja Mohanty, Thu. 11:25

PR D 105 (2022) 072004

## keV-scale sterile neutrino search:

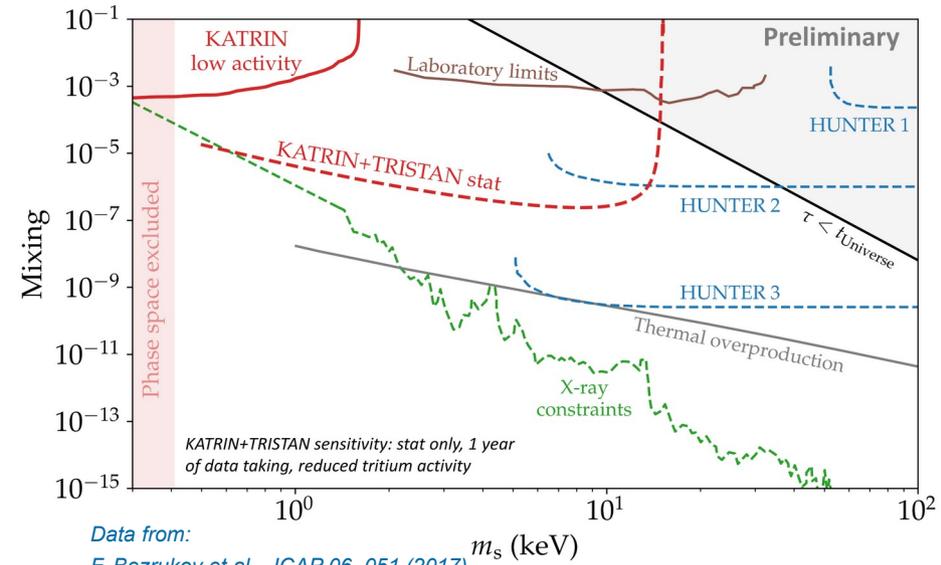
- sterile neutrinos on keV-scale are a viable candidate for dark matter in the universe (WDM)
- First scans deep into the  $\beta$ -spectrum during FT campaign at 0.5% c.d. [arXiv:2207.06337v1 \(2022\)](https://arxiv.org/abs/2207.06337v1)
- high-sensitivity search requires new high-rate detector system (TRISTAN) to handle huge electron rates from WGTS over large spectral range

## TRISTAN project in KATRIN:

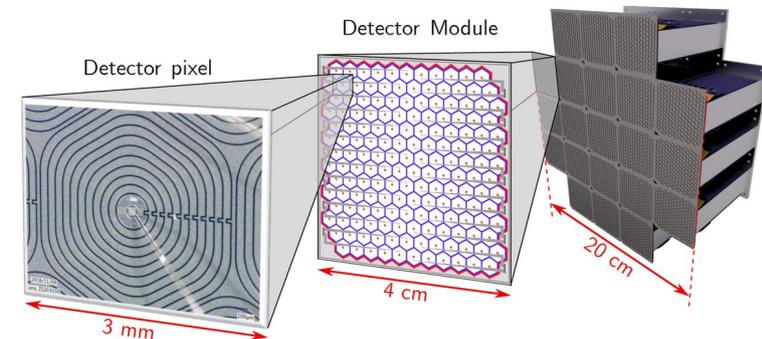
- novel multi-pixel Silicon Drift Detector array
- large count rates
- excellent energy resolution
- prototypes installed as monitoring devices @ KATRIN
- target sensitivity:  $\sin^2\theta < 10^{-6}$



see Talks by Anthony Onillon, Wed. 9:40  
and Daniela Spreng, Wed. 10:50

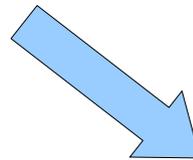
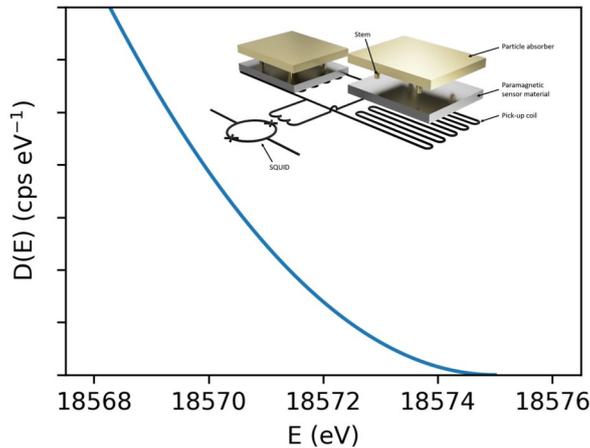
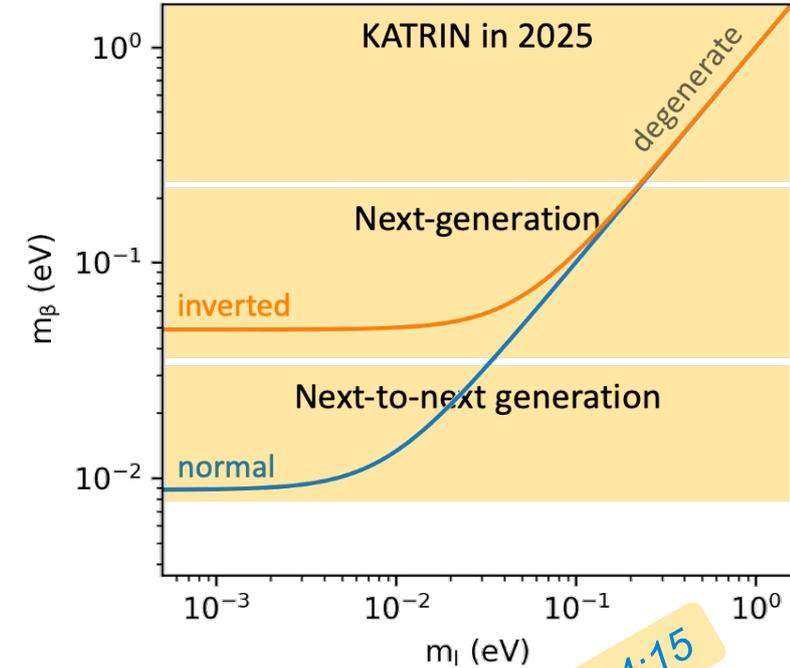
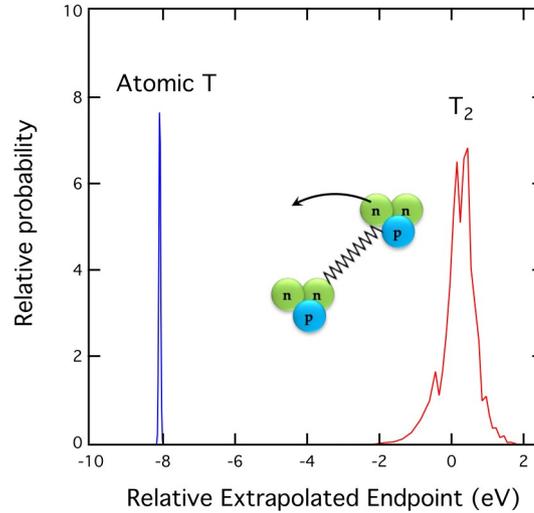


Data from:  
 F. Bezrukov et al., JCAP 06, 051 (2017)  
 J. N. Abdurashitov et al., JETP Letters 105, 12 (2017)  
 F. Benso et al., Phys. Rev. D 100, 115035 (2019)  
 C. J. Martoff et al., Quantum Sci. Technol. 6 024008 (2021)  
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## Advancing further with kinematic neutrino mass measurements

- New source concepts: remove spectral broadening due to molecular FSD by going to an atomic tritium source
- Novel detector technologies required for a differential measurement with  $O(eV)$  resolution



# KATRIN++

see Talk by Magnus Schösser, Tue. 14:15

- Studies of  $\beta$ -decay kinematics offer a model-independent way to determine the neutrino mass, complementary to cosmology and  $0\nu\beta\beta$  searches
- The KATRIN experiment has finalized the analysis of the first two science runs and published the first sub-eV neutrino mass limit with  $m_\nu < 0.8 \text{ eV}$
- Several improvements allowed to strongly reduce experimental background and systematic uncertainties
- Analysis of KNM3 to KNM5 science runs ongoing, analysis release expected in summer
- KATRIN has the capability to study several physics topics beyond neutrino mass:
  - eV-scale sterile neutrinos (first upper limits published)
  - keV scale sterile neutrinos (future project with new focal plane detector TRISTAN)
  - upper limit on local relic neutrino overdensity
  - investigations of Lorentz invariance
  - search for exotic weak interactions

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