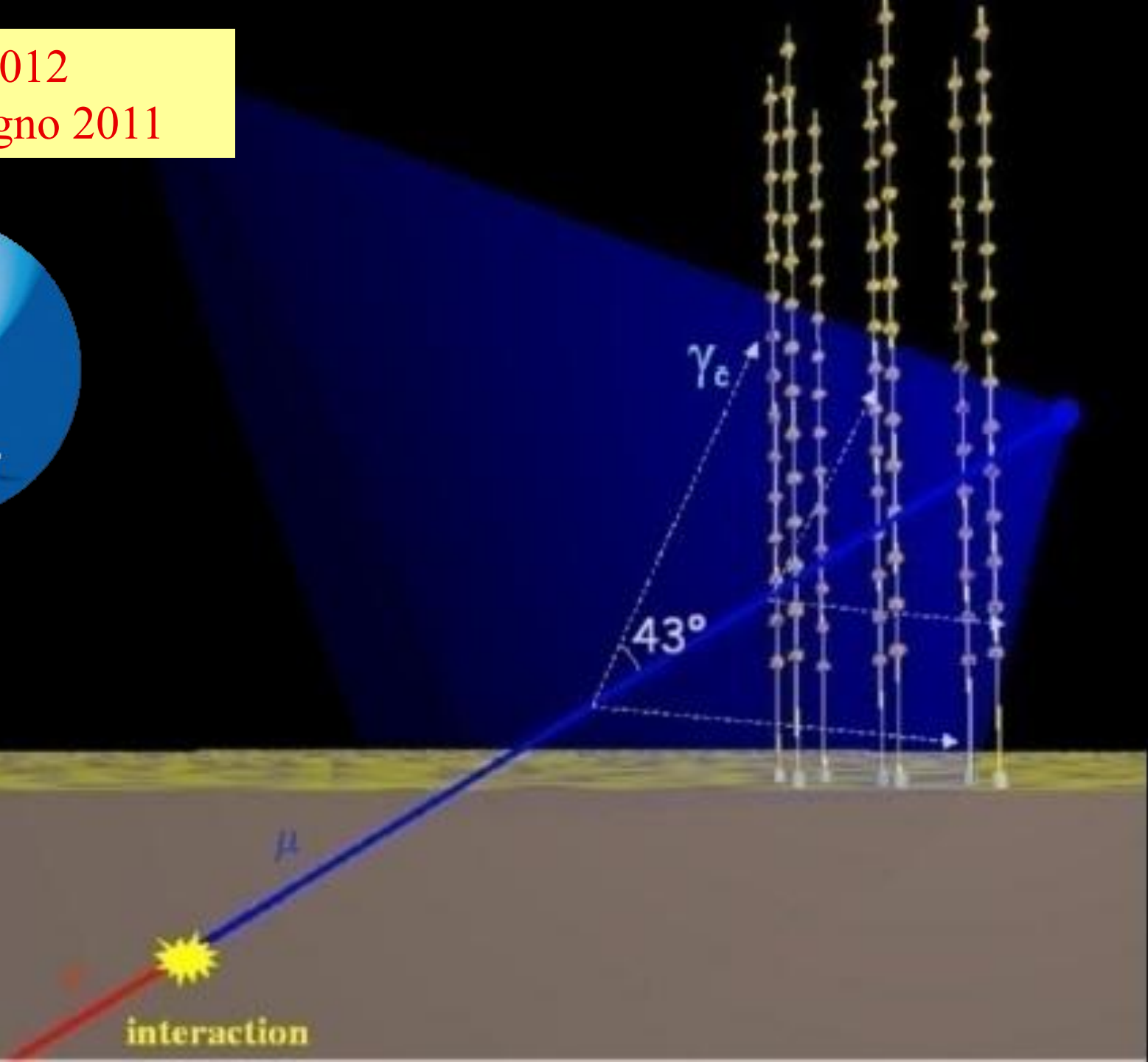
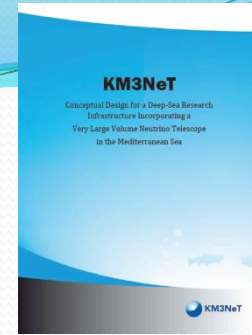
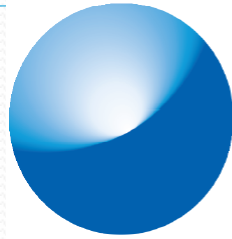
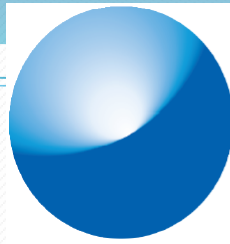


Preventivi 2012
Pisa 22 Giugno 2011

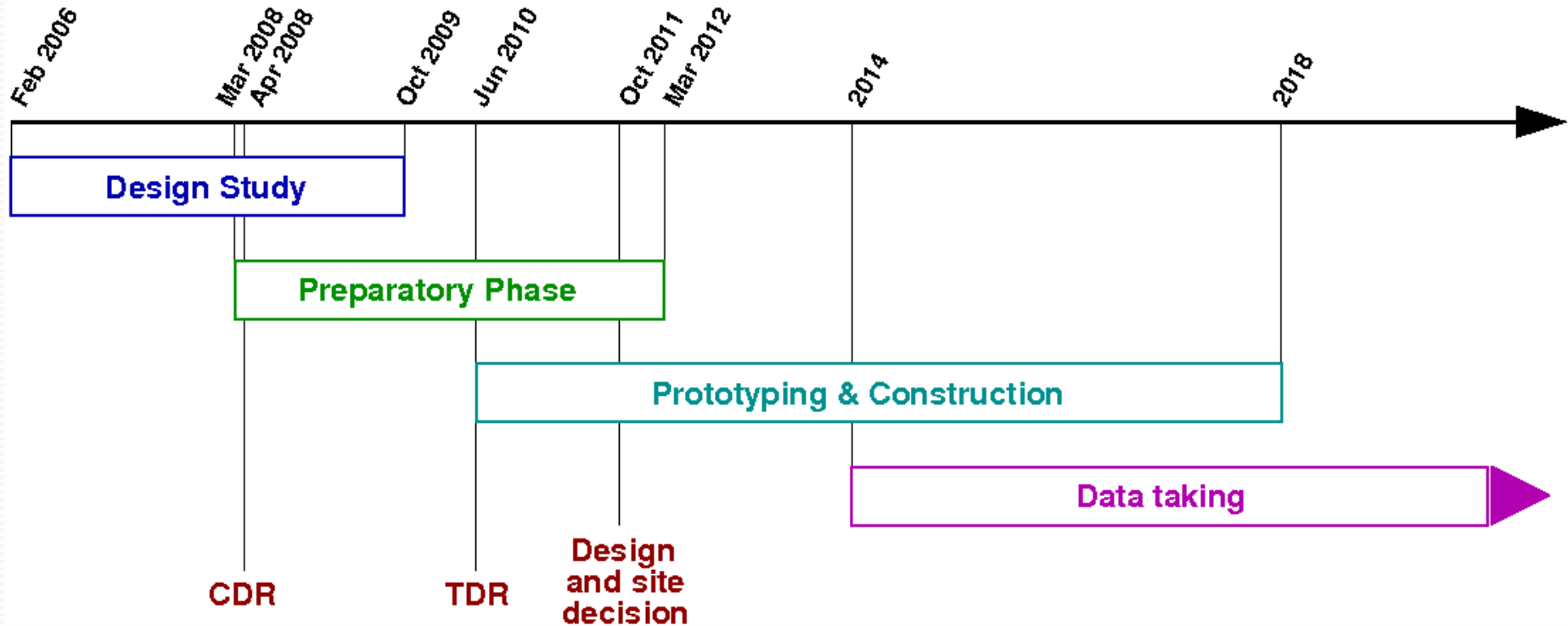


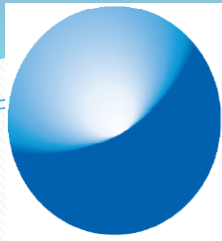


- Central physics goals:
 - Neutrino Astronomy under the Mediterranean Sea
 - Investigate neutrino “point sources” in the 100 GeV-1 PeV energy range
 - Complement IceCube field of view
 - Instrumented volume $> 5 \text{ km}^3$
- Implementation requirements:
 - Construction time ≤ 5 years
 - Operation over at least 10 years without “major maintenance”
- KM3NeT consortium consists of 40 European institutes, including those in Antares, Nemo and Nestor, from 10 countries (Cyprus, France, Germany, Greece, Ireland, Italy, The Netherlands, Rumania, Spain, U.K)
- KM3NeT is included in the ESFRI and ASPERA roadmaps
- Design Study (2006-2009) funded by the EU VIth Framework Program
- Conceptual Design Report (ISBN 978-90-6488-033-9) and Technical Design Report (ISBN 978-90-6488-031-5) available:
www.km3net.org/public.php
- KM3NeT PreparatoryPhase (2008-2012) funded by the EU VIIth Framework Program
 Final design, production plans for the detector elements and infrastructure features. In-situ prototype validation is underway. Legal, governance and funding aspects are also under study.



KM3NeT





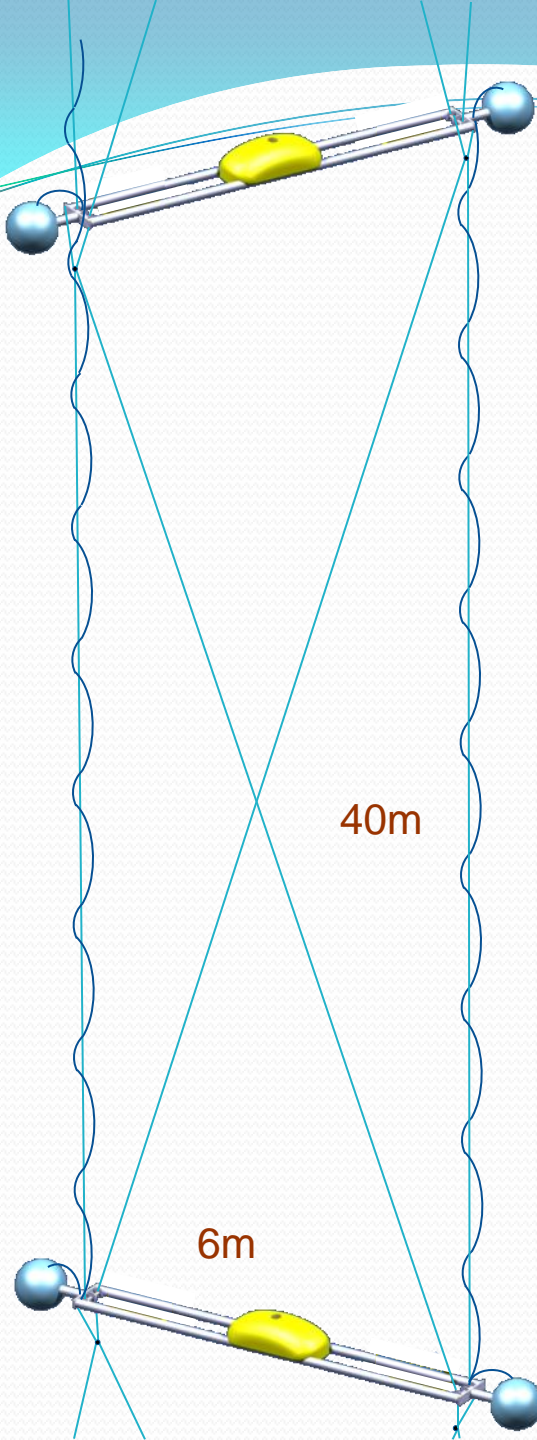
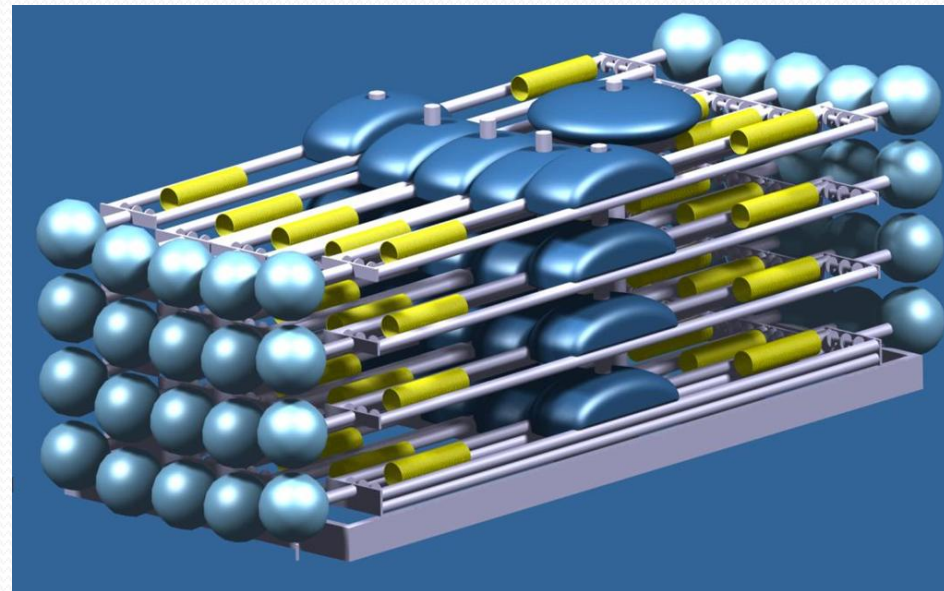
KM3NeT

Problems:

- *the ambient hydrostatic pressure;*
 - *the corrosive environment of the seawater;*
 - *the distance from shore for the communication;*
 - *the force on the structure due to the sea currents;*
 - *the backgrounds due to downward going muons;*
 - *the backgrounds due to ^{40}K decay.*
-
- *optimal angular resolution of the reconstructed muon;*
 - *large sensitive area facing the Galactic centre.*

The packed flexible tower (20 storeys)

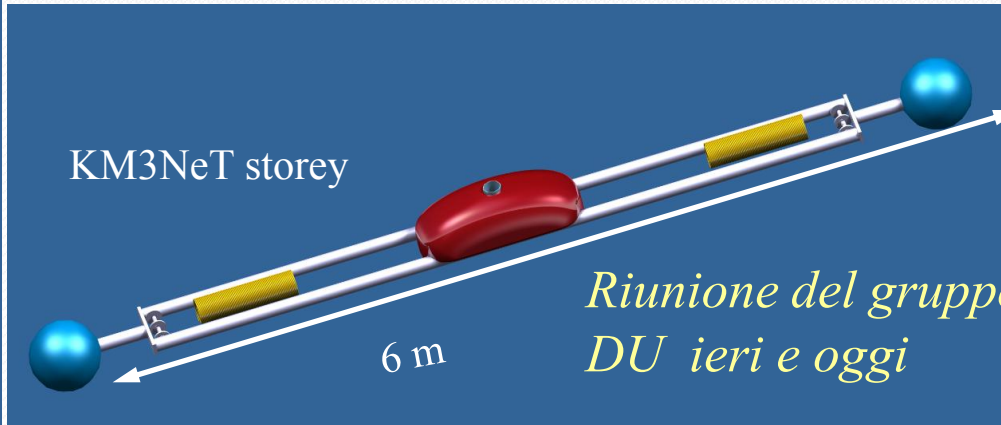
- Compact package
- Self unfurling
- Connection to seabed network by Remotely Operated Vehicle



Storeys	20
Height	900m
Compact Package	6 x 2.5 x 2.5m
Top drift @ 30 cm/s	~120 m
Total buoyancy	~10 kN
EO Cable	2 x 6.35 mm OD

KM3Net Detection Unit

20 storeys / 40 m distance (DU height ~ 900 m)
 storey: bar of 6 m, with 2 multi-PMT OM
 DOM: 31 x 3" PMT



KM3Net "Building Block" configuration

DU	St/DU	OM/St.	PMT/OM	PMT
154	20	2	31	190 960

Multi-PMT Optical Module

Self-contained "plug-and-play" module (17" pressure-resistant sphere)

- Photo-sensors 31 (19+12) 3" PMTs Equivalent of 4 x 8" PMTs
- Includes: All read-out/control electronics and calibration devices
- Single colour point to point connection via DWDM between each OM and the shore station.



KM3Net
 DOM: multi-PMT OM

Distinguish single from multiple photon hits:

- Photon counting = PMT counting
- Background rejection - ⁴⁰K

Looking upward:

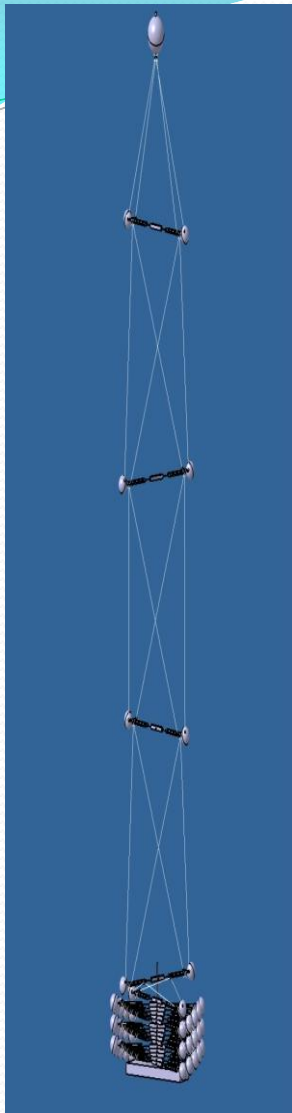
- Background rejection - atmospheric muons
- More uniform angular acceptance

Directionality:

- Signal photons from one side

Ageing:

- lower gain ~10⁶
- charge spread over multiple dynode chains

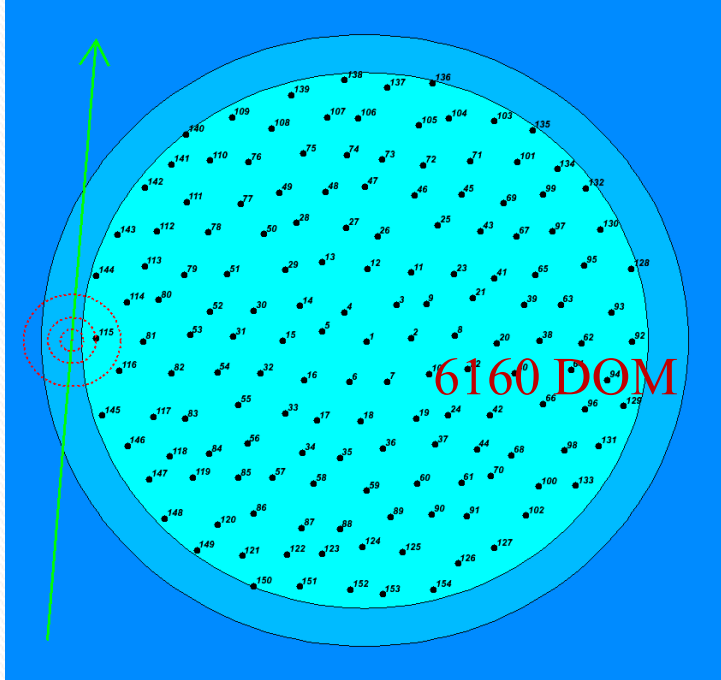


KM3Net DU

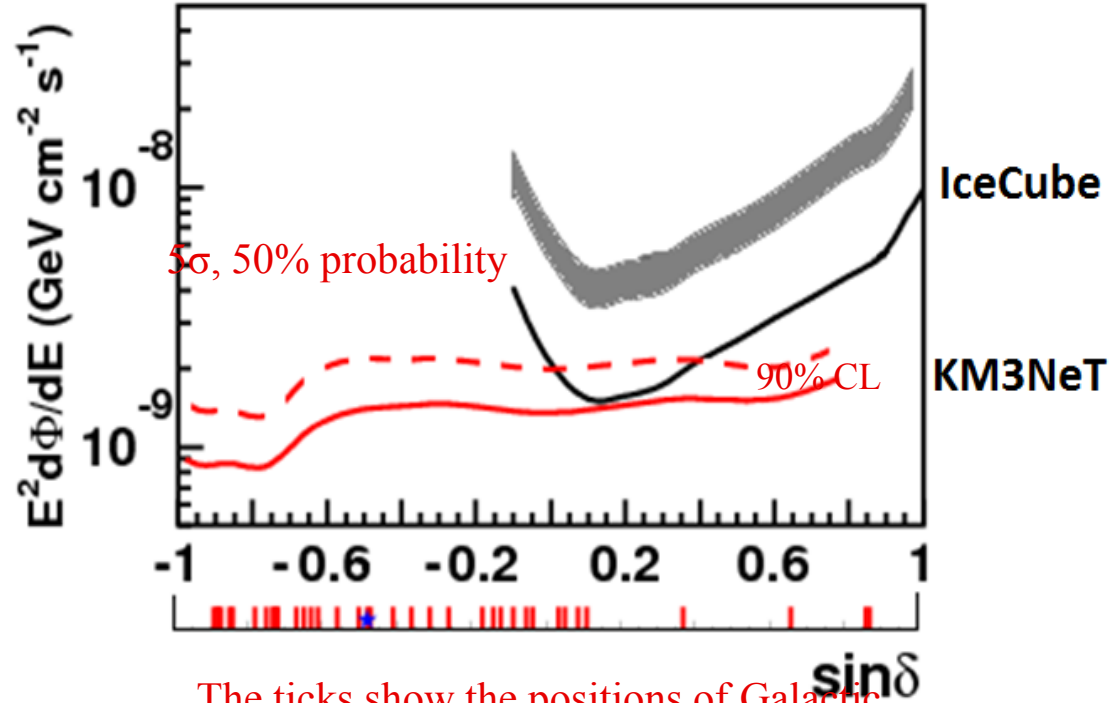
Flux sensitivity to neutrino point source
One year of observation

Building Block:

Currently considered option:
“Randomized Hexagon”



← ANTARES today: $\sim 10^{-7} \text{ GeV cm}^{-2} \text{ s}^{-1}$



The ticks show the positions of Galactic gamma ray sources. The position of the Galactic Centre is indicated by a blue star.

Surface area = $\pi R^2 = 4.2 \text{ km}^2$

$R = 1160 \text{ m}$

Instrumented volume = $\pi R^2 h$

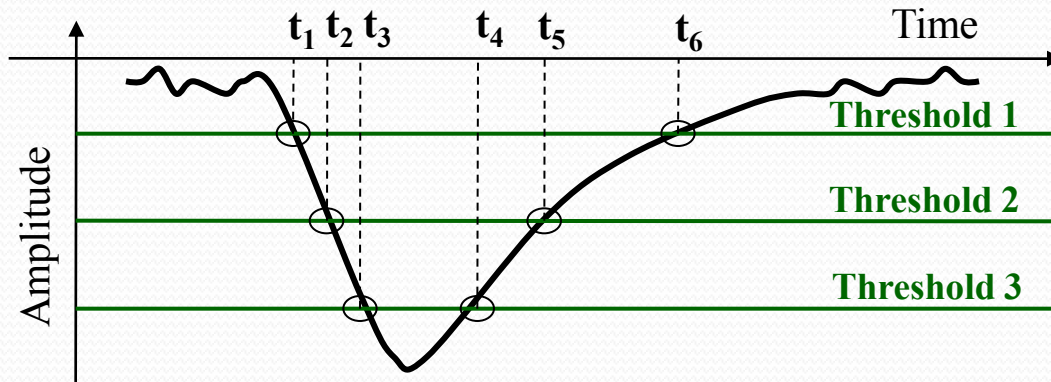
$h = 760 \text{ m (19x40)}$

$V_{\text{inst}} \approx 3 \text{ km}^3$

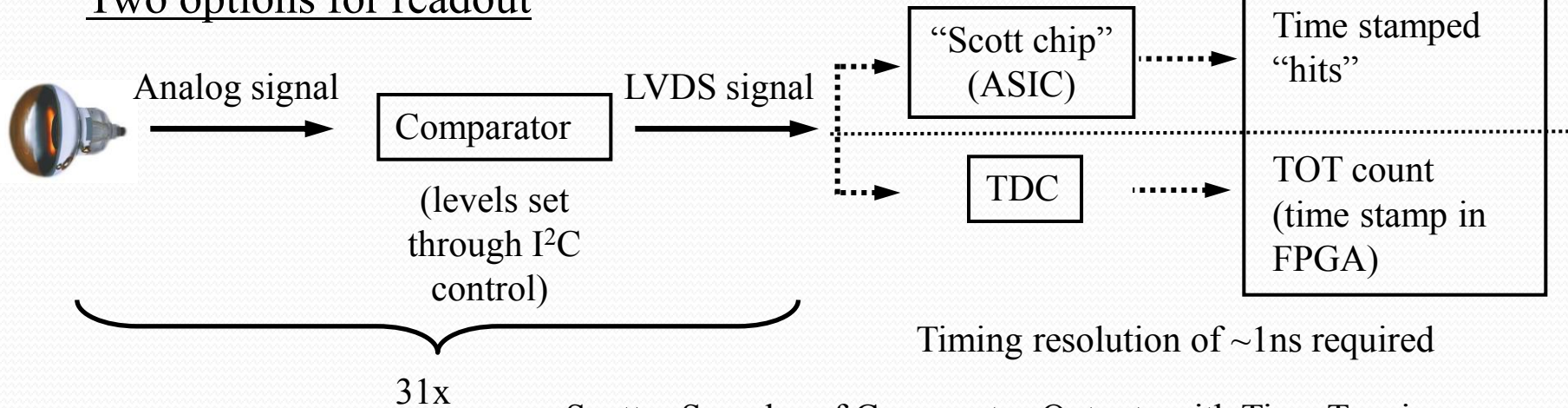
Building Blocks will make up
KM3NeT?

Front End Electronics

Time over threshold (TOT):



Two options for readout

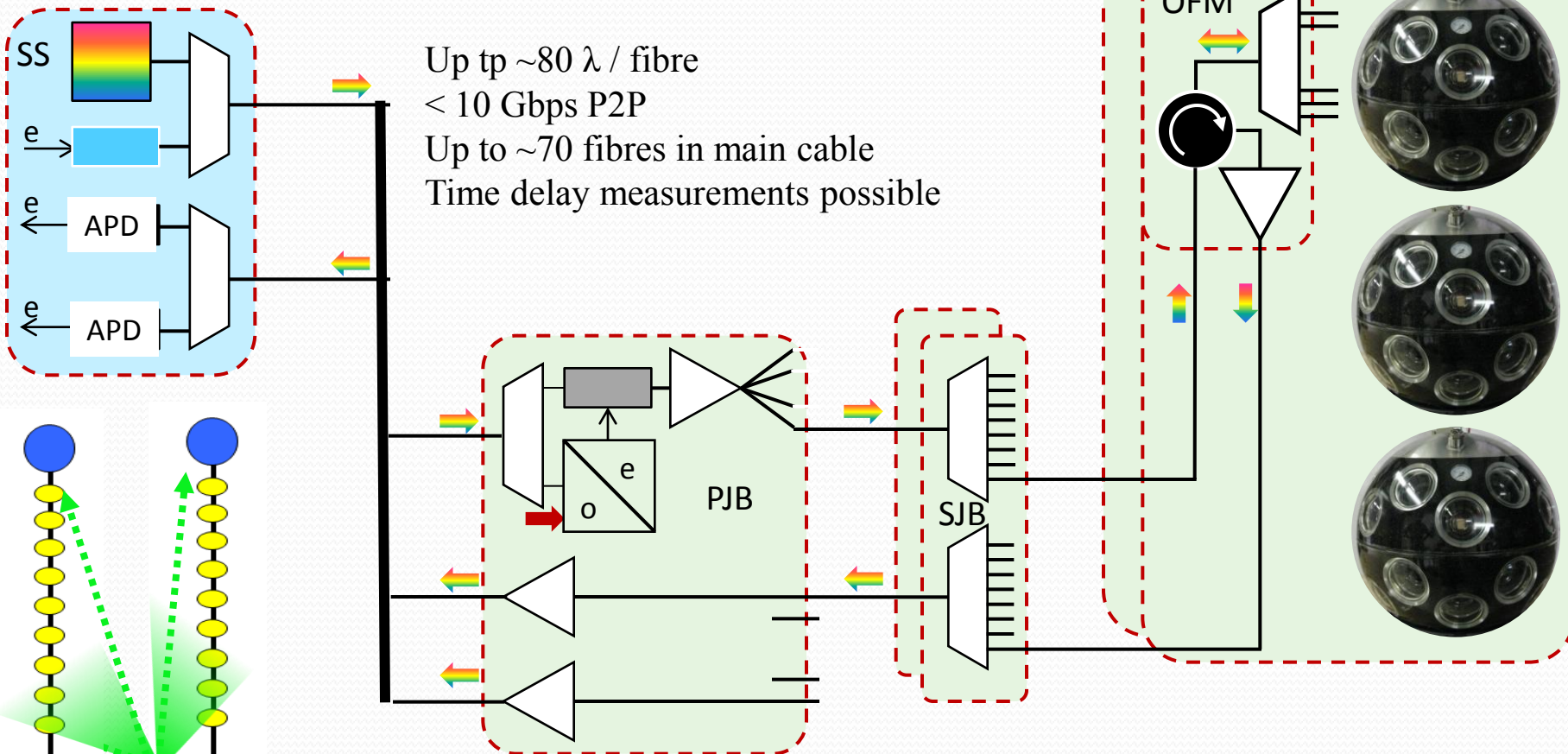


Scott = Sampler of Comparator Outputs with Time Tagging

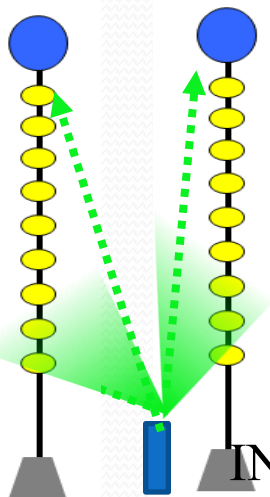
TDC = Time to Digital Converter

Optical Network

Dense Wavelength Division Multiplexing:
Following ITU Grid Specification:
C (1530-1570 nm) or L band (1570-1610 nm)
with 25GHz separation



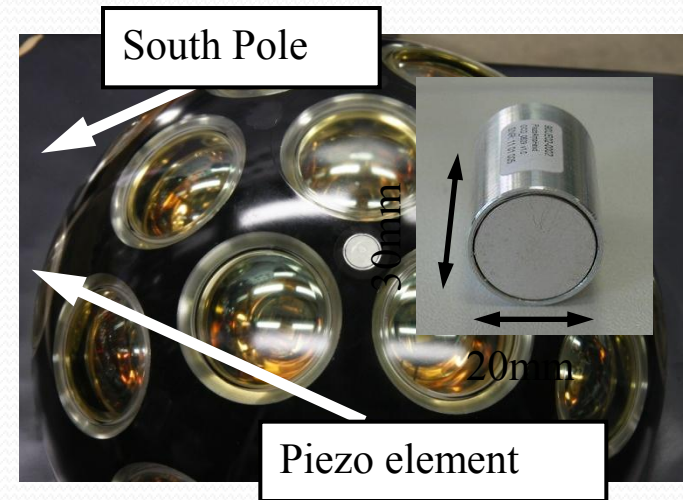
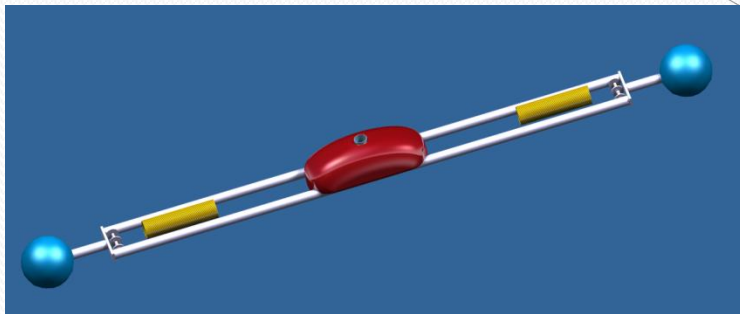
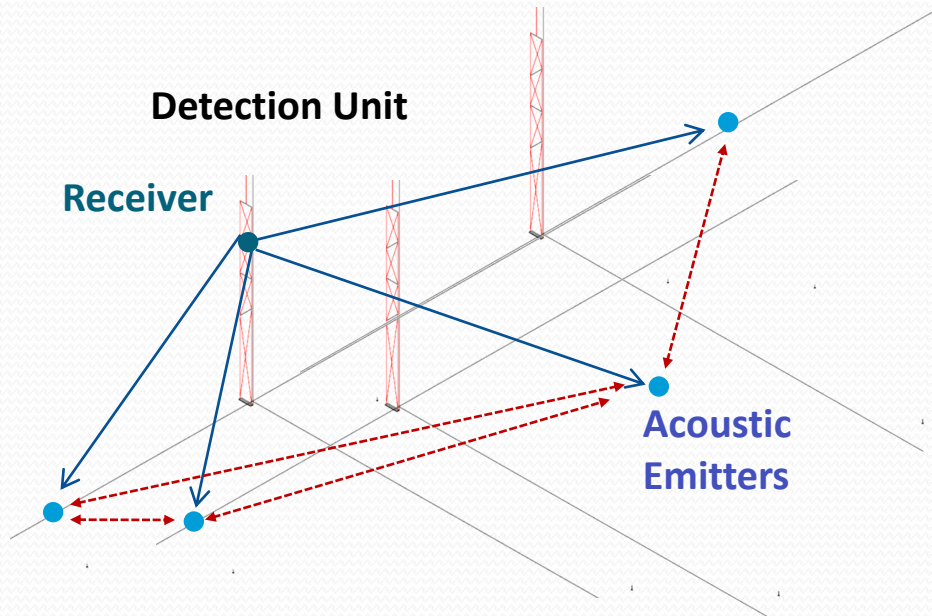
Up to $\sim 80 \lambda$ / fibre
< 10 Gbps P2P
Up to ~ 70 fibres in main cable
Time delay measurements possible



INTER DU Calibration Laser Beacons @ 532 nm

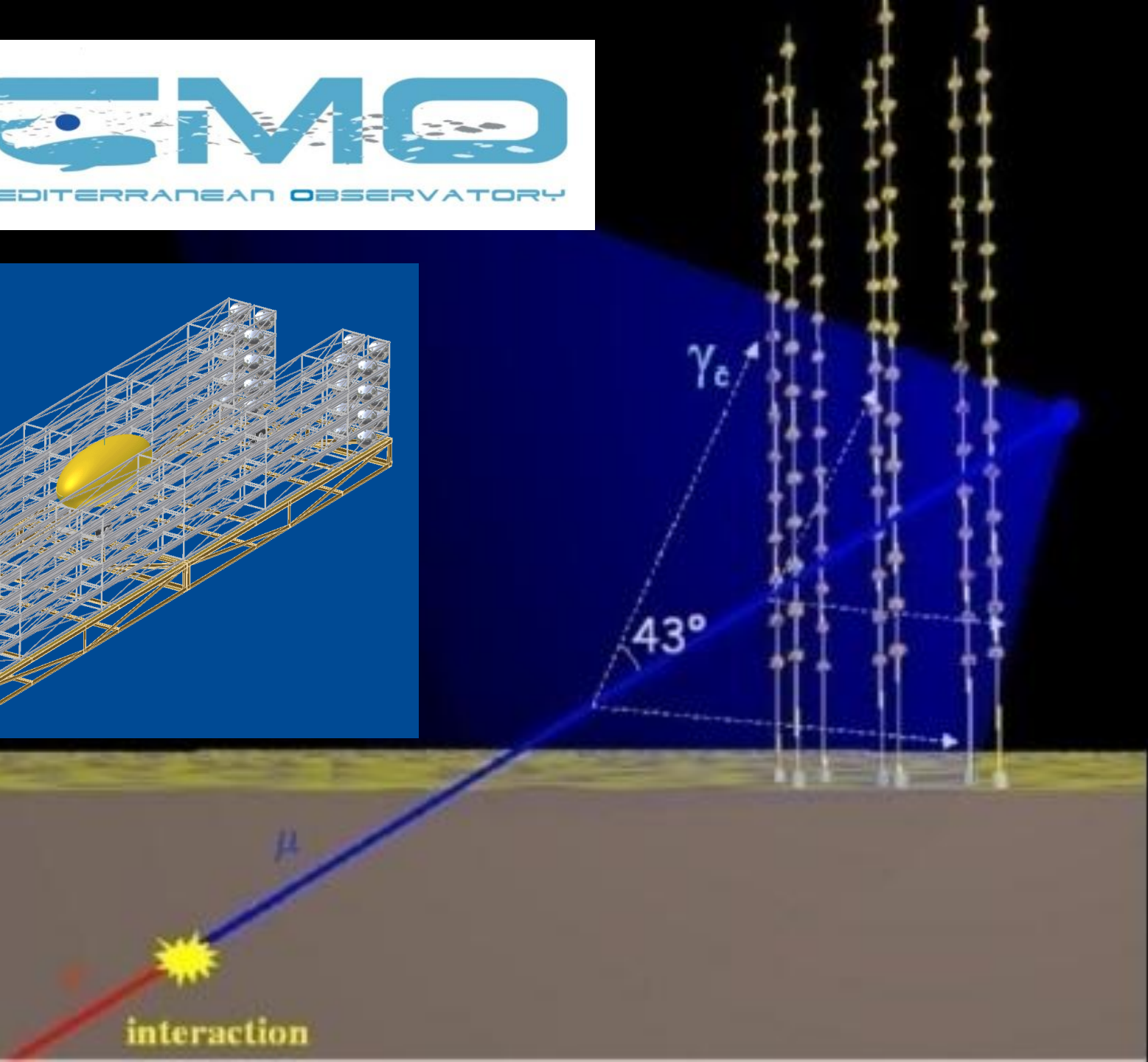
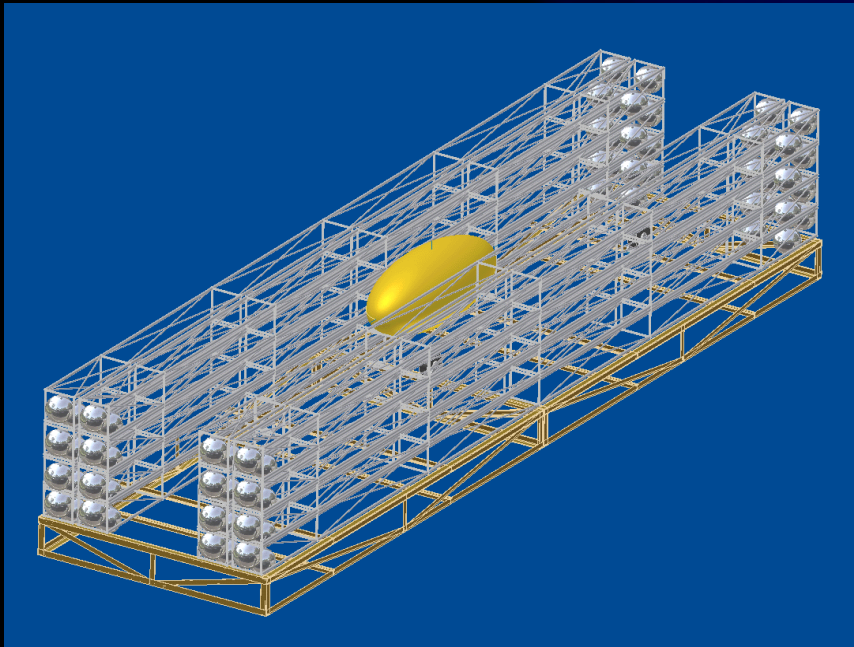
Acoustic Position Calibration

Options “piezo in sphere” and standard hydrophones investigated



NeMO

NEUTRINO MEDITERRANEAN OBSERVATORY



Pubblicazione del bando per il potenziamento delle infrastrutture (fondi PON)

Scadenza 21-8-2011 Inporto massimo 45 milioni di euro

In preparazione una proposta per la realizzazione di una infrastruttura di ricerca

che individua il sito di Portopalo (Capo Passero) come uno dei

networked KM3NeT nodes

[1 Meuro da Miur su progetti speciali]

Continua l'attività NEMO phase II: in costruzione, immersione fine 2011

Torre con 8 piani ($\ell=6\text{m}$, $d=40\text{ m}$), 4 OM (PMT Hamamatsu 10", sfera da 13") per piano

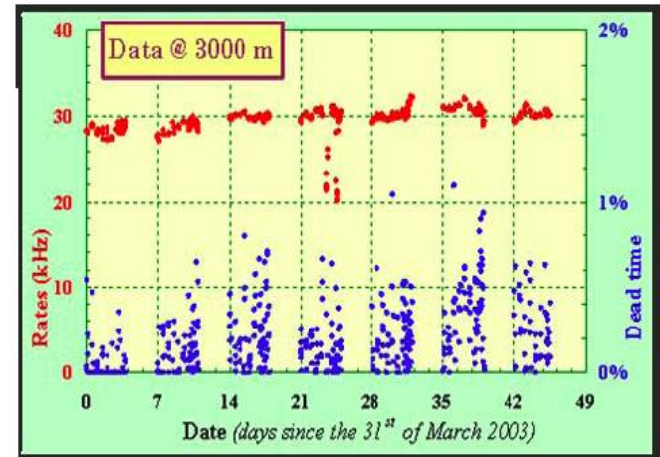
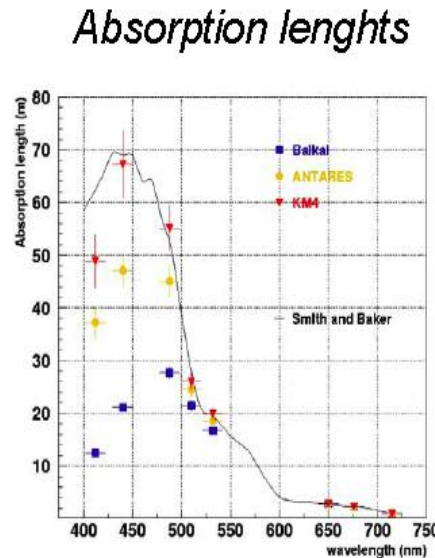
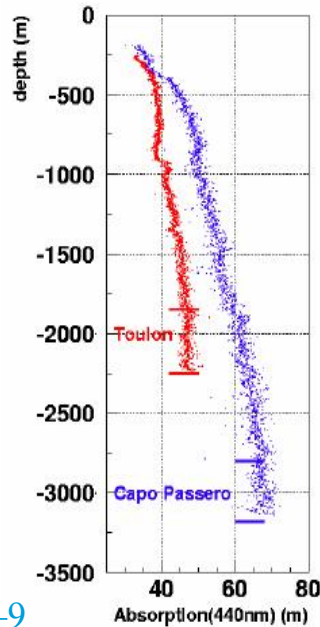
Uno o due piani per test (DOM etc...)

Completa di strumenti oceanografici e Idrofoni per la calibrazione spaziale

Extensive site exploration of Mediterranean Sea:

selected Capo Passero (depth 3500 m, shore distance 100 km)

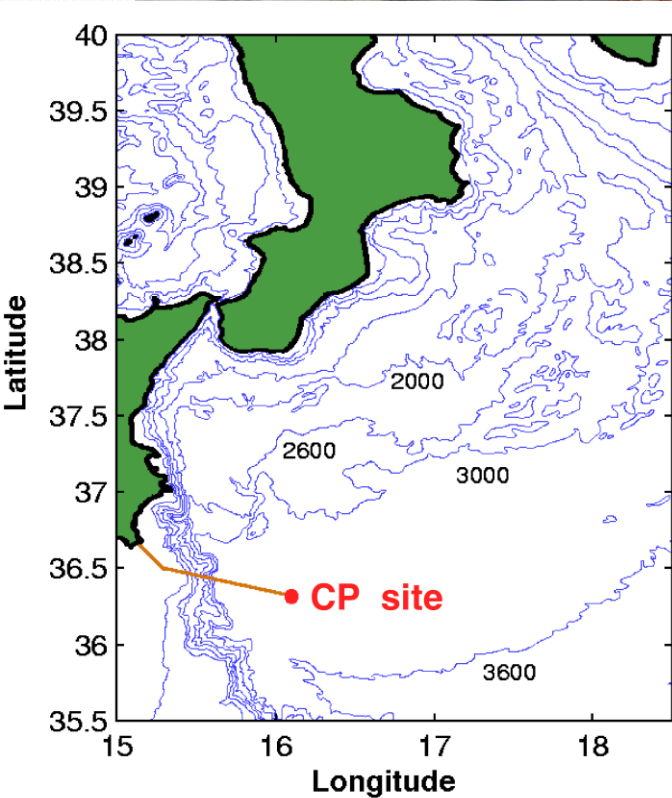
- best optical properties out of investigated sites
 - $L_a \sim 70\text{m} @ 440\text{nm}$
- no seasonal variations of water optical properties
- extremely low background from bioluminescence
 - $f \sim 30 \text{ kHz} (\varnothing_{\text{PMT}}=10'')$
- deep-sea water currents are low (3cm/s avg.) and stable
- wide abyssal plain, far from the shelf break, allows for possible reconfigurations of the detector layout



➤ PMT: 10''
➤ Thres: ~.5 SPE

Fraction of time
R > 200 kHz

Portopalo

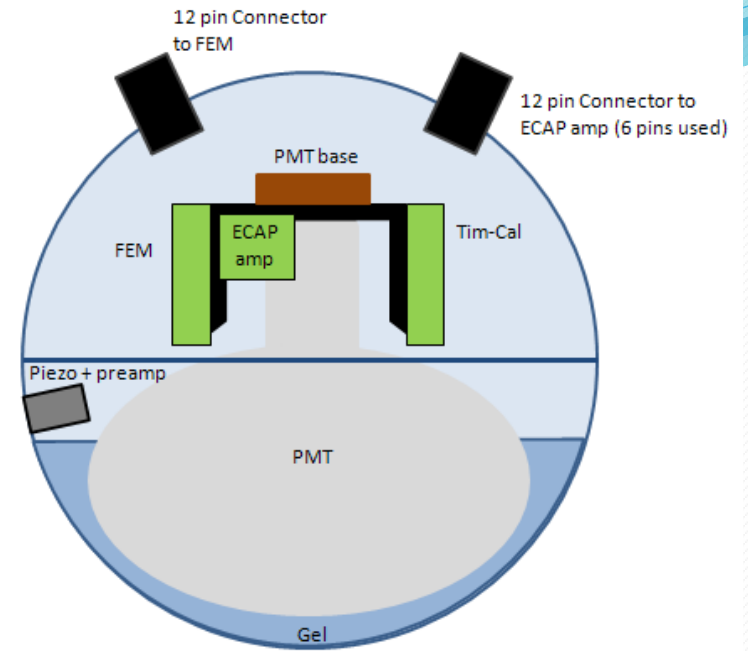
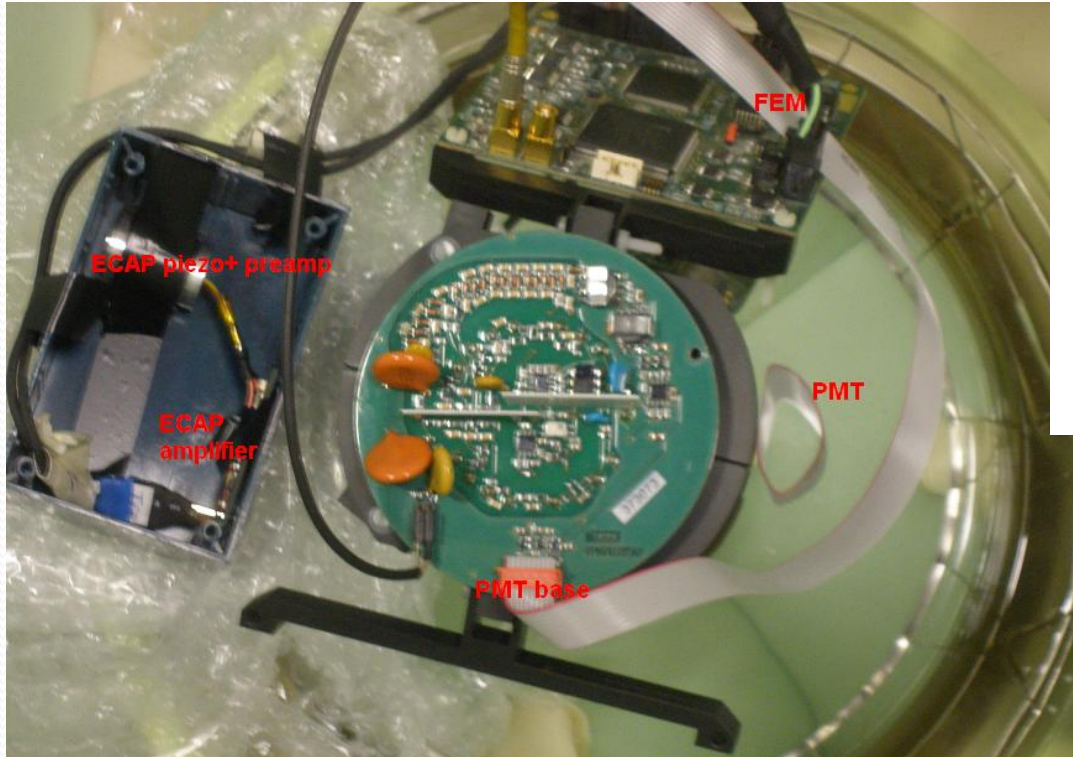


Cable and Power System installed end tested

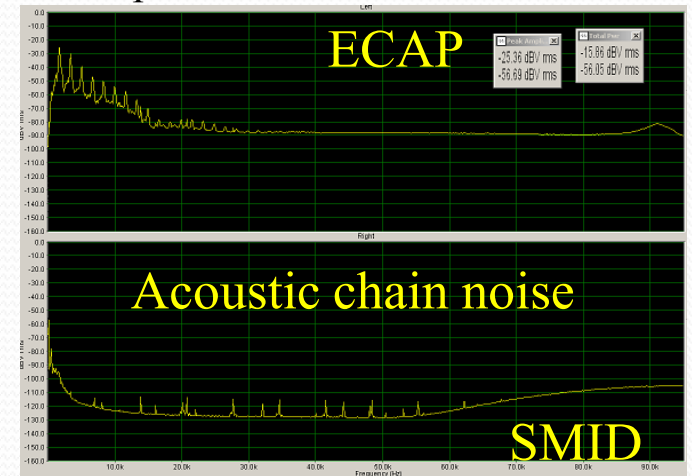
Ethernet: ok

Nemo phase II:
Settembre 2011 →

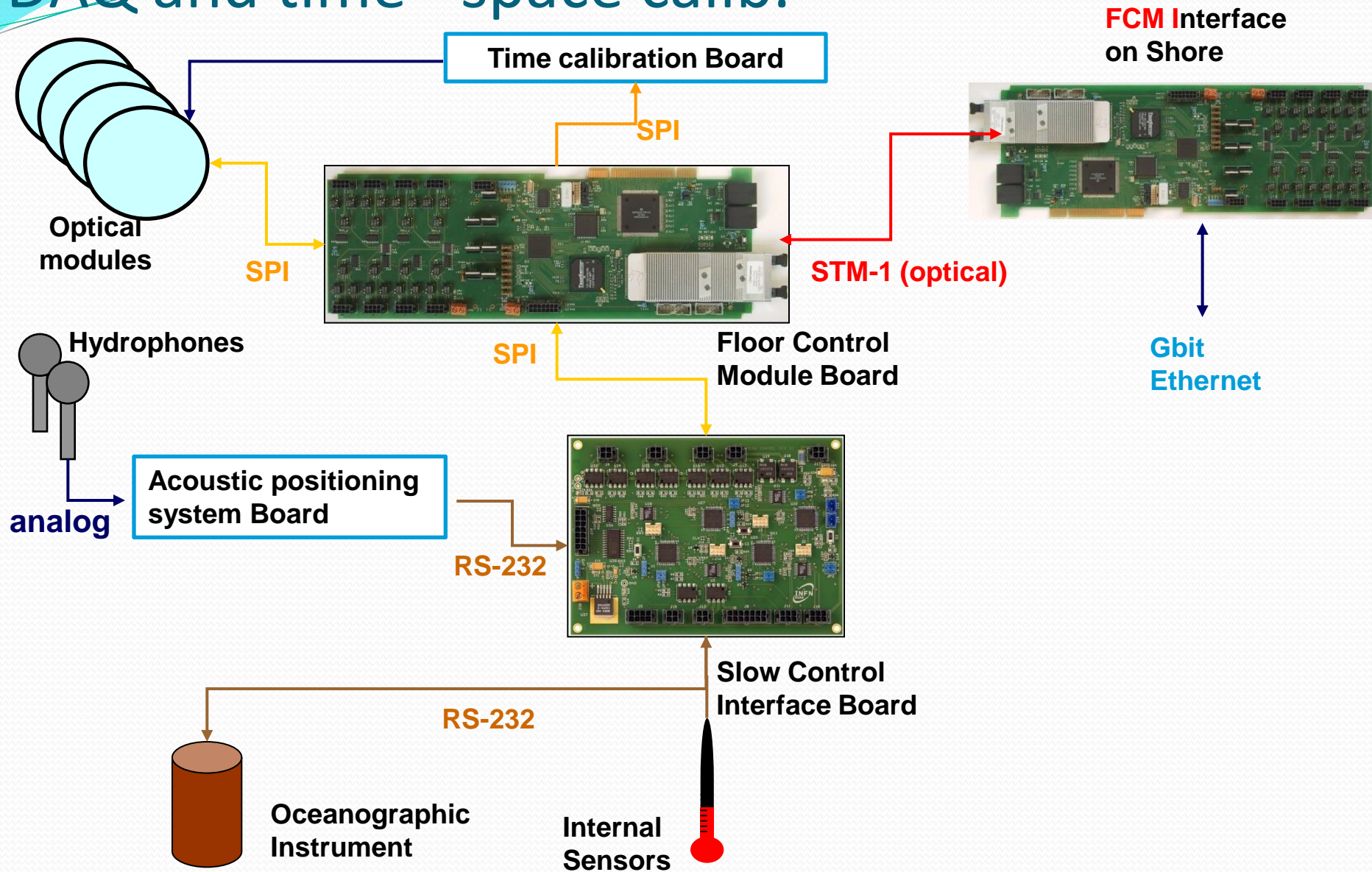
KM3 2012 → ??



Studio di nuove soluzioni per la calibrazione spaziale e temporale



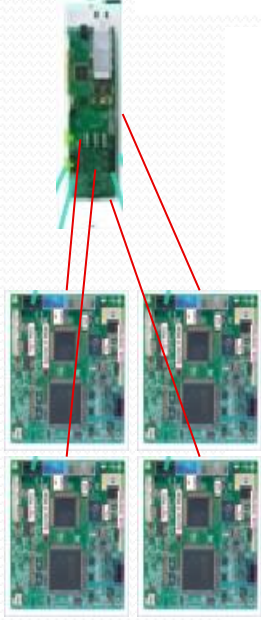
DAQ and time - space calib.



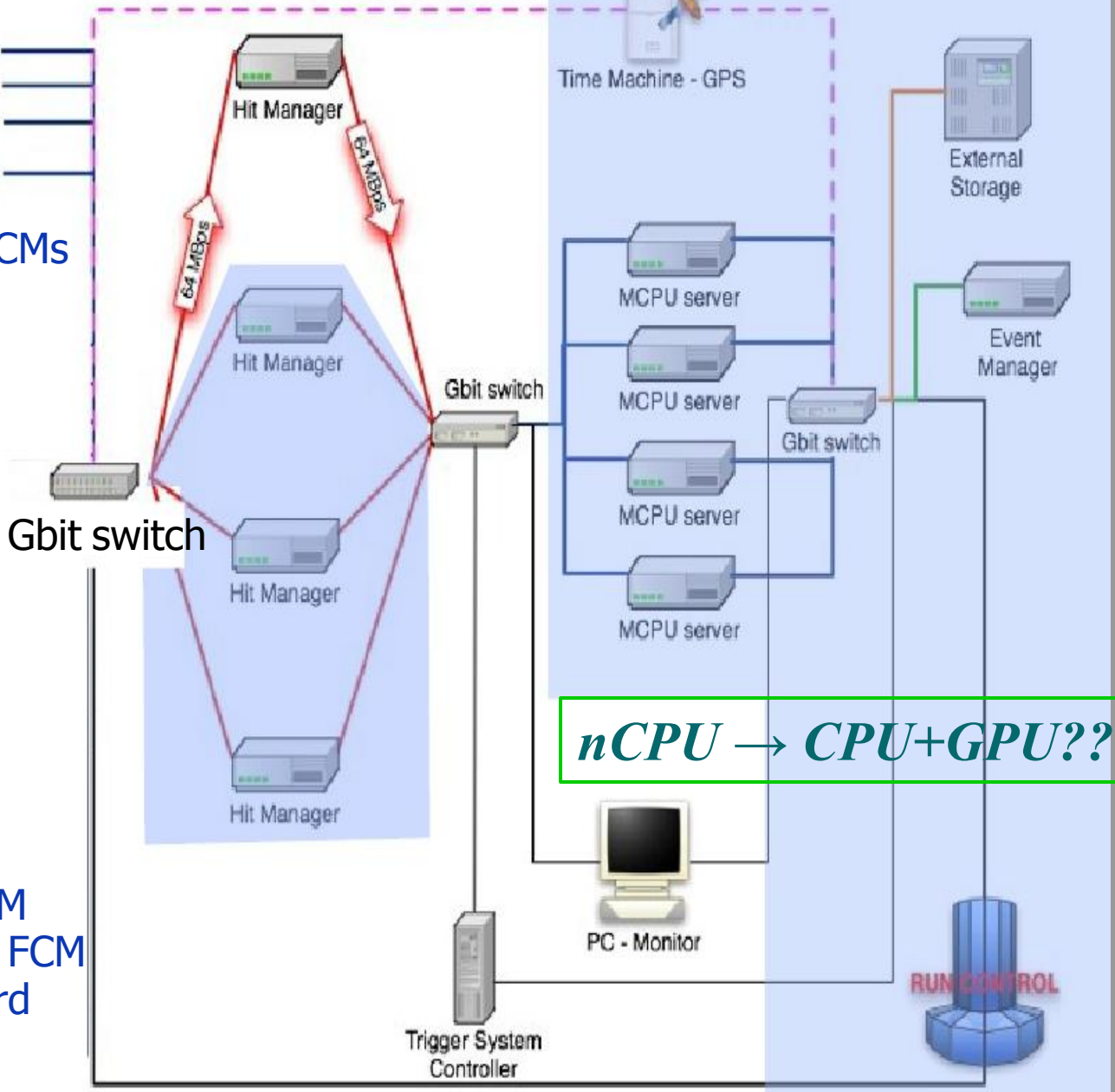
NETWORK OTTICO

9 Shore FCMs

9 Sea FCMs



The 4 FEM
all to 1 Sea FCM
+ Acu-Board

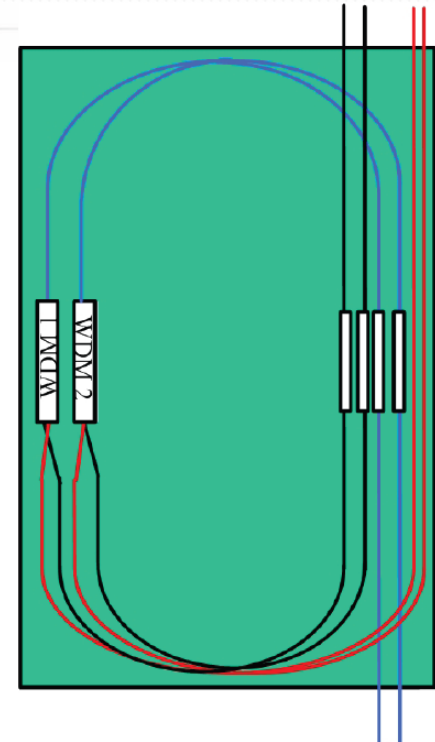
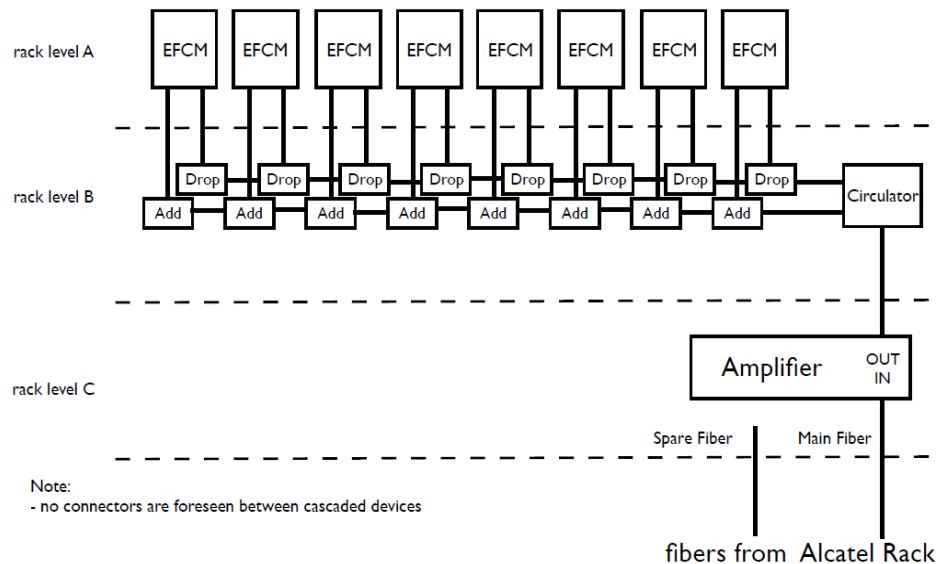


nCPU → *CPU+GPU??*

NEMO FASE-2: NETWORK OTTICO

- Costruito il modulo ottico che con filtri MUX/DEMUXD connette i Transceivers delle 9 schede eFCM (con lunghezze d'onda definite) al circolatore ottico DWDM
- Possediamo la tecnologia per effettuare le giunzioni ottiche e gli strumenti ottici per verificare la bontà delle connessioni e la loro attenuazione
- Gli ADD&Drop sono giuntati ad un connettore standard ottico SC. Si è verificato che siano conformi alle specifiche.
- Un pannello frontale permette la connessione con bretelle fra amplificatore Raman (da e verso il mare) e schede eFCM

N2 Shore Station - Optical Network Rack



Trigger on GPU

Montecarlo (muoni da neutrini + 50 KHz di background):

4 TTS da 192 ms di dati (torre di 16 piani)

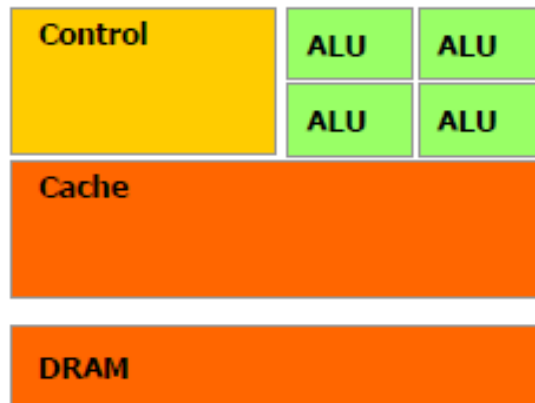
Ciascuna TTs è suddivisa in miniTTs da 12ms

In ogni miniTTs gli hits) sono ordinati secondo il numero di PMT e per tempo.

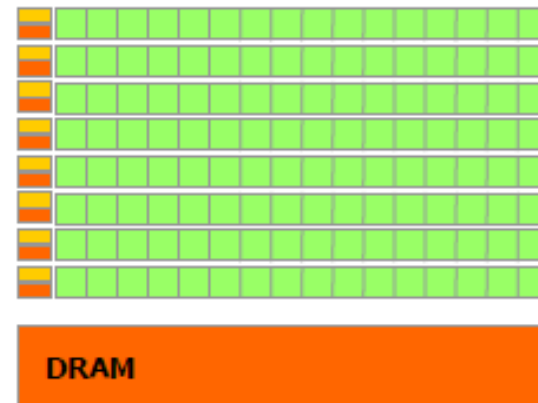
Le tracce (≥ 3 hit, hit spread $< 3.5 \mu\text{s}$) da muoni sono distribuite nelle TTs e hanno fra loro la stessa distanza temporale ($100 \mu\text{s} \gg 3.5 \mu\text{s}$)

Nella GPU sfruttando il calcolo parallelo:

- gli hit vengono ordinati temporalmente mescolando i 64 PMT*
- si studia un algoritmo di trigger che individui la traccia nel background.*



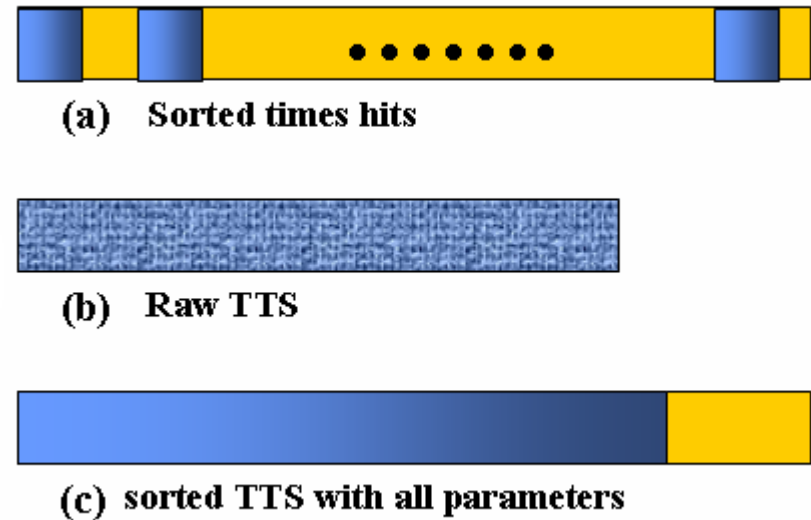
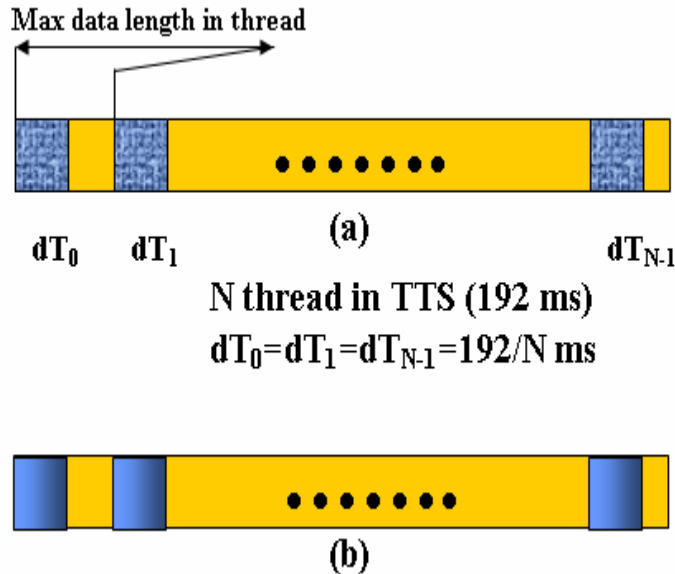
CPU



GPU

I processi nella GPU sono organizzati in blocchi di threads

TTS Reading and Sorting



Send times and relative hit positions to GPU

- divide the hits in time intervals (*number of intervals = number of threads per block*)
- sort data using quicksort algorithm.

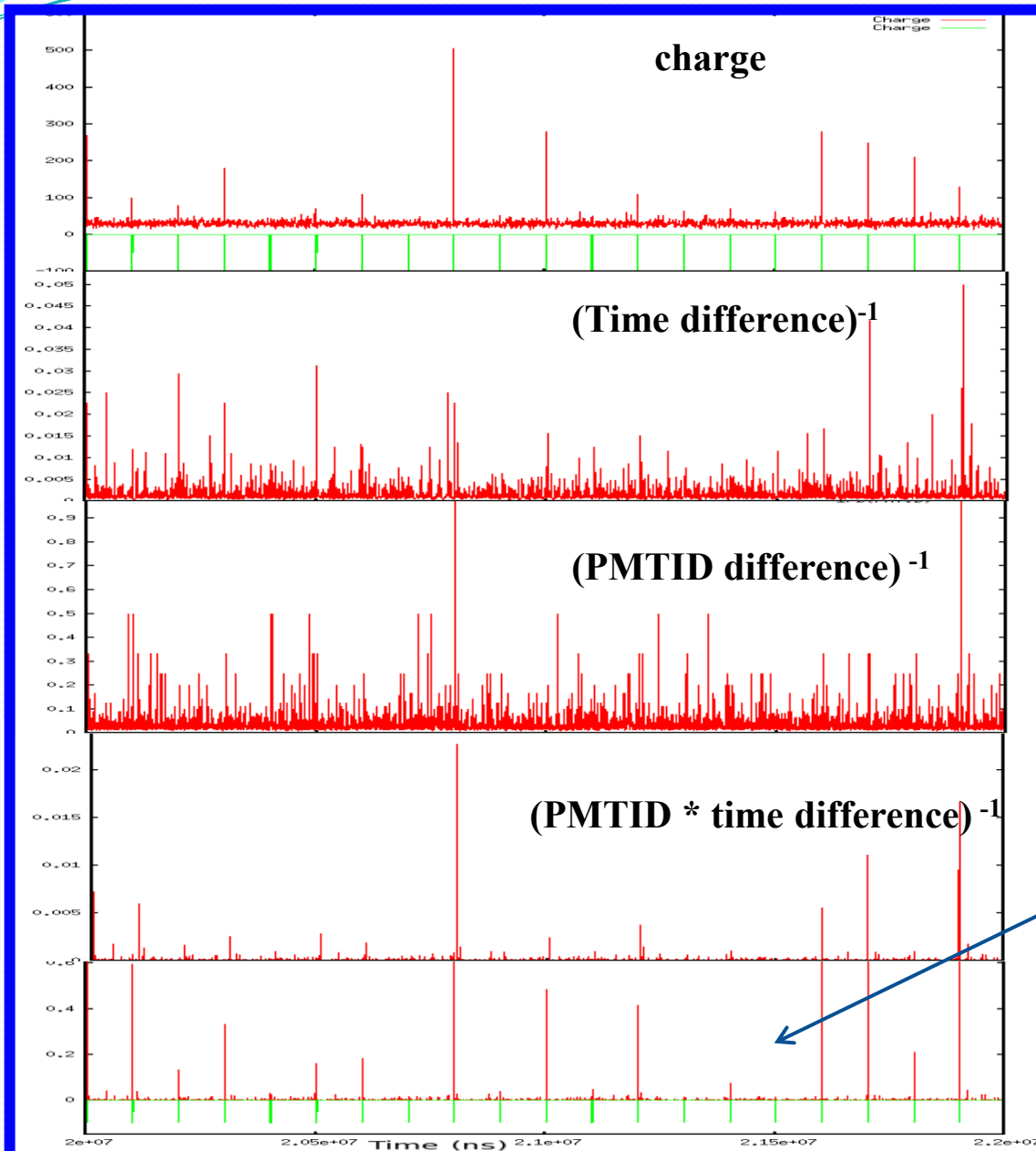
→ CPU group all data → the TTS is ready for trigger

Trigger algorithms in GPU

After the sorting step, we have know an equivalent TTS, which we call it TTS_s . the paramters to be worked on by triggers are: time hits $t_{h[i]}$, its charges $Q_{h[i]}$, and PMT_{id} the work of our first trigger is as follow:

- each thread will work on portion of data, where three tasks to do:
- each hit $h[i]$, and the two next, we calculate the sum of its charges $Q_m = Q_i + Q_{i+1} + Q_{i+2}$.
- for each hit $h[i]$, we calculate the time difference δT between the current hit $t_{h[i]}$, and $t_{h[i+2]}$.
- for each hit $h[i]$, we calculate sum of pmt_{id} differences $\equiv id_m$ e.g. $|id_{h[i]} - id_{h[i+1]}|$, and $|id_{h[i]} - id_{h[i+2]}|$.
- the minimum values of time difference and sum of the pmt_{id} differences would be considered as a trigger for an event, hence $\frac{1}{\delta T \times id_m}$ can be used as trigger.
- to include the charge effect, we will use $\frac{Q_m}{\delta T \times id_m}$.

Trigger algorithms in GPU



ΔT for 4 TTS = **768ms**

Ordering and trigger search
250ms

Algoritmi semplici:
5 somme, 1 prodotto, 1 una divisione

$$\frac{\text{Charge}}{(\text{PMTID} * \text{time difference})}$$

Software per la gestione della documentazione

- **Software Web-Server** (su INFN-Pisa)
- **Compilatore di documenti di test** tramite interfaccia grafica guidata (lavora su template definiti e approvati dalla collaborazione)
- **DB delle componenti di progetto**, strutturato ad albero (componente/sottocomponente), con la possibilità di allegare alla singola unità tutta la documentazione del proprio ciclo vita
- L' inserimento della documentazione/nuova componente/modifiche/test/upgrade..etc comporta l'invio automatico alla mailing list definita
- Differenti profili d'accesso (Admin, User..)
- Differenti modalità di interrogazione
- Manuale Utente

=> necessario per il QA e RA

Attività 2012

- KM3:
 - progettazione meccanica per KM3
 - partecipazione costruzione e test del prototipo della DU
 - QA - RA

- Nemo Phase II
 - fine della fase d'integrazione della torre ad 8 piani
 - sistema di acquisizione e trigger/on-line
 - continua lo sviluppo di un trigger con GPU
 - acquisizione ed analisi dati
 - QA - RA

Percentuali 2012

Andrea	Betti	contratto Univ
Nicolò	Beverini	PA
Bachir	Bouhadeb	Assegno Univ
Mauro	Morganti	PS Accademia
Andrea	Verlicchi	Ric. Accademia
Calogero	Sollima	Art. 2222 in scadenza
Enrico	Maccioni	Tecnico Univ.
Fabio	Stefani	Assegno Univ
Francesco	Francesconi	Tecnico Univ.
Stefano	Galeotti	INFN
Fabrizio	Raffaelli	INFN
Giuseppe	Terreni	INFN

Antonio Marinelli non è più associabile

NEMO + Antares : Percentuali da decidere

NEMO Pisa: preventivo 2012

	MI	ME	Cons	App.	Inv	T.	Tot.
Riunioni di collaborazione: 3CM*2persone* 1,1 k€ Tec. Board: 6TB*1persone*0,7 k€	6,6						
	4,2						
integrazione torre, test, presa dati: 8missioni*1persone*1settimana*1,7k€	13,2						
Meeting di lavoro: 8 missioni * 1 persona * 0,5 k€	4,0						
Missioni per contatti con istituti stranieri (convergenza)		8,0					
Consumo per costruzione e integrazione (ripartizione spese)			4,0				
<div style="border: 2px solid blue; padding: 10px; margin: 10px auto; width: 80%;"> <p><i>I preventivi verranno discussi all'inizio di luglio. Sono solo indicate le previsioni di spesa per: trasferte, consumo e trasporti</i></p> </div>							
Trasporti e Assicurazioni						2,0	
Totale	28,0	8,0	4,0	0,0	0,0	2,0	42,0

- **Progettazione mecc.** ~ 3mu
Contributo alla progettazione della DU di KM3
- **Officina mecc.** ~ 4 mu
Partecipazione a fine integrazione e test
- **Elettronica** > 0
E' sufficiente l'Ing. Stefani
- **Calcolo**
G.Terreni che svolto un lavoro fondamentale seguendo tutte le fasi del progetto del DAQ a terra: da eFCM a Data Storage