### **PMTs:**

### **Ongoing Analysis & Future**

CYGNO Collaboration Meeting

Frascati, 4-6 December, 2023

### **David Marques**

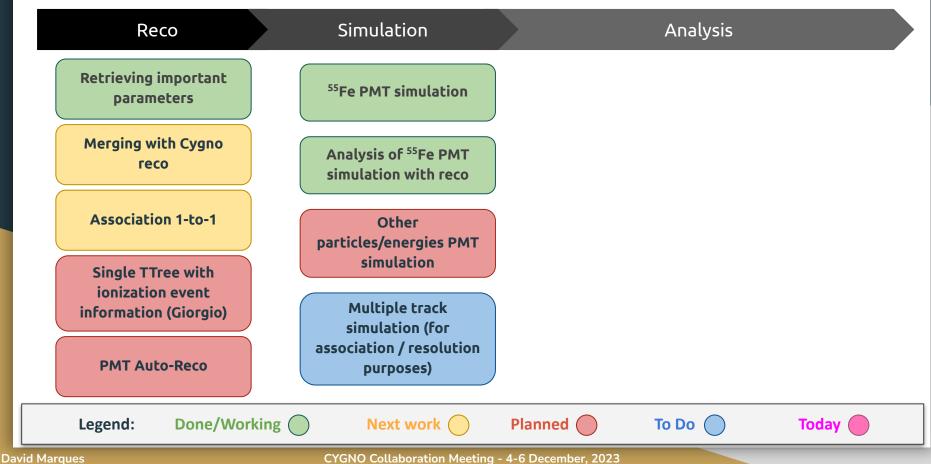
& PMT analysis working group

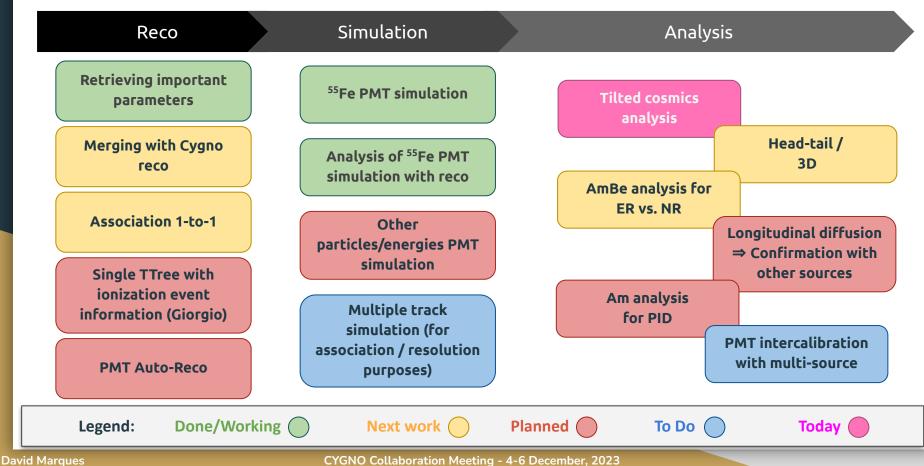
### Contents

- 1. Overview of PMT works
- 2. Time over Threshold explanation
- **3.** Tilted cosmics analysis
  - a. Motivation
  - b. Setup
  - c. Results and discussion
    - i. Follow up?
- 4. Next analysis
  - a. LIME ⇒ Am(Be); PID; 3D
  - b. MANGO  $\Rightarrow$  NID longitudinal diffusion





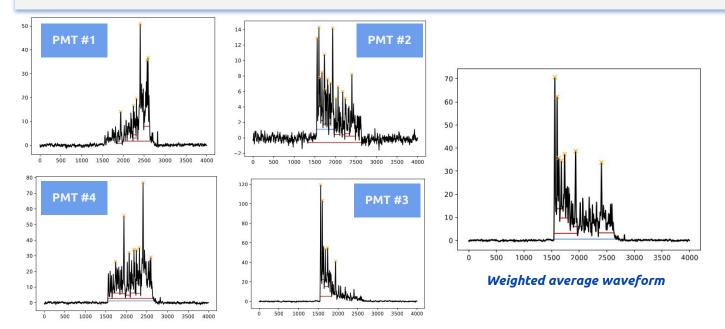




### **Time over Threshold**

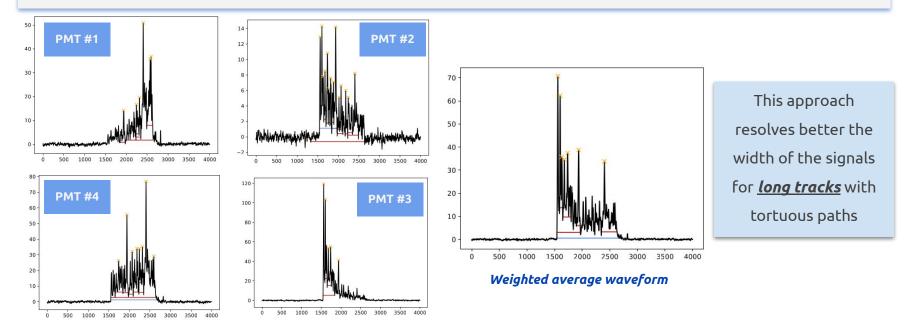
## **Time over Threshold**

- Measurement of the **<u>time length</u>** of the signal which is **<u>above a given threshold</u>**.
  - Not trivial when each PMT sees a different signal intensity and tracks can have very complicated paths
  - I do a **weighted average based on waveform's SNR** ⇒ Only correct for <u>timing</u> purposes



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#### Motivation & Math:

#### Possible results / analysis:

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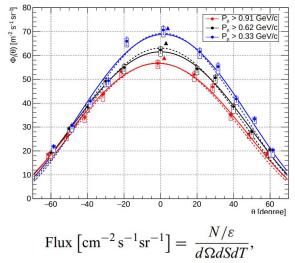
- This measurement presents a *clear dataset* with tracks with *well-defined orientation* and energy deposit (MIP)
  - We have a specific <u>range of possible angles</u> of entering LIME (given by geometry of LIME + scintillators)
  - PMT measures the <u>Time over Threshold</u>
    - Multiplied by v<sub>drift e-</sub> gives the ∆z
  - Height of LIME (c1) is known (33 cm)
    - The tracks inclination (α) will be tan<sup>-1</sup> (Δz/c1)

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- We can <u>compare</u> it with the geometrical accepted angles.
- We can **<u>calculate the flux</u>** and compare it with the **cosmic**

**muons angle distribution** at ground (  $\propto \cos^2(\Theta)^*$  )





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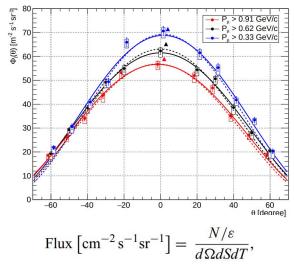
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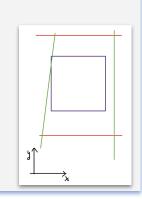


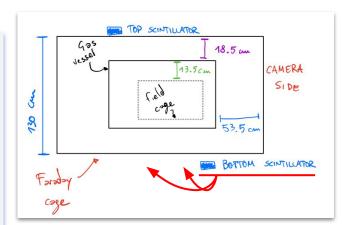
Gives us a measurement of PMT Reco / ToT efficiency First CYGNO 3D analysis (on a distribution basis)

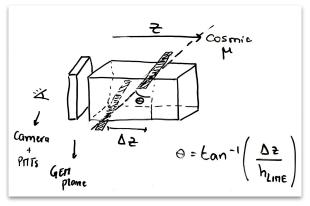
#### Setup:

Dimensions:

Accuracy:







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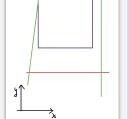
- → Two scintillator bars were placed on top and bottom of LIME
- → LIME DAQ triggered by coincidence of two scintillators
  - 3 different scintillator position were used
  - By geometry, only certains angles are possible

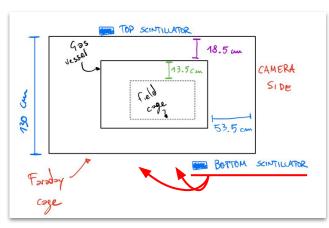
#### **Dimensions:**

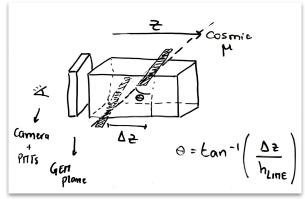
- → LIME:
  - Z: 50 cm ; Y: 33 cm ; X: 33 cm
- → Scintillator:
  - Z: ~5 cm ; Y: ~2 cm ; X: > 33 cm

#### Accuracy:

- This configuration actually allows for cosmics to enter and trigger **from the side** of LIME ⇒ Creates **long tails**
- **Random coincidences** from radioactivity or secondary particles are also possible







#### Analysis method:

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- Retrieve *Time over threshold* with weighted average
- Multiply with e-velocity in our gas to get *travelled Z* 
  - $\circ$  v<sub>e</sub> = 5.471 cm/us @ 800 V/cm, from Giorgio's thesis.
- Calculate angle: *theta = tan<sup>-1</sup> ((ToT\*1000) \* v<sub>e</sub> / 33cm )*

#### Analysis method:

- Retrieve *Time over threshold* with weighted average
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- Calculate angle: *theta = tan<sup>-1</sup> ( (ToT\*1000) \* v<sub>e-</sub> / 33cm )*
- Compare obtained angle with geometrical accepted one
- Calculate *muon flux* using <u>interaction rate</u> and compare with a cos<sup>2</sup>(theta) distribution

Time over threshold distribution

12000

10000

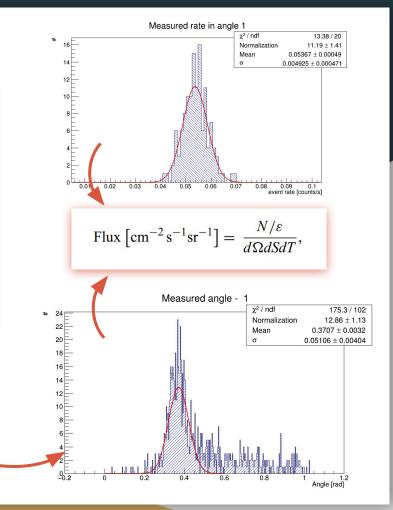
14000 16000

Time over threshold [ns]

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40 E



15

2000 4000

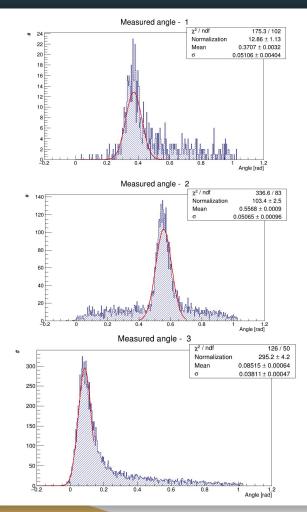
6000

8000

Z [cm]

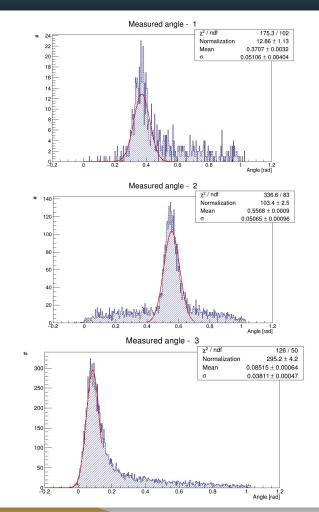
Travelled Z distribution

#### **Results - Angles comparison:**



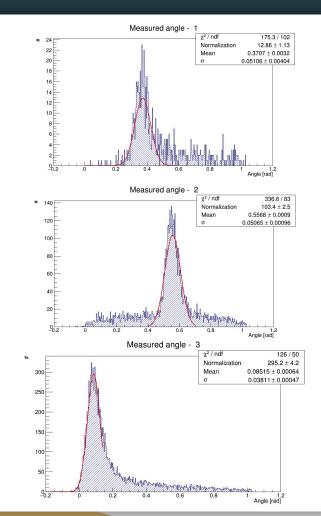
### **Results - Angles comparison:**

- Measured angles:
  - 21.24 deg
  - 31.90 deg
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- Measured angles:
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- Geometry allowed angles (hand calculated):
  - o 20.27 24.04 deg
  - 28.64 31.93 deg
  - 0 deg 4.4 deg (Parallel case is special)
- Good agreement at first order
- Long tails visible as expected



**Results - Flux measurement:** 

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• Event rate retrieved from logbook

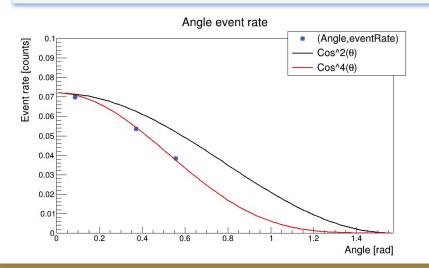
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Dead-times don't need to be considered

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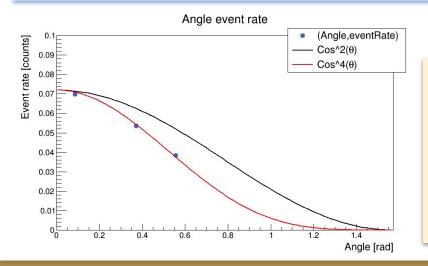
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#### <u>Results - Flux measurement:</u>

- Event rate retrieved from logbook
  - Only one trigger per event, easy calculation  $\Rightarrow$

Dead-times don't need to be considered



	Data clearly fits a cos <sup>4</sup> instead of a cos <sup>2</sup>						
	$\downarrow$						
what are we missing?							
	Let's revisit the theory						

Flux 
$$\left[ \operatorname{cm}^{-2} \operatorname{s}^{-1} \operatorname{sr}^{-1} \right] = \frac{N/\varepsilon}{d\Omega dS dT},$$

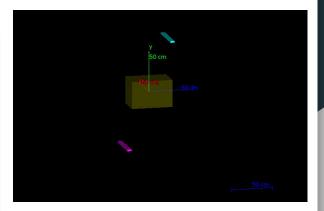
$$d\Omega = \sin heta\, d heta\, darphi, \qquad \Omega = \iint_S rac{\hat r \cdot \hat n}{r^2}\, dS \ = \iint_S \sin heta\, d heta\, darphi,$$

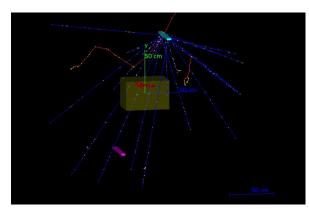
The <u>muon flux</u> takes into account the <u>acceptance / geometry</u> <u>factor</u> of the detectors ↓ Each of our configurations is different ⇒ To properly retrieve the cos<sup>2</sup> dependency of the flux, one needs to do this calculation.

Simulation - Flux measurement:

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- Basic **<u>GEANT4 simulation</u>** was created (with some help from Samuele)
- Particle (2 GeV muon) shot randomly from top scintillator with flat direction distribution





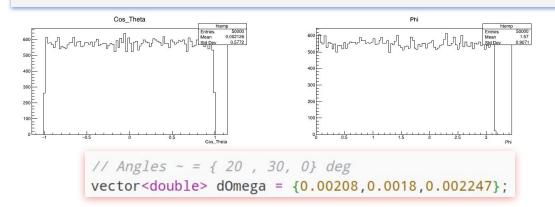


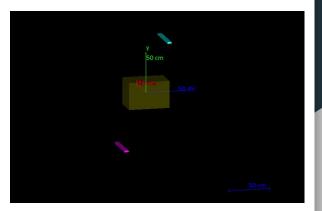
#### CYGNO Collaboration Meeting - 4-6 December, 2023

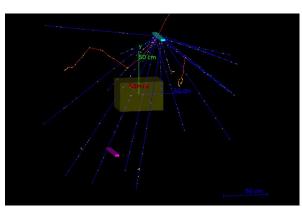
50000 1.57 0,9071

### Simulation - Flux measurement:

- Basic **<u>GEANT4 simulation</u>** was created (with some help from Samuele)
- Particle (2 GeV muon) shot randomly from top scintillator with flat direction distribution
- **<u>Geometrical acceptance</u>** calculated from ratio between triple coincidence and total shot particles
  - **<u>Recalculated muon flux</u>**





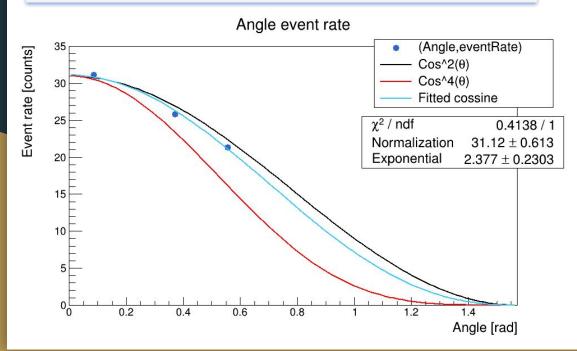


**Results - Flux measurement:** 

• Final result:

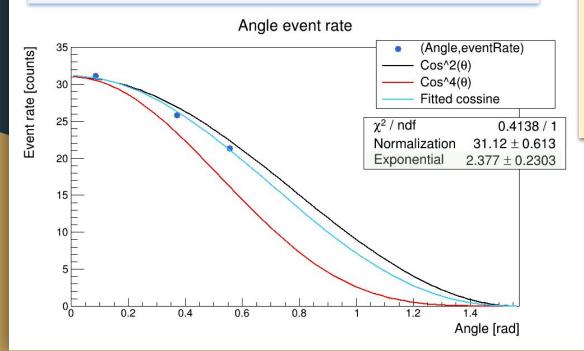
### **Results - Flux measurement:**

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• Final result:



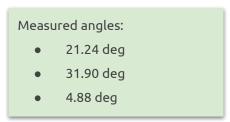
Doesn't fit a cos<sup>4</sup> anymore ↓ More consistent with a cos<sup>2</sup> distribution as expected from literature (missing error bars) ↓ Still not the real muon flux since I miss the area (dA) normalization ⇒ With this we can compare with real data

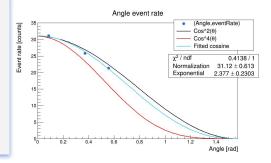
Authors	$P_c$ (GV)	Alt. (m)	$P_{\mu}$ (GeV/c)	n value	$\begin{array}{c} \Phi_I(0^\circ) \\ (m^2 \ sr \ s)^{-1} \end{array}$
Pethuraj et al. [34]	17.6	160	$\geq 0.11$	$2.00{\pm}0.16$	$70.07 \pm 5.26$
Sogarwal et al. [35]	16.38	SL	$\geq 0.25$	$2.10{\pm}0.25$	$66.70 \pm 1.54$
S. Pal et al. [36]	16	SL	$\geq 0.28$	$2.15{\pm}0.01$	$62.17 \pm 0.05$
Bhattacharyya [37]	14	24	$\geq 0.4$	$1.91 \pm 0.1$	-
			≥1.	$1.85{\pm}0.11$	-
Arneodo et al. [38]	14	SL	$\geq 0.04$	$1.91{\pm}0.18$	$75.4 \pm 1.4$
Present data	9.6	38	<u>≥0.33</u>	$1.82 {\pm} 0.11$	68.77±1.94
			$\geq 0.62$	$1.72{\pm}0.10$	61.49±1.44
			$\geq 0.91$	$1.72{\pm}0.10$	$56.66 \pm 1.60$
CRY [25]	9.6	SL	≥0.33	2.02	69.26
			$\geq 0.62$	1.95	63.02
			$\geq 0.91$	1.87	56.80
Riggi et al. [33]	8	3100	$\geq 0.2$	$1.83{\pm}0.13$	83.±8
Judge and Nash [39]	2.5	SL	$\geq 0.7$	$1.96 \pm 0.22$	

### **Conclusions:**

Geometry allowed angles (hand calculated):

- 20.27 24.04 deg
- 28.64 31.93 deg
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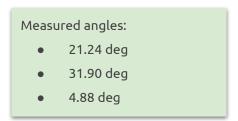


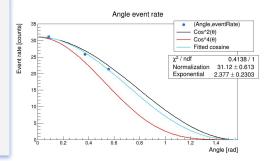
### **Conclusions:**

- Results are at first order satisfactory.
  - **ToT is a reliable variable** and is already in the CYGNO **reconstruction**
- Interesting study on the CYGNO PMT analysis ⇒ Could eventually be used in a paper
- Some caveats:
  - A small correction on the ToT was applied since theta = 0 does \*not\* produce a ToT = 0 (Perhaps a **longitudinal**  $\sigma_0$  for MIPs is necessary?)
  - There were some uncertainties regarding the dimensions of the scintillators
    - I will redo the simulations and *add more statistics*
- \*If\* you think this would be relevant, the same study could be performed with GIN, adding more angles, and measuring longitudinal diffusion with MIP.

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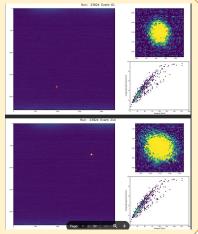




**Finalization** of Tilted Cosmics study

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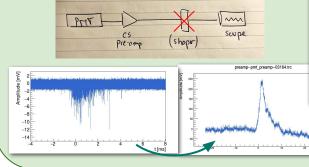
AmBe dataset analysis ↓ Attempts on PID and dE/dx analysis ↓ Already have the "most likely NR" data selected from Matteo



### **Finalization** of Tilted Cosmics study

Negative Ion Drift ⇒ ⇒ <u>Longitudinal diffusion</u>

Stronger and <u>independent</u> confirmation of **below thermal** behaviour of NID

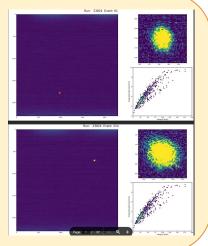




AmBe dataset analysis ↓ Attempts on PID and dE/dx analysis

 $\downarrow$ 

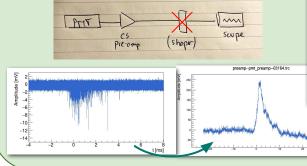
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Negative Ion Drift ⇒ ⇒ <u>Longitudinal diffusion</u> ↓ Stronger and <u>independent</u> confirmation

of *below thermal* behaviour of NID

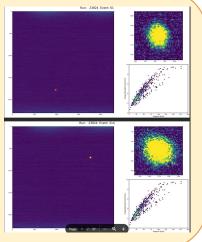


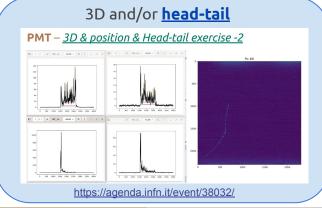


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### Thank you for

### your attention!

The CYGNO Project counts with the collaboration of several international researchers coming from:



( And given the amount of portuguese speakers, we could also think of changing the name of the experiment to <u>Cisne</u> )