

Einstein Telescope Mock Data Challenge

ET: Scienza e Tecnologia in Italia, Assisi, 20-23 February 2024

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NextGenerationEU





Why ET MDC

 Training on simulated data to test methods and pipelines

 Find out the limitations of current methods

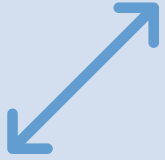
 Encourage the community to develop new tools

 Provide a common dataset for comparison of analysis methods

 Assess the science potential with ET

 Assess the requirements for computing infrastructure

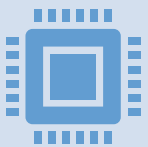
What's in the ET MDC



Series of MDCs with increasing complexity



First MDC with Gaussian coloured noise and a cosmological population of CBCs (more details in the next slide)



Next MDCs with all type of sources, glitches, correlated noise (need to be implemented in the simulation code)

First ET MDC

1 month duration



Gaussian colored noise
(ET-D 10 km, $f_{\text{Min}}=5\text{Hz}$),
triangle



Population of BBHs (10%),
BNSs (87%) and NSBHs
(3%) (see CoBA) with
isotropic distribution in the
sky.



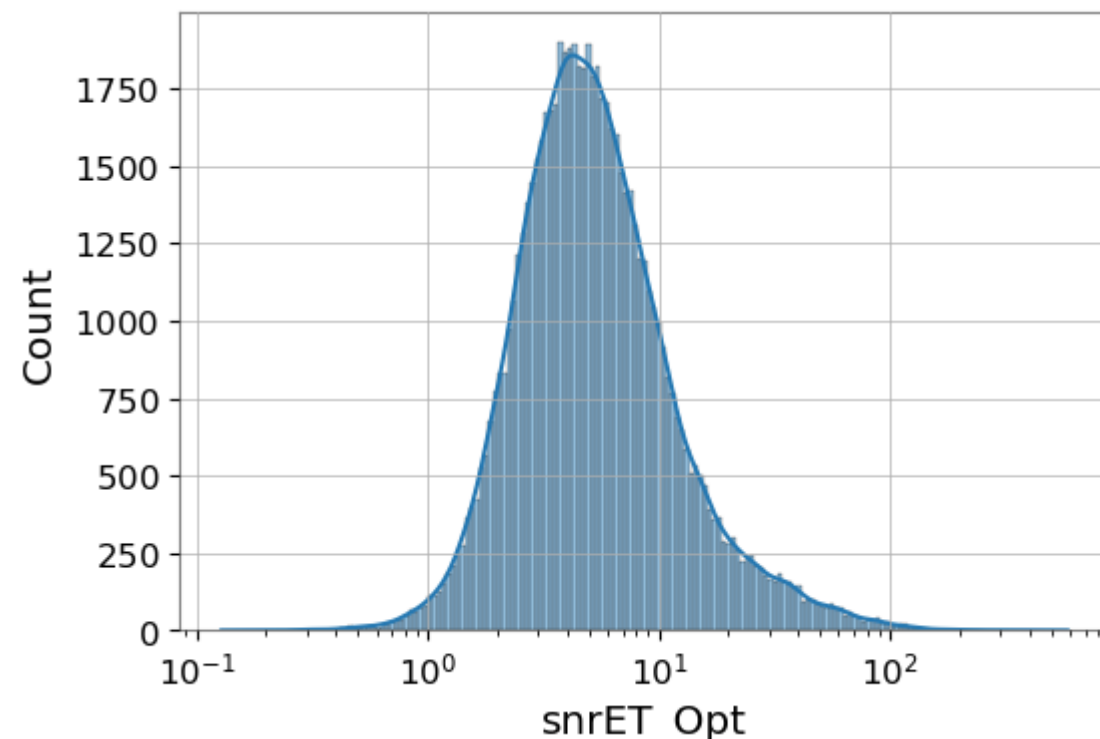
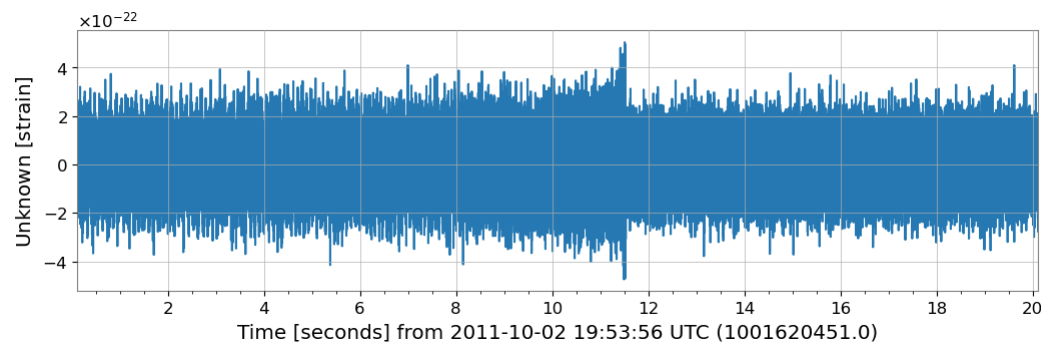
CBC Waveforms:
IMRPhenomXPHM for
BBHs and BHNSs, and
IMRPhenomPv2 with tidal
effects NRTidalv2_v for
BNSs.

Tania Regimbau, Thomas Dent, Walter Del Pozzo, Stefanos Giampanis, Tjonnje G. F. Li, Craig Robinson, Chris Van Den Broeck, Duncan Meacher, Carl Rodriguez, B. S. Sathyaprakash, and Katarzyna Wójcik
Phys. Rev. D **86**, 122001 – Published 3 December 2012

Some Statistics about the injected signals

SNR>8: 11551 BNSs , 537 BHNSs, 6119 BBHs,

SNR>12: 4048 BNSs , 238 BHNSs, 5228 BBHs



Challenges

Beginner

- Recovery of high-SNR signals within given time windows SNR = 597, 386, 383 (BNS), 374, 343, 306

Expert

- Parameter estimation of ultra-high SNR BBH signals
- Long duration binary neutron stars
- Overlapping signals

The Data



[/cvmfs/et-gw.osgstorage.org/et-gw/PUBLIC/MDC1](https://cvmfs.et-gw.osgstorage.org/et-gw/PUBLIC/MDC1)

Instructions here: <https://wiki.et-gw.eu/EIB/SoftwareFrameworks/WebHome?validationkey=e2698d03b6eff5856cfab4654d3fbfe5>

Frame files for E1, E2, E3 and E0 (set 0: noise only, set 1: noise+GWs)

1300 frames per detector of length 2048s and sampling rate 8192 Hz (1.3 TB) + frames for Cosmic Explorer, CEA and CEB

Text files with lists with the source parameters and expected SNR

CernVM File System (CVMFS)

- The CernVM File System (CernVM-FS) provides a scalable, reliable and low- maintenance software distribution service...<https://cvmfs.readthedocs.io/en/stable/>
- If installed in your environment you can easily access the data as they were on local disks

- ESCAPE Data Lake

- ESCAPE AAI

The jupyter Environment

- The ESCAPE Virtual Research Environment
 - <https://jhub-vre.cern.ch/>
- We will find the data under the path /cvmfs/et-gw.osgstorage.org/et-gw/PUBLIC/MDC1
- You can find the notebooks at the github repo:
 - <https://github.com/elenacuoco/ET-MDC-Tutorials>

Notebooks

Read and
Plot the
Data

Estimate
the Power
Spectral
Density

Whiten the
data

Transform
the data in
Time-
Frequency
domain

Brief
example of
Matched
Filter
Application