

Finanziato dall'Unione europea NextGenerationEU







Exotic Quark Decay in the 331 Model: LHC Prospects and Computational Techniques G. Corcella, C. Corianò, <u>D. Melle</u>, P. H. Frampton



ICSC Italian Research Center on High-Performance Computing, Big Data and Quantum Computing

Missione 4 • Istruzione e Ricerca









Introduction

Why there are three families in the Standard Model? Need to go beyond...

331 Model(s) are a possible answers:

- $SU(3)_C \times SU(3)_L \times U(1)_X$
- Fermions are in triplets and antitriples of $SU(3)_L$
- Generations are not treated democratically
- Anomaly cancellation and QCD asymptotic freedom constrains the number of families to three
- $SU(3)_C \times SU(3)_L \times U(1)_X \rightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \rightarrow SU(3)_C \times U(1)_{em}$
- Involved scalar potential with 26 parameter (at least)
 → Difficulties in finding valid benchmarks









Benchmarcks

How to find valid benchmarks for a BSM model?

-SARAH/SPheno/HiggsBounds -SSP automate the chain

In the case of the 331 model with 3 scalar triplets and a sextet

$M_{h_1} = 125.31$	$M_{h_2} = 6799.5$	$M_{h_3} = 10490$	$M_{h_4} = 17673$	$M_{h_5} = 17737$
$\mathcal{M}_{Ah_1} = 10489$	$\mathcal{M}_{Ah_2} = 17735$	$M_{H_1^{\pm\pm}} = 10572$	$M_{H_2^{\pm\pm}} = 17674$	$M_{H_{2}^{\pm\pm}} = 17795$
$M_{H_1^{\pm}} = 2566.4$	$M_{H_2^{\pm}} = 104900$	$M_{H_2^{\pm}}^{-} = 17674$	$M_{H_4^{\pm}}^2 = 17736$	${ m M}_T^{-}=5656.9$
${ m M}_{D_1} = 5656.9$	$M_{D_2} = 5656.9$	$M_{Y^\pm}=1328.0$	$M_{Y^{\pm\pm}} = 1325.6$	$\mathcal{M}_{Z'} = 4150.4$

How to do collider phenomenology (LHC, FCC)?

-MadGraph5 -MadAnalysis5









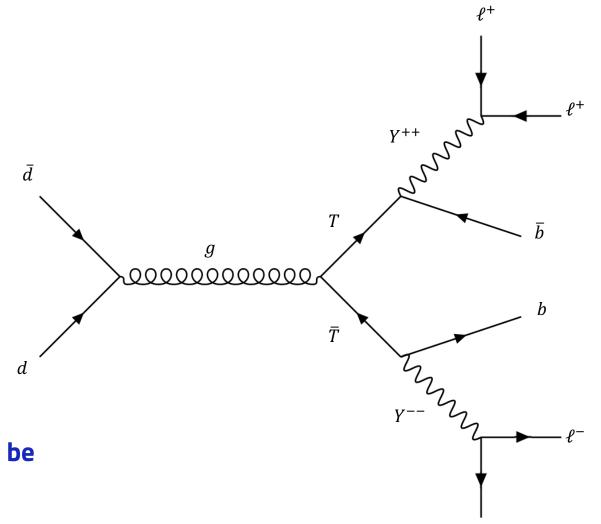
Exotic Quark Decay

The charge operator has a free parameter β In the case of Frampton's version $\beta = \sqrt{3}$

- •
- T has charge $\frac{5}{3}e$ Y⁺⁺ has charge 2e •

-Monte Carlo simulation -Data Analysis

Parton Shower and Hadronization have yet to be done









0.0225

0.0200

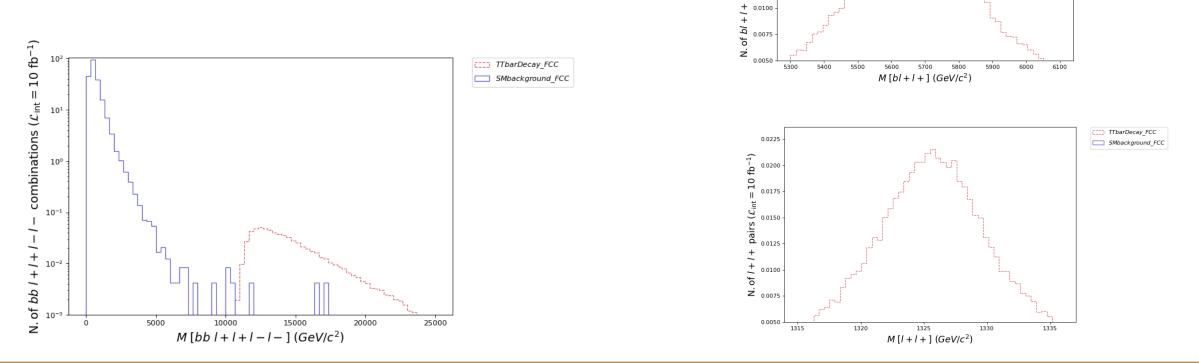


TTbarDecay FCC

SMbackground FCC

Some preliminary results





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Conclusions and future developments

Improving the benchmark search for deeper analysis of a complicated potential with many scalars and parameters

Analysis of possible first order phase transitions in BSM models with involved scalar sector like 331 Model(s):

- Nucleation temperature
- Phase transition strength
- Rate of the phase transition
- Bubble wall velocity

In order to obtain the predicted spectrum of the cosmological GW background