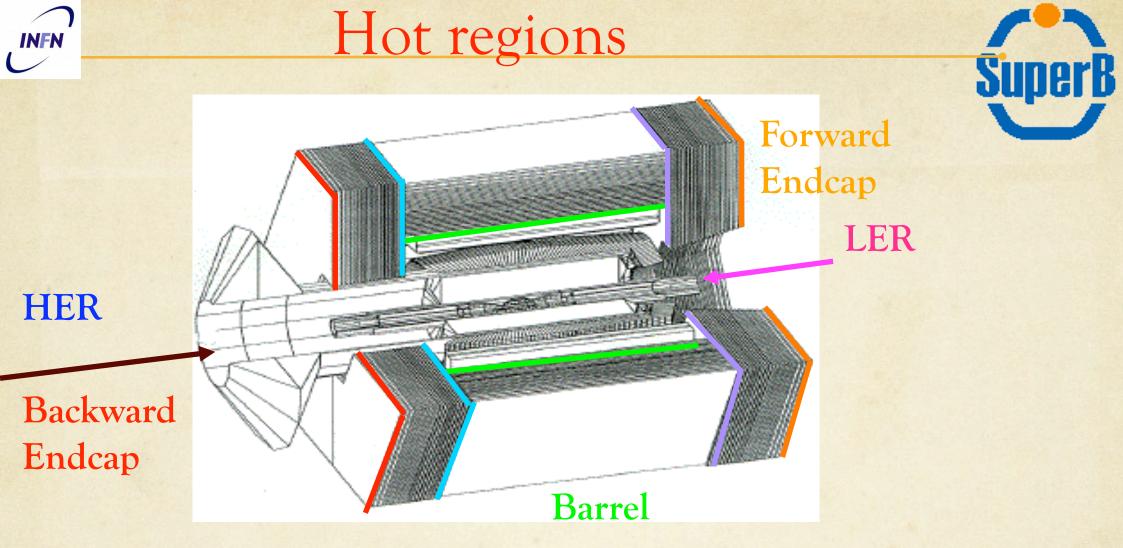


• IFR Background Status Valentina Santoro INFN Ferrara

09/13/2011

INFN



Barrel: innermost layers, mostly neutrons and photonsFWD encaps (hottest region) : inner layer and outer layers (BEAM halo), neutrons, electron and photonsBWD encaps: inner layer and small radii

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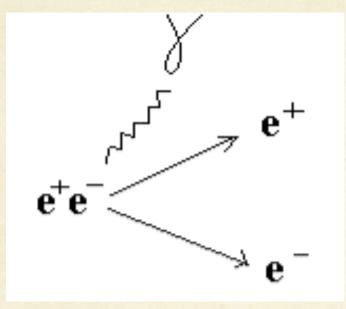
Outline

SuperB

- Radiative BhaBha events (Elba 2011 production):
 - Study of the neutrons, photons and electron background crossing the IFR
 - Study of the the neutrons and electron background crossing the IFR FEEs boards

- Touschek scattering events:
 - Study of the neutrons, photons and electron background crossing the IFR

Radiative BhaBha background crossing Super the IFR detector



Why we don't like Radiative BhaBha:

This background scales with the luminosity

The high energy γ coming from radiative Bhabha strikes a beamline elements and showers secondaries with energies ranging from sub-MeV to several tens of MeV, including MeV-energy neutron producted via the giant dipole resonance

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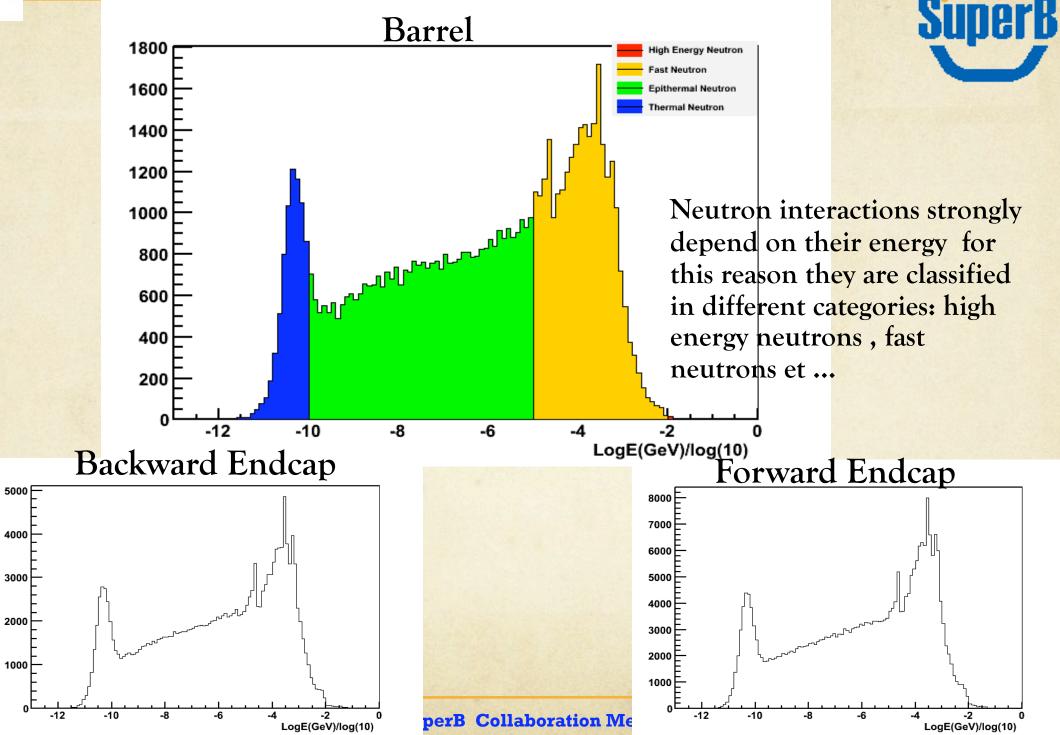


Why do we have to worry about neutrons

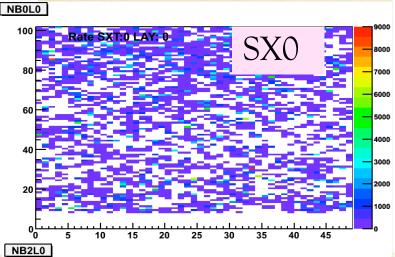
- Neutrons damage silicon devices → Neutrons damage SiPM
- The silicon damage function has a strong dependance on the energy spectrum therefore we scaled all the doses in this presentation to 1MeV equivalent accordingly to ASTM E 722 93.



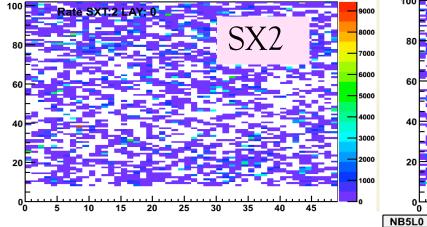
Neutron Energy Distributions

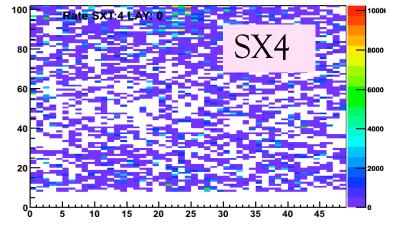


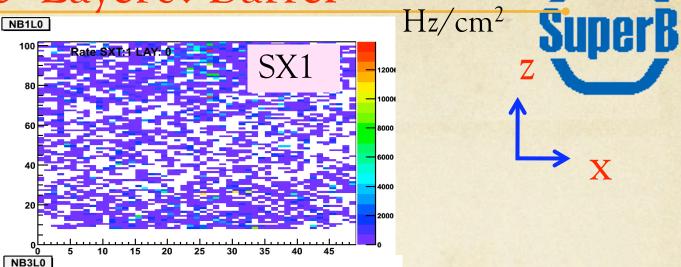


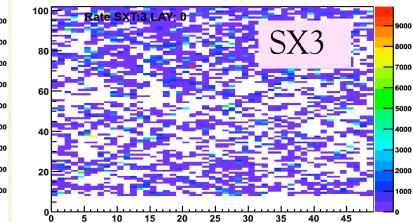


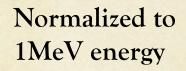
INFN



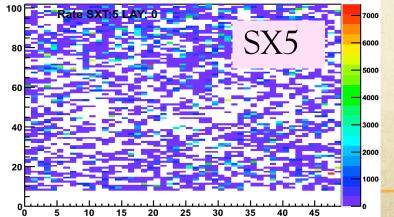


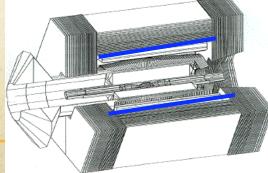






X



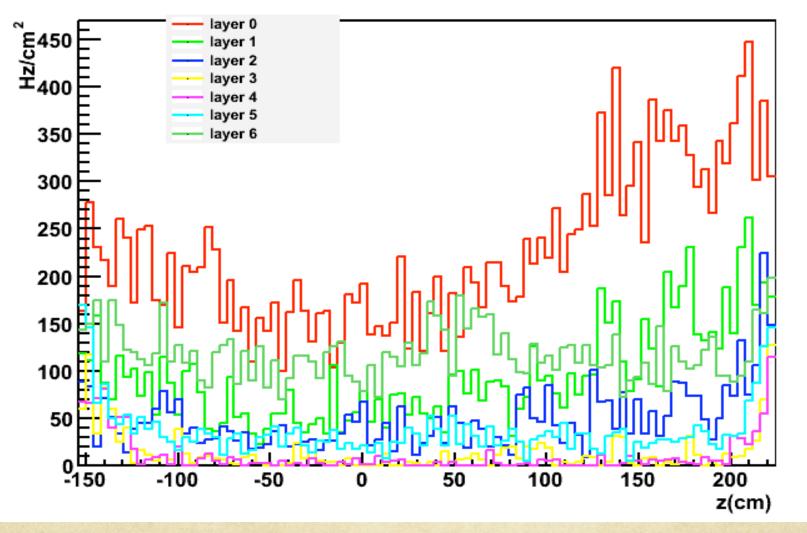




Rate for Barrel Different layers



Rate vs Z-coordinate



Rate of $450 \text{Hz/cm}^2 \rightarrow \text{about } 3x10^9 \text{ neutrons/cm}^2$ for a year

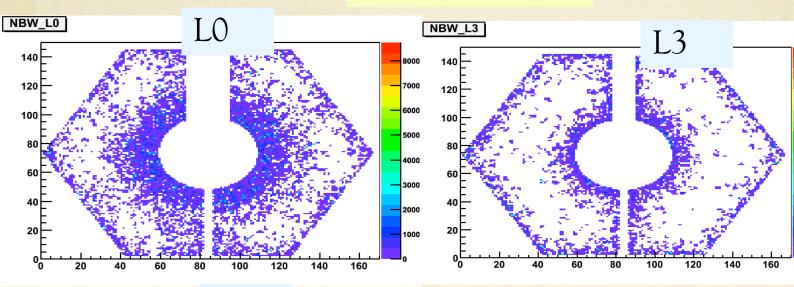
Valentina Santoro

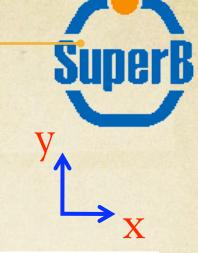


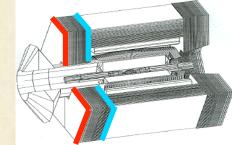
NFW_L0

Endcap

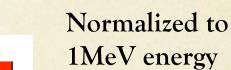
BWD ENDCAP

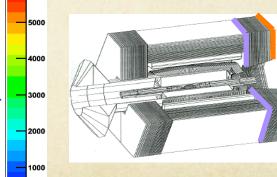






FWD ENDCAPL3

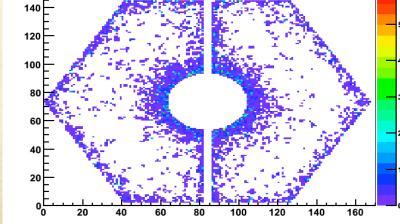




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LO

NFW_L3



Valentina Santoro

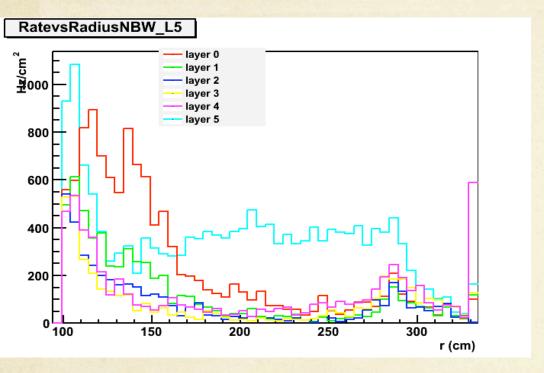


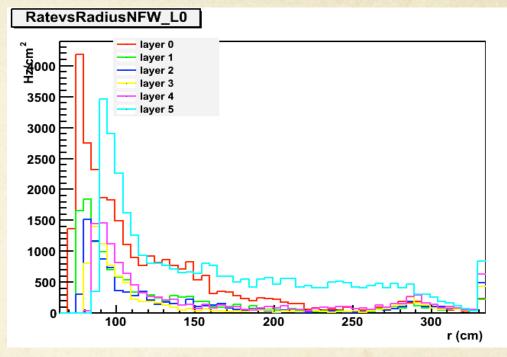
Rate vs Radius for Endcap for Different layers

SuperB

BWD ENDCAP

FWD ENDCAP





Rate of 600Hz/cm² -> 4x10⁹ neutrons for a year Rate of 3000Hz/cm² -> 2x10¹⁰ neutrons for a year







Why do we have to worry about photons

- High Energy Photons convert in e⁺e⁻ that produce signal in the detector
- Visible photons can produce fake signal



INFN

5000

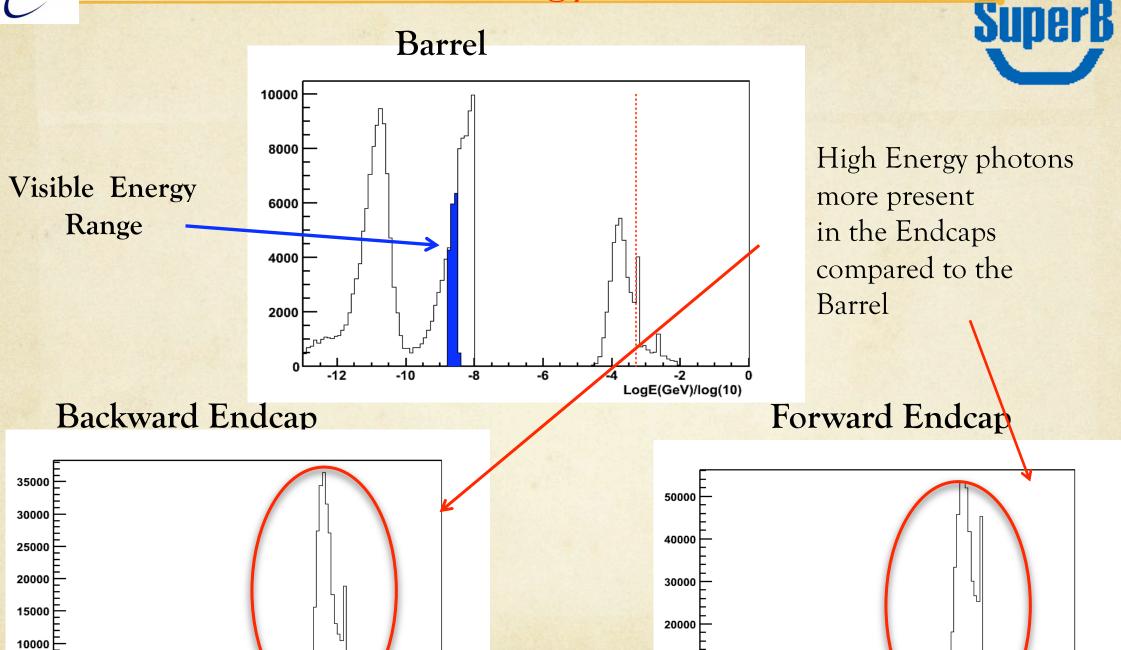
-8

.10

-6

-2

LogE(GeV)/log(10)



erB Collaboration

10000

-12

-8

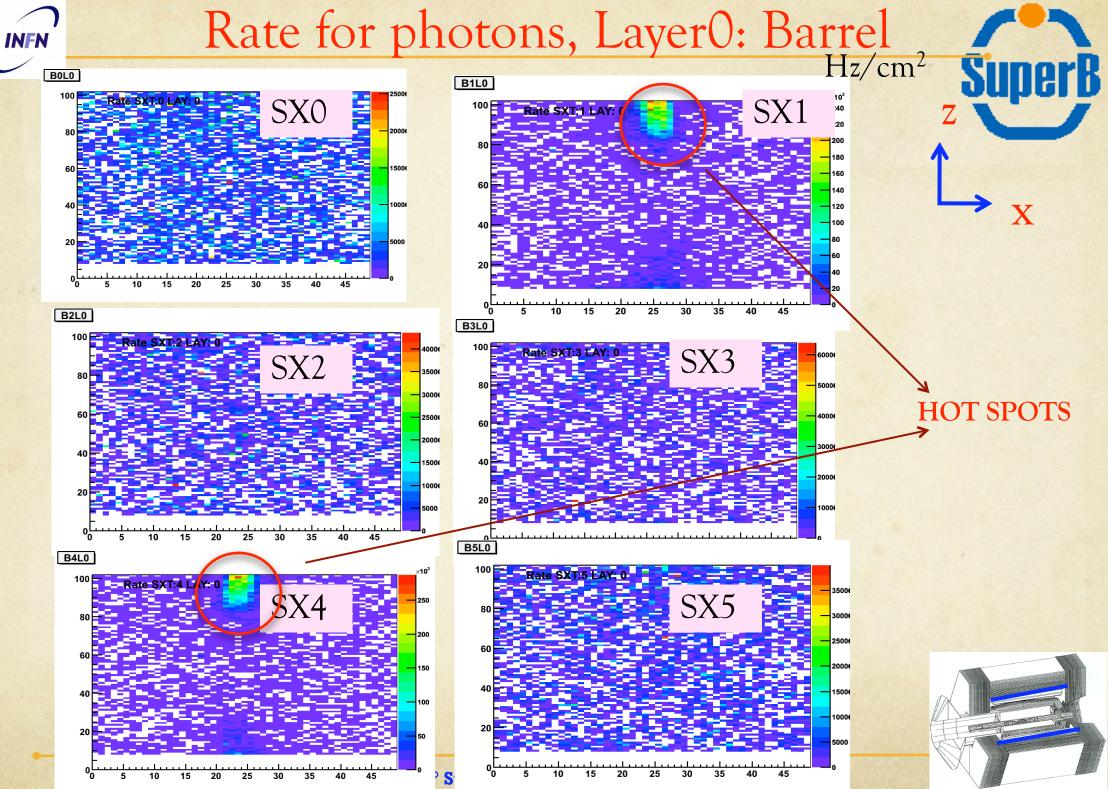
-10

-6

-4

-2

LogE(GeV)/log(10)



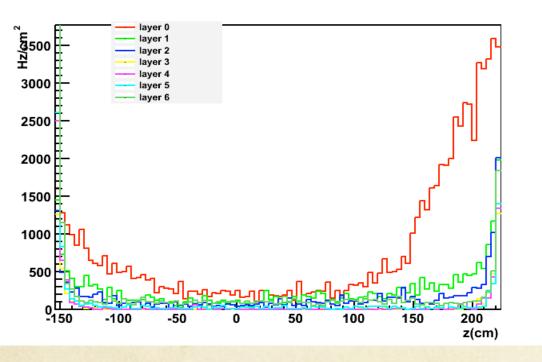


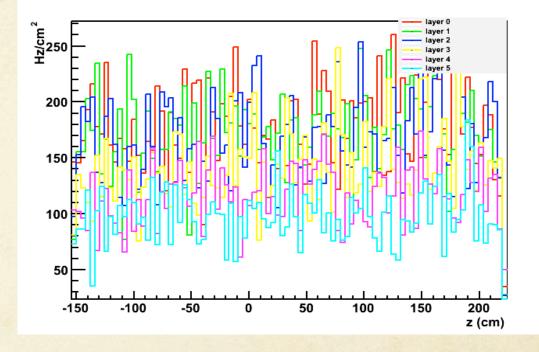
Rate for Barrel Different layers

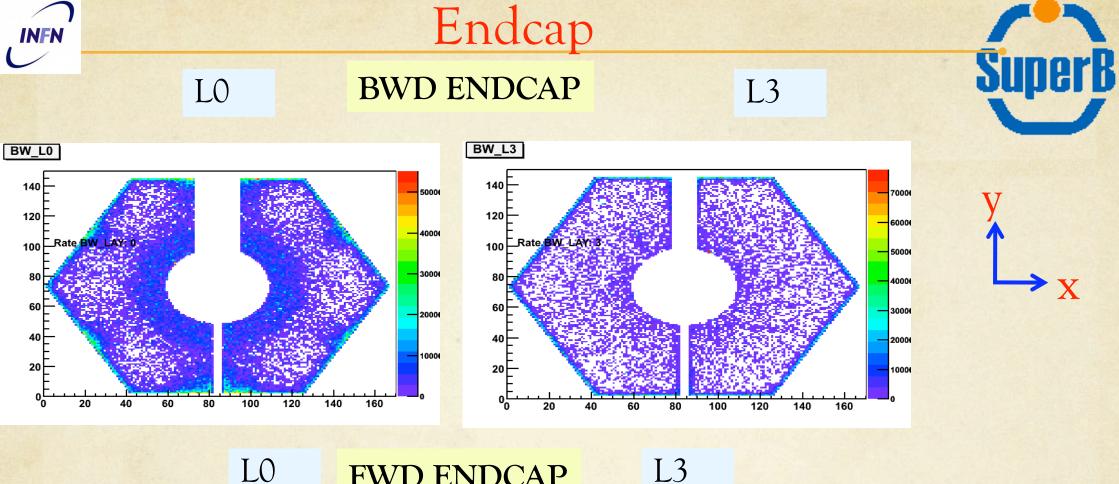


Rate vs Z-coordinate for all the γ

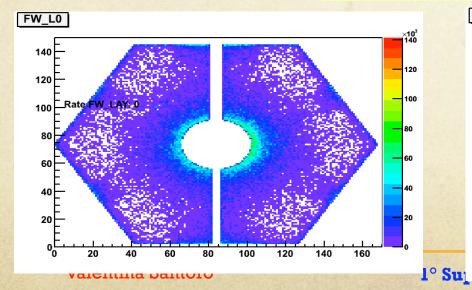
Rate vs Z-coordinate for visible γ

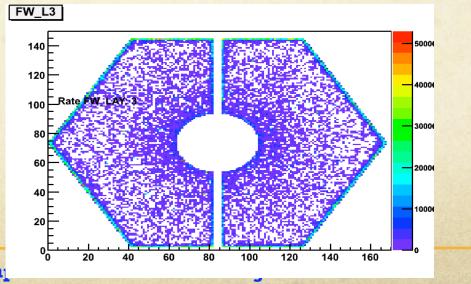






L3 **FWD ENDCAP**







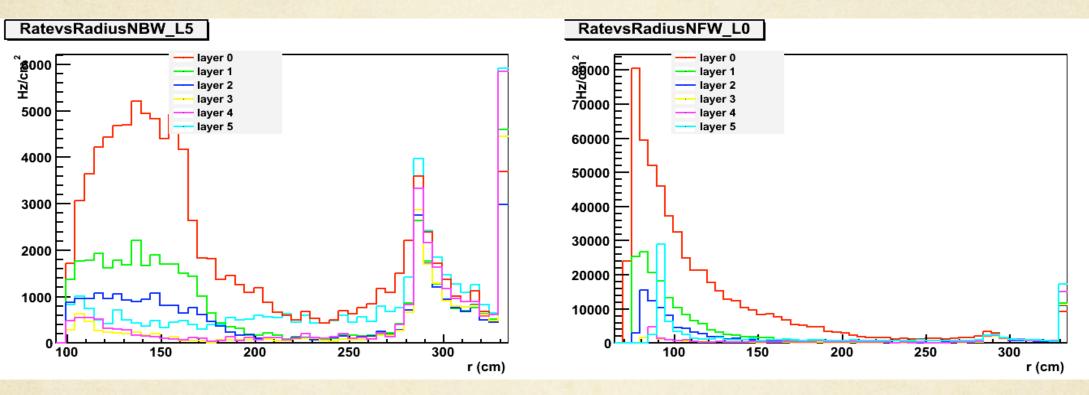
Rate for Endcap for Different layers

Rate vs radius



BWD ENDCAP

FWD ENDCAP





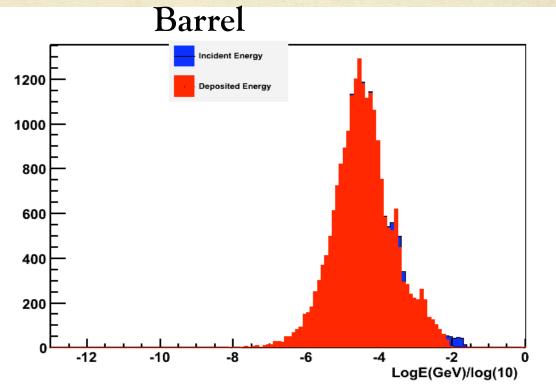




Why do we have to worry about electrons

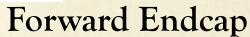
• Electrons are charged particle and can produce signals

Electron Energy Distributions

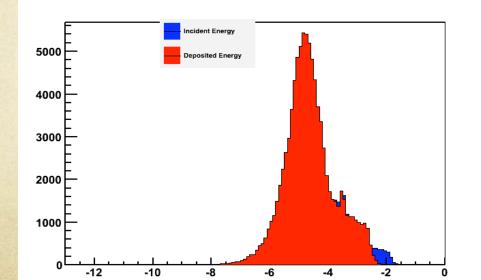


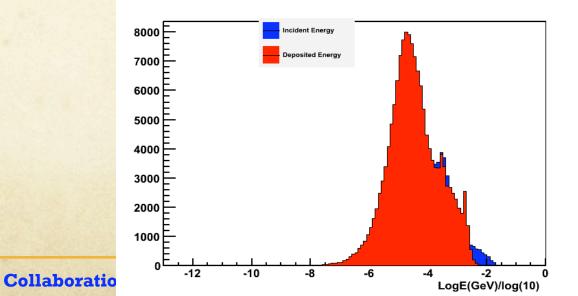
Backward Endcap

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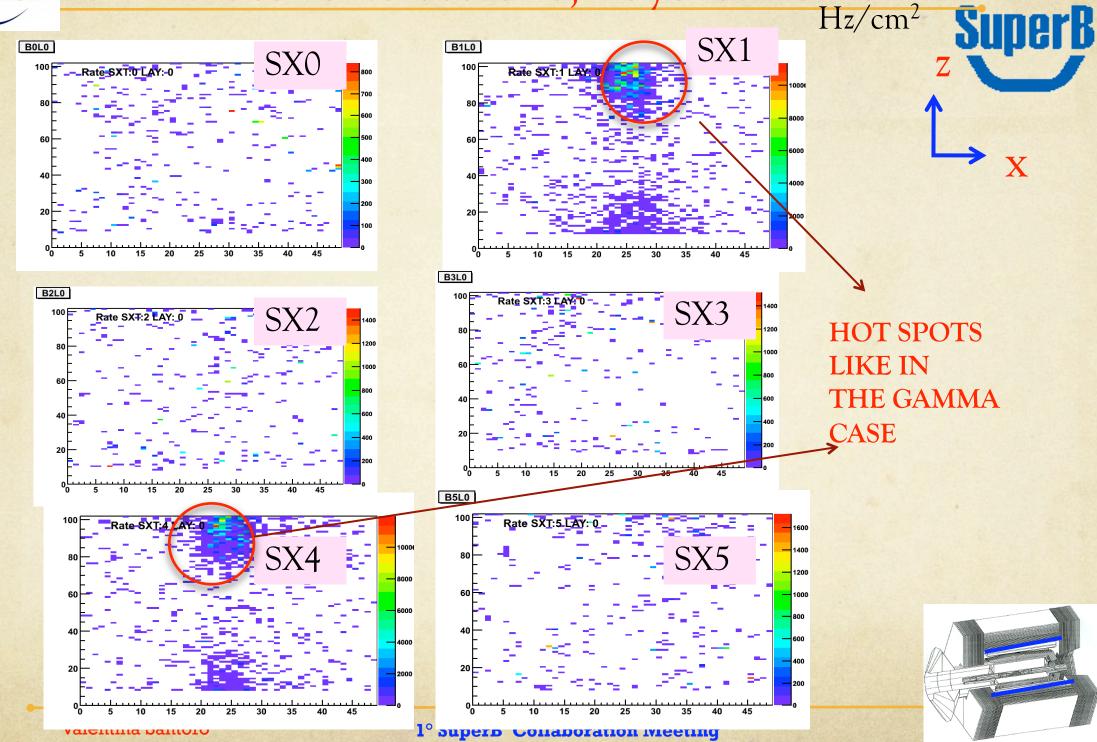
SuperB

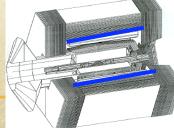




Rate for electrons, LayerO: Barrel Hz/cm²

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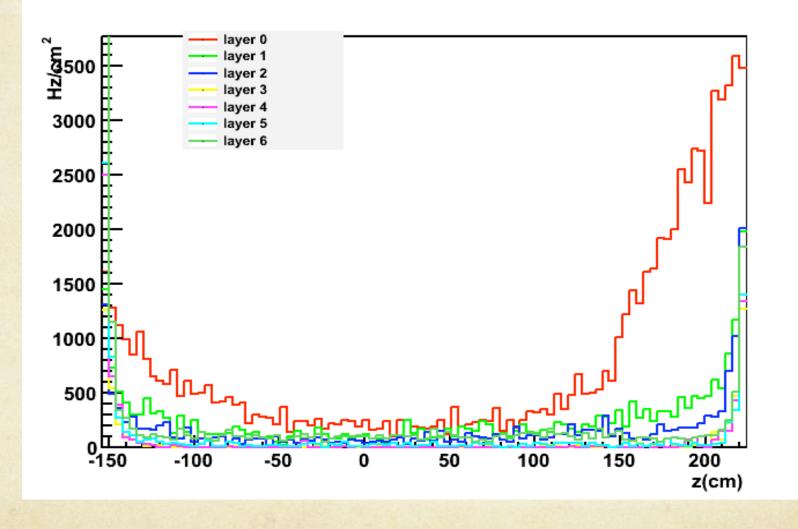
Χ

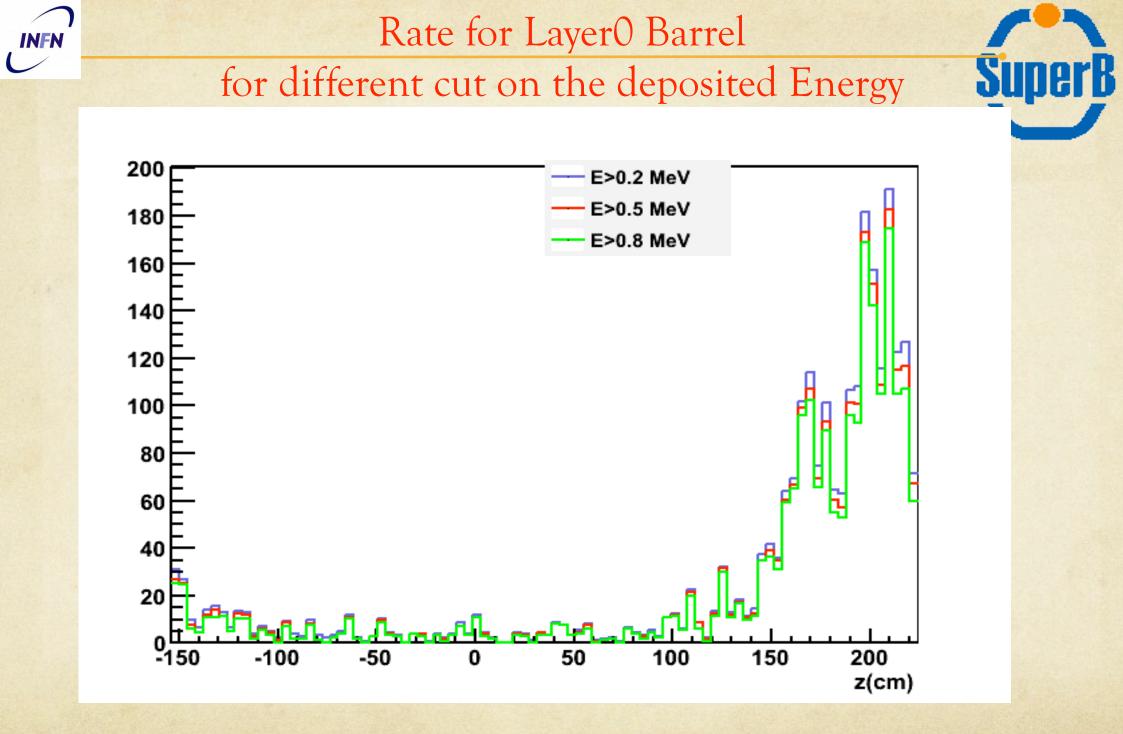


Rate for Barrel Different layers

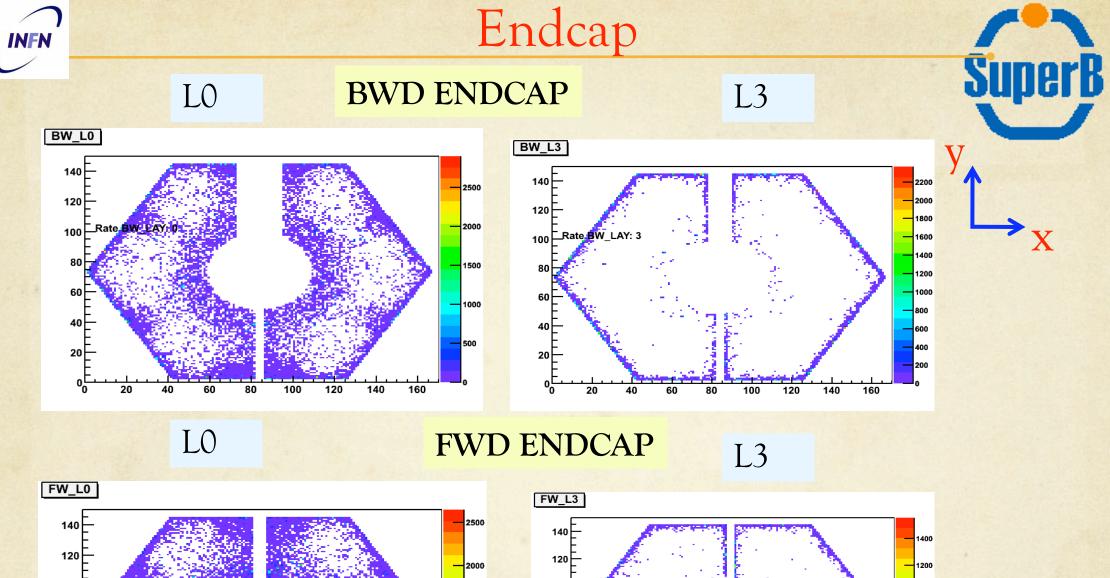


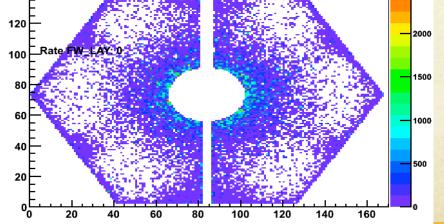
Rate vs Z-coordinate

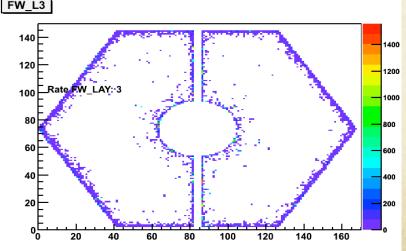




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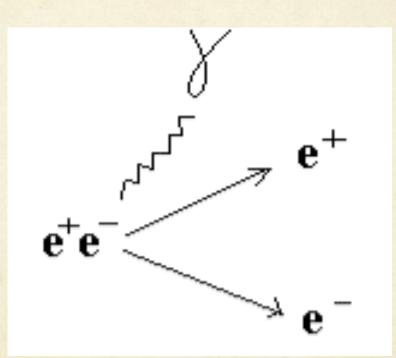


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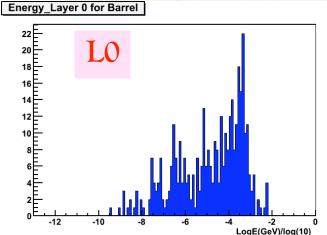
Radiative BhaBha background crossing the IFR FEE boards



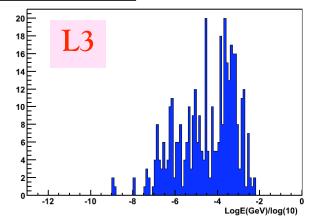


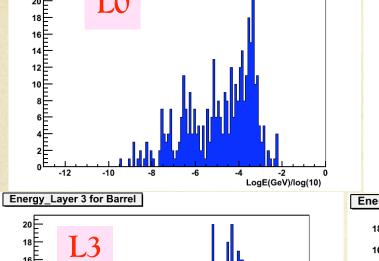
Valentina Santoro

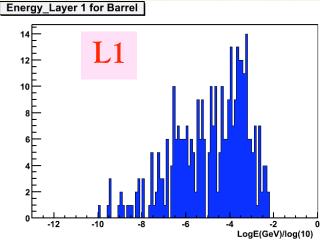
Neutron Energy Distribution for Barrel Electronics



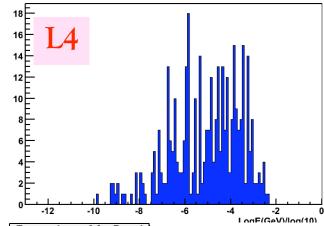
INFN

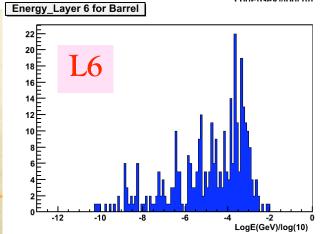




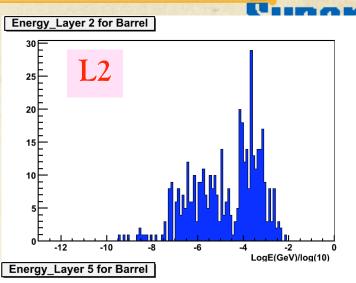


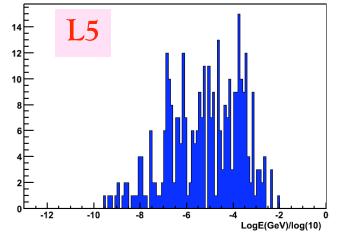
Energy_Layer 4 for Barrel





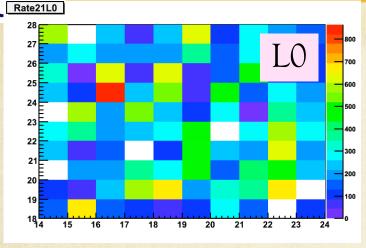
ng



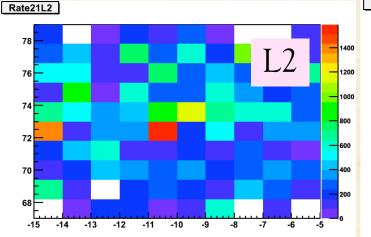


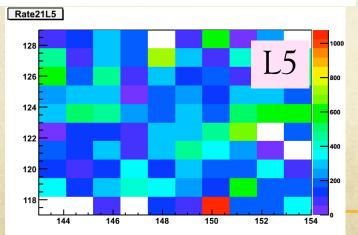
Valentina Santoro

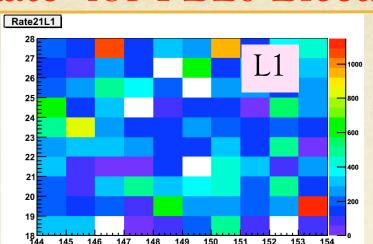
Neutron Rate for FEEs Electronics Barrel

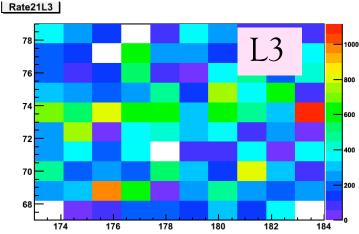


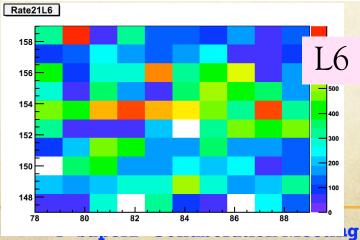
INFN

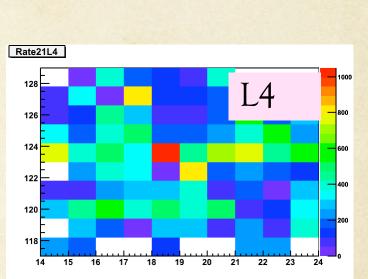












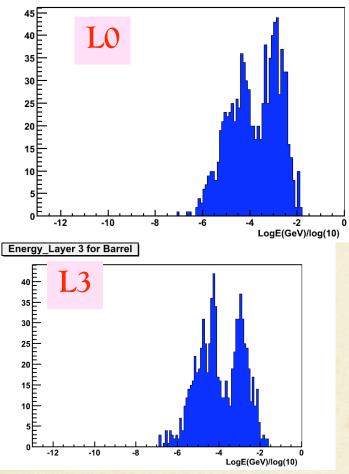
 Hz/cm^2

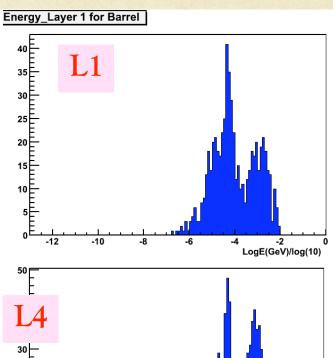
To see the plots with better resolution http://www.fe.infn.it./ ~santoro/SuperB/ Background/Neutrons/ Touschek.html rB

INFN

Energy_Layer 0 for Barrel

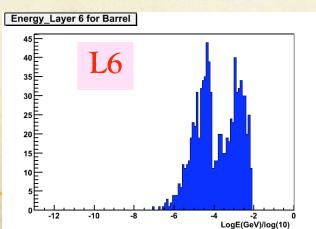
Electrons Energy Distribution for Barrel FEE





-6

-4



-10

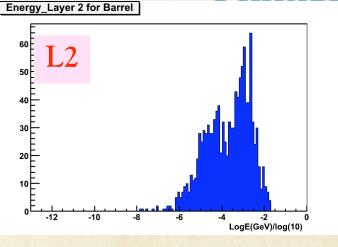
-8

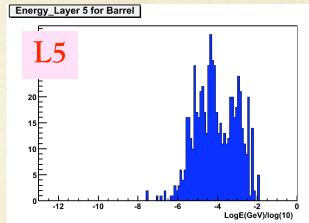
20

10

n

-12





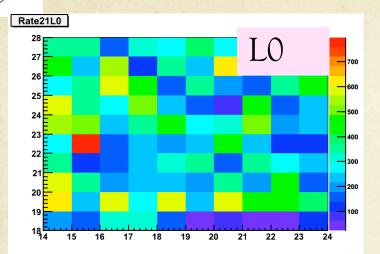
To see the plots with better resolution http://www.fe.infn.it./ ~santoro/SuperB/ Background/Electrons/ Touschek.html

Ieeting

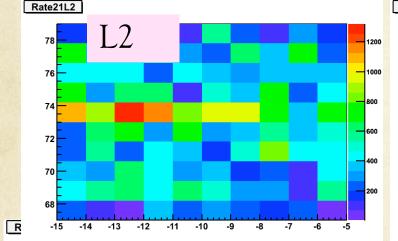
-2 LogE(GeV)/log(10)

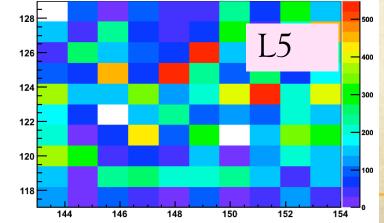
Valentina Santoro

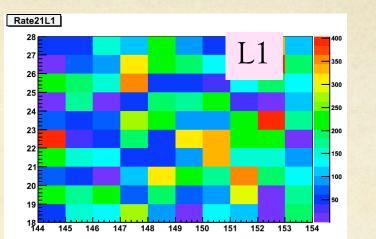
Electrons Rate for FEE Electronics Barrel

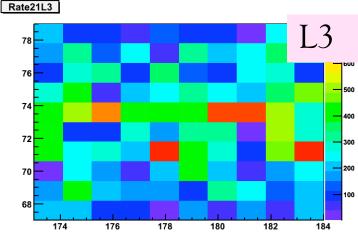


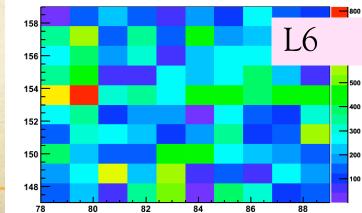
INFN



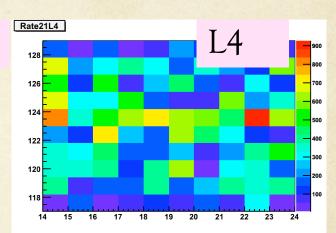








Hz/cm²



To see the plots with better resolution http://www.fe.infn.it./ ~santoro/SuperB/ Background/Electrons/ Touschek.html



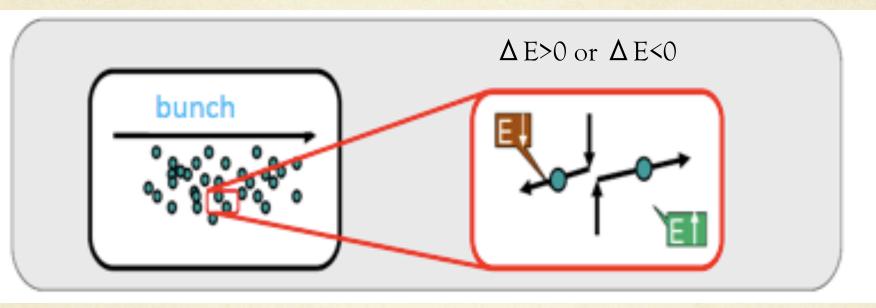


- We have the same information for the FEEs in the Forward and Backward Endcaps
- The code for analyzing the background on the FEEs is now in place if you need a detailed information (absorbed dose, particle flux) please let me know

INFN



Touschek scattering results from a Coulomb collision of two relativistic electrons in a particle beam, producing an instantaneous change in particle energy



Why we don't like Touschek events:

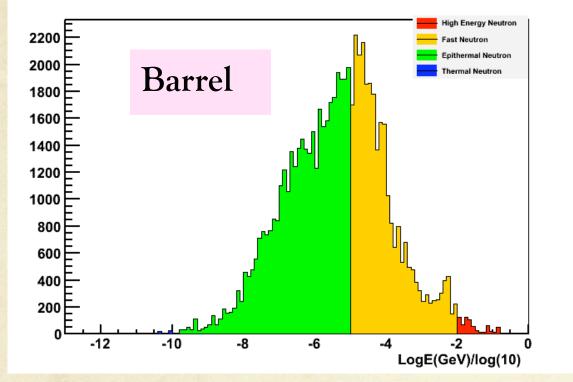
Scattered e⁺/e⁻ goes off trajectory -> lost at beam pipe wall near IP ->creates shower ->reach detector

The Touschek events in this presentation come from the HER

Valentina Santoro



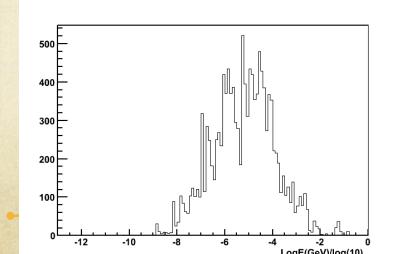
Neutron Energy Distributions for Touschek events



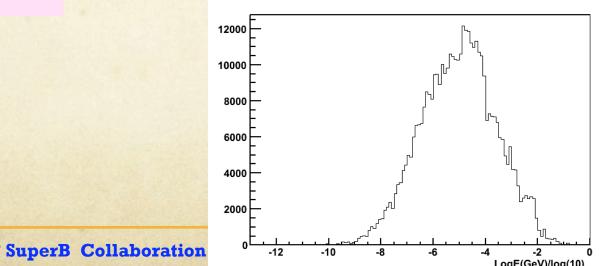
Neutron interactions strongly depend on their energy for this reason they are classified in different categories: high energy neutrons, fast neutrons et ...

SuperB

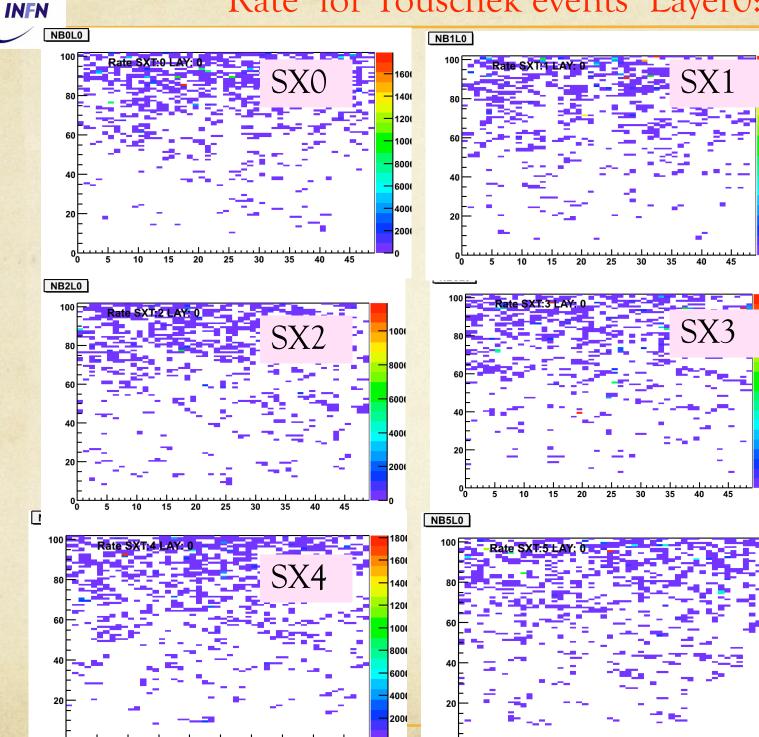
Forward Endcap



Backward Endcap



Rate for Touschek events LayerO: Barrel Hz/cm² SuperB



Normalized to 1MeV energy

X

500(

400(

300(

200(

600(

300(

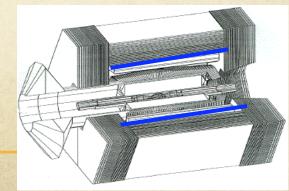
100(

700(

500(

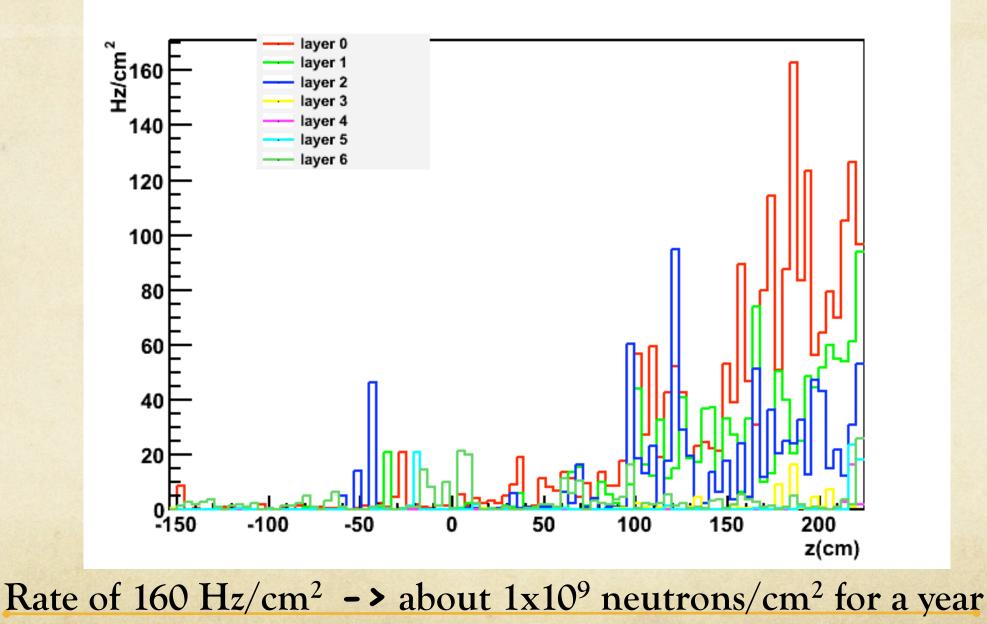
400(

300(





Rate vs Z-coordinate



Valentina Santoro

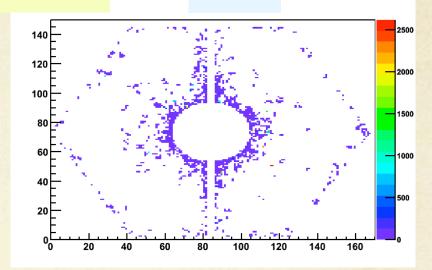
1° SuperB Collaboration Meeting

SuperB

FWD Endcap Touschek events

LO **FWD ENDCAP** NFW_L0

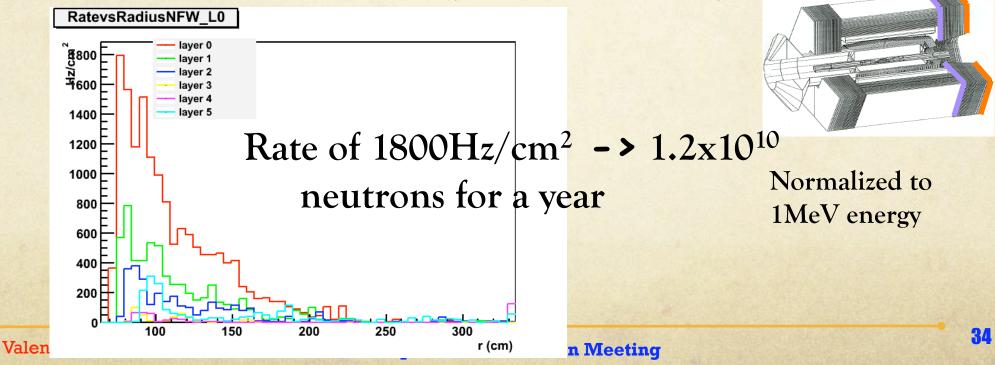
INFN



L3

SuperB

Rate vs Radius for FWD Endcap for Different layers









What we have done:

- Neutrons background crossing the IFR studied using the Elba 2011 Production
- Photons and Electron backgrounds has been studied in details for the first time
- Background from Electrons and Neutrons on FEEs boards studied
- Touschek background studied for the HER

What we will have to do

- Touschek background studies for the LER (one week)
- Add shielding for the Endcap outer layers (one week)
- Estimate the FEEs doses (few days)





For additional plots and information

http://www.fe.infn.it./~santoro/SuperB/Background/