

# The Minimal Massive Majoron

Xavier Ponce Díaz

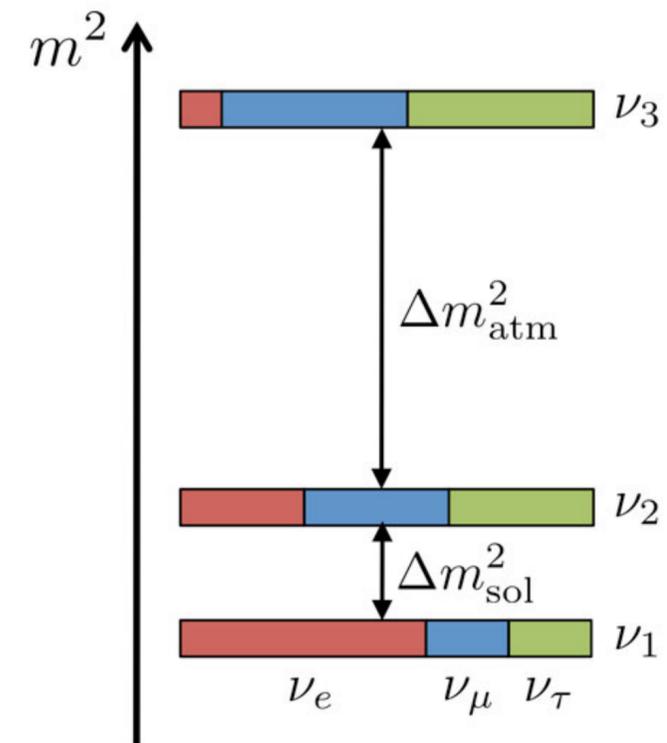
based on: [arXiv:2312.13417](https://arxiv.org/abs/2312.13417),  
with Arturo de Giorgi, Luca Merlo and Stefano Rigolin



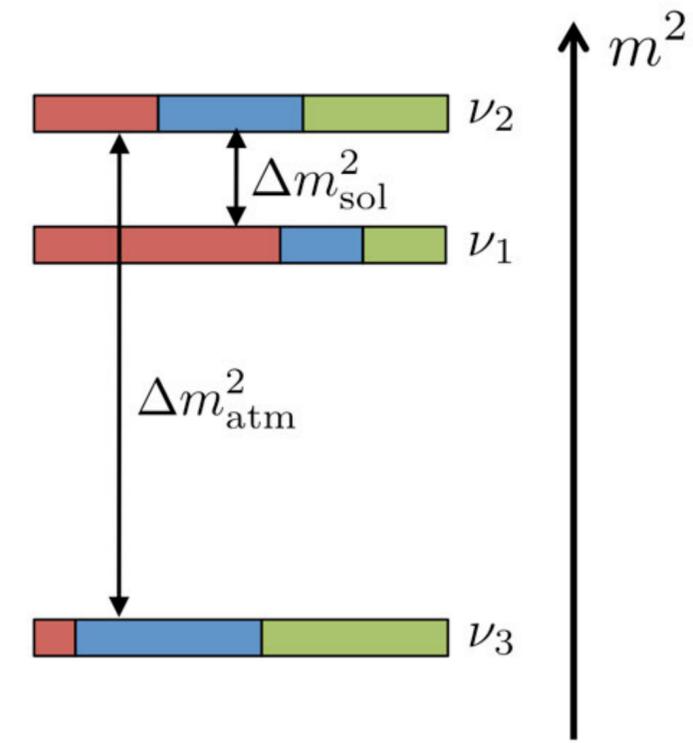
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DI PADOVA

# The Majoron

normal hierarchy (NH)



inverted hierarchy (IH)



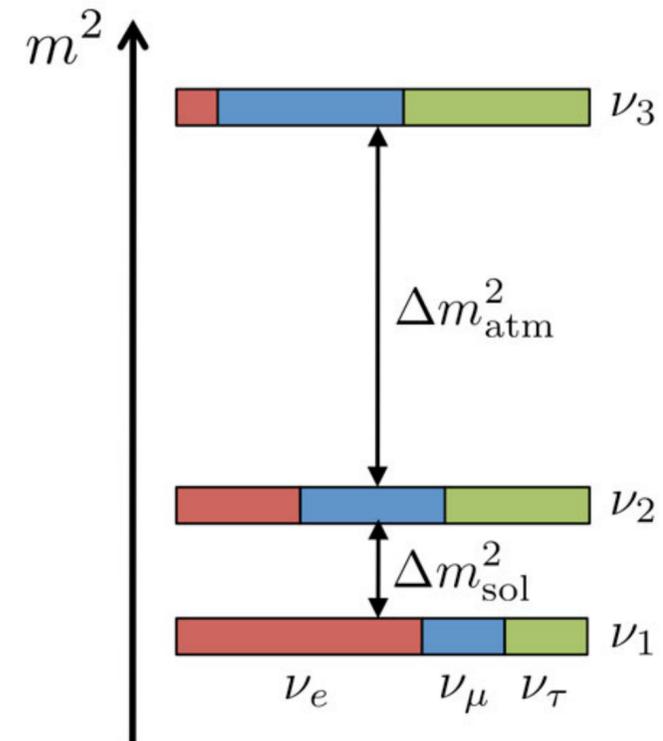
# The Majoron

Add RH neutrinos, singlets under SM gauge group:

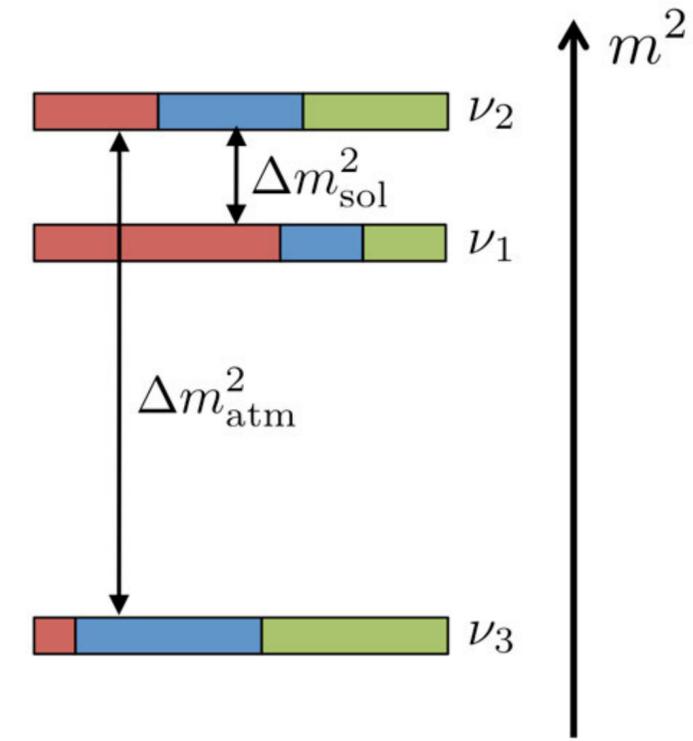
$$\mathcal{L}_{\text{type-I}} = \bar{\ell} \hat{Y}_N N_R H + \frac{1}{2} \overline{N_R^c} \hat{\Lambda} N_R + \text{h.c.}$$

$$\hat{m}_\nu \simeq \frac{v^2}{2} \hat{Y}_N \hat{\Lambda}^{-1} \hat{Y}_N^T$$

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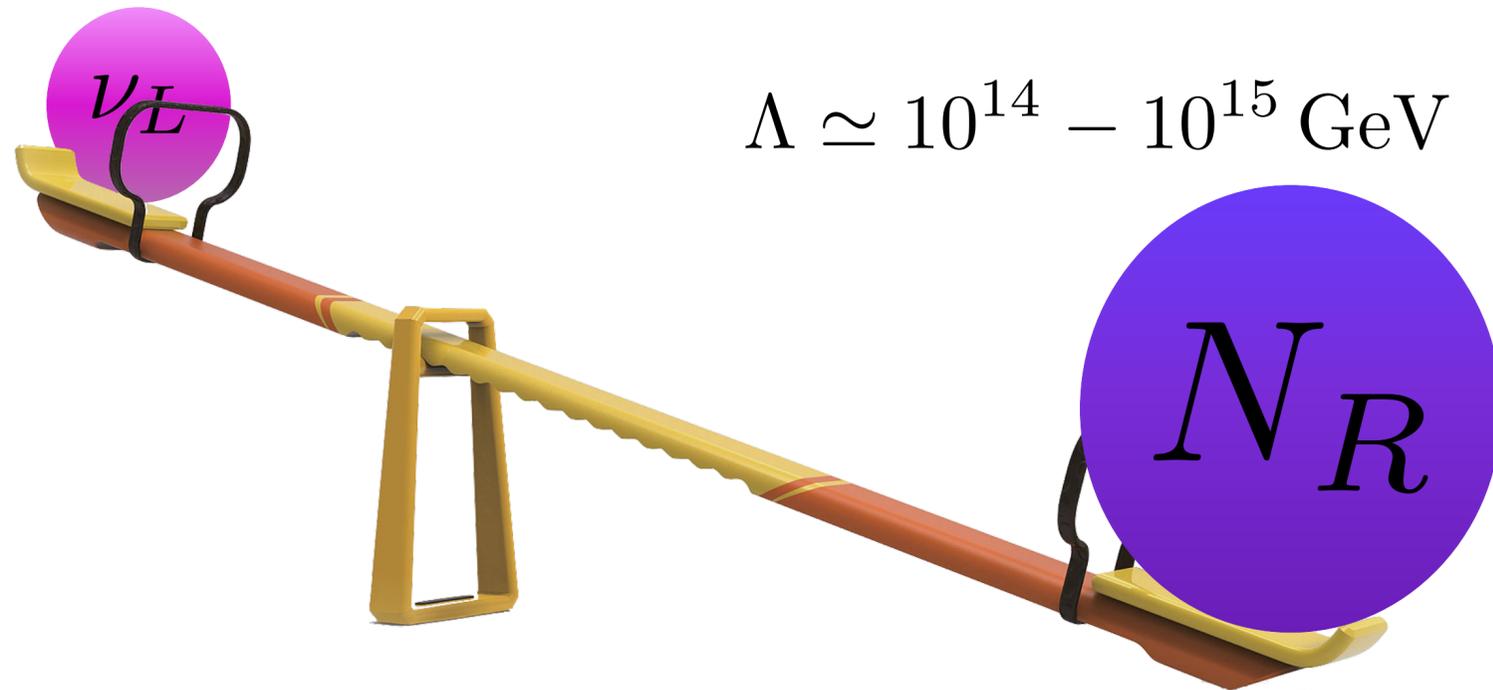
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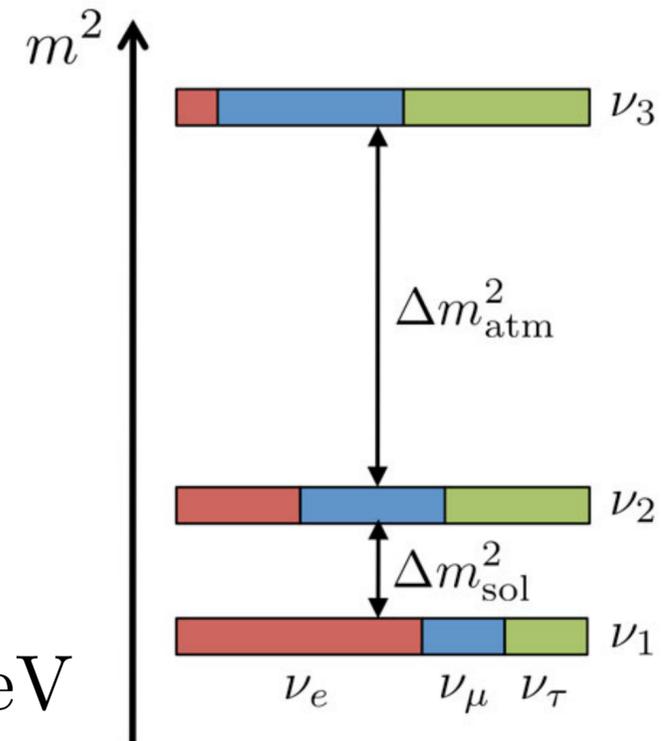
$$\mathcal{L}_{\text{type-I}} = \bar{\ell} \hat{Y}_N N_R H + \frac{1}{2} \overline{N_R^c} \hat{\Lambda} N_R + \text{h.c.}$$

$$m_\nu \lesssim 1 \text{ eV} \quad \hat{m}_\nu \simeq \frac{v^2}{2} \hat{Y}_N \hat{\Lambda}^{-1} \hat{Y}_N^T$$

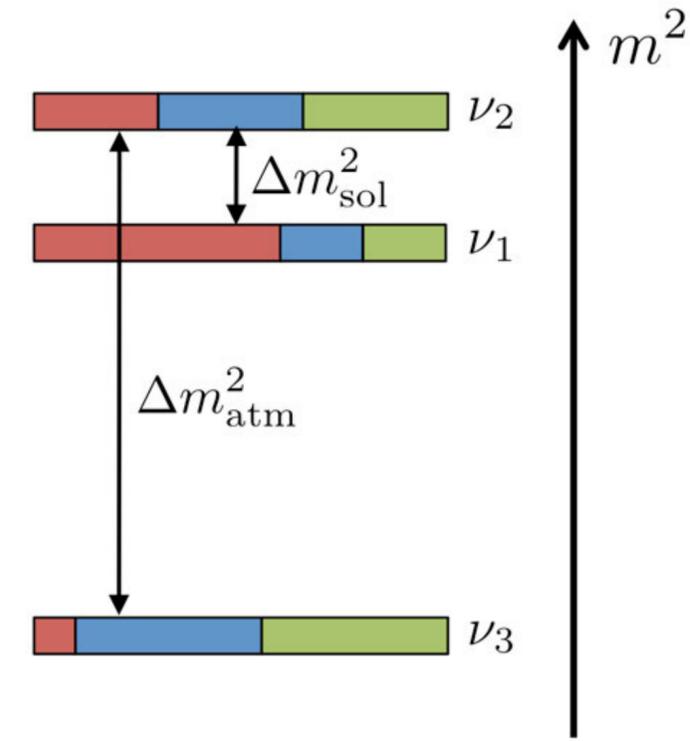
$$\Lambda \simeq 10^{14} - 10^{15} \text{ GeV}$$



normal hierarchy (NH)

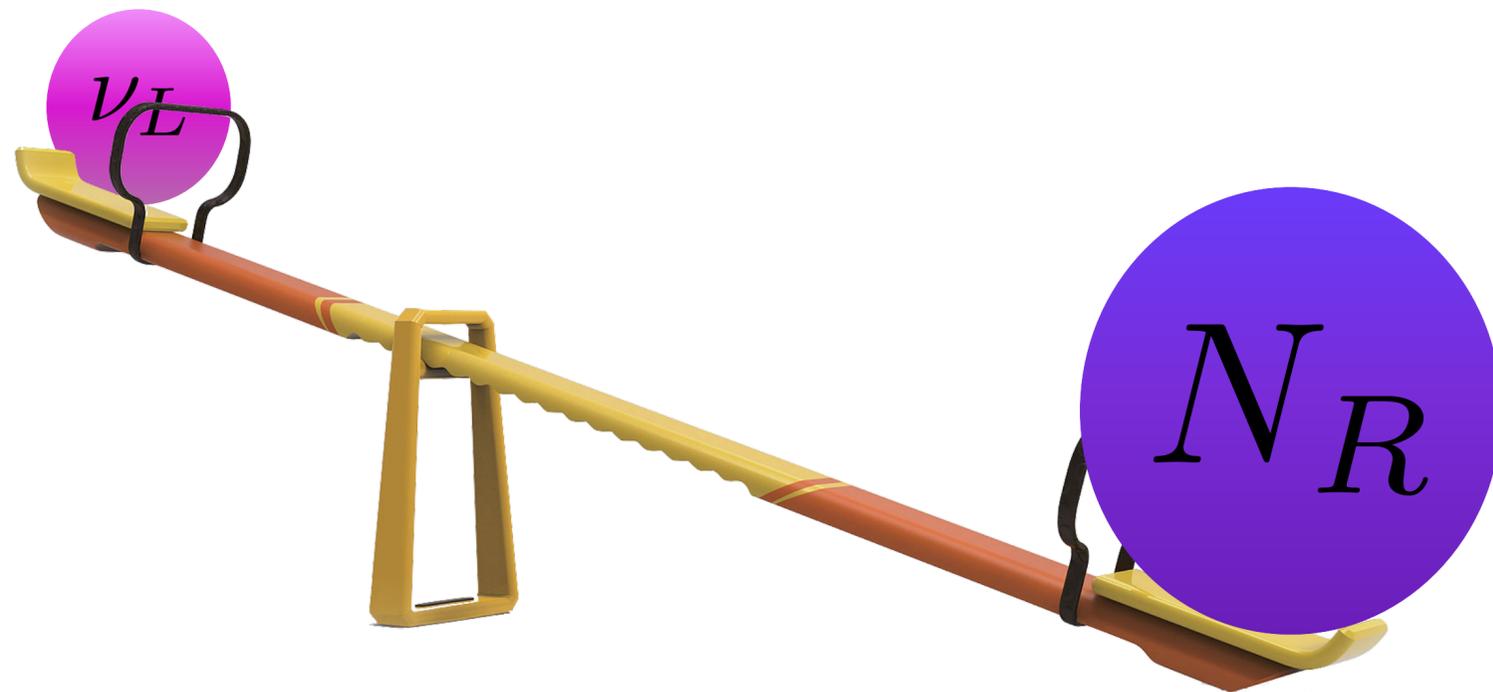


inverted hierarchy (IH)



# The Majoron

$$\mathcal{L}_{\text{type-I}} = \bar{\ell} \hat{Y}_N N_R H + \frac{1}{2} \overline{N_R^c} \hat{Y}_{NN} \phi N_R + \text{h.c.}$$



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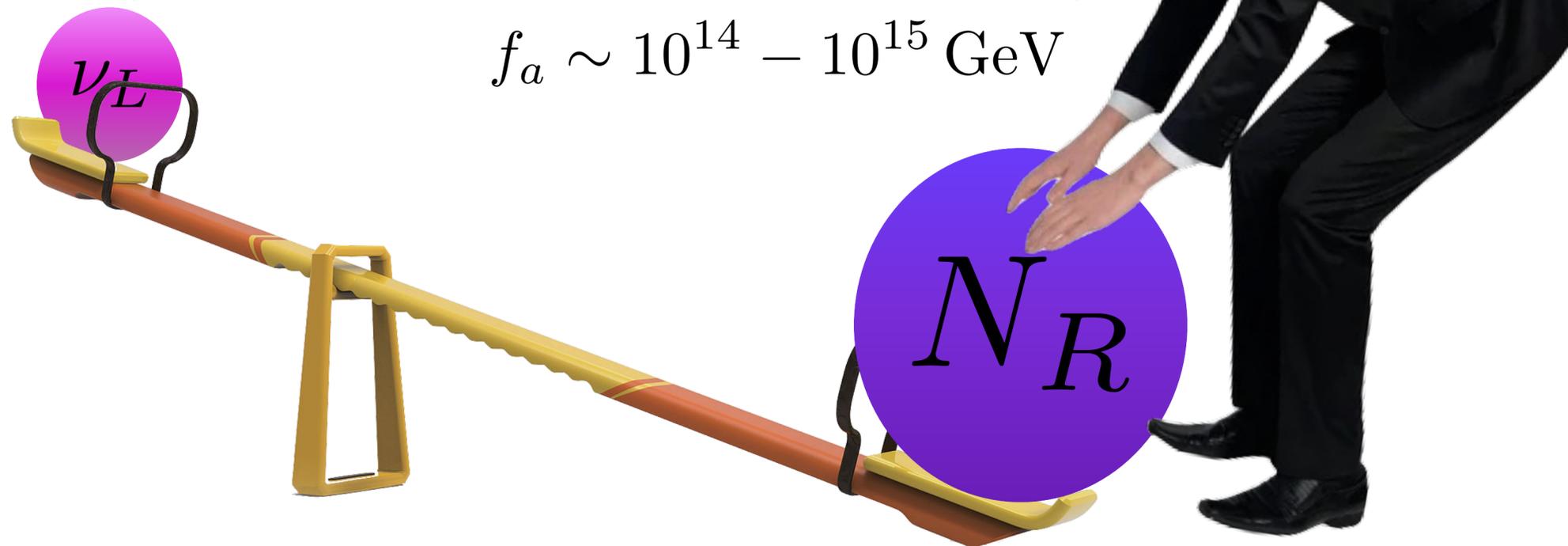
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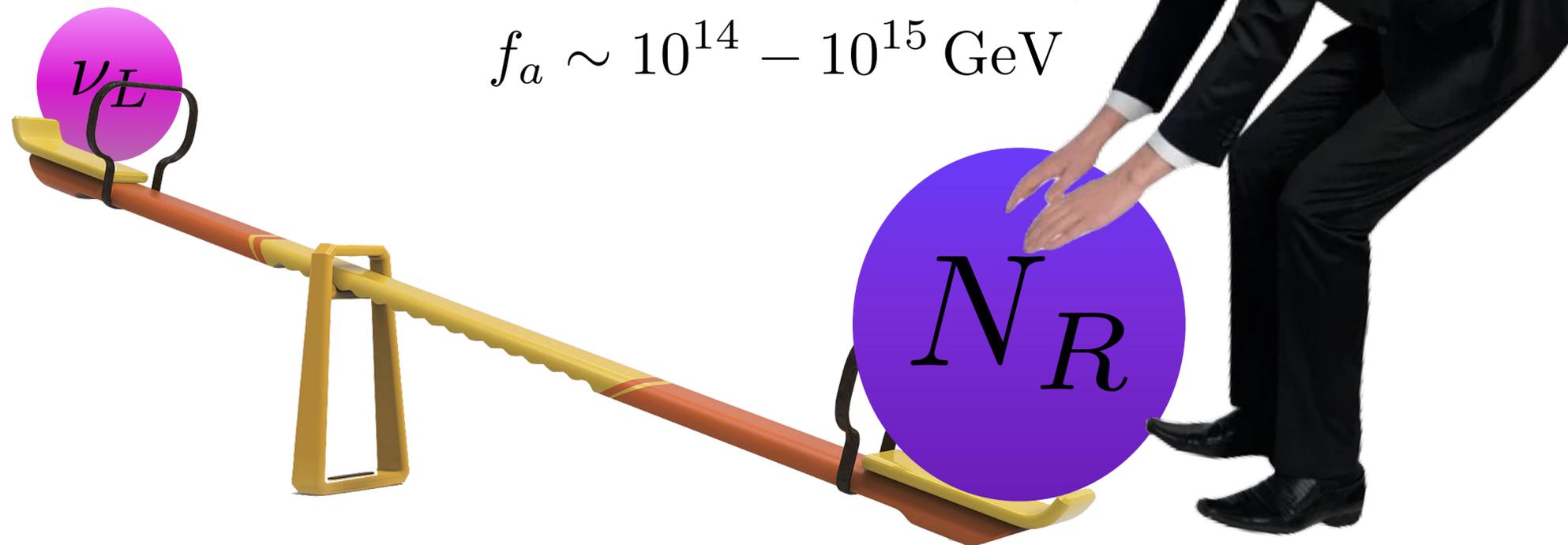
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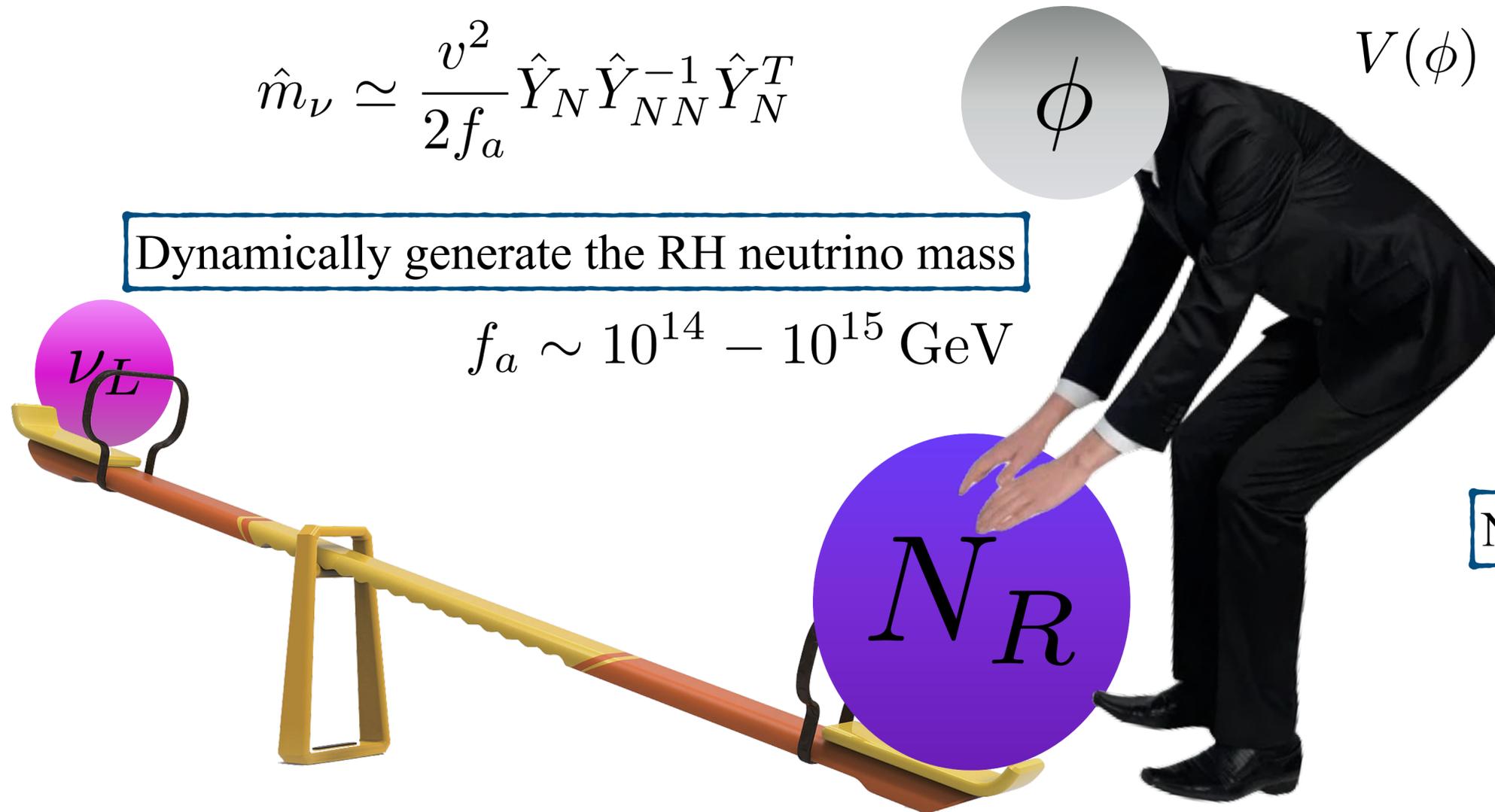
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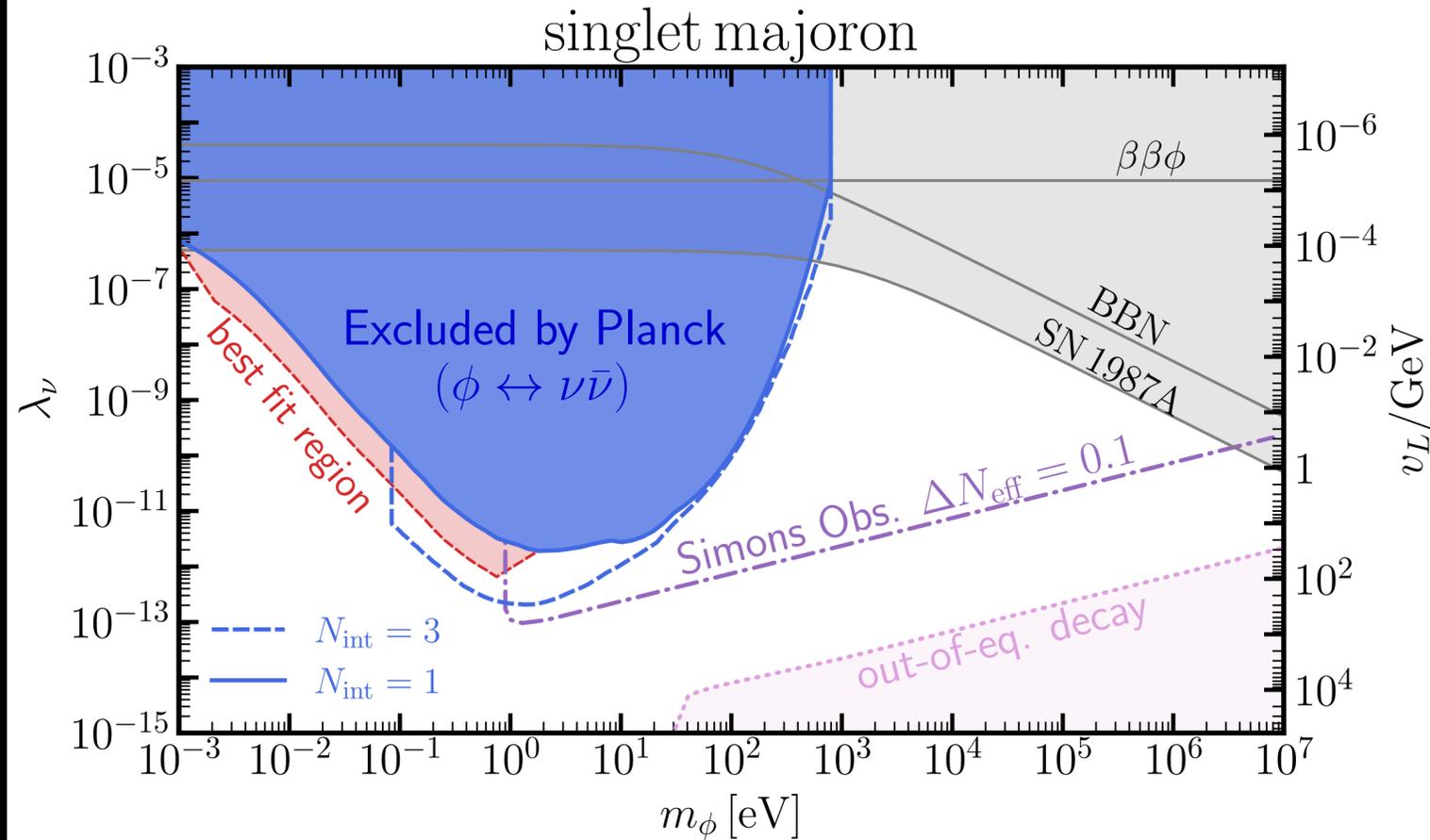
Natural light Dark Matter candidate



# Motivation

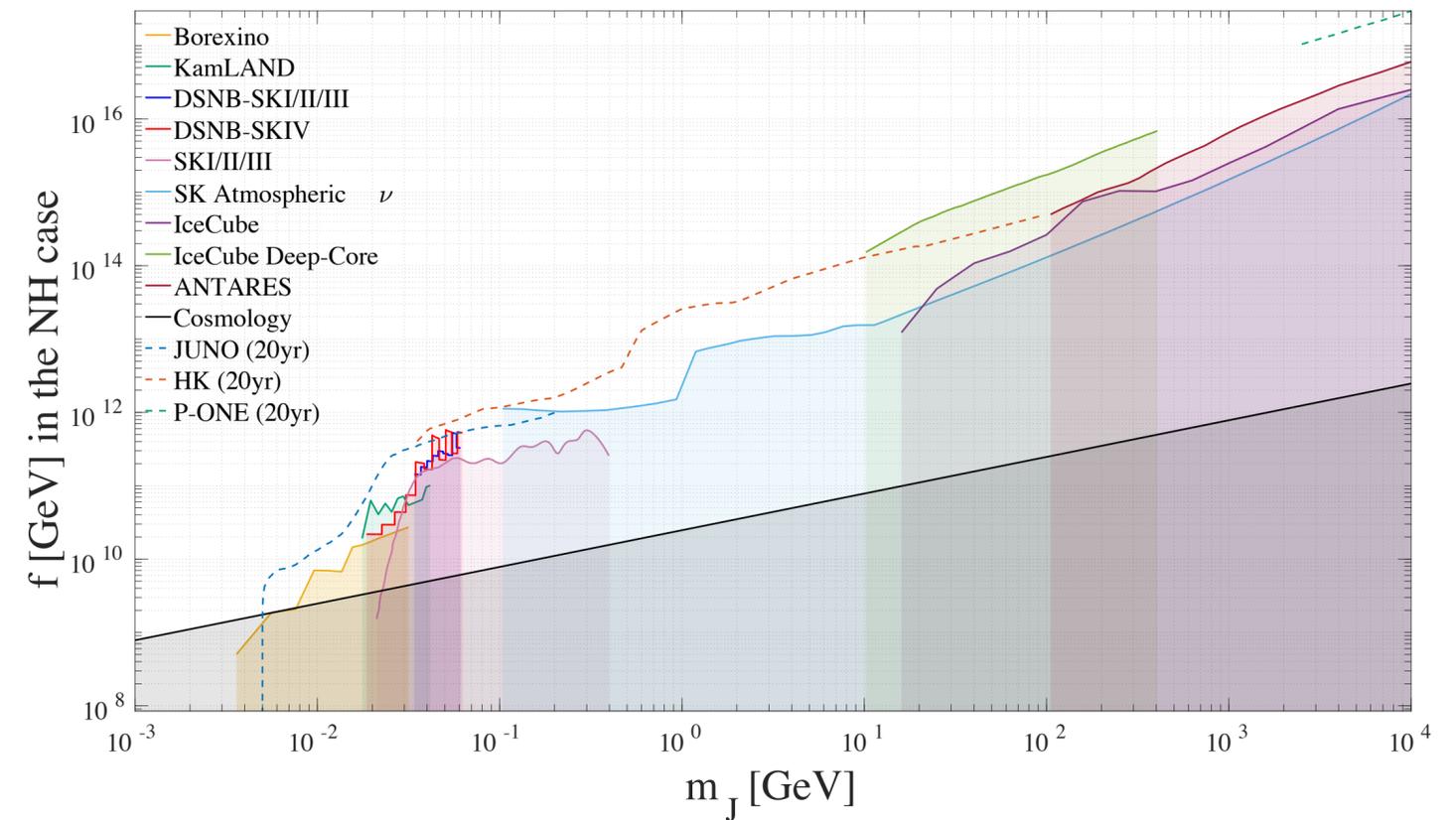
$$m_a^2 = ?^2$$

## Cosmological probes



Sandner, Escudero, Witte, [2305.01692]

## Terrestrial Experiments



Akita, Niibo, [2304.04430]

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- Gravity effects - explicit breaking in the potential:

**Explicit breaking ↔ massive Majoron**

$$V_{\text{grav.}} = g_{2m+n} \frac{|\phi|^{2m} \phi^n}{M_{\text{Pl}}^{2m+n-4}} + \text{h.c.}$$

Alonso, Urbano, [[1706.07415](#)]

Akhmedov, Berezhiani, Mohapatra, Senjanović [[hep-ph/9209285](#)]

Babu, Rothstein, Seckel [[hep-ph/9301213](#)]

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$$m_a^2 \sim \frac{1}{8\pi^2} \frac{\Lambda_{11}\Lambda_{22}\Lambda_{12}\Lambda_{21}}{f_a^2} \log \frac{f_a^2}{\mu^2}$$

$$m_a^2 \sim \frac{v^2}{8\pi^2} \frac{\Lambda_{ij}\Lambda_{kl}(\hat{Y}_N)_{\alpha 1}(\hat{Y}_N)_{\alpha 2}}{f_a^2} \log \frac{f_a^2}{\mu^2}$$

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Frigerio, Hambye, Massó [[1107.4564](#)]

# Neutrino Mass Models 101

$$-\mathcal{L}_{\nu M} \supset \frac{1}{2} \bar{\chi}_L \mathcal{M}_\chi \chi_L^c \quad \chi_L = \begin{pmatrix} \nu_L \\ N_R^c \\ S_R^c \end{pmatrix}$$

# Neutrino Mass Models 101

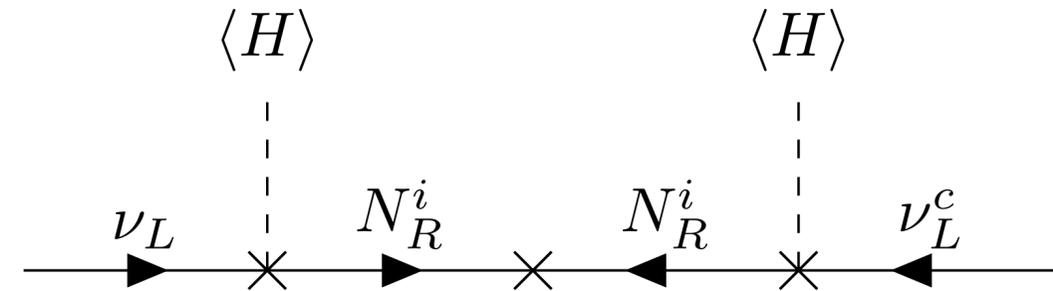
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Type-I Seesaw

$$\mathcal{M}_\chi^{\text{Type-I}} = \begin{pmatrix} 0 & m_N & m_S \\ m_N^T & \Lambda_{NN} & \Lambda_{NS} \\ m_S^T & \Lambda_{NS} & \Lambda_{SS} \end{pmatrix}$$

$$m_\nu^{\text{type-I}} \simeq -m_i \hat{\Lambda}_{ij}^{-1} m_j^T$$

$$\Lambda \sim 10^{14} \text{ GeV}$$



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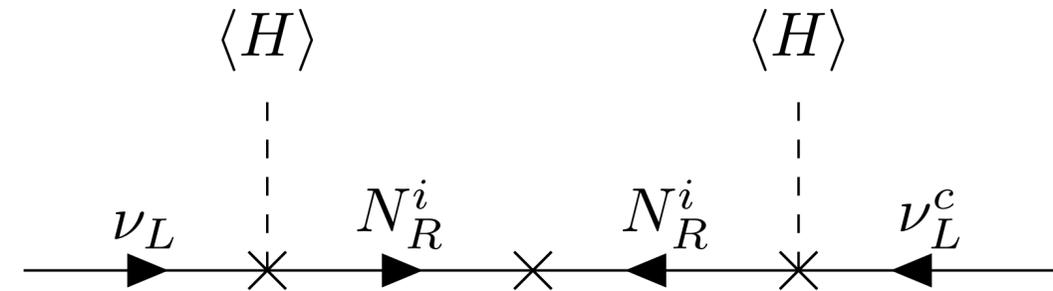
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$m_S, \Lambda_{NN}, \Lambda_{SS}$   
break L

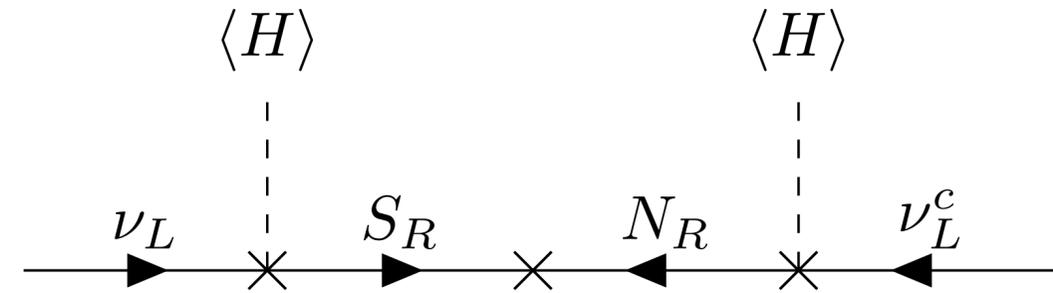
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Type-I Seesaw

Linear Seesaw

$$\mathcal{M}_\chi^{\text{Type-I}} = \begin{pmatrix} 0 & m_N & m_S \\ m_N^T & \Lambda_{NN} & \Lambda_{NS} \\ m_S^T & \Lambda_{NS} & \Lambda_{SS} \end{pmatrix}$$

$$\mathcal{M}_\chi^{\text{LSS}} = \begin{pmatrix} 0 & m_N & \epsilon m_S \\ m_N^T & 0 & \Lambda_{NS} \\ \epsilon m_S^T & \Lambda_{NS} & 0 \end{pmatrix}$$

$m_S, \Lambda_{NN}, \Lambda_{SS}$   
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$$L(\ell_L) = L(N_R) = -L(S_R)$$

$$m_\nu^{\text{type-I}} \simeq -m_i \hat{\Lambda}_{ij}^{-1} m_j^T$$

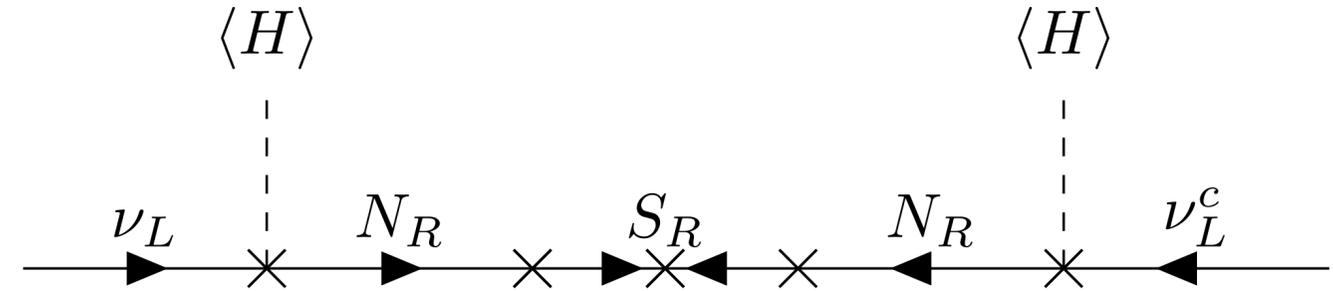
$$\Lambda \sim 10^{14} \text{ GeV}$$

$$m_\nu^{\text{LSS}} \simeq -\epsilon \frac{m_S m_N^T + m_N m_S^T}{\Lambda_{NS}}$$

$$\epsilon m_S \sim 10 \text{ eV}, \quad \Lambda_{NS} \sim 1 \text{ TeV}$$

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Type-I Seesaw

Linear Seesaw

Inverse Seesaw

$$\mathcal{M}_\chi^{\text{Type-I}} = \begin{pmatrix} 0 & m_N & m_S \\ m_N^T & \Lambda_{NN} & \Lambda_{NS} \\ m_S^T & \Lambda_{NS} & \Lambda_{SS} \end{pmatrix}$$

$$\mathcal{M}_\chi^{\text{LSS}} = \begin{pmatrix} 0 & m_N & \epsilon m_S \\ m_N^T & 0 & \Lambda_{NS} \\ \epsilon m_S^T & \Lambda_{NS} & 0 \end{pmatrix}$$

$$\mathcal{M}_\chi^{\text{ISS}} = \begin{pmatrix} 0 & m_N & 0 \\ m_N^T & 0 & \Lambda_{NS} \\ 0 & \Lambda_{NS} & \mu \end{pmatrix}$$

$m_S, \Lambda_{NN}, \Lambda_{SS}$   
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$$\epsilon m_S \sim 10 \text{ eV}, \quad \Lambda_{NS} \sim 1 \text{ TeV}$$

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Remember  
Need an explicit breaking to  
generate a Majoron mass

Type-I Seesaw +  $\phi$

Linear Seesaw +  $\phi$

Inverse Seesaw +  $\phi$

$$\mathcal{M}_\chi^{\text{Type-I}} = \begin{pmatrix} 0 & m_N & m_S \\ m_N^T & Y_{NN}\phi & Y_{NS}\phi \\ m_S^T & Y_{NS}\phi & Y_{SS}\phi \end{pmatrix}$$

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$$PQ(\ell_L) = PQ(N_R) = PQ(S_R) = -\frac{1}{2} PQ(\phi)$$

$\epsilon m_S$   
 $\mu \equiv \epsilon Y_{SS} \langle \phi \rangle$  No longer breaking terms!

# Majoron Neutrino Mass Models

**The list of models:** We identify all the possible non-redundant PQ symmetries

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Up to trivial exchanges:  $\phi \longleftrightarrow \phi^*$   $N_R \longleftrightarrow S_R$

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Two different scales:  $f_a, \Lambda$

Extra requirements to fix the two scales

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Extra requirements to fix the two scales  $m_\nu \simeq -\mu \frac{m_N m_N^T}{\Lambda_{NS}^2} \sim Y_{SS} \frac{m_N m_N^T}{Y_{NS}^2 f_a}$

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$m_N m_N^T$  is a rank-1 matrix, it can only produce 1 massive light neutrino

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$$-\mathcal{L}_{\text{PQ}}^{\text{mmM}} = \bar{\ell}_L \tilde{H} Y_N N_R + \frac{Y_{NS}}{2} \phi^* (\overline{N_R^c} S_R + \overline{S_R^c} N_R) + \frac{Y_{NN}}{2} \phi \overline{N_R^c} N_R + \text{h.c.} \quad \mathcal{M}_\chi = \begin{pmatrix} 0 & m_N & 0 \\ m_N^T & Y_{NN}\phi & Y_{NS}\phi^* \\ 0 & Y_{NS}\phi^* & 0 \end{pmatrix},$$

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 -\mathcal{L}_{\text{PQ}}^{\text{mmM}} &= \bar{\ell}_L \tilde{H} Y_N N_R + \frac{Y_{NS}}{2} \phi^* (\overline{N_R^c} S_R + \overline{S_R^c} N_R) + \frac{Y_{NN}}{2} \phi \overline{N_R^c} N_R + \text{h.c.} \\
 -\mathcal{L}_{\epsilon\text{PQ}}^{\text{mmM}} &= \epsilon \bar{\ell}_L \tilde{H} Y_S S_R + \text{h.c.},
 \end{aligned}
 \quad \mathcal{M}_\chi = \begin{pmatrix} 0 & m_N & \epsilon m_S \\ m_N^T & Y_{NN} \phi & Y_{NS} \phi^* \\ \epsilon m_S^T & Y_{NS} \phi^* & 0 \end{pmatrix}$$

Minimality:

i) Minimal number of fields to account for  
the correct neutrino spectrum  
 $N_R, S_R + \phi$

Light neutrino masses can be correctly generated

$$m_\nu^{\text{TL}} = -\epsilon \frac{m_S m_N^T + m_N m_S^T}{\Lambda_{NS}} + \epsilon^2 \frac{\Lambda_{NN}}{\Lambda_{NS}} \frac{m_S m_S^T}{\Lambda_{NS}},$$

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 Other breaking-effects are assumed to be subleading

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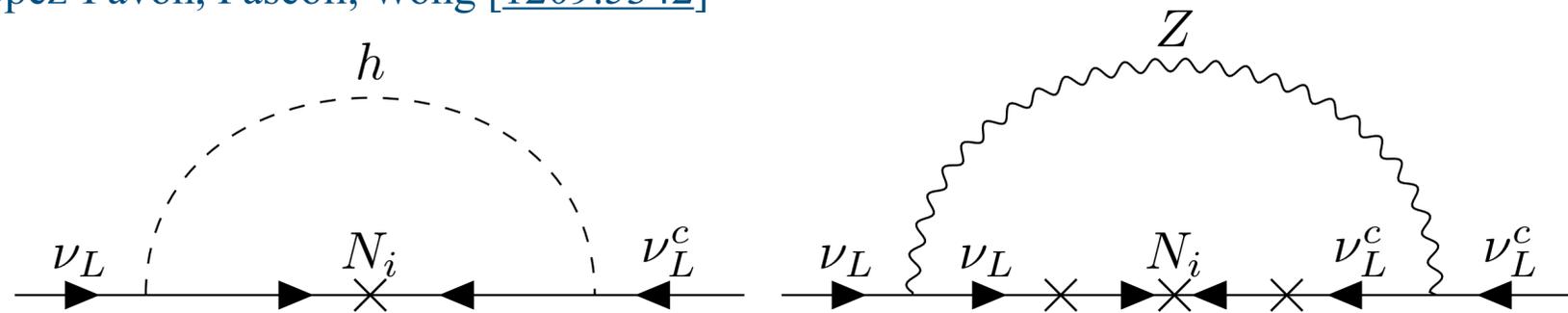
$$\epsilon \ll Y_i \sim [10^{-2}, 1]$$

iii) Only one scale  $\Lambda \sim f_a = \langle \phi \rangle$

$$\Lambda_{NN} \equiv \frac{f_a}{\sqrt{2}} Y_{NN}, \quad \Lambda_{NS} \equiv \frac{f_a}{\sqrt{2}} Y_{NS}$$

# Radiative Neutrino Masses

López-Pavón, Pascoli, Wong [1209.5342]

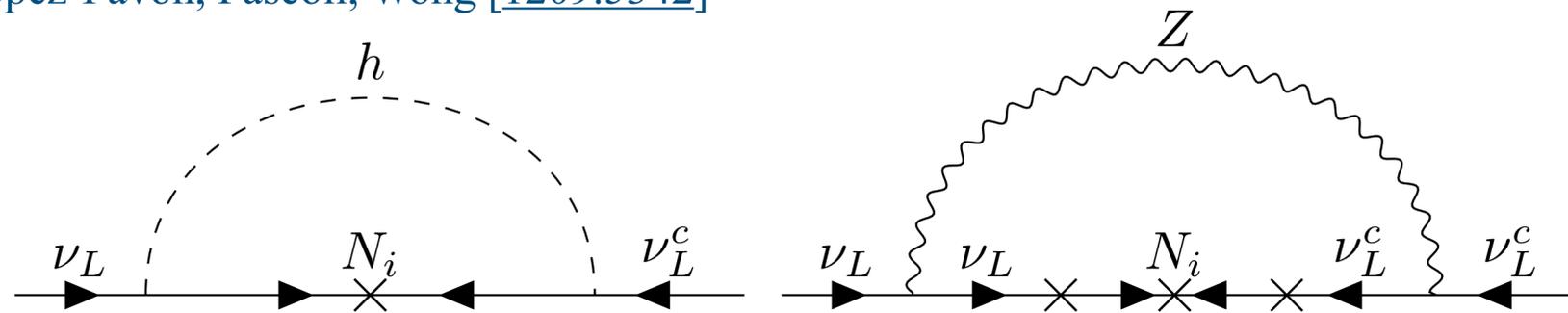


Two massive RH neutrinos

$$M_{N,S} = \frac{\Lambda_{NS}}{2} \left[ \sqrt{4 + \left( \frac{\Lambda_{NN}}{\Lambda_{NS}} \right)^2} \mp \left( \frac{\Lambda_{NN}}{\Lambda_{NS}} \right) \right]$$

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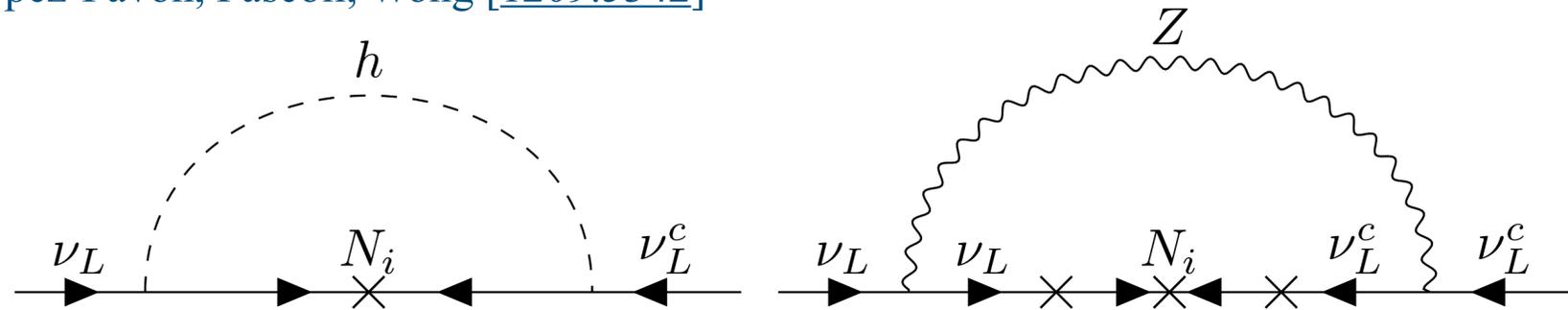
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*1-loop contribution*

Need to tame the loop contribution!

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López-Pavón, Pascoli, Wong [[1209.5342](#)]

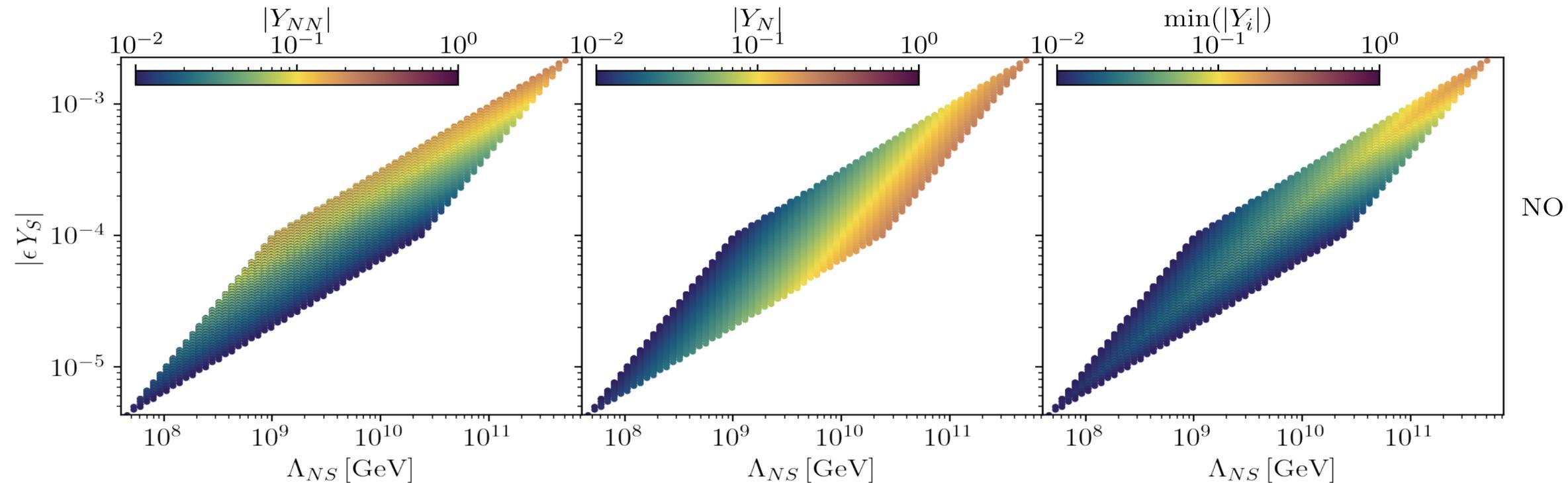


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# The Minimal Massive Majoron

Using the Coleman-Weinberg potential:

$$V_{\text{CW}} = -\frac{1}{2} \times \frac{1}{16\pi^2} \text{Tr} \left[ (\mathcal{M}_\chi \mathcal{M}_\chi^\dagger)^2 \left( \log \left( \frac{\mathcal{M}_\chi \mathcal{M}_\chi^\dagger}{\mu_R^2} \right) - \frac{3}{2} \right) \right] \quad \text{with } \mathcal{M}_\chi = \mathcal{M}_\chi(\phi, H)$$

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$$V_{\text{CW}} = -\epsilon \frac{Y_{NS} Y_{NN} |Y_N \cdot Y_S| v^2 f_a^2}{16\pi^2} \cos \left( \vartheta_\eta + \frac{2a}{f_a} \right) \times \quad \vartheta_\eta \equiv \arg(Y_N^\dagger Y_S)$$
$$\times \left[ \frac{(Y_{NN}^2 + 2Y_{NS}^2)}{Y_{NN} \sqrt{Y_{NN}^2 + 4Y_{NS}^2}} \text{Arcoth} \left( \frac{(Y_{NN}^2 + 2Y_{NS}^2)}{Y_{NN} \sqrt{Y_{NN}^2 + 4Y_{NS}^2}} \right) + \left( \log \left( \frac{Y_{NS}^2 f_a^2}{2\mu_R^2} \right) - 1 \right) \right]$$

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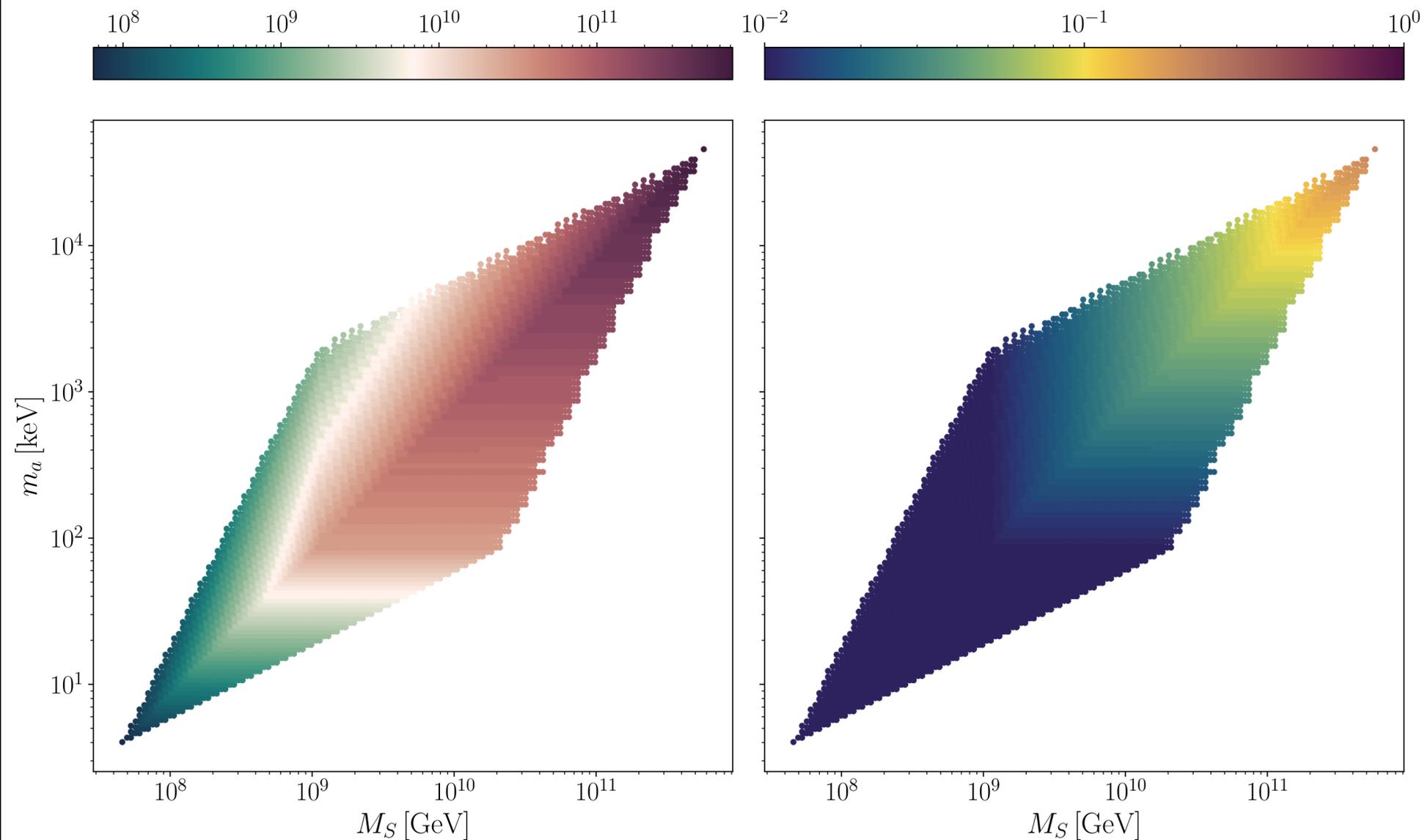
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$$m_a^2 = \frac{|\epsilon m_S \cdot m_N|}{\pi^2} \frac{\sqrt{M_N M_S}}{M_N + M_S} \left[ \frac{(M_S^2 + M_N^2)}{f_a^2} \log \left( \frac{M_S}{M_N} \right) + \frac{(M_S^2 - M_N^2)}{f_a^2} \left( \log \left( \frac{M_N M_S}{\mu_R^2} \right) - 1 \right) \right]$$

$$m_a^2 \sim \epsilon \frac{v^2}{\pi^2} \ll f_a^2$$

# The Minimal Massive Majoron

$$m_a^2 \simeq \frac{\sqrt{|\eta|} |\Delta m_{32}^2|}{2\pi^2} \frac{M_N M_S}{M_N + M_S} \left[ \frac{(M_S^2 + M_N^2)}{f_a^2} \log\left(\frac{M_S}{M_N}\right) + \frac{(M_S^2 - M_N^2)}{f_a^2} \left( \log\left(\frac{M_N M_S}{\mu_R^2}\right) - 1 \right) \right]$$



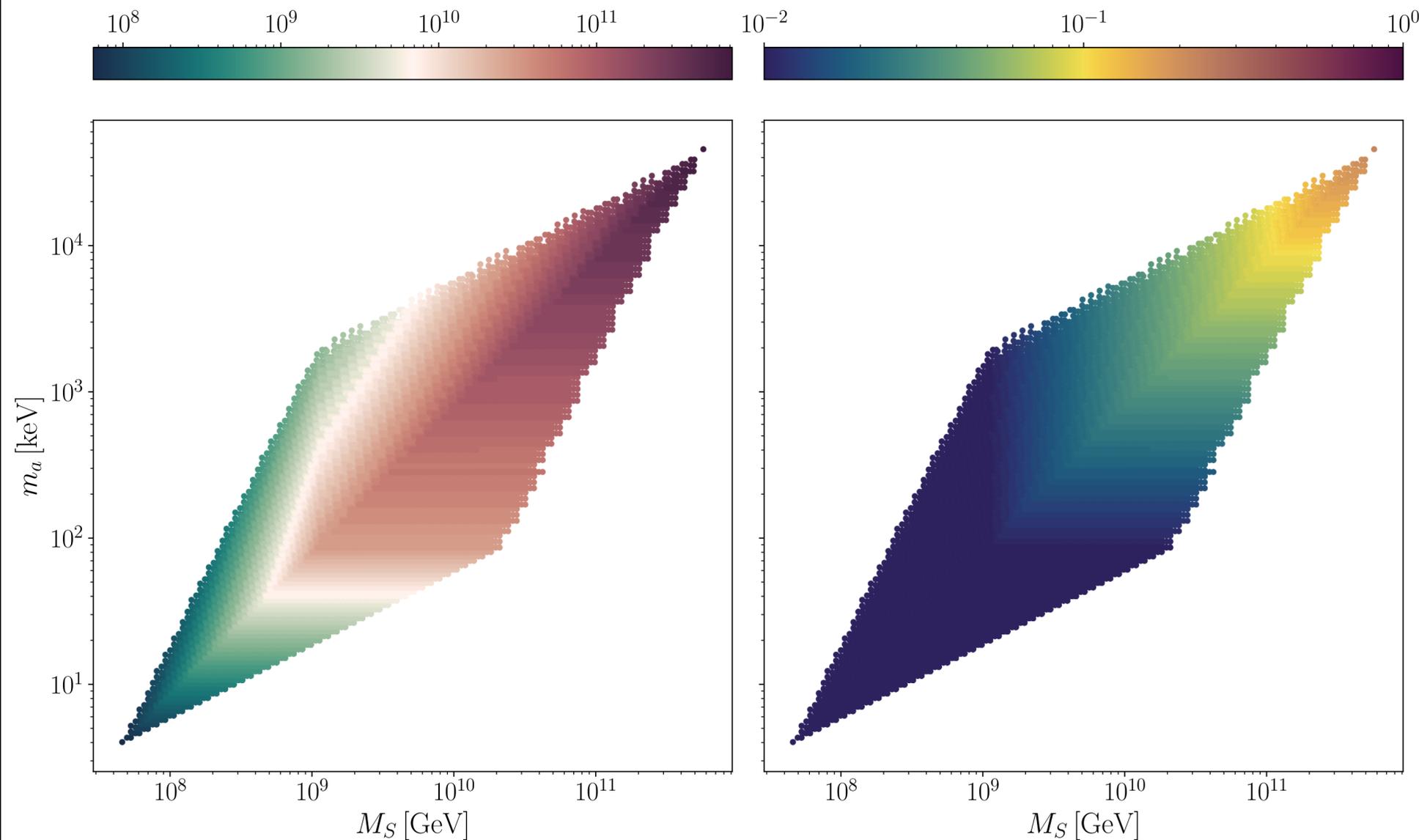
$$|\eta| \equiv \frac{1 - \sqrt{r}}{1 + \sqrt{r}}, \quad \text{for the NO}$$

$$r \equiv \frac{|\Delta m_{\text{sol.}}^2|}{|\Delta m_{\text{atm.}}^2|}$$

Gavela, Hambie, Hernández, Hernández [0906.2461]

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The Minimal Massive Majoron can be found

$$f_a \sim [10^8, 10^{11}] \text{ GeV}$$

$$m_a \sim [1 \text{ keV}, 10 \text{ MeV}]$$

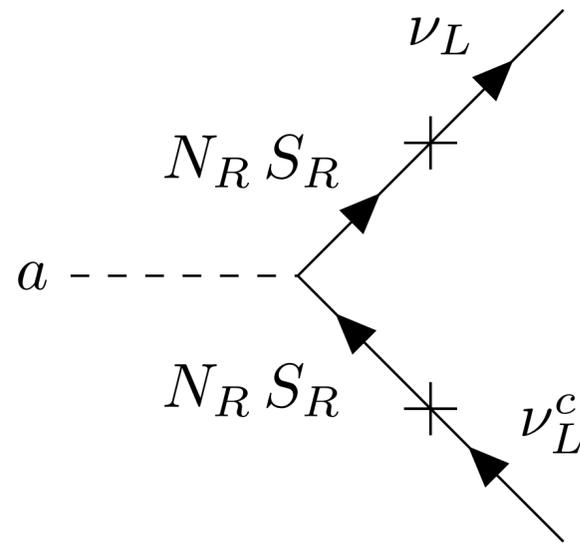
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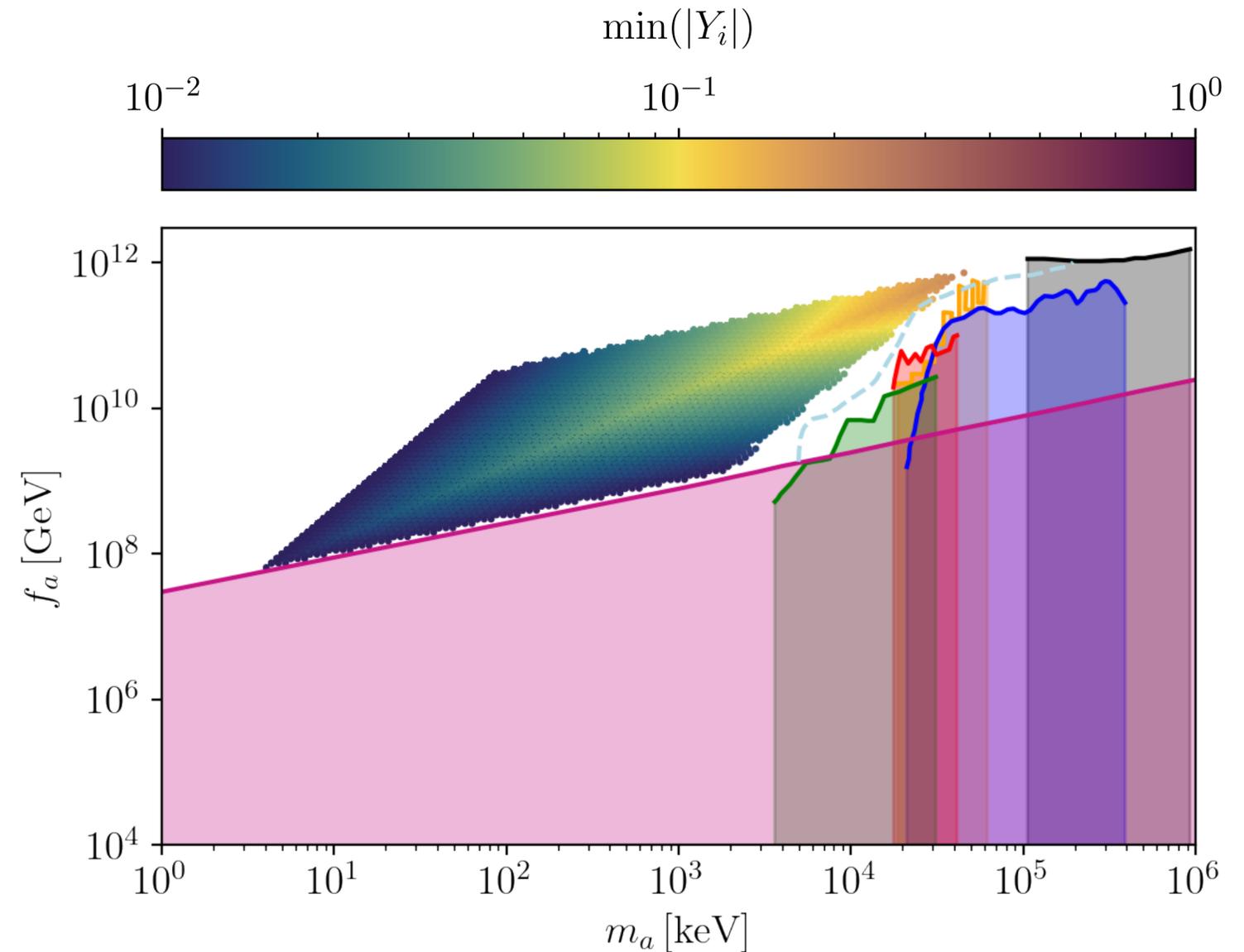
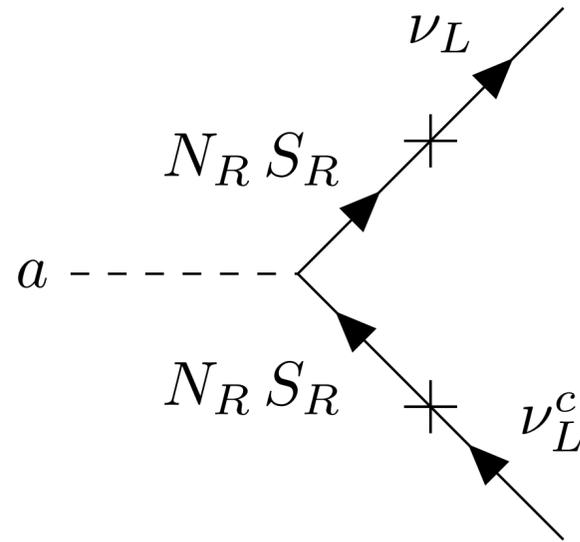
# Phenomenology of the Majoron

$$\mathcal{L}_{a\nu\nu} \simeq -\frac{ia}{2f_a} \bar{\nu}_L m_\nu \gamma_5 \nu_L^c$$



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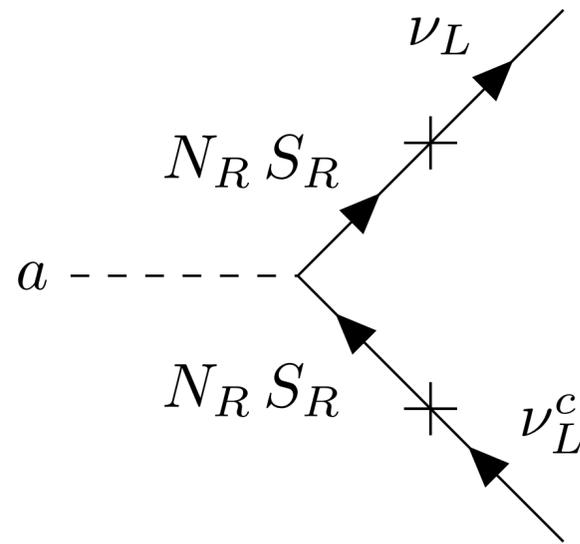
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Akita, Niibo, [2304.04430]

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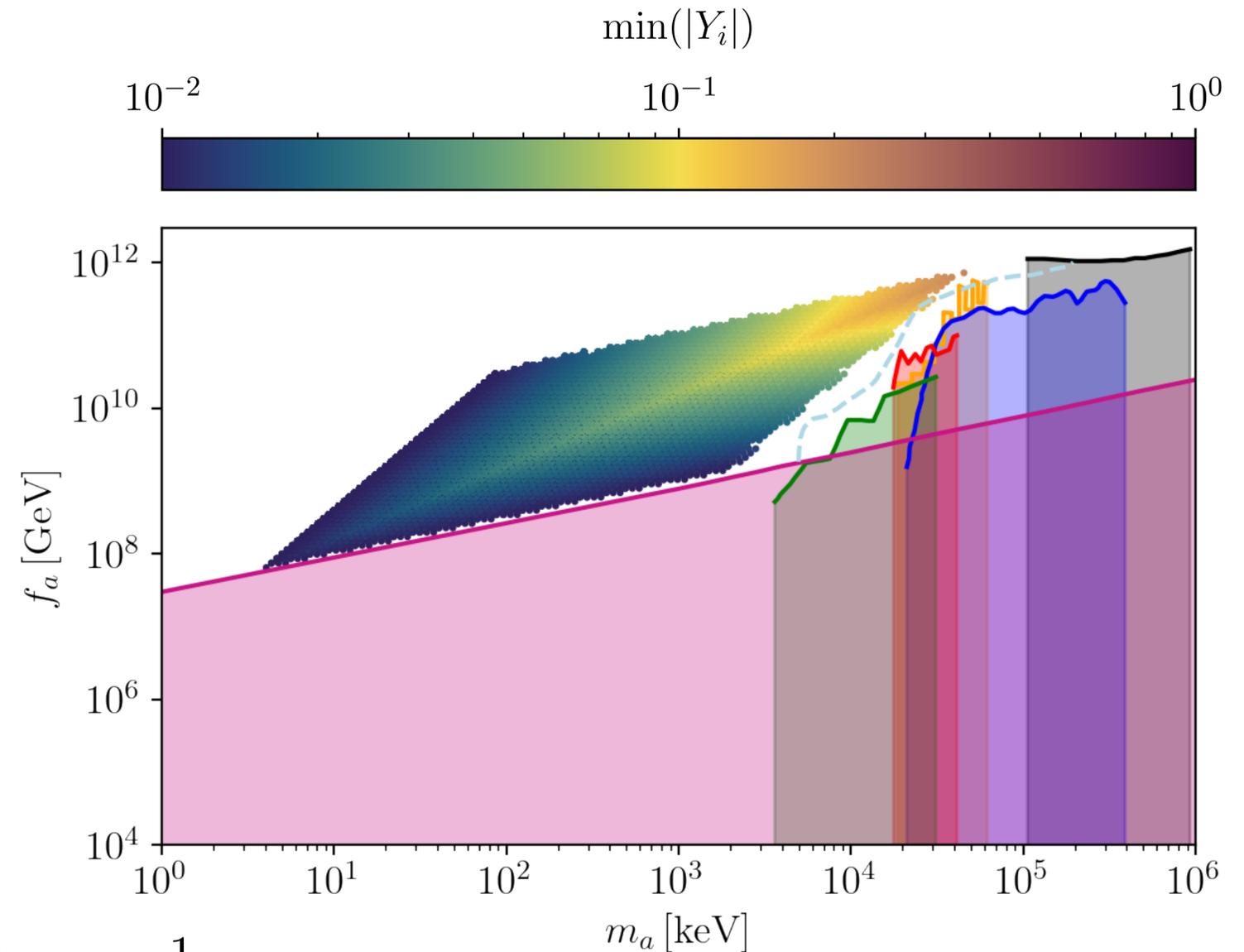
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$$\Gamma(a \rightarrow 2\nu) \simeq \frac{m_a}{16\pi f_a^2} \sum_{i=1}^3 m_i^2 \sim$$

$$\frac{1}{3 \times 10^{19} \text{ sec}} \left( \frac{m_a}{1 \text{ MeV}} \right) \left( \frac{10^9 \text{ GeV}}{f_a} \right)^2 \left( \frac{\sum m_i^2}{10^{-3} \text{ eV}^2} \right) \lesssim \frac{1}{10^{19} \text{ sec}}$$

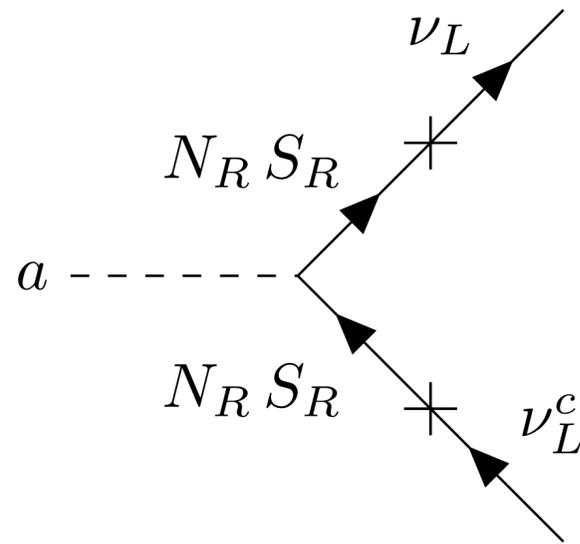
— From CMB +BAO



Akita, Niibo, [2304.04430]

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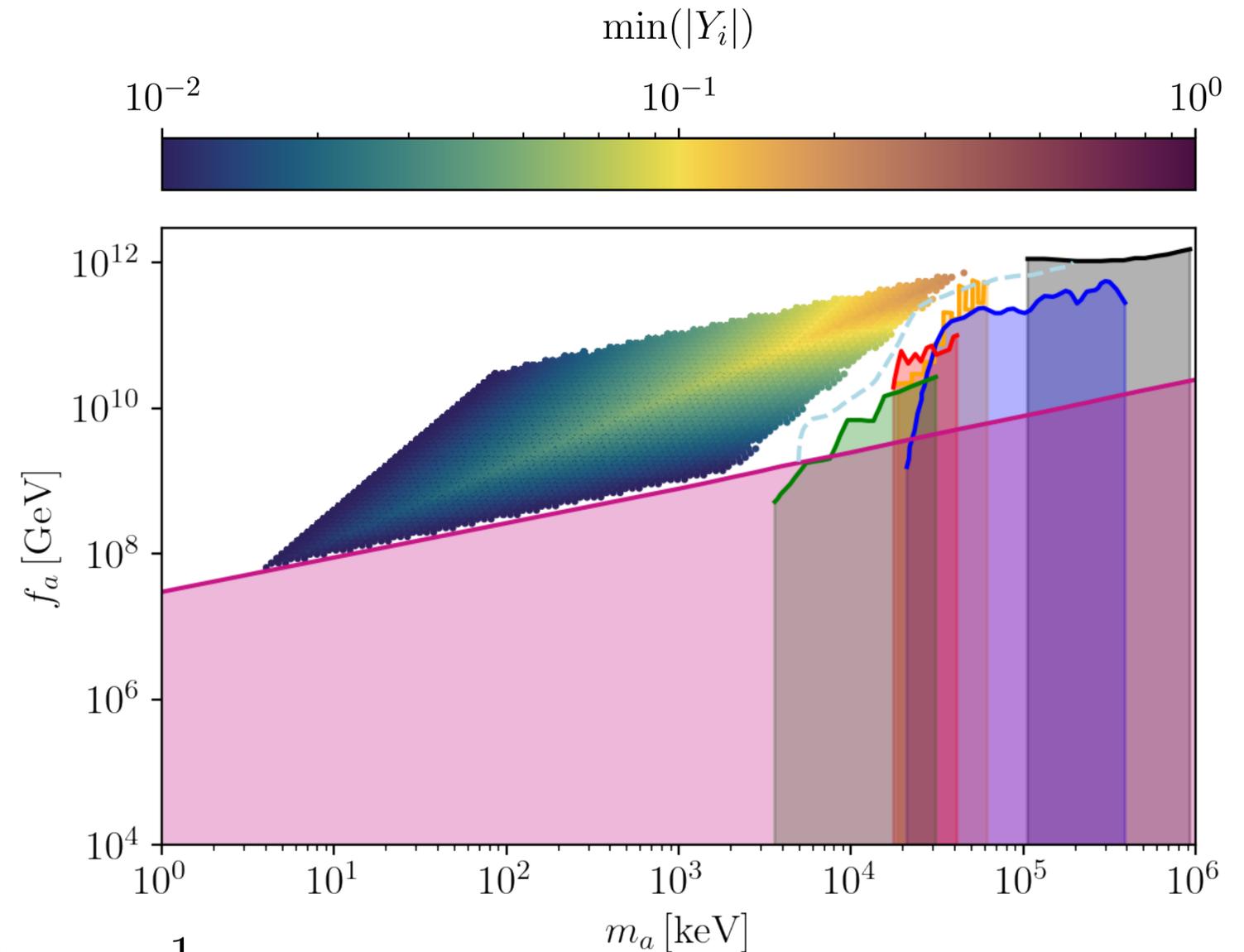
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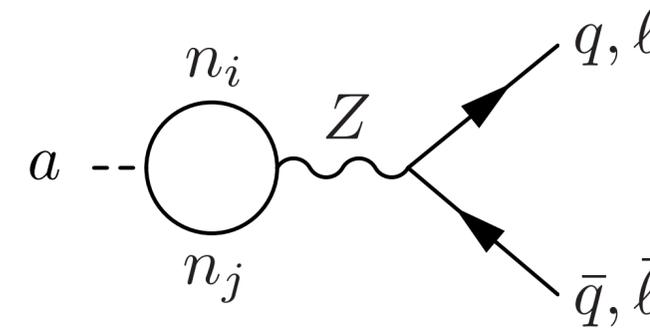
- Borexino
- Kamland
- SK
- SK Reinterpreted
- Atmospheric Neutrinos
- Juno (20yr)

Akita, Niibo, [2304.04430]

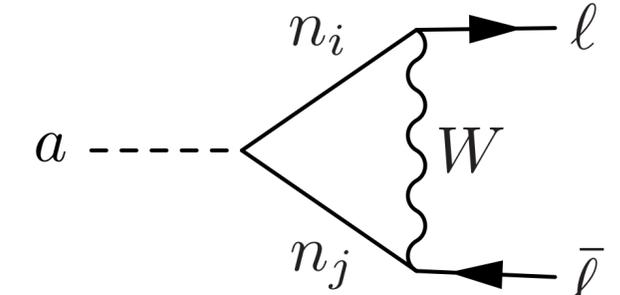
# Phenomenology of the Majoron

$$\mathcal{L}_{all} \simeq \frac{i a}{16\pi^2 f_a} \bar{\ell} (M_\ell \text{tr} [K] \gamma_5 + 2 M_\ell K P_L - 2 K M_\ell P_R) \ell$$

$$K \equiv \frac{\hat{m}\hat{m}^\dagger}{v^2} = \frac{m_N m_N^\dagger + \epsilon^2 m_S m_S^\dagger}{v^2}$$



Heeck, Patel [1909.02029]

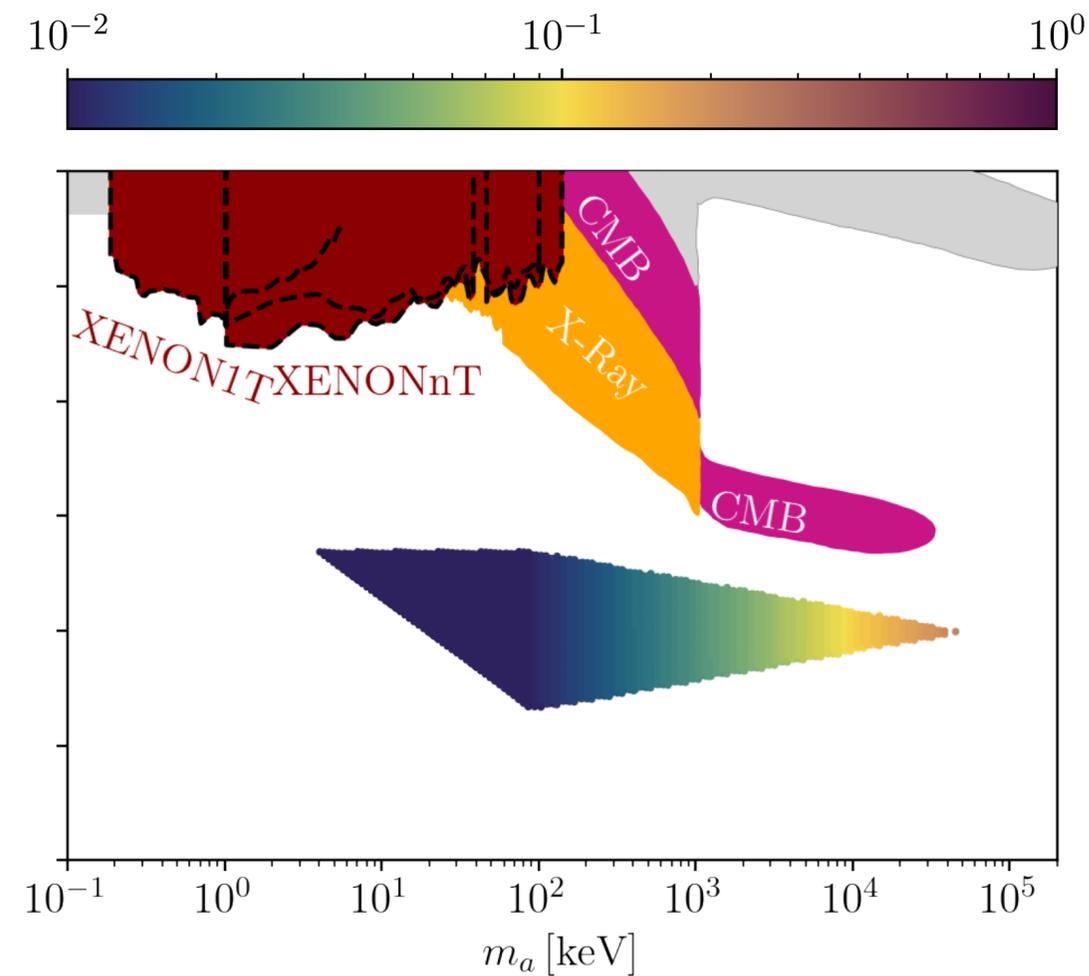
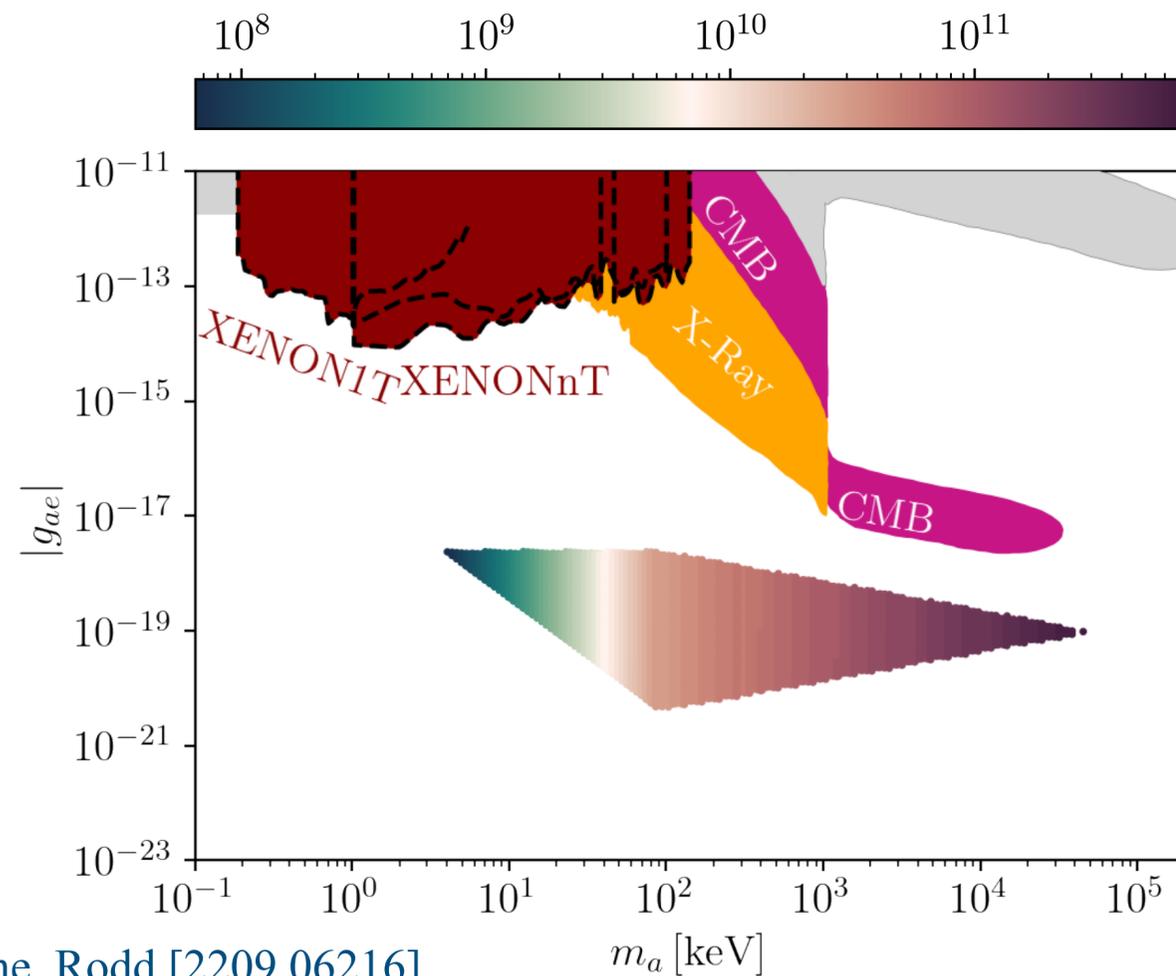
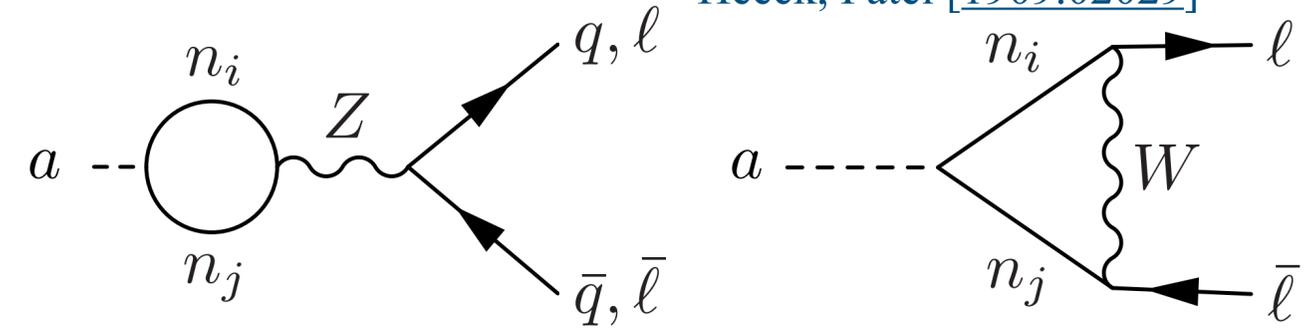


# Phenomenology of the Majoron

$$\mathcal{L}_{aee} \simeq i \frac{m_e}{16\pi^2 f_a} (\text{tr}K - 2(K)_{ee}) a \bar{e} \gamma_5 e \equiv i g_{ae} a \bar{e} \gamma_5 e$$

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Heeck, Patel [1909.02029]



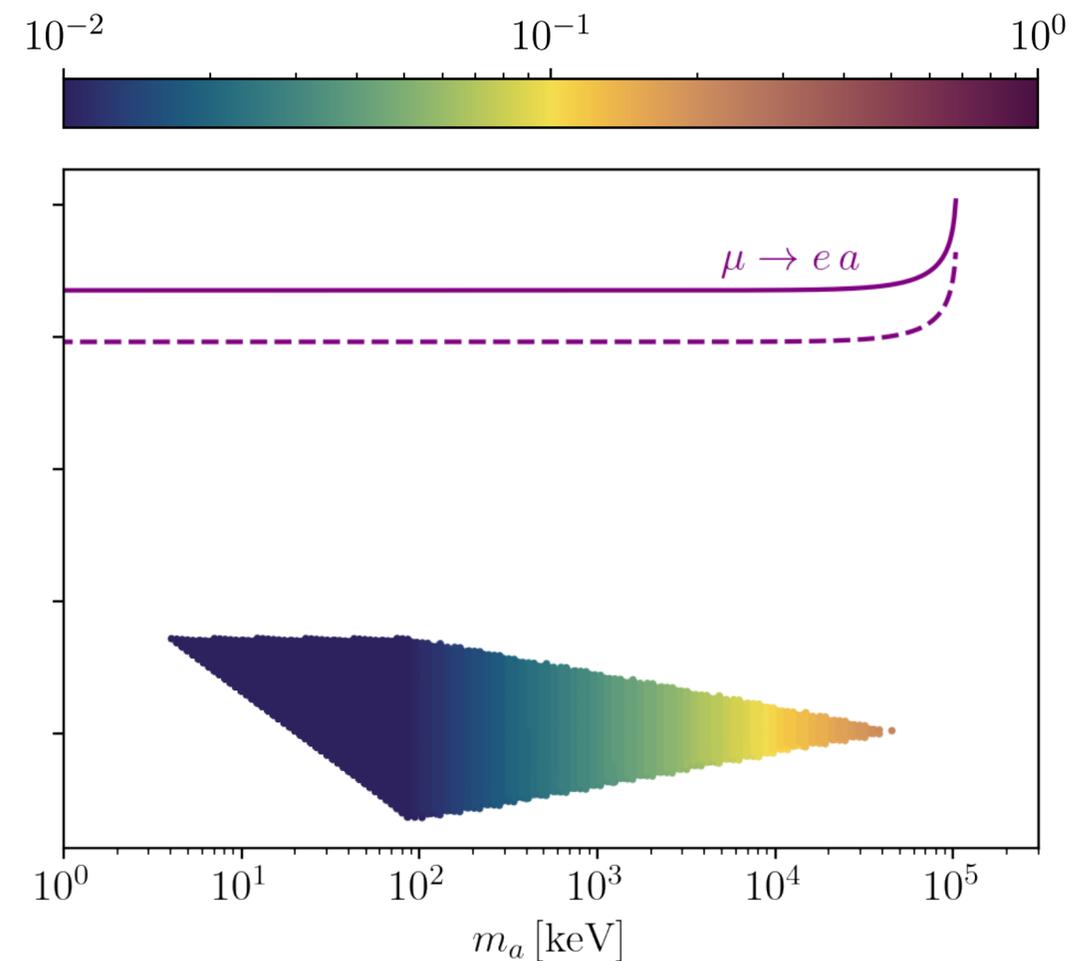
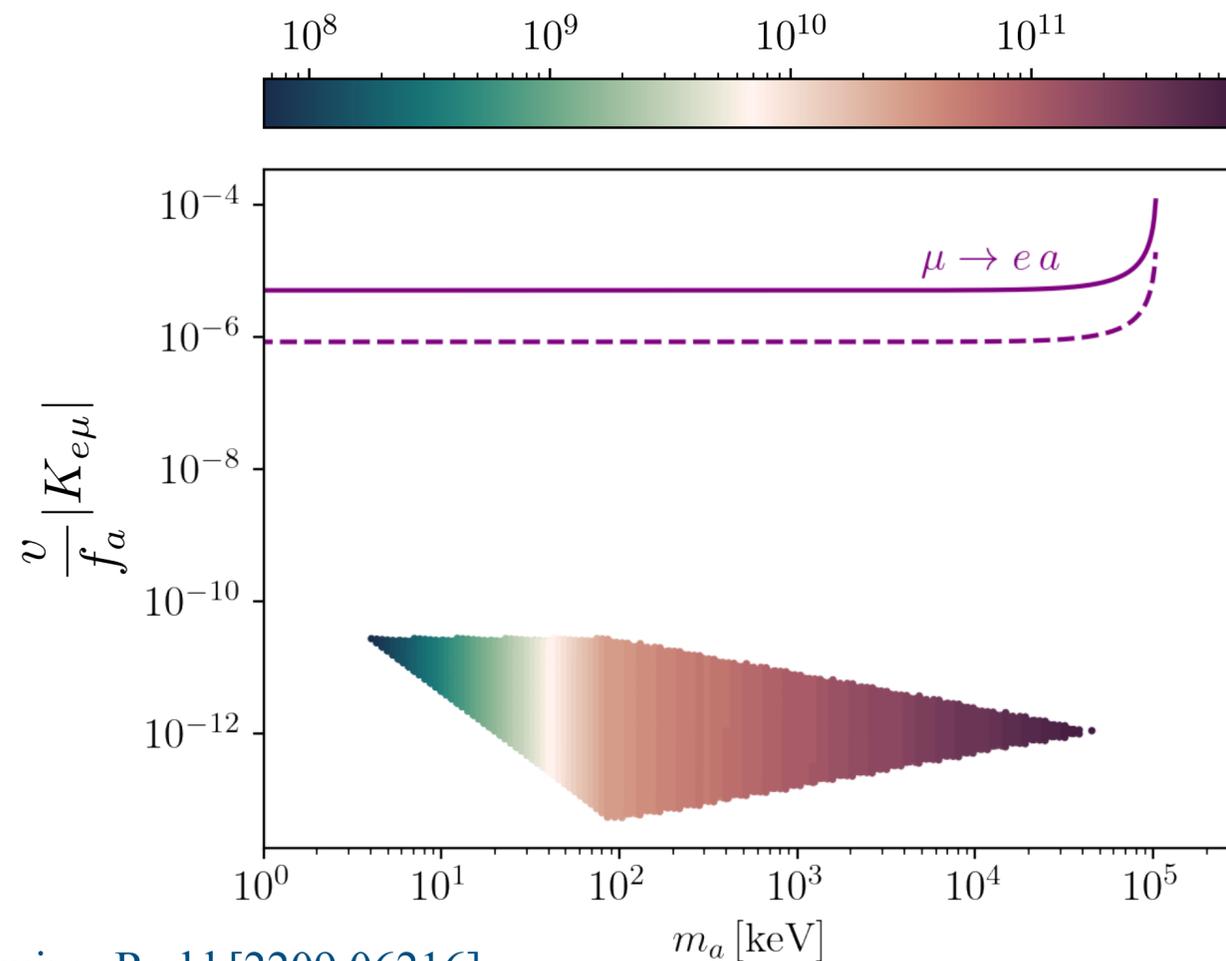
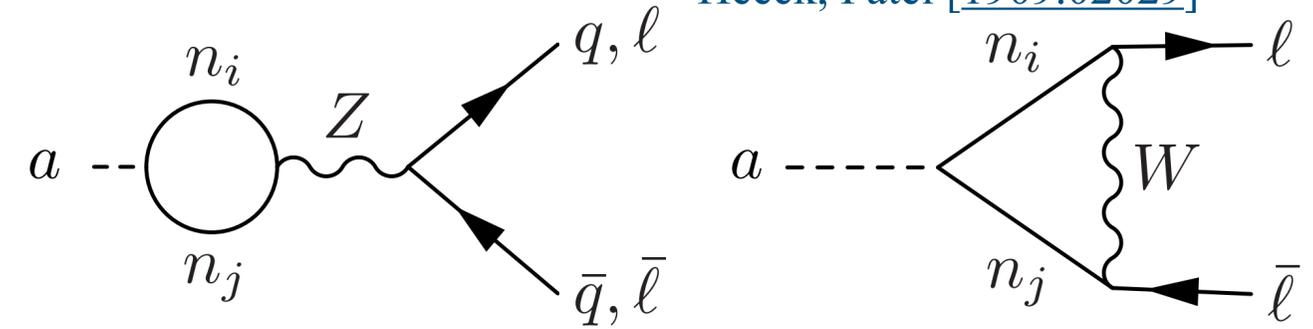
Langhoff, Outmezguine, Rodd [2209.06216]

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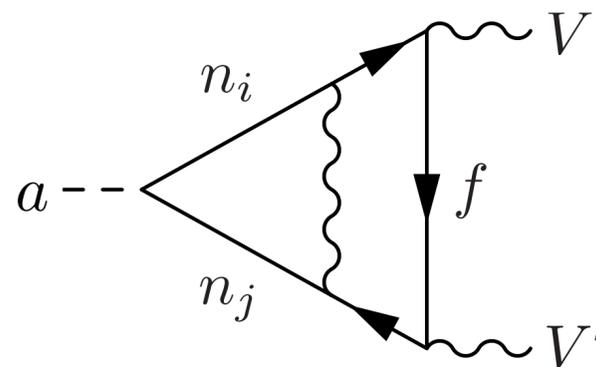
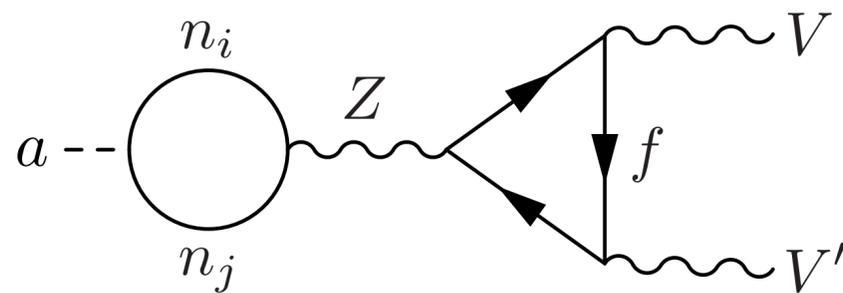


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Heeck, Patel [1909.02029]

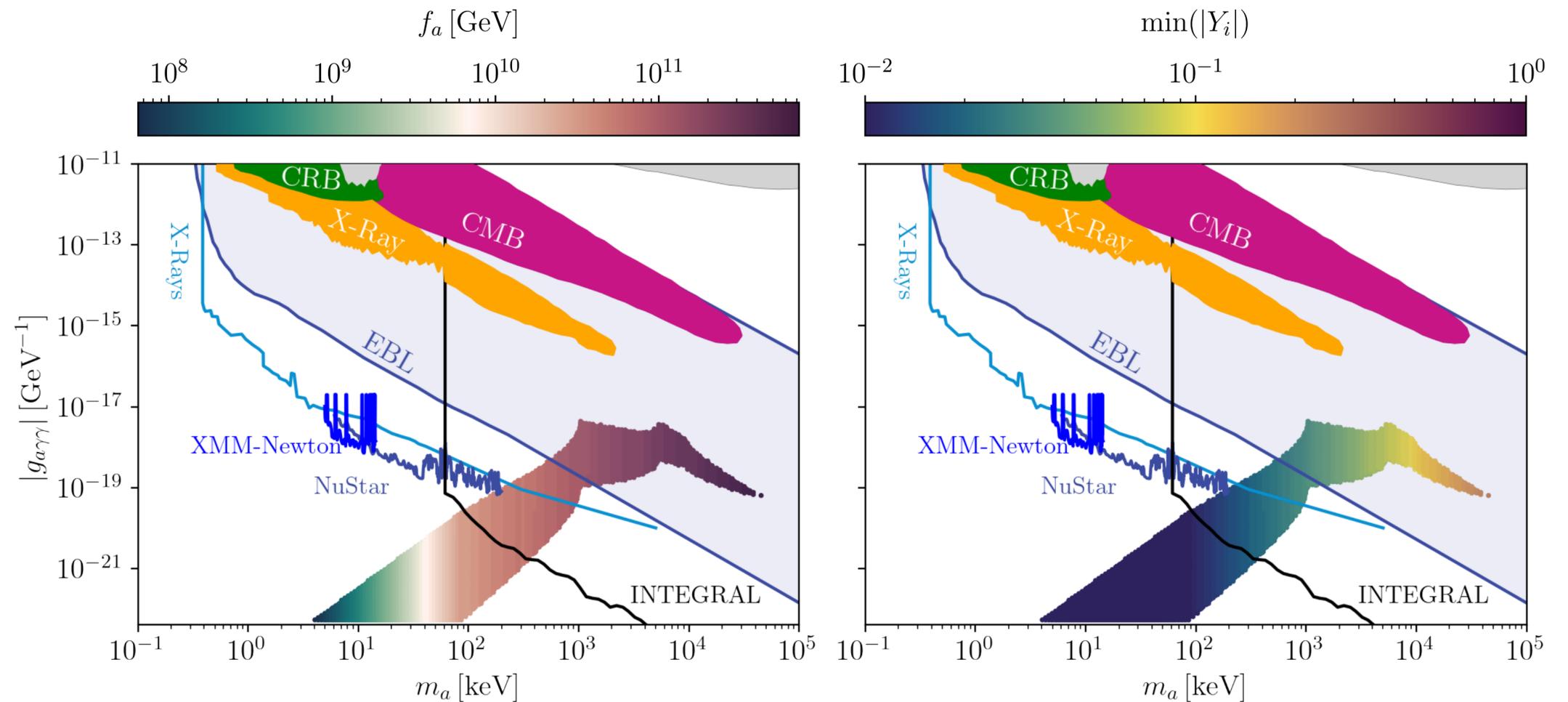
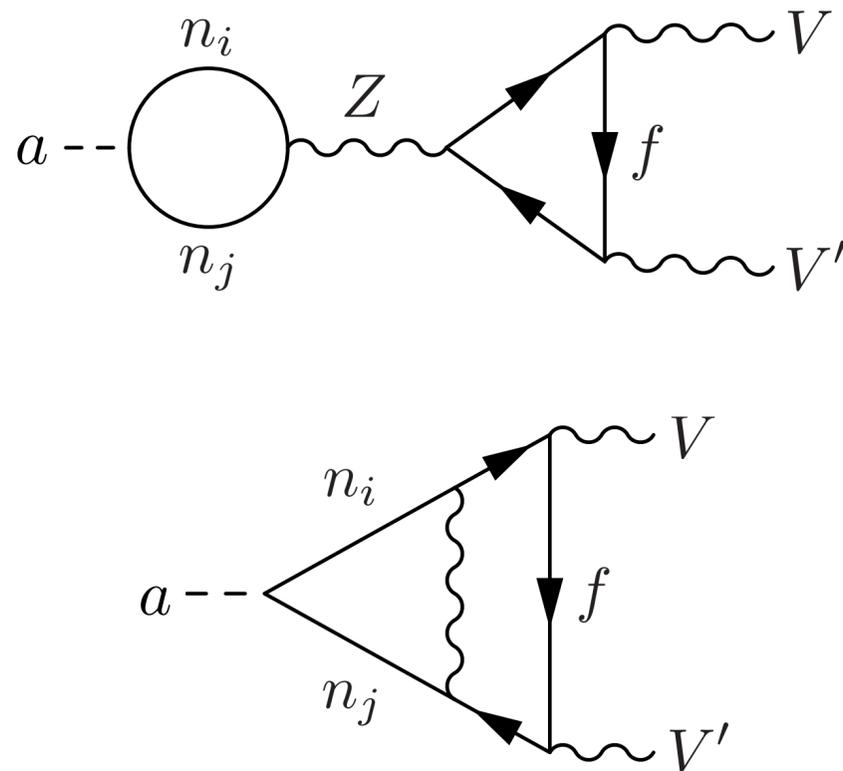
$$\mathcal{L}_{a\gamma\gamma} = -\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu} \quad g_{a\gamma\gamma} \simeq -\frac{\alpha_{\text{em}}}{8\pi^3 f_a} \left[ \text{tr} K \sum_f N_c^f Q_f^2 T_3^f h \left( \frac{m_a^2}{4m_f^2} \right) + \sum_{\ell=e,\mu,\tau} K_{\ell\ell} h \left( \frac{m_a^2}{4m_\ell^2} \right) \right]$$



# Phenomenology of the Majoron

Heeck, Patel [1909.02029]

$$\mathcal{L}_{a\gamma\gamma} = -\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu} \quad g_{a\gamma\gamma} \simeq -\frac{\alpha_{\text{em}}}{8\pi^3 f_a} \left[ \text{tr} K \sum_f N_c^f Q_f^2 T_3^f h \left( \frac{m_a^2}{4m_f^2} \right) + \sum_{\ell=e,\mu,\tau} K_{\ell\ell} h \left( \frac{m_a^2}{4m_\ell^2} \right) \right]$$



Cadamuro, Redondo [1110.2895]

Langhoff, Outmezguine, Rodd [2209.06216]

# Conclusions

- We investigate all possible textures of non-redundant models
- Imposing a set of **minimal** conditions we are able to link together the masses of the Majoron and light neutrinos.
- Obtaining a prediction for the parameter space!
- This model can be tested via Cosmology + neutrino detectors, if this particle is Dark Matter!

$$f_a \sim [10^8, 10^{11}] \text{ GeV}$$

$$m_a \sim [1 \text{ keV}, 10 \text{ MeV}]$$

Thanks for your attention!

# Acknowledgements



This project has received funding from the European Union's **Horizon 2020** research and innovation programme under the **Marie Skłodowska-Curie** grant agreement **No 860881**.

**Back up**

# Light Neutrino Masses

One can factorise the mass-scale contributions from the flavour structure:

$$m_\nu = m_\nu^{\text{TL}} + \delta m_\nu^{1L} \equiv m_{T_1} (\mathbf{u}\mathbf{v}^T + \mathbf{v}\mathbf{u}^T) + m_{T_2} \mathbf{v}\mathbf{v}^T + m_L \mathbf{u}\mathbf{u}^T$$

When  $m_{T_1} \gg m_{T_2}$ ,  $\delta m_\nu^{1L}$   
Is fixed by the PMNS

$$m_{T_1} = -\epsilon \frac{|Y_N| |Y_S| v_{\text{EW}}^2}{2\Lambda_{NS}}, \quad m_{T_2} = \epsilon^2 \frac{\Lambda_{NN}}{\Lambda_{NS}} \frac{|Y_S|^2 v_{\text{EW}}^2}{2\Lambda_{NS}}, \quad m_L = \frac{|Y_N|^2}{16\pi^2} \frac{M_H^2 + 3M_Z^2}{M_N + M_S} \log \left( \frac{M_S}{M_N} \right).$$

$$|m_\pm|^2 = \frac{1}{2} \left[ m_C^2 - \tau^2 (2m_{T_1}^2 - m_{T_2}^2 - m_L^2) \pm \sqrt{(m_C^2 - \tau^2 (2m_{T_1}^2 - m_{T_2}^2 - m_L^2))^2 - 4\tau^4 (m_{T_1}^2 - m_L m_{T_2})^2} \right],$$

$$m_C \equiv |2m_{T_1} + \eta m_L + \eta^* m_{T_2}|, \quad \eta \equiv \mathbf{u}^\dagger \mathbf{v} \equiv |\eta| e^{i\vartheta_\eta}, \quad \tau^2 \equiv 1 - |\eta|^2.$$

$$\text{NO:} \quad |m_1|^2 = 0, \quad |m_2|^2 = |m_-|^2, \quad |m_3|^2 = |m_+|^2,$$

$$\text{IO:} \quad |m_1|^2 = |m_-|^2, \quad |m_2|^2 = |m_+|^2, \quad |m_3|^2 = 0.$$

# Majoron Mass: Loop Computation



$$\mathcal{L}_{\text{Yuk}} = \bar{\ell}_L H Y_e e_R + \bar{\ell}_L \tilde{H} Y_N N_R + \epsilon \bar{\ell}_L \tilde{H} Y_S S_R + \frac{Y_{NS}}{2} \phi^* (\overline{N_R^c} S_R + \overline{S_R^c} N_R) + \frac{Y_{NN}}{2} \phi \overline{N_R^c} N_R$$

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Perform the field redefinition:  $\{N_R, \ell_L, e_R\} \rightarrow \{N_R, \ell_L, e_R\} e^{-ia/(2f_a)}$ ,  $S_R \rightarrow S_R e^{3ia/(2f_a)}$

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# Majoron Mass: Loop Computation

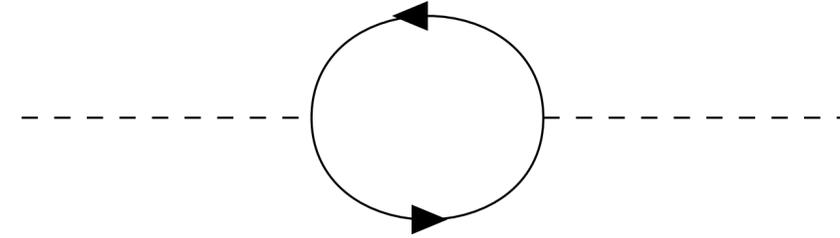
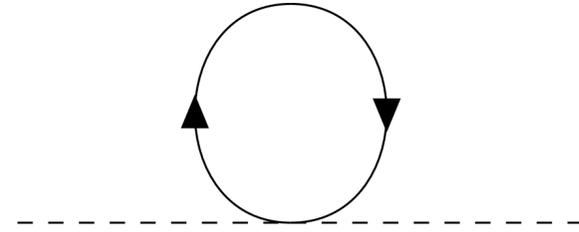


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# Majoron Mass: Loop Computation



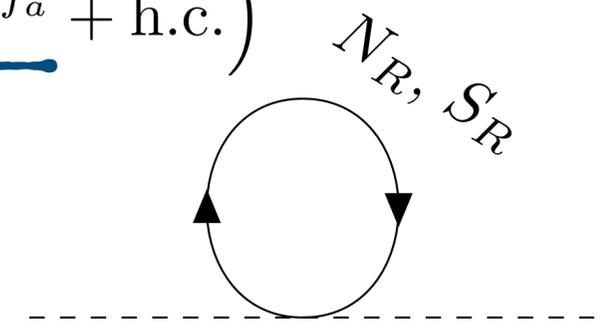
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After diagonalization

$$\mathcal{L}_a \supset \frac{|\epsilon m_S \cdot m_N|}{2\sqrt{M_N M_S}(M_N + M_S)} \left( M_N \bar{S}_R^c S_R - M_S \bar{N}_R^c N_R \right) \frac{a^2}{f_a^2} + \text{h.c.}$$



Problem reduced to only two loops!