Galaxies far, far away...

Exploring the Universe with Weak Gravitational Lensing and Galaxy Clustering

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Northeastern University Dark Energy Survey, Dark Energy Science Collaboration La Thuile, March 2024



Photometric (imaging) surveys



Dark Energy Survey





Geometry and Growth





(Virgo simulations: Jenkins+ 1998)

on 8 Mpc

density fluctuations $\longrightarrow \sigma_8(z)$ $S_8 \equiv \sigma_8 (\Omega_{\rm m}/0.3)^{0.5}$

Concordance Cosmology?



Park & Rozo 2019

Outline

- Cosmology with galaxy surveys: combining weak lensing and galaxy clustering
- Dark Energy Survey (DES, Year 3) and joint analysis with the Kilo Degree Survey (KiDS)
- Modeling galaxy observables
- The future of galaxy surveys

Lensing and large-scale structure



DES Year 1: Elvin-Poole+ 2018; Chang+ 2018

galaxy density

 κ_E ; 0.2 < z < 1.3 lensing shear/convergence

Combining probes



Dark Energy Survey

- DECam (520 Mpix) on 4m Blanco Telescope, Cerro Tololo, Chile
- 1/8 of sky (5000 deg²)
- 6 year mission, 525 nights, completed Jan 2019. Y3 is full area. *Y6 analysis in progress*
- grizY filters (photometric redshifts)
- ~300 million galaxies (0 < z < 2)
 Y3: 100 million with WL shapes
 Y6: deeper and has a second shear measurement method





DES Y3 correlation functions

DES Collaboration Key Paper 2022 (arXiv:2105.13549)





Y3 results: Consistency (?) with Planck (in **A**CDM)



 1.5σ parameter tension vs. 2.2σ in Y1

Y3 results: Hubble tension

~4 σ tension with SH0ES



Comparison to other surveys

Comparison to other surveys

Cosmic shear only DES + KiDS 2023 (arXiv:2305.17173)

Combining surveys: DES + KiDS

Cosmic shear - Hybrid analysis pipeline DES + KiDS 2023 (arXiv:2305.17173)

Combining surveys: DES + KiDS

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DES Y3 correlation functions

Modeling galaxy observables

(MassiveBlack II: Khandai+ 2014; Tenneti+ 2014a,b)

Modeling galaxy observables

z = 0.06

Galaxy bias:

$$\delta_{\rm g} \neq \delta_{\rm m}$$

Galaxy intrinsic alignments (IA):

$$\begin{split} \gamma^{\rm obs} &= \gamma^{\rm G} + \gamma^{\rm I} + \epsilon_{\rm n} \\ \langle \gamma_i^{\rm obs} \gamma_j^{\rm obs} \rangle &= \langle \gamma_i^{\rm G} \gamma_j^{\rm G} \rangle + \langle \gamma_i^{\rm G} \gamma_j^{\rm I} \rangle + \langle \gamma_i^{\rm I} \gamma_j^{\rm I} \rangle \end{split}$$

(MassiveBlack II: Khandai+ 2014; Tenneti+ 2014a,b)

Effective perturbative expansions

galaxy positions (biasing)

$$\delta_g(x) = b_1 \delta_m(x) + b_2 \delta_m^2(x) + b_s s^2(x) + \cdots$$

$$\gamma_{ij}^{I} = C_{1}s_{ij} + C_{2}(s_{ik}s_{kj}) + C_{\delta}(\delta s_{ij}) + C_{t}t_{ij} + \cdots$$

galaxy shapes (intrinsic alignments), e.g. TATT model

Implementation with FAST-PT and the Core Cosmology Library: McEwen, Fang, Hirata, JB 2016; Fang, JB, McEwen, Hirata 2017 Chisari et al (LSST-DESC) 2019 FAST-PT on github: JoeMcEwen/FAST-PT CCL on github: LSSTDESC/CCL

20 Mpc/h

e.g. McDonald & Roy 2009; JB+ 2015; 2019; Schmitz, Hirata, JB+ 2019; Vlah+ 2020

Simulating galaxy bias and IA with semi-analytic methods

Halotools-IA Halotools on github: astropy/halotools

Van Alfen, Campbell, JB, Lanusse, Leonard, Hearin+ 2023

In progress: Building a neural net-based emulator for simulation-based modeling

The future present is exciting!

Dark Energy Survey

Kilo Degree Survey

Hyper Suprime Cam

Vera C. Rubin Observatory

- Legacy Survey of Space and Time (LSST)
- LSSTCam (3.2 Gpix) on 8.4m telescope, Cerro Pachón, Chile
- ~1/2 of sky (18-25k deg²)
- 10 year dedicated survey, starting next year
- ~20 billion galaxies
- 8 science collaborations: "Dark Energy" (DESC), also includes dark matter, gravity, inflation, neutrinos, etc.

Beyond two-point statistics

N. Jeffrey, Dark Energy Survey

Dark matter map from DES weak lensing

Beyond two-point statistics

Extracting more information:

- Peaks in the mass map (e.g. Zürcher et al 2022)
- The distribution of convergence, including moments (e.g. Gatti et al 2022)
- Three-point statistics (e.g. Secco et al 2022)
- Field-level inference (e.g. Bayer, Seljak, Modi 2023)
- Other novel statistics (e.g. Gatti et al 2023)

Conclusions

- Studies of weak lensing and galaxy clustering are a powerful cosmological probe for current and future projects.
- Astrophysical modeling will be critical for future analyses, e.g. galaxy IA and biasing.
- New experiments starting now, combined with new statistical and modeling approaches, make the coming decade a very exciting time for observational cosmology!

Thank you!

Extra Slides

"Spectroscopic" surveys

Dark Energy Spectroscopic Instrument

galaxy (over)density field

Robot positioners assembled here!

Combining surveys: DES + KiDS

"In both our mock and data studies, the most significant changes arise from the choice of intrinsic alignment (IA) model."

Combining surveys: DES + KiDS

Cosmic shear - Hybrid analysis pipeline DES + KiDS 2023 (arXiv:2305.17173)

Implementation with FAST-PT

McEwen, Fang, Hirata, JB 2016; Fang, JB, McEwen, Hirata 2017 **FAST-PT on github: JoeMcEwen/FAST-PT**

$$I(k) = \int \frac{d^3 q_1}{(2\pi)^3} K(\hat{q}_1 \cdot \hat{q}_2, \hat{q}_1 \cdot \hat{k}, \hat{q}_2 \cdot \hat{k}, q_1, q_2) P(q_1) P(q_2)$$

$$f(k) = \int \frac{d^3 q_1}{(2\pi)^3} \mathcal{P}_{\ell}(\hat{q}_1 \cdot \hat{q}_2) \mathcal{P}_{\ell_1}(\hat{k} \cdot \hat{q}_2) \mathcal{P}_{\ell_2}(\hat{k} \cdot \hat{q}_1) q_1^{\alpha} q_2^{\beta} P(q_1) P(q_2)$$

$$\bigvee$$

$$J_{J_1 J_2}^{\alpha \beta}(r) \equiv \left[\int_0^{\infty} dq_1 \ q_1^{2+\alpha} P(q_1) j_{J_1}(q_1 r) \right] \left[\int_0^{\infty} dq_2 \ q_2^{2+\beta} P(q_2) j_{J_2}(q_2 r) \right]$$

- Python; easy to use and integrate into other code.
- In DES and LSST analysis software.
- Euclid, Roman in progress.
- DESI? IR Res for BAO and improved RSD in progress. 31

The Core Cosmology Library

Chisari, ..., JB+ 2019 (DESC Collaboration)

CCL on github: LSSTDESC/CCL

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es	PT_demo.ipynt ×									
Ē	8	+	Ж	60		C	Code	V		
Running	Preliminaries Let's just begin by setting up a cosmology and some biases									
Commands	<pre>In [2]: # Cosmology cosmo = ccl.Cosmology(Omega_c=0.27, Omega_b=0.045, h=0.67, A_s=2.1e-9, n # Biases for number counts b_1 = 2.0 # constant values for now b_2 = 1.0</pre>							_c=0.27, Omega_b=0.045, h=0.67, A_s=2.1e-9, n_s=0.96)		
Cell Tools				b_s = # Bia c_1 = c_2 = c_d =	<pre>b_s = 1.0 # Biases for IAs c_1 = 1. c_2 = 0.5 c_d = 0.5</pre>					
Tabs	ulations									
		In	[3]:	# Num	<pre># Number counts ptt_g = pt.PTNumberCountsTracer(b1=b_1, b2=b_2, bs=b_s)</pre>					

ptt_g = pt.PTNumberCountsTracer(b1=b_1, b2=b_2, bs=b_s)
Intrinsic alignments
ptt_i = pt.PTIntrinsicAlignmentTracer(c1=c_1, c2=c_2, cdelta=c_d)
Matter
ptt_m = pt.PTMatterTracer()

 Core theory calculations for LSST (and other) analyses

 Incorporates FAST-PT for nonlinear calculations.

Why go beyond linear theory?

Combining probes

e.g. DES Y3 2023; Krolewski+ 2021 (unWISE + Planck)

Semi-analytic sims consistent with hydro sims

Van Alfen, Campbell, JB, Lanusse, Leonard, Hearin+ 2023