



# New Physics in the third generation: current status and future prospects

Claudia Cornella (JGU Mainz)

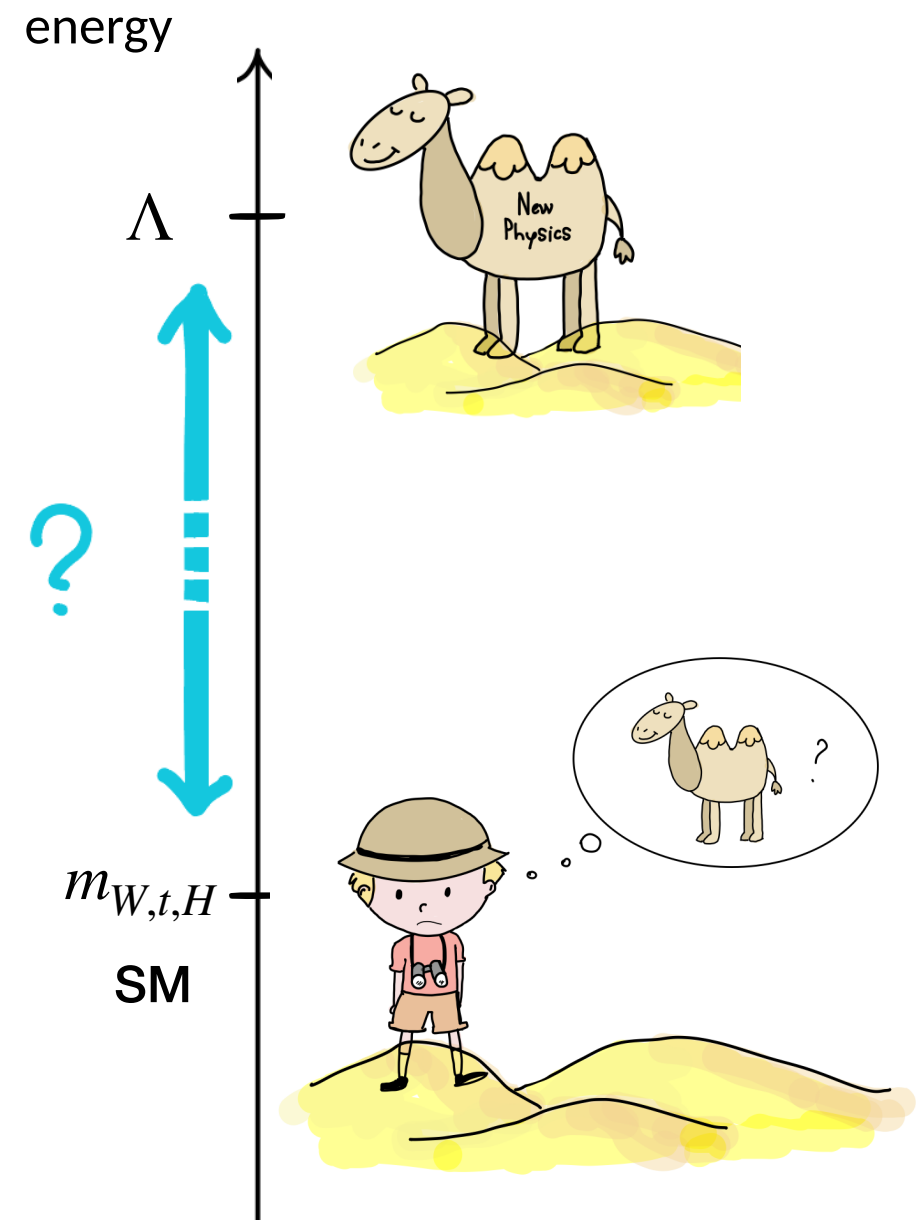
based on 2311.00020 with L. Allwicher, G. Isidori, and B. Stefanek

# The scale of New Physics

We have many reasons to think that the SM must be extended at higher energies. But **how high?**

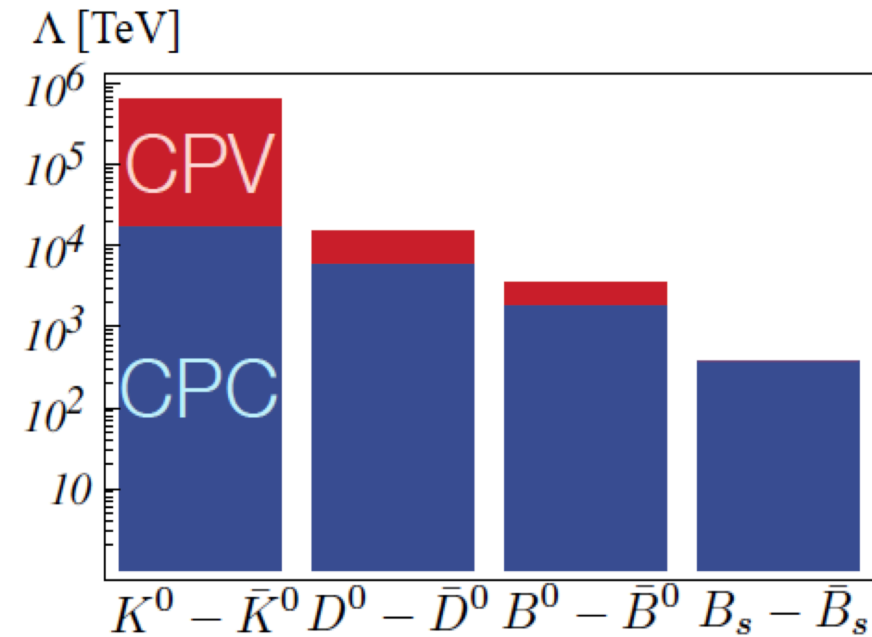
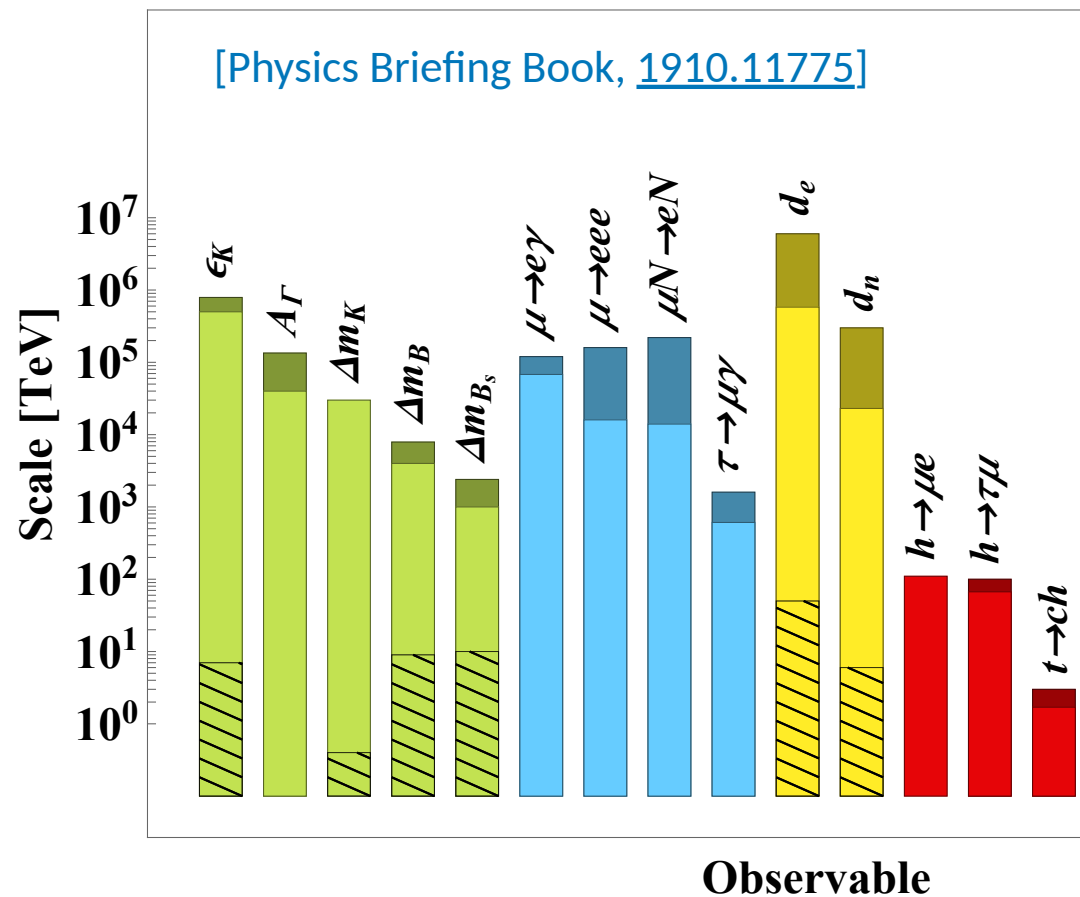
In absence of direct evidence, we rely on the **SMEFT**:

With data we place constraints on the coefficients of SMEFT operators, and interpret them as **constraints on an (effective) NP scale**.



# The scale of New Physics

With  $O(1)$  NP couplings, bounds on **flavor-violating** operators point to **huge scales**:



...but in realistic models these couplings can be suppressed, and give much looser constraints!

Making **educated assumptions about the NP structure** and translating them into selection rules in the SMEFT can provide a more informative interpretation of bounds!

# Goal

Here: focus on **models where NP predominantly couples to the third generation.**

1. **How low** can the energy **scale of new physics** be for these class of models, and which conditions make this possible?
2. **How will the bounds** on these models **change** in the **future**?  
(considering up-coming flavor and collider data, and, more long term, a future e+e- collider like the FCC-ee)



# The SM flavor puzzle and the U(2) symmetry

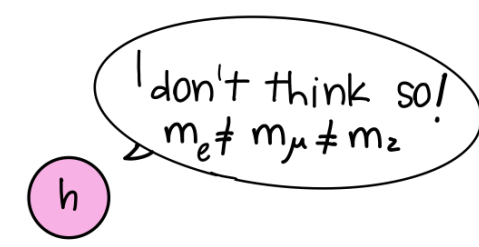
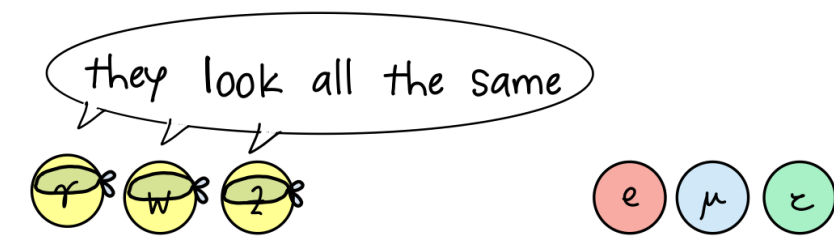
Models where NP couples mostly to the **3rd family** are well-motivated: the 3rd generation plays a special role in the **hierarchy problem** and the **flavor puzzle**.

The **gauge** sector of the SM is **flavor blind**, and has a large accidental symmetry:

$$\mathcal{G}_F = U(3)^5 \equiv U(3)_q \times U(3)_u \times U(3)_d \times U(3)_\ell \times U(3)_e$$

**Yukawa** interactions **break** this symmetry in a specific way:

$$M_{e,d,u} = \begin{bmatrix} \square & & \\ & \square & \\ & & \blacksquare \end{bmatrix} \quad V_{\text{CKM}} = \begin{bmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{bmatrix} = \begin{bmatrix} \blacksquare & \square & \square \\ \square & \blacksquare & \square \\ \square & \square & \blacksquare \end{bmatrix}$$



$$U^5(3) \rightarrow U(2)^5 \equiv U(2)_q \times U(2)_u \times U(2)_d \times U(2)_\ell \times U(2)_e$$

$$\psi = ((\psi_1 \ \psi_2) \ \psi_3)$$

[Barbieri et al. 2022, Isidori, Straub 2012]

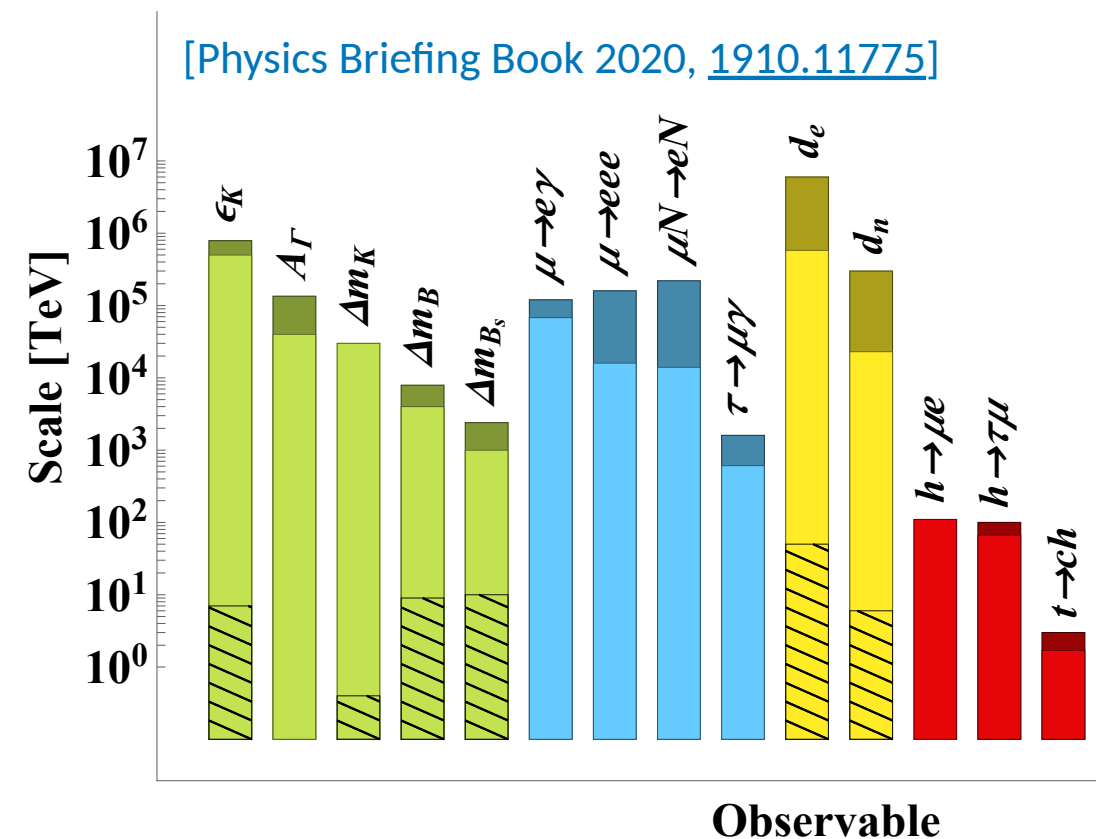
# The New Physics flavor puzzle

## The NP flavor puzzle:

Flavor is just an accidental symmetry: nothing forbids it to be badly violated in the UV. Then why don't we observe sizeable non-standard flavor-violating effects?

Either because the scale of these interaction is astronomically high, or because the couplings of these operators are small.

In either case, the only **unambiguous** message of these bounds is that **there is no large breaking of  $U(2)^5$  at nearby scales.**



**$U(2)^5$  is a good symmetry also of the SMEFT!**

# $U(2)^5$ vs MFV

A way to allow for TeV NP while protecting it from flavor bounds was to assume **Minimal Flavor Violation**.

- Yukawas are the only sources of  $G_f=U(3)^5$  breaking also beyond the SM.
- by construction, MFV gives little to no effect in flavor-changing processes.
- MFV describes (perturbations around) **flavor-universal NP**  $\textcircled{1} = \textcircled{2} = \textcircled{3}$

In particular, it does *not* suppress NP couplings to valence quarks....

....Now **LHC data push the scale of MFV NP to scales  $\gtrsim 10$  TeV!**

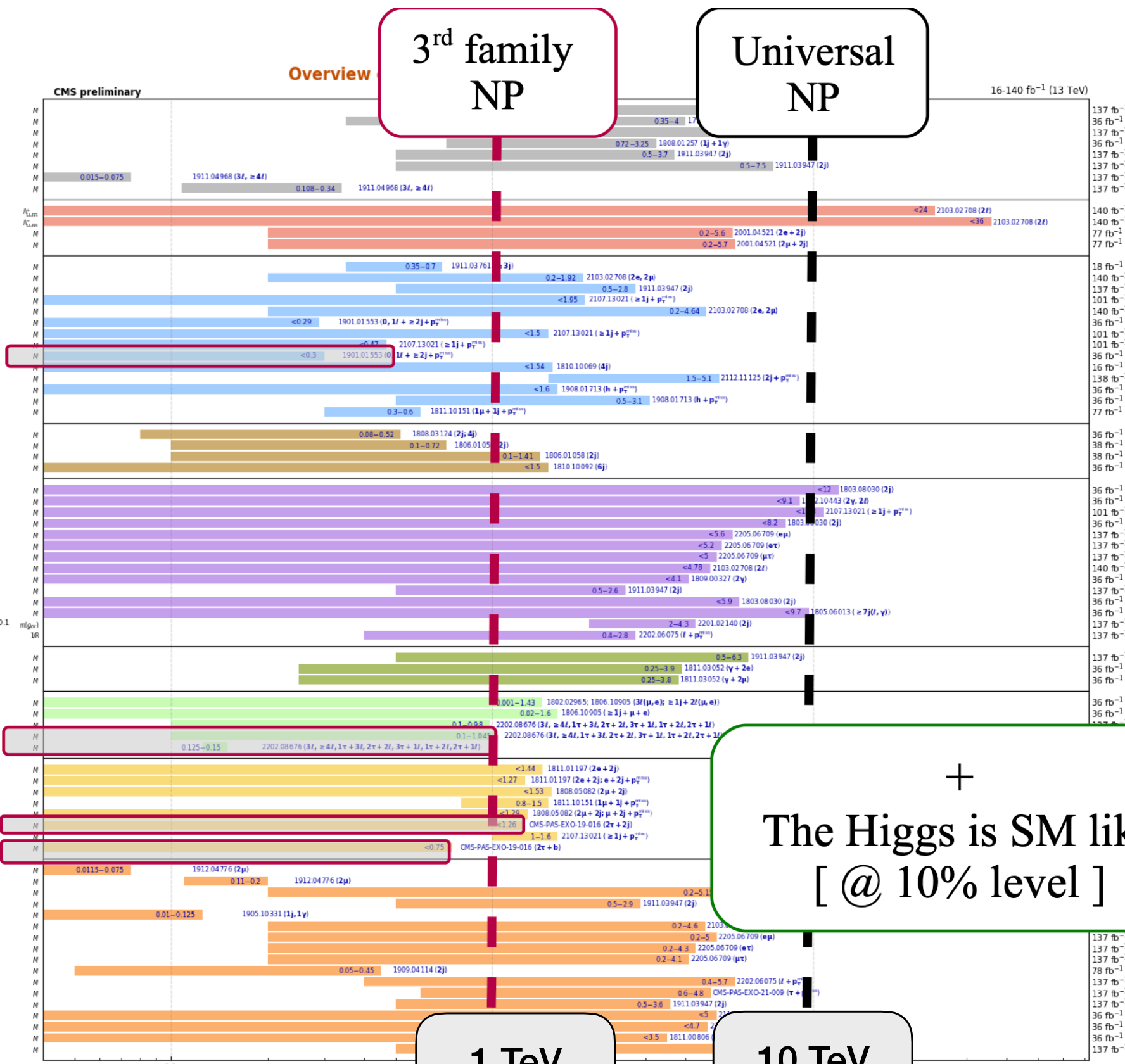
By contrast,  $U(2)^5$  describes **flavor non-universal NP**, placing a clear distinction between light and heavy generations.  $\textcircled{1} = \textcircled{2} \neq \textcircled{3}$

Different NP couplings for light families make it possible to suppress couplings to valence quarks and relax direct search bounds to  **$\sim 1$  TeV**

# Status of high-energy searches

@ICHEP2022

<b>Other</b>	String resonance Zy resonance Wy resonance Higgs y resonance Color Octet Scalar, $k_2^2 = 1/2$ Scalar Diquark $t\bar{t} + \phi$ , pseudoscalar (scalar), $g_{\text{top}}^2 \times \text{BR}(\phi \rightarrow 2l) > = 0.03(0.04)$ $t\bar{t} + \phi$ , pseudoscalar (scalar), $g_{\text{top}}^2 \times \text{BR}(\phi \rightarrow 2l) > = 0.03(0.04)$
<b>Contact Interactions</b>	quark compositeness ( $ll$ ), $\eta_{L,RR} = 1$ quark compositeness ( $ll$ ), $\eta_{L,RR} = -1$ Excited Lepton Contact Interaction Excited Lepton Contact Interaction
<b>Dark Matter</b>	vector mediator ( $qq$ ), $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV vector mediator ( $ll$ ), $g_s = 0.1, g_{\text{DM}} = 1, g_l = 0.01, m_\nu > 1$ TeV (axial-)vector mediator ( $qq$ ), $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV (axial-)vector mediator ( $ll$ ), $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV (axial-)vector mediator ( $ll$ ), $g_s = 0.1, g_{\text{DM}} = 1, g_l = 0.1, m_\nu > m_{\text{charm}}/2$ scalar mediator ( $+t\bar{t}$ ), $g_s = 1, g_{\text{DM}} = 1, m_\nu = 1$ GeV scalar mediator (fermion portal), $\lambda_s = 1, m_\nu = 1$ GeV pseudoscalar mediator ( $+t\bar{t}$ ), $g_s = 1, g_{\text{DM}} = 1, m_\nu = 1$ GeV pseudoscalar mediator ( $+t\bar{t}$ ), $g_s = 1, g_{\text{DM}} = 1, m_\nu = 1$ GeV complex sc. med. (dark QCD), $m_{\text{dark}} = 5$ GeV, $c\tau_{\text{dark}} = 25$ mm Z' mediator (dark QCD), $m_{\text{dark}} = 20$ GeV, $r_{\text{dark}} = 0.3, \alpha_{\text{dark}} = \alpha_{\text{SM}}$ Baryonic Z', $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV Z' - 2HDM, $g_Z = 0.8, g_{\text{DM}} = 1, \tan\beta = 1, m_\nu = 100$ GeV Leptoquark mediator, $\beta = 1, B = 0.1, A_{\text{L,DM}} = 0.1, 800 < M_{LQ} < 1500$ GeV
<b>RPV</b>	RPV stop to 4 quarks RPV squark to 4 quarks RPV gluino to 4 quarks RPV gluinos to 3 quarks
<b>Extra Dimensions</b>	ADD ( $jj$ ) HLZ, $n_{\text{ED}} = 3$ ADD ( $yy, ll$ ) HLZ, $n_{\text{ED}} = 3$ ADD $G_{XX}$ emission, $n_{\text{ED}} = 2$ ADD QBH ( $jj$ ), $n_{\text{ED}} = 6$ ADD QBH ( $jj$ ), $n_{\text{ED}} = 4$ ADD QBH ( $jj$ ), $n_{\text{ED}} = 4$ ADD QBH ( $jj$ ), $n_{\text{ED}} = 4$ RS $G_{XX}(ll)$ , $k/\bar{M}_P = 0.1$ RS $G_{XX}(yy)$ , $k/\bar{M}_P = 0.1$ RS $G_{XX}(qq, gg)$ , $k/\bar{M}_P = 0.1$ RS QBH ( $jj$ ), $n_{\text{ED}} = 1$ non-rotating BH, $M_0 = 4$ TeV, $n_{\text{ED}} = 6$ 3-brane WED $g_{XX}(\phi + g + gg)$ , $g_{\text{UV}} = 6, g_{\text{IR}} = 3, \epsilon = 0.5, m(\phi)/m(g_{XX}) = 0.1$ split-UED, $\mu \geq 2$ TeV
<b>Excited Fermions</b>	excited light quark ( $qq$ ), $\Lambda = m_{\text{exc}}^*$ excited electron, $f_e = f = F = 1, \Lambda = m_{\text{exc}}^*$ excited muon, $f_e = f = F = 1, \Lambda = m_{\text{exc}}^*$
<b>Heavy Fermions</b>	$v\text{MSM},  V_{cb} ^2 = 1.0,  V_{cb} ^2 = 1.0$ $v\text{MSM},  V_{cb} ^2/ V_{cb} ^2 +  V_{cb} ^2 = 1.0$ Type-III seesaw heavy fermions, Flavor-democratic Vector like taus, Doublet Vector like taus, Singlet
<b>Leptoquarks</b>	scalar LQ (pair prod.), coupling to 1 <sup>st</sup> gen. fermions, $\beta = 1$ scalar LQ (pair prod.), coupling to 1 <sup>st</sup> gen. fermions, $\beta = 0.5$ scalar LQ (pair prod.), coupling to 2 <sup>nd</sup> gen. fermions, $\beta = 1$ scalar LQ (pair prod.), coupling to 2 <sup>nd</sup> gen. fermions, $\beta = 0.5$ scalar LQ (pair prod.), coupling to 3 <sup>rd</sup> gen. fermions, $\beta = 1$ scalar LQ (single prod.), coupling to 1 <sup>st</sup> gen. fermions, $\beta = 0, \lambda = 1$ scalar LQ (single prod.), coupling to 3 <sup>rd</sup> gen. fermions, $\beta = 1, \lambda = 1$
<b>Heavy Gauge Bosons</b>	Z <sub>1</sub> , narrow resonance Z <sub>2</sub> , narrow resonance SSM Z' ( $ll$ ) SSM Z' ( $qq$ ) Z' ( $qq$ ) Superstring Z' <sub>1</sub> LFV Z', BR( $\mu\mu$ ) = 10% LFV Z', BR( $\tau\tau$ ) = 10% LFV Z', BR( $\mu\tau$ ) = 10% Leptophobic Z' SSM W' ( $lv$ ) SSM W' ( $\tau\nu$ ) SSM W' ( $qq$ ) LRSM W <sub>1</sub> ( $\mu N_c$ ), $M_{W_1} = 0.5M_{\text{SM}}$ LRSM W <sub>2</sub> ( $e N_c$ ), $M_{W_2} = 0.5M_{\text{SM}}$ LRSM W <sub>3</sub> ( $\tau N_c$ ), $M_{W_3} = 0.5M_{\text{SM}}$ Axigluon, Coloron, $\cot\theta = 1$



3<sup>rd</sup> family NP

Universal NP

+  
The Higgs is SM like  
[ @ 10% level ]

1 TeV

10 TeV

Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not included).



# Flavor non-universal interactions

These considerations translate into model-building ideas!

For a while, attempts to extend the SM implicitly assumed:

- TeV-scale **flavor-universal** NP stabilising the Higgs
- flavor dynamics originates at some  $\Lambda \gg \text{TeV}$

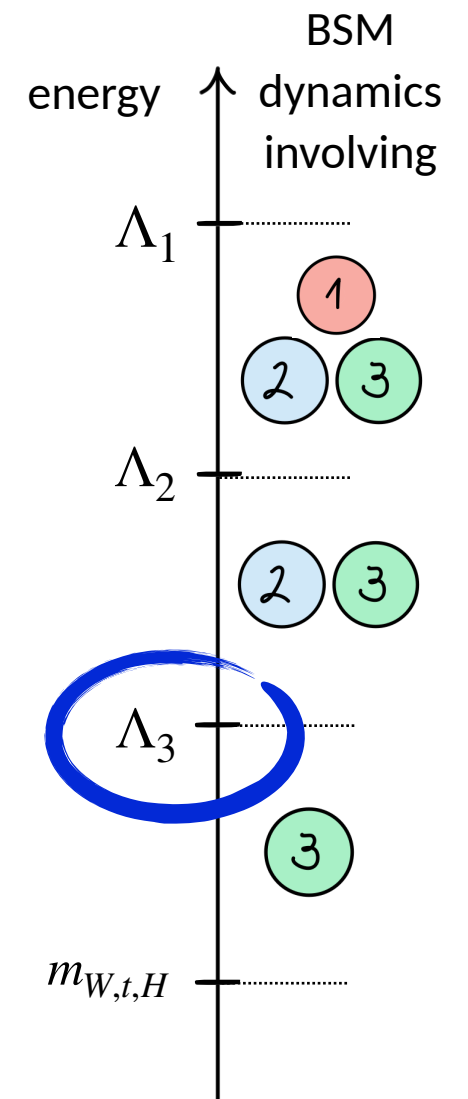
Now **flavor non-universal** interactions are gaining momentum.

[Dvali, Shifman, '00, Panico, Pomarol 1603.06609;...Bordone, CC, Fuentes, Isidori 1712.01368; Barbieri, 2103.15635; Davighi, Isidori, 2303.01520; Davighi, Stefanek, 2305.16280, Greljo, Thomsen 2309.11547...]

- The 3 families are *not* identical up to very high energies.

*Multiscale picture:* non-universal interactions acting on the  $i$ -th family switch on at  $\Lambda_1 \gg \Lambda_2 \gg \Lambda_3 \gg m_W$

- interactions distinguishing light vs 3rd family emerge first @  $\Lambda_3$



# The $U(2)$ symmetric SMEFT

$U(2)^5$  is an **efficient organising principle**:

- SMEFT with 3 generations has  $1350 + 1149 = 2499$  independent WCs at dim-6.
- In the exact  $U(2)^5$  limit, this is reduced to  $124 + 23 = 147$  independent WCs.

Here we focus on the CP-conserving case.

Operators	$U(2)^5$ [terms summed up to different orders]													
	Exact		$\mathcal{O}(V^1)$		$\mathcal{O}(V^2)$		$\mathcal{O}(V^1, \Delta^1)$		$\mathcal{O}(V^2, \Delta^1)$		$\mathcal{O}(V^2, \Delta^1 V^1)$		$\mathcal{O}(V^3, \Delta^1 V^1)$	
Class 1–4	9	6	9	6	9	6	9	6	9	6	9	6	9	6
$\psi^2 H^3$	3	3	6	6	6	6	9	9	9	9	12	12	12	12
$\psi^2 XH$	8	8	16	16	16	16	24	24	24	24	32	32	32	32
$\psi^2 H^2 D$	15	1	19	5	23	5	19	5	23	5	28	10	28	10
$(\bar{L}L)(\bar{L}L)$	23	–	40	17	67	24	40	17	67	24	67	24	74	31
$(\bar{R}R)(\bar{R}R)$	29	–	29	–	29	–	29	–	29	–	53	24	53	24
$(\bar{L}L)(\bar{R}R)$	32	–	48	16	64	16	53	21	69	21	90	42	90	42
$(\bar{L}R)(\bar{R}L)$	1	1	3	3	4	4	5	5	6	6	10	10	10	10
$(\bar{L}R)(\bar{L}R)$	4	4	12	12	16	16	24	24	28	28	48	48	48	48
<b>total:</b>	<b>124</b>	<b>23</b>	182	81	234	93	212	111	264	123	349	208	356	215

Table 6: Number of independent operators in the SMEFT assuming a minimally broken  $U(2)^5$  symmetry, including breaking terms up to  $\mathcal{O}(V^3, \Delta^1 V^1)$ . Notations as in Table 1.

[D. A. Faroughy, G. Isidori, F. Wilsch, K. Yamamoto, [arXiv:2005.05366](https://arxiv.org/abs/2005.05366)]

# The flavor rotation

What is the **third generation** in the SMEFT?

Non-trivial to define for the LH quark doublet because of the CKM misalignment!

In the interaction basis where the dim-6 SMEFT operators are  $U(2)^5$  symmetric, the **3rd generation quark doublet** is somewhere **in-between** the **down-aligned** and the **up-aligned** case.

$$\begin{pmatrix} t_L \\ V_{td}d_L + V_{ts}s_L + V_{tb}b_L \end{pmatrix} = \begin{matrix} q_t \\ q_3 \\ q_b \end{matrix} = \begin{pmatrix} V_{ub}^* u_L + V_{cb}^* c_L + V_{tb}^* t_L \\ b_L \end{pmatrix}$$

We can describe this **misalignment** in terms of a single **angle** in the 2-3 sector,  $\theta \sim V_{cb} \epsilon_F$ .

# Observables

## EWPO

- W-pole observables [V. Bresó-Pla, A. Falkowski, M. González-Alonso, [2103.12074](#)]
- Z-pole observables [L. Allwicher, G. Isidori, J. M. Lizana, N. Selimovic, B. Stefanek, [2302.11584](#)]
- Higgs signal strengths + LFU tests in  $\tau$ -decays

## Flavor

- $\Delta F = 1$  ( $B \rightarrow X_s \gamma$ ,  $B \rightarrow K \nu \bar{\nu}$ ,  $K \rightarrow \pi \nu \bar{\nu}$ ,  $B \rightarrow K^{(*)} \mu^+ \mu^-$ ,  $B_{s,d} \rightarrow \mu^+ \mu^-$ )
- $\Delta F = 2$  ( $B_{s,d}$ - mixing,  $K$ - mixing,  $D$ - mixing)
- Charged-current  $b \rightarrow c, u$  transitions ( $R_D, R_{D^*}, B_{u,c} \rightarrow \tau \nu$ )

## Collider

- LHC Drell-Yan  $pp \rightarrow \ell \ell$  and mono-lepton  $pp \rightarrow \ell \nu$  [L. Allwicher, D. A. Faroughy, F. Jaffredo, O. Sumensari, F. Wilsch, [2207.10756](#)]
- LHC 4-quark observables
- LEP 4-lepton  $ee \rightarrow \ell \ell$  [Ethier, Magni, Maltoni, Mantani, Nocera, Rojo, Slade, Vryonidou, Zhang, [2105.00006](#)]

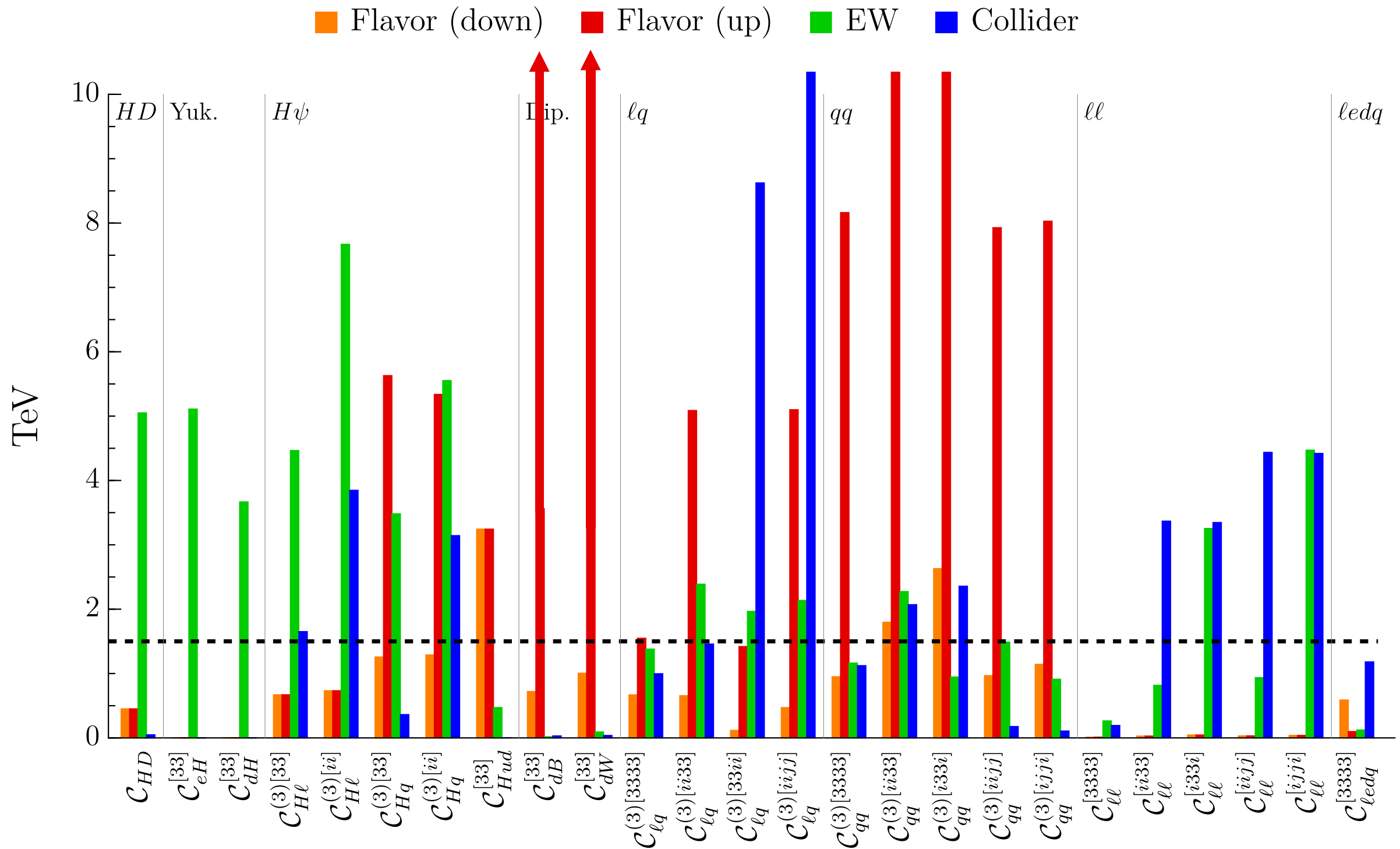




# Analysis strategy

- Run all WCs to a reference scale  $\Lambda = 3 \text{ TeV}$ .
- For LEFT running, LEFT-SMEFT matching and SMEFT running we use DSixTools, which allows us to work analytically in the WCs also beyond leading log.
- Once all observables have been expressed in terms of SMEFT WCs at the high scale, we impose the  $U(2)^5$  symmetry.
- We construct the combined likelihood from collider, EW, and flavour observables as a function of the 124 WCs of the  $U(2)^5$ -symmetric (and CP conserving) SMEFT, and switch them on one at a time to get lower bound on the NP scale.

# Results



# Results

Strong **complementarity** between 3 sectors.

Out of 124 bounds, 46 are dominated by **EWPO**, 42 by **collider**, 36 by **flavor**

- the strongest bounds in the **EW sector** are 5 - 10 TeV for operators with one or more Higgs fields.
- the strongest bounds from **collider data** are 5 - 20 TeV for 4-fermion operators with 1st-family quarks and leptons.

Operators with 3rd-family fermions get milder bounds,  $\sim 1$  TeV.

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$
$\mathcal{C}_{H\ell}^{(1)[33]}$	0.1	0.1	4.4	1.6
$\mathcal{C}_{H\ell}^{(1)[ii]}$	0.7	0.7	7.6	3.
$\mathcal{C}_{H\ell}^{(3)[33]}$	0.7	0.7	4.5	1.7
$\mathcal{C}_{H\ell}^{(3)[ii]}$	0.7	0.7	7.7	3.8
$\mathcal{C}_{He}^{[33]}$	-	-	3.8	1.5
$\mathcal{C}_{He}^{[ii]}$	0.9	0.9	6.6	2.7
$\mathcal{C}_{\ell q}^{(3)[3333]}$	0.7	1.5	1.4	1.
$\mathcal{C}_{\ell q}^{(3)[ii33]}$	0.7	5.1	2.4	1.5
$\mathcal{C}_{\ell q}^{(3)[33ii]}$	0.1	1.4	2.	8.6
$\mathcal{C}_{\ell q}^{(3)[iijj]}$	0.5	5.1	2.1	22.5

# Results

Strong **complementarity** between 3 sectors.

Out of 124 bounds, 46 are dominated by **EWPO**, 42 by **collider**, 36 by **flavor**

For operators contributing to **flavor-violating** observables, U(2) is quite effective in reducing the associated scales.

- Still, certain operators get bounds of 5 - 10 TeV, especially in the **up-aligned** scenario, similarly to MFV.
- **Down alignment** can relax these bounds down to ~ few TeV.

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$
$\mathcal{C}_{qq}^{(1)[3333]}$	1.	7.8	1.6	1.1
$\mathcal{C}_{qq}^{(1)[ii33]}$	1.3	11.2	0.9	1.5
$\mathcal{C}_{qq}^{(1)[i33i]}$	2.5	11.3	0.7	1.6
$\mathcal{C}_{qq}^{(1)[iijj]}$	0.9	8.1	0.4	-
$\mathcal{C}_{qq}^{(1)[ijji]}$	1.1	8.1	0.5	-



# Results

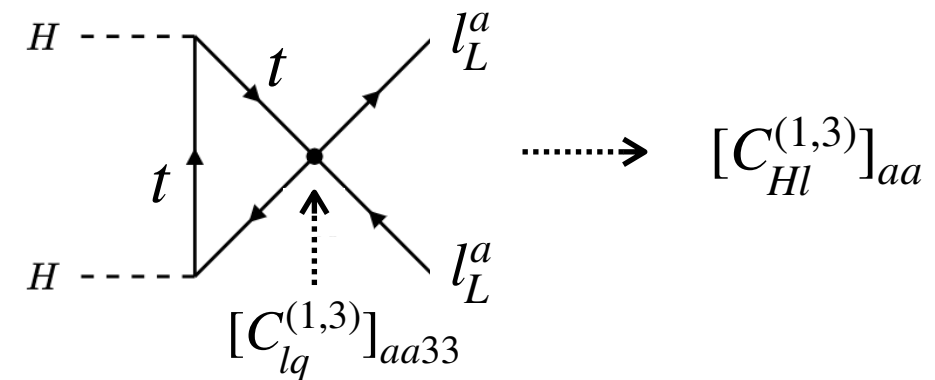
- **RG effects in the EW sector are very important.**

Without running, only 16 operators enter the EW fit.

With running, 123 out of 124 operators enter the EW fit.

44 get bounds stronger than 1 TeV!

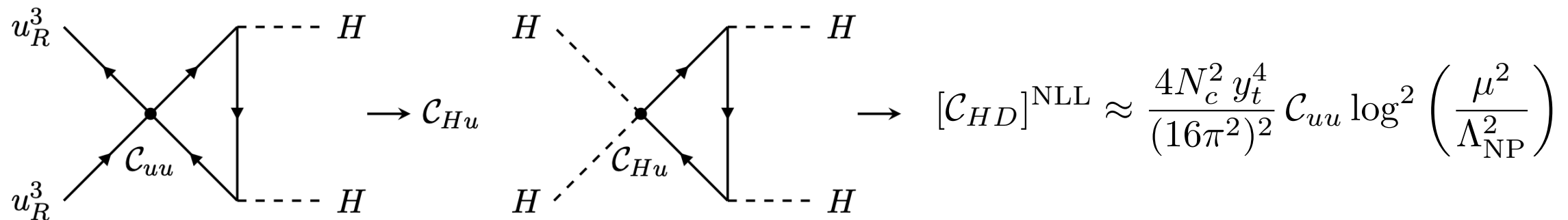
these are operators w/ 3rd-family quarks running with  $y_t$  into operators directly constrained by Z-pole obs.



- **Going beyond LL when solving RGEs is also important.**

NLL effects can change bounds by 30%

Example:  $[O_{uu}]_{3333}$  enters the EW fit only at NLL by mixing with  $O_{HD}$



# The hypothesis of NP in the 3rd generation

Until now, we have used  $U(2)^5$  without other assumptions.

$U(2)^5$  does not specify whether NP interacts more with light or 3rd-family fermions: it just distinguishes among them and protects against flavor violation in the light families.

**Now** focus on the well-motivated case where **NP couples mostly to the 3rd family**:

- WCs of operators w/light fields get a suppression  $\epsilon_q, \epsilon_l$  for each light quark & lepton:

$$C_{qe}^{[ijj]} = \frac{\epsilon_q^2 \epsilon_l^2}{\Lambda^2}$$

Additional assumptions:

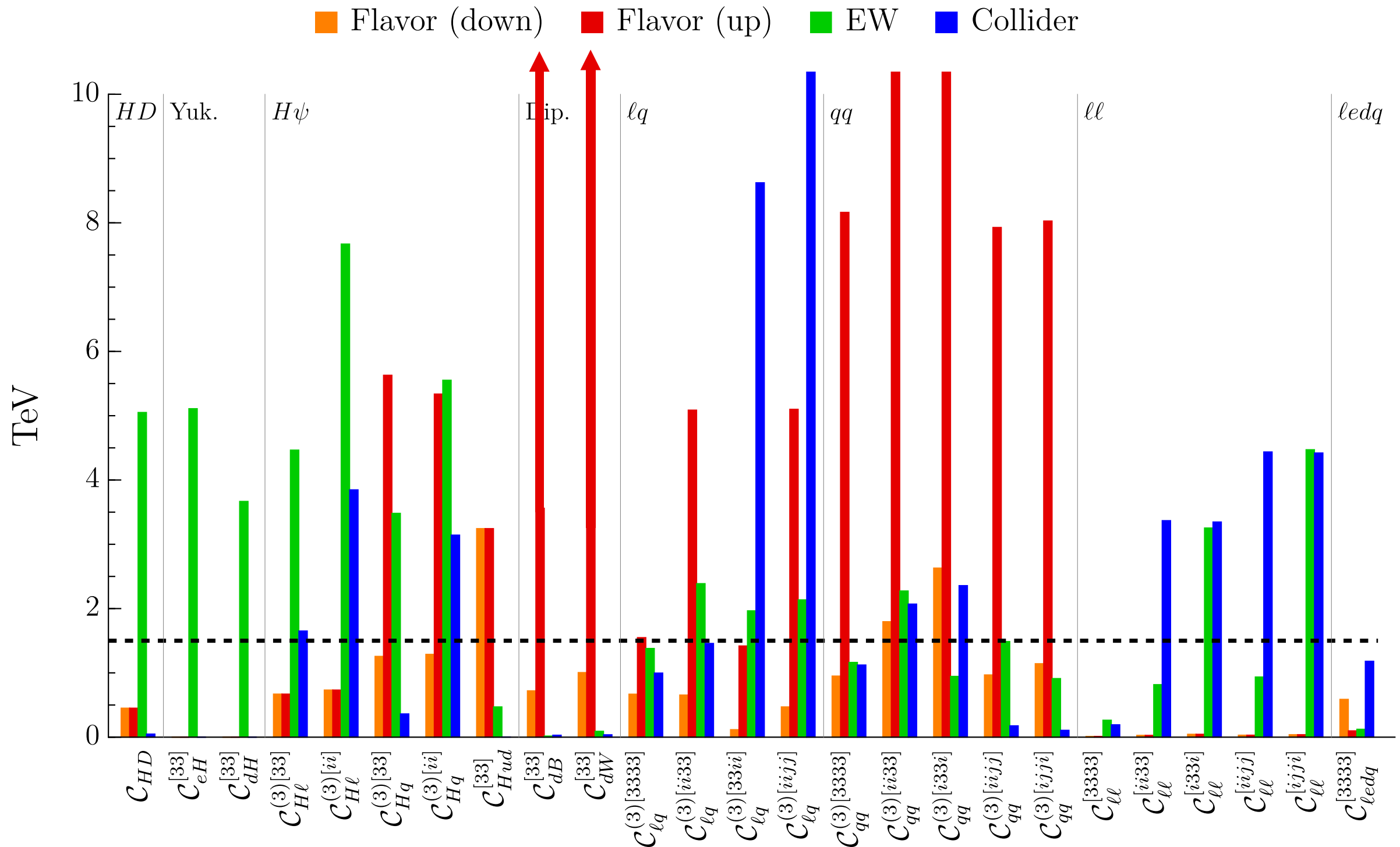
- WCs of operators with Higgs fields gets a suppression  $\epsilon_H$  for each Higgs
- operators w/field strengths are loop generated  $\Rightarrow$  suppressed by  $\epsilon_{\text{loop}} = \prod_i \frac{g_i}{16\pi^2}$

Only 4-fermion operators with 3rd family fields only are unsuppressed.

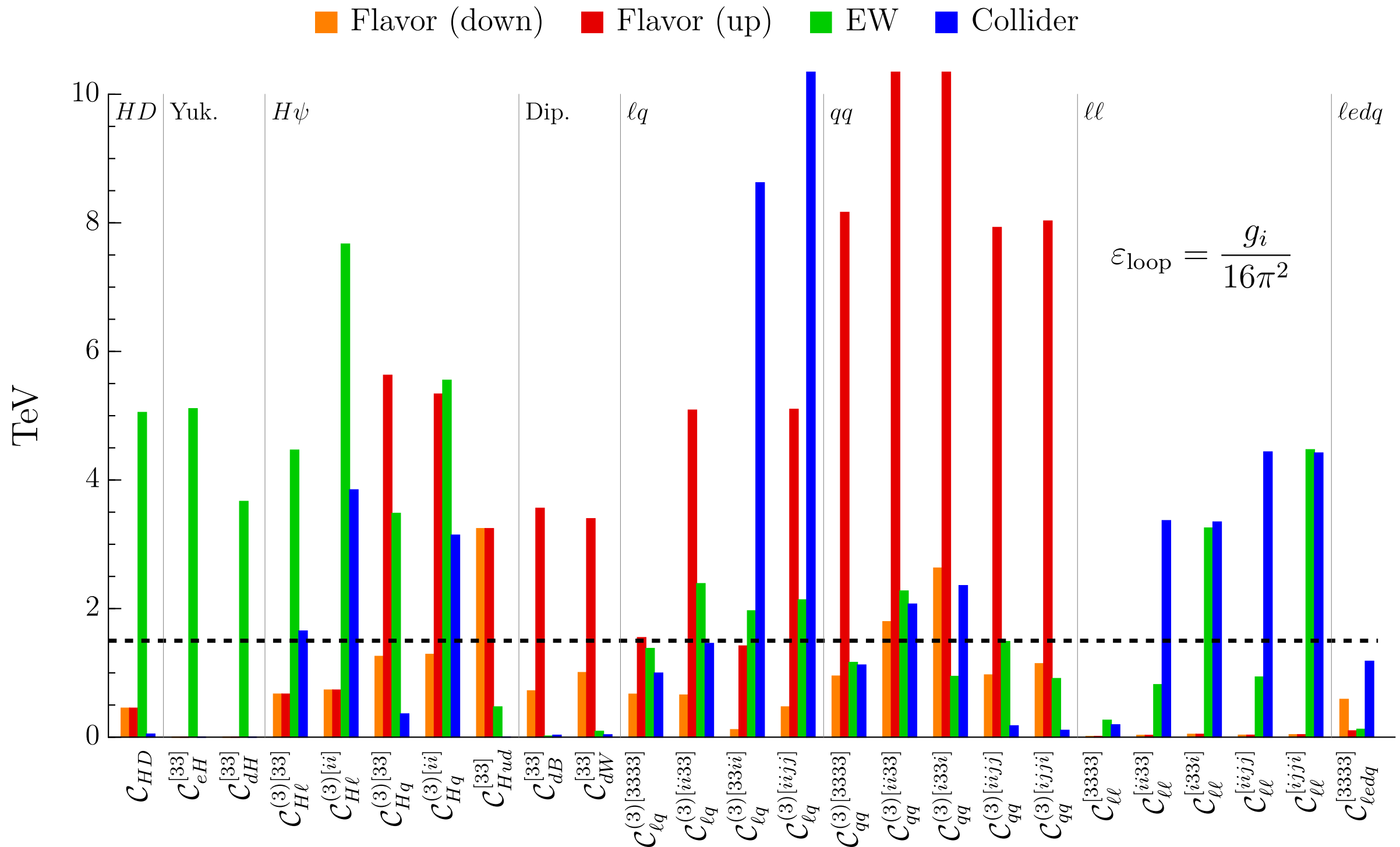
For them,  $\Lambda \sim 1.5$  TeV.

**Can we make the bounds on ALL other operators compatible with 1.5 TeV for reasonable values for the suppression factors  $\epsilon_q, \epsilon_l$ , and  $\epsilon_H$ ?**

# The hypothesis of NP in the 3rd generation

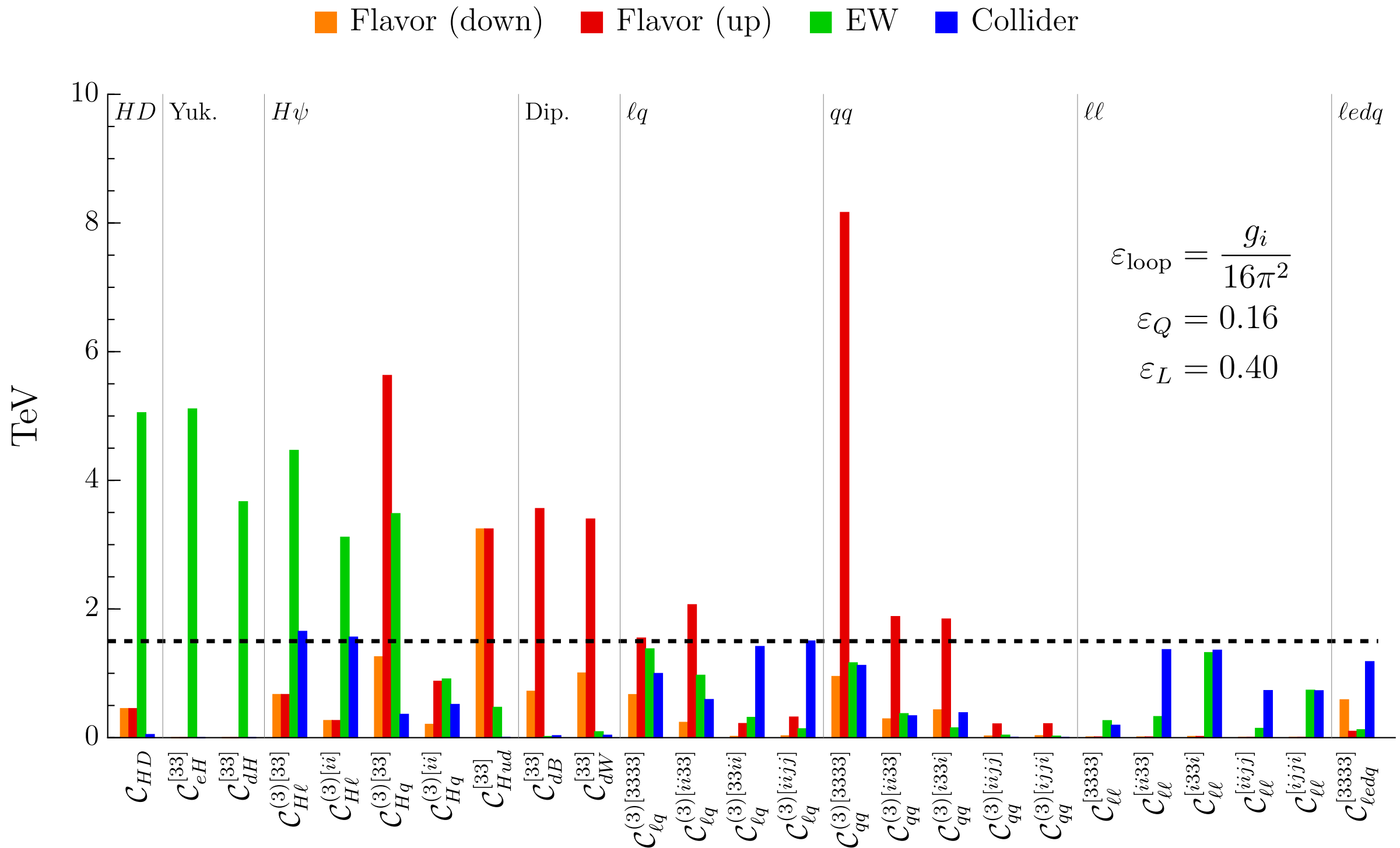


# The hypothesis of NP in the 3rd generation

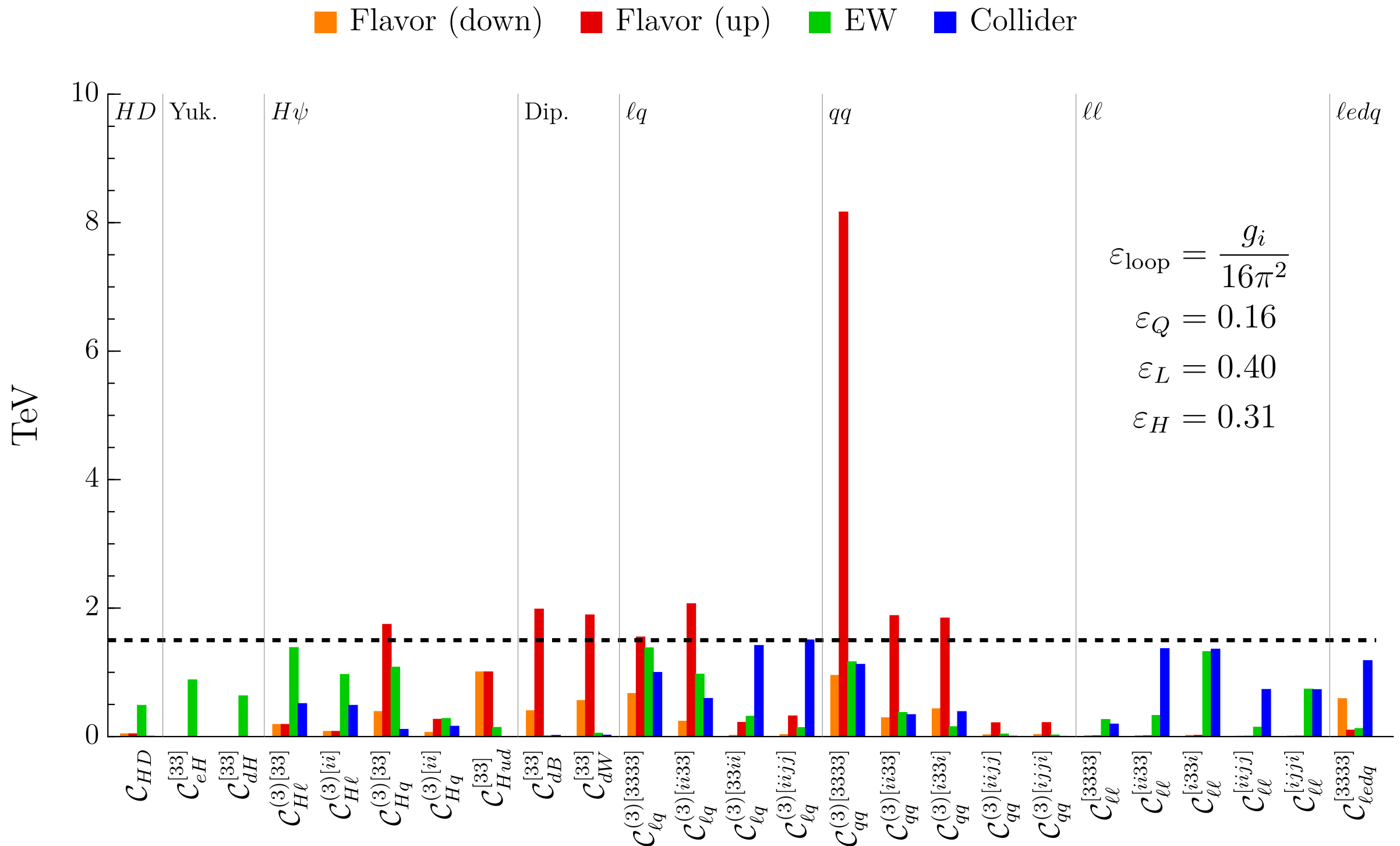




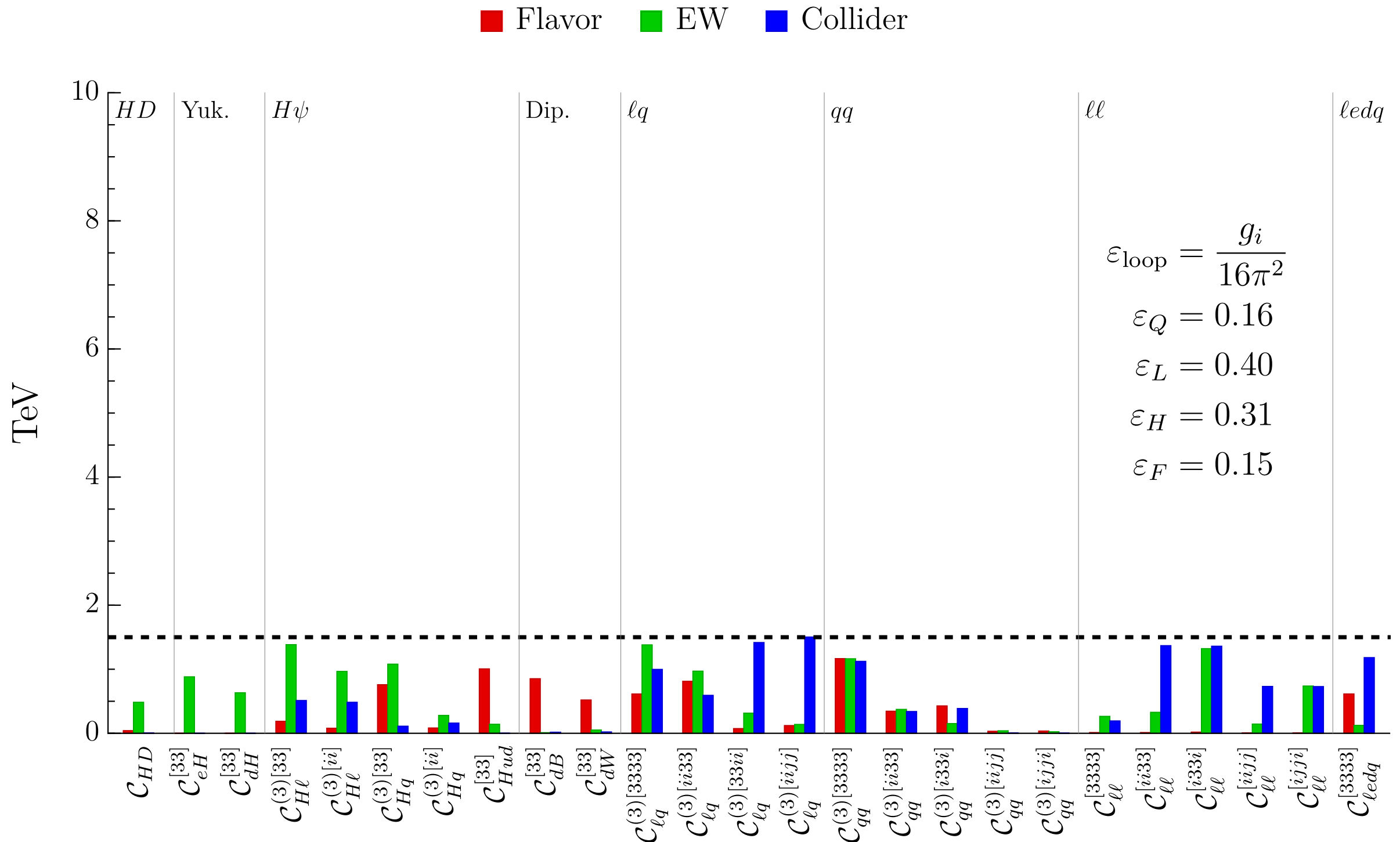
# The hypothesis of NP in the 3rd generation



# The hypothesis of NP in the 3rd generation



# The hypothesis of NP in the 3rd generation



# The hypothesis of NP in the 3rd generation

New Physics mainly coupled to the 3rd generation compatible with all current data can exist at scales as low as 1.5 TeV under these conditions:

$$\varepsilon_q \leq 0.16, \quad \varepsilon_l \leq 0.40, \quad \varepsilon_H \leq 0.31, \quad \varepsilon_F \leq 0.15$$

The precise numbers are not “special”, but give a semi-quantitative **indication** of the general UV conditions NP models must meet to exist at nearby scales.

These conditions are **simple** to realise & **radiatively stable**:  
we can envision realistic SM extensions with NP predominantly coupled to  
the 3rd generation right at the TeV scale!

# Projections for FCC-ee

The expected improvements for Z- and W-pole observables, Higgs and tau decays are available from the literature.

[J. De Blas, G. Durieux, C. Grojean, J. Gu and A. Paul, [1907.04311](#), A. Blondel and P. Janot, [2106.13885](#), Snowmass [2203.06520](#)]

Tera Z- pole run:  **$10^5$  more Z bosons than LEP**, so statistics can improve by up to a factor 300.

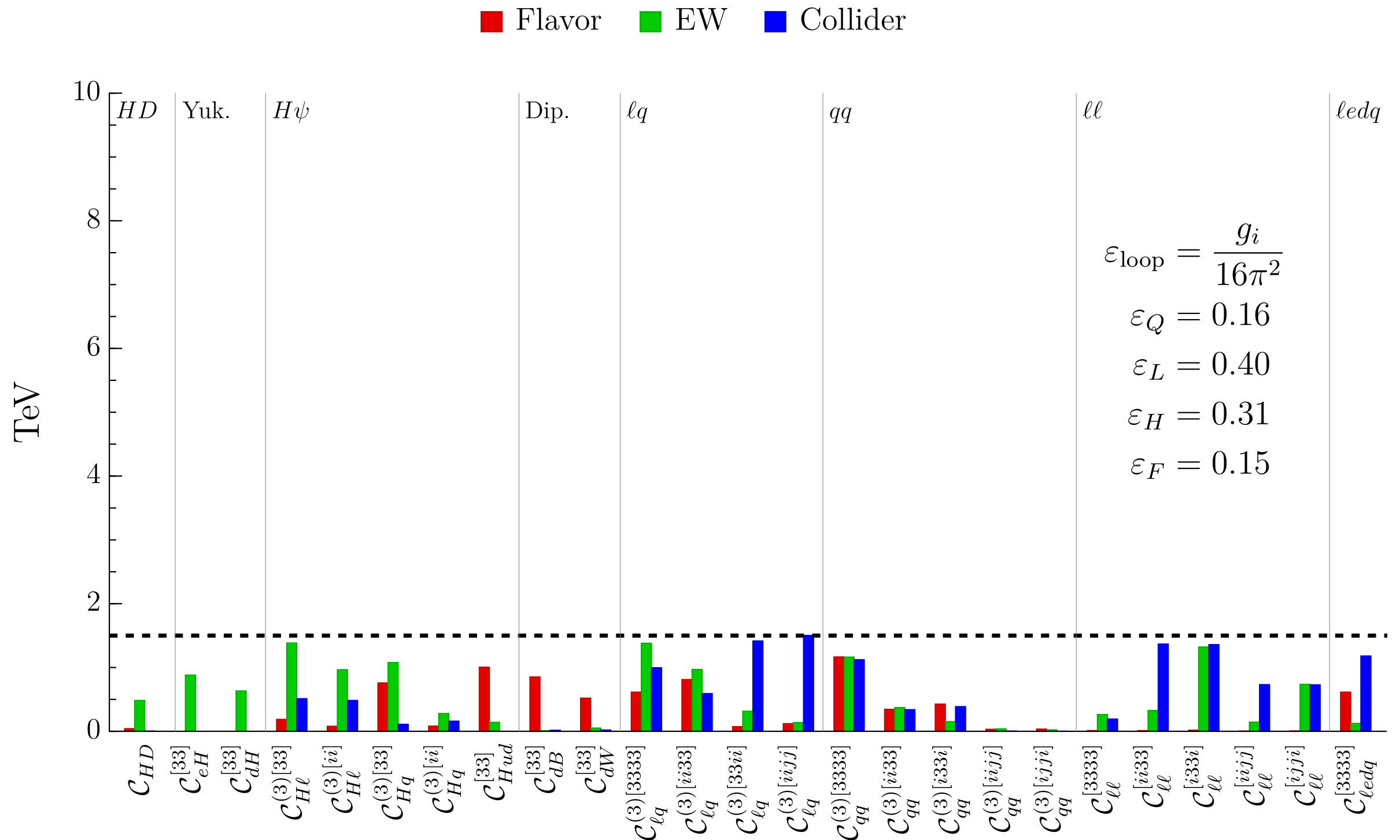
In practice, **leptonic** (**hadronic**) obs. improve by a factor **10-100** (**10**).

To build a projected EW likelihood for FCC-ee:

- Exp. values set to the SM
- error reduction as tabulated in the literature

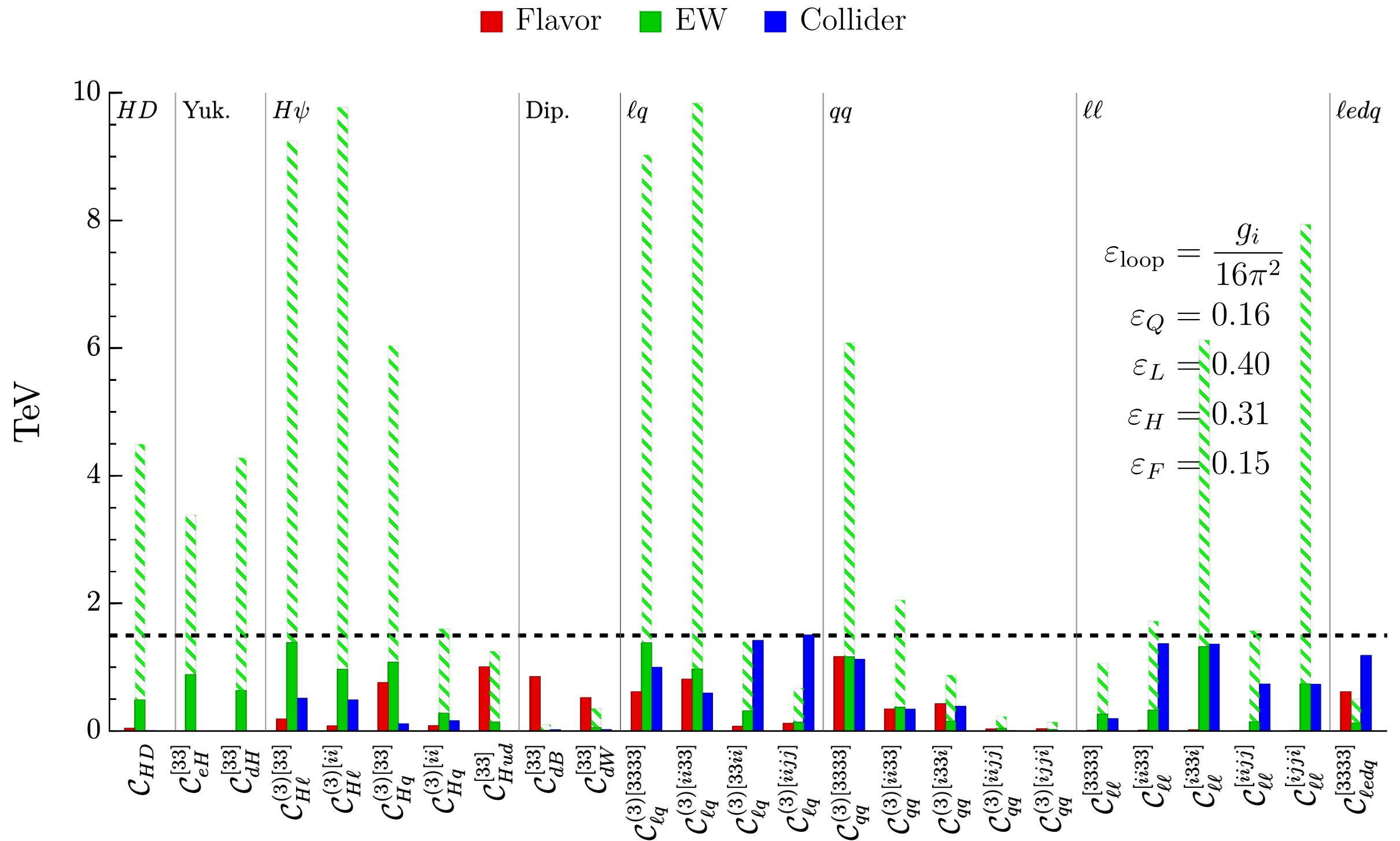
Observable	Proj. Error Reduction
$\Gamma_Z$	23
$\sigma_{\text{had}}^0$	7.4
$R_b$	10.2
$R_c$	11.6
$A_{\text{FB}}^{0,b}$	15.5
$A_{\text{FB}}^{0,c}$	15.4
$A_b$	7.13
$A_c$	5.05
$R_e$	8.03
$R_\mu$	31.8
$R_\tau$	21.7
$A_{\text{FB}}^{0,e}$	30.8
$A_{\text{FB}}^{0,\mu}$	26.7
$A_{\text{FB}}^{0,\tau}$	21
$A_e^{**}$	130
$A_\mu^{**}$	680
$A_\tau^{**}$	340

# FCC and 3rd generation NP

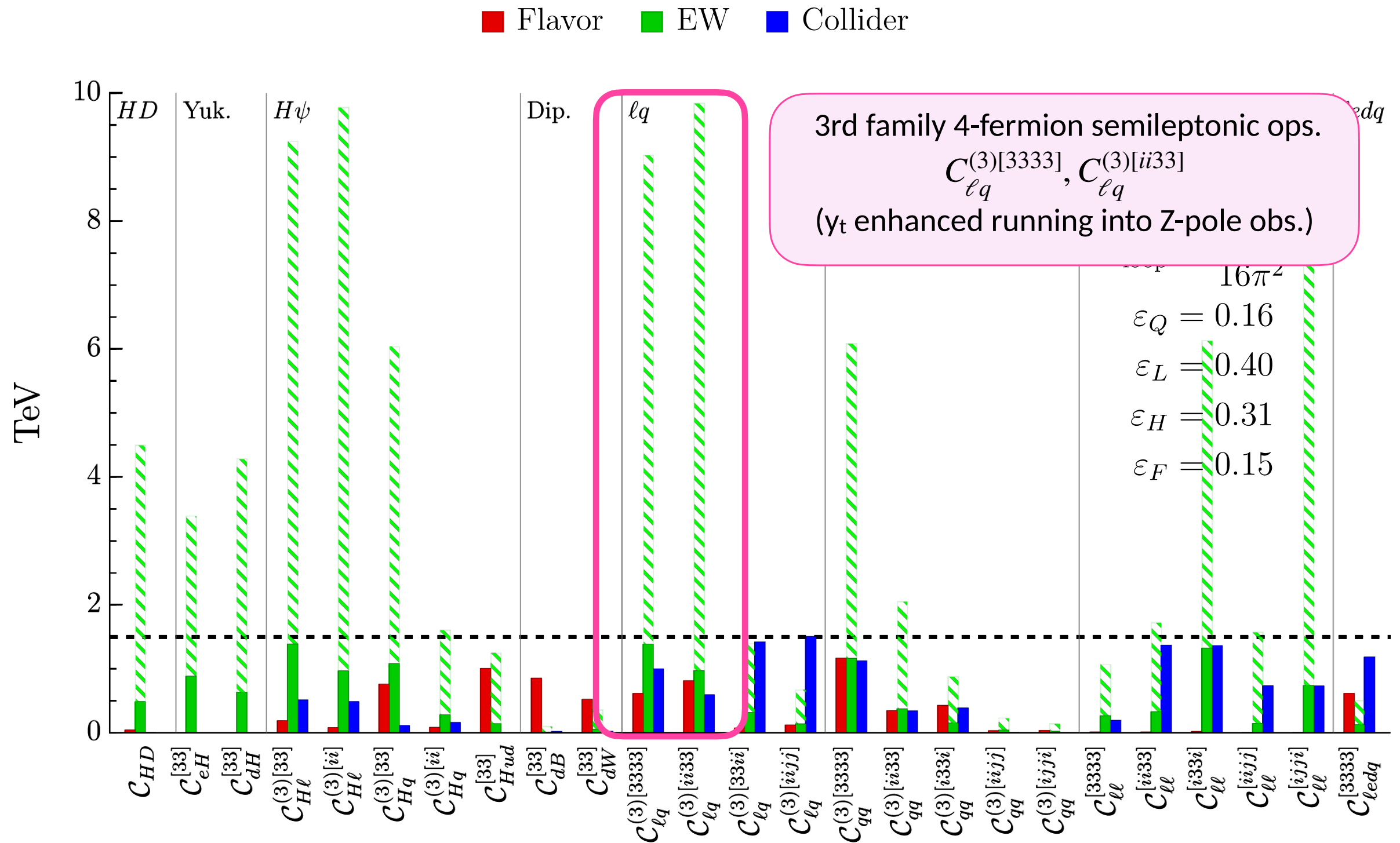




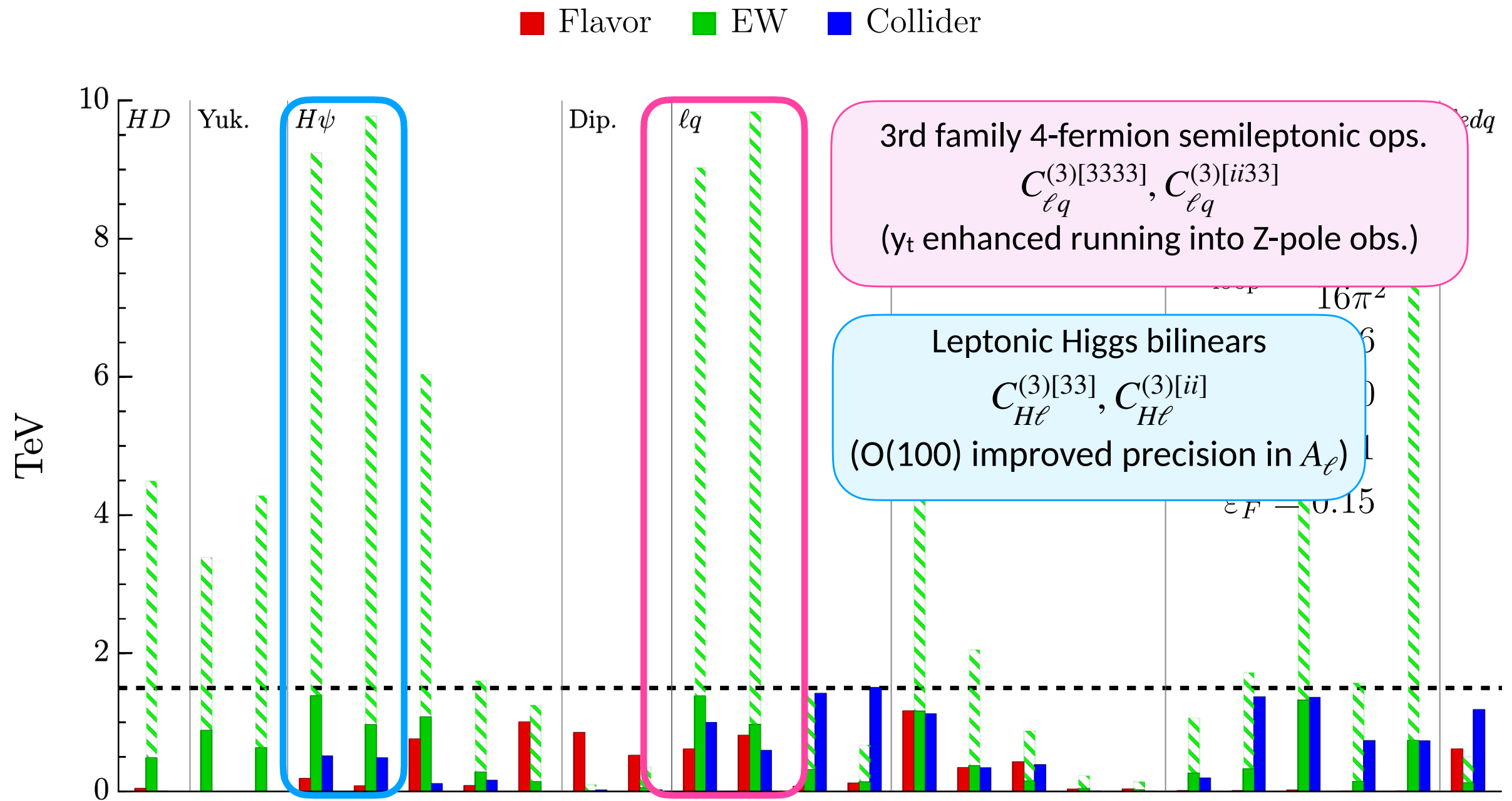
# FCC and 3rd generation NP



# FCC and 3rd generation NP



# FCC and 3rd generation NP



**FCC-ee could probe third-generation New Physics up to ~ 10 TeV!**

# Rare decays and 3rd generation NP

More short-term, improvements in flavor and collider observables can help us probe this scenario. Consider the **rare decays**  $B \rightarrow K\nu\bar{\nu}$  and  $K \rightarrow \pi\nu\bar{\nu}$ .

$$\frac{\mathcal{B}(B^+ \rightarrow K^+ \nu\bar{\nu})_{\text{exp}}}{\mathcal{B}(B^+ \rightarrow K^+ \nu\bar{\nu})_{\text{SM}}} = 2.8 \pm 0.8,$$



[Exp: combination from Belle II @EPS 2023]

~3 $\sigma$  tension with the SM

$$\frac{\mathcal{B}(K^+ \rightarrow \pi^+ \nu\bar{\nu})_{\text{exp}}}{\mathcal{B}(K^+ \rightarrow \pi^+ \nu\bar{\nu})_{\text{SM}}} = 1.23 \pm 0.39$$

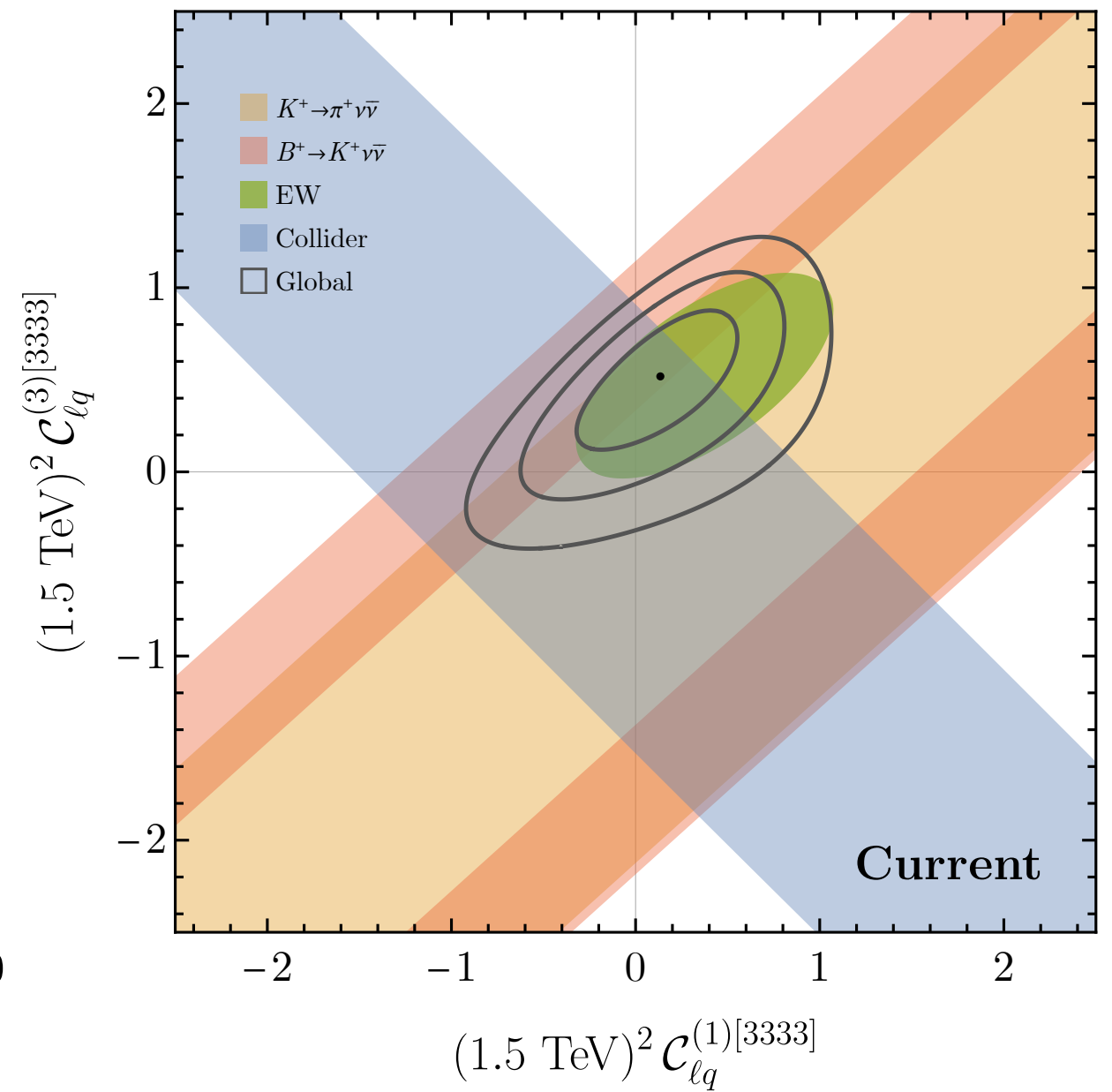
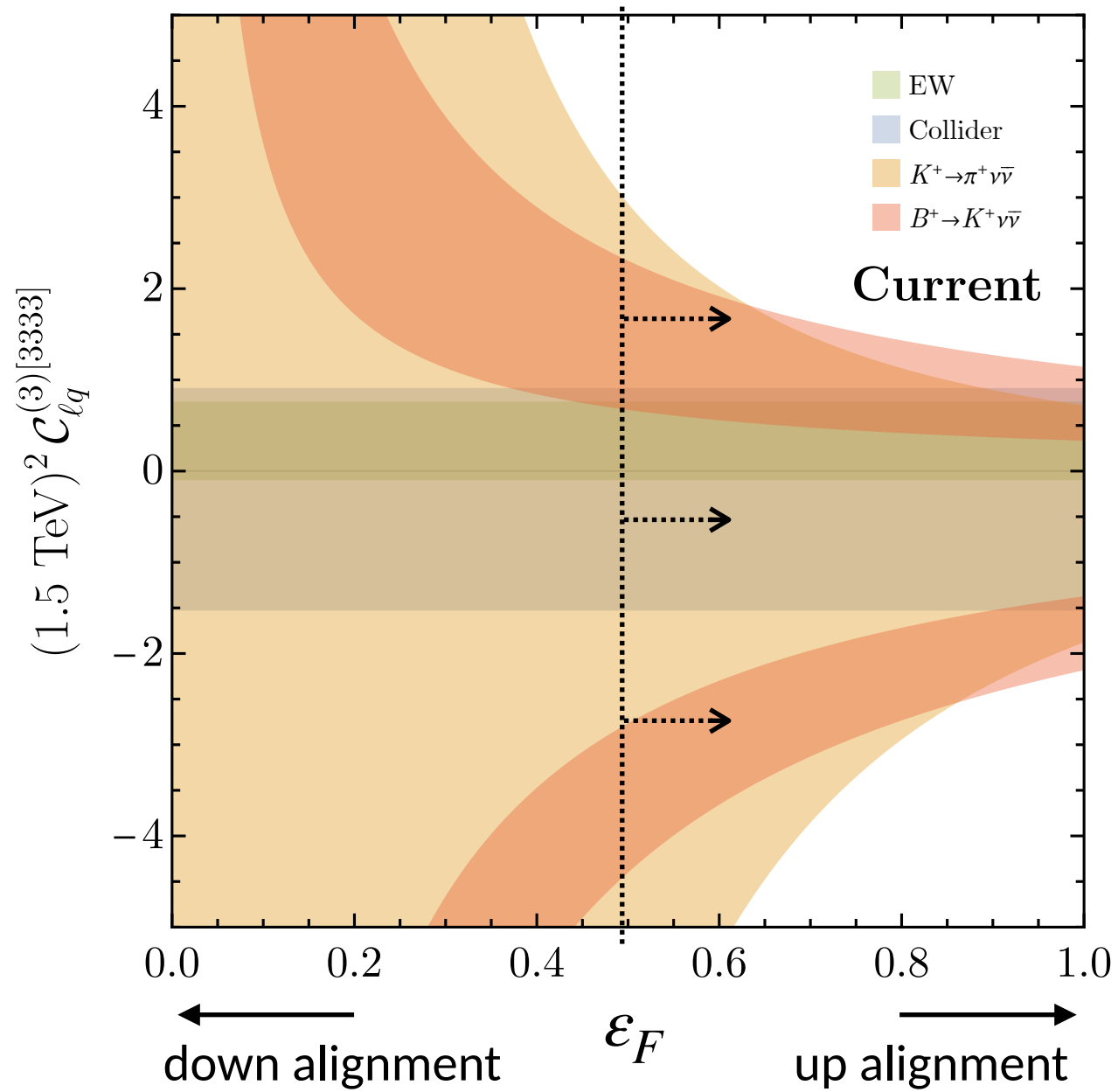
[Exp: NA62 2021; SM: Buras et al. 2015]

Compatible with the SM at 1 $\sigma$

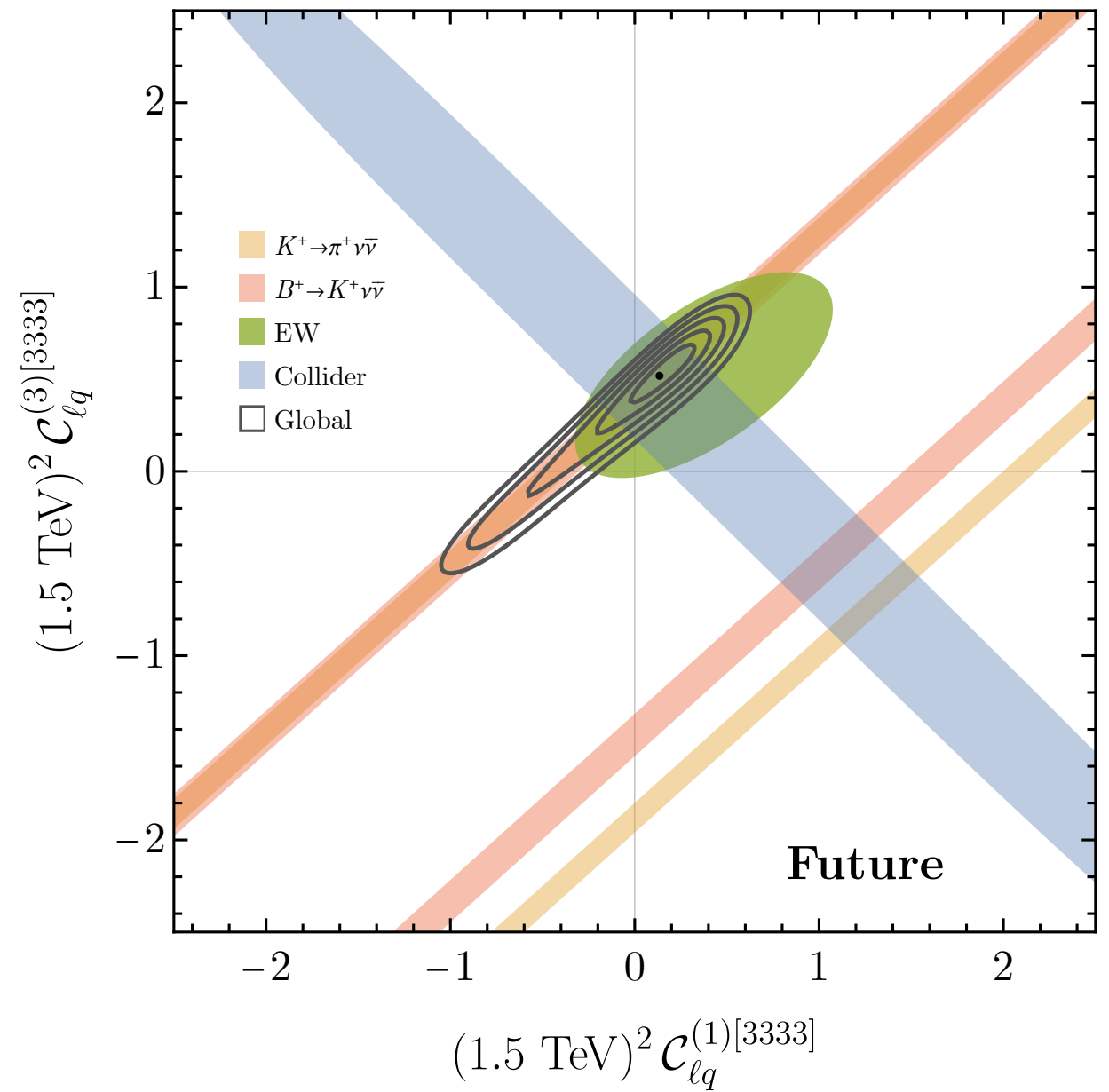
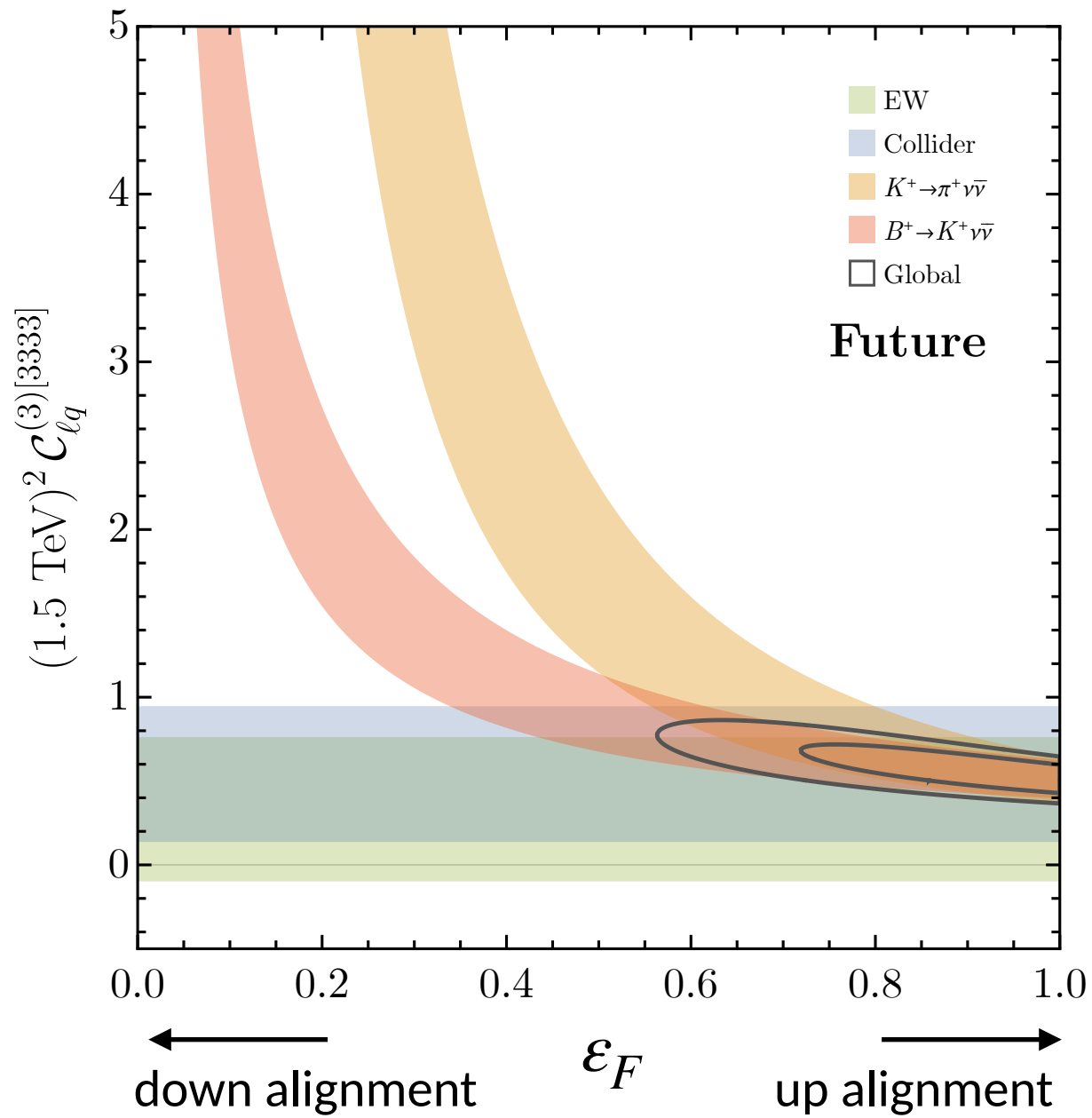


- theoretically clean
- significant improvements expected in the next years:  
Belle II will measure  $B \rightarrow K\nu\bar{\nu}$  @ 10%, and NA62(HIKE)  $K \rightarrow \pi\nu\bar{\nu}$  @ 15%(5%)
- sensitive to a limited number of EFT operators:  $C_{\ell q}^{(3)[3333]}$ ,  $C_{\ell q}^{(1)[3333]}$
- scale differently with the alignment parameter  $\varepsilon_F$

# Rare decays and 3rd generation NP: current data



# Rare decays and 3rd generation NP: projections





# Conclusions

We investigated NP scenarios characterized by a  $U(2)^5$  symmetry acting on the light families. We included EW, flavor, and collider data, and accounted for RG effects.

Our main focus was **NP coupled mostly to the 3rd generation**, because of its strong theoretical motivation.

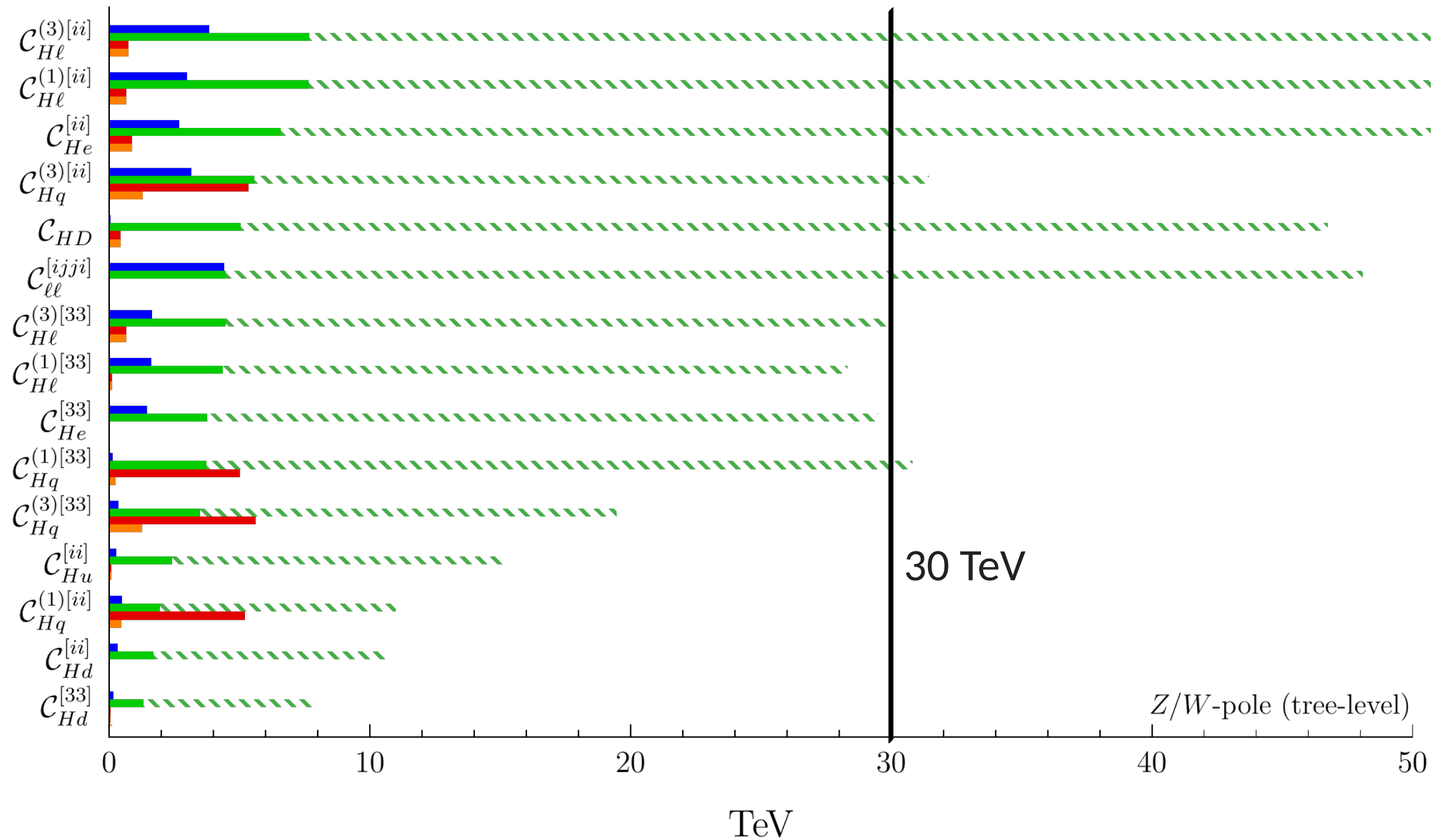
1. **How low** can the energy **scale of new physics** be for these class of models, and which conditions make this possible?
2. **How will the bounds** on these models **change** in the **future**?

1. **NP in the 3rd family is compatible with a scale as low as 1.5 TeV** under simple, non-tuned assumptions. Well-motivated NP models can be nearby!
2. A future tera-Z machine like **FCC-ee** can probe these scenarios up to **10 TeV**. **Precision flavor measurement** can provide complementary information, e.g.  $B \rightarrow K\nu\bar{\nu}$  and  $K \rightarrow \pi\nu\bar{\nu}$  can help determine the flavor alignment.

**Back-up slides**

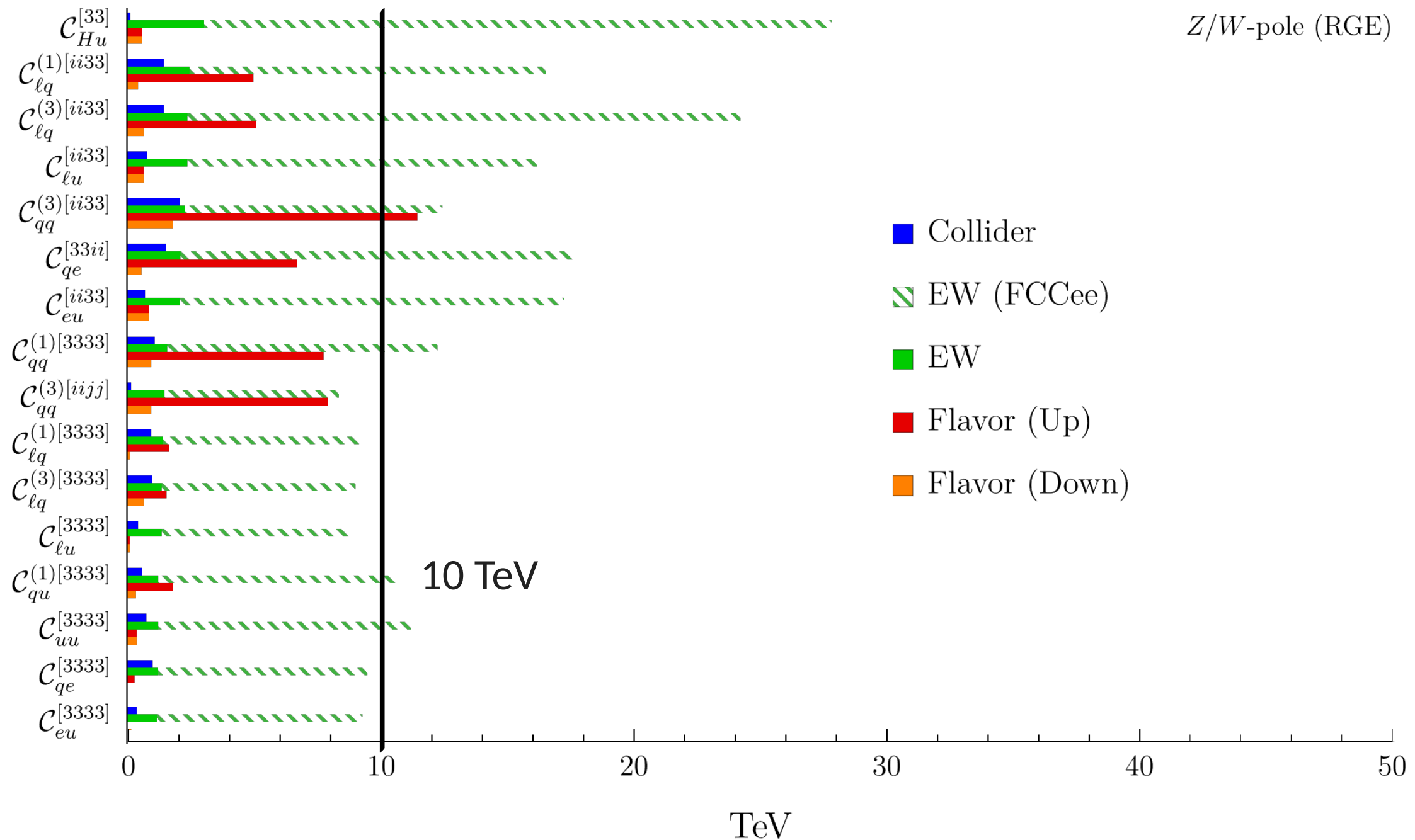
# FCC-ee for “generic” U(2) NP (no suppression factors)

- Operators entering Z-pole observables at tree-level get bounds of 30-50 TeV



# FCC-ee for “generic” U(2) NP (no suppression factors)

- Operators entering Z-pole observables at tree-level get bounds of 30-50 TeV
- 4-fermion operators involving third-family quarks get bounds ~ 10 TeV,



# FCC-ee for “generic” $U(2)$ NP (no suppression factors)

- Operators entering Z-pole observables at tree-level get bounds of 30-50 TeV
- 4-fermion operators involving third-family quarks get bounds  $\sim 10$  TeV,

Two comments:

- A future EW precision machine such as FCC-ee is a great way to probe NP with sizeable couplings to the Higgs
- NP that does not couple directly to the Higgs but does couple to the 3rd generation can be probed up to effective scales of about 10 TeV

**FCC-ee can push most of the existing bounds on NP from the EW sector by one order of magnitude!**

# Higgs bi-fermion operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{H\ell}^{(1)[33]}$	0.1	0.1	4.4	1.6	4.3	$R_\tau$	4.3	$R_\tau$
$\mathcal{C}_{H\ell}^{(1)[ii]}$	0.7	0.7	7.6	3.	7.8	$\sigma_{\text{had}}$	7.8	$\sigma_{\text{had}}$
$\mathcal{C}_{H\ell}^{(3)[33]}$	0.7	0.7	4.5	1.7	4.4	$R_\tau$	4.4	$R_\tau$
$\mathcal{C}_{H\ell}^{(3)[ii]}$	0.7	0.7	7.7	3.8	7.7	$\sigma_{\text{had}}$	7.7	$\sigma_{\text{had}}$
$\mathcal{C}_{He}^{[33]}$	-	-	3.8	1.5	3.7	$R_\tau$	3.7	$R_\tau$
$\mathcal{C}_{He}^{[ii]}$	0.9	0.9	6.6	2.7	6.7	$\sigma_{\text{had}}$	6.7	$\sigma_{\text{had}}$
$\mathcal{C}_{Hq}^{(1)[33]}$	0.3	5.	3.7	0.1	3.7	$\Gamma_Z$	5.1	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{Hq}^{(1)[ii]}$	0.5	5.2	1.9	0.5	2.	$R_c$	5.4	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{Hq}^{(3)[33]}$	1.3	5.6	3.5	0.4	3.4	$R_b$	5.5	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{Hq}^{(3)[ii]}$	1.3	5.3	5.6	3.1	5.7	$R_\tau$	7.7	$\Gamma_Z$
$\mathcal{C}_{Hd}^{[33]}$	-	-	1.3	0.2	1.3	$R_b$	1.3	$R_b$
$\mathcal{C}_{Hd}^{[ii]}$	-	-	1.7	0.3	1.7	$R_\tau$	1.7	$R_\tau$
$\mathcal{C}_{Hu}^{[33]}$	0.6	0.6	3.	0.1	3.1	$A_b^{\text{FB}}$	3.1	$A_b^{\text{FB}}$
$\mathcal{C}_{Hu}^{[ii]}$	-	-	2.4	0.3	2.4	$R_\tau$	2.4	$R_\tau$

# 3H and dipole operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{eH}^{[33]}$	-	-	5.1	-	5.1	$H \rightarrow \tau\tau$	5.1	$H \rightarrow \tau\tau$
$\mathcal{C}_{uH}^{[33]}$	-	-	0.2	-	0.2	$H \rightarrow \tau\tau$	0.2	$H \rightarrow \tau\tau$
$\mathcal{C}_{dH}^{[33]}$	-	-	3.7	-	3.7	$H \rightarrow bb$	3.7	$H \rightarrow bb$
$\mathcal{C}_{Hud}^{[33]}$	3.2	3.2	0.5	-	3.2	$B \rightarrow X_s\gamma$	3.2	$B \rightarrow X_s\gamma$
$\mathcal{C}_{eB}^{[33]}$	-	-	0.2	1.2	1.2	$pp \rightarrow \tau\tau$	1.2	$pp \rightarrow \tau\tau$
$\mathcal{C}_{uB}^{[33]}$	0.7	0.8	2.4	1.9	2.7	$A_b^{\text{FB}}$	2.7	$A_b^{\text{FB}}$
$\mathcal{C}_{dB}^{[33]}$	15.2	74.8	0.4	0.7	15.2	$B \rightarrow X_s\gamma$	74.8	$B \rightarrow X_s\gamma$
$\mathcal{C}_{eW}^{[33]}$	-	-	1.	1.9	1.8	$pp \rightarrow \tau\nu$	1.8	$pp \rightarrow \tau\nu$
$\mathcal{C}_{uW}^{[33]}$	0.5	0.9	2.3	3.6	3.7	QuarkDipoles	3.8	QuarkDipoles
$\mathcal{C}_{dW}^{[33]}$	15.7	53.	1.4	0.6	15.7	$B \rightarrow X_s\gamma$	53.	$B \rightarrow X_s\gamma$
$\mathcal{C}_{uG}^{[33]}$	0.1	0.3	0.5	2.7	2.7	QuarkDipoles	2.7	QuarkDipoles
$\mathcal{C}_{dG}^{[33]}$	4.	25.5	0.3	-	4.	$B \rightarrow X_s\gamma$	25.5	$B \rightarrow X_s\gamma$



# Scalar and tensor operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{ledq}^{[3333]}$	0.6	-	0.1	1.2	1.1	$pp \rightarrow \tau\tau$	1.2	$pp \rightarrow \tau\tau$
$\mathcal{C}_{quqd}^{(1)[3333]}$	1.8	5.5	1.7	0.4	2.2	$B \rightarrow X_s \gamma$	5.5	$B \rightarrow X_s \gamma$
$\mathcal{C}_{quqd}^{(8)[3333]}$	1.	5.1	0.7	0.2	1.	$B \rightarrow X_s \gamma$	5.1	$B \rightarrow X_s \gamma$
$\mathcal{C}_{lequ}^{(1)[3333]}$	-	-	2.1	-	2.1	$H \rightarrow \tau\tau$	2.1	$H \rightarrow \tau\tau$
$\mathcal{C}_{lequ}^{(3)[3333]}$	-	-	0.8	-	0.8	$H \rightarrow \tau\tau$	0.8	$H \rightarrow \tau\tau$

# LLLL vector operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{\ell\ell}^{[3333]}$	-	-	0.3	0.2	0.3	$\sigma_{\text{had}}$	0.3	$\sigma_{\text{had}}$
$\mathcal{C}_{\ell\ell}^{[ii33]}$	-	-	0.8	3.4	3.3	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	3.3	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell\ell}^{[i33i]}$	-	-	3.3	3.3	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell\ell}^{[iijj]}$	-	-	0.9	4.4	4.4	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	4.4	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell\ell}^{[ijji]}$	-	-	4.5	4.4	4.9	$A_b^{\text{FB}}$	4.9	$A_b^{\text{FB}}$
$\mathcal{C}_{qq}^{(1)[3333]}$	1.	7.8	1.6	1.1	1.7	$\Gamma_Z$	7.6	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[ii33]}$	1.3	11.2	0.9	1.5	1.7	FourQuarksTop	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[i33i]}$	2.5	11.3	0.7	1.6	2.6	$B_s \rightarrow \mu\mu$	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[iijj]}$	0.9	8.1	0.4	-	0.9	$\text{Im}(C_D)$	8.1	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[ijji]}$	1.1	8.1	0.5	-	1.	$\text{Im}(C_D)$	8.1	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[3333]}$	1.	8.2	1.2	1.1	1.5	$m_W$	8.2	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[ii33]}$	1.8	11.5	2.3	2.1	3.	$R_b$	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[i33i]}$	2.6	11.2	0.9	2.4	3.1	$B_s \rightarrow \mu\mu$	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[iijj]}$	1.	7.9	1.5	0.2	1.5	$R_\tau$	7.9	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[ijji]}$	1.1	8.	0.9	0.1	1.2	$K^+ \rightarrow \pi^+\nu\bar{\nu}$	8.	$ C_{B_s} $
$\mathcal{C}_{\ell q}^{(1)[3333]}$	0.1	1.7	1.4	1.	1.4	$R_\tau$	1.6	$K^+ \rightarrow \pi^+\nu\bar{\nu}$
$\mathcal{C}_{\ell q}^{(1)[ii33]}$	0.4	5.	2.5	1.5	2.5	$\sigma_{\text{had}}$	5.1	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{\ell q}^{(1)[33ii]}$	-	1.6	0.3	3.4	3.4	$pp \rightarrow \tau\tau$	3.4	$pp \rightarrow \tau\tau$
$\mathcal{C}_{\ell q}^{(1)[iijj]}$	0.5	5.	0.5	5.4	5.4	$pp \rightarrow \mu\mu$	5.6	$pp \rightarrow \mu\mu$
$\mathcal{C}_{\ell q}^{(3)[3333]}$	0.7	1.5	1.4	1.	1.6	$R_\tau$	1.6	$K^+ \rightarrow \pi^+\nu\bar{\nu}$
$\mathcal{C}_{\ell q}^{(3)[ii33]}$	0.7	5.1	2.4	1.5	2.5	$A_b^{\text{FB}}$	5.	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{\ell q}^{(3)[33ii]}$	0.1	1.4	2.	8.6	8.8	$pp \rightarrow \tau\nu$	8.7	$pp \rightarrow \tau\nu$
$\mathcal{C}_{\ell q}^{(3)[iijj]}$	0.5	5.1	2.1	22.5	22.5	$pp \rightarrow \mu\nu$	23.7	$pp \rightarrow \mu\nu$

# RRRR vector operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{ee}^{[3333]}$	-	-	0.3	0.2	0.3	$R_\tau$	0.3	$R_\tau$
$\mathcal{C}_{ee}^{[ii33]}$	-	-	0.7	3.2	3.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	3.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{ee}^{[ijjj]}$	-	-	0.8	4.2	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{uu}^{[3333]}$	0.4	0.4	1.2	0.8	1.3	$A_b^{\text{FB}}$	1.3	$A_b^{\text{FB}}$
$\mathcal{C}_{uu}^{[ii33]}$	0.1	0.1	1.1	1.3	1.4	FourQuarksTop	1.4	FourQuarksTop
$\mathcal{C}_{uu}^{[i33i]}$	-	-	0.5	1.3	1.4	FourQuarksTop	1.4	FourQuarksTop
$\mathcal{C}_{uu}^{[ijjj]}$	-	-	0.3	-	0.3	$R_\tau$	0.3	$R_\tau$
$\mathcal{C}_{uu}^{[ijji]}$	-	-	0.3	-	0.3	$R_\tau$	0.3	$R_\tau$
$\mathcal{C}_{dd}^{[3333]}$	-	-	-	-	-	$R_b$	-	$R_b$
$\mathcal{C}_{dd}^{[ii33]}$	-	-	0.1	-	0.1	$R_\tau$	0.1	$R_\tau$
$\mathcal{C}_{dd}^{[i33i]}$	-	-	-	-	-	$\Gamma_Z$	-	$\Gamma_Z$
$\mathcal{C}_{dd}^{[ijjj]}$	-	-	0.2	-	0.2	$R_\tau$	0.2	$R_\tau$
$\mathcal{C}_{dd}^{[ijji]}$	-	-	0.1	-	0.1	$R_\tau$	0.1	$R_\tau$
$\mathcal{C}_{eu}^{[3333]}$	-	-	1.2	0.4	1.2	$R_\tau$	1.2	$R_\tau$
$\mathcal{C}_{eu}^{[ii33]}$	0.9	0.9	2.1	0.7	2.2	$\sigma_{\text{had}}$	2.2	$\sigma_{\text{had}}$
$\mathcal{C}_{eu}^{[33ii]}$	-	-	0.3	2.8	2.8	$pp \rightarrow \tau\tau$	2.8	$pp \rightarrow \tau\tau$
$\mathcal{C}_{eu}^{[ijjj]}$	-	-	0.6	7.4	7.4	$pp \rightarrow ee$	7.4	$pp \rightarrow ee$
$\mathcal{C}_{ed}^{[3333]}$	-	-	0.2	1.	1.	$pp \rightarrow \tau\tau$	1.	$pp \rightarrow \tau\tau$
$\mathcal{C}_{ed}^{[ii33]}$	-	-	0.3	1.5	1.5	$pp \rightarrow \mu\mu$	1.5	$pp \rightarrow \mu\mu$
$\mathcal{C}_{ed}^{[33ii]}$	-	-	0.2	2.8	2.8	$pp \rightarrow \tau\tau$	2.8	$pp \rightarrow \tau\tau$
$\mathcal{C}_{ed}^{[ijjj]}$	-	-	0.4	4.4	4.4	$pp \rightarrow \mu\mu$	4.4	$pp \rightarrow \mu\mu$
$\mathcal{C}_{ud}^{(1)[3333]}$	0.1	0.1	0.4	0.3	0.4	$R_b$	0.4	$R_b$
$\mathcal{C}_{ud}^{(1)[ii33]}$	-	-	0.1	-	0.1	$R_\tau$	0.1	$R_\tau$
$\mathcal{C}_{ud}^{(1)[33ii]}$	-	-	0.5	1.2	1.2	FourQuarksTop	1.2	FourQuarksTop
$\mathcal{C}_{ud}^{(1)[ijjj]}$	-	-	0.2	-	0.2	$R_\tau$	0.2	$R_\tau$
$\mathcal{C}_{ud}^{(8)[3333]}$	0.1	0.1	-	0.2	0.2	FourQuarksBottom	0.2	FourQuarksBottom
$\mathcal{C}_{ud}^{(8)[ii33]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{ud}^{(8)[33ii]}$	-	-	0.1	0.7	0.7	FourQuarksTop	0.7	FourQuarksTop
$\mathcal{C}_{ud}^{(8)[ijjj]}$	-	-	-	-	-	-	-	-

# LLRR vector operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$C_{le}^{[3333]}$	-	-	0.2	0.1	0.2	$A_\tau$	0.2	$A_\tau$
$C_{le}^{[ii33]}$	-	-	0.4	2.	1.9	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	1.9	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$C_{le}^{[33ii]}$	-	-	0.3	1.9	2.	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	2.	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$C_{le}^{[iijj]}$	-	-	0.5	3.8	3.8	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	3.8	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$C_{lu}^{[3333]}$	0.1	0.1	1.4	0.4	1.3	$R_\tau$	1.3	$R_\tau$
$C_{lu}^{[ii33]}$	0.7	0.7	2.4	0.8	2.3	$\sigma_{\text{had}}$	2.3	$\sigma_{\text{had}}$
$C_{lu}^{[33ii]}$	-	-	0.4	3.1	3.1	$pp \rightarrow \tau\tau$	3.1	$pp \rightarrow \tau\tau$
$C_{lu}^{[iijj]}$	-	-	0.7	5.2	5.2	$pp \rightarrow \mu\mu$	5.2	$pp \rightarrow \mu\mu$
$C_{ld}^{[3333]}$	-	-	0.2	1.	1.	$pp \rightarrow \tau\tau$	1.	$pp \rightarrow \tau\tau$
$C_{ld}^{[ii33]}$	-	-	0.3	1.5	1.5	$pp \rightarrow \mu\mu$	1.5	$pp \rightarrow \mu\mu$
$C_{ld}^{[33ii]}$	-	-	0.3	3.	3.	$pp \rightarrow \tau\tau$	3.	$pp \rightarrow \tau\tau$
$C_{ld}^{[iijj]}$	-	-	0.5	4.7	4.7	$pp \rightarrow \mu\mu$	4.7	$pp \rightarrow \mu\mu$
$C_{qe}^{[3333]}$	-	0.3	1.2	1.	1.3	$R_\tau$	1.2	$R_\tau$
$C_{qe}^{[33ii]}$	0.6	6.7	2.1	1.5	2.2	$\sigma_{\text{had}}$	6.7	$B_s \rightarrow \mu\mu$
$C_{qe}^{[ii33]}$	-	0.3	0.2	3.7	3.7	$pp \rightarrow \tau\tau$	3.7	$pp \rightarrow \tau\tau$
$C_{qe}^{[iijj]}$	-	-	0.4	6.	6.	$pp \rightarrow \mu\mu$	6.	$pp \rightarrow \mu\mu$
$C_{qu}^{(1)[3333]}$	0.3	1.8	1.2	0.6	1.3	$\Gamma_Z$	1.7	$B_s \rightarrow \mu\mu$
$C_{qu}^{(1)[ii33]}$	0.3	1.8	0.6	1.6	1.6	FourQuarksTop	2.1	$B_s \rightarrow \mu\mu$
$C_{qu}^{(1)[33ii]}$	-	0.6	0.8	1.4	1.4	FourQuarksTop	1.2	FourQuarksTop
$C_{qu}^{(1)[iijj]}$	-	0.6	0.2	-	0.2	$R_\tau$	0.6	$ C_{Bd} $
$C_{qu}^{(8)[3333]}$	0.2	0.7	0.1	0.4	0.4	FourQuarksTop	0.7	$ C_{Bs} $
$C_{qu}^{(8)[ii33]}$	0.3	0.7	0.1	1.2	1.2	FourQuarksTop	1.2	FourQuarksTop
$C_{qu}^{(8)[33ii]}$	-	0.1	0.2	0.8	0.8	FourQuarksTop	0.8	FourQuarksTop
$C_{qu}^{(8)[iijj]}$	-	0.1	-	-	-	$R_\tau$	0.1	$C_9^U$
$C_{qd}^{(1)[3333]}$	0.2	0.3	0.4	0.3	0.3	$R_b$	0.3	$R_b$
$C_{qd}^{(1)[ii33]}$	-	0.3	0.1	-	-	$R_\tau$	0.3	$B_s \rightarrow \mu\mu$
$C_{qd}^{(1)[33ii]}$	-	0.4	0.6	1.3	1.2	FourQuarksTop	1.1	FourQuarksTop
$C_{qd}^{(1)[iijj]}$	-	0.4	0.2	-	0.2	$R_\tau$	0.4	$B_s \rightarrow \mu\mu$
$C_{qd}^{(8)[3333]}$	-	-	-	0.2	0.2	FourQuarksBottom	0.2	FourQuarksBottom
$C_{qd}^{(8)[ii33]}$	0.1	-	-	-	0.1	$B \rightarrow X_s\gamma$	-	$B \rightarrow X_s\gamma$
$C_{qd}^{(8)[33ii]}$	-	-	0.1	0.7	0.7	FourQuarksTop	0.7	FourQuarksTop
$C_{qd}^{(8)[iijj]}$	-	-	-	-	-	$R_\tau$	-	$ C_{Bs} $

# Bosonic operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	$\Lambda_{\text{EW}}$	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_H$	-	-	-	-	-	-	-	-
$\mathcal{C}_{H\Box}$	0.2	0.2	0.6	0.1	0.6	$A_b^{\text{FB}}$	0.6	$A_b^{\text{FB}}$
$\mathcal{C}_{HD}$	0.5	0.5	5.1	-	5.	$A_b^{\text{FB}}$	5.	$A_b^{\text{FB}}$
$\mathcal{C}_{HG}$	0.8	0.8	0.4	-	0.9	$B \rightarrow X_s \gamma$	0.9	$B \rightarrow X_s \gamma$
$\mathcal{C}_{HB}$	0.5	0.5	0.9	-	0.9	$A_b^{\text{FB}}$	0.9	$A_b^{\text{FB}}$
$\mathcal{C}_{HW}$	0.7	0.7	0.9	-	1.	$A_b^{\text{FB}}$	1.	$A_b^{\text{FB}}$
$\mathcal{C}_{HWB}$	1.	1.	9.	-	9.	$A_b^{\text{FB}}$	9.	$A_b^{\text{FB}}$
$\mathcal{C}_G$	1.1	1.1	0.1	-	1.1	$B \rightarrow X_s \gamma$	1.1	$B \rightarrow X_s \gamma$
$\mathcal{C}_W$	0.3	0.3	0.9	-	0.9	$A_b^{\text{FB}}$	0.9	$A_b^{\text{FB}}$