

Recent advances in axion thermal production

La Thuile 2024

Luca Di Luzio



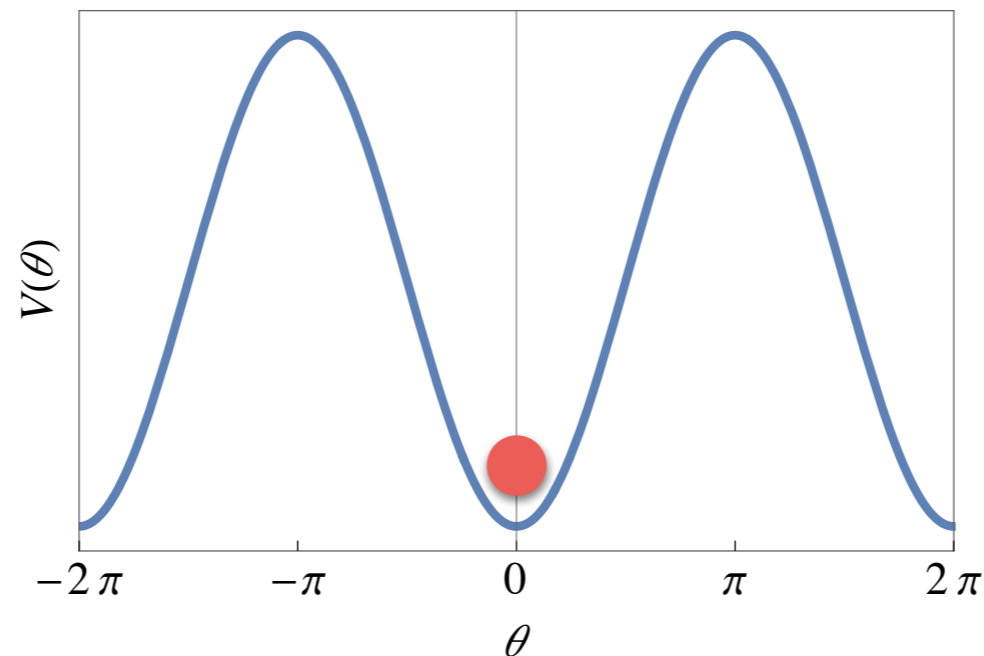
The QCD axion

- Introduced to address the strong CP problem

[Peccei, Quinn '77, Weinberg '78, Wilczek '78]

$$\delta\mathcal{L}_{\text{QCD}} = \theta \frac{g_s^2}{32\pi^2} G\tilde{G} \quad |\theta| \lesssim 10^{-10}$$

- promote θ to a dynamical field (**axion**): $\theta \rightarrow \frac{a}{f_a}$
- it acquires a potential and relaxes dynamically to zero



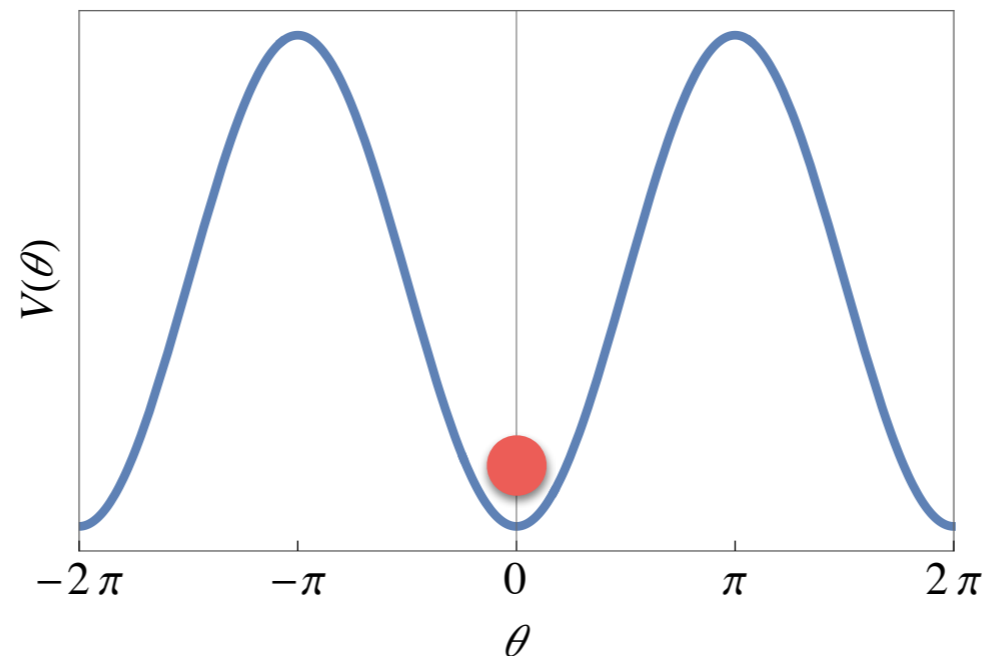
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- Introduced to address the strong CP problem [Peccei, Quinn '77, Weinberg '78, Wilczek '78]
- Unavoidably contributes to the energy density of the universe

Ω_{DM} (non-thermal production)

i) misalignment mechanism (axion oscillations)

[Preskill, Wise, Wilczek '83,
Abbott, Sikivie '83,
Dine, Fischler '83]



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Ω_{DM} (non-thermal production)

i) *misalignment mechanism (axion oscillations)*

[Preskill, Wise, Wilczek '83,
Abbott, Sikivie '83,
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ii) *topological defects (axion strings, ...)*

[Davies '86, Harari Sikivie '87, ...]



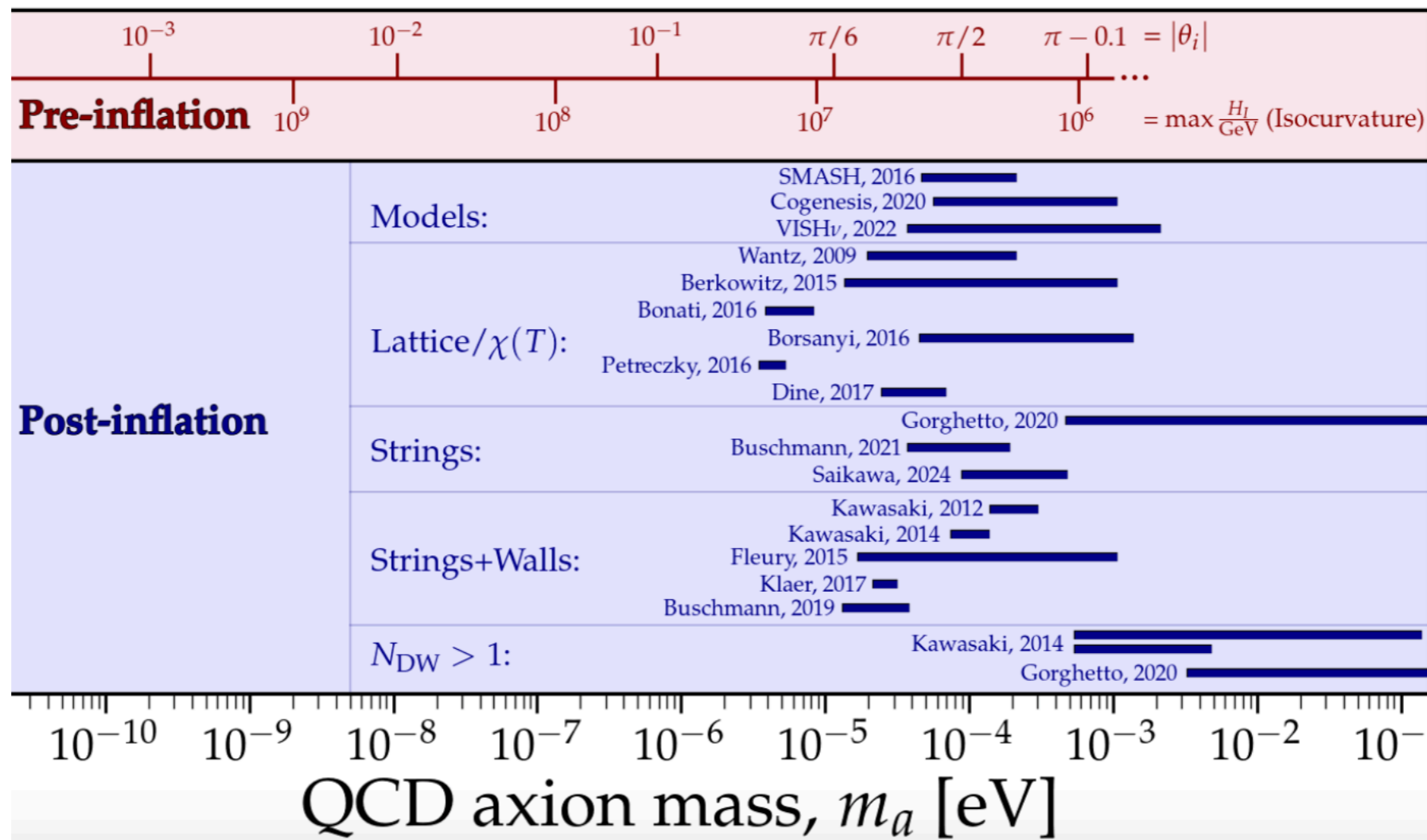
absent if PQ symmetry is broken before inflation (Pre-inflationary)

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[From <https://cajohare.github.io/AxionLimits>]




The QCD axion

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Ω_{DM} (non-thermal production)

Ω_{rad} (thermal production) [Turner '87, ...]

$$\rho_{\text{rad}} = \rho_{\gamma} + \rho_{\nu} + \rho_a = \left[1 + \frac{7}{8} \left(\frac{T_{\nu}}{T_{\gamma}} \right)^4 N_{\text{eff}}^{\text{SM}} + \frac{1}{2} \left(\frac{T_a}{T_{\gamma}} \right)^4 \right] \rho_{\gamma} \equiv \left[1 + \frac{7}{8} \left(\frac{T_{\nu}}{T_{\gamma}} \right)^4 N_{\text{eff}} \right] \rho_{\gamma}$$

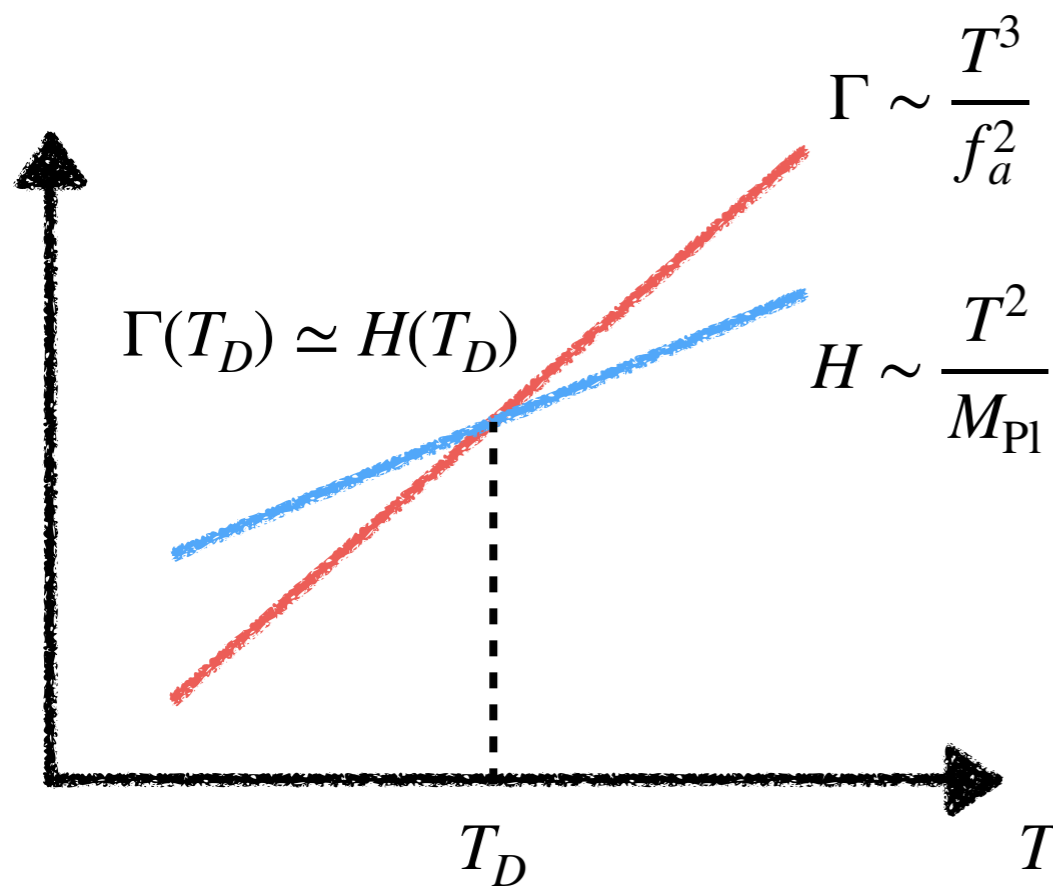

$$\Delta N_{\text{eff}} \equiv N_{\text{eff}} - N_{\text{eff}}^{\text{SM}} = \frac{4}{7} \left(\frac{T_a}{T_{\nu}} \right)^4 \simeq 0.027 \left(\frac{106.75}{g_S(T_D)} \right)^{4/3}$$

Axion thermal production

- Axions can be thermally produced in the early universe

$$\mathcal{L}_a = \frac{1}{2}(\partial_\mu a)^2 + \frac{a}{f_a} \frac{\alpha_s}{8\pi} G\tilde{G} + \dots$$

$$T_D \sim \frac{f_a^2}{M_p} \sim \Lambda_{\text{QCD}} \left(\frac{f_a}{10^8 \text{ GeV}} \right)^2$$

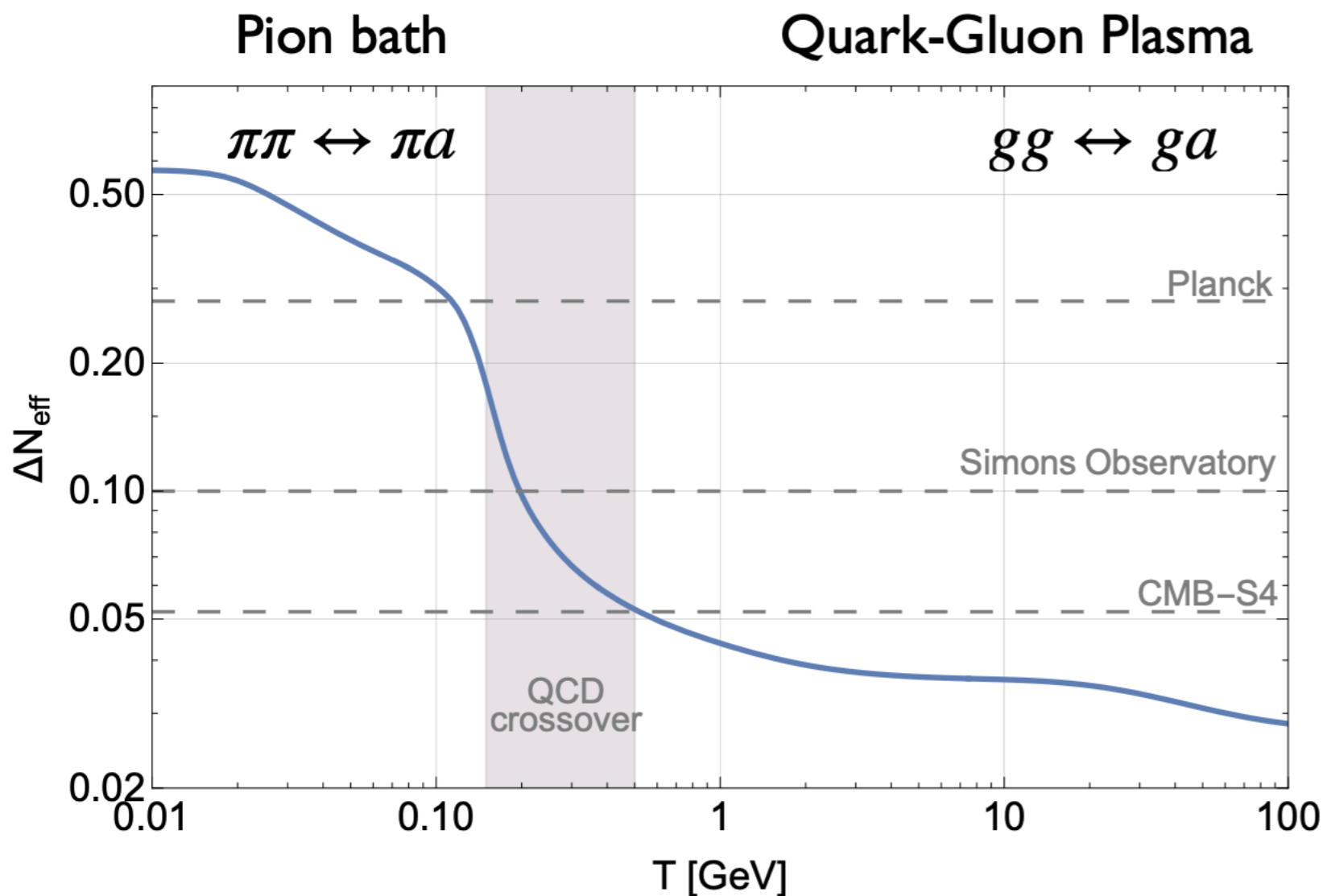


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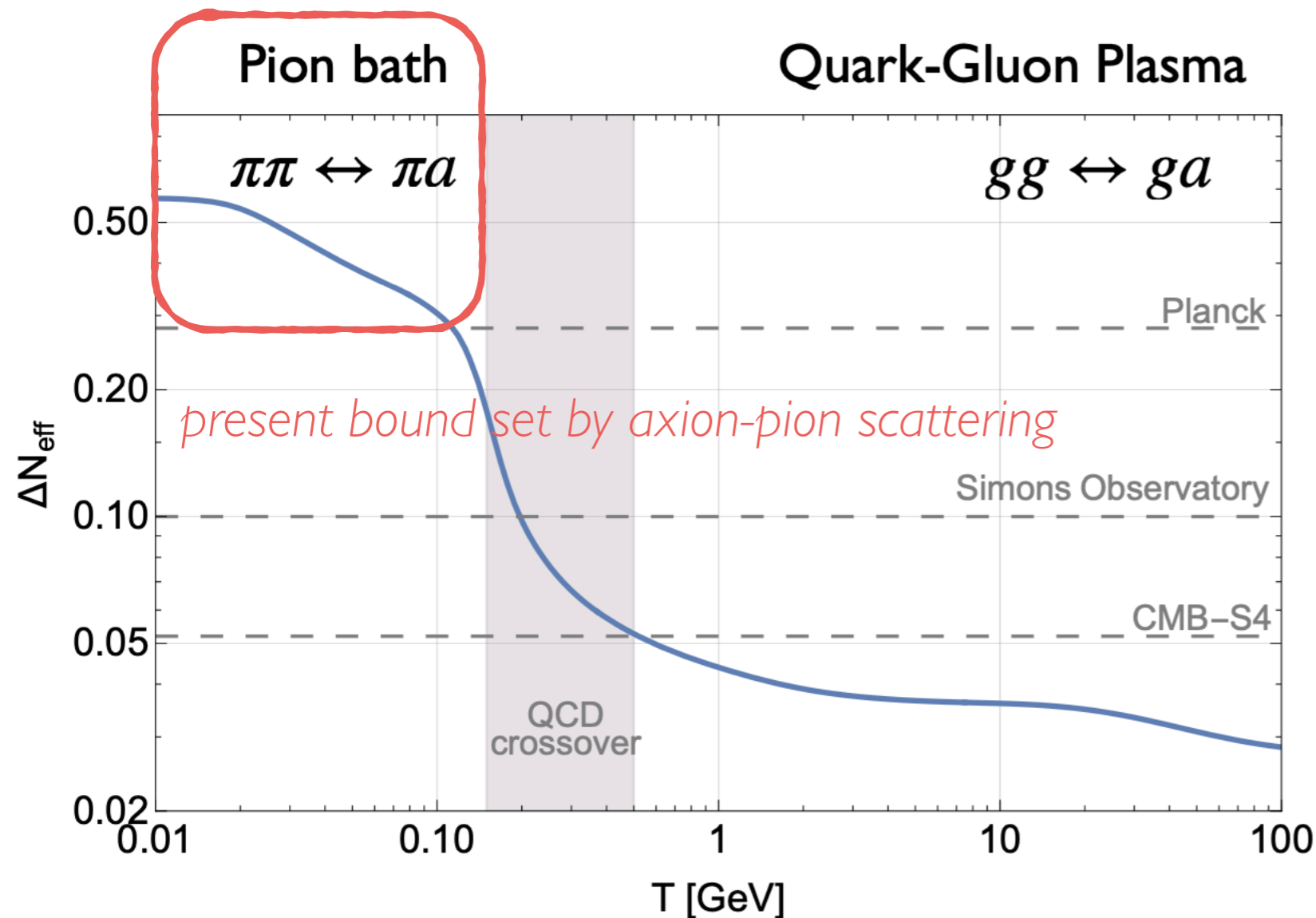
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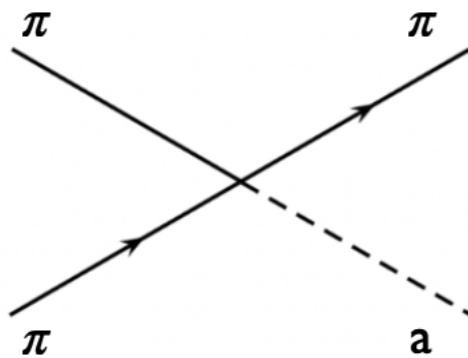
Axion-pion scattering

- Common lore circa 2021

$$\mathcal{L}_a = \frac{1}{2}(\partial_\mu a)^2 + \frac{a}{f_a} \frac{\alpha_s}{8\pi} G\tilde{G} + \dots \quad \longrightarrow \quad \frac{C_{a\pi}}{f_a f_\pi} \partial a \partial \pi \pi \pi \quad C_{a\pi} = \frac{1}{3} \frac{m_d - m_u}{m_u + m_d} + \dots$$

$$\sum |\mathcal{M}|_{\text{LO}}^2 = \left(\frac{C_{a\pi}}{f_a f_\pi} \right)^2 \frac{9}{4} [s^2 + t^2 + u^2 - 3m_\pi^4]$$

[Chang, Choi - hep-ph/9306216
Hannestad, Mirizzi, Raffelt - hep-ph/0504059]



$$\begin{aligned} \Gamma_a &= \frac{1}{n_a^{\text{eq}}} \int \frac{d^3 \mathbf{p}_1}{(2\pi)^3 2E_1} \frac{d^3 \mathbf{p}_2}{(2\pi)^3 2E_2} \frac{d^3 \mathbf{p}_3}{(2\pi)^3 2E_3} \frac{d^3 \mathbf{p}_4}{(2\pi)^3 2E_4} \\ &\times \sum |\mathcal{M}|^2 (2\pi)^4 \delta^4(p_1 + p_2 - p_3 - p_4) \\ &\times f_1 f_2 (1 + f_3)(1 + f_4) \\ &= \left(\frac{C_{a\pi}}{f_a f_\pi} \right)^2 0.212 T^5 [h_{\text{LO}}(m_\pi/T)] \end{aligned}$$

Axion-pion scattering

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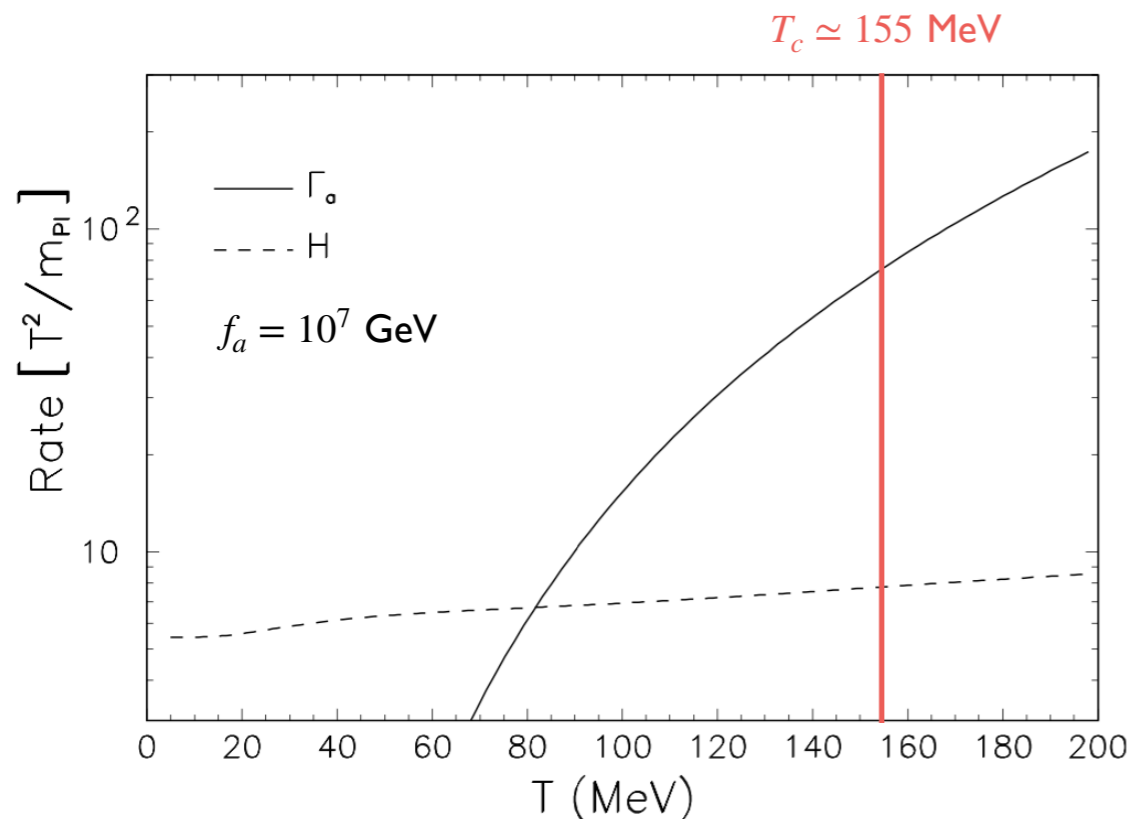
$$\mathcal{L}_a = \frac{1}{2}(\partial_\mu a)^2 + \frac{a}{f_a} \frac{\alpha_s}{8\pi} G\tilde{G} + \dots$$



$$\frac{C_{a\pi}}{f_a f_\pi} \partial a \partial \pi \pi \pi$$

$$C_{a\pi} = \frac{1}{3} \frac{m_d - m_u}{m_u + m_d} + \dots$$

[Chang, Choi - hep-ph/9306216
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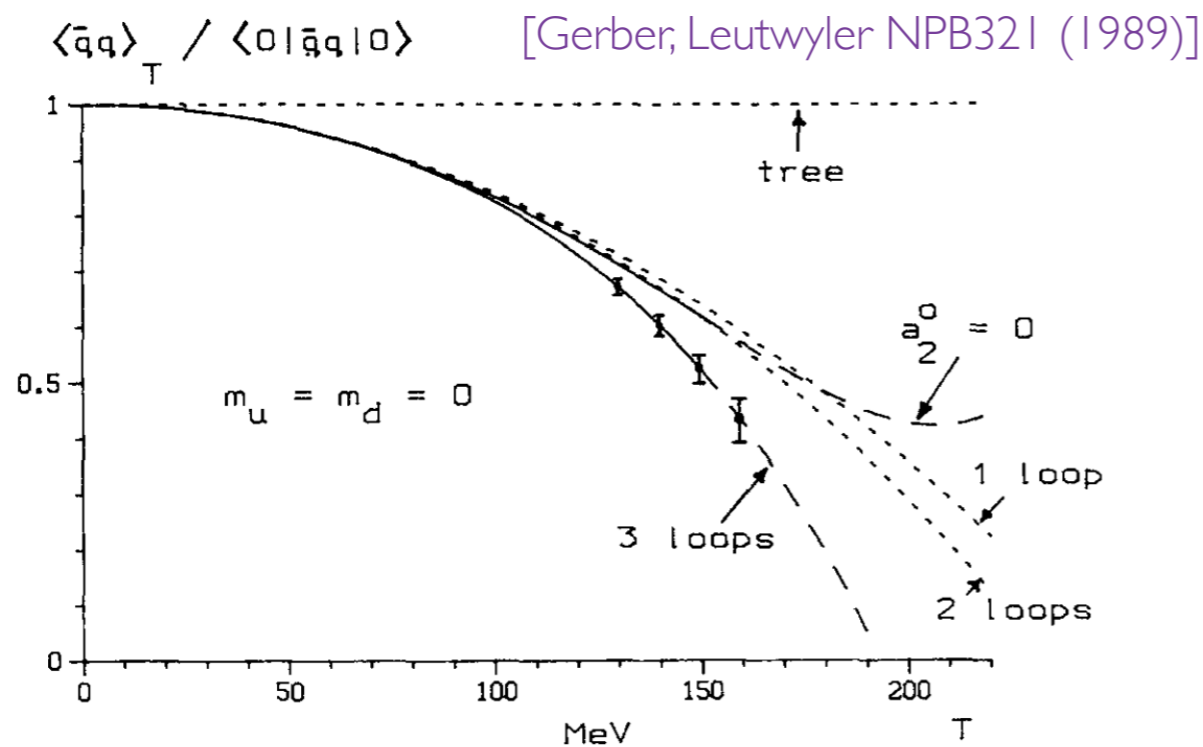


f_a [GeV]	T_D [MeV]	$g_*(T_D)$	n_a [cm^{-3}]
3×10^3	13.37	10.84	74.14
1×10^4	15.30	10.93	73.50
3×10^4	17.63	11.10	72.39
1×10^5	21.21	11.46	70.11
3×10^5	26.06	12.06	66.63
1×10^6	34.75	13.15	61.08
3×10^6	49.12	14.54	55.24
1×10^7	81.61	16.43	48.88
3×10^7	145.31	21.10	38.08

ChiPT breakdown @ finite T

- Qualitative argument #1: melting of chiral condensate

$$\langle \bar{q}q \rangle = \langle 0 | \bar{q}q | 0 \rangle \left(1 - \frac{T^2}{8F^2} - \frac{T^4}{384F^4} - \frac{T^6}{288F^6} \ln \frac{\Lambda_q}{T} + O(T^8) \right)$$

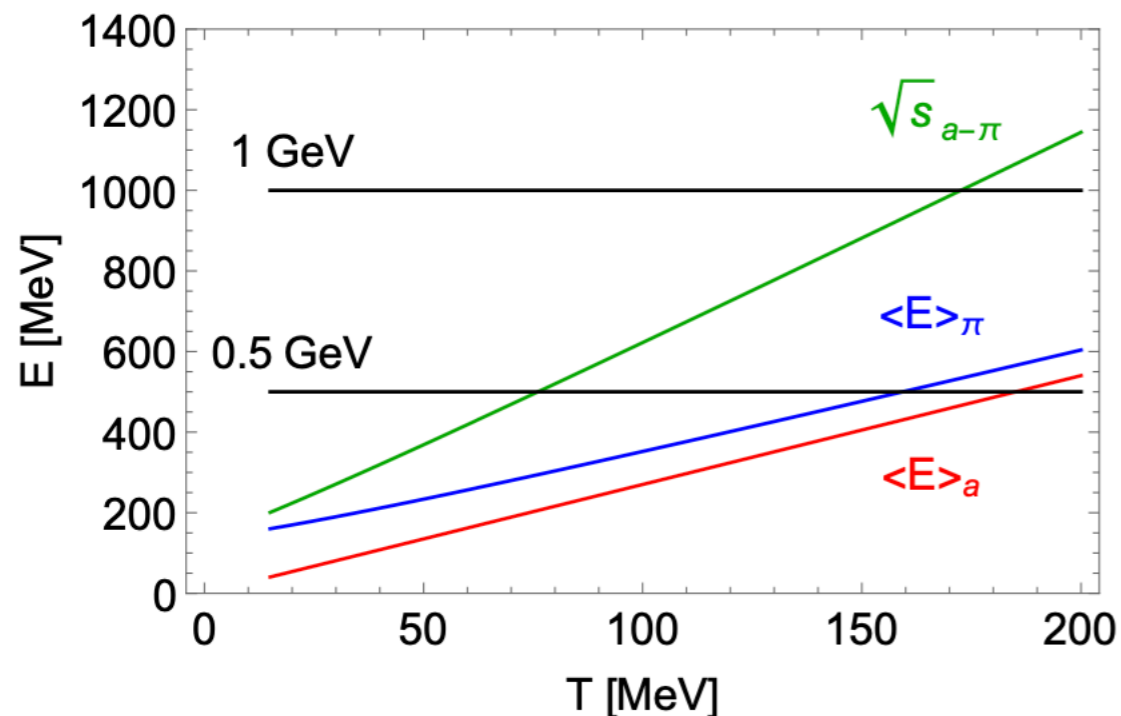


ChiPT breaks down much earlier than T of chiral symmetry restoration

ChiPT breakdown @ finite T

- Qualitative argument #2: scattering energy in the thermal bath

$$\langle E_a \rangle \sim \frac{\rho_a(T)}{n_a(T)} = \frac{\frac{\pi^2}{30} T^4}{\frac{\zeta(3)}{\pi^2} T^3} \simeq 2.7 T \quad \langle E_\pi \rangle \sim \frac{\rho_\pi(T)}{n_\pi(T)} = \dots$$



→ $\sqrt{s_{a-\pi}} \sim \langle E_\pi \rangle + \langle E_a \rangle \simeq 500 \text{ MeV} @ T \simeq 70 \text{ MeV}$

Axion-pion scattering at NLO

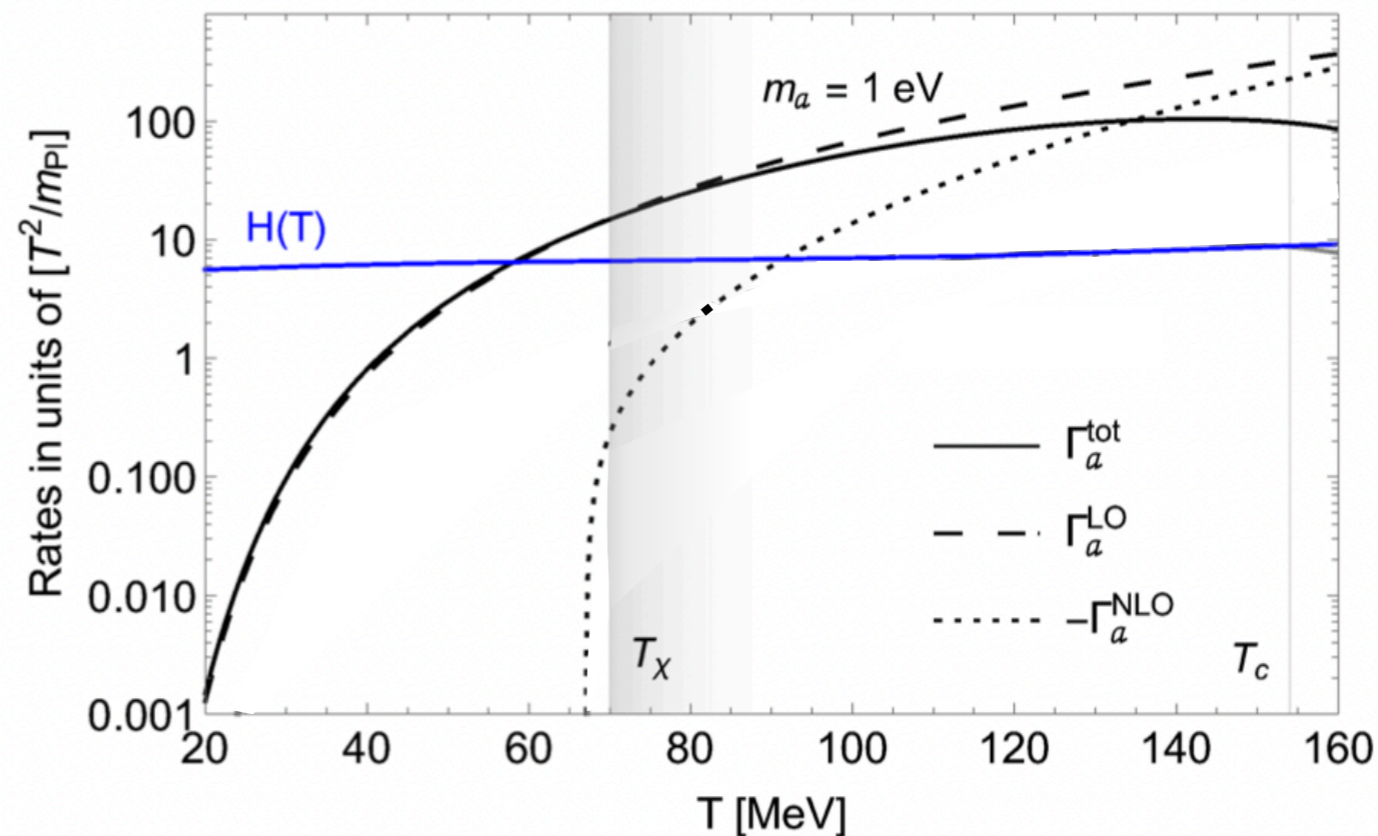
- NLO thermalization rate

[LDL, Piazza, Martinelli - 2101.10330]

[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli - 2211.05073]

$$\sum |\mathcal{M}|^2 = |\mathcal{M}_{\text{LO}}|^2 + 2\text{Re}[\mathcal{M}_{\text{LO}}\mathcal{M}_{\text{NLO}}^*]$$

$$\Gamma_a(T) = \left(\frac{C_{a\pi}}{f_a f_\pi}\right)^2 0.163 T^5 \left[h_{\text{LO}}(m_\pi/T) - 0.251 \frac{T^2}{f_\pi^2} h_{\text{NLO}}(m_\pi/T) \right]$$

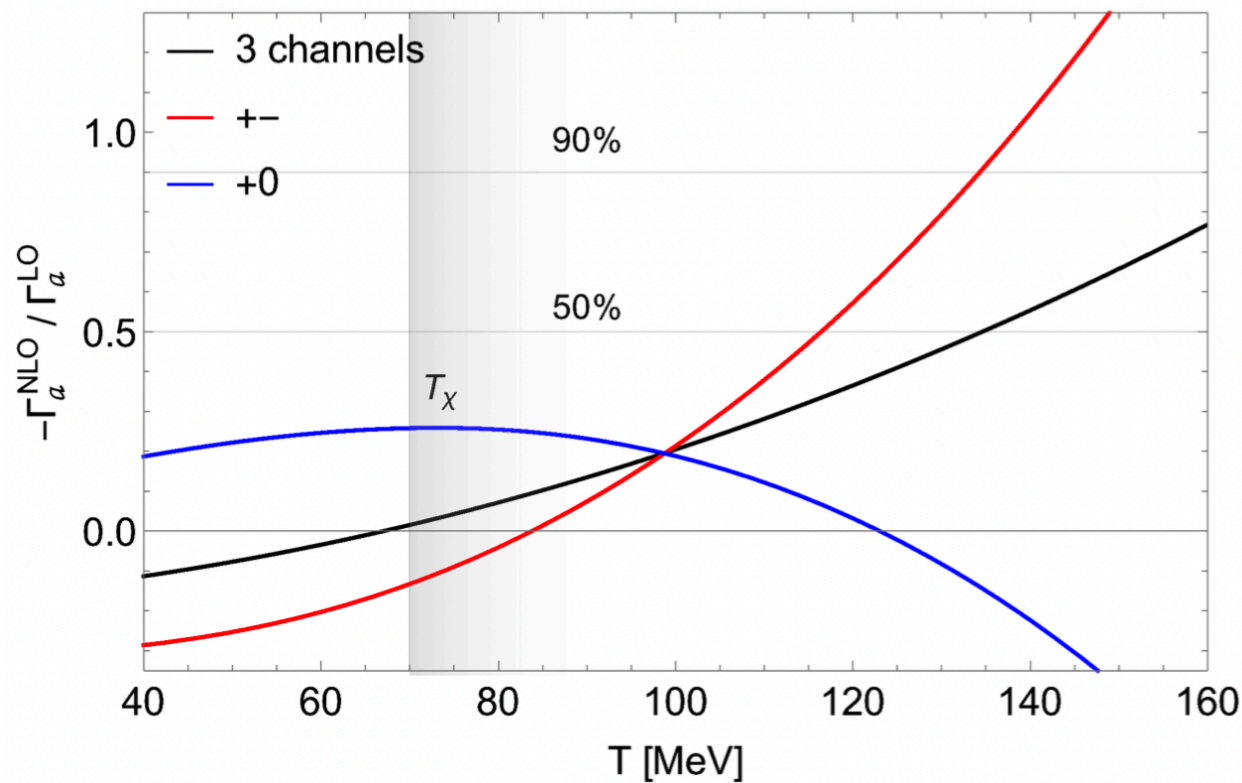


Axion-pion scattering at NLO

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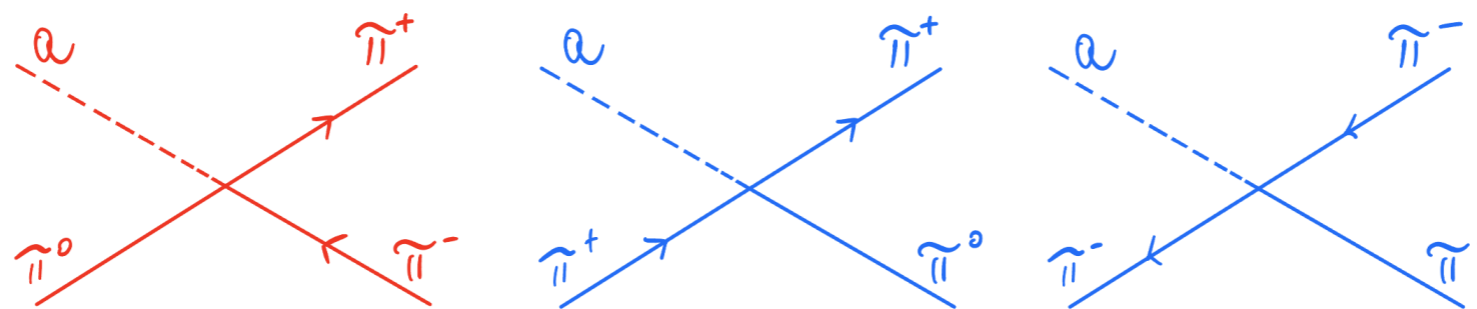
[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli - 2211.05073]



NLO correction to total Γ reaches 50% \times LO at $T \simeq 135$ MeV (due to accidental cancellations)

More realistic estimate of ChiPT validity by looking at first exclusive channel with large NLO correction

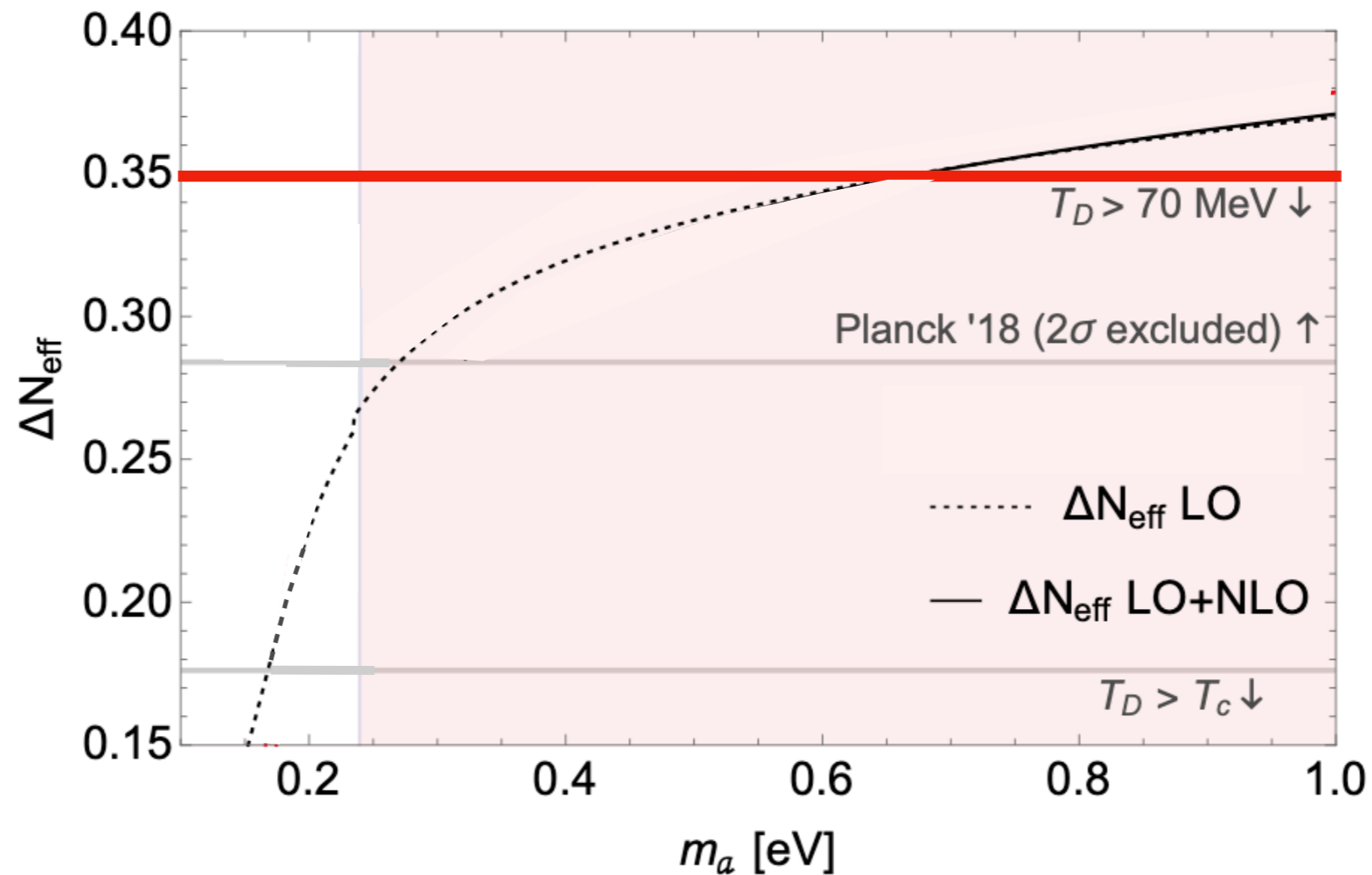
In $\pi^+\pi^0$ large correction at $T_\chi \simeq 70$ MeV



ΔN_{eff} with NLO corrections

[LDL, Piazza, Martinelli - 2101.10330]

[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli - 2211.05073]



present Planck bound is beyond the region of validity of ChIPT !

Beyond ChiPT

- Inverse Amplitude Method (IAM)

[Truong - PRL61 (1988), ...]

i) project scattering amplitude on partial waves, with angular momentum \mathbf{J} and iso-spin \mathbf{I}

$$\mathcal{M} \rightarrow A_{IJ} \qquad A_{IJ} = A_{IJ}^{(2)} + A_{IJ}^{(4)} + \dots \quad (\text{ChiPT expansion})$$

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ii) construct IAM amplitude, based on a *dispersion relation*

[See Salas-Bernárdez et al - 2010.13709 for theoretical uncertainties of IAM]

$$A_{IJ}^{\text{IAM}} = \frac{A_{IJ}^{(2)}}{1 - A_{IJ}^{(4)}/A_{IJ}^{(2)}} \quad \leftarrow \quad \frac{1}{A_{IJ}} = \frac{1}{A_{IJ}^{(2)} + A_{IJ}^{(4)} + \dots}$$



LECs from fit to $\pi\pi$ scattering data ($\mathcal{O}(10\%)$ error)

[Dobado, Pelaez - hep-ph/9604416]

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$$A_{IJ}^{\text{IAM}} = \frac{A_{IJ}^{(2)}}{1 - A_{IJ}^{(4)}/A_{IJ}^{(2)}}$$



- exact elastic unitarity
- reproduces σ and ρ resonances
- matches to ChiPT at low energy



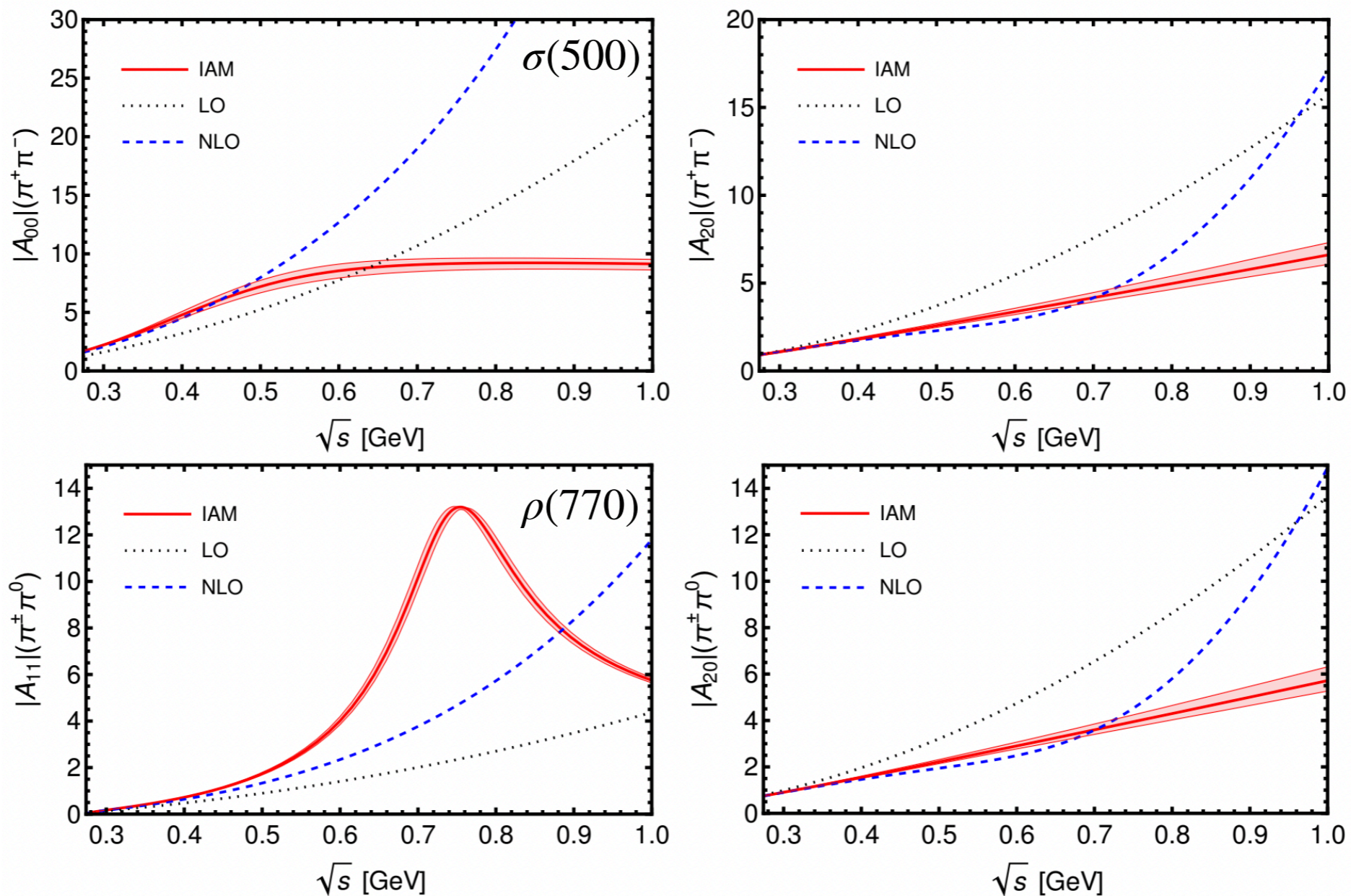
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Partial wave amplitudes

- Growth in energy of ChiPT amplitudes tamed by unitarization

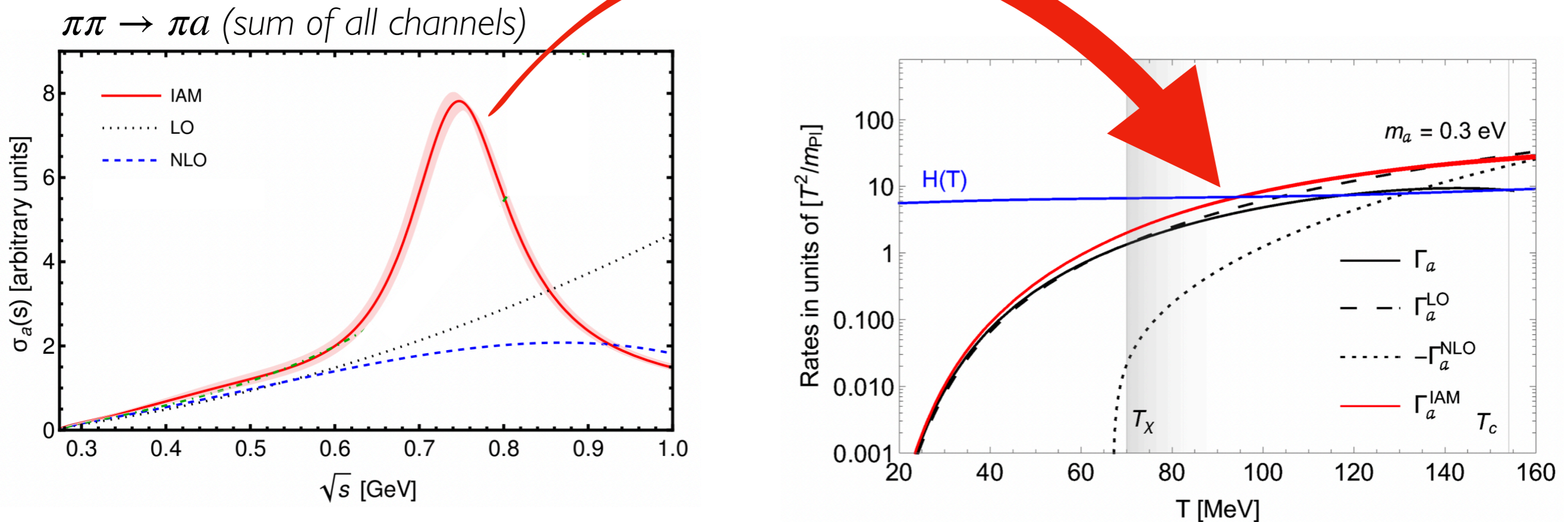
[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli - 2211.05073]



IAM thermalization rate

- Feature of ρ resonance (770 MeV) manifests around $T \sim 100$ MeV

[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli - 2211.05073]

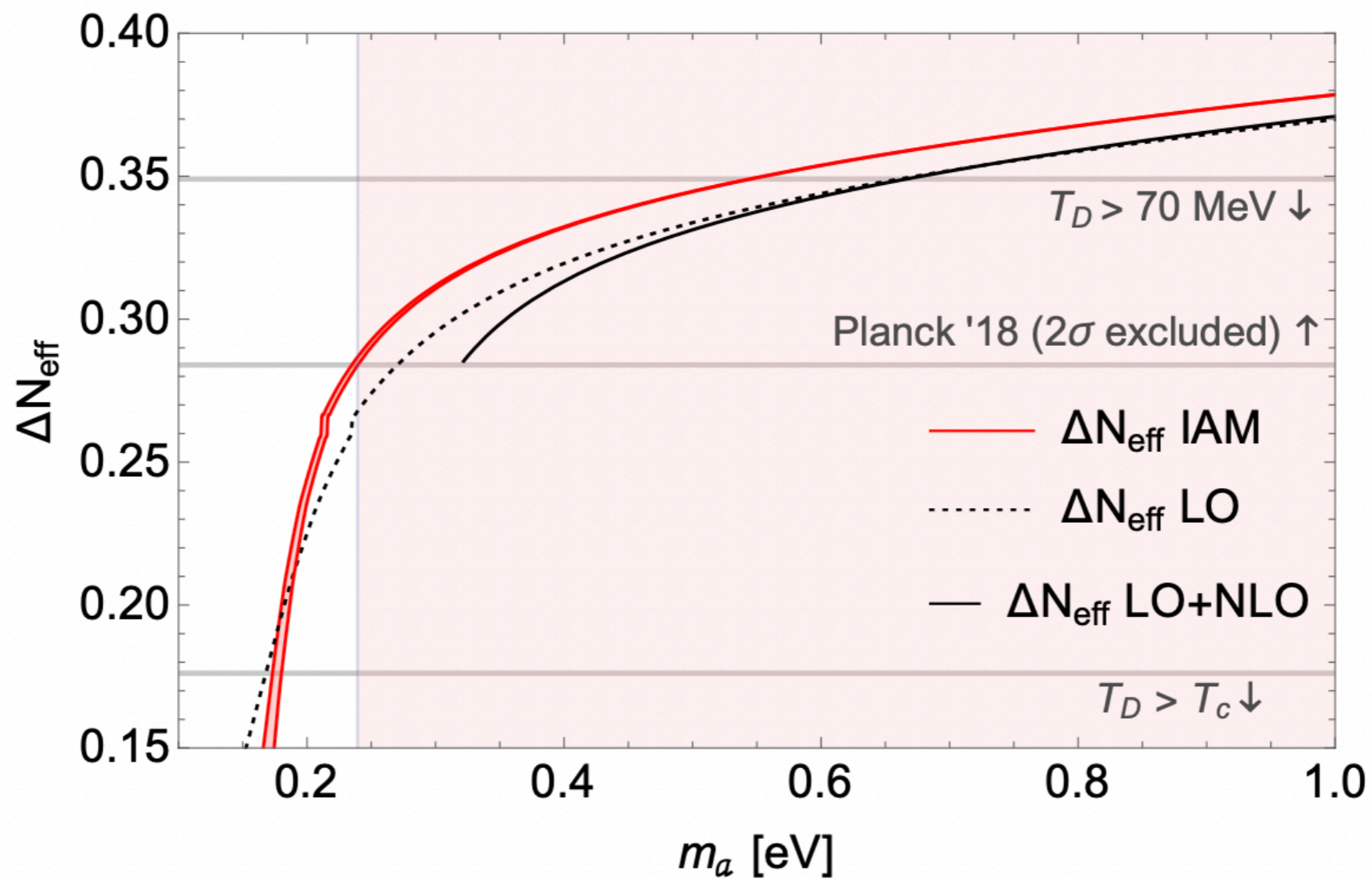


[Similar results in Notari, Rompineve, Villadoro - 2211.03799 based on a different unitarization approach]

ΔN_{eff} : IAM vs LO

- IAM bound: $m_a \lesssim 0.24$ eV

[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli 22 | I.05073]



Towards a robust bound

- Axion thermalization rate at $T \lesssim T_C \simeq 155 \text{ MeV}$

i) thermal corrections in axion-pion scattering

[Recently addressed in Wang, Guo, Zhou - 2312.15240. They amount to an $\mathcal{O}(10\%)$ shift on the bound on the axion mass]

ii) 3-flavour analysis to extend IAM above $\sqrt{s} \sim 800 \text{ MeV}$ (see backup slides)

iii) check contribution extra channels (kaons, etc)

[$K\pi \rightarrow Ka$ implemented at LO in ChiPT in Notari, Rompineve, Villadoro - 2211.03799. Comparable to $\pi\pi \rightarrow \pi a$ at $T \sim T_C$]

iv) axion couplings model-dependency

Towards a robust bound

- Axion thermalization rate at $T \lesssim T_C \simeq 155 \text{ MeV}$
- Boltzmann equation

$$\frac{dY}{d \log x} = (Y^{\text{eq}} - Y) \frac{\Gamma}{H} \left(1 - \frac{1}{3} \frac{d \log g_{*S}}{d \log x} \right)$$

[Salvio, Strumia, Xue -1310.6982, ...]



Momentum dependence is important:

1. x-section depends on momenta, which decouple at different times
2. # of d.o.f. decreases rapidly around T_C , higher momenta less diluted
3. production might be never in thermal equilibrium

$$\frac{\partial \mathcal{F}_a}{\partial t} - H |\mathbf{k}| \frac{\partial \mathcal{F}_a}{\partial |\mathbf{k}|} = \Gamma_a (\mathcal{F}_a^{\text{eq}} - \mathcal{F}_a)$$

[Notari, Rompineve, Villadoro - 2211.03799
Bianchini, Grilli di Cortona, Valli - 2310.08169]

$$\Gamma_a \equiv \text{Im}(\Pi_a^{\text{R}}) / (E f_a^2)$$

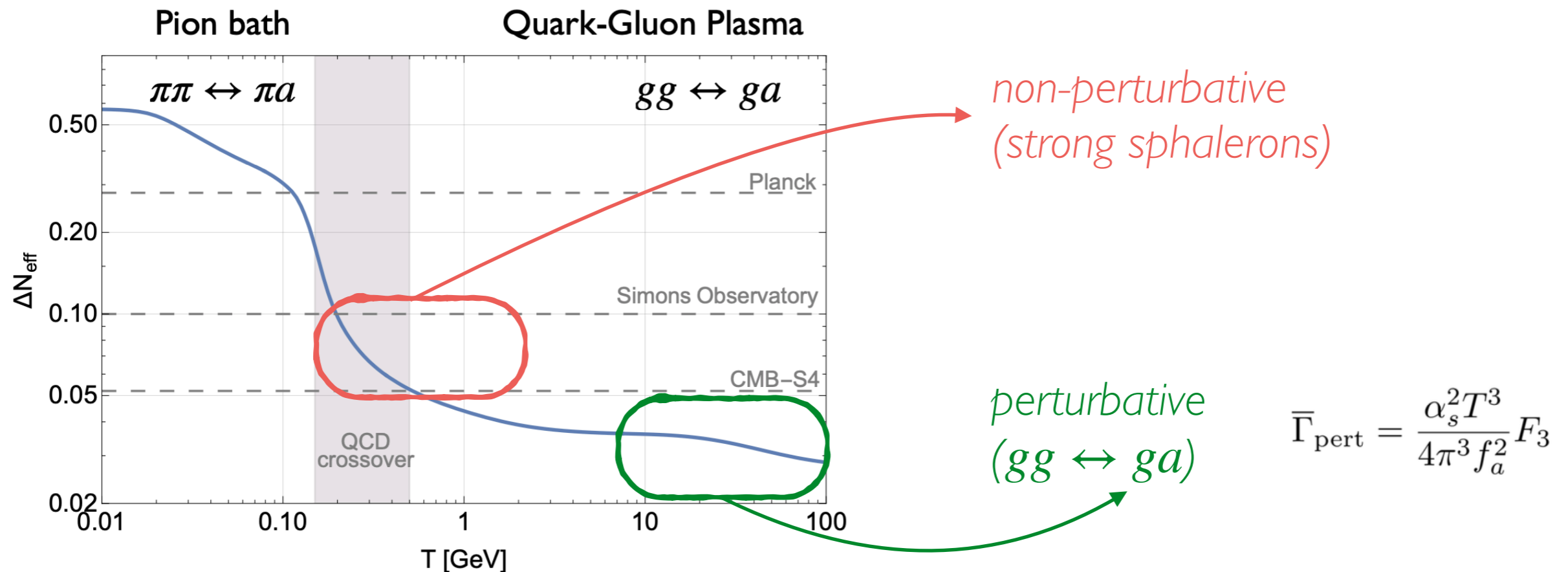
$$\Pi_a^{\text{R}} = i \int d^4 x e^{i x k} \langle [Q(x), Q(0)] \Theta(t) \rangle$$

$$Q \equiv \alpha_s / (8\pi) G \tilde{G}$$

Towards a robust bound

- Axion thermalization rate at $T \lesssim T_C \simeq 155$ MeV
- Boltzmann equation
- Axion thermalization rate at $T \gtrsim T_C \simeq 155$ MeV

➔ needed to assess sensitivity of future CMB surveys

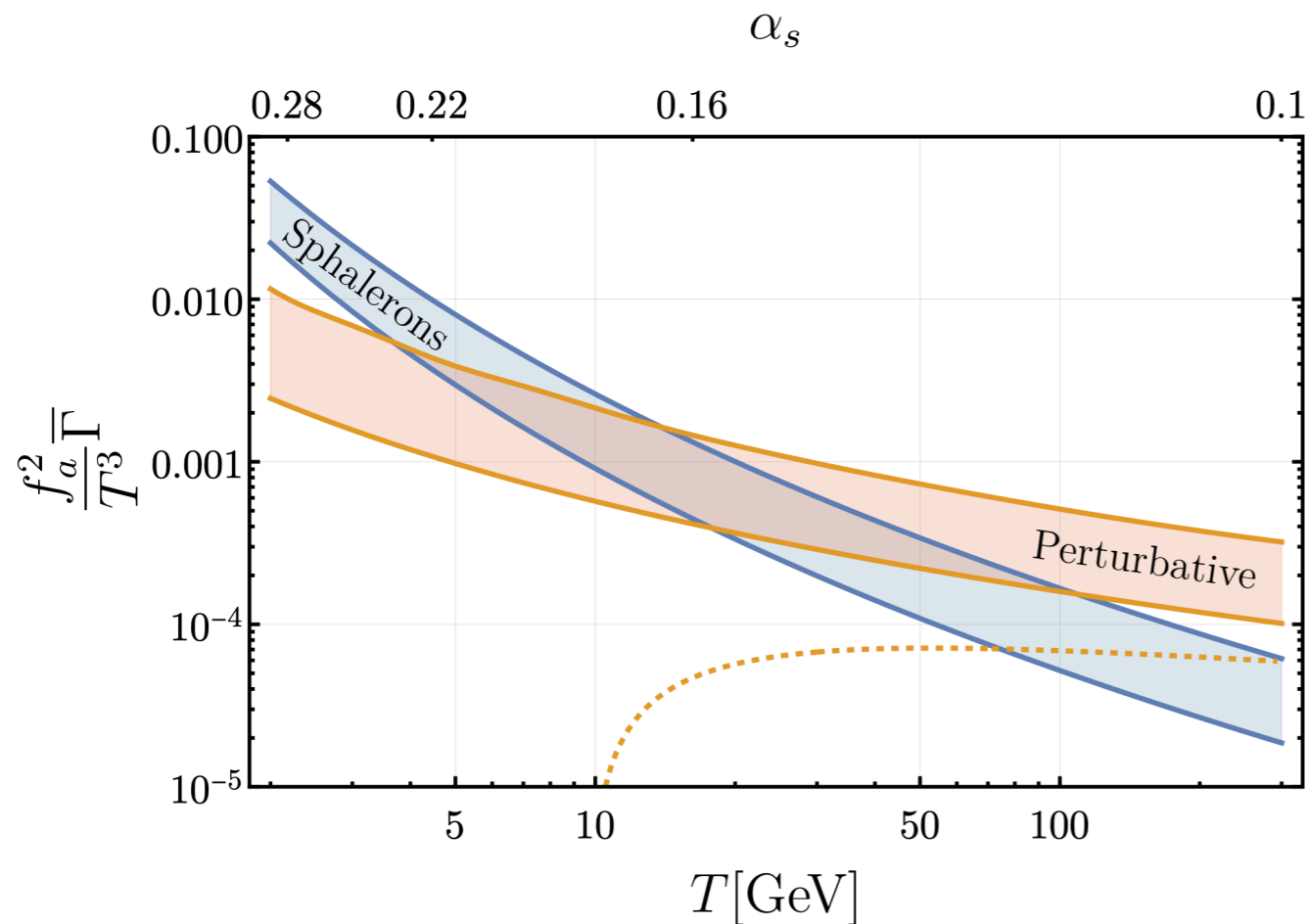


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[D'Eramo, Hajkarim, Yun - 2211.03799 → interpolation strategy between perturbative and ChiPT regions]

[Notari, Rompineve, Villadoro - 2211.03799 → strong sphalerons dominate until $T \sim 10$ GeV]



*non-perturbative
(strong sphalerons)*

*perturbative
($gg \leftrightarrow ga$)*

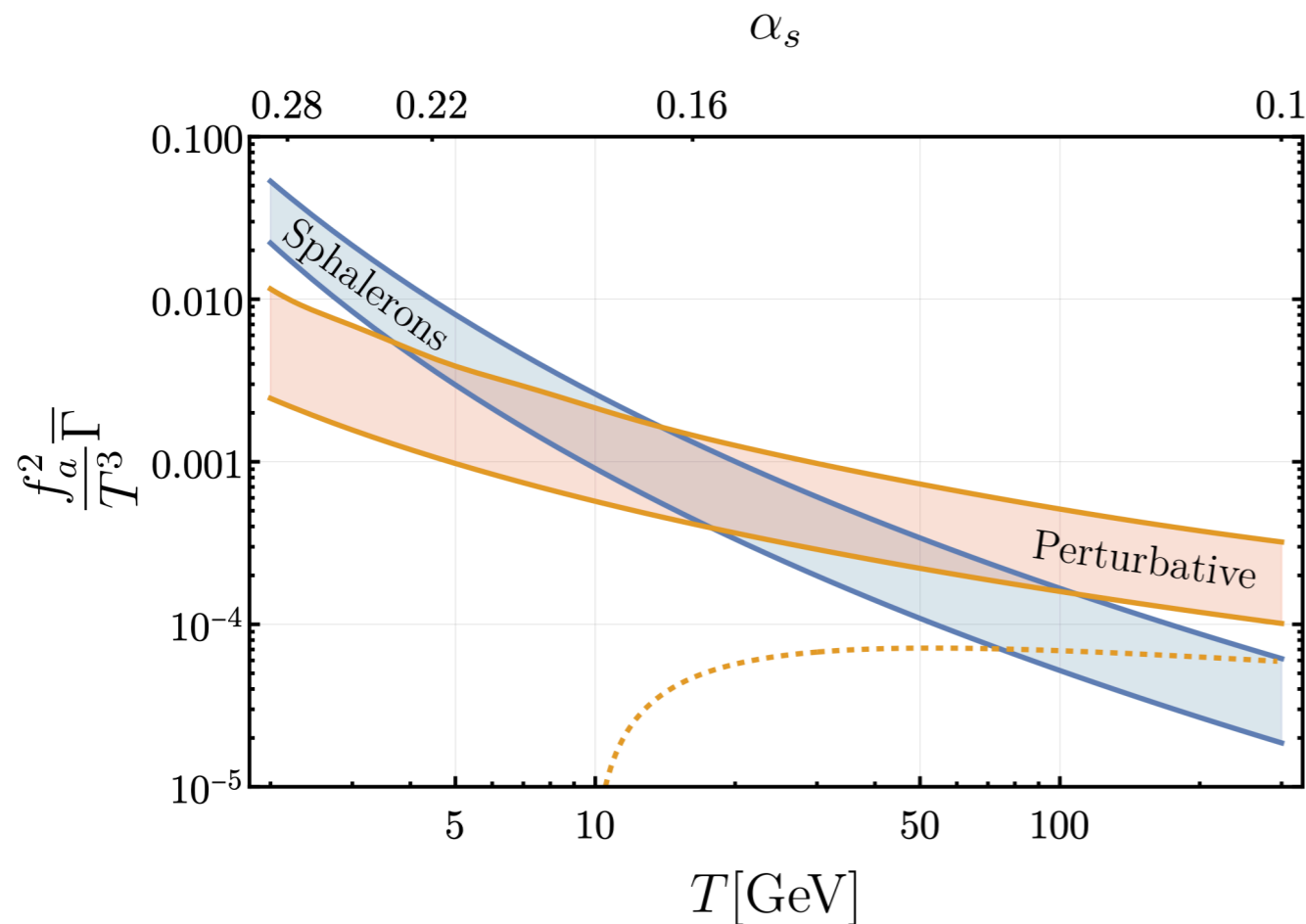
$$\bar{\Gamma}_{\text{pert}} = \frac{\alpha_s^2 T^3}{4\pi^3 f_a^2} F_3$$

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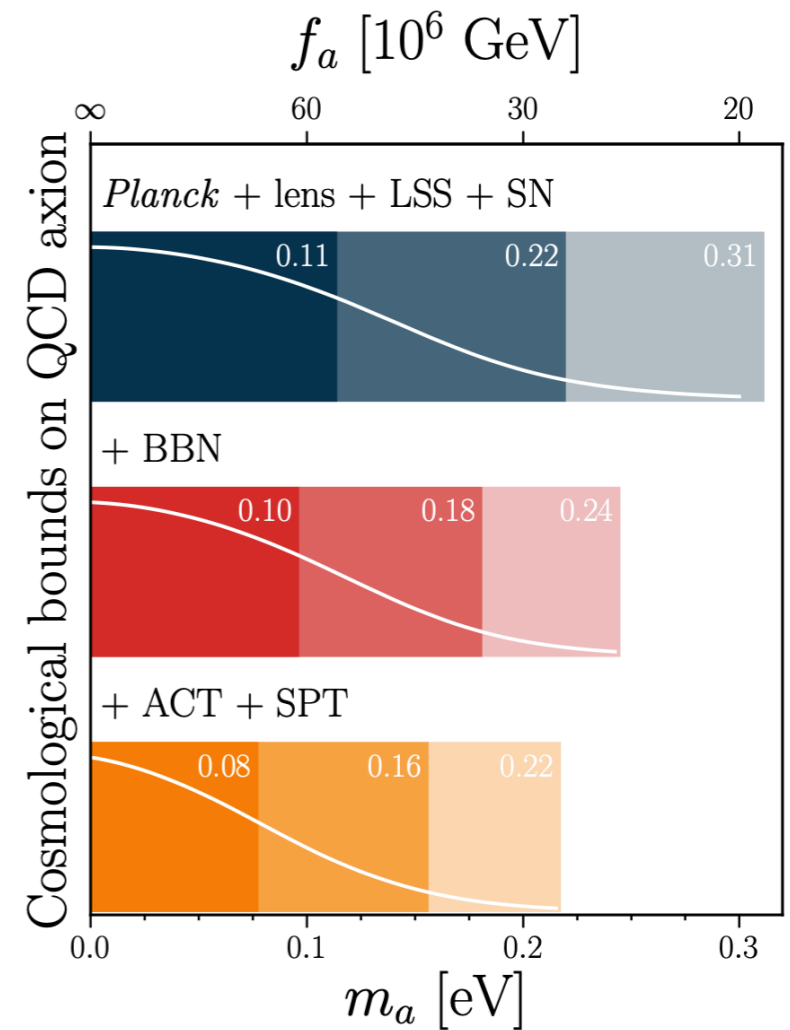
[Bonanno, D'Angelo, D'Elia, Naviglio, Maio
- 2308.01287 → first $N_f = 2+1$ lattice QCD
result at $k = 0$ for $200 \text{ MeV} \lesssim T \lesssim 600 \text{ MeV}$]

$$\Gamma_{\text{Sphal}} = \lim_{\substack{V_s \rightarrow \infty \\ t_M \rightarrow \infty}} \frac{1}{V_s t_M} \left\langle \left[\int_0^{t_M} dt'_M \int_{V_s} d^3x q(t'_M, \vec{x}) \right]^2 \right\rangle$$

Towards a robust bound

- Axion thermalization rate at $T \lesssim T_C \simeq 155$ MeV
- Boltzmann equation
- Axion thermalization rate at $T \gtrsim T_C \simeq 155$ MeV
- Cosmological observables

[...
Caloni, Gerbino, Lattanzi, Visinelli - 2205.01637
D'Eramo, Di Valentino, Giarè, Hajkarim, Melchiorri, Mena, Renzi, Yue - 2205.07849
Di Valentino, Gariazzo, Giarè, Melchiorri, Mena, Renzi - 2212.11926
Bianchini, Grilli di Cortona, Valli - 2310.08169]



[Bianchini, Grilli di Cortona, Valli - 2310.08169]

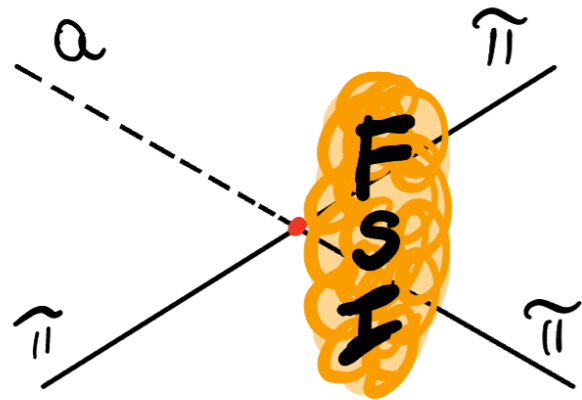
Conclusions

- Resurgence of interest in **axion thermal production**
- Robust bound in the ballpark of $m_a \lesssim 0.2$ eV
- Future CMB exp.'s will provide an axion discovery channel beyond astrophysical limits

Backup slides

Elementary Watson !

- $\pi\pi$ final state interactions (FSI) are resonant (ρ in $I = J = 1$ and σ in $I = J = 0$)



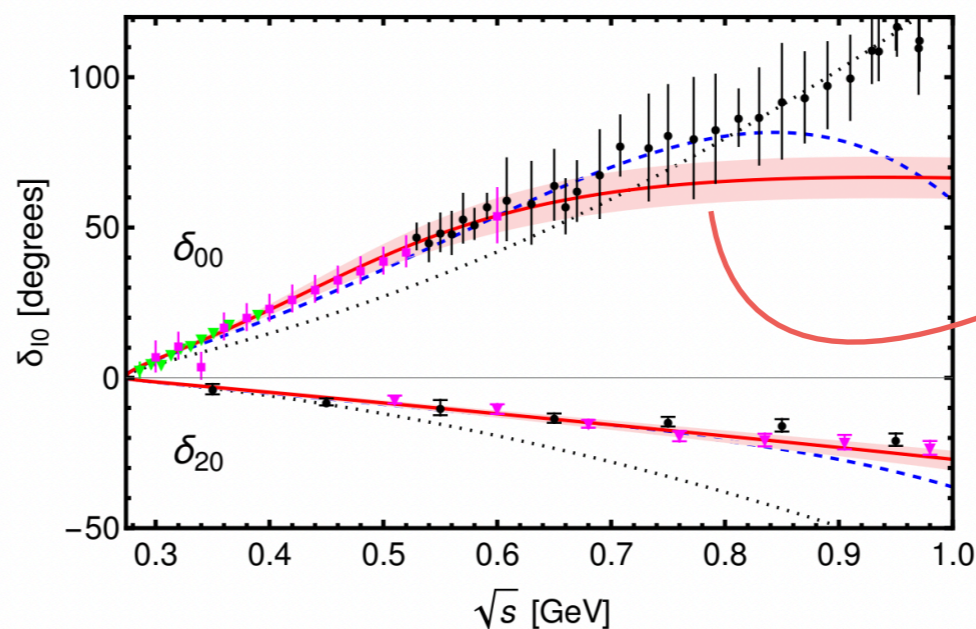
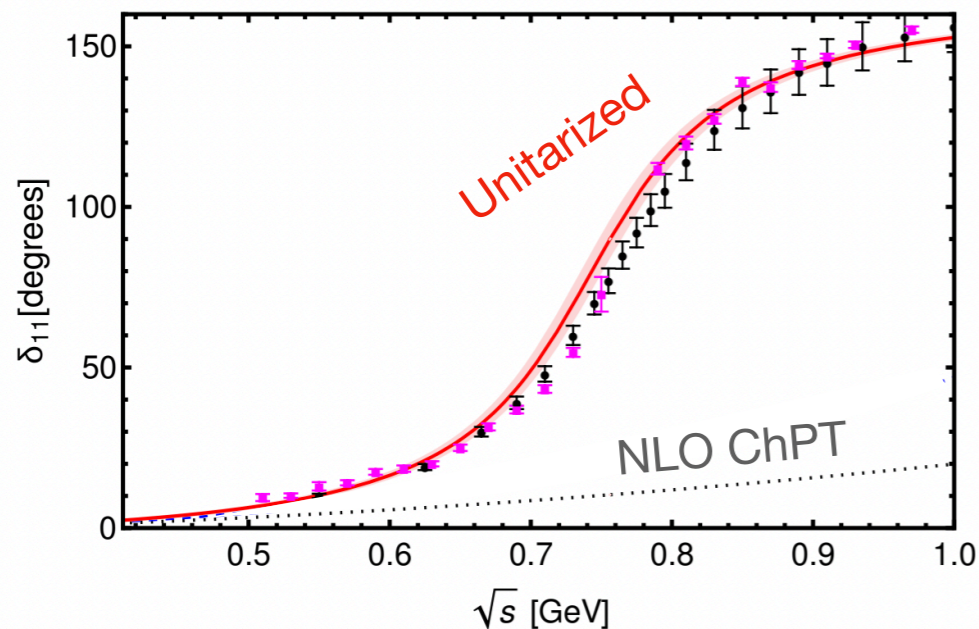
$$\text{Im } T_{IJ} = \frac{\sigma(s)}{32\pi} |T_{IJ}|^2 \quad \longleftrightarrow \quad T_{IJ} = \frac{32\pi}{\sigma(s)} e^{i\delta_{IJ}} \sin \delta_{IJ}$$

[Watson - Phys. Rev. 88 (1952)]

Unitarity \longrightarrow Watson's theorem: $(\delta_{a\pi})_{IJ} = (\delta_{\pi\pi})_{IJ}$

- Phase shifts from IAM reproduce $\pi\pi$ data up to $\sqrt{s} \simeq 800$ MeV

[LDL, Camalich, Martinelli, Oller, Piazza, Martinelli - 2211.05073]



K-Kbar threshold - $f_0(980)$ requires 3-flavour analysis