

# All-sky and Galactic Ridge diffuse astrophysical neutrino flux search with KM3NeT/ARCA6-8-19-21 data



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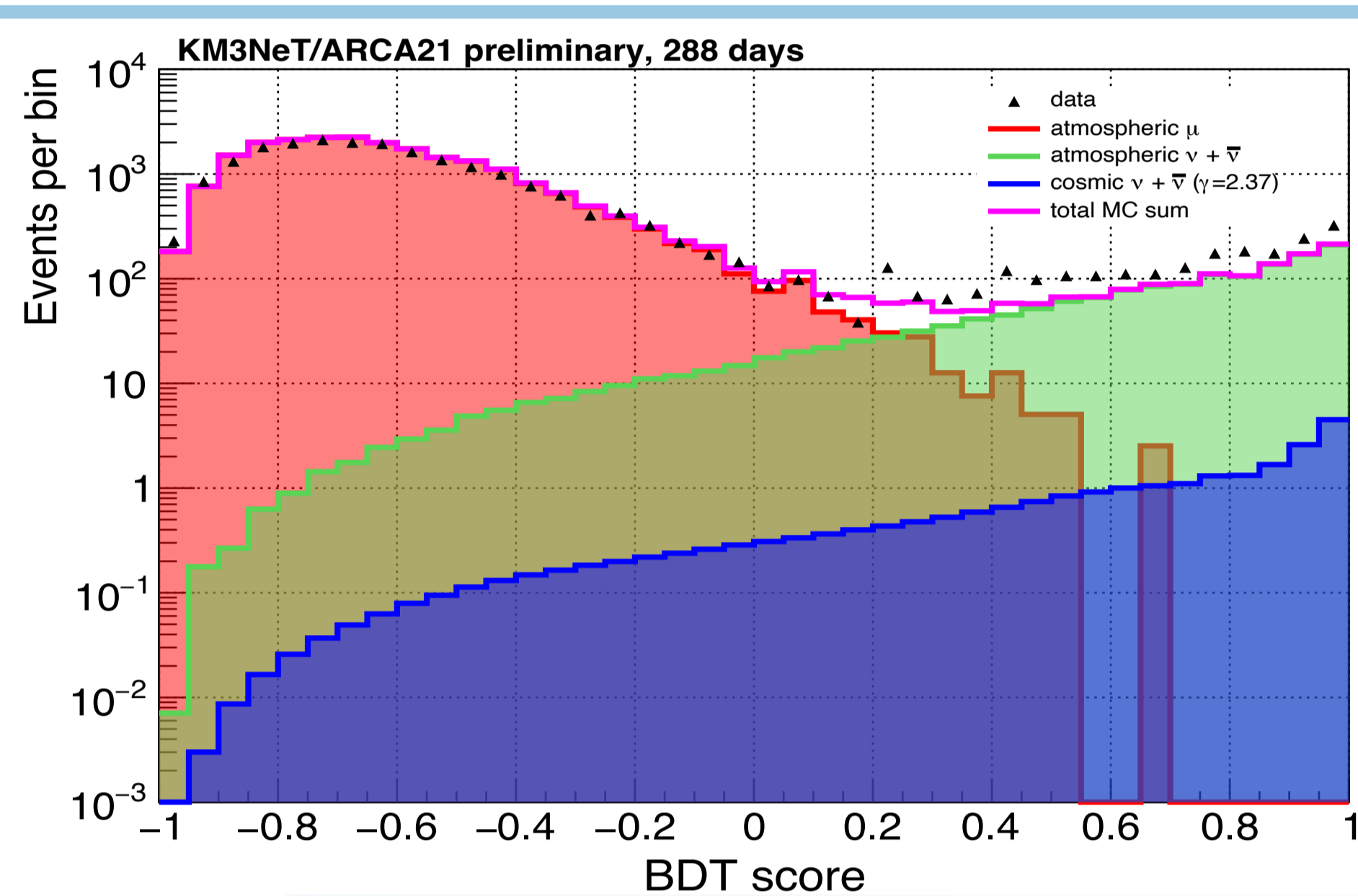
## KM3NeT detectors

- Network of Cherenkov detectors [1]:
  - ➔ KM3NeT/ARCA optimized for high-energy (TeV-PeV) astrophysical neutrinos;
  - ➔ KM3NeT/ORCA optimized for few-GeV atmospheric neutrinos;
- 31 x 3" PMTs hosted in pressure resistant glass sphere: Digital Optical Module (DOM);
- 18 DOMs anchored at the seafloor with buoyancy at the top form a detection unit (DU);
- 115 DUs evenly spaced form a building block.

	ARCA	ORCA
Location	Italy	France
N. building building blocks	2	1
DU distance	90 m	23 m
DOM spacing	36 m	9 m
DU height	~ 800 m	~ 200 m
Instrumented mass (Mton)	2*650	7
Depth	3500 m	2500 m



The KM3NeT DOM



## Event selection & BDT

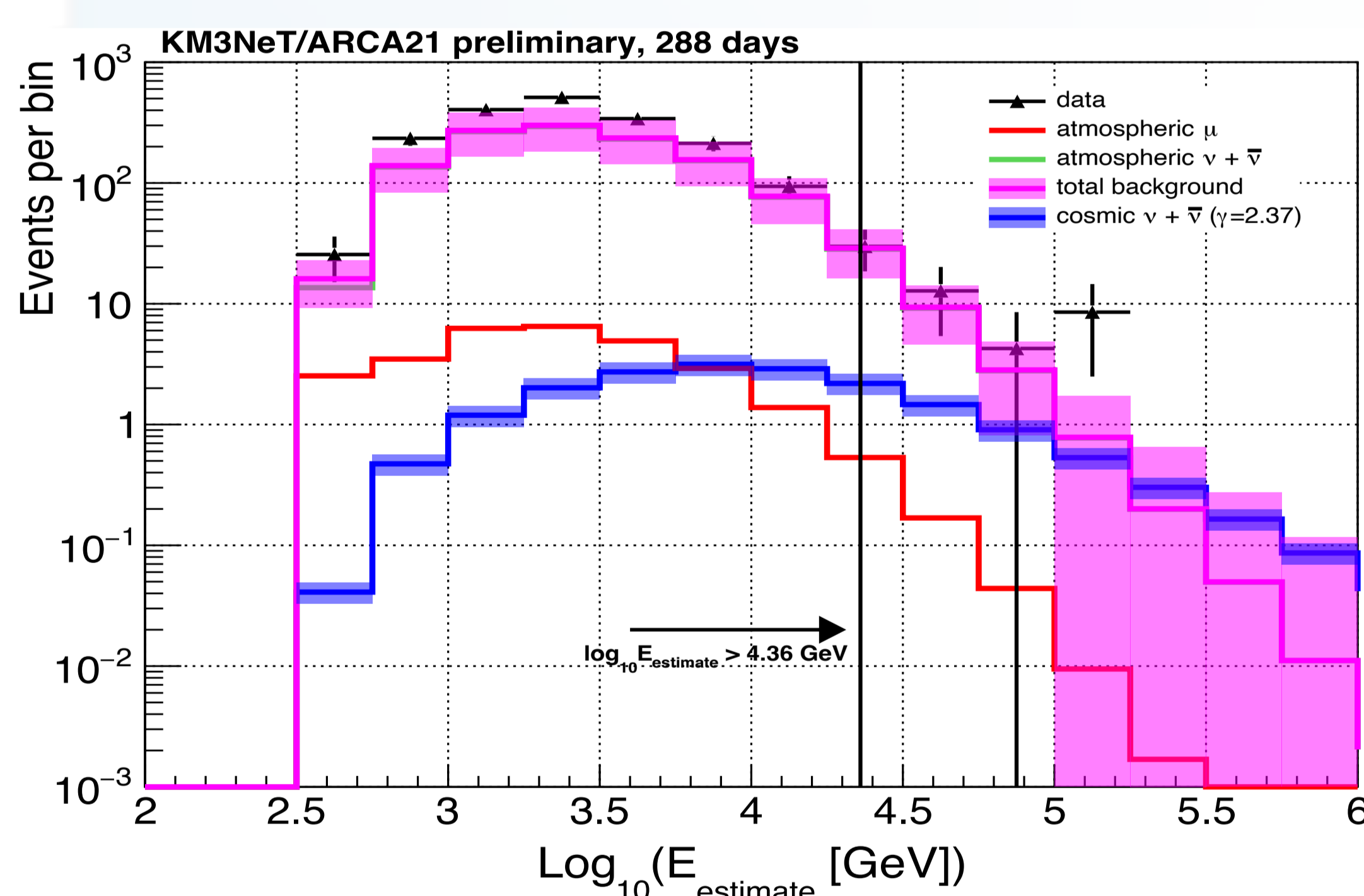
Events reconstructed as up-going tracks are selected based on trigger level information and on variables indicating the quality of the track reconstruction.

A Boosted Decision Tree (BDT) classifier, using ROOT TMVA<sup>[2]</sup> was applied to suppress the atmospheric muon background. Separate BDT trainings have been performed for the ARCA6/8 and ARCA19/21 geometries.

Cut optimization based on MRF minimization<sup>[3]</sup>.

## All-sky

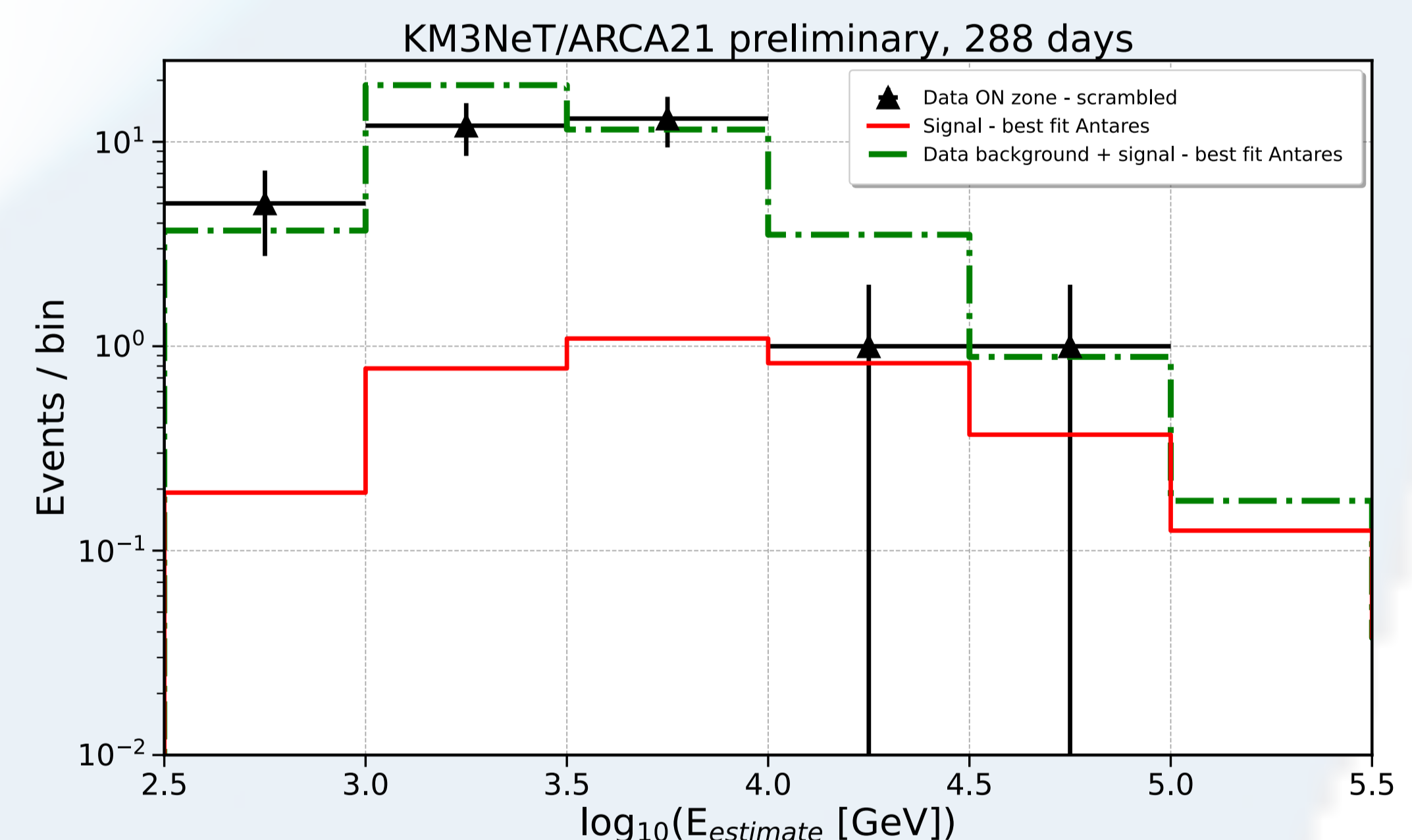
An all-flavor cosmic neutrino search (both CC+NC) is performed. In this work the flux is parameterized with an unbroken power-law. As a baseline IceCube's reported values<sup>[4]</sup>  $\phi_0 = 1.44$  and  $\gamma = 2.37$  are used.



## Galactic ridge

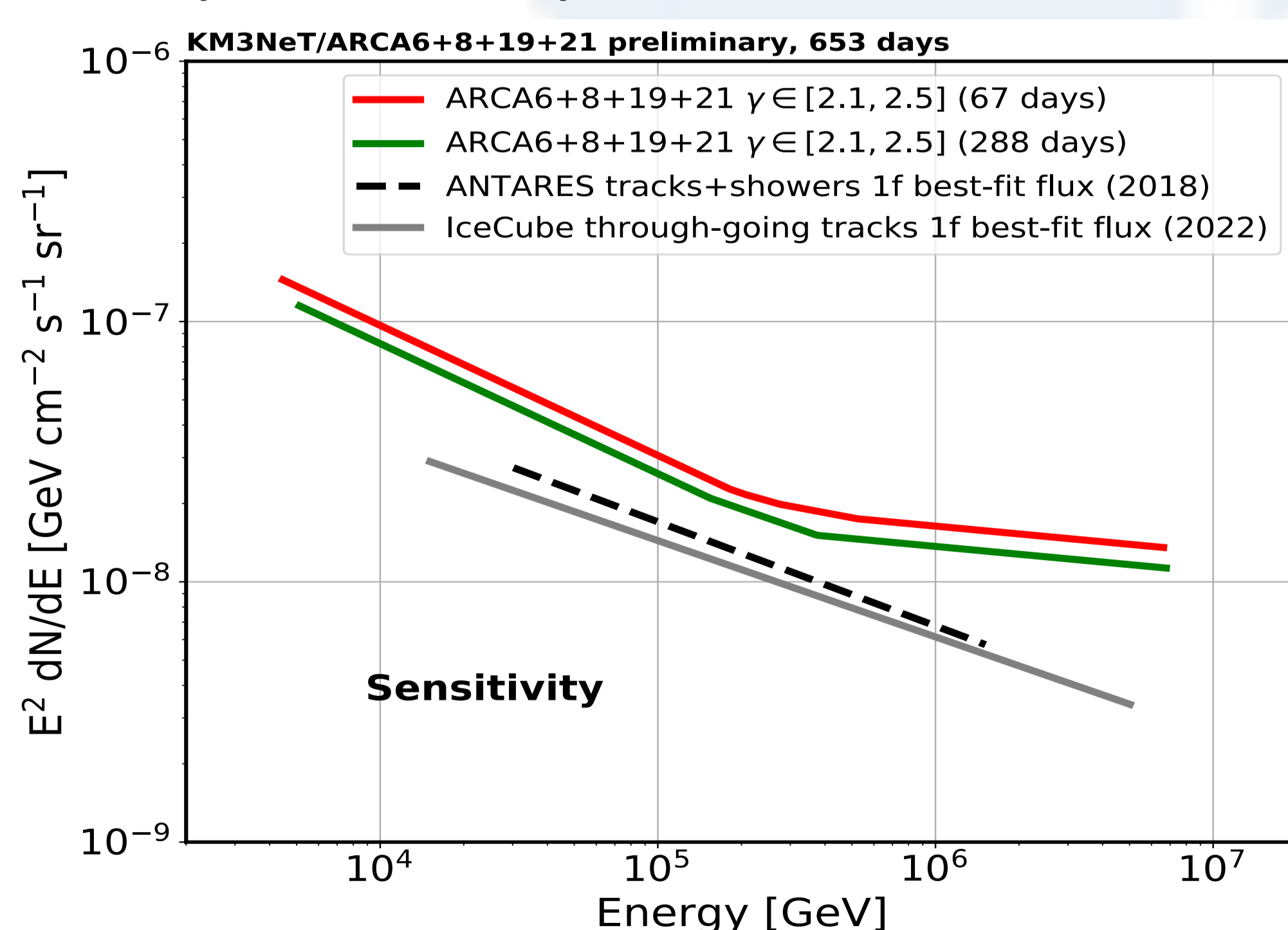
Searching for a possible excess of events coming from an extended region, with Galactic coordinates  $|l| < 30^\circ$  and  $|b| < 2^\circ$ , namely the Galactic Ridge (ON zone).

Data-driven background extracted from OFF zone region: same sky coverage but shifted in right ascension. Signal modeled assuming an unbroken power-law. More detail on the analysis procedure in<sup>[5]</sup>.

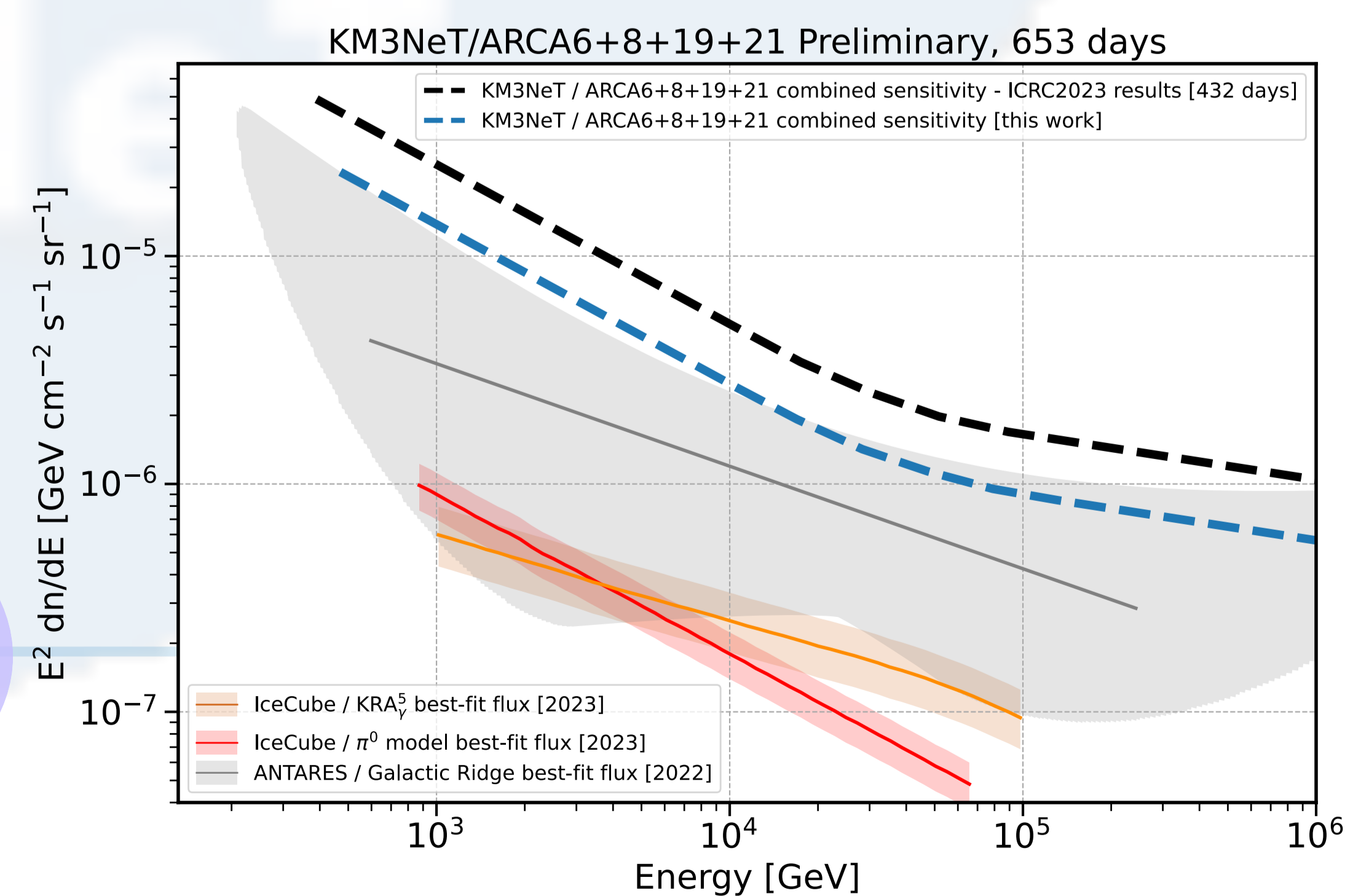


## Sensitivity

A Bayesian approach, described in<sup>[5]</sup> is used to extract sensitivities and upper limits for different spectral indices. The convolution of sensitivities at 90% C.I. for ARCA6+8+19+21, as a function of energy, for a subset of selected spectral indices is shown. The improvement in sensitivity reported in this contribution with respect to ICRC23, is due to an increase of the livetime analyzed: from 67 days to the whole ARCA21 data taking period, accounting in total 288 days.



KM3NeT/ARCA is rapidly evolving approaching ANTARES and IceCube fitted fluxes



In the plot above, the KM3NeT/ARCA sensitivity, for the two searches, is compared to the ANTARES and IceCube best fits.

KM3NeT detectors are collecting more and more data and the instrumented volume is growing. KM3NeT/ARCA is currently taking data with 28 DUs.

## References:

- [1] KM3NeT collaboration, Letter of intent for KM3NeT 2.0, *J. Phys. G* 43 (2016) 084001;
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- [4] R. Abbasi et al. 2022 *ApJ* 928 50, DOI:10.3847/1538-4357/ac4d29;
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Vasileios Tsourapis acknowledges the support of the Hellenic Foundation for Research and Innovation (HFRI) under the 3rd Call for HFRI PhD Fellowships (Fellowship Number: 5403).