


NEOS-II updates

June 21 @ , Milano

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on behalf of NEOS-II Collaboration



Neutrino Experiment for Oscillation at SBL

Key features

- A single, homogeneous IBD detector at a fixed distance (~24 m) from a commercial reactor.
- No self L resolution
- Oscillation analysis depending on existing model/longer baseline data.

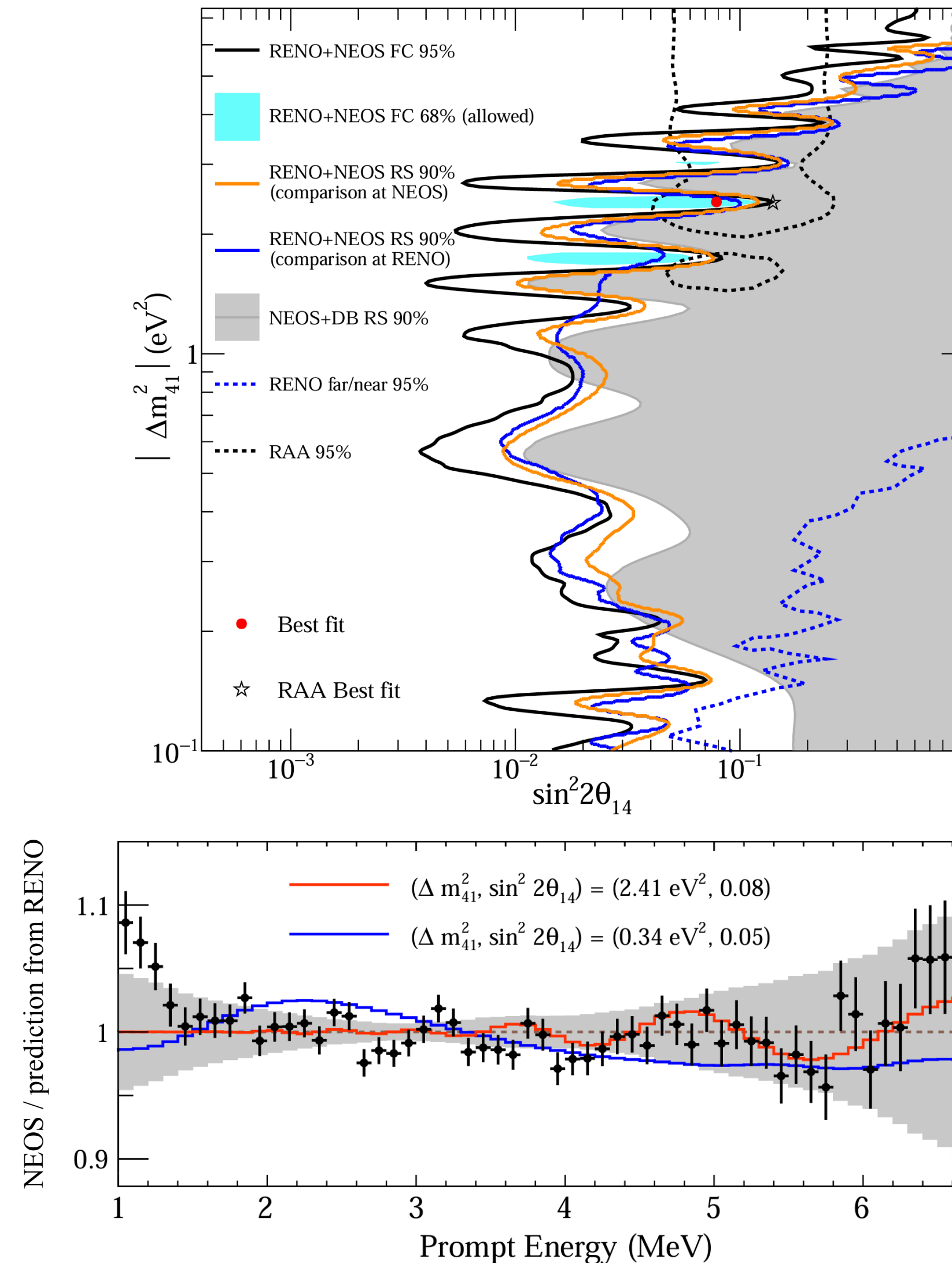
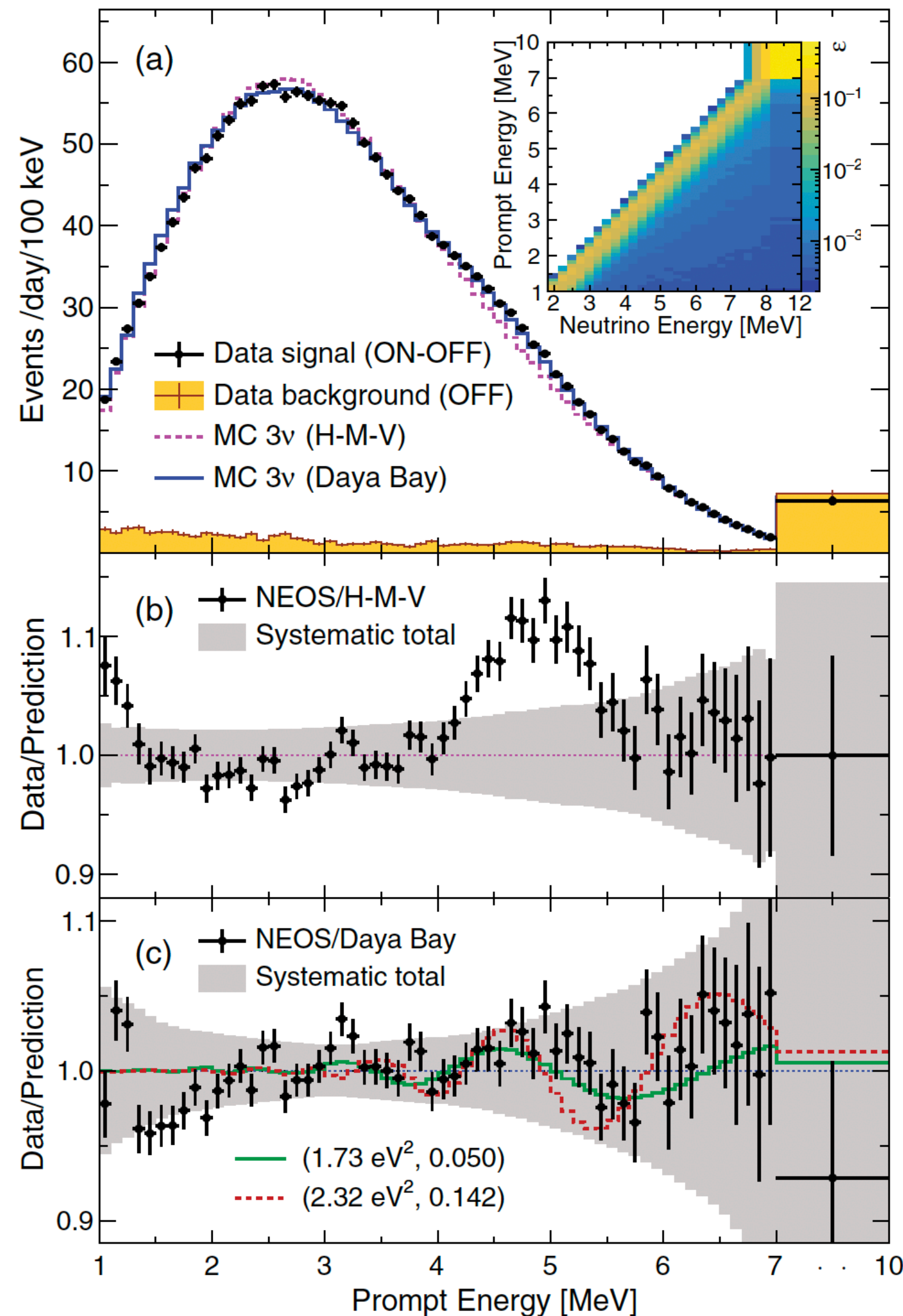
Outline

- Reminder: NEOS-I and experimental apparatus
- NEOS-II goals
- NEOS-II result updates:
 - Extraction of antineutrinos from ^{235}U and ^{239}Pu .
 - Short baseline oscillation

NEOS-I result on SBL oscillation

PRL 118 (2017) 121802
PRD 105 (2022) L111101

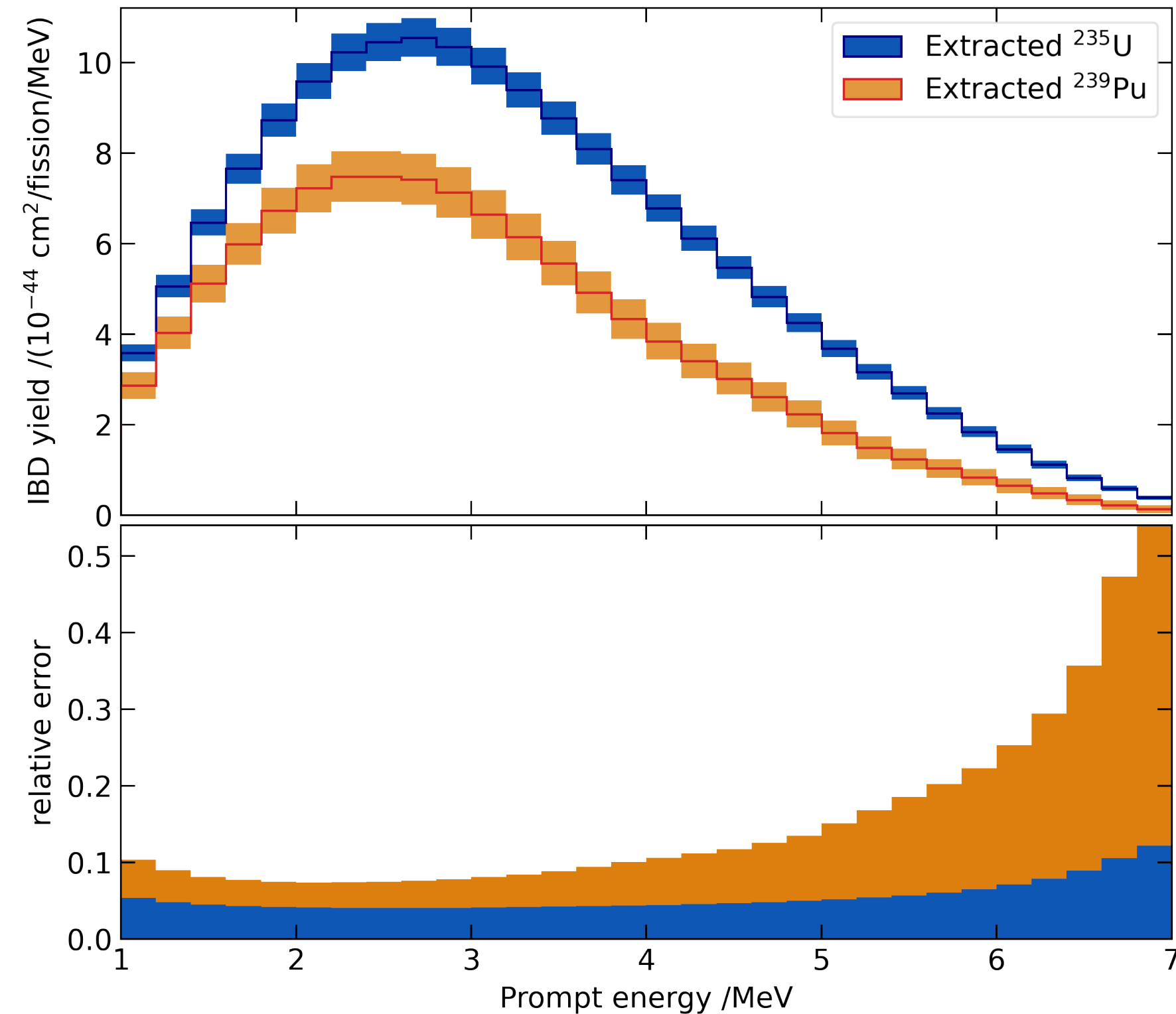
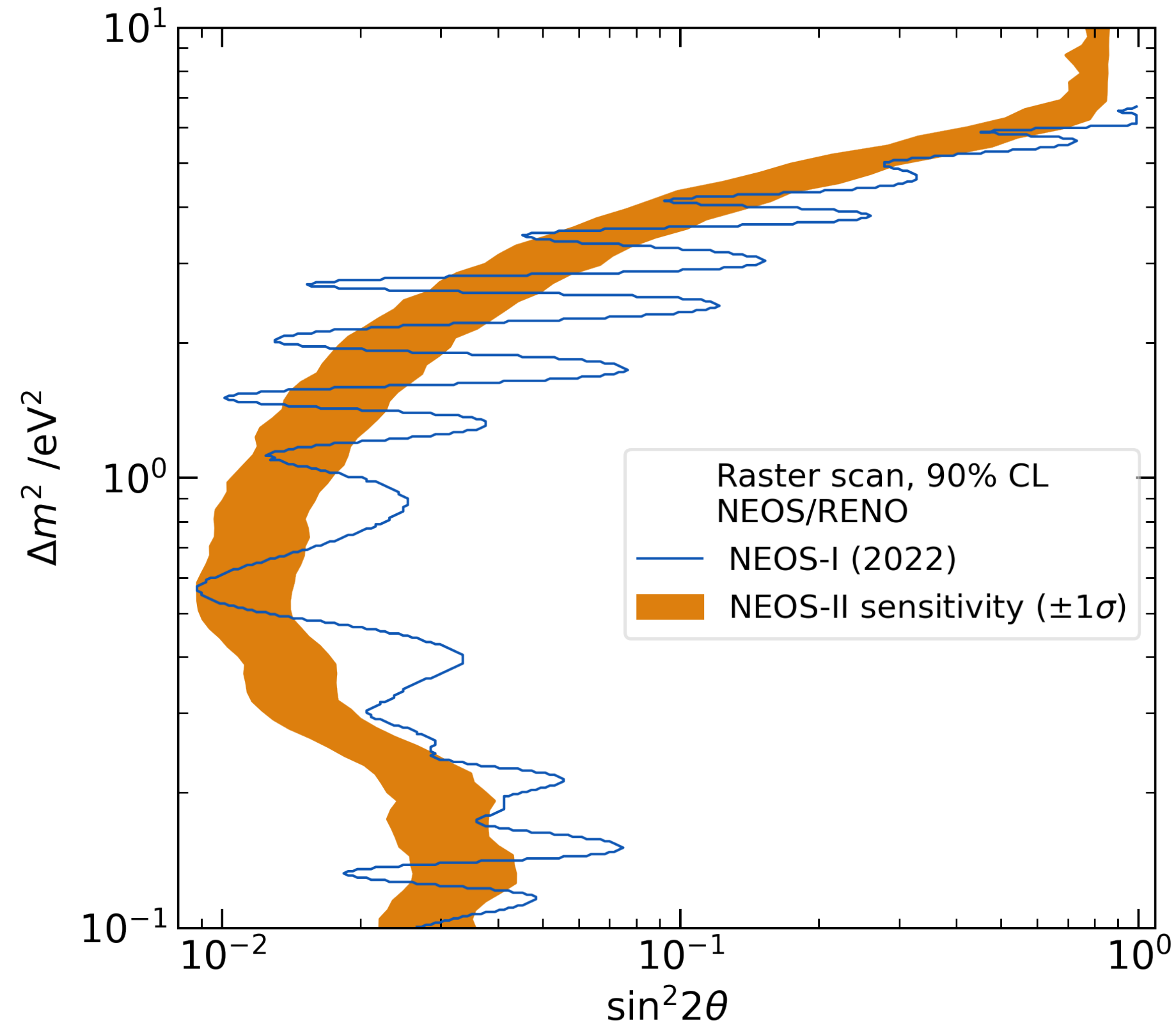
Comparison with Huber-Mueller, Daya Bay, and RENO



- Reactor ON: 180 days, OFF: 46 days.
- Daily IBD candidate rate: 1977 (ON) / 85 (OFF), S/B~22.
- vs Daya Bay reference:
 - Best fit at: $(0.05, 1.7 \text{ eV}^2)$
 $\Delta\chi^2 = 6.5, p\text{-value} = 22\%$.
- vs RENO reference:
 - Considering oscillation at RENO, too.
 - Best fit at: $(0.08, 2.4 \text{ eV}^2)$
 $\Delta\chi^2 = 8.4, p\text{-value} = 8.2\%$.
- Distortion or “bump” not totally disappeared after the Daya Bay/RENO normalization.
 - Differences in fission fractions?

From NEOS-I to NEOS-II

Utilizing a full fuel cycle of a sole commercial reactor



- Improving sensitivity for the SBL oscillation: adding more significance to the NEOS-I best fit(s), or to find out that it was an statistical fluctuation?
- Reactor- ν spectrum evolution in time: extraction of the primary isotopes' ($^{235}\text{U}/^{239}\text{Pu}$) contribution.

Hanbit Nuclear Power Plant

Yeonggwang (靈光: ghost light), Korea

黃海
Yellow sea

RENO
near
detector

1

2

3

4

5

6

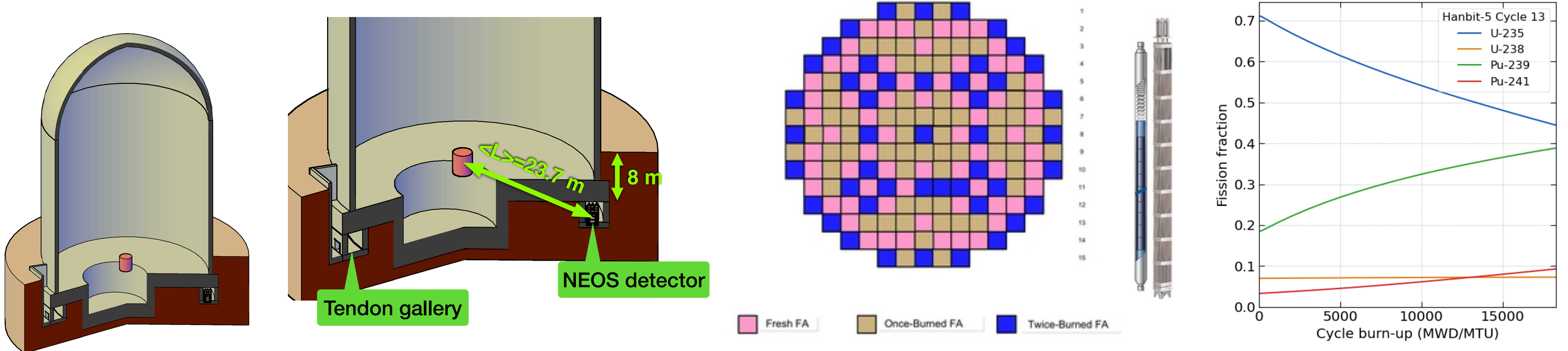
NEOS

NEON
(Poster#493)

Neutrino source: Hanbit-5 reactor

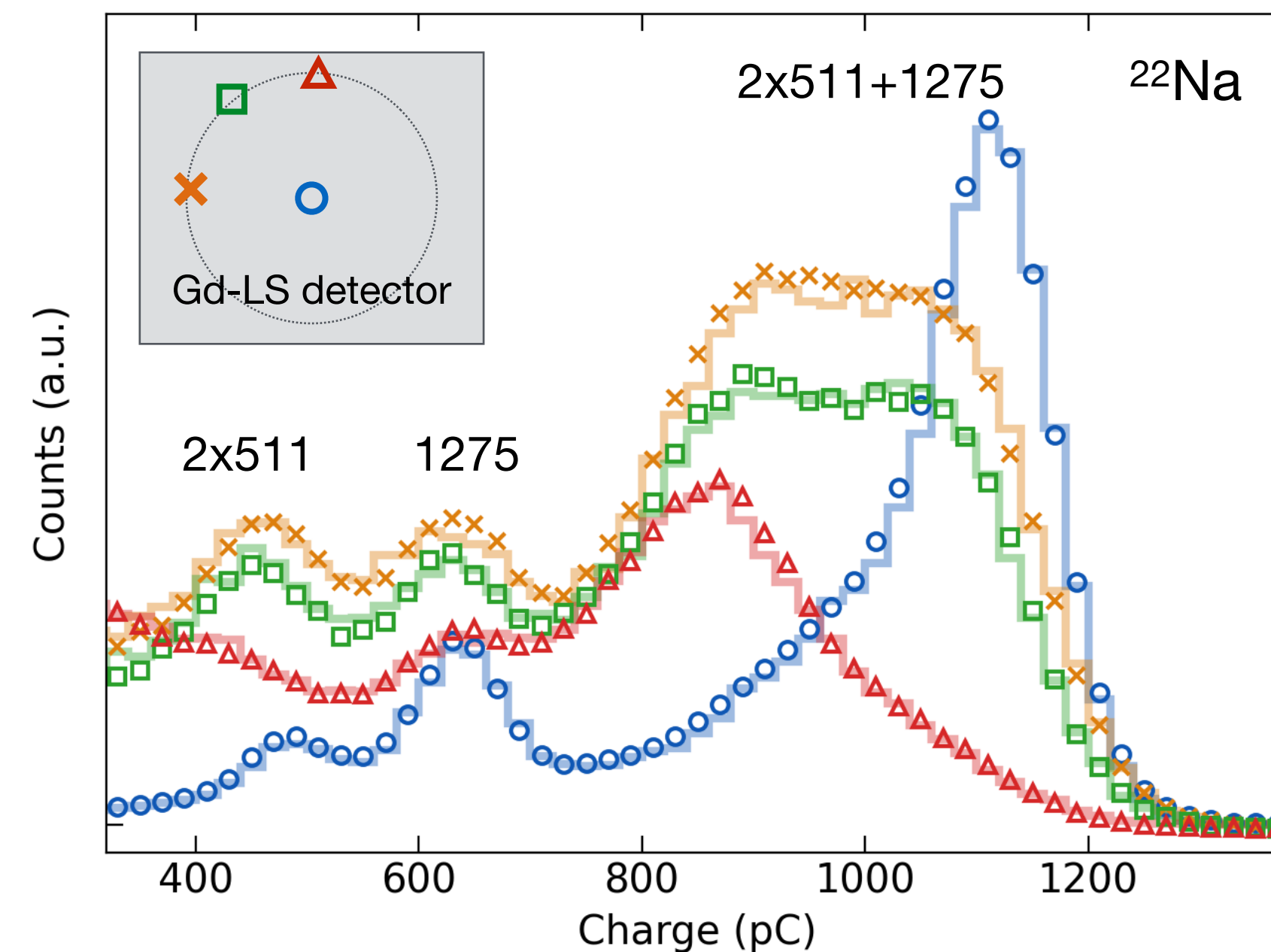
Low-enriched-uranium fuel / commercial / thermal power 2.8 GW

- A typical burn-up cycle takes about 1.5 years.
 - 1/3 of the fuel assemblies changed to new ones for a new cycle.
- Active core size— diameter: 3.1 m, height: 3.8 m.
- NEOS detector located 23.7 m away from the active core center.
- The nearest neighboring core, Hanbit-4&6, at 256 m away, less than 1% contribution of the main source.



Detector

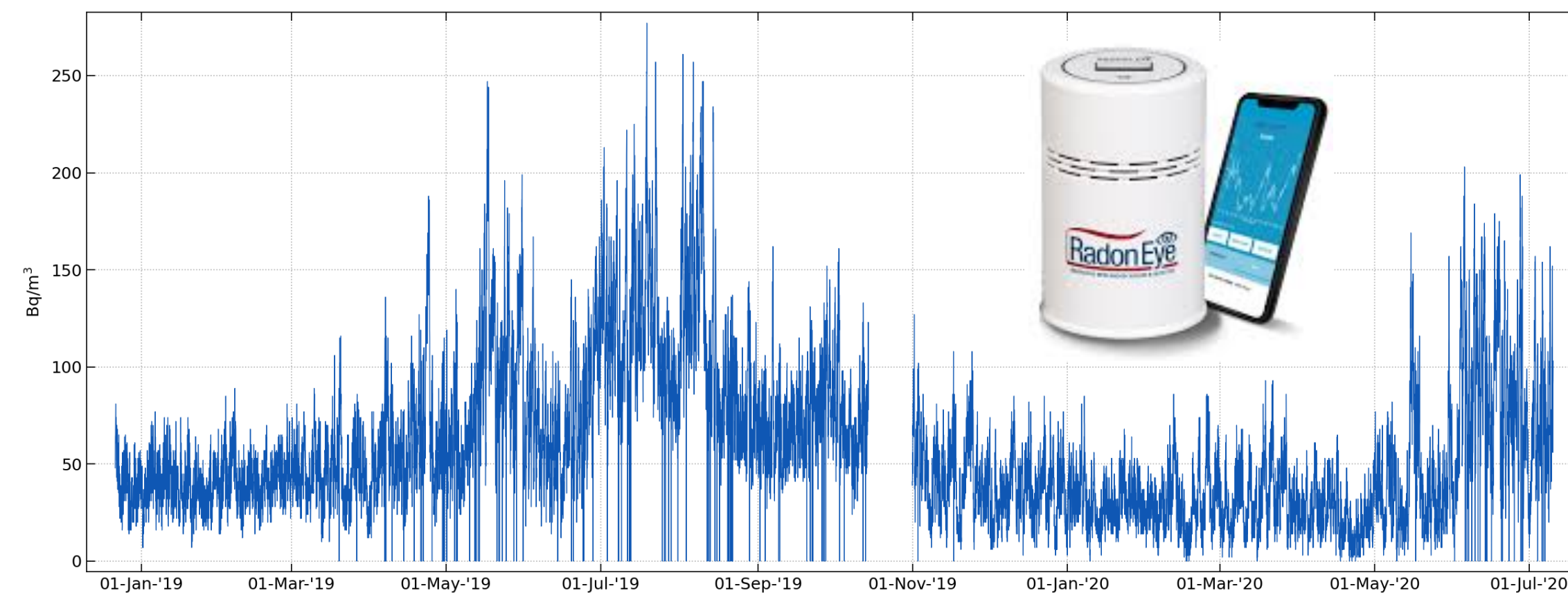
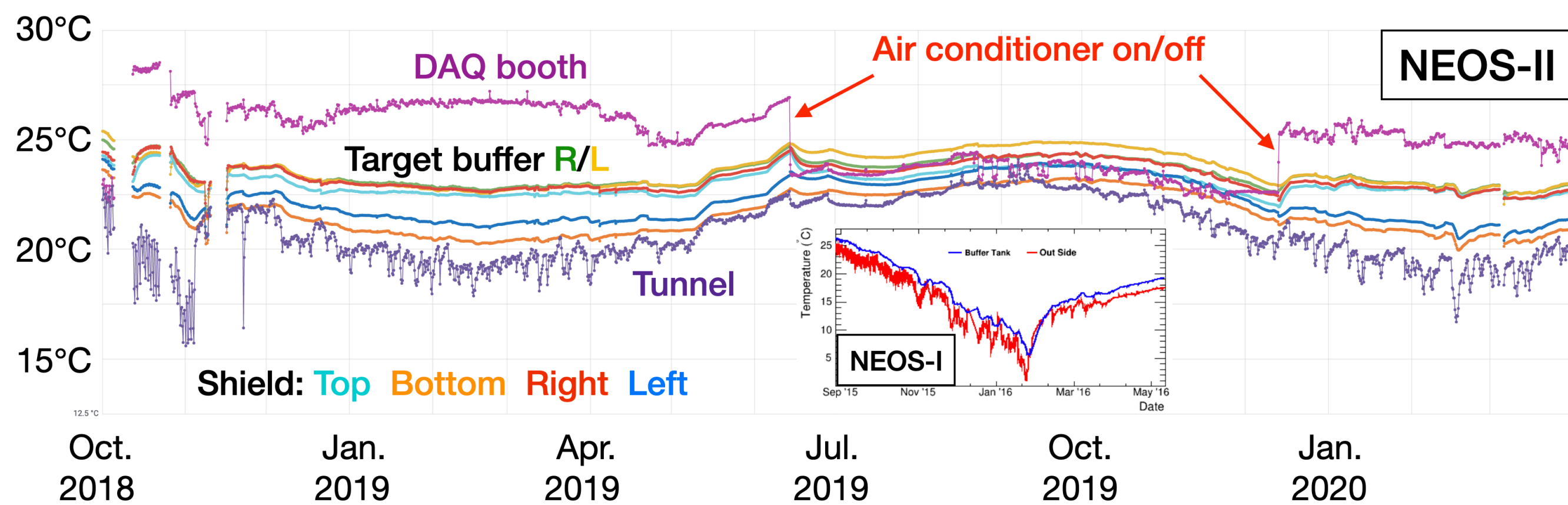
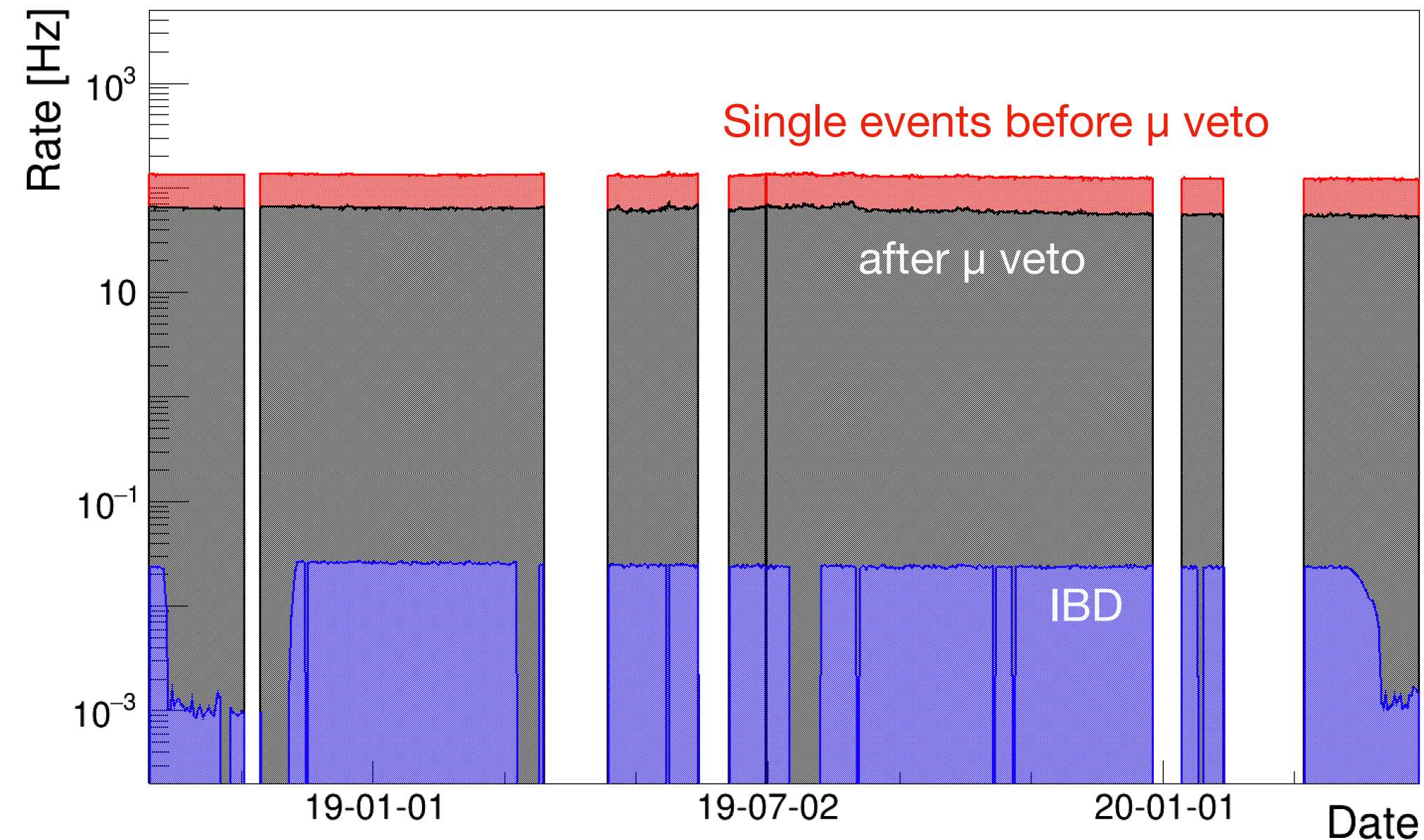
1 kL of homogeneous Gd-LS for IBD target + shield



- Refurbished from NEOS-I detector.
 - Gd-LS newly produced—Gd: 0.5%wt.
 - LAB-based:Ultima Gold F = 9:1 for PSD.
- Target seen by 2×19 8-inch PMTs.
- B-PE (10 cm), Pb (10 cm), Muon counter.
- Calibration: ^{22}Na , ^{60}Co , ^{137}Cs , PoBe, ^{252}Cf
- No γ -catcher layer:
 - Important to properly estimate γ -escaping effect.
- 2-dimensional source calibration.

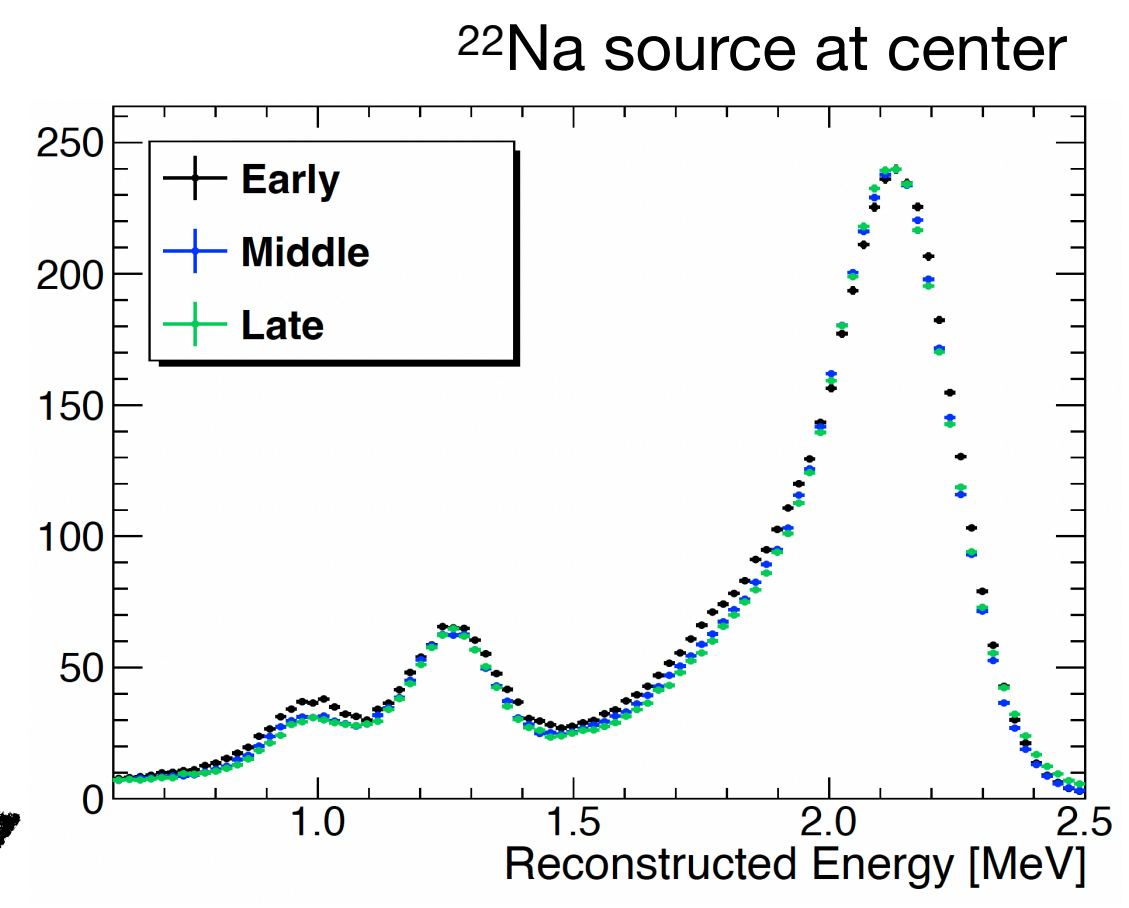
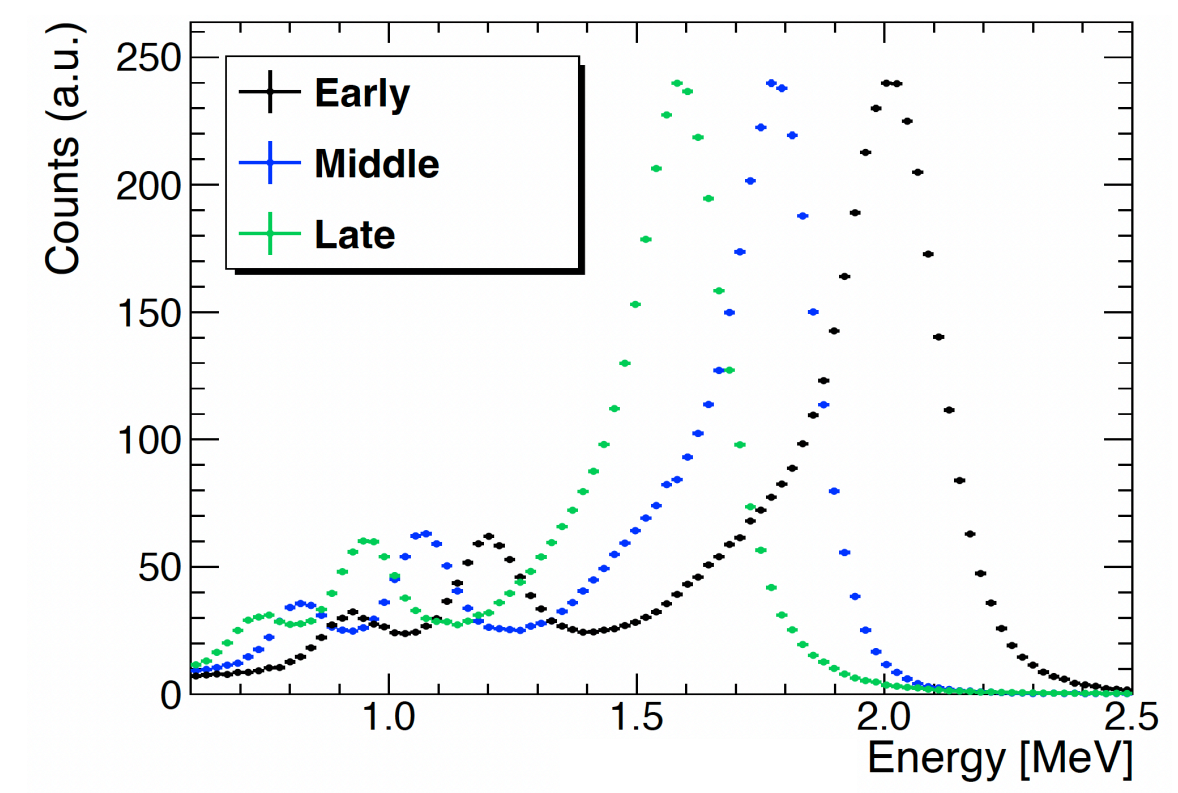
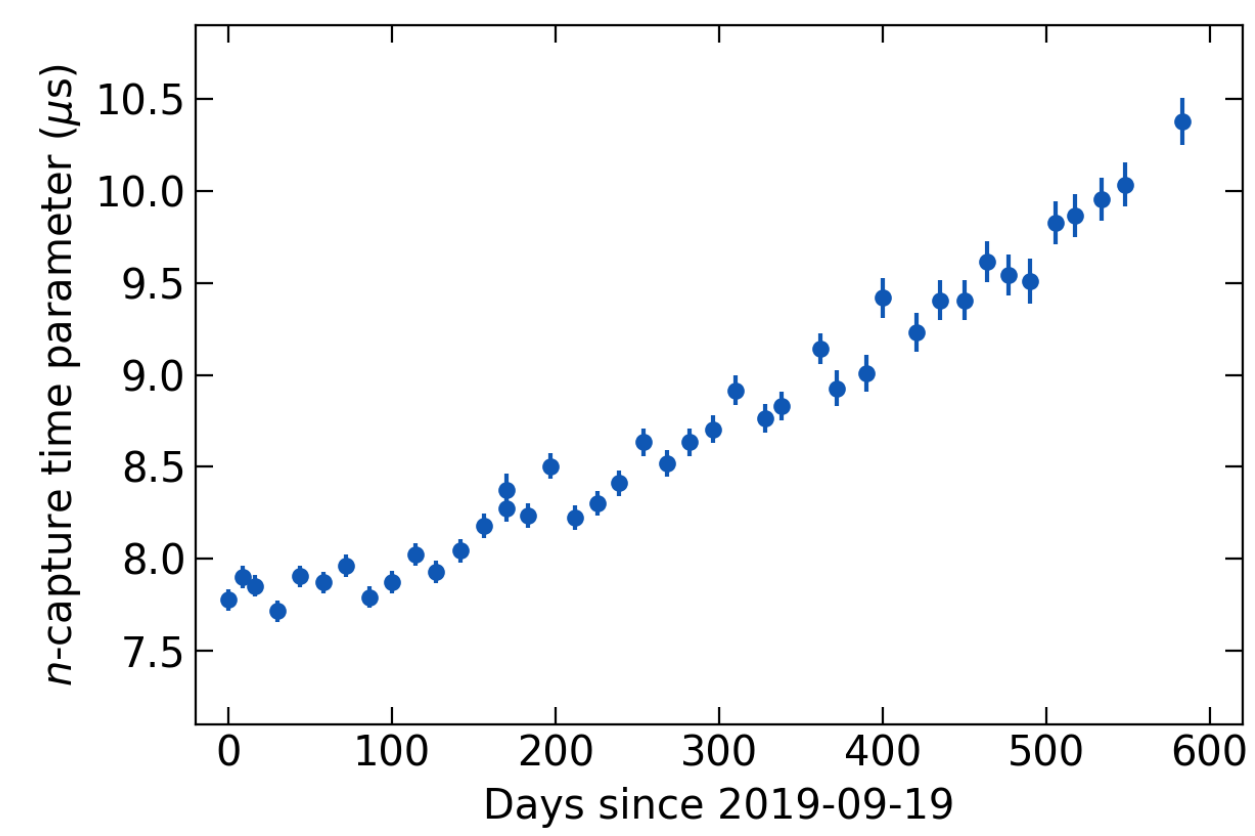
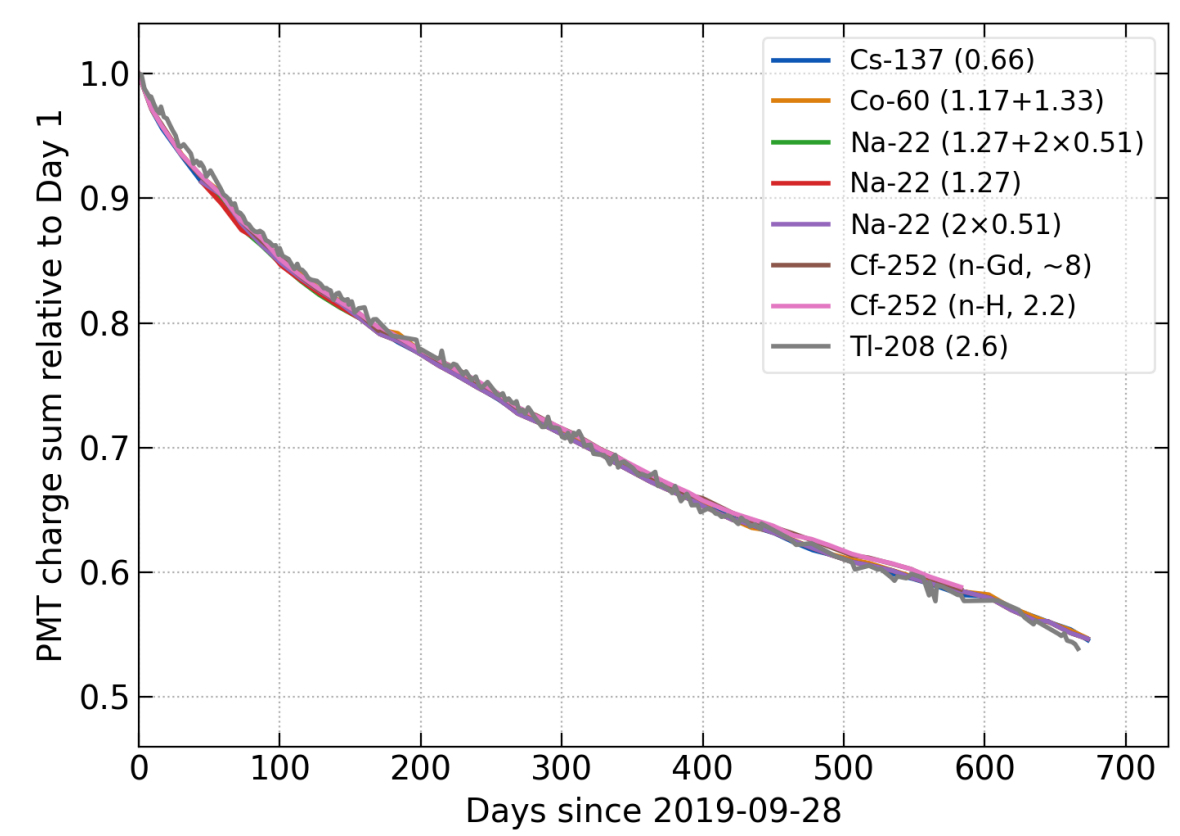
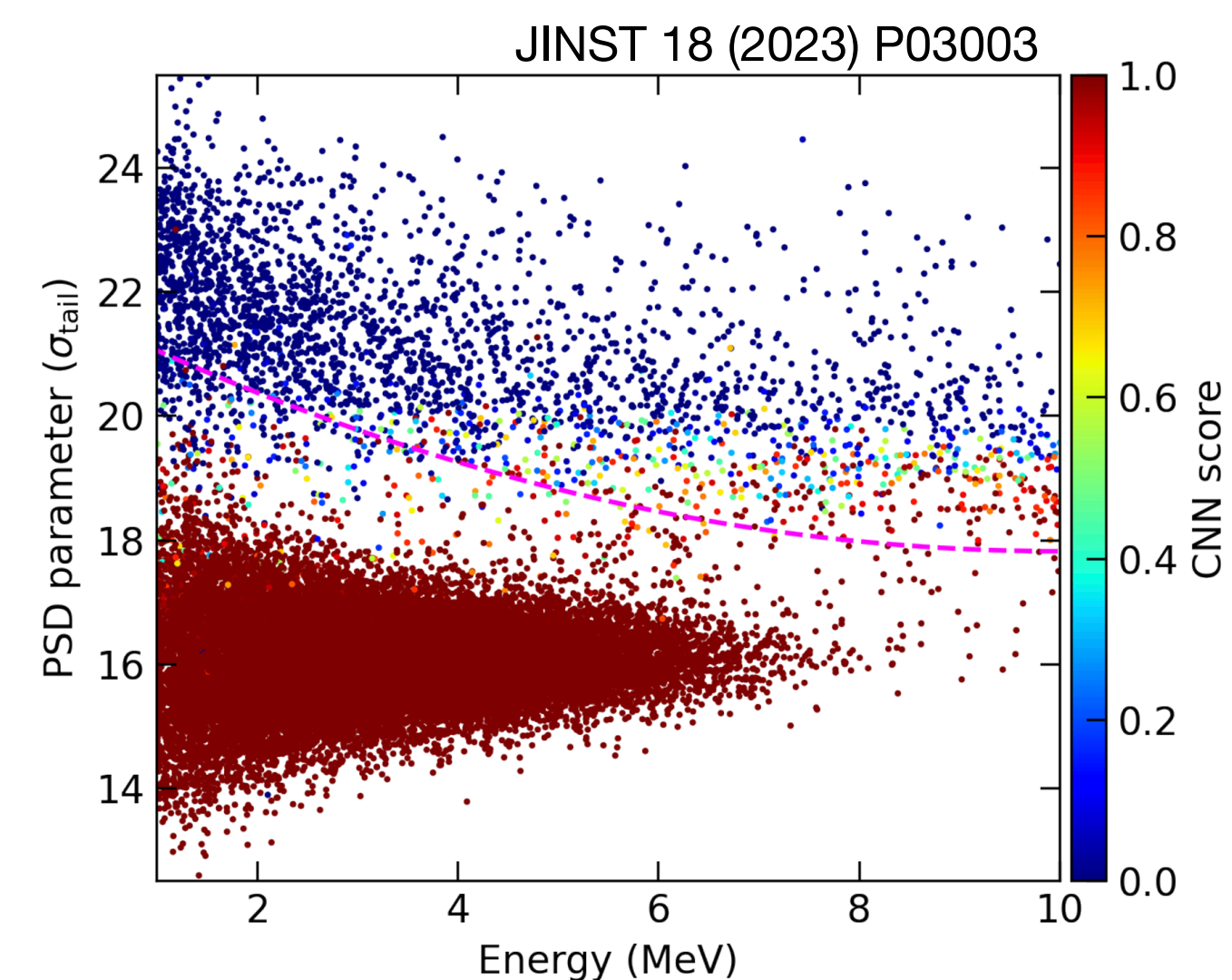
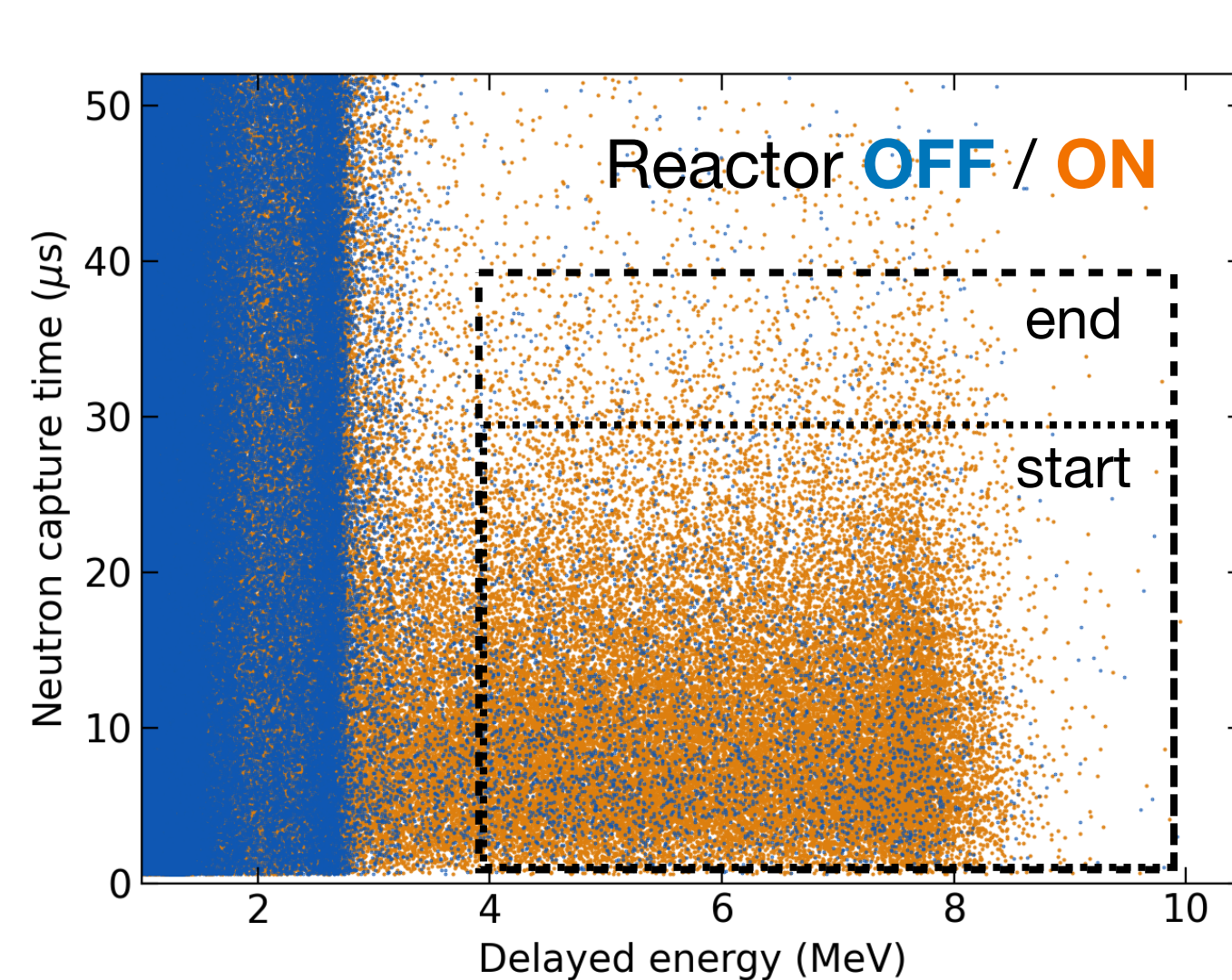
Data taking

- From Sep 2018 to Oct 2020.
- DAQ duty cycle ~ 90%.
 - Physics data for analysis ~ 80%:
 - With the reactor active (ON): 388 days,
 - With the reactor inactive (OFF): 43+74 days.
- Slow monitoring: temperature, radon in the air



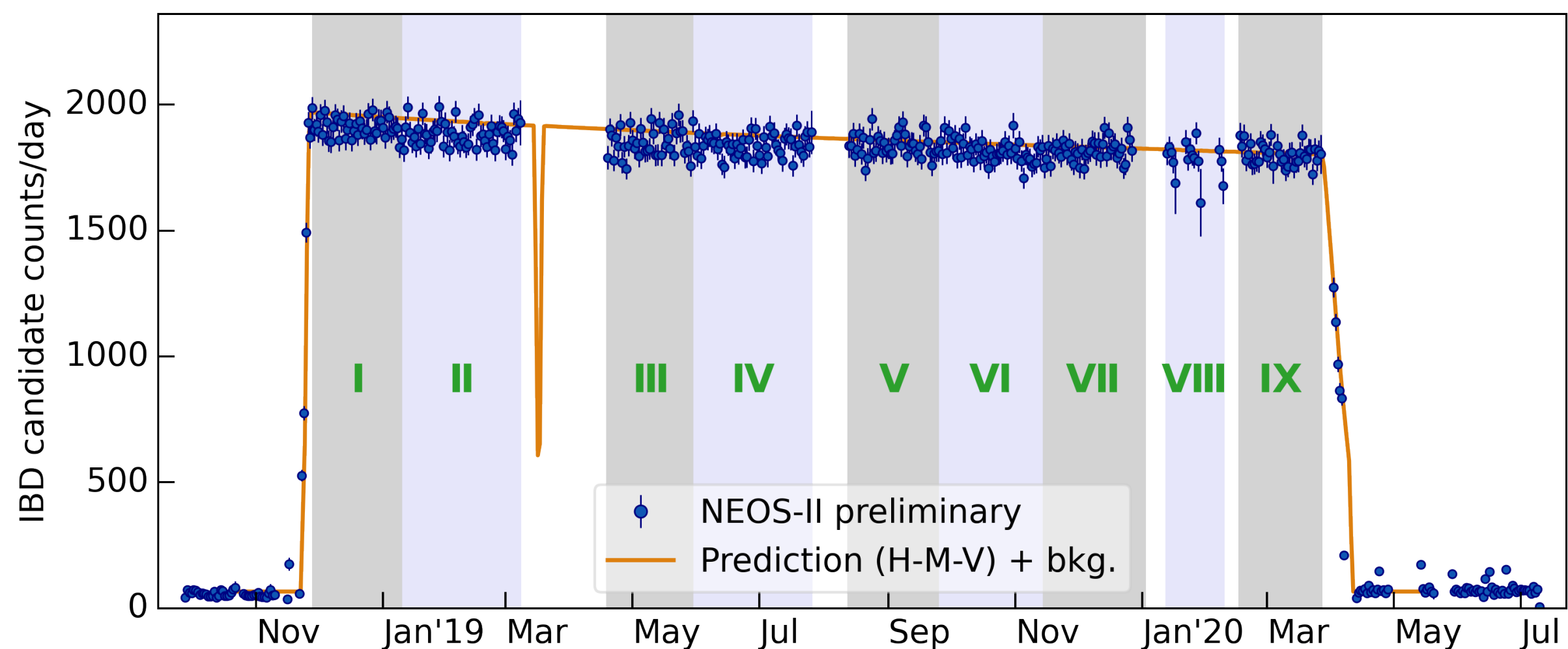
Analysis highlights

- Corrections for:
 - Non-uniform detector response
 - Light output decrease
- Energy smearing: $\sigma \sim 6.5\%$ at 1 MeV.
- Cut evolution utilizing simulation
 - n-capture time and delayed energy
 - PSD: CNN for different data groups
 - Negligible effect on prompt shape

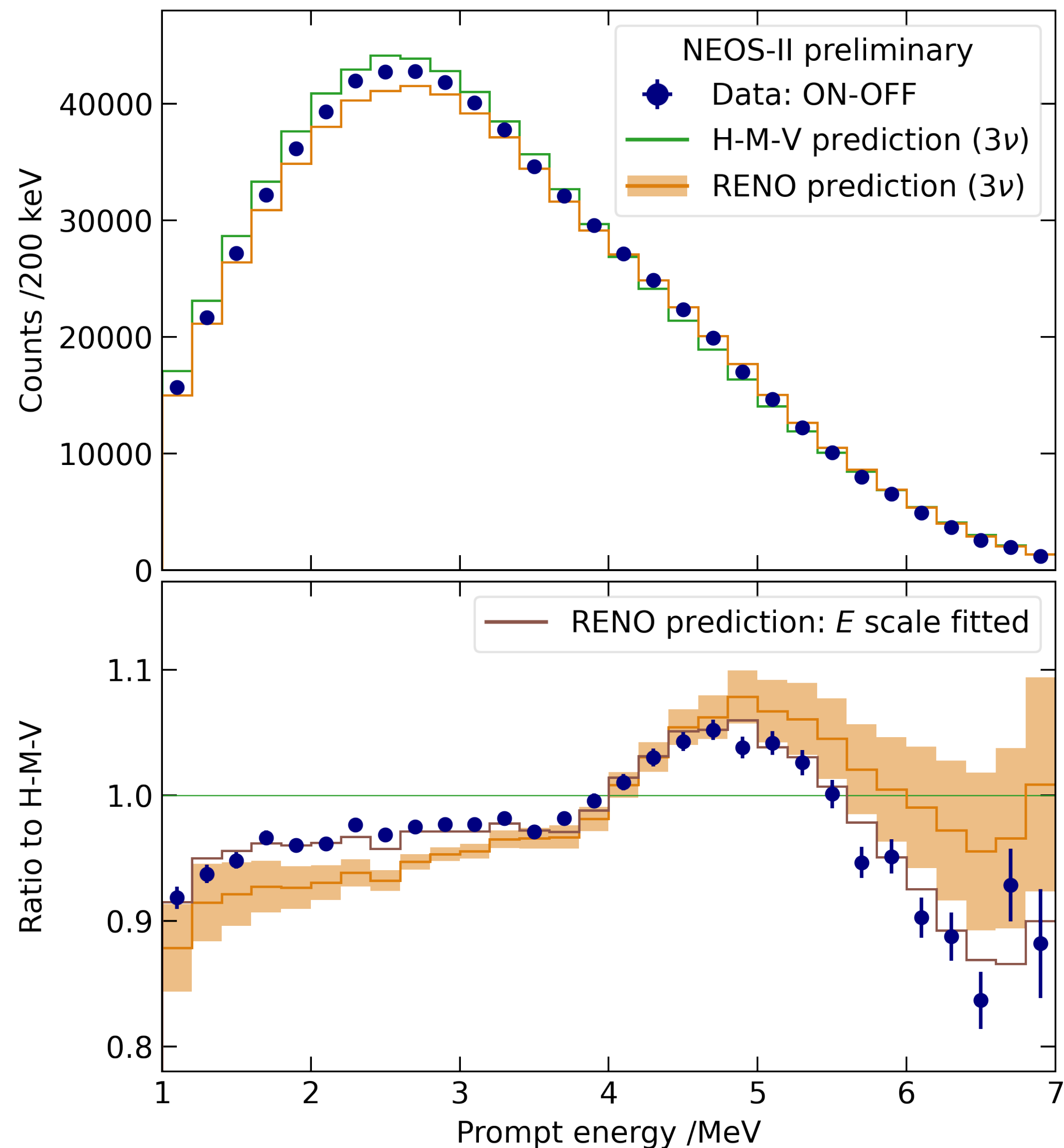


Gain correction+
energy smearing

IBD rate and spectrum

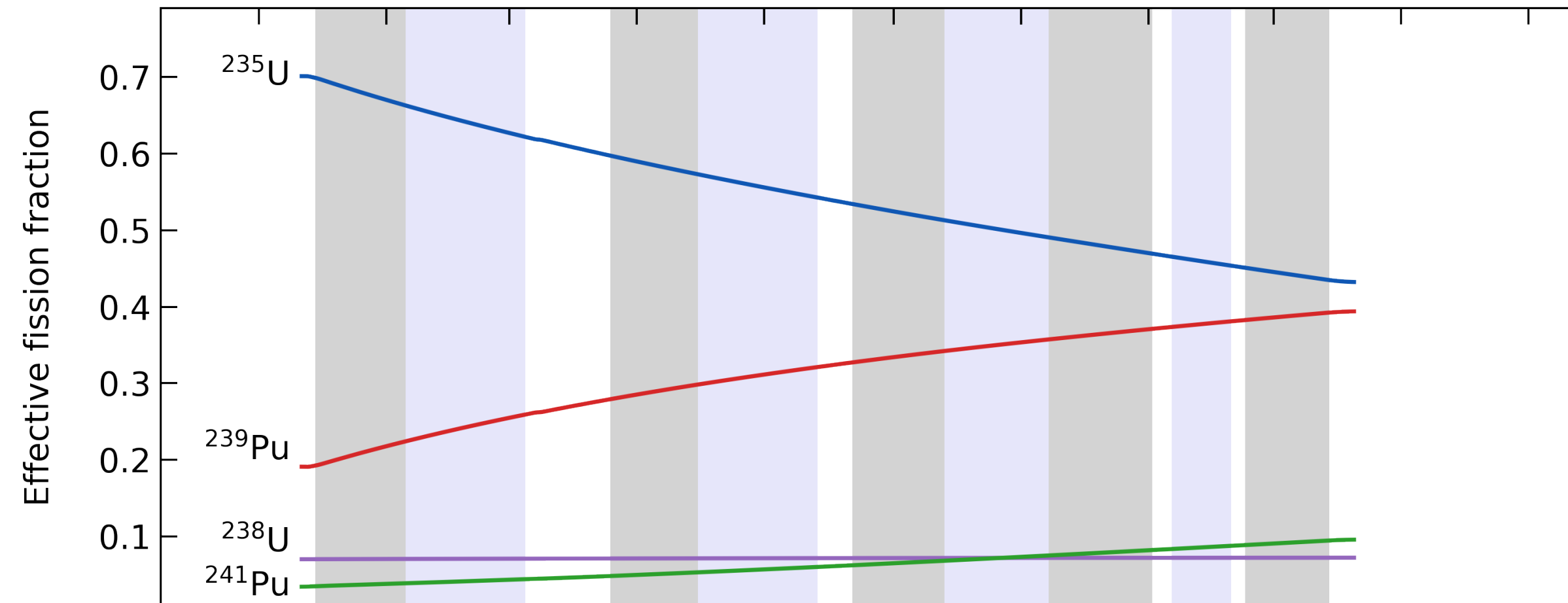
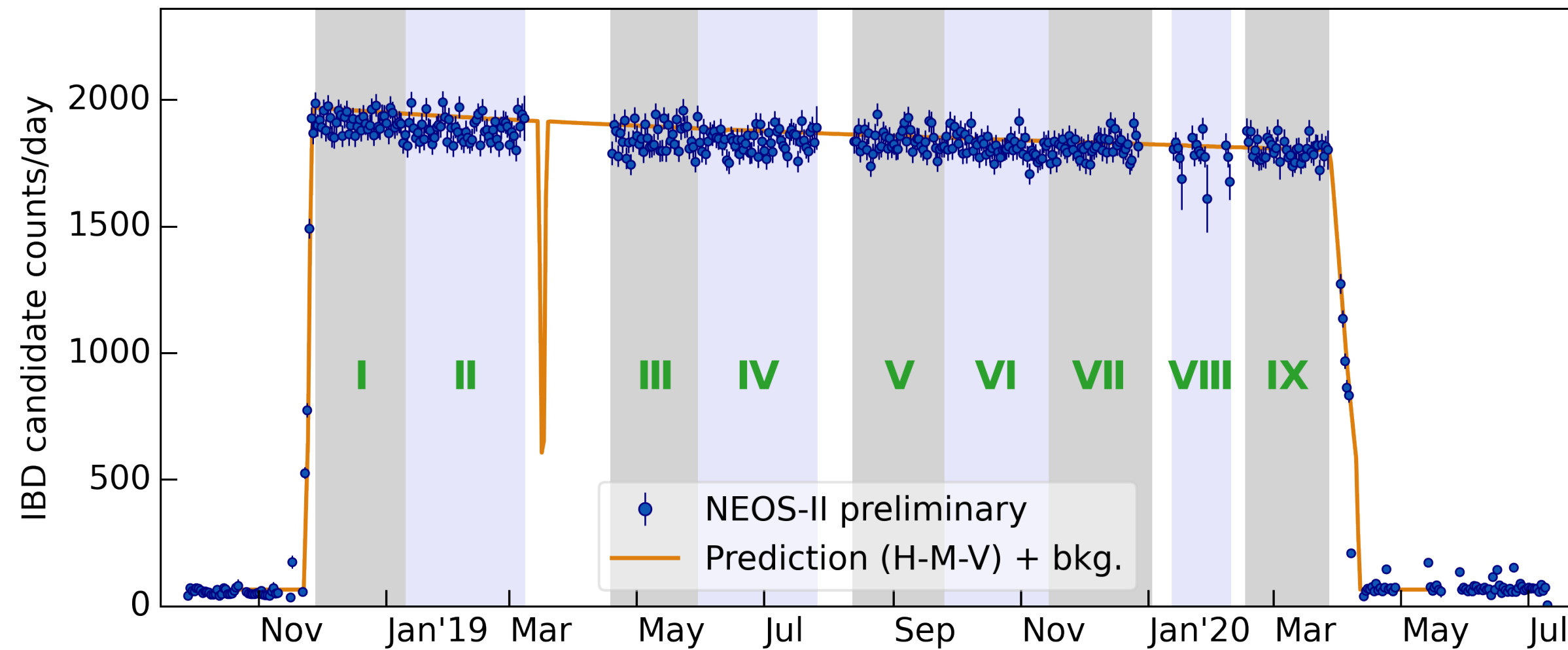


- IBD candidate event rate:
 - OFF: 67 /day
 - ON: ~ 1930 (start) \rightarrow ~ 1800 /day (end).
- Rate lies between Huber-Mueller-Vogel and RENO.
 - $-2.7 \pm 2.5\%$ of $\phi(\text{Huber+Mueller}) \times \sigma(\text{Vogel})$
- Energy scale mismatch.
 - E scale parameter released in the oscillation analysis.



IBD yield and its evolution

Extraction of ^{235}U , ^{239}Pu components



*Hanbit-6 contribution (~0.8%) is ignored in the prediction plot. Hanbit-3,4 were off.

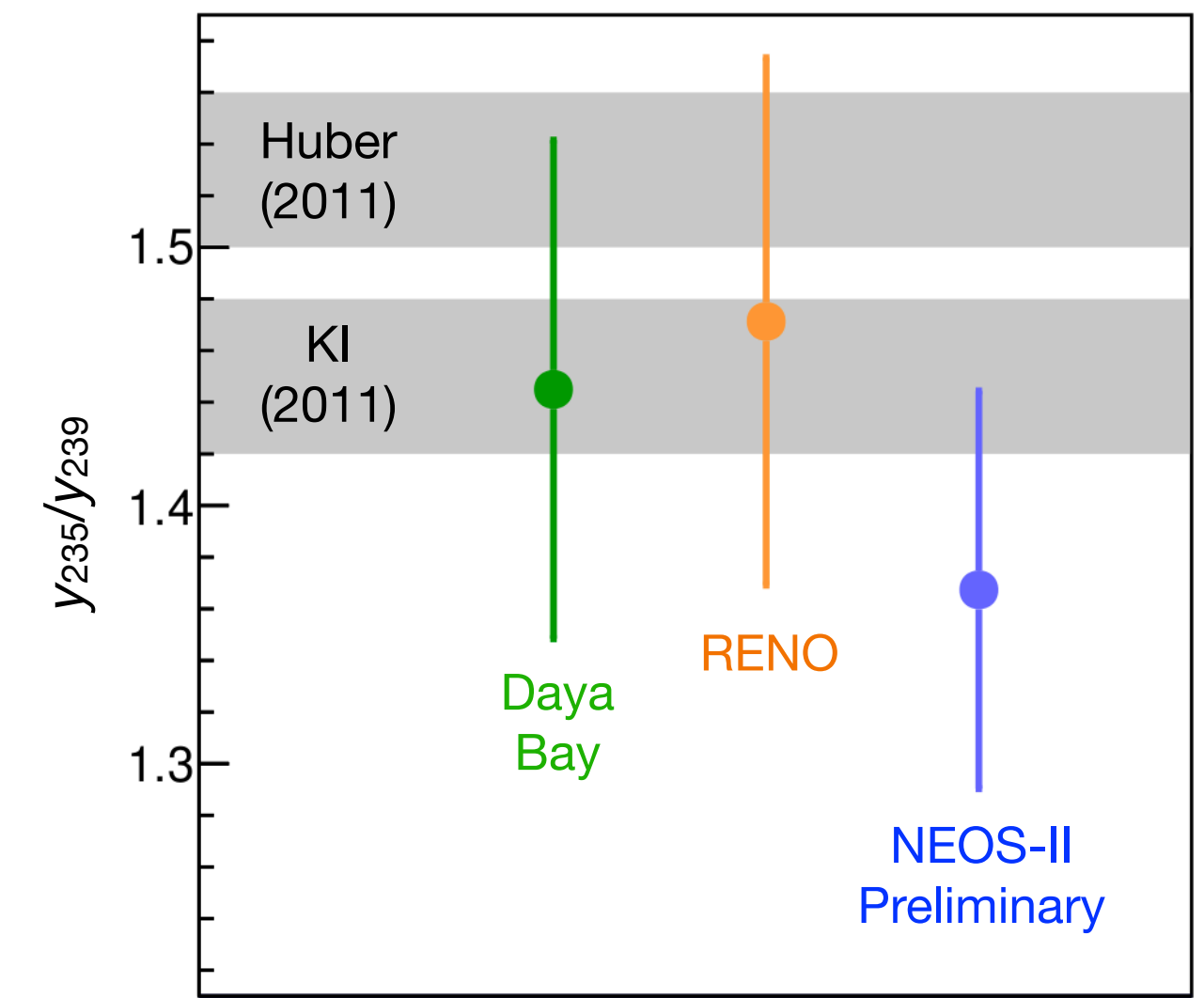
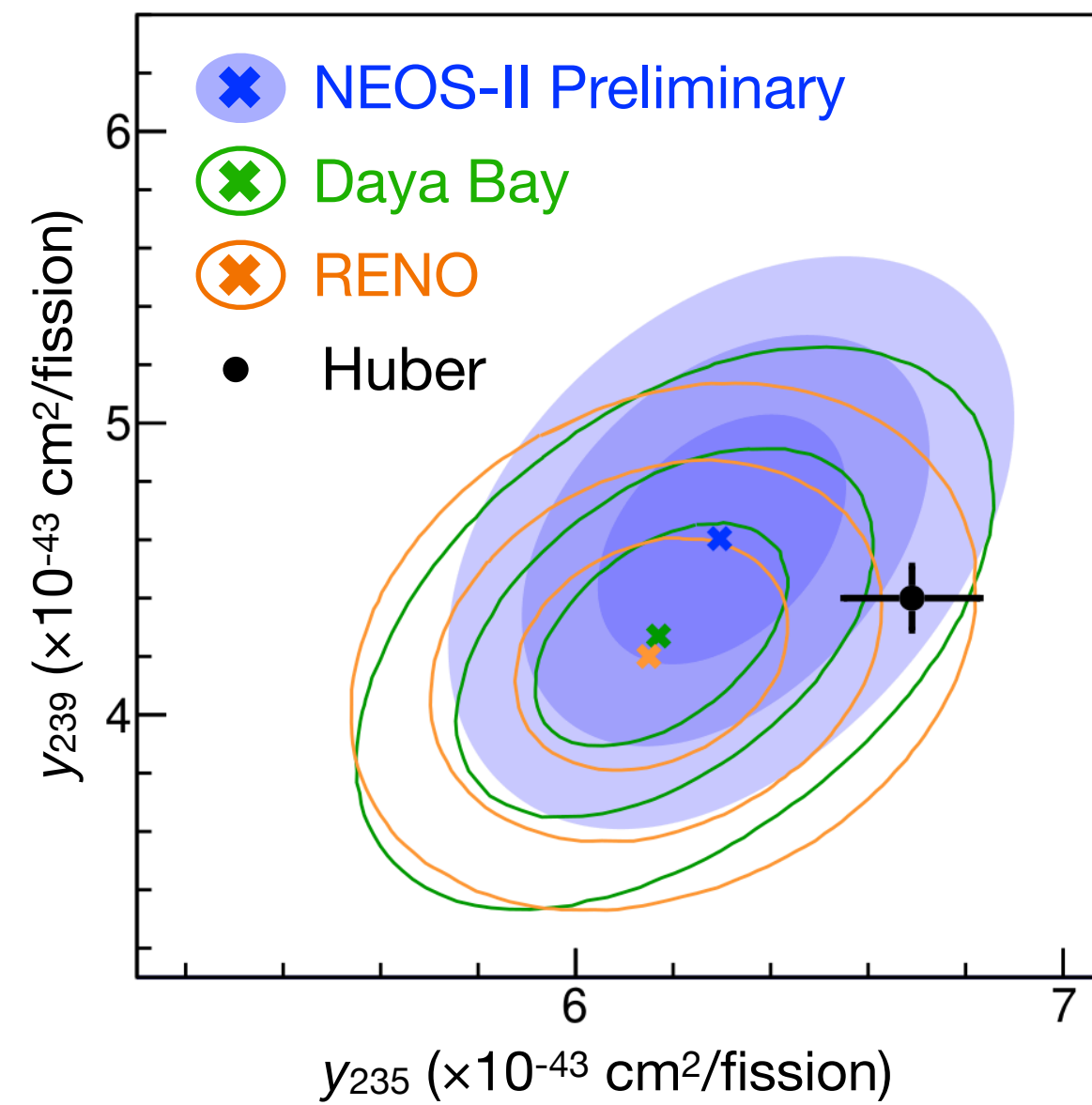
- IBD yield (10^{-43} cm²/fission):

$$y_{235} = 6.32 \pm 0.17,$$

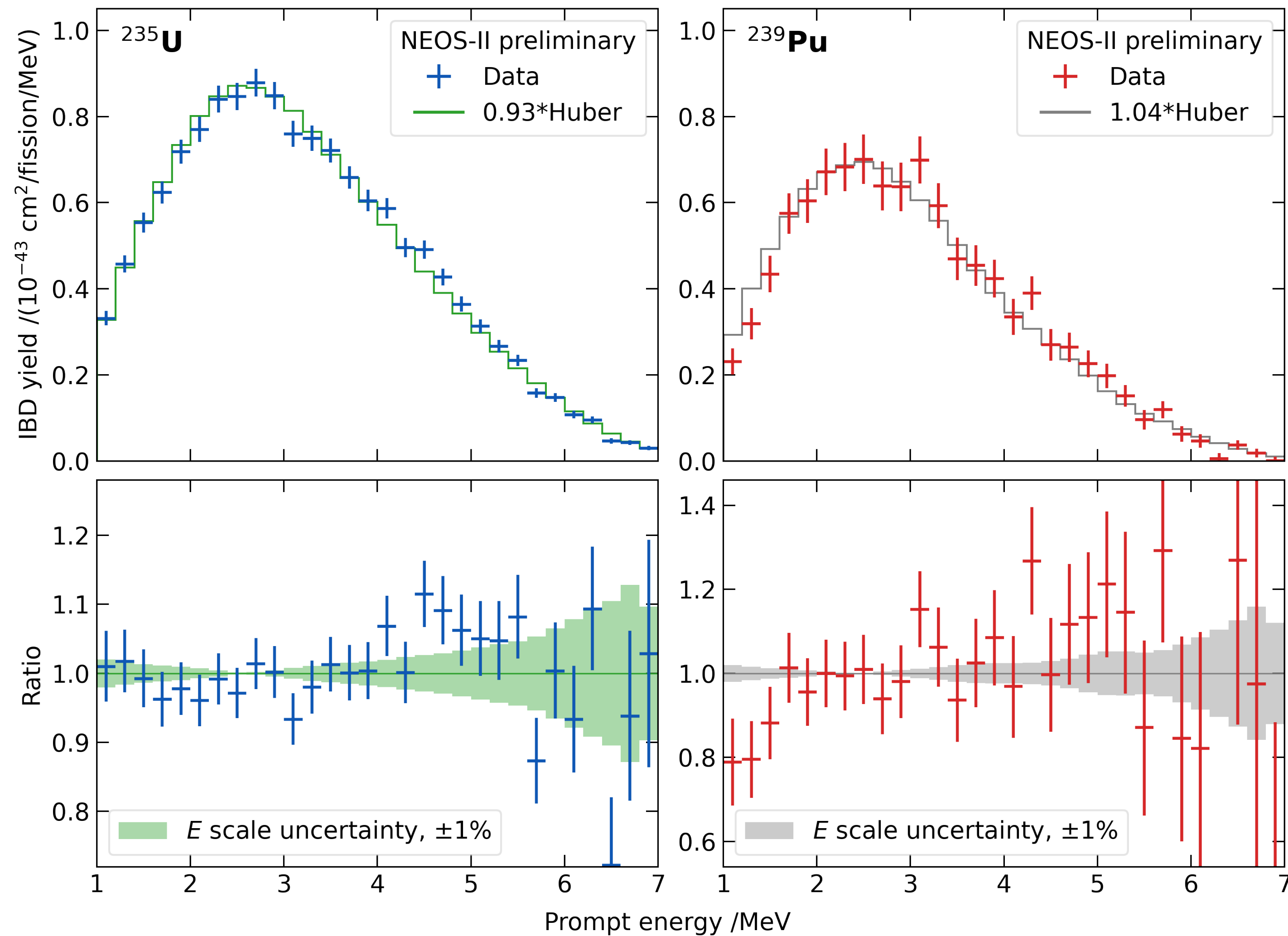
$$y_{239} = 4.70 \pm 0.28.$$

- $y_{235}/y_{239} = 1.37 \pm 0.08.$

- y_{238}/y_{241} : constrained at Huber-Muellers $\pm 10\%.$

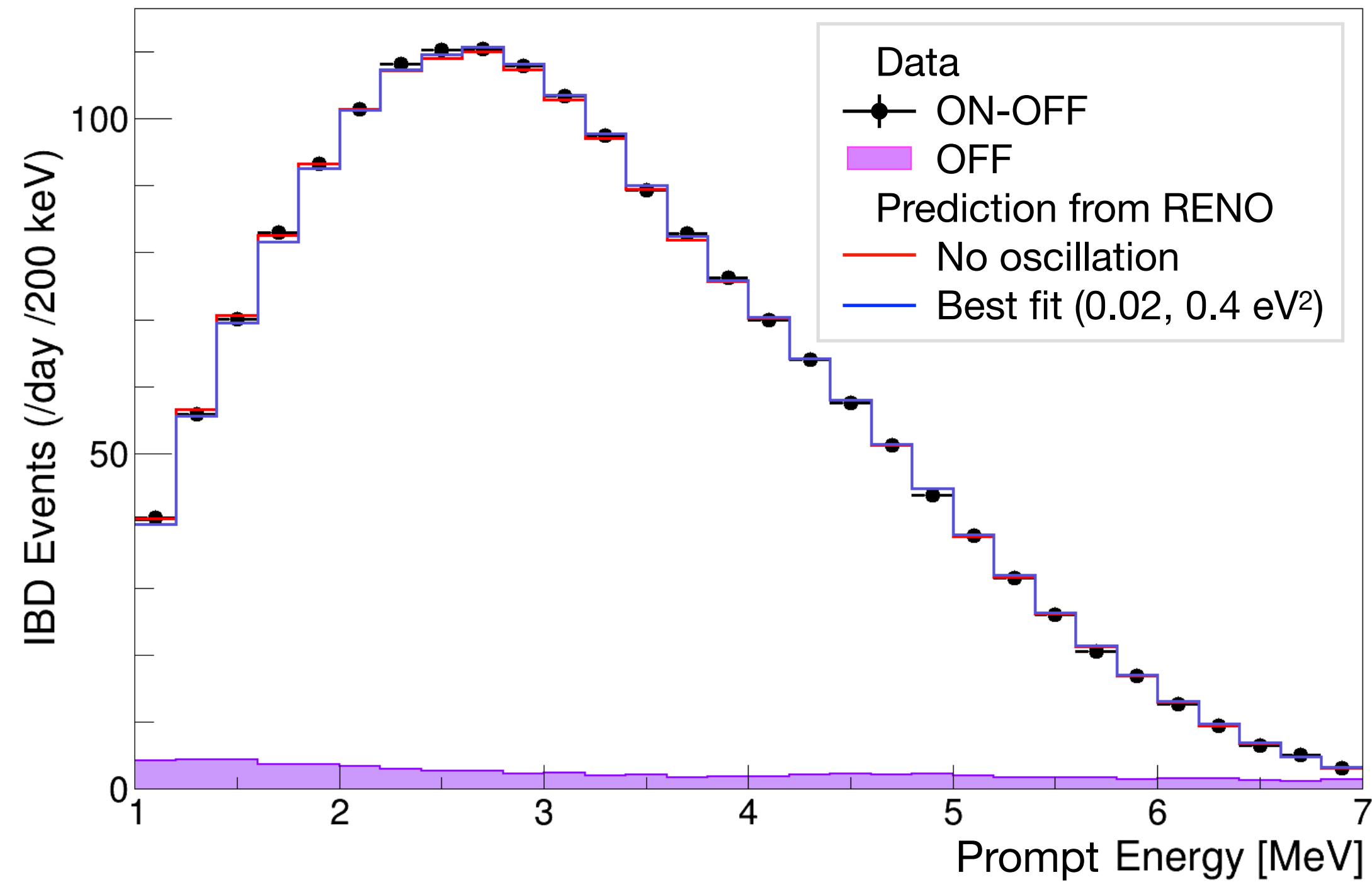


IBD spectrum decomposition

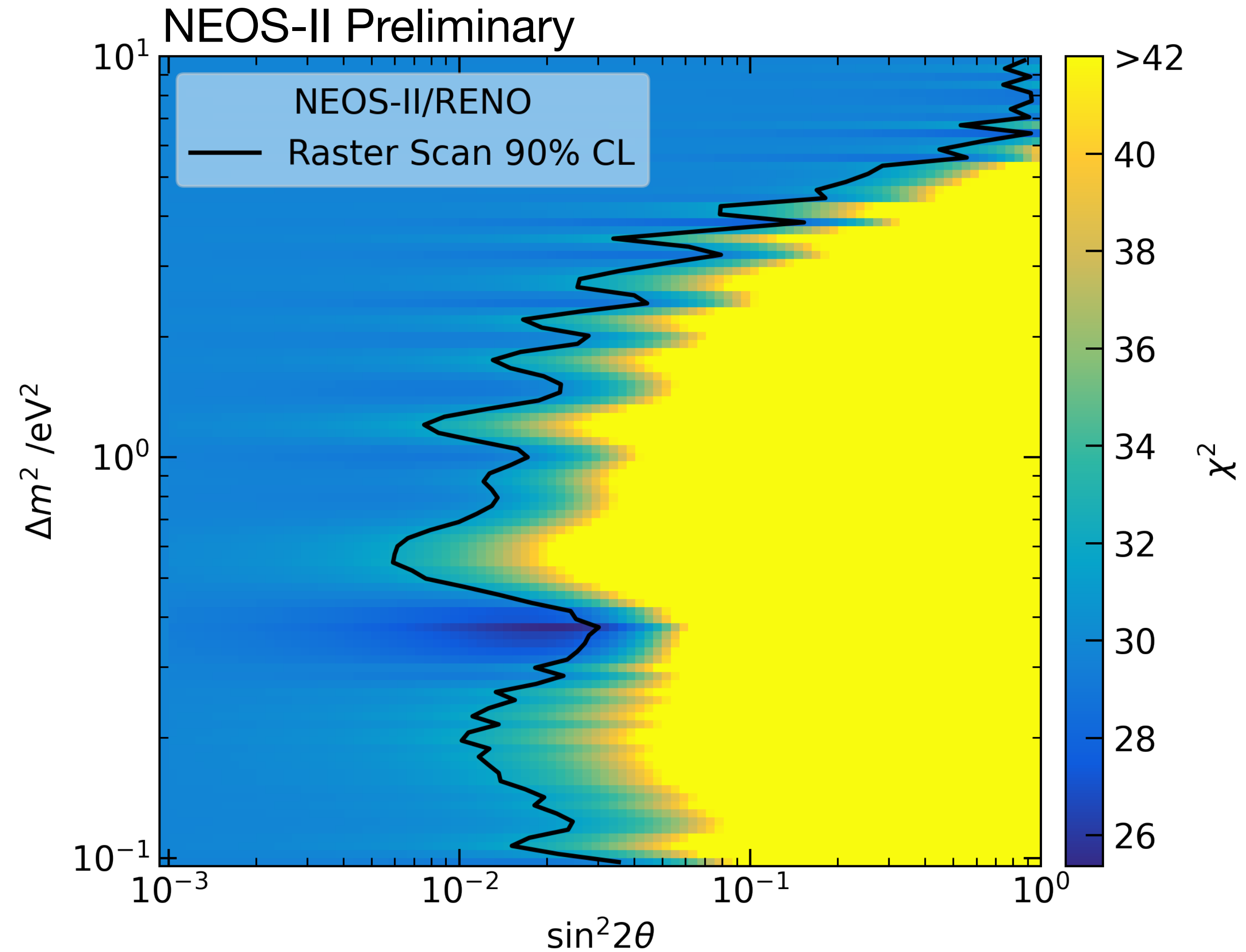


- Large statistical fluctuations in the ^{239}Pu spectrum.
- Inconclusive for a sole isotope responsible for the bump.

SBL oscillation: negative result

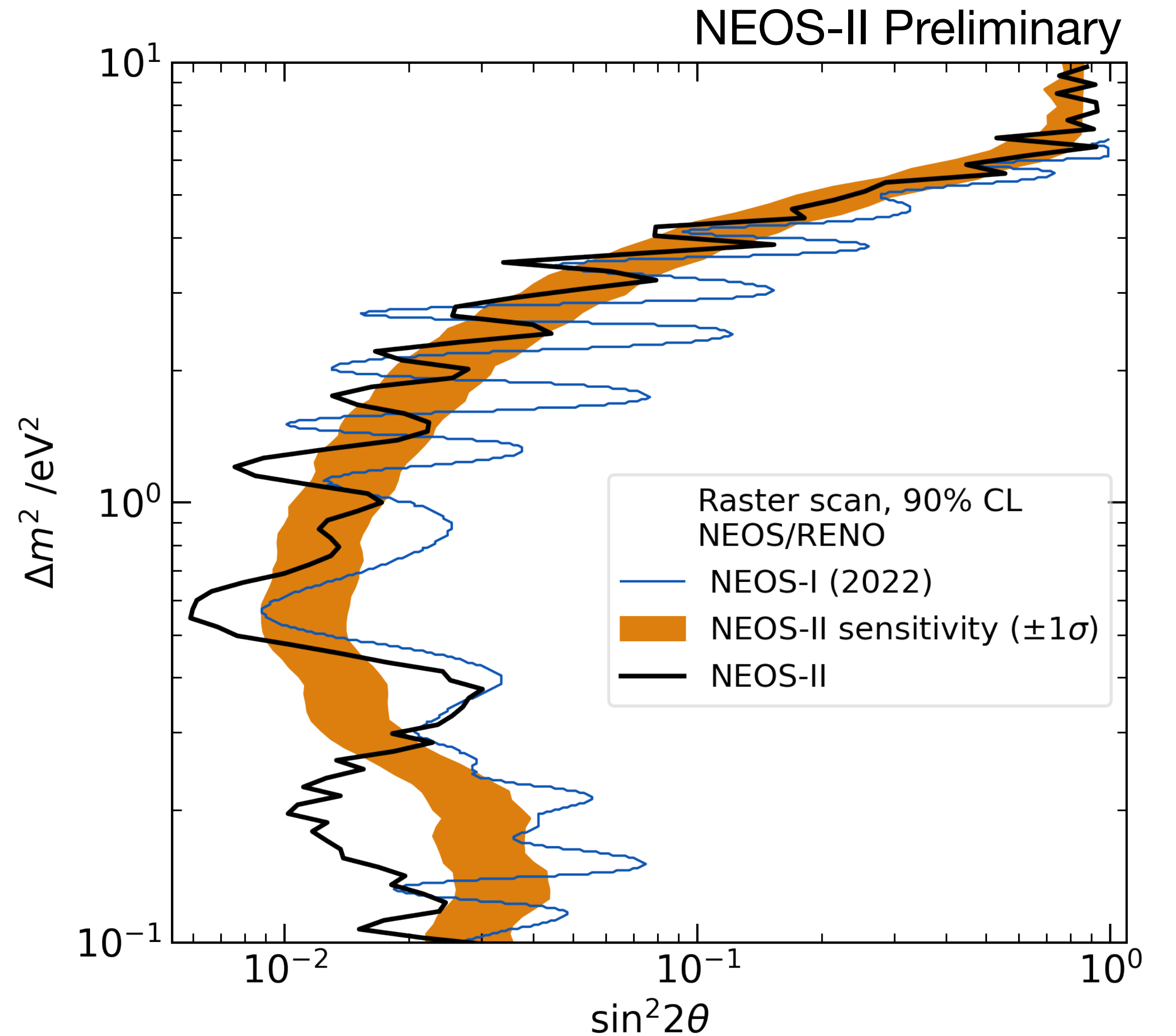
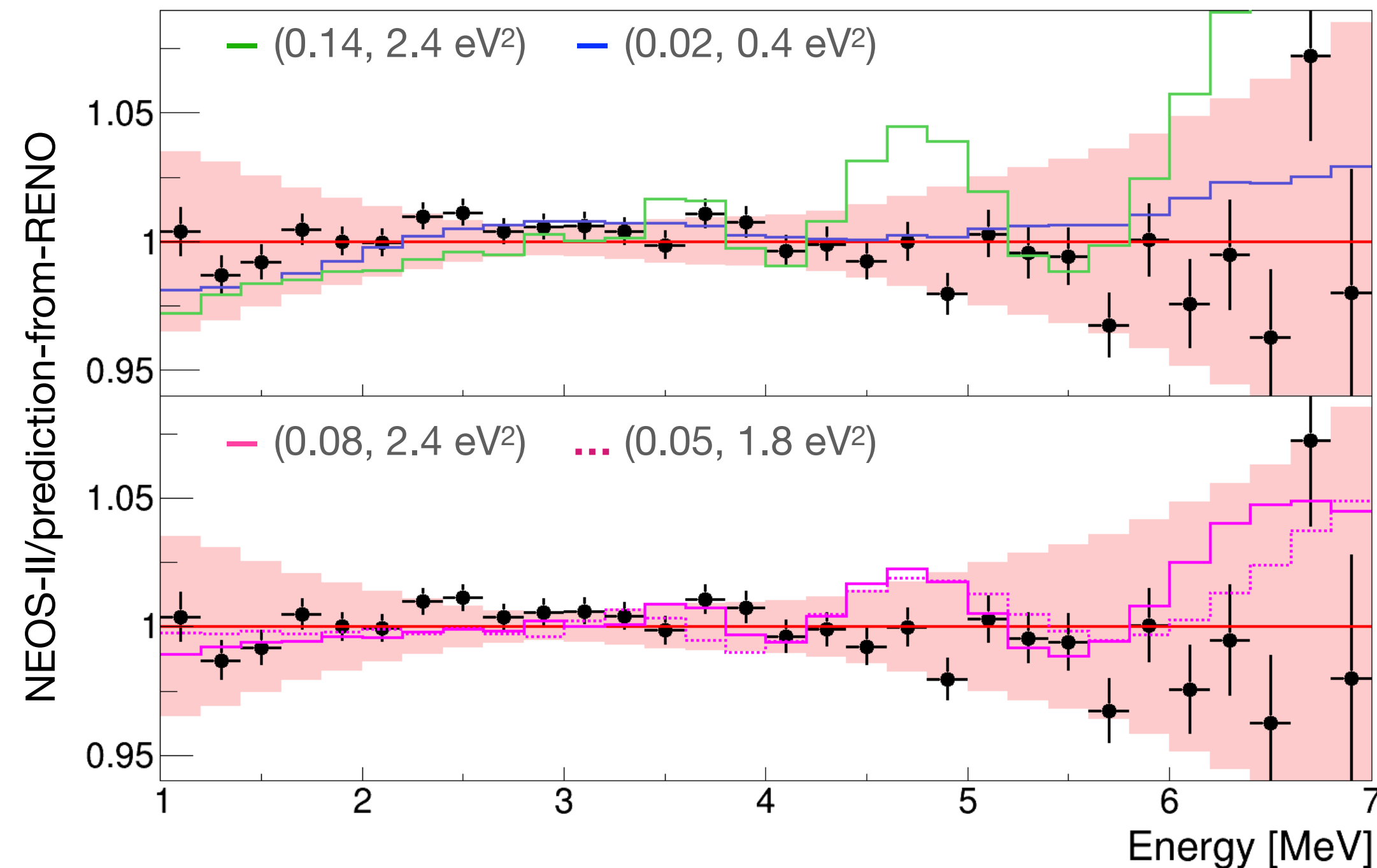


- Best fit at $(\sin^2 2\theta, \Delta m^2) = (0.02, 0.4 \text{ eV}^2)$,
- $\Delta\chi^2 = \chi^2_{3\nu} - \chi^2_{4\nu} = 29.7 - 25.4 = 4.3$



Comparison with NEOS-I result

$(\sin^2 2\theta, \Delta m^2 / \text{eV}^2)$	χ^2	Note
x	29.7	3v
(0.02, 0.4)	25.4	Best fit (this work)
(0.08, 2.4)	44.7	NEOS-I/RENO 1st best
(0.05, 1.8)	46.1	NEOS-I/RENO 2nd best
(0.14, 2.4)	77.1	RAA best



Summary

Results are being finalized

- Measured IBD yield: $-2.7_{\pm 2.5\%}$ of Huber-Mueller-Vogel prediction.
- $(y_{235}, y_{239}) = (6.32_{\pm 0.17}, 4.70_{\pm 0.28}) \times 10^{-43} \text{ cm}^2/\text{fission}$, $y_{235}/y_{239} = 1.37_{\pm 0.08}$.
- No evidence for the short baseline oscillation from the $3+1\nu$ model.

