



# JSNS<sup>2</sup> Status & Results

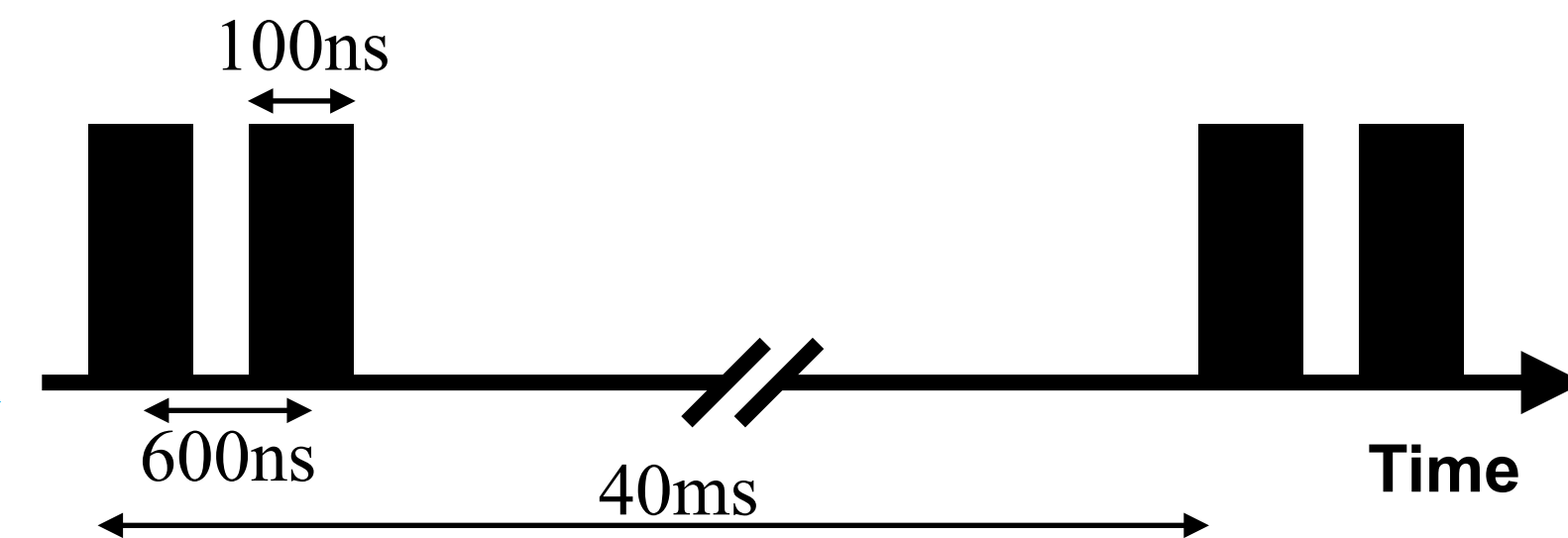
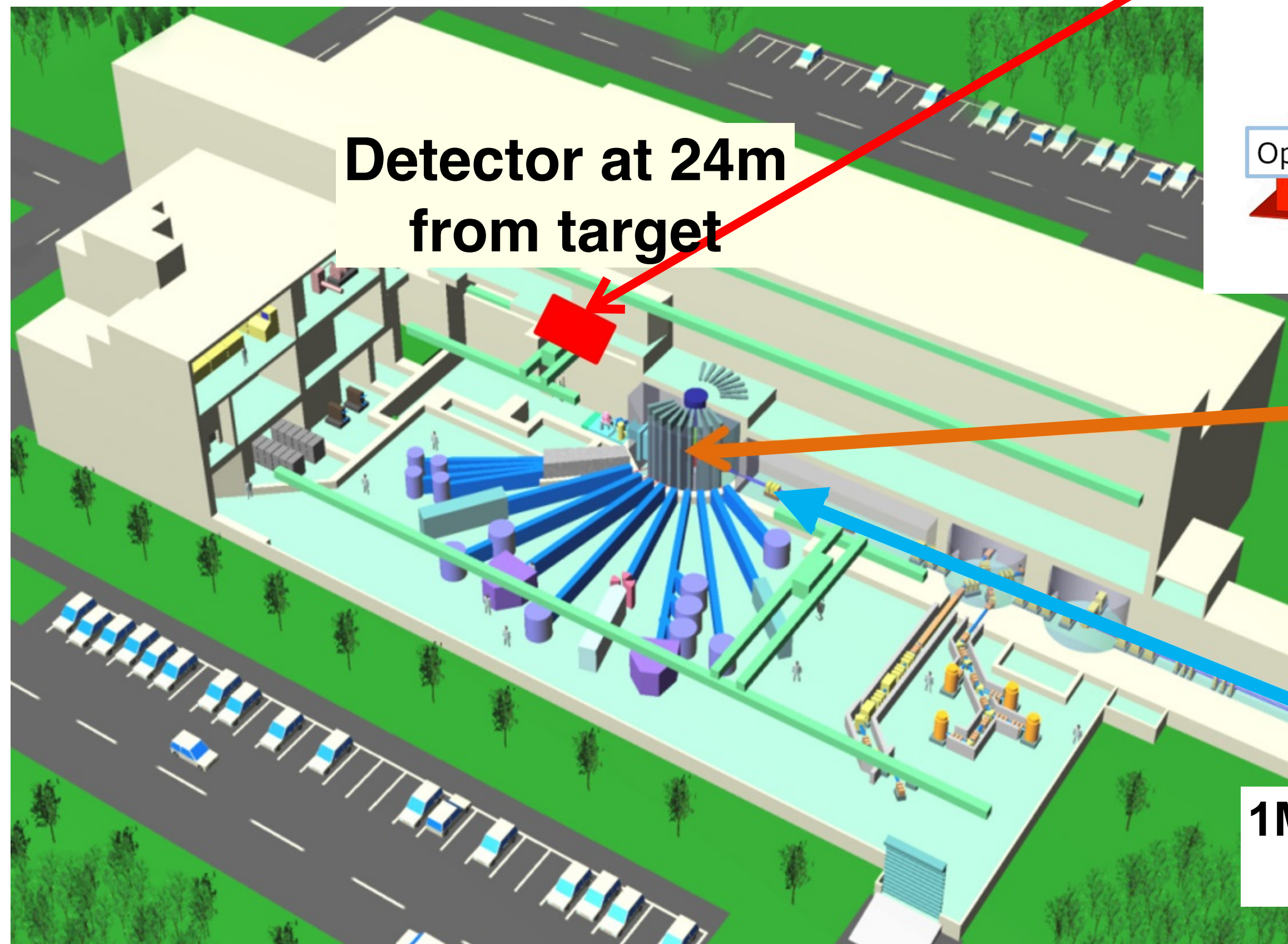
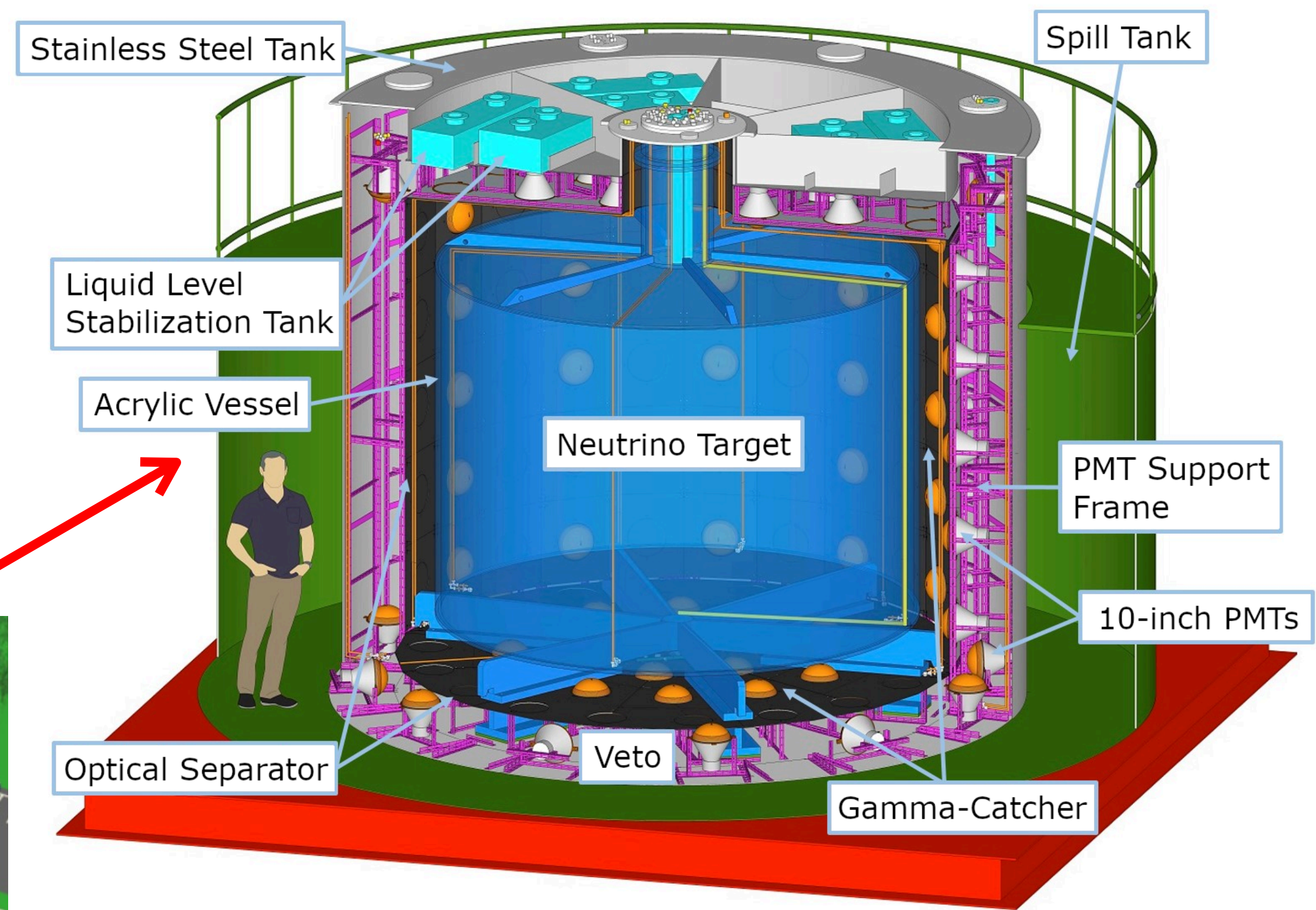


Eric Marzec for the JSNS<sup>2</sup> Collaboration  
University of Michigan

June 17 2024

# Detector & Facility

- Located at Materials & Life Sciences Facility (MLF) at J-PARC
- Detector is 17tons Gd-LS w/ 10% DIN & 31tons LS outer region
- Observed by 96 10-inch inner PMTs & 24 veto PMTs

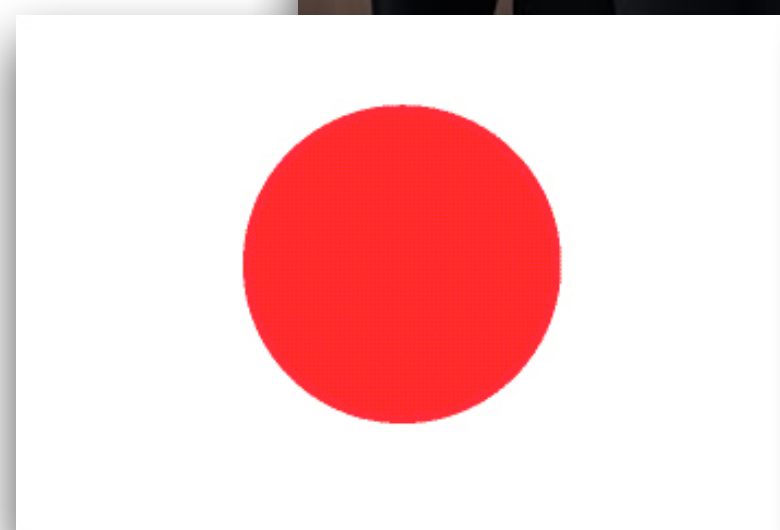


1MW, 3GeV pulsed Proton beam

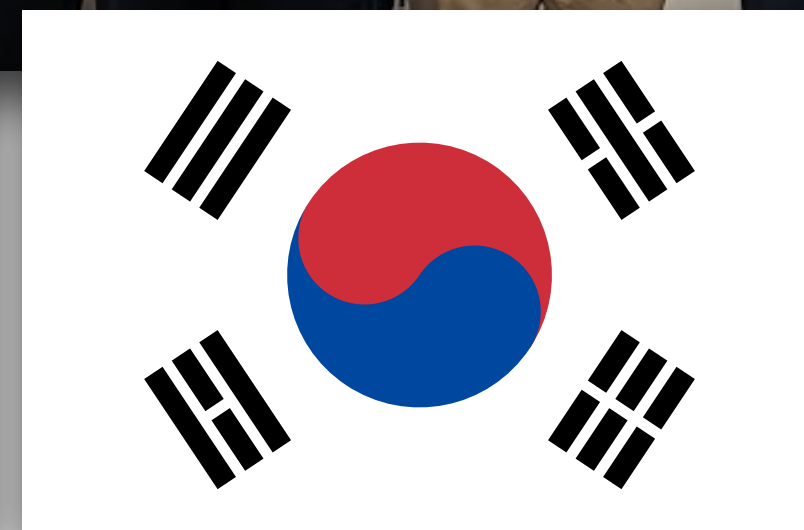
# JSNS<sup>2</sup> & JSNS<sup>2</sup>-II Collaboration

JSNS<sup>2</sup> Collaboration - 61 Members across 23 institutions & 5 countries

JSNS<sup>2</sup> Collaboration - February 2024



KEK  
JAEA  
J-PARC  
Tsukuba University  
Osaka University  
Tohoku University  
Kitasato University  
Kyoto Sangyo University



Soongsil University  
Dongshin University  
Seoyeong University  
Kyung Hee University  
Gwangju Institute of Science and Technology  
Seoul National University of Science and Technology  
Sungkyunkwan University  
Chonnam National University  
Jeonbuk National University  
Kyungpook National University



Brookhaven National Laboratory  
University of Michigan  
University of Utah



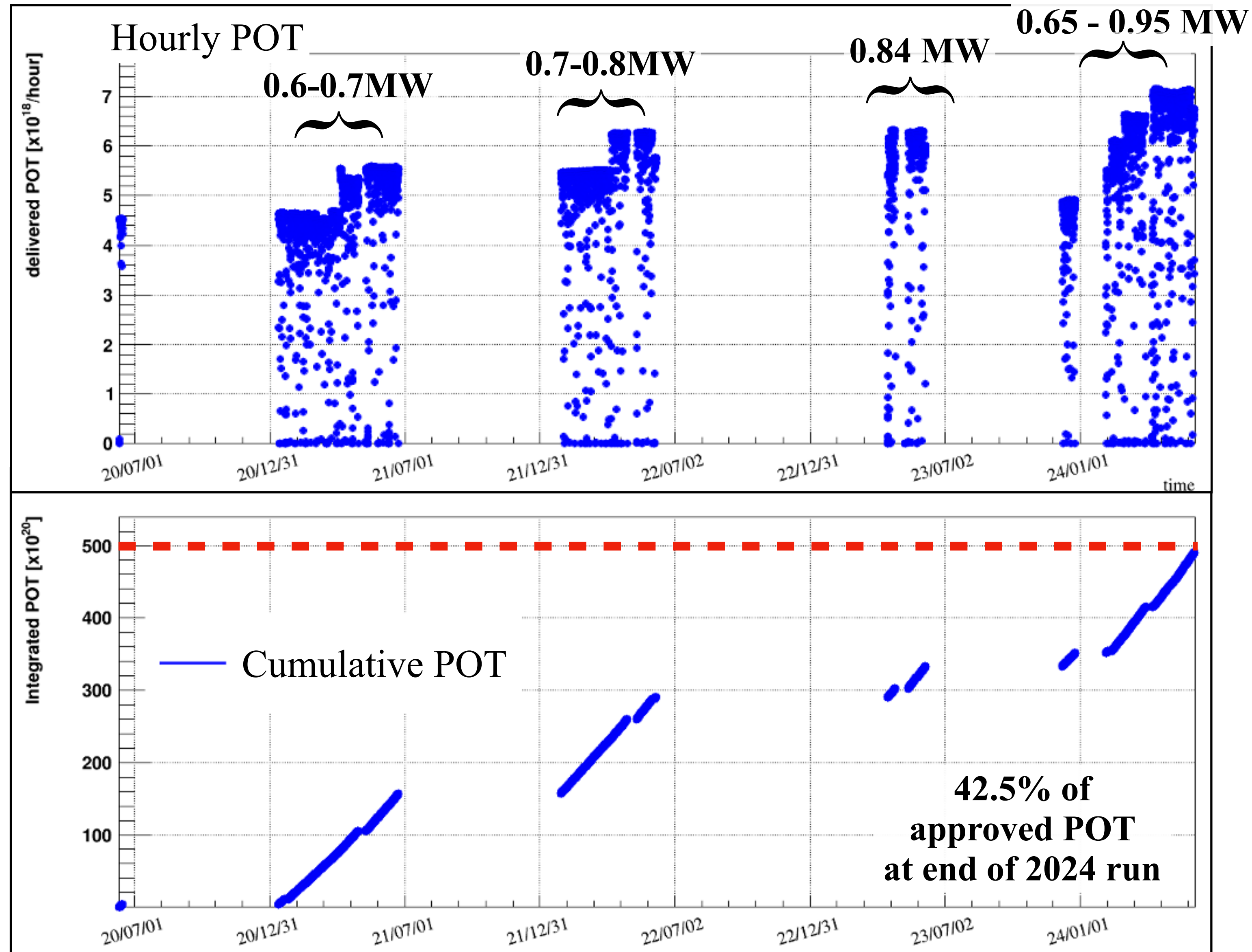
Sun Yat-sen University



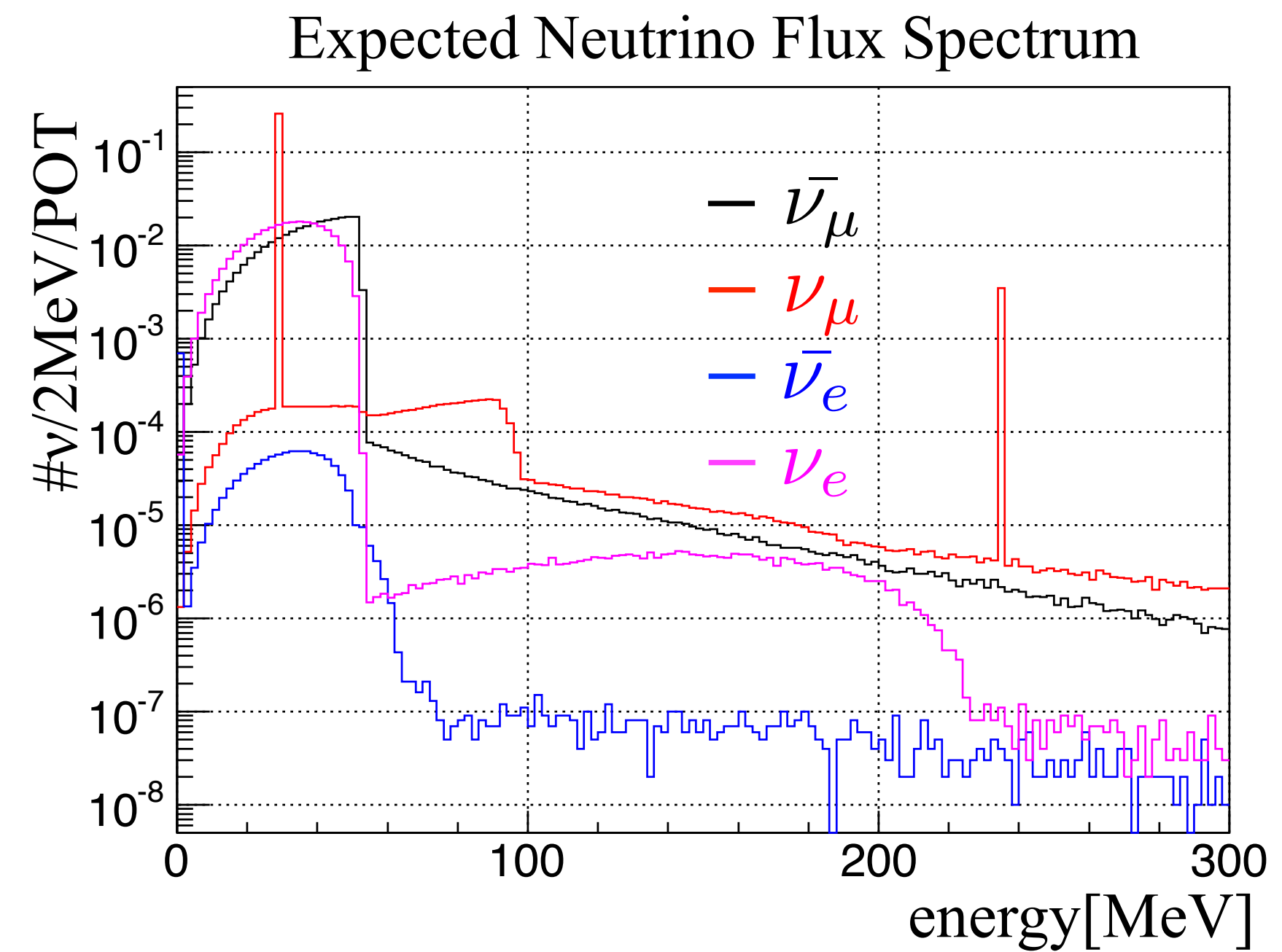
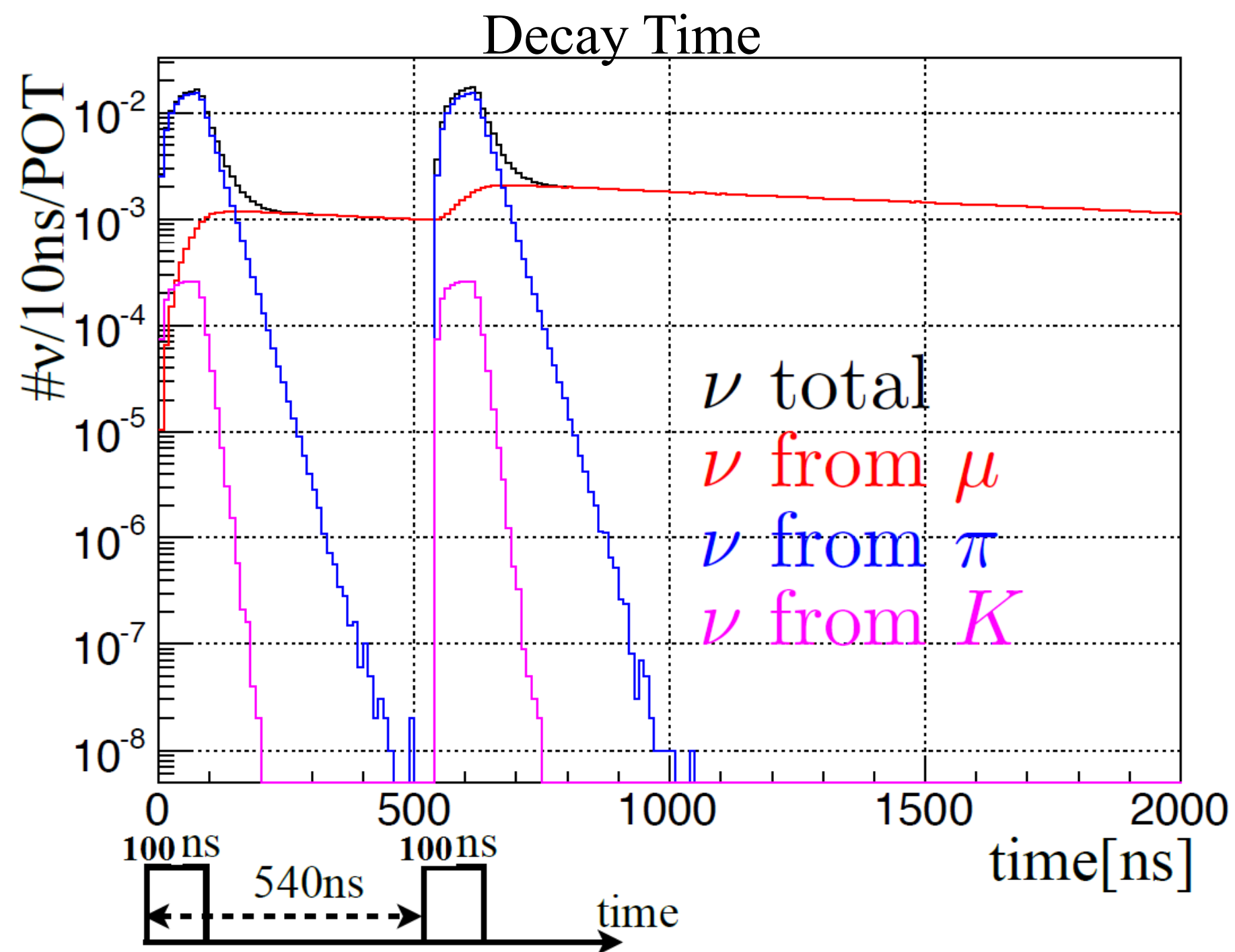
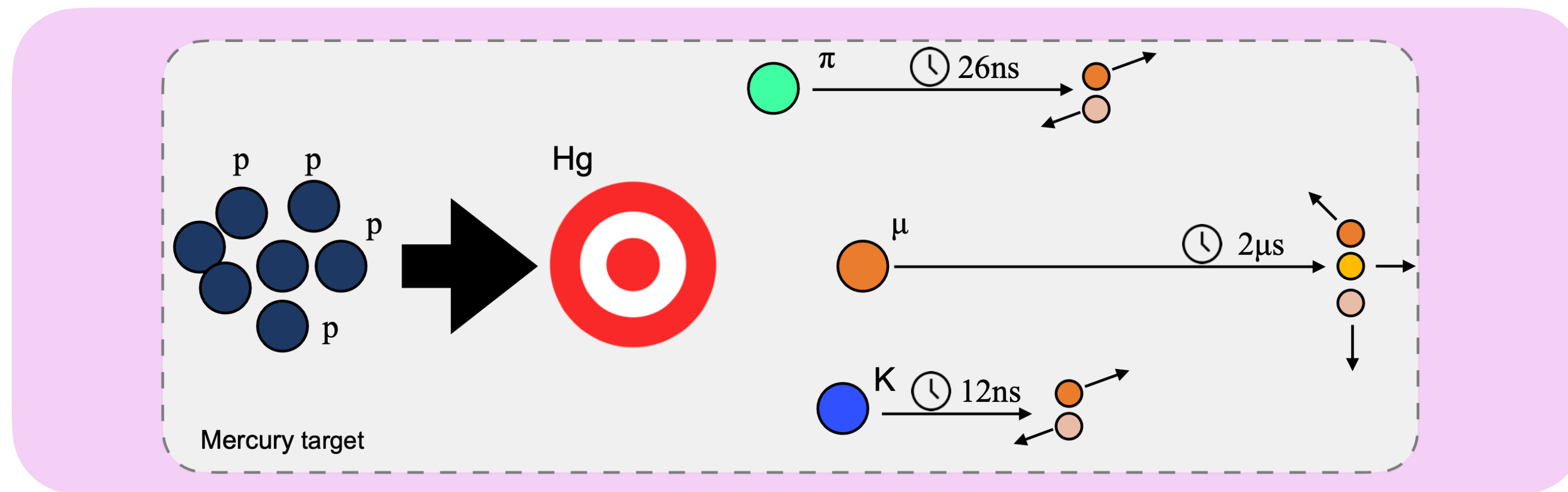
University of Sussex

# JSNS<sup>2</sup> Data Taking

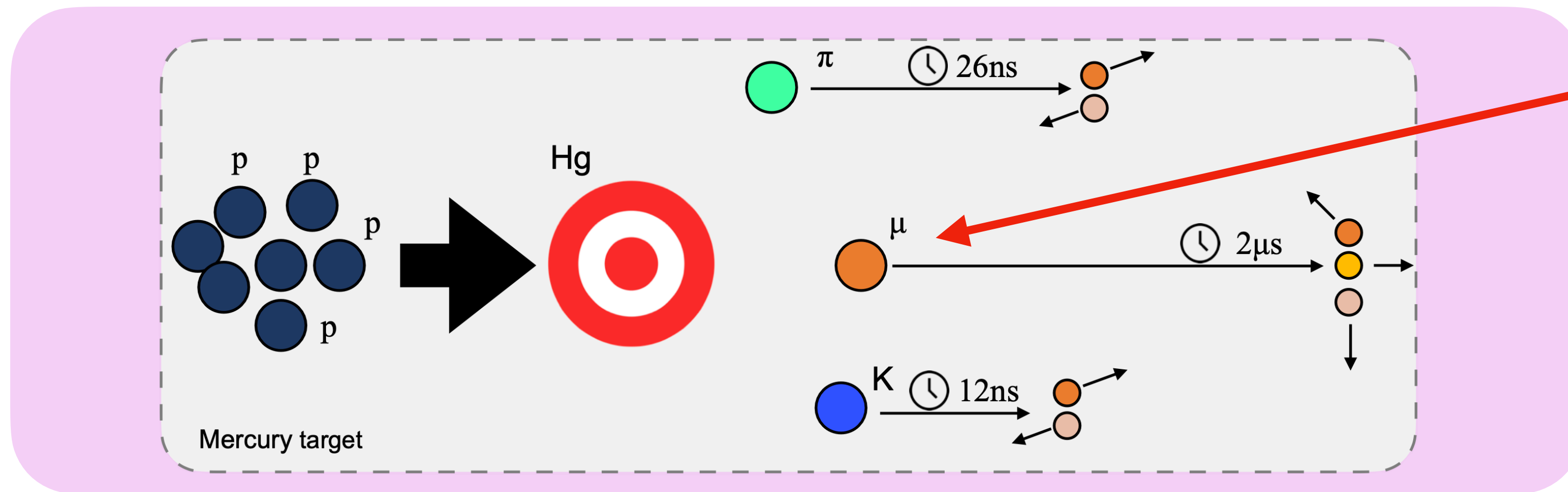
- Data taking began in spring 2020. First physics run was in 2021.
- Recently finished our 2024 data taking
- 1MW beam power design goal
- **Consistent 950kW beam delivery!**
- Data taking will resume in December



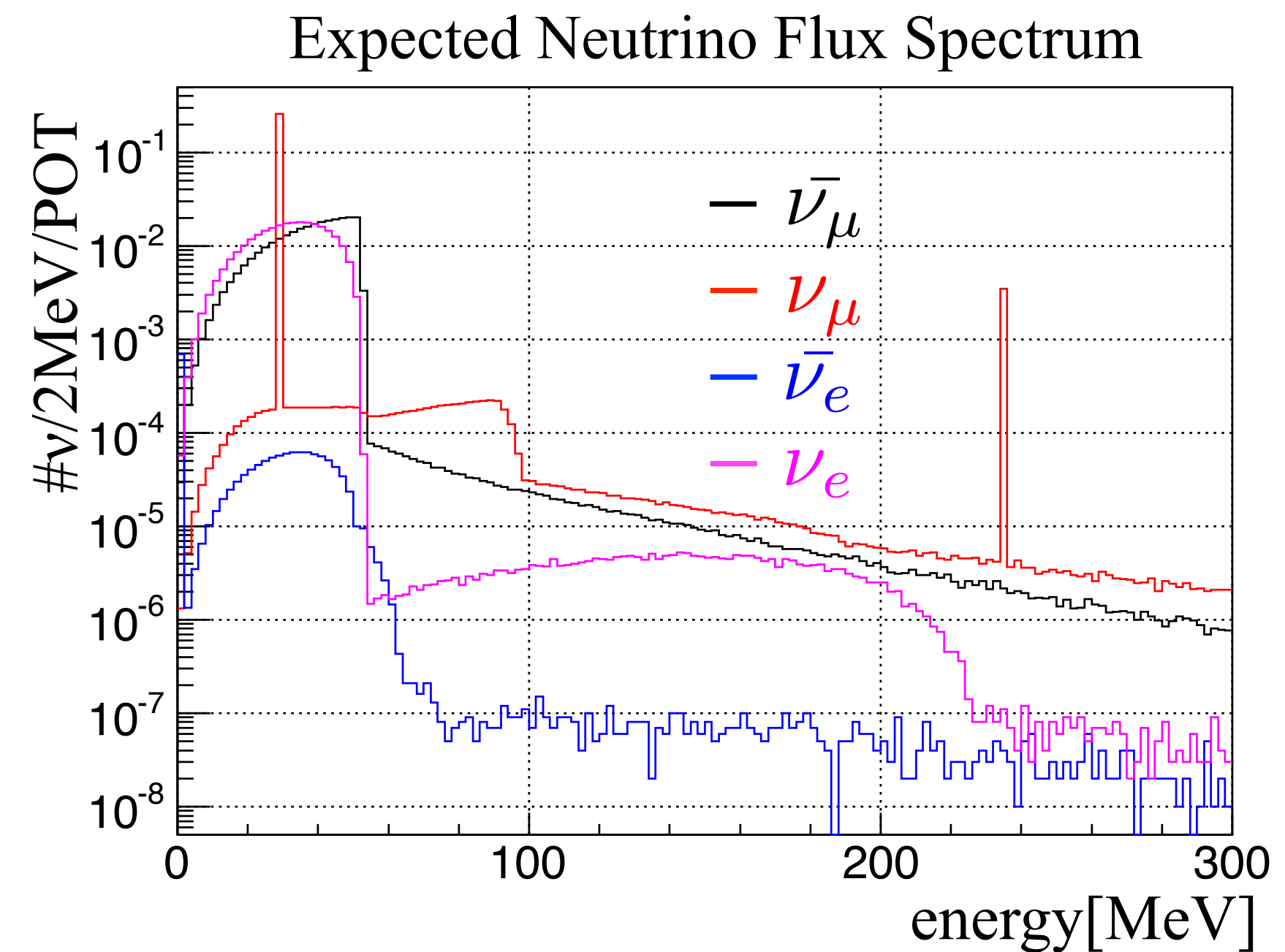
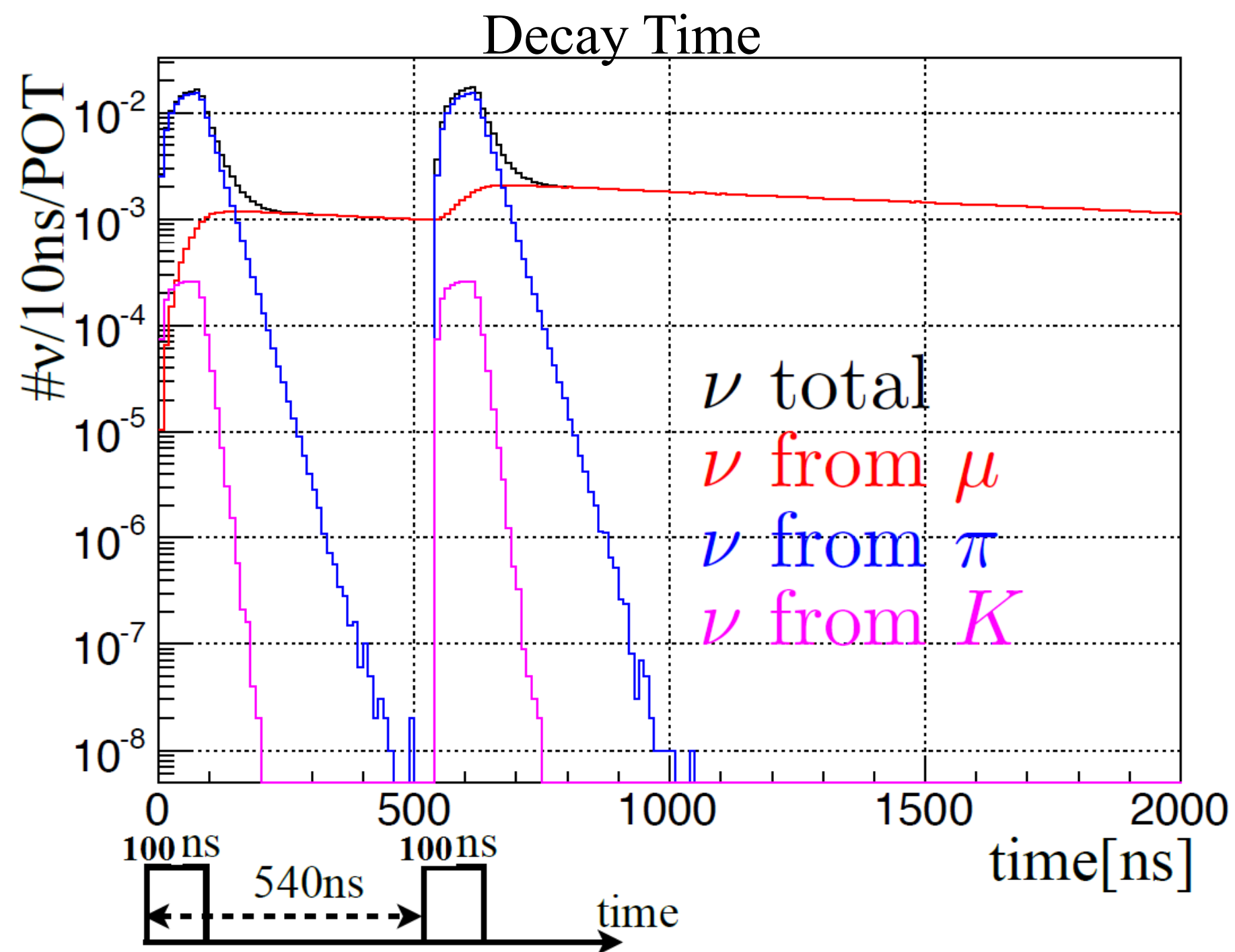
# Neutrino Signals



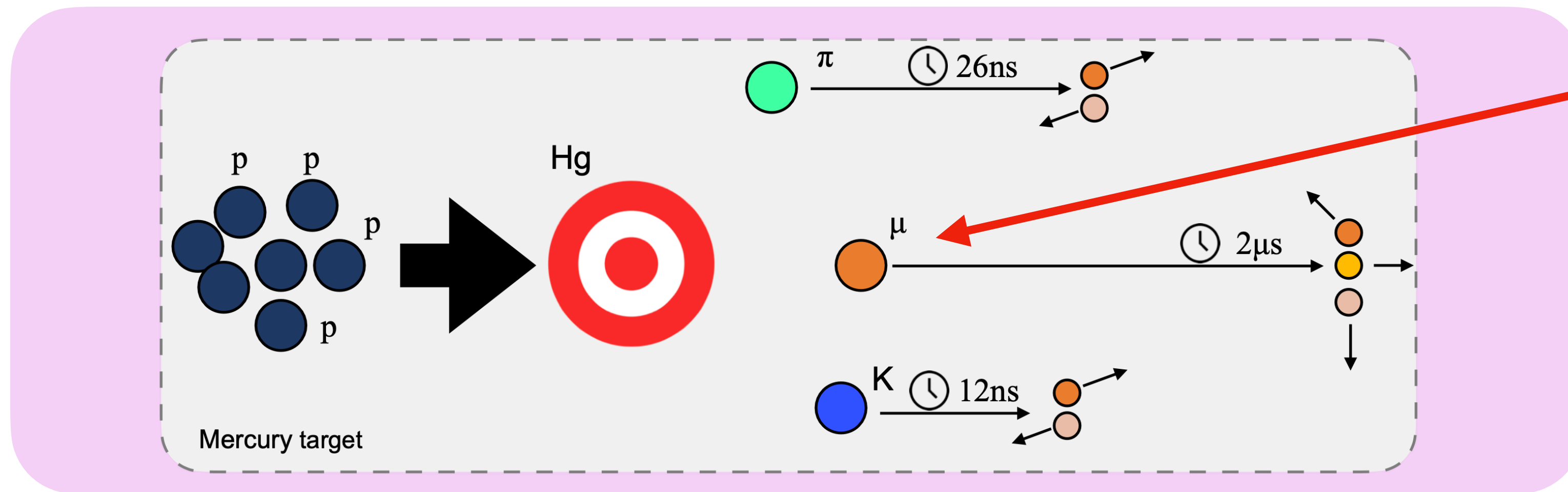
# Neutrino Signals



**Look for short baseline  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations with neutrinos from muon decay**

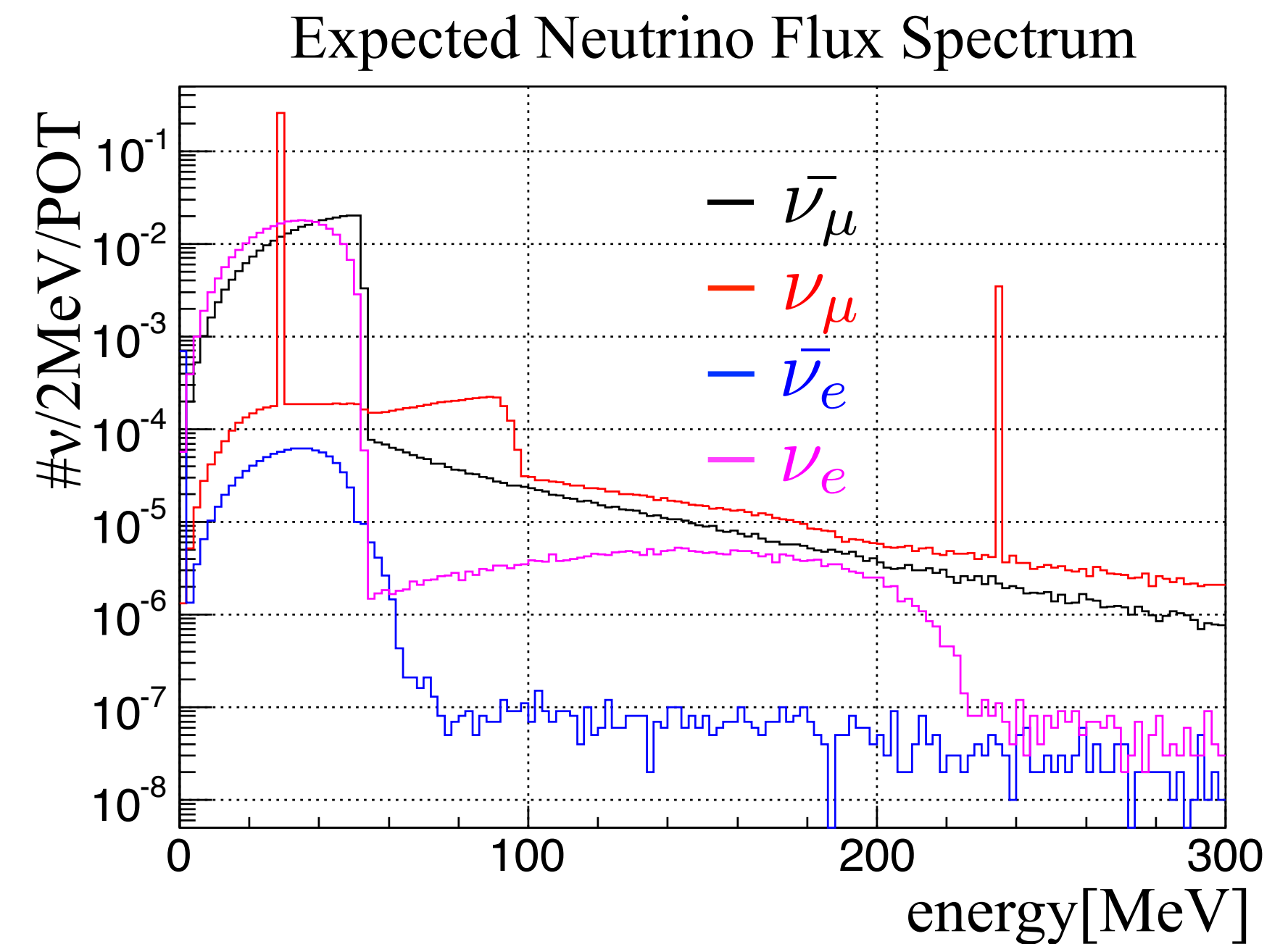
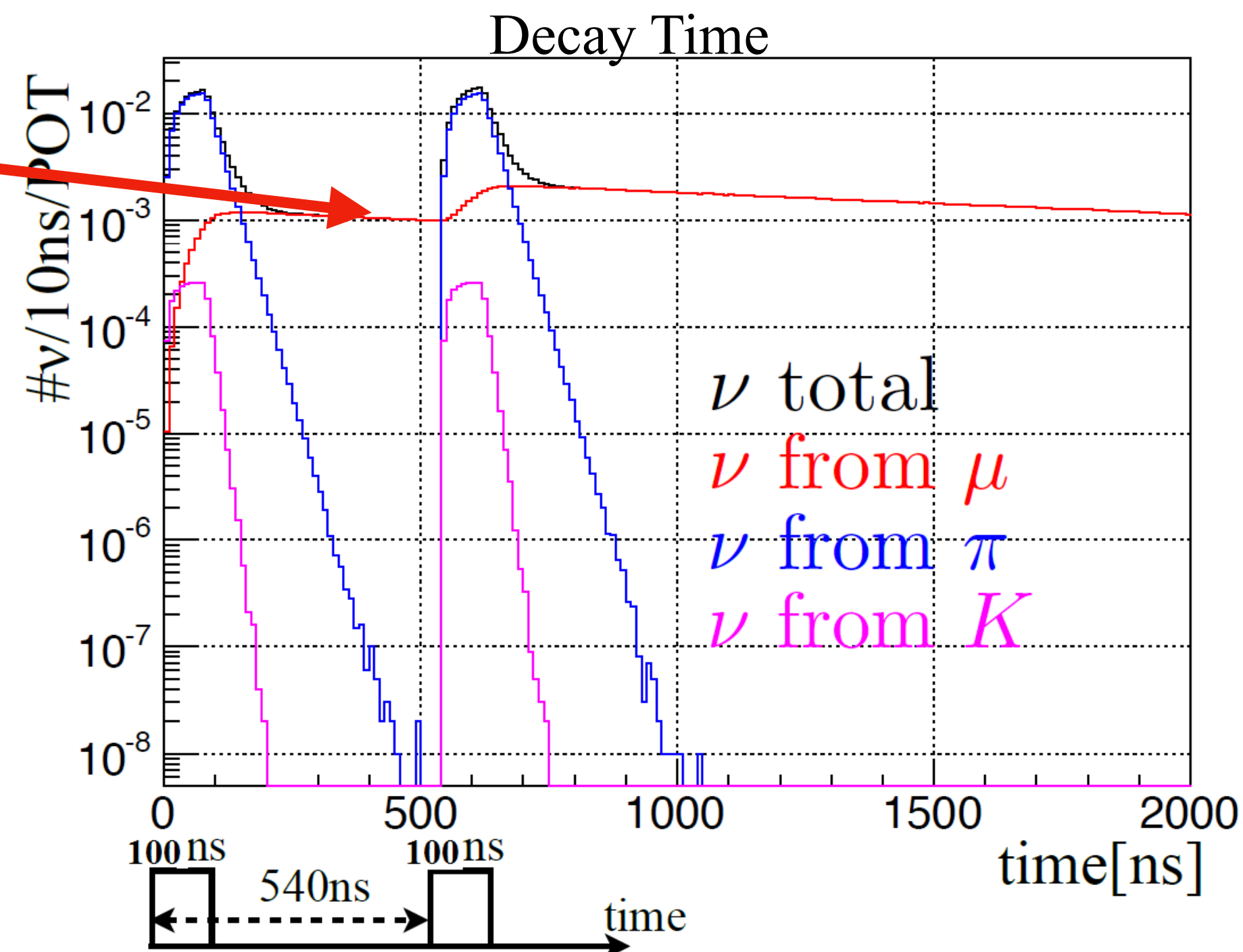


# Neutrino Signals

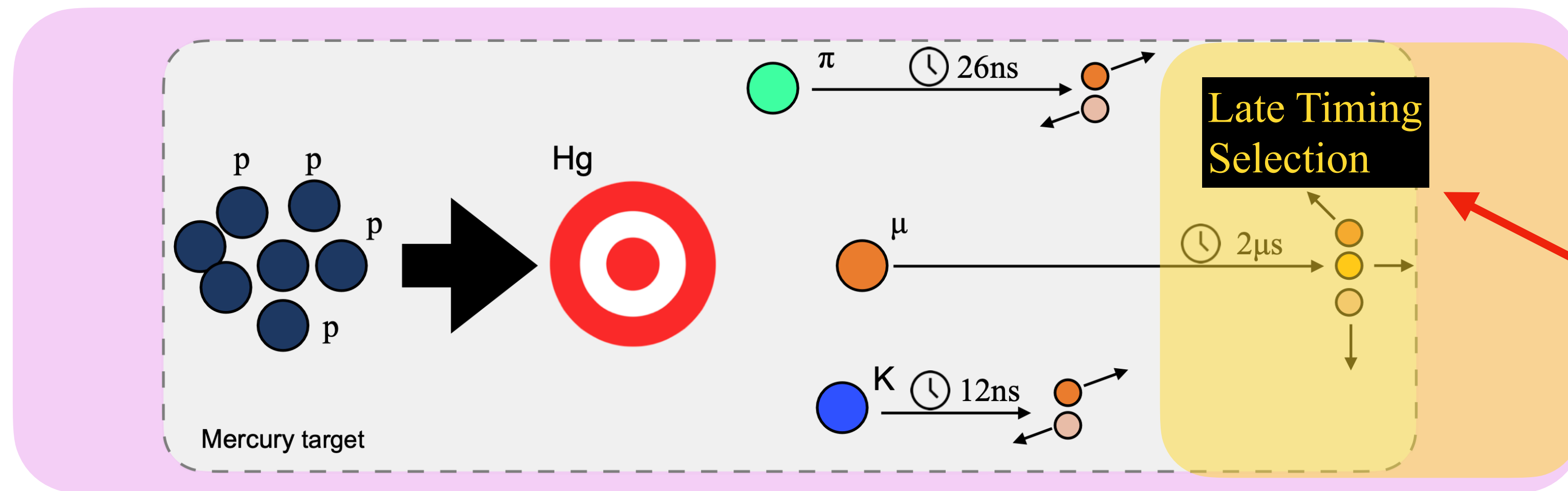


**Look for short baseline  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations with neutrinos from muon decay**

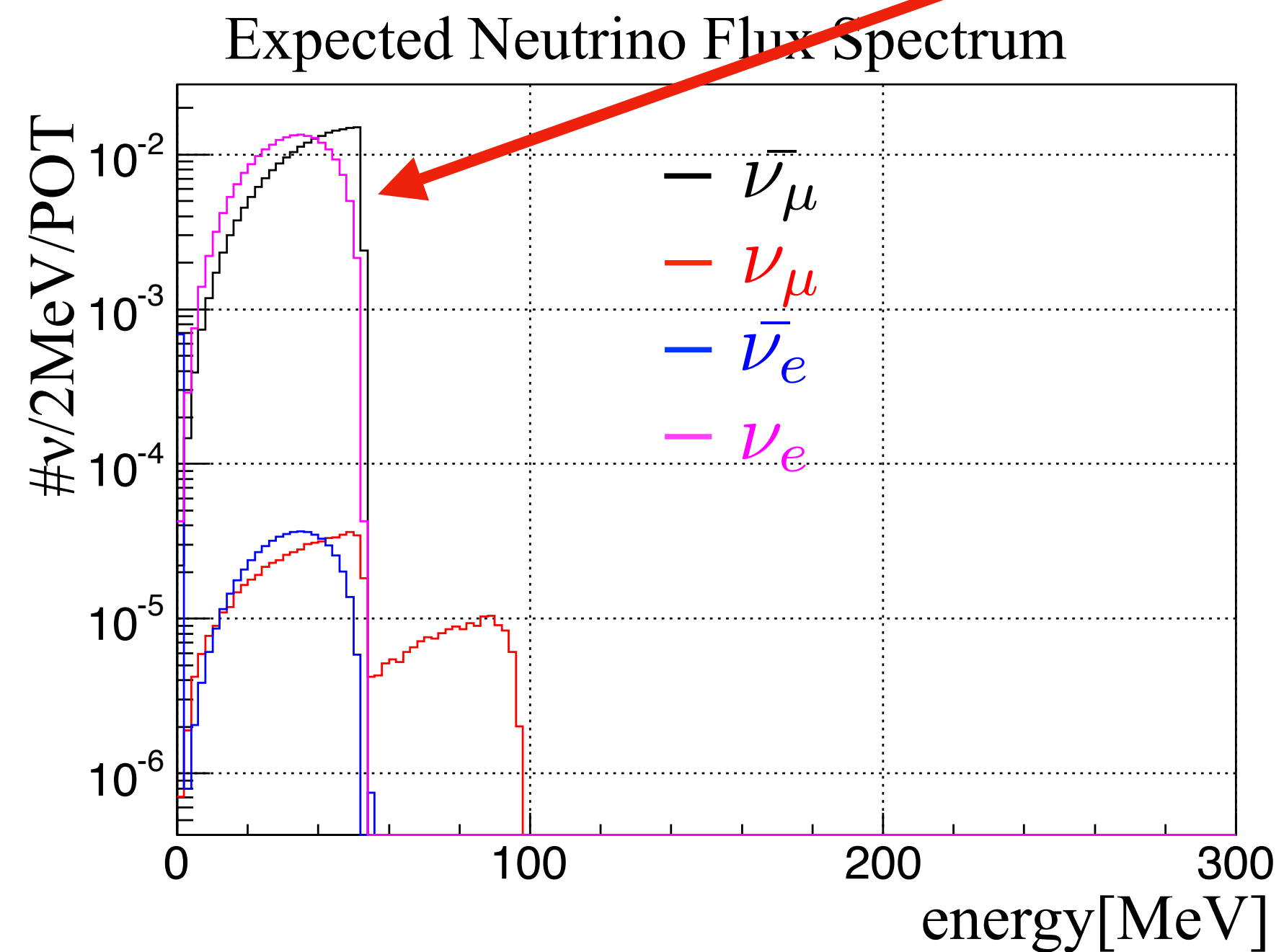
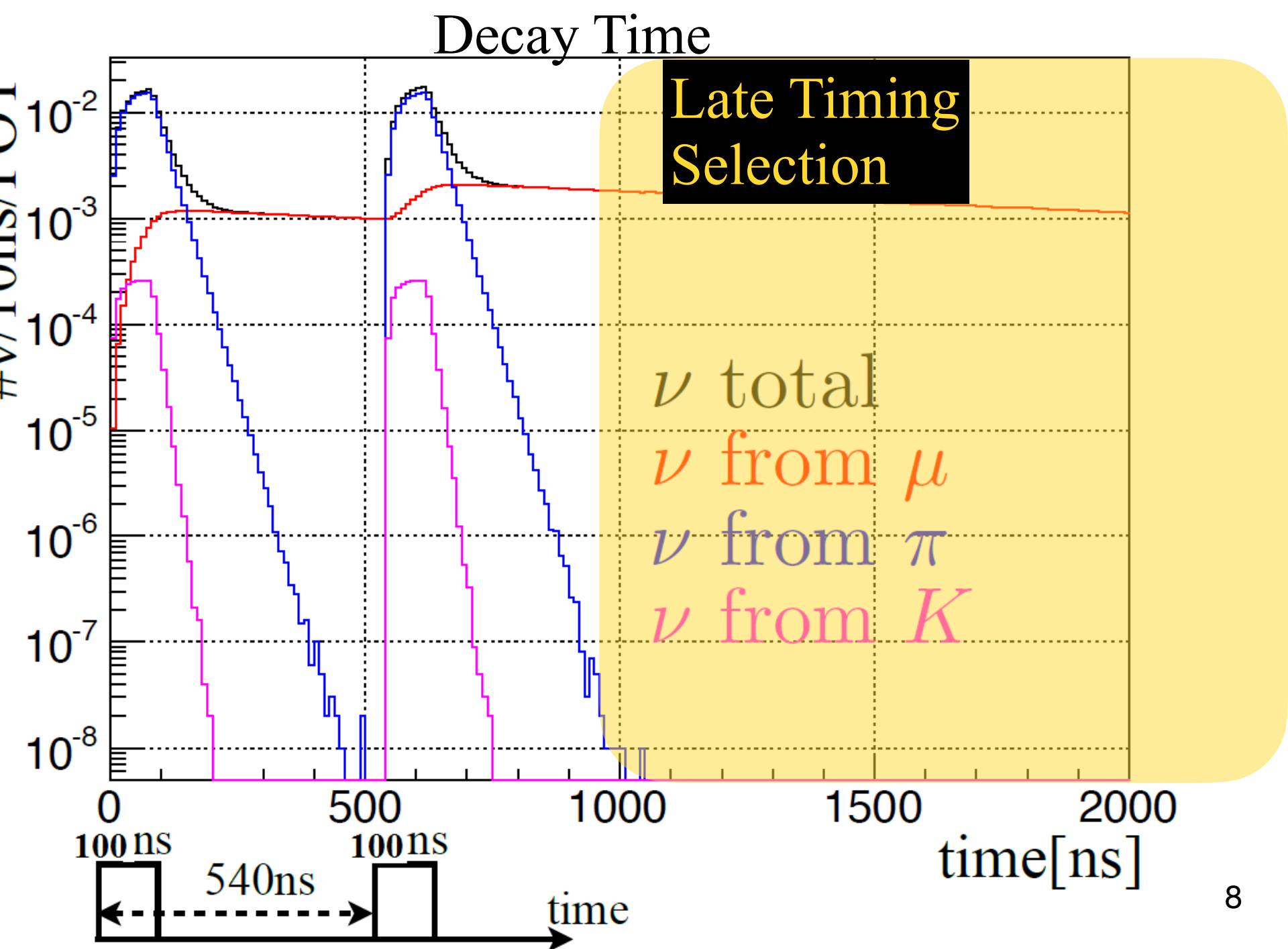
**Muon's long lifetime separates its neutrinos from pion/kaon neutrinos**



# Neutrino Signals



Selecting for late times only removes decay-in-flight & pion/kaon neutrinos

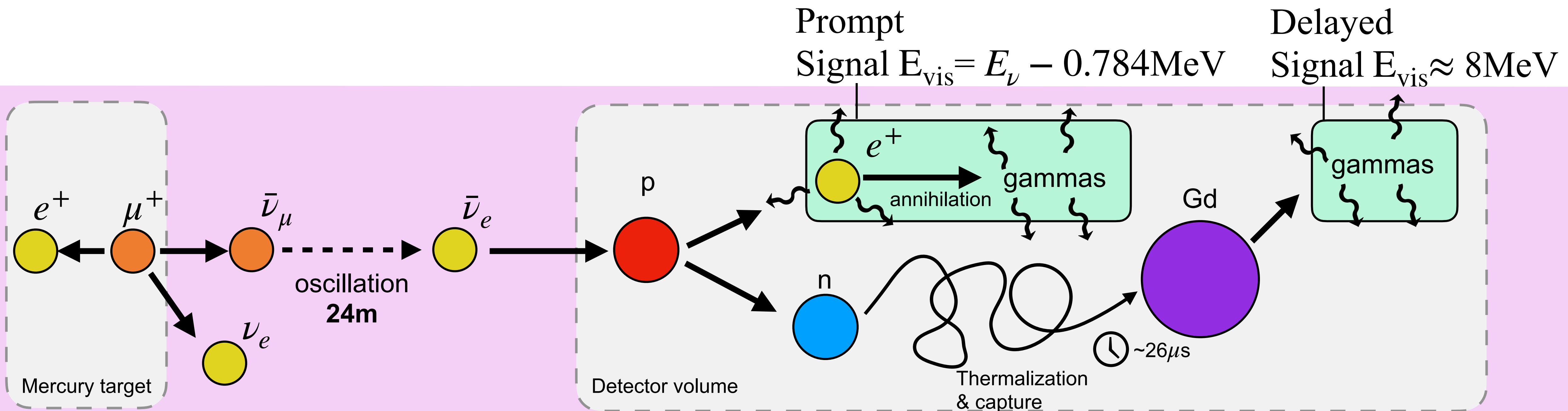




# Sterile Neutrino Signal

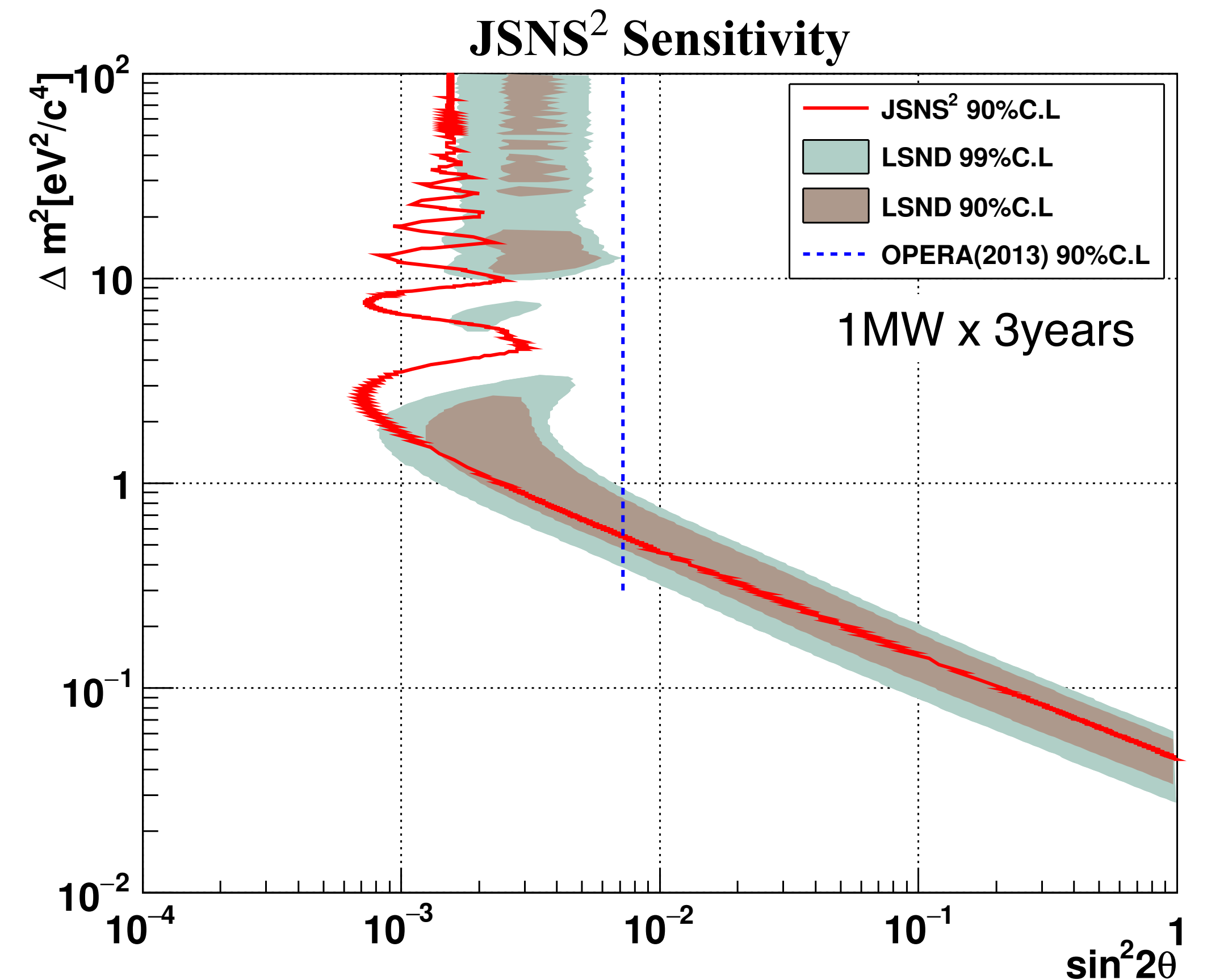
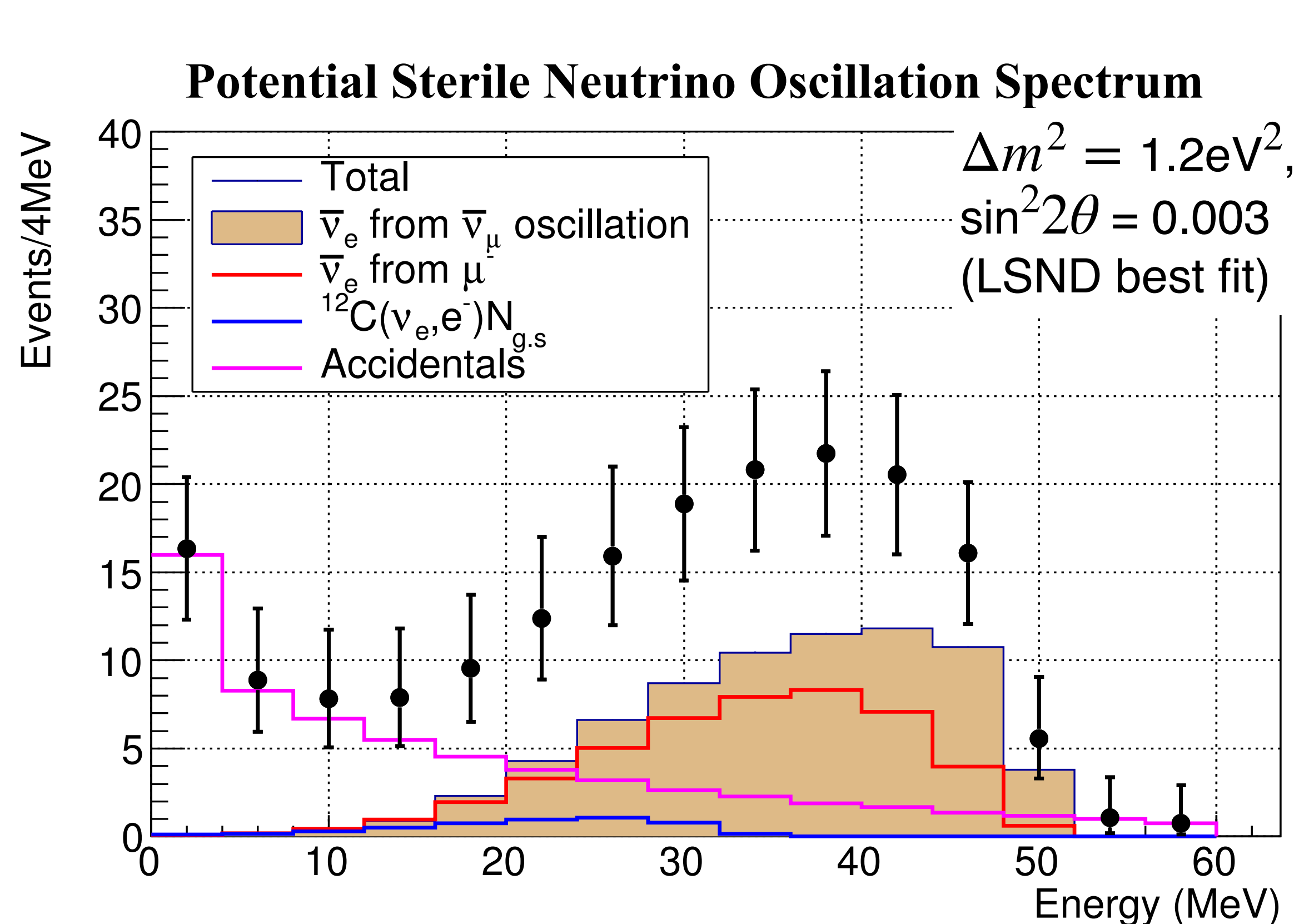
- Detect  $\bar{\nu}_e$  via inverse beta decay (IBD) interactions.
- Coincidence selection...

	Timing	Energy
<b>Prompt</b>	$1.5 < t < 10\mu\text{s}$	$20 < E < 60 \text{ MeV}$
<b>Delayed</b>	$\Delta t < 100\mu\text{s}$	$7 < E < 12 \text{ MeV}$



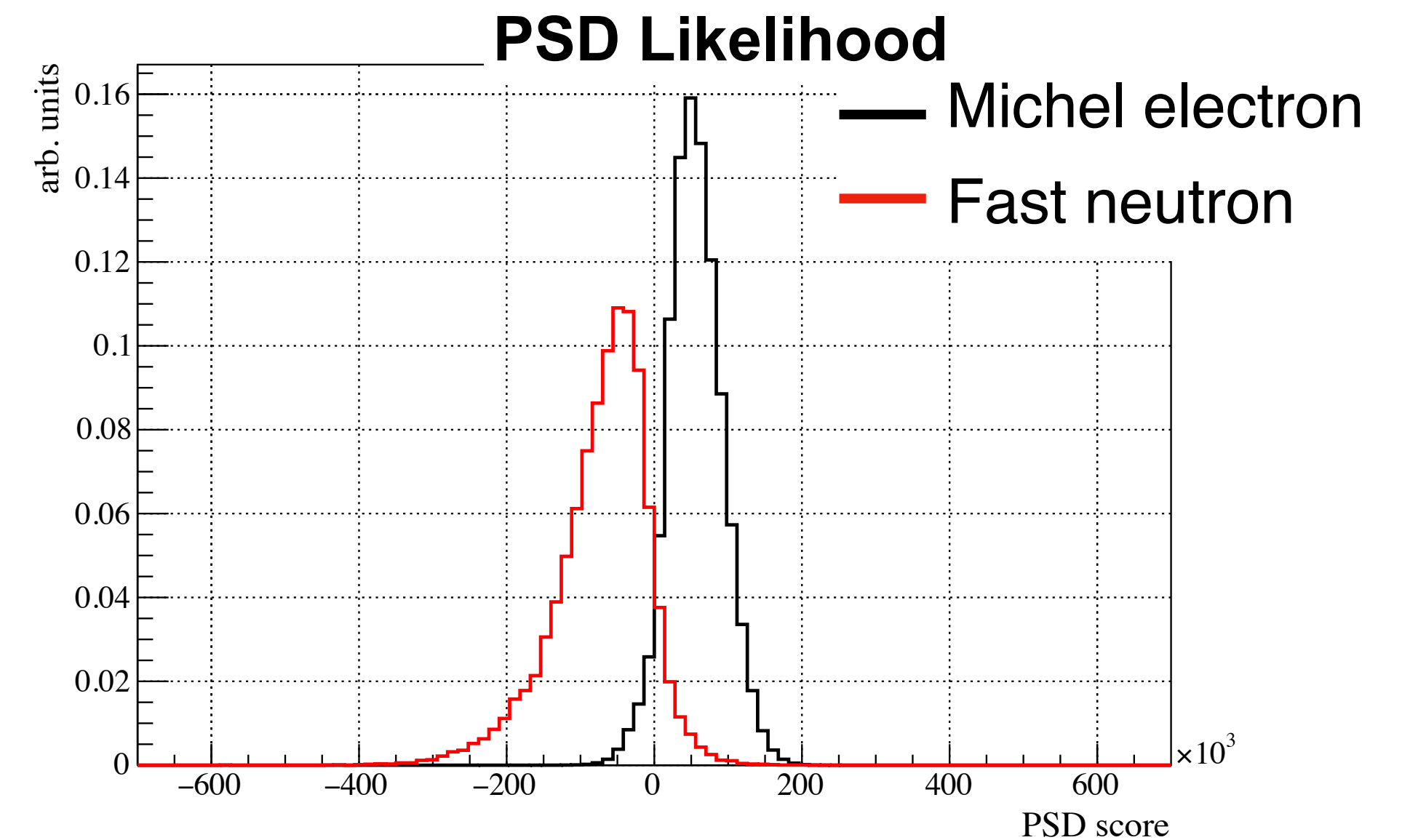
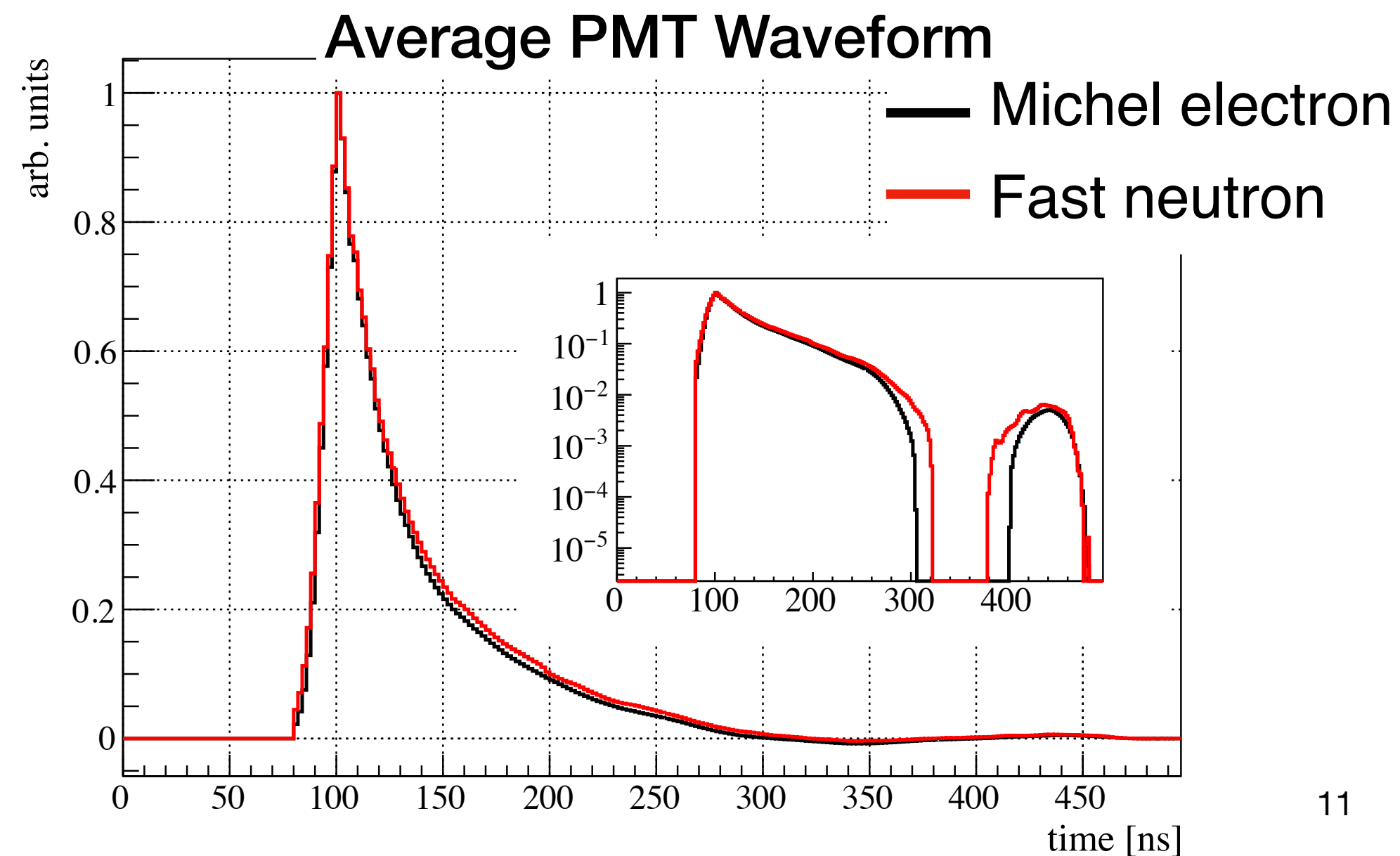
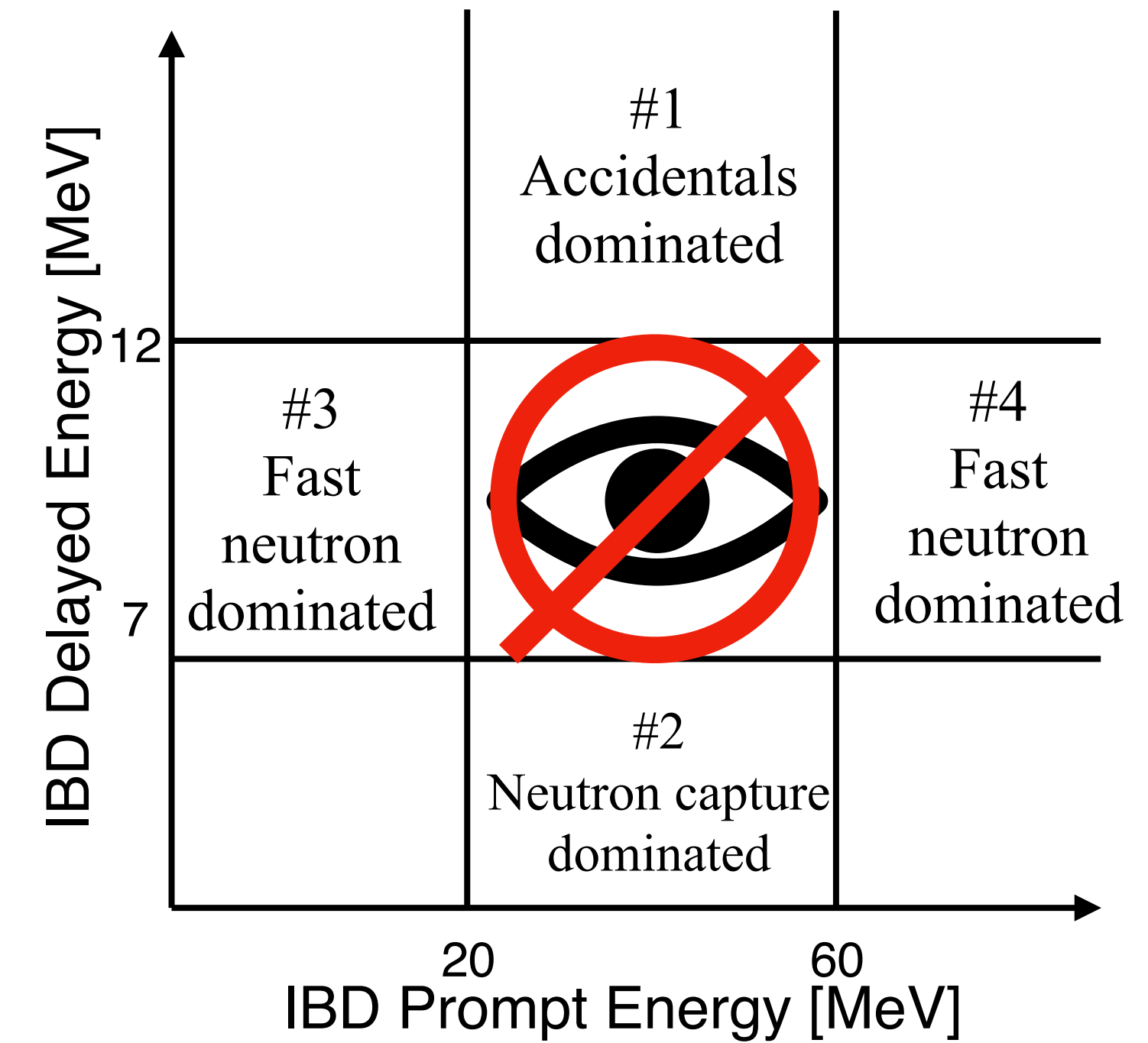
# Expected Sterile Neutrino Sensitivity

- With 1MW x 3years observed POT (currently acquired 42.5%) we'll cover the majority of the LSND observation region at 90%CL

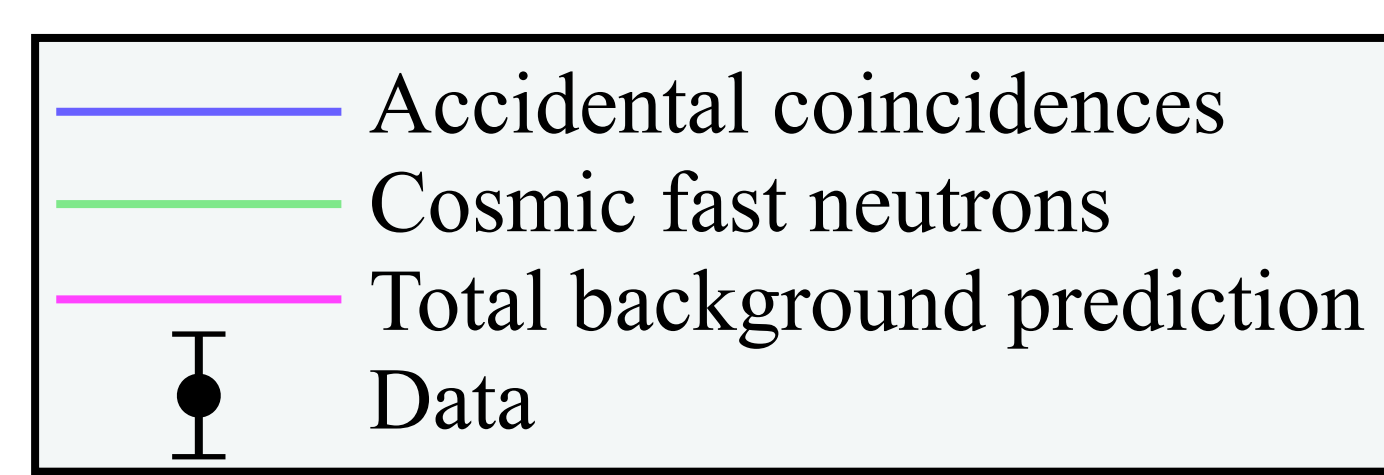
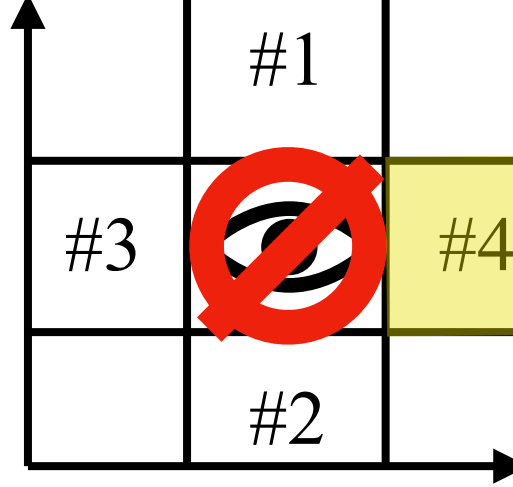


# Sterile Neutrino Search Status

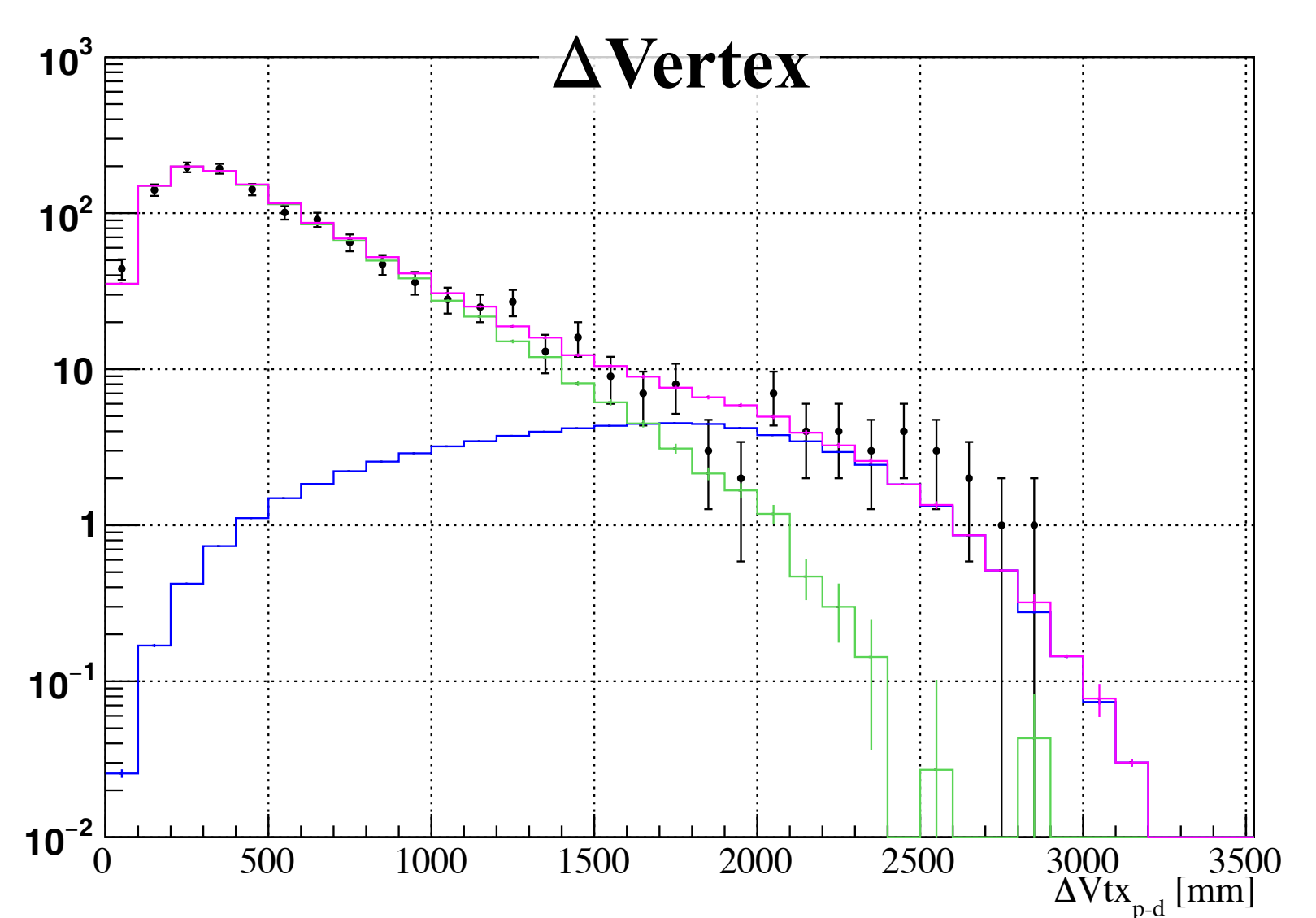
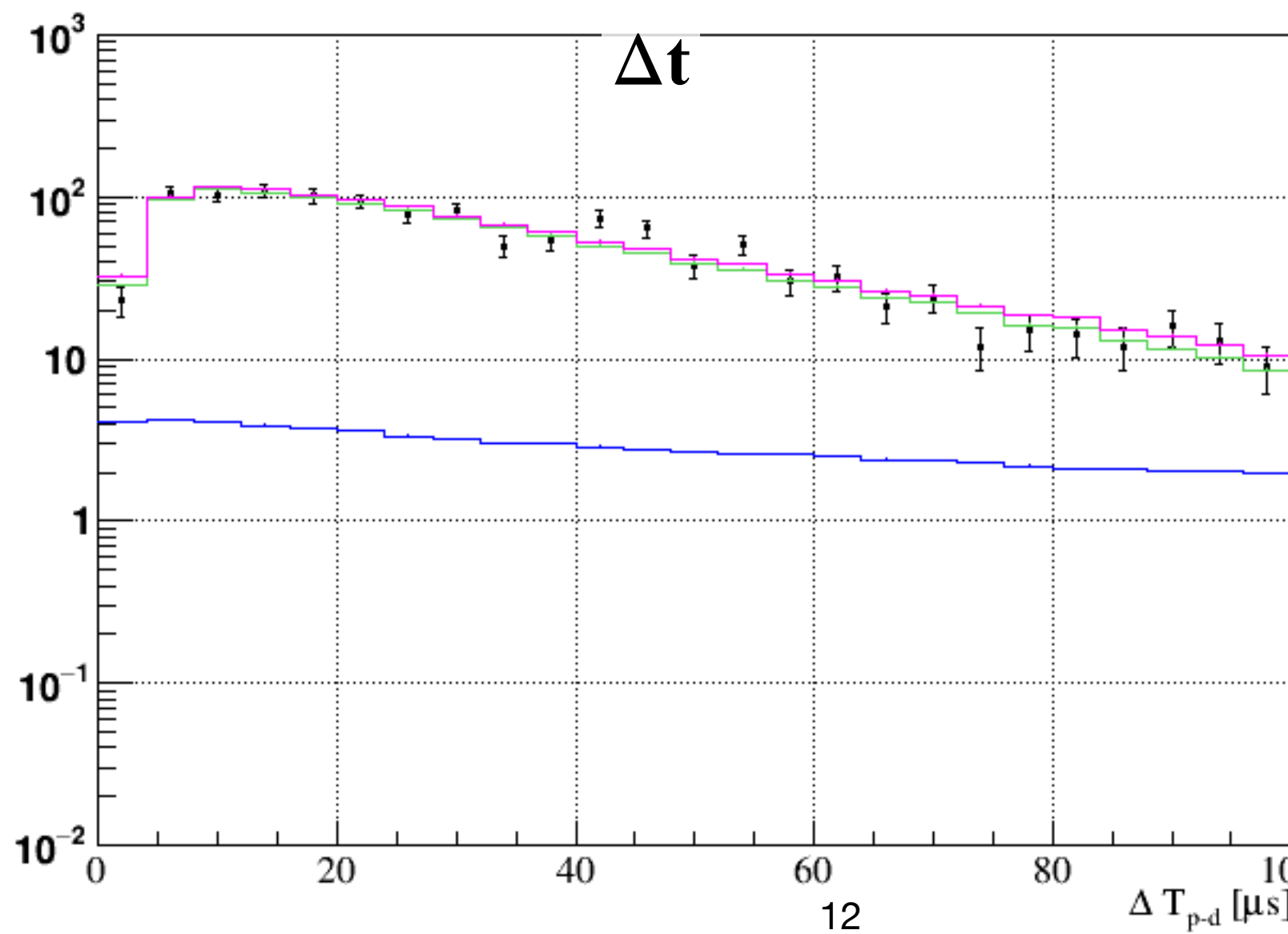
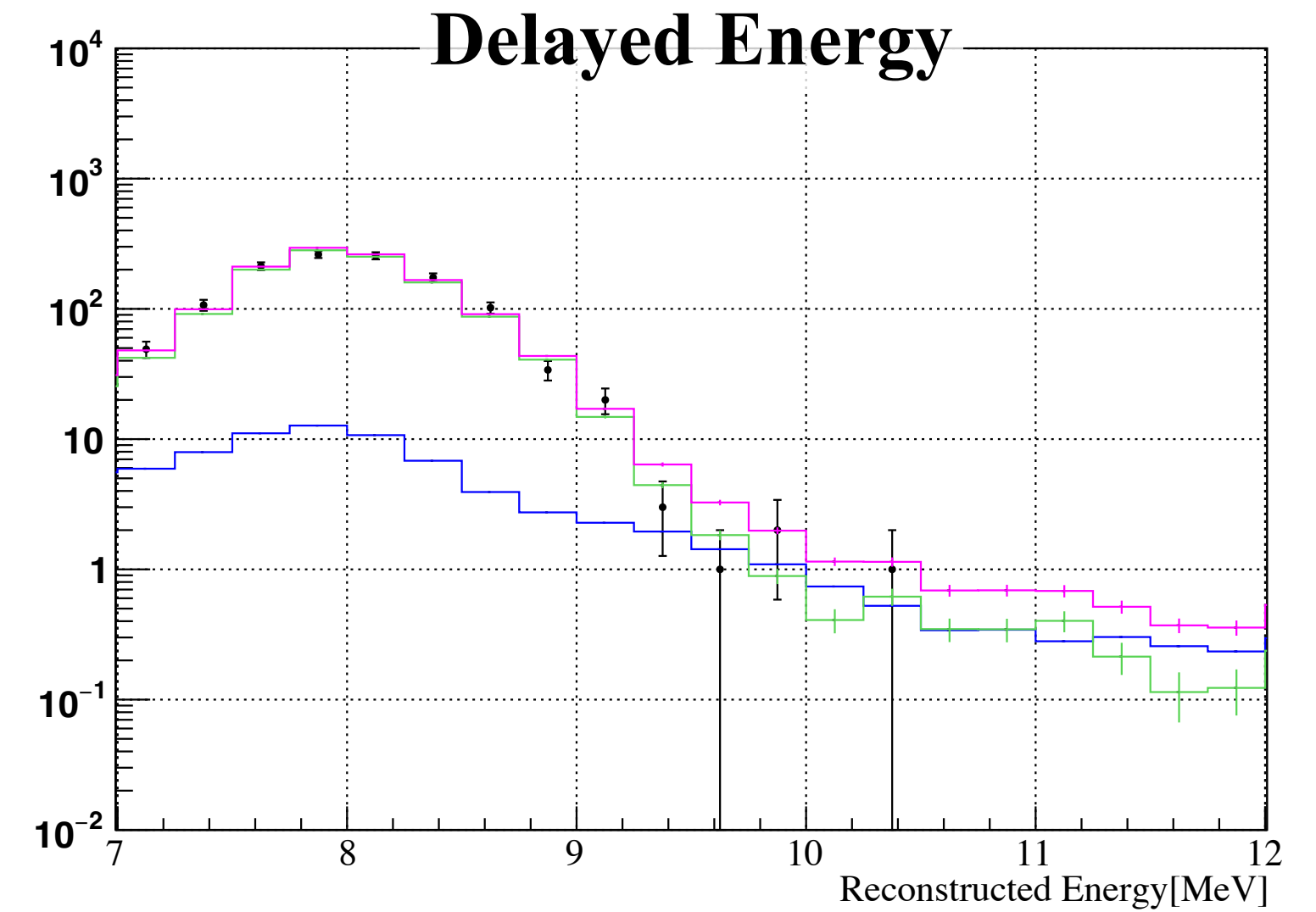
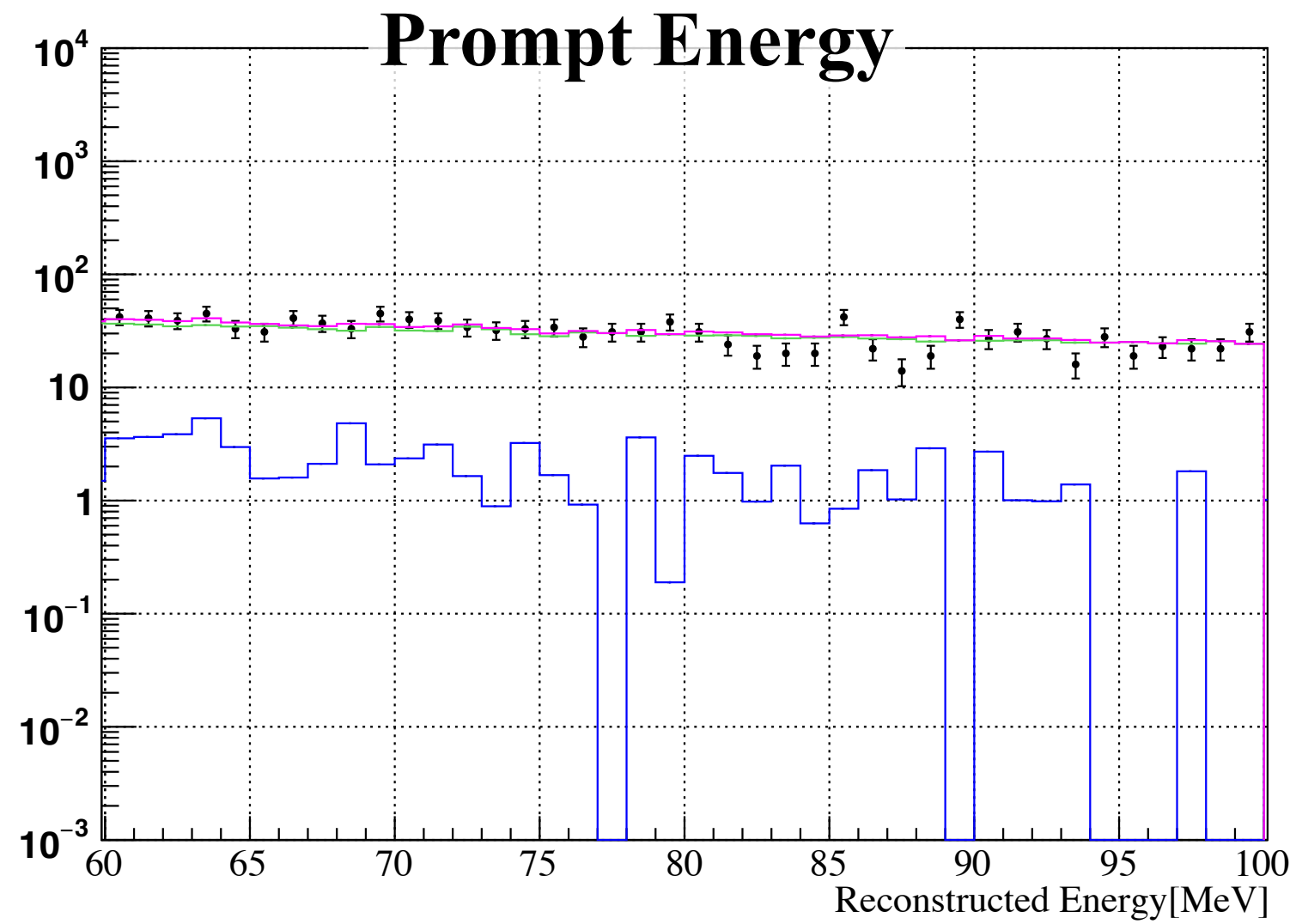
- We define a blinded signal energy window
- Sideband regions used to develop the background estimates and compare to observation
- Cosmic fast neutrons are our most pernicious background.
  - We can reject them using pulse-shape discrimination (PSD).
  - Achieves 95% rejection with only 10% signal loss (arXiv:2404.03679)



# Side-band #4, Prompt Energy > 60 MeV Delayed Energy 7-12 MeV



- All background estimates are data driven
- Accidental rates are estimated via “spill-shift” method
  - Prompt-delayed pairs from different beam spills are purposefully mis-paired.
- Cosmic fast neutron rate is estimated by looking in late time window (>1ms after beam)
- **Observed data is consistent with accidentals+neutron background estimate**
- Other sidebands are being studied similarly

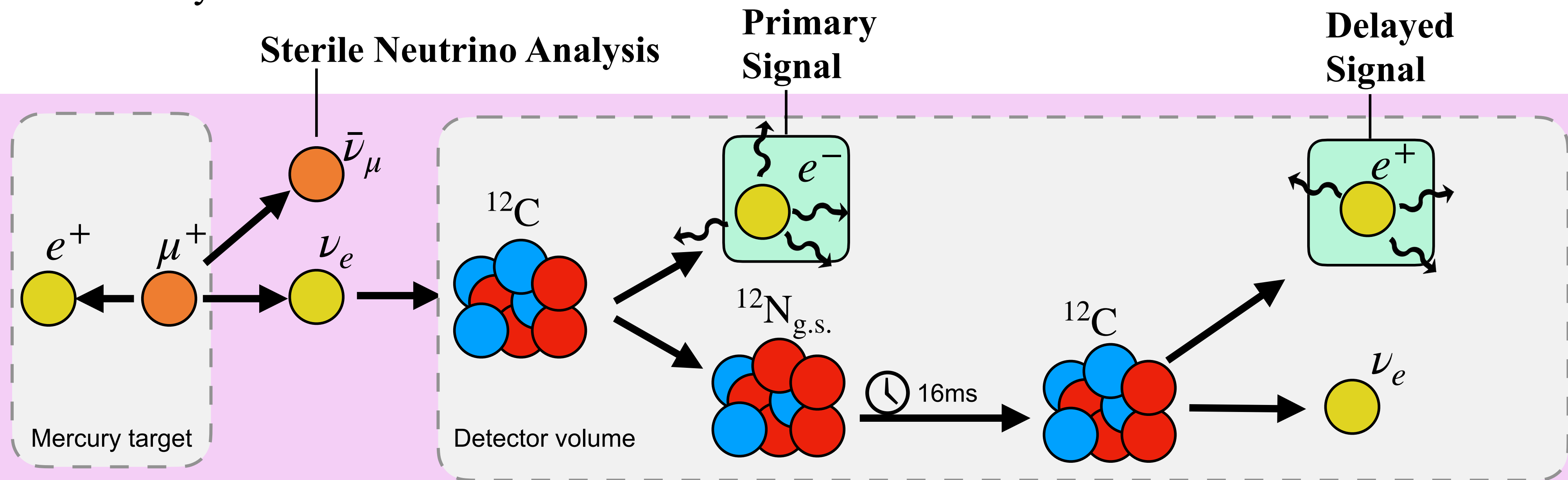


	Observed	Estimated
<b>Accidentals</b>	-	71.6±1.2
<b>Fast Neutrons</b>	-	1178±4.4
<b>Total</b>	1224±35.0	1250±4.6

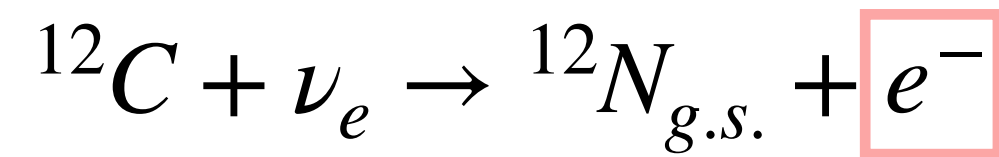
# $\bar{\nu}_\mu$ Flux Normalization from Carbon to Nitrogen ground state (CNgs) Measurement

- Production:  $\mu^+ \rightarrow e^+ + \boxed{\nu_e} + \boxed{\bar{\nu}_\mu}$ 

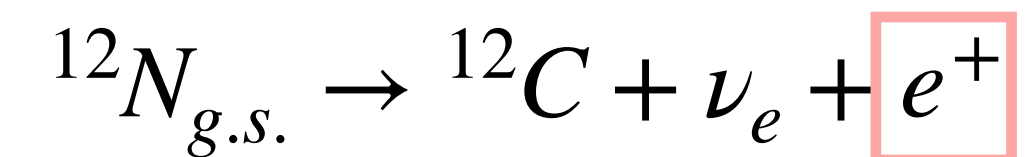
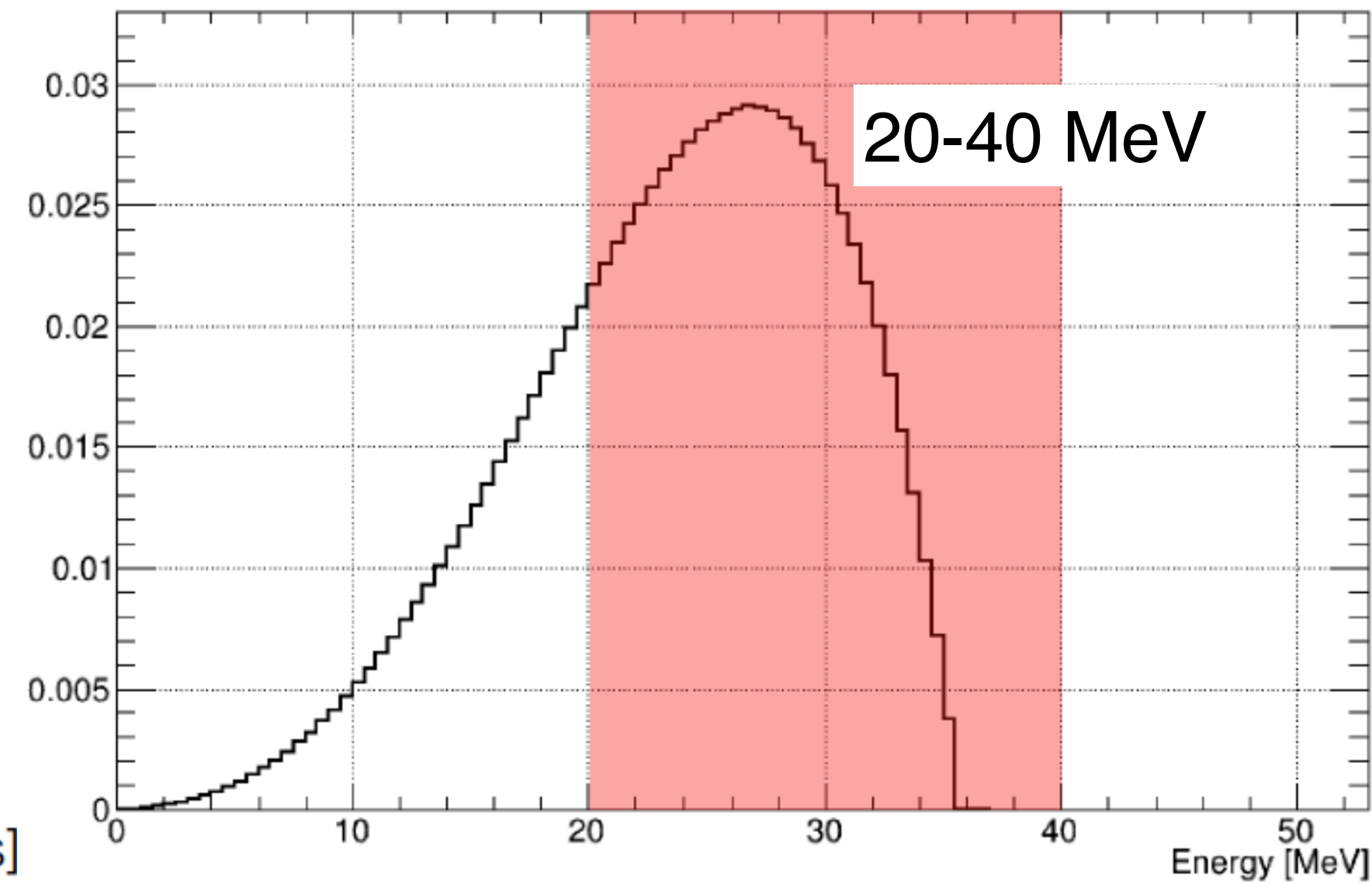
CNgs      IBD/Sterile  
 $\nu_e$      $\bar{\nu}_\mu$
- Detection:  $^{12}\text{C} + \nu_e \rightarrow ^{12}\text{N}_{g.s.} + e^-$  followed by  $^{12}\text{N}_{g.s.} \rightarrow ^{12}\text{C} + \nu_e + e^+$
- **CNgs measurement provides a data-driven estimate of the  $\bar{\nu}_\mu$  normalization**
- Avoid reliance on uncertain/complicated beam production calculations  $\rightarrow$  improves sterile neutrino sensitivity



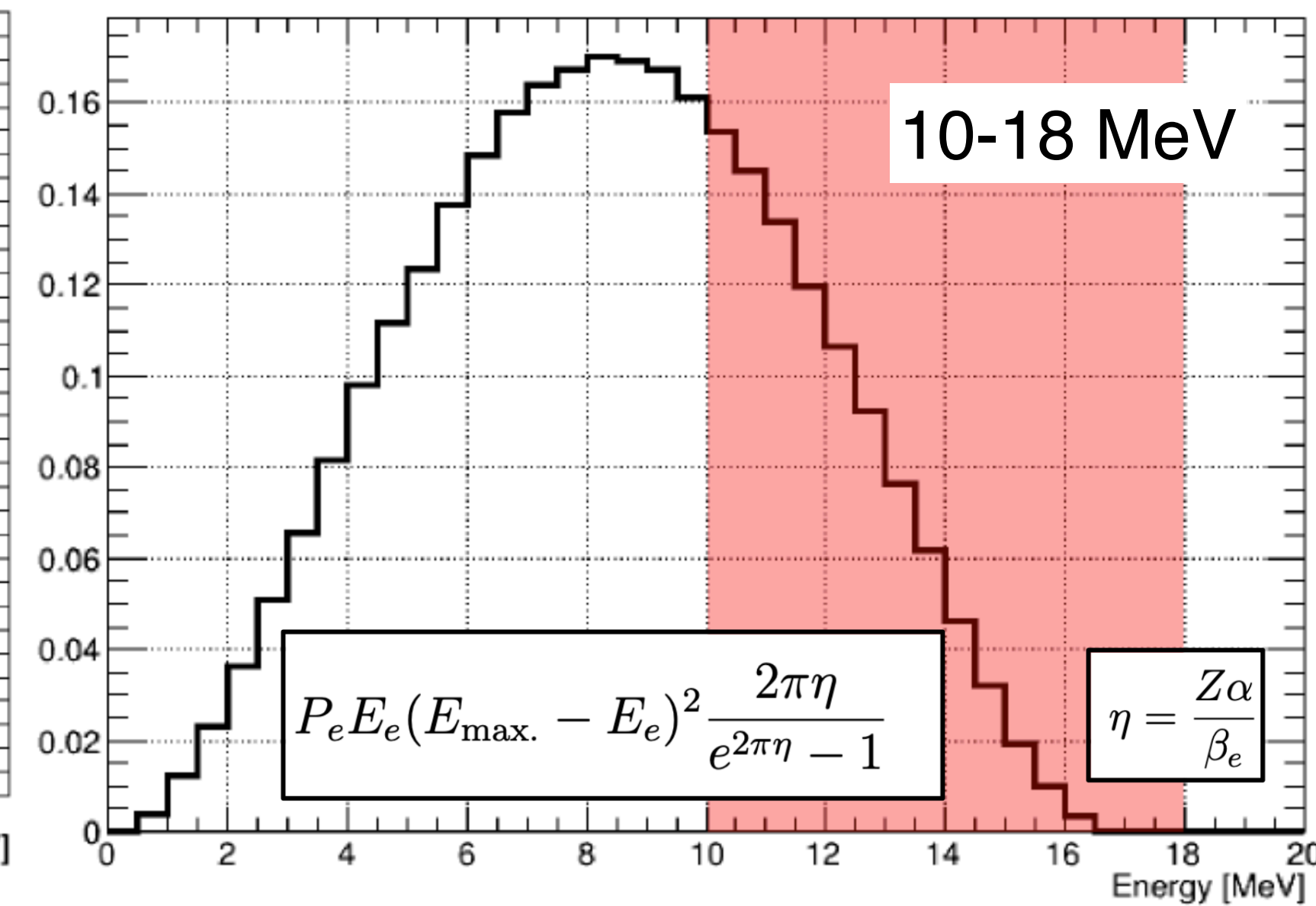
# CNgs Dataset & Event Selection



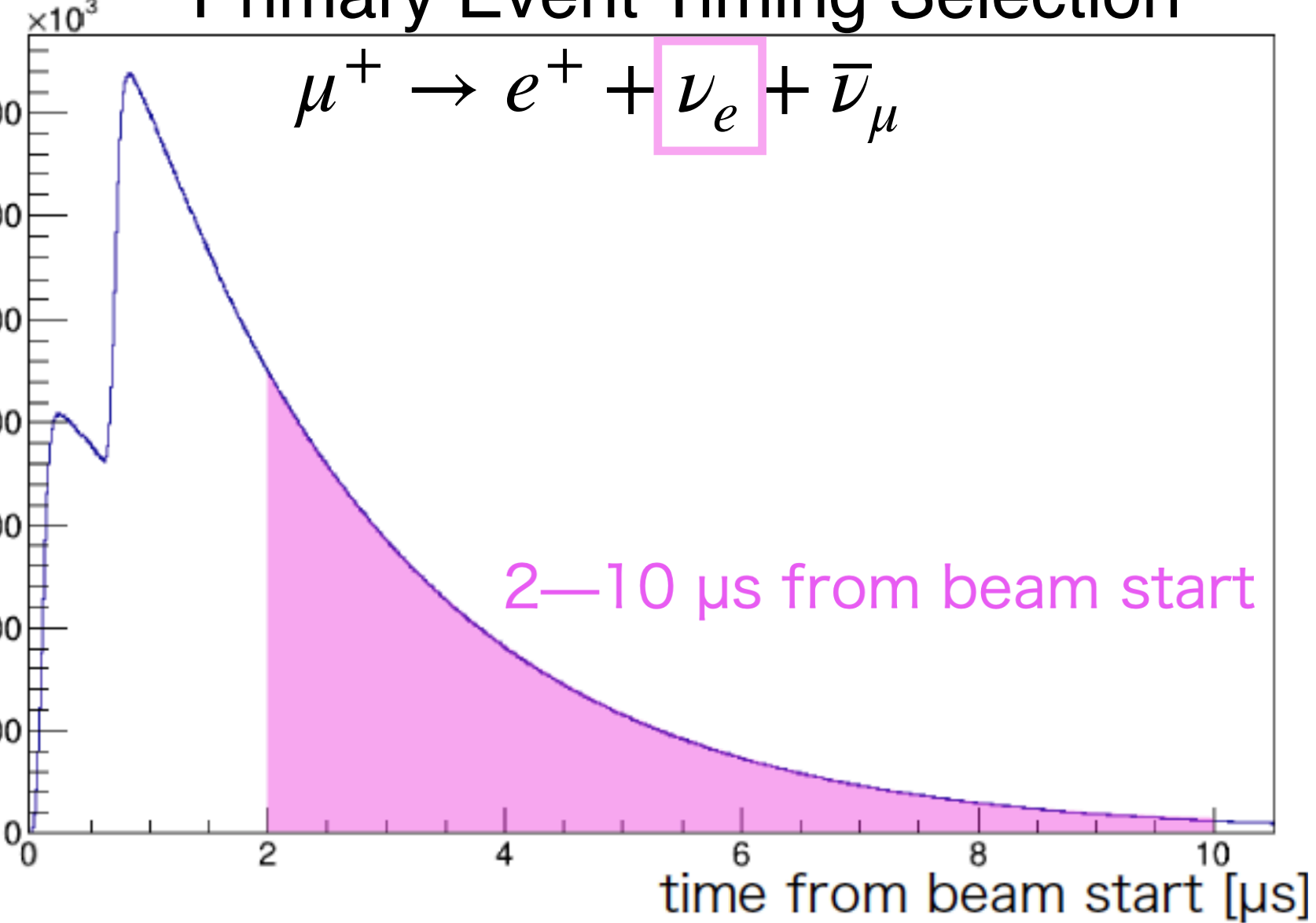
Primary Event Energy Selection



Follower Event Energy Selection



Primary Event Timing Selection



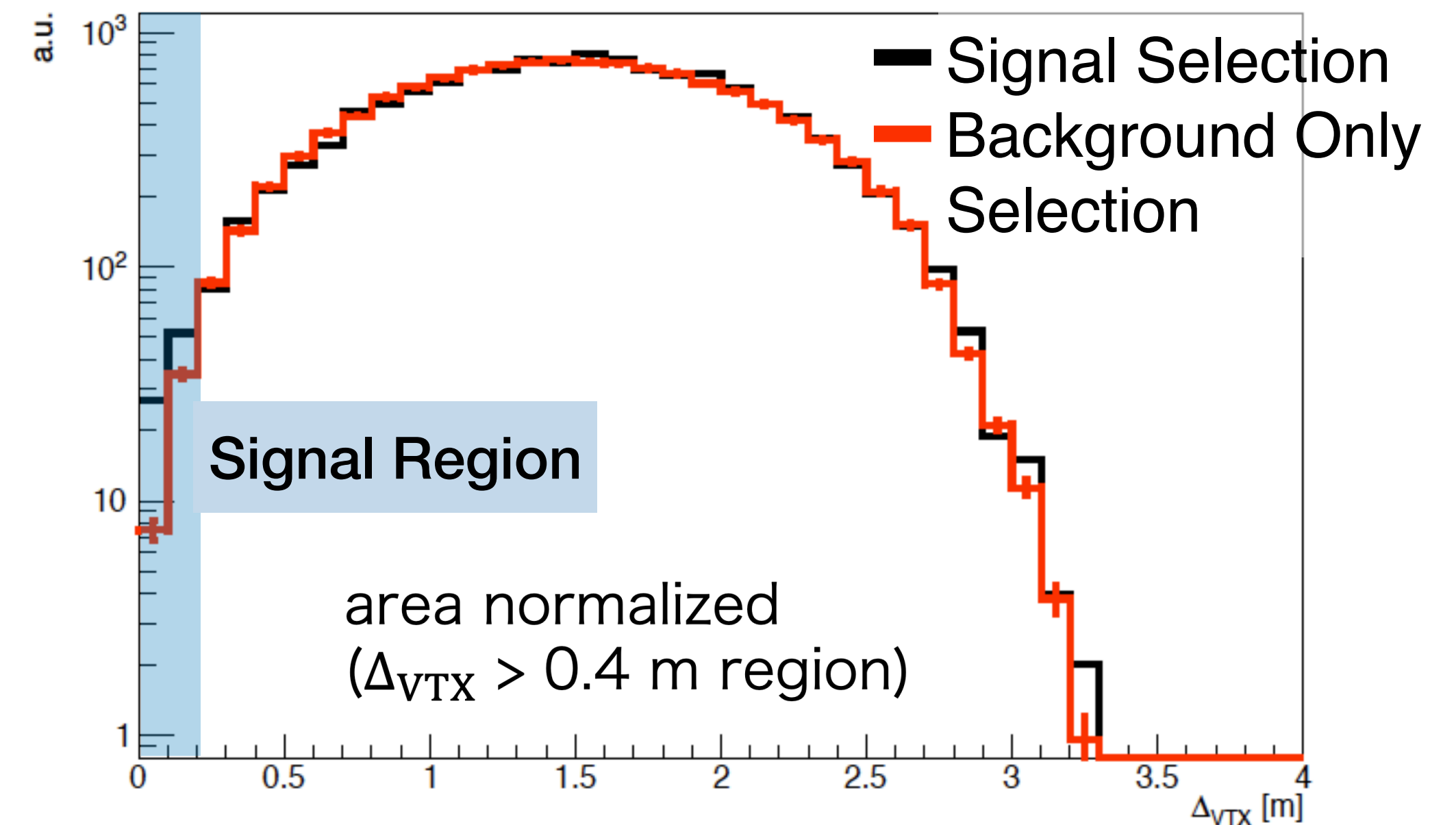
- 2021 & 2022 beam datasets are analyzed ( $2.2 \times 10^{22}$  POT)
- Event pairing selection:  
 $0.2\text{ms} < \Delta t < 12\text{ms}$  (2021),  $25\text{ms}$ (2022),  
 $\Delta\text{Vertex} < 200\text{mm}$

**79 events pass selection**

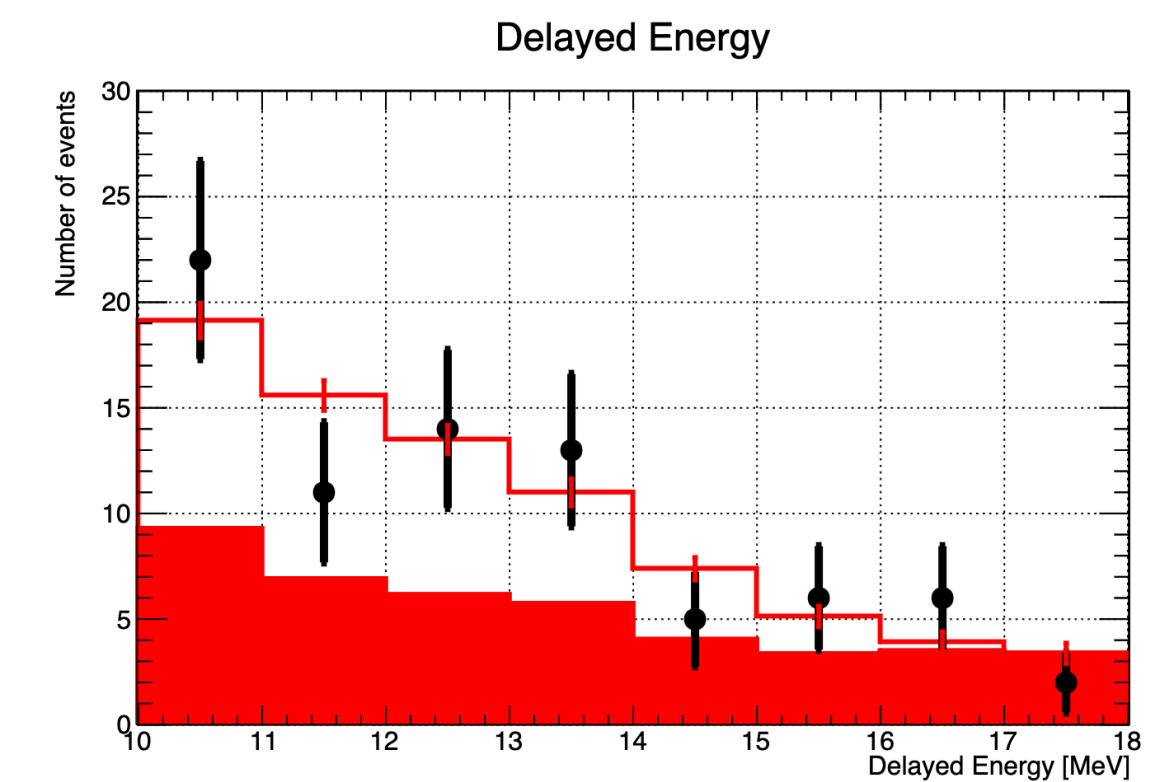
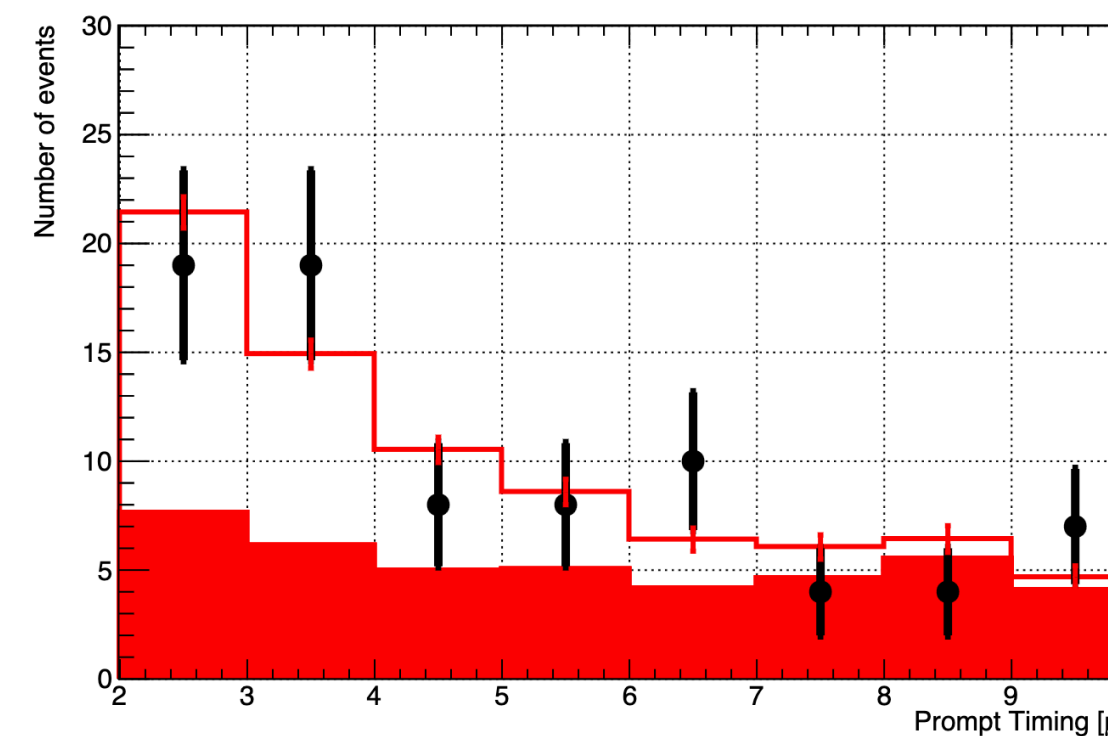
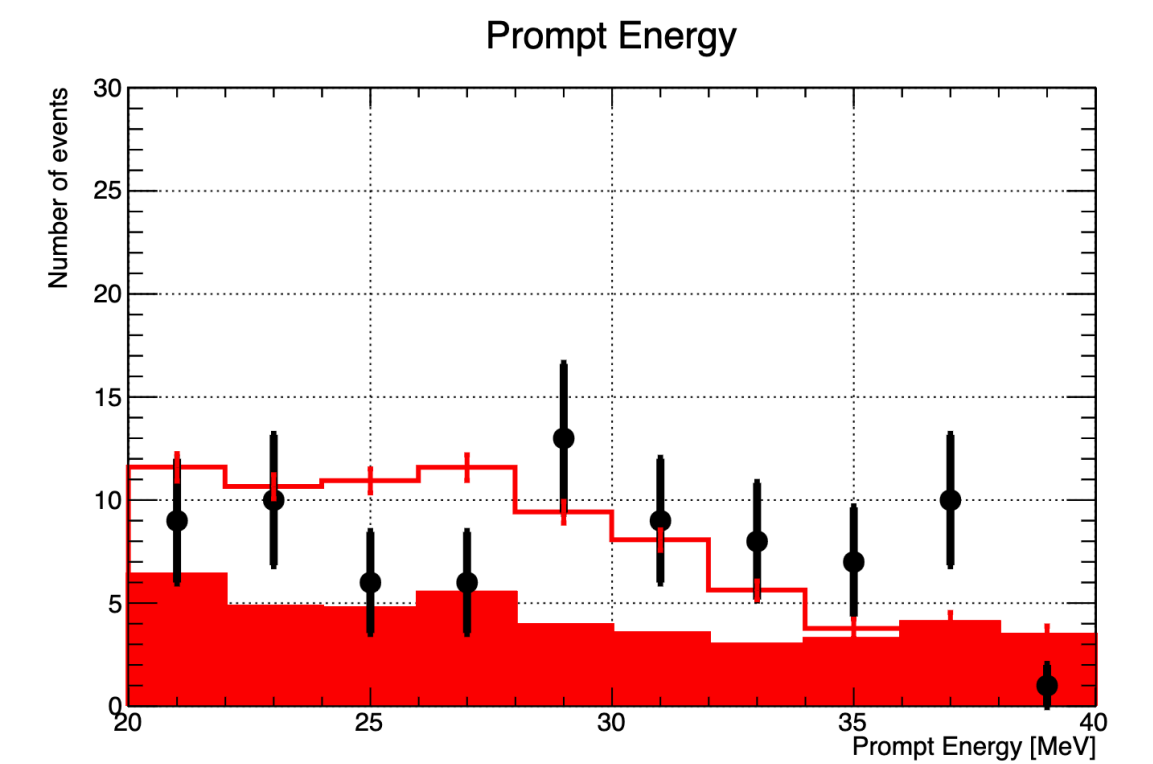
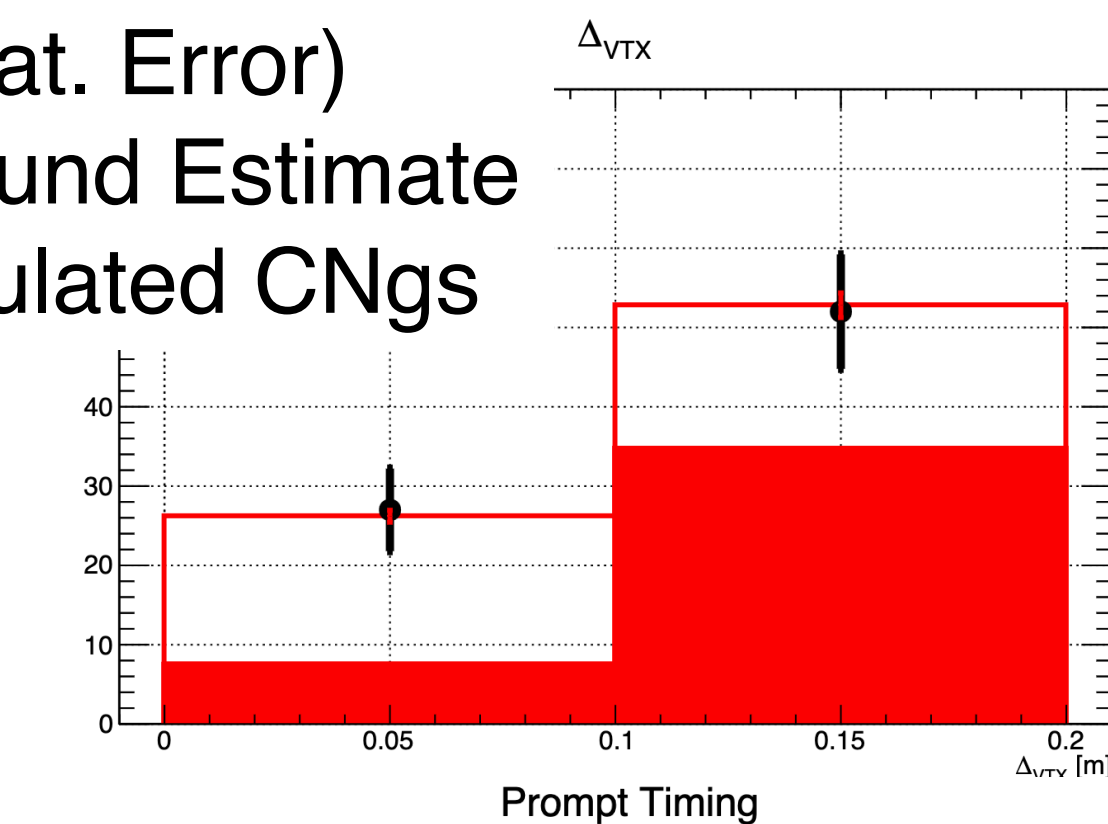
# CNGs Backgrounds

Estimated background rate is:  
 **$42.2 \pm 6.5(\text{stat.}) \pm 1.7(\text{syst.})$**

- Dominant background source is from accidental coincidences.
- Background PDFs are estimated by performing event selection in a late timing window ( $>1\text{ms}$  after beam)
- Normalization determined from fit in high  $\Delta$ Vertex region
- Reject background only hypothesis with  $p=2.9 \times 10^{-7}$



- Data (Stat. Error)
- Background Estimate
- MC Simulated CNGs

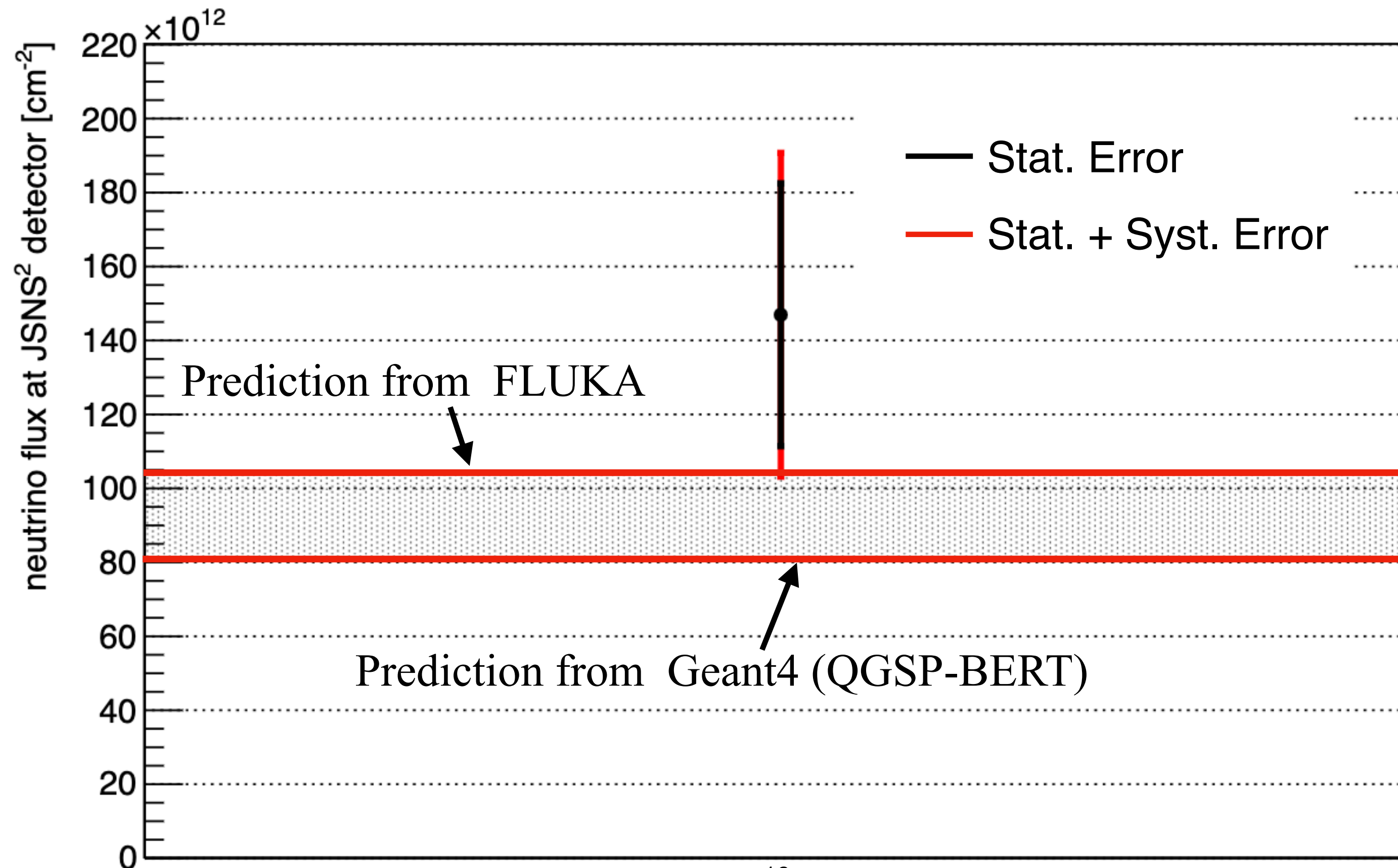


# CNgs Results

- Estimated detection efficiency:  $0.0588 \pm 0.0021$
- Combined LSND<sup>1</sup> + KARMEN<sup>2</sup> cross section estimate:  
 $9.1 \pm 0.7 \times 10^{-42} \text{cm}^2$
- JSNS<sup>2</sup> Measured neutrino flux:  
 $\Phi_{\nu_e} = \Phi_{\bar{\nu}_\mu} = \underline{1.47 \pm 0.36(\text{stat.}) \pm 0.26(\text{syst.}) \times 10^{14} \text{cm}^{-2}}$

$$N_{CNGS} = \frac{\text{Observed Event Rate}}{4\pi r^2} = \frac{\text{Neutrino Flux } \Phi_{\nu_e} \text{ POT}}{4\pi r^2} \frac{\text{Selection Efficiency } \epsilon}{\text{Cross Section } \sigma} N_C$$

Number of Carbons

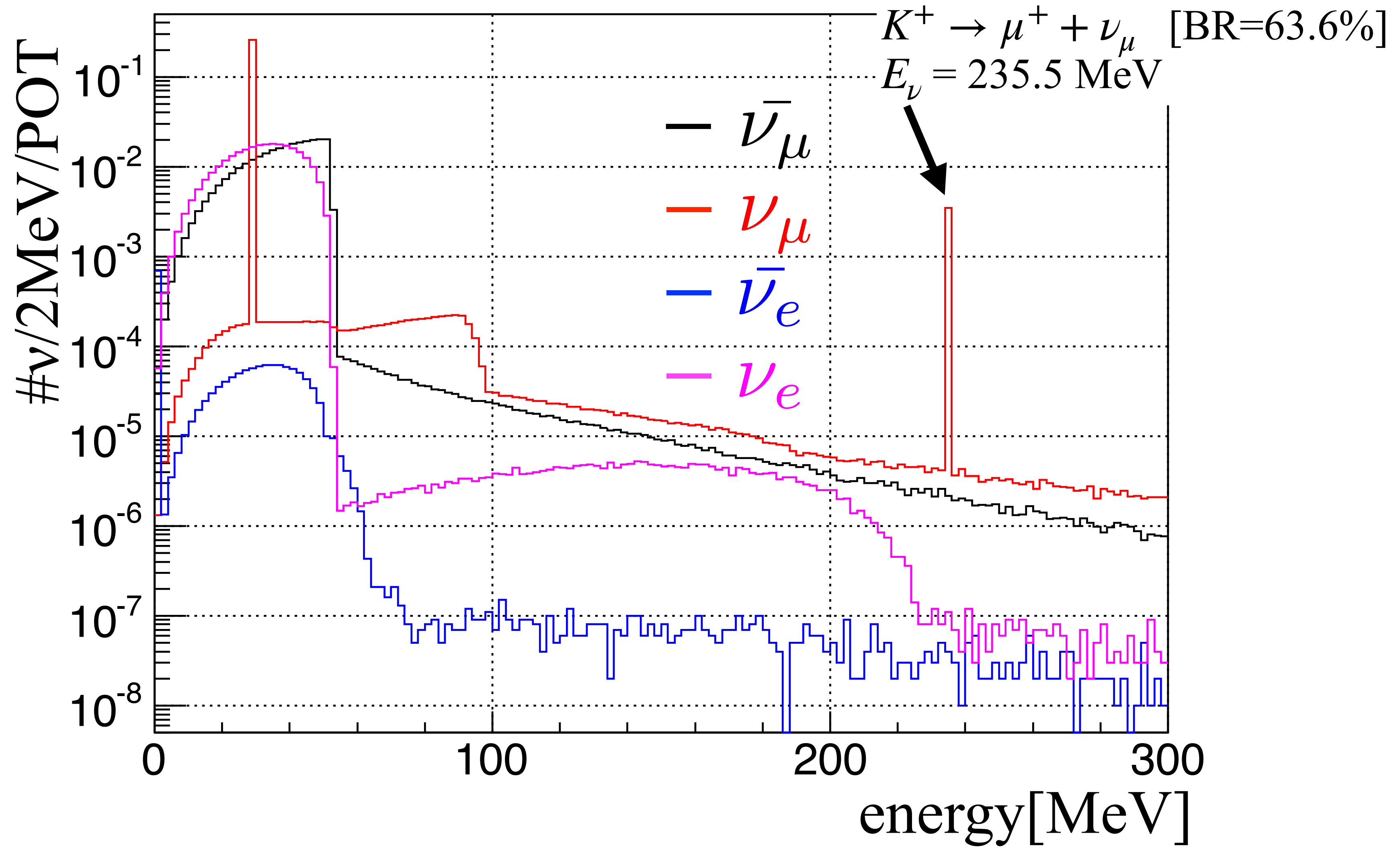


<sup>1</sup>LSND CNgs measurement:  
Phys. Rev. C 64:065501 (2001)

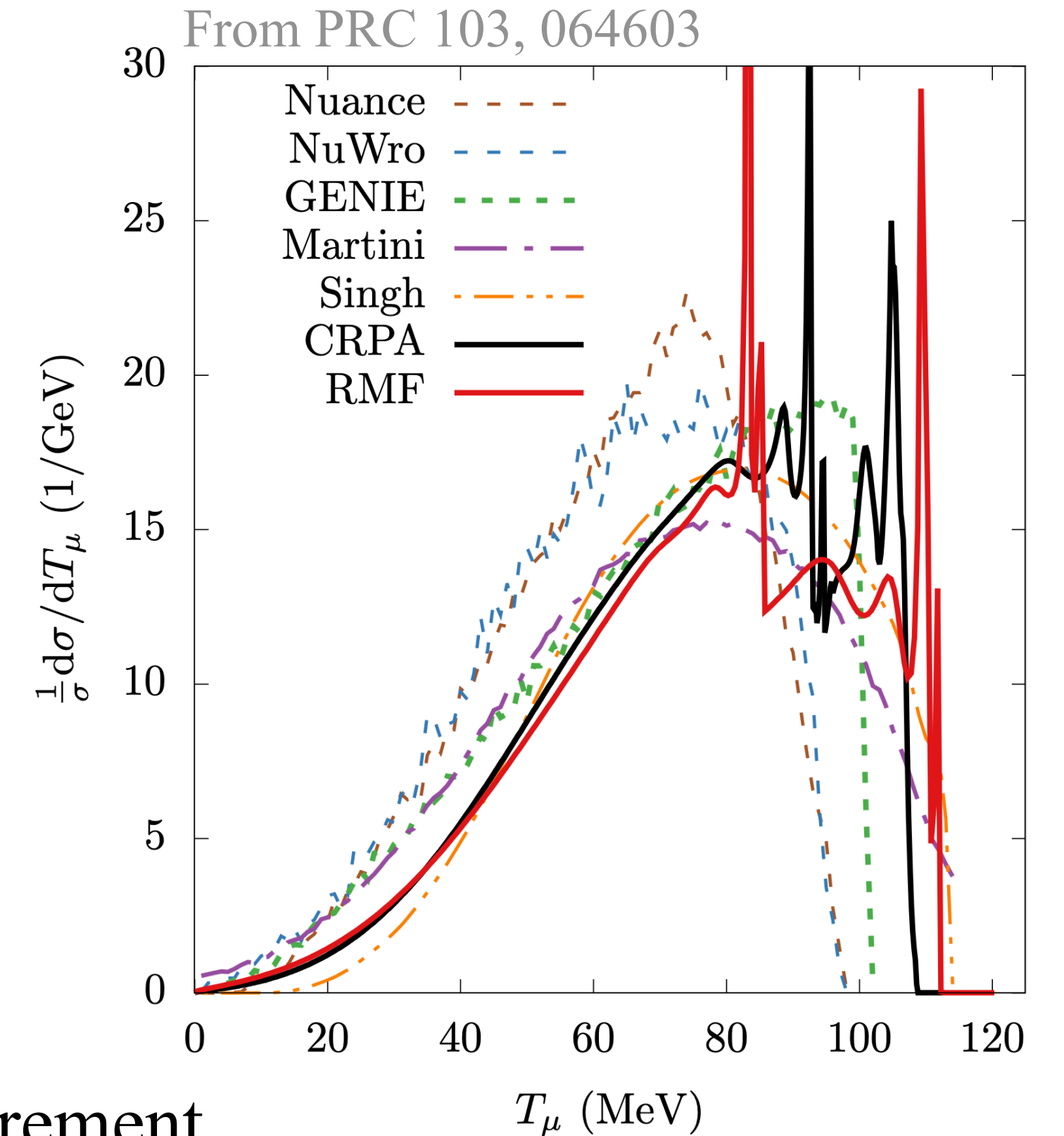
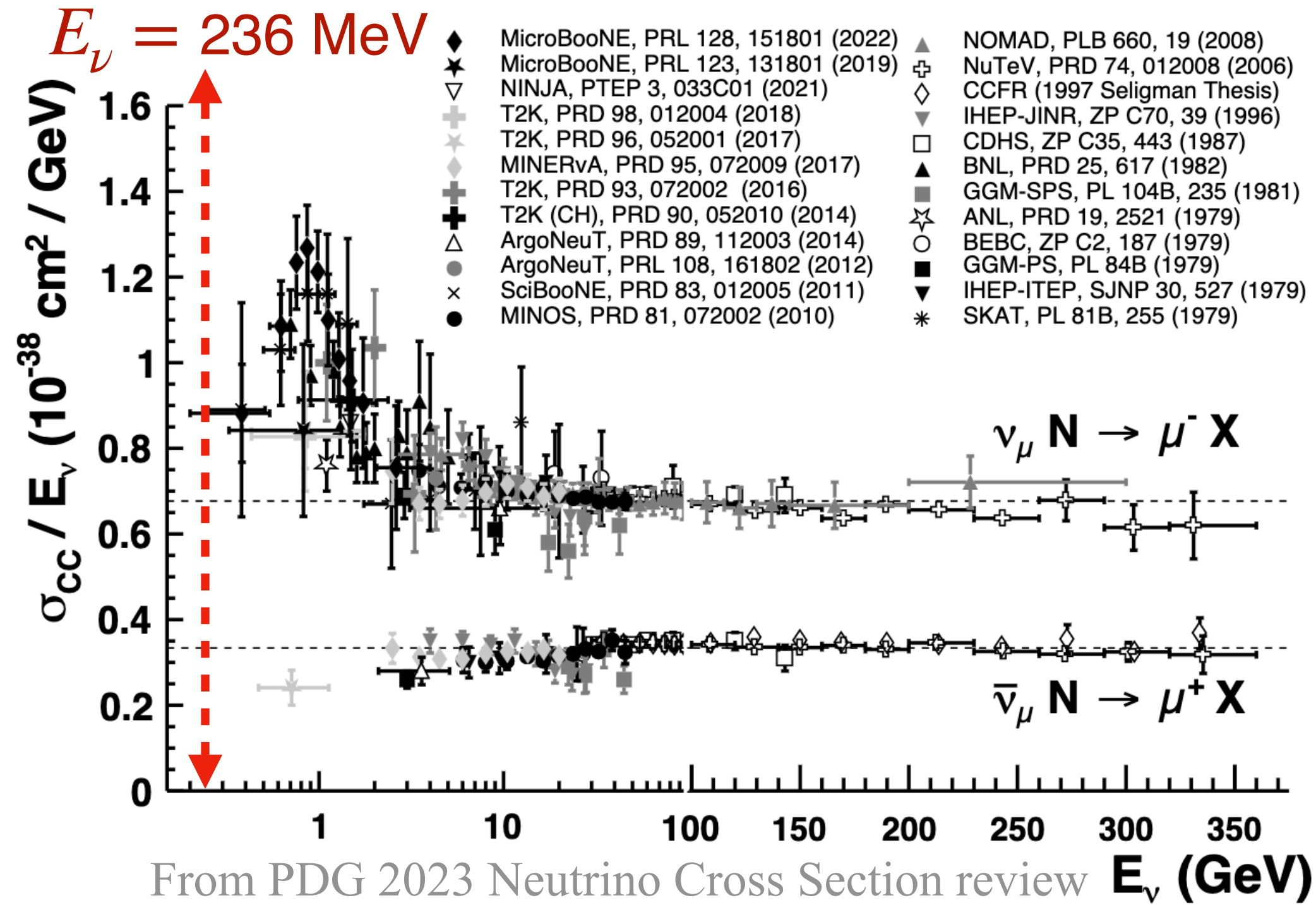
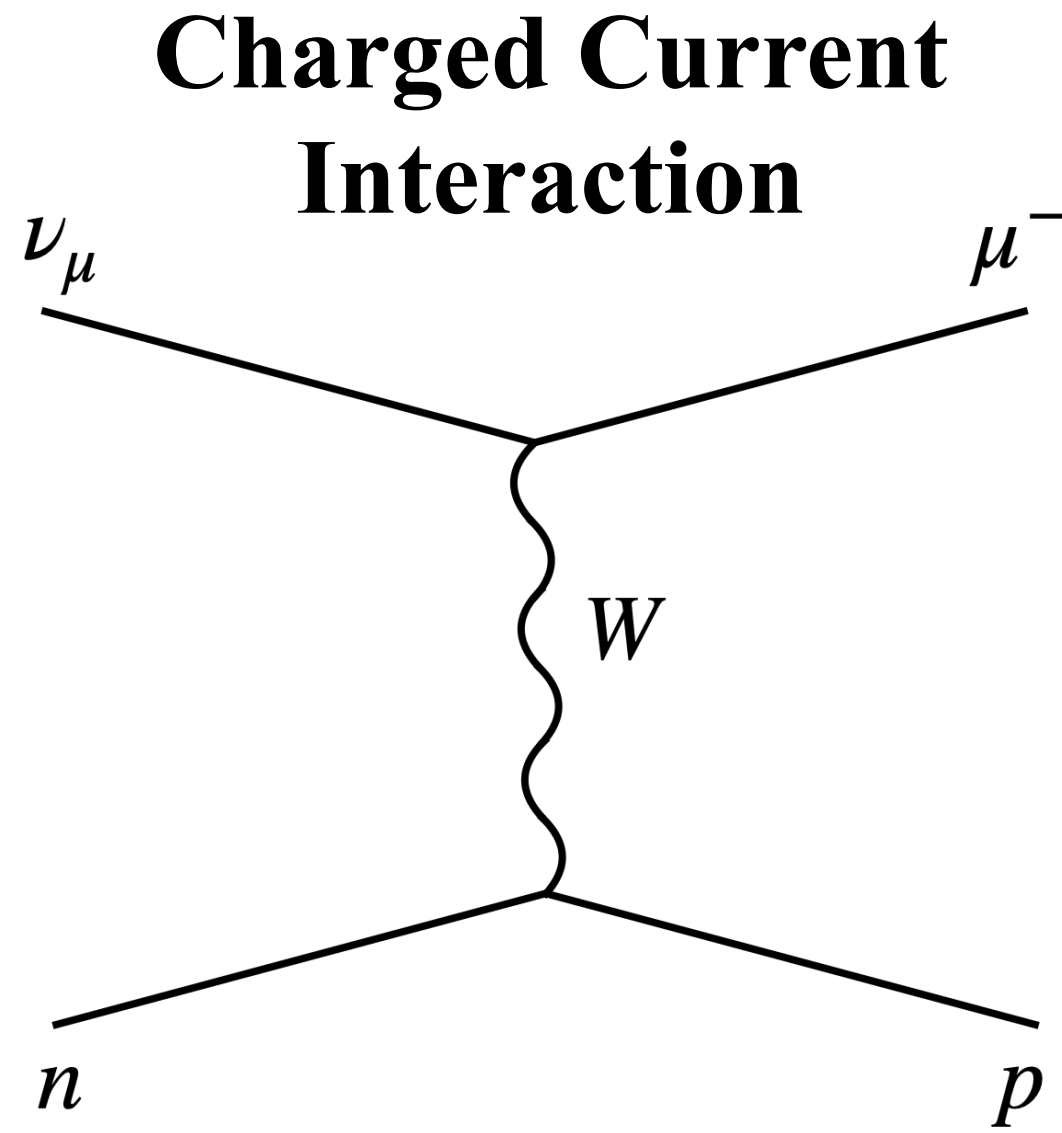
<sup>2</sup>KARMEN CNgs measurement:  
Nucl. Phys. 40, 183 (1998)



# Kaon Decay At Rest (KDAR)



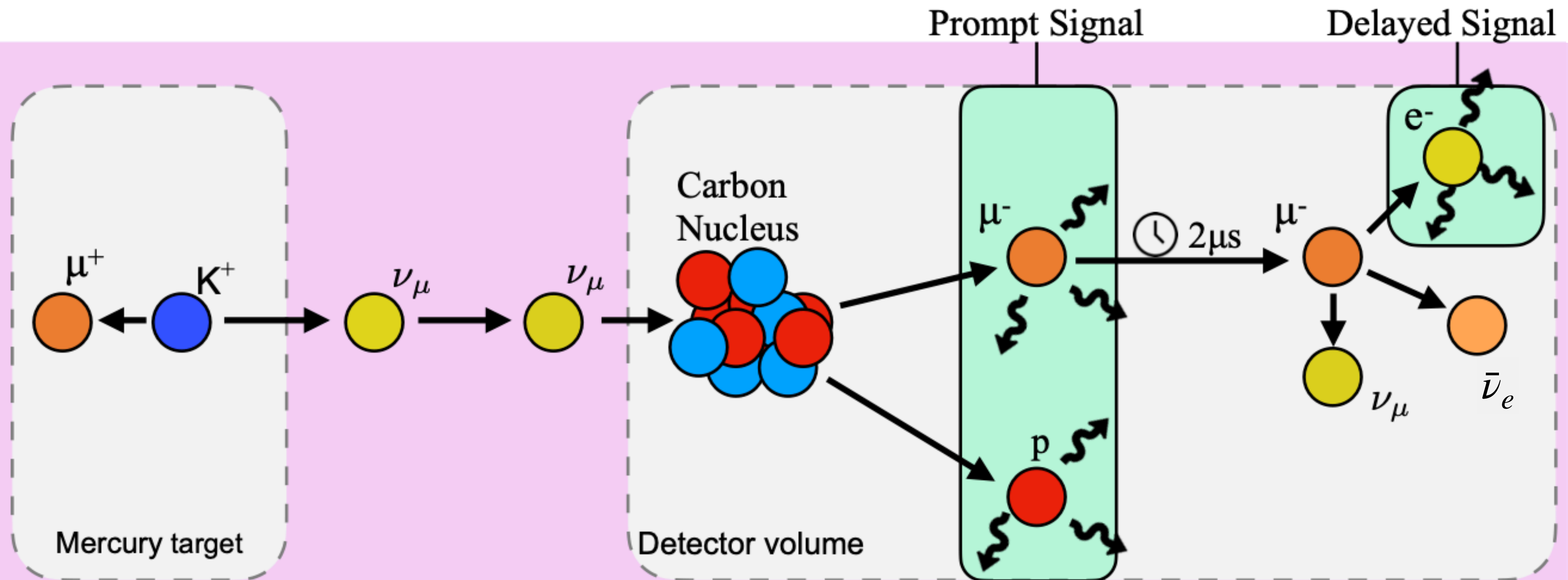
# KDAR



- KDAR neutrinos have a KNOWN energy, ideal for a cross section measurement
- Few cross-section measurements below 1GeV
- Nuclear models disagree significantly about the expected cross section & kinematics
- Only one prior KDAR measurement exists from MiniBoone (PRL 120, 141802)

# KDAR Neutrino Signal

- Kaons decay quickly ( $\tau=12.4\text{ns}$ ). KDAR interactions are prompt with the beam
- The prompt interaction produces a muon and (typically) 1 proton.
- Ensuing muon decay produces a secondary event



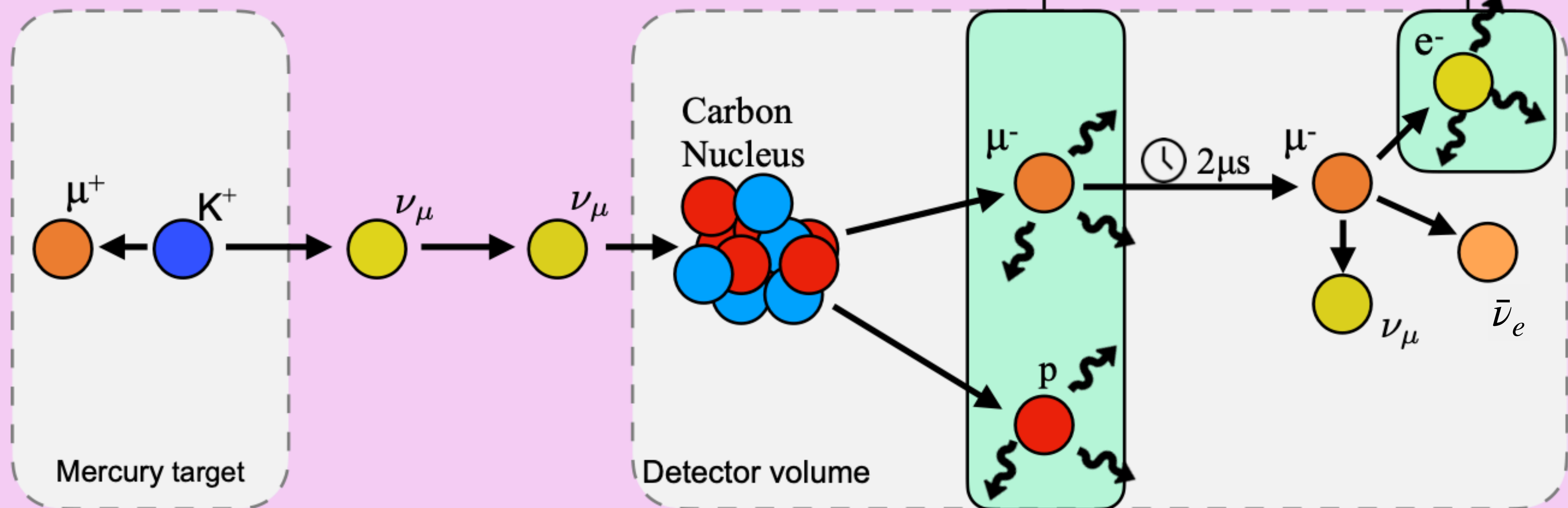
# KDAR Neutrino Signal

“Visible energy” is our observed variable

$$E_{\text{vis}} = T_{\mu} + \sum T_p$$

Prompt Signal

Delayed Signal

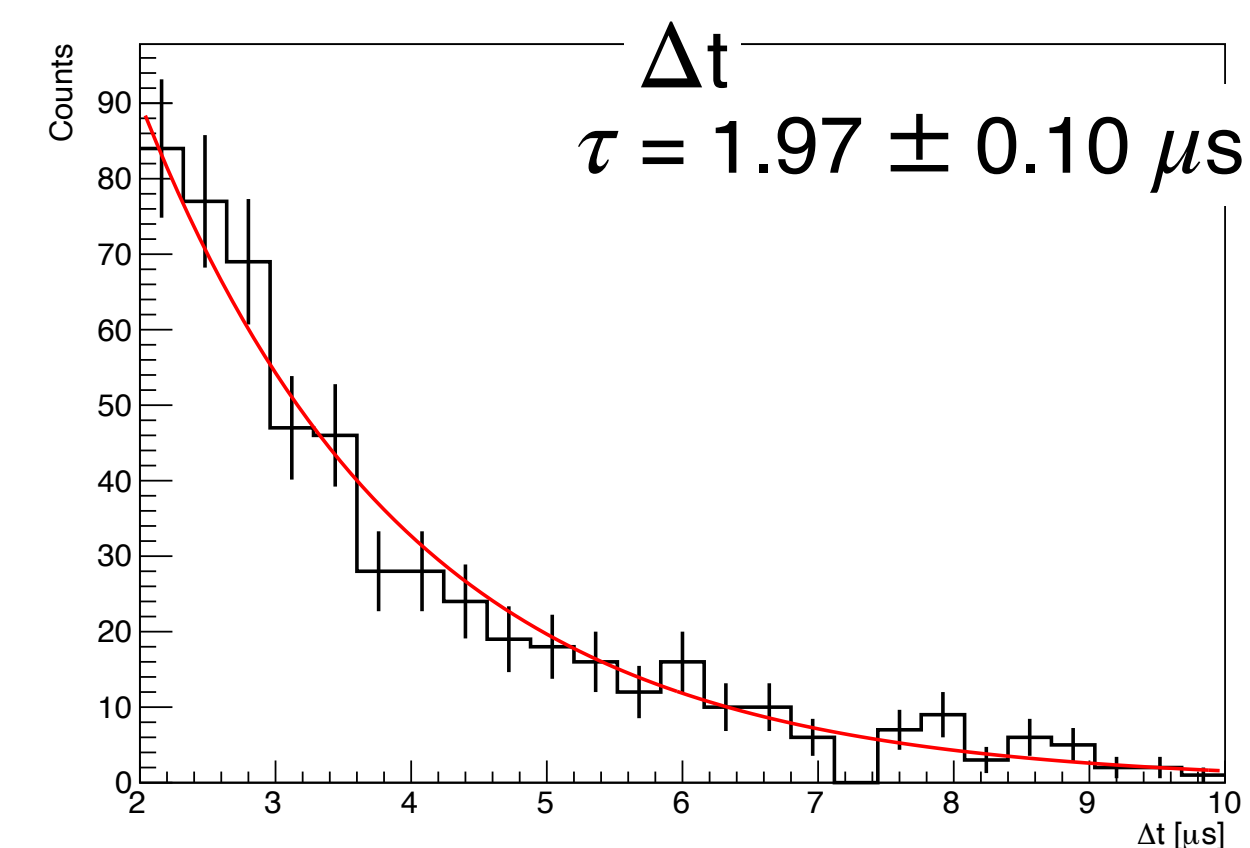
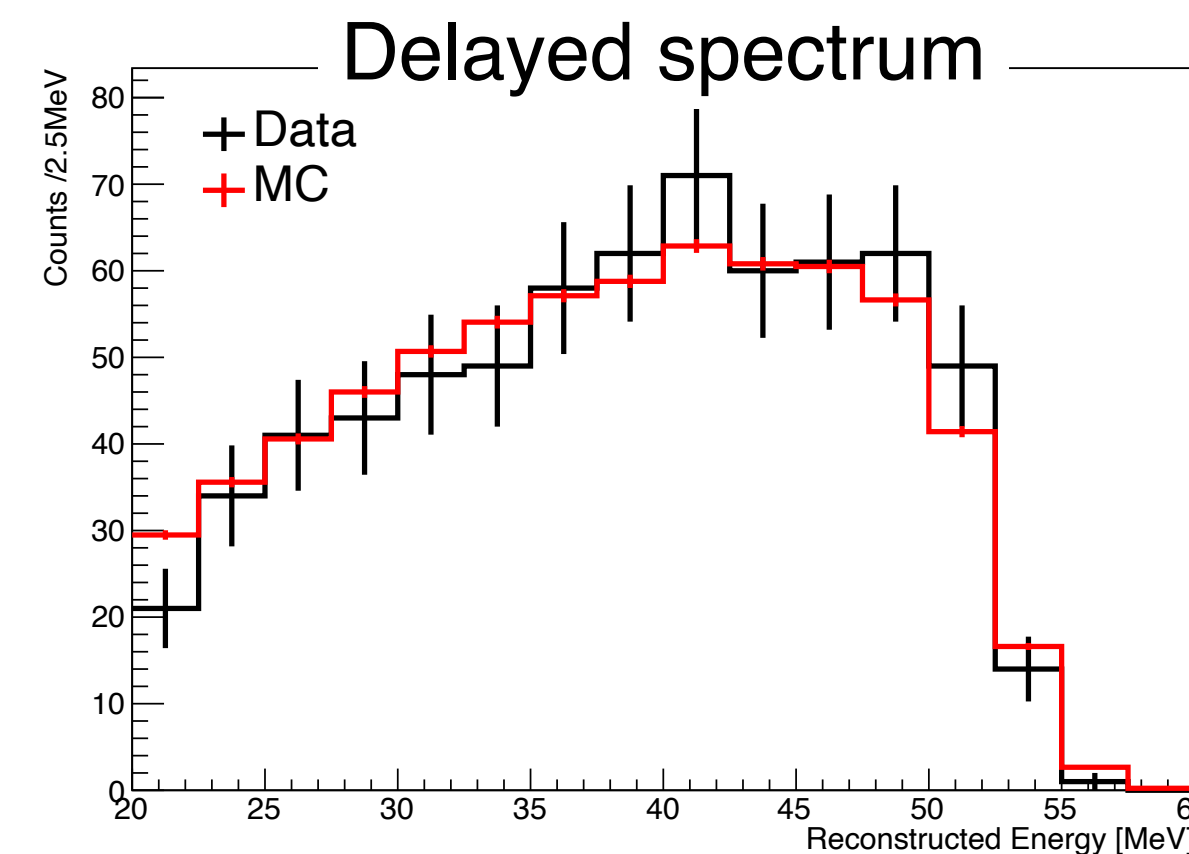
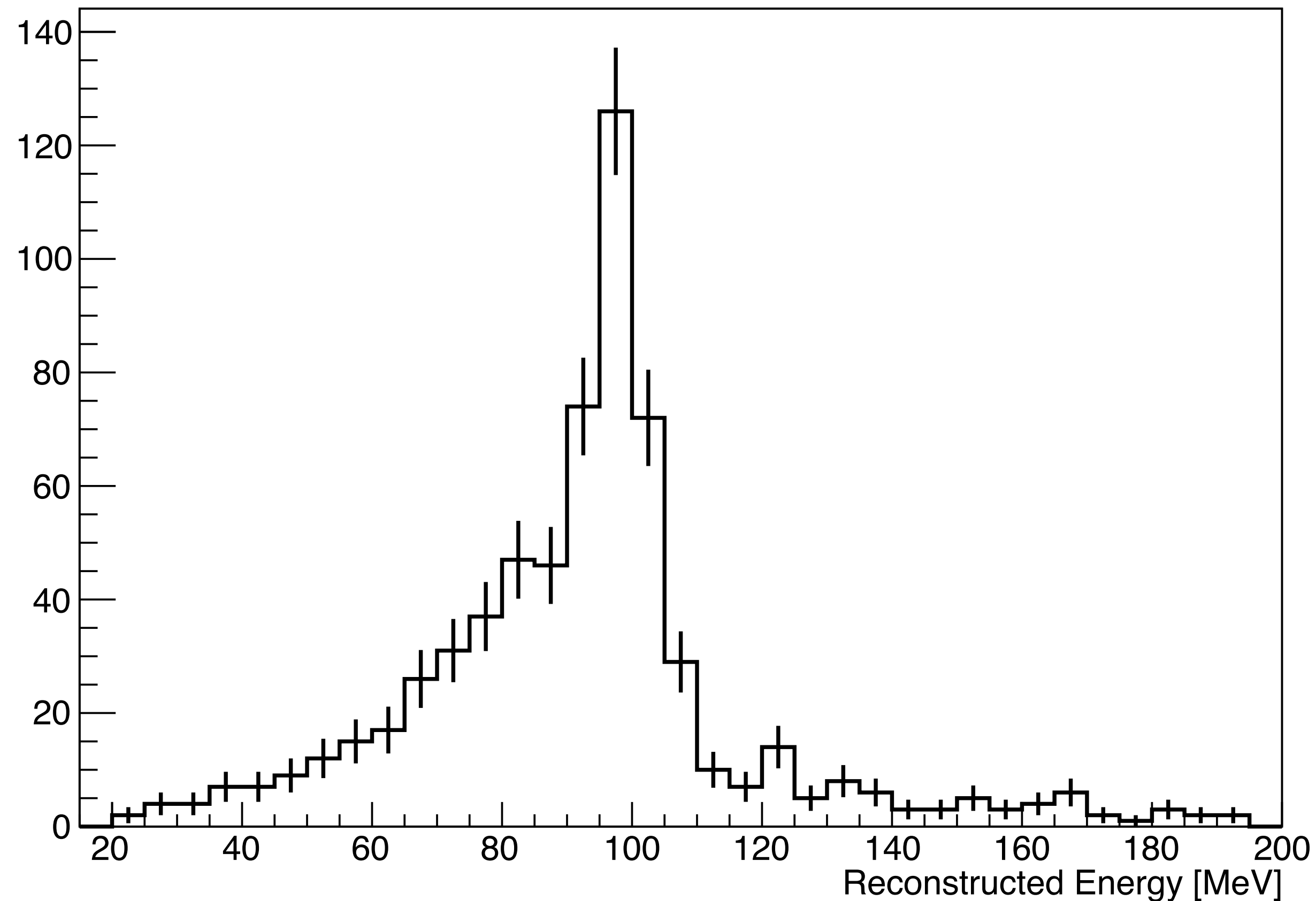


# KDAR Dataset

- 2021 dataset analyzed,  $1.4 \times 10^{22}$  POT
- **621 events pass all cuts**
- Selection criteria...

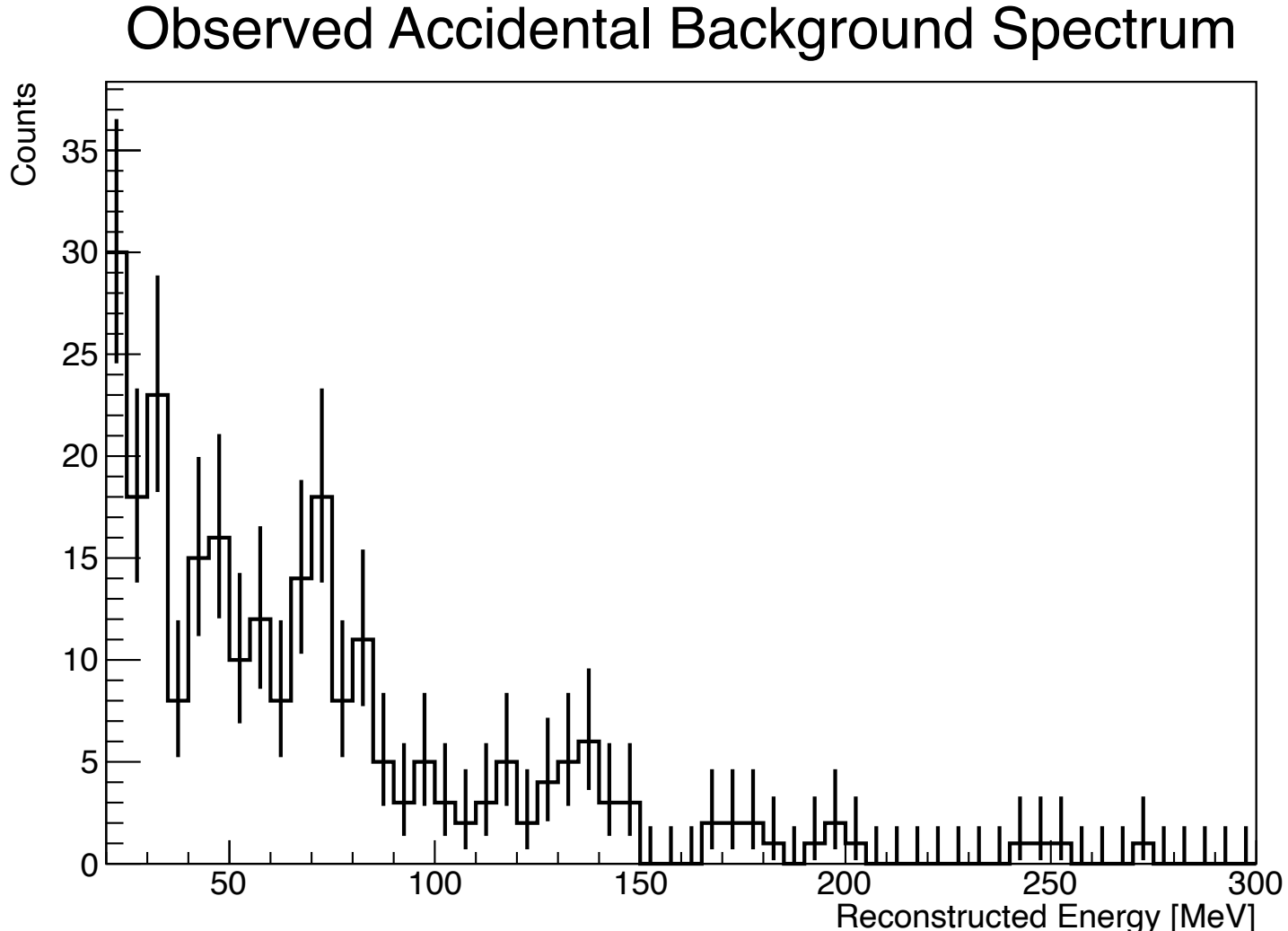
	Prompt	Delayed
<b>Energy</b>	20-150 MeV	20-60 MeV
<b>Timing</b>	2x150ns Beam centered windows	$\Delta t < 10\mu\text{s}$
<b>Position</b>	Fiducial Volume: R<1400mm -1000mm < z < 500mm	$\Delta\text{Vertex} < 300\text{mm}$

Energy spectrum of prompt events passing KDAR selection

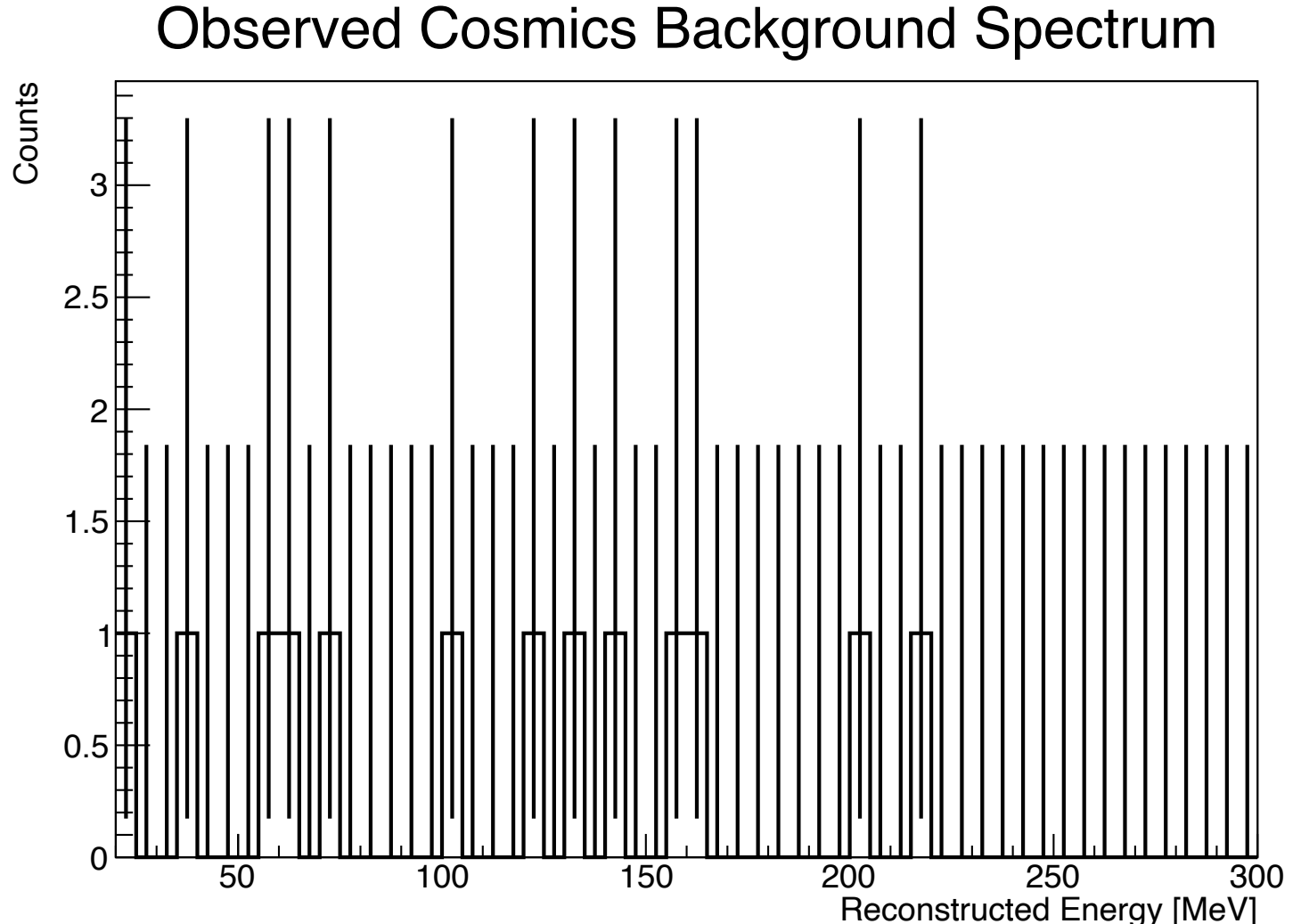


# KDAR Backgrounds

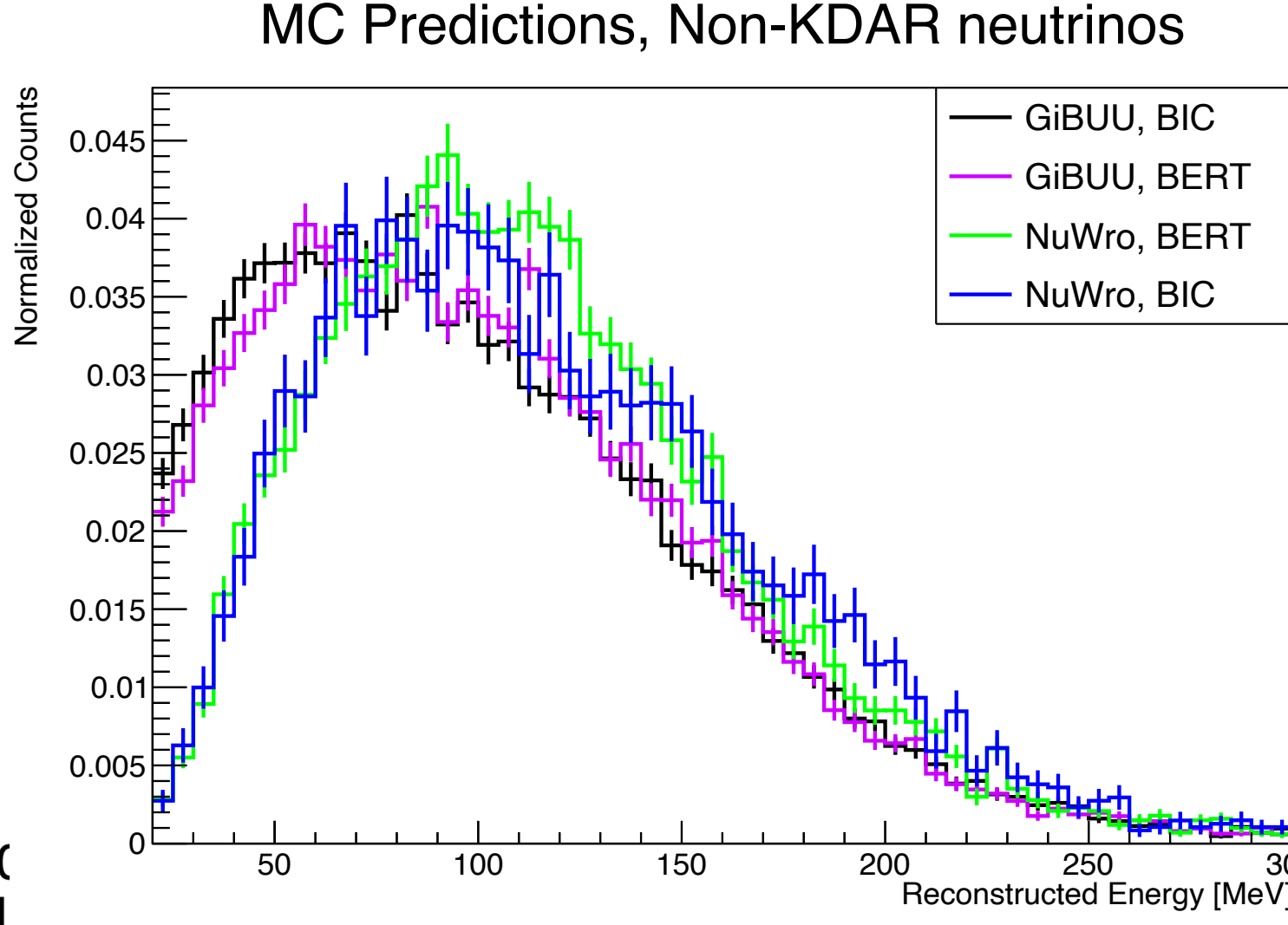
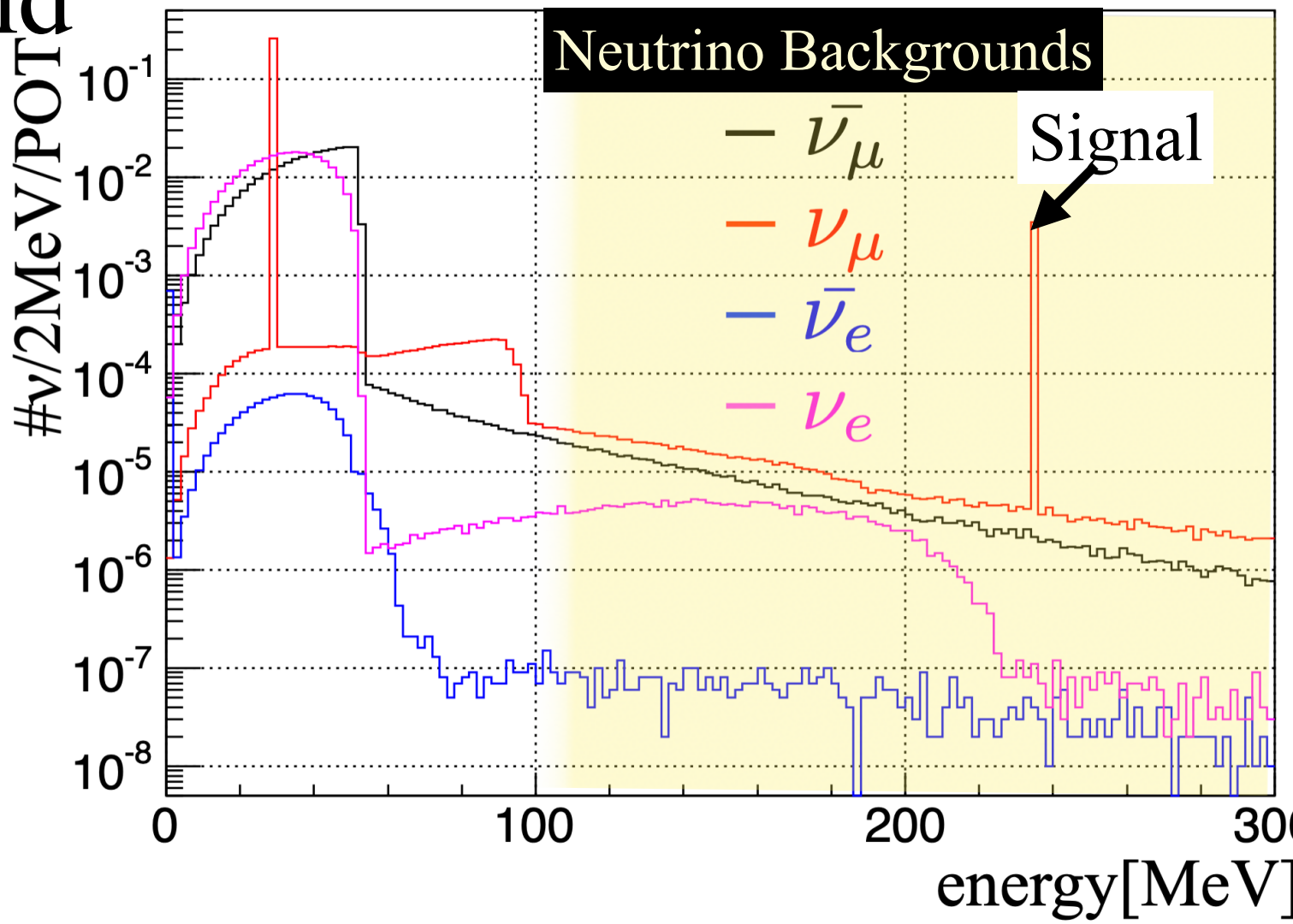
- **Dominant background source is pion decay-in-flight (DIF) neutrinos**
- DIF background spectral shape estimated with MC
- Both NuWro & GiBUU event generators are used for DIF background simulation
- Normalization estimated using the kinematically disallowed ( $>150\text{MeV}$ ) portion of the KDAR spectrum



< 1.14 events in KDAR dataset

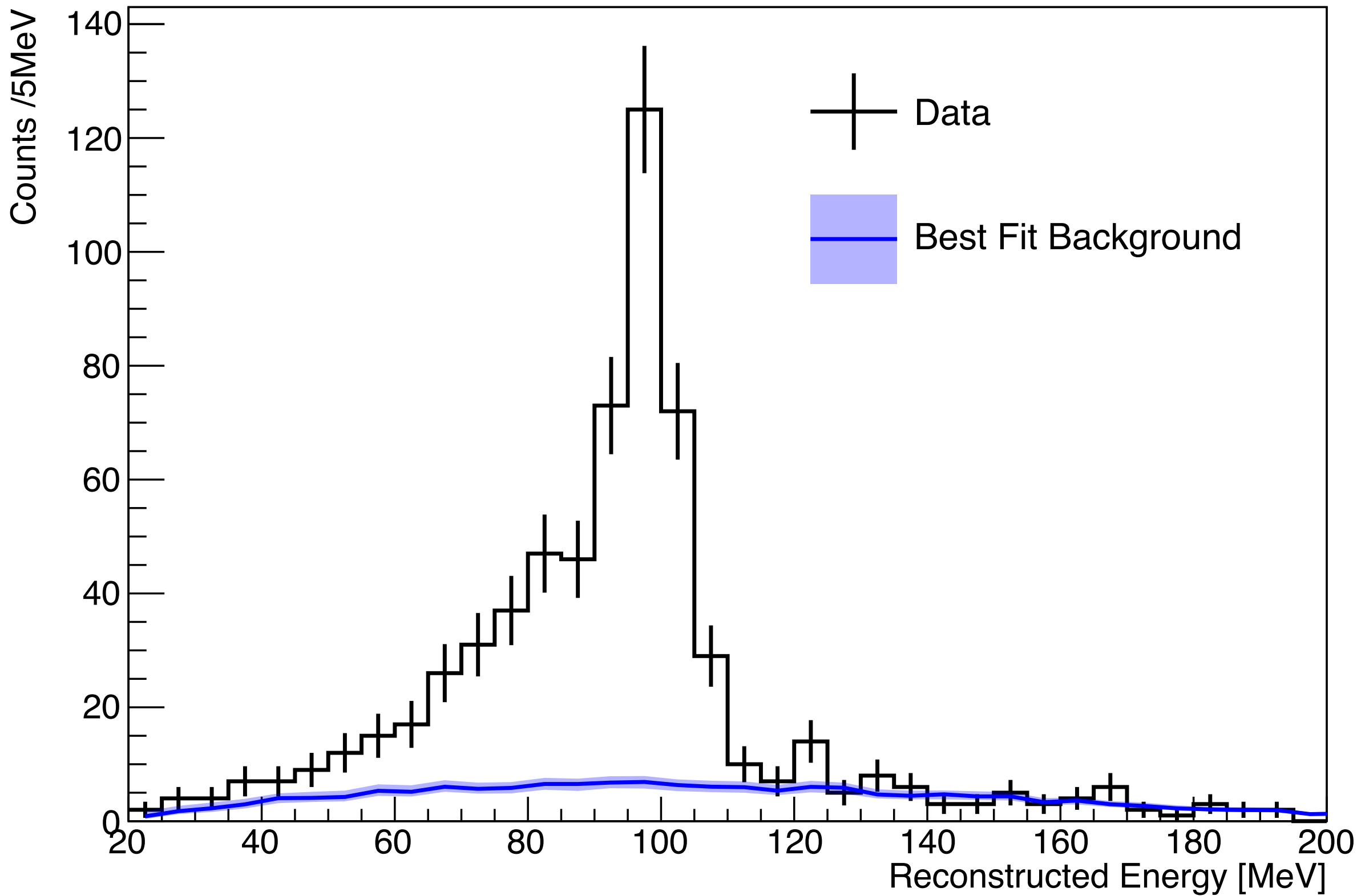


< 6.57 events in KDAR dataset

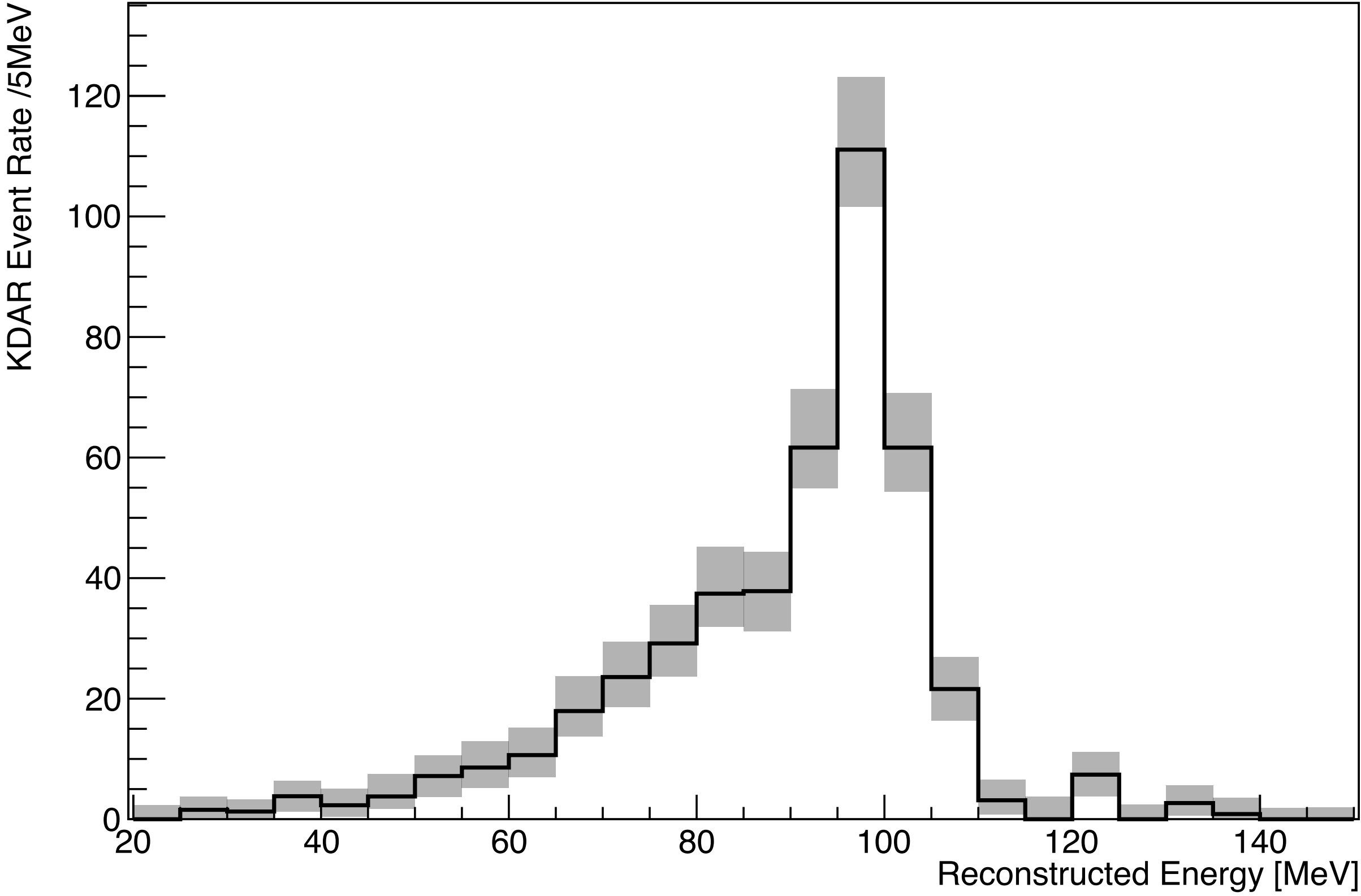


# KDAR Signal Extraction

- Best fit signal rate:  $492.2 \pm 29.1$  events
- Background rate:  $131.1 \pm 19.2$  events(20-150MeV)

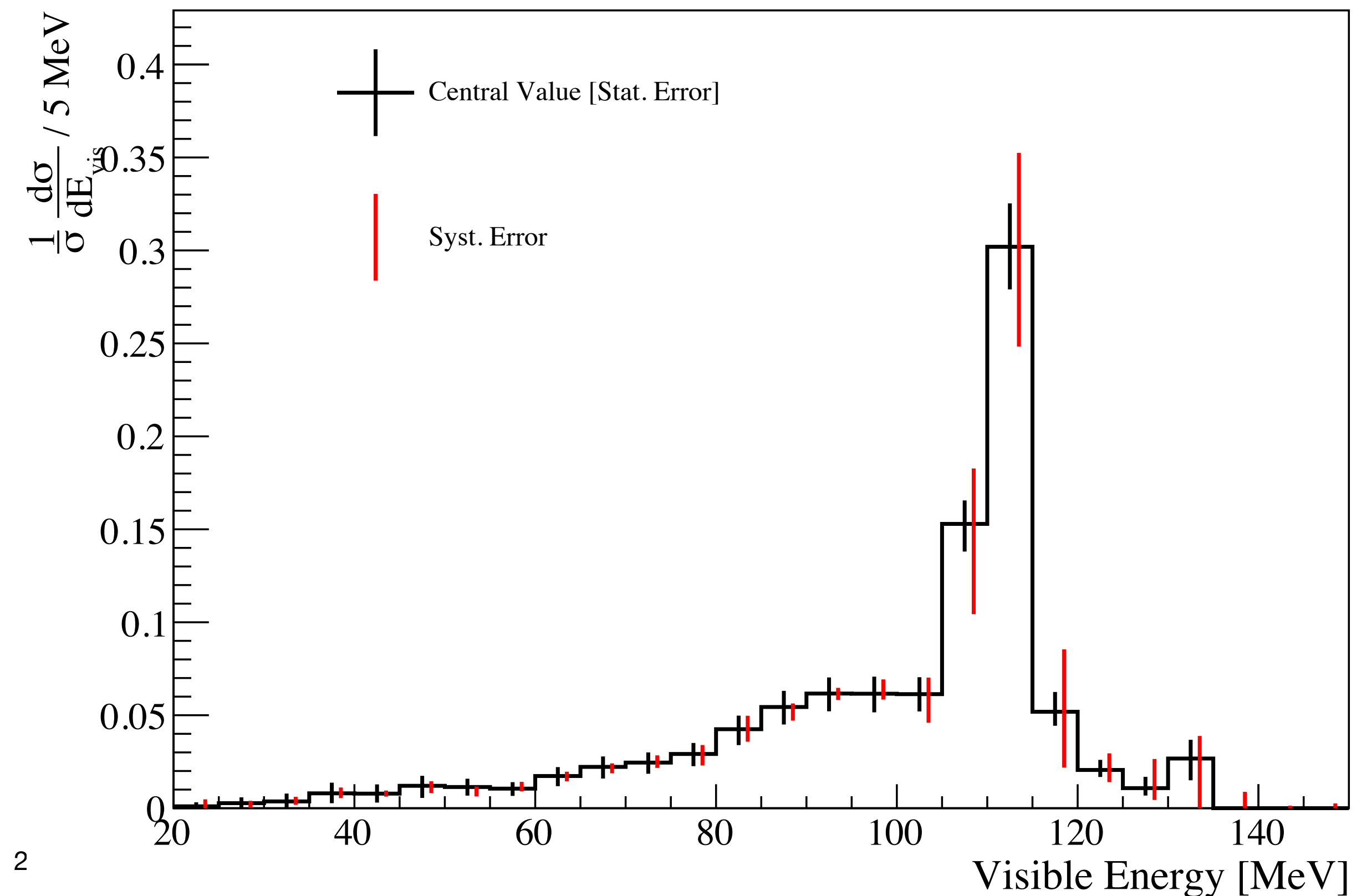
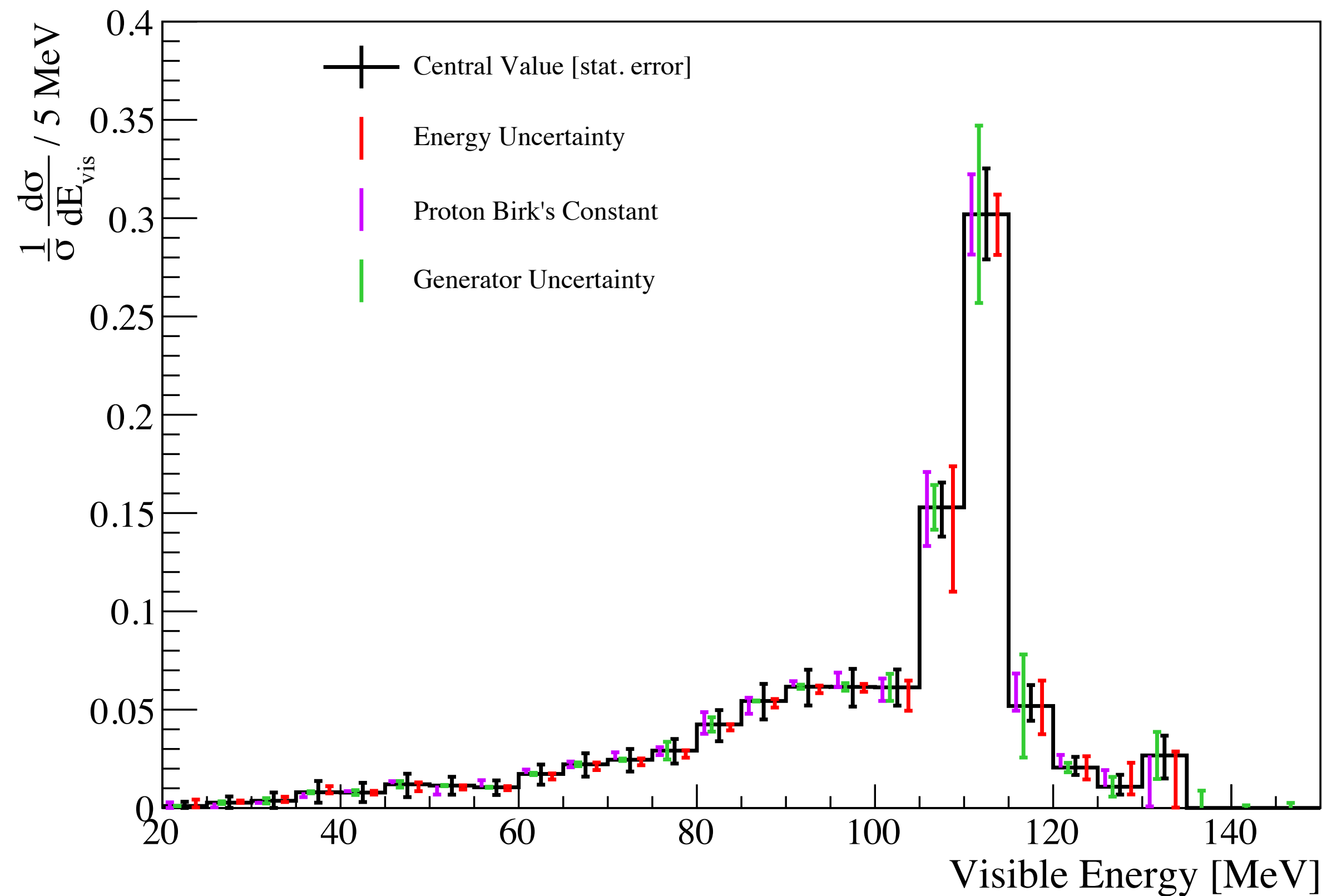


Background subtracted KDAR spectrum



# KDAR Results

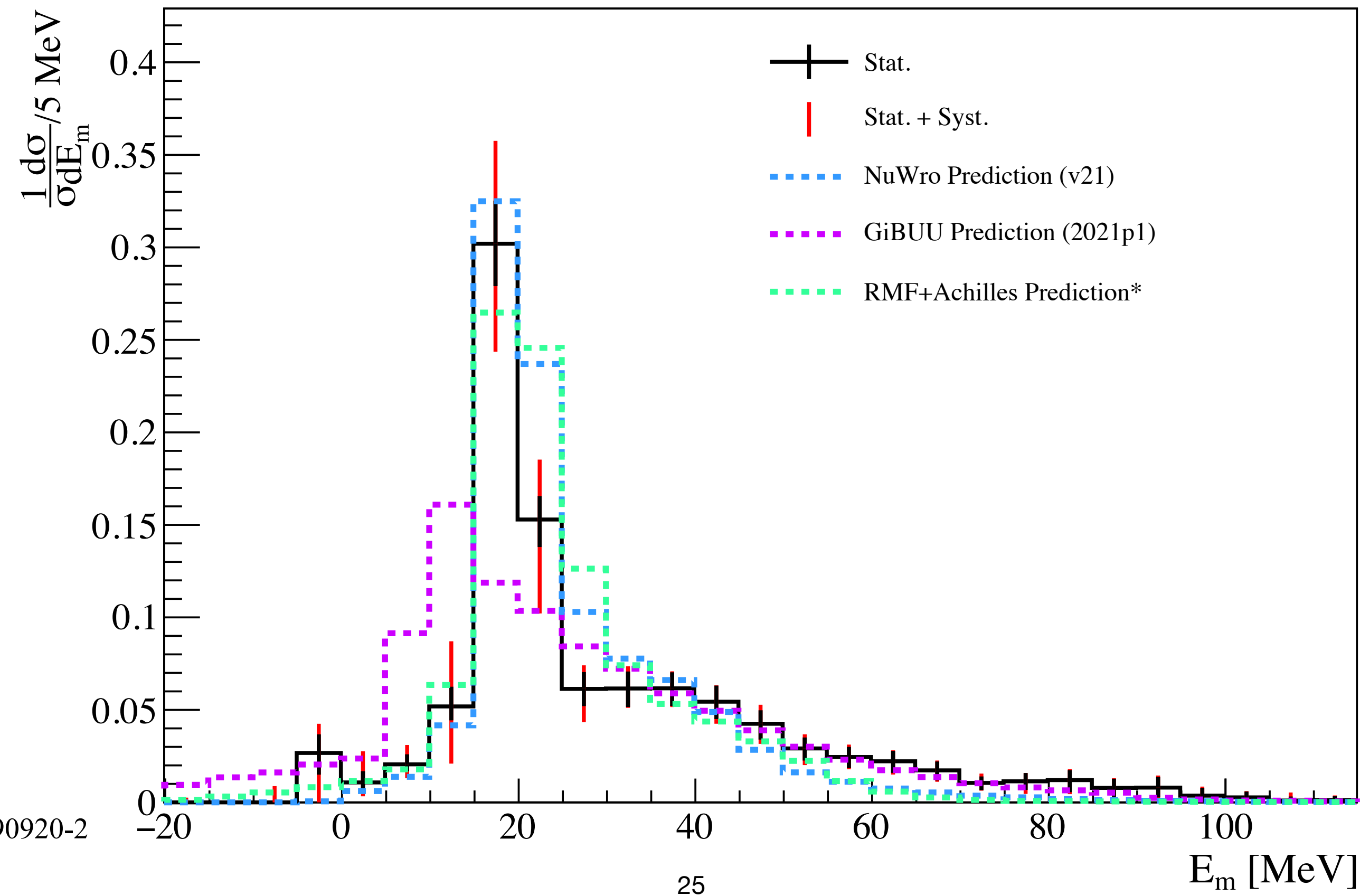
- We use Iterative Bayes (D'Agostini) unfolding, detector response estimated using NuWro & GiBUU event generators.
- **Shape only measurement.** Kaon production rates are too uncertain to estimate the integrated cross section.
- The largest source of error is the “generator” systematic.
- Measurement will improve with better modeling and additional statistics





# First KDAR Missing Energy Measurement

- This result can also be in terms of “missing energy”
- $E_{miss} = E_\nu - m_\mu - E_{vis}$
- The neutrino energy is KNOWN. Any unobserved/extra energy was absorbed/provided by the nucleus



\*Provided by Alexis Nikolakopoulos[1-4].

[1] RMF: arxiv:1904.10696, arxiv:2104.01701

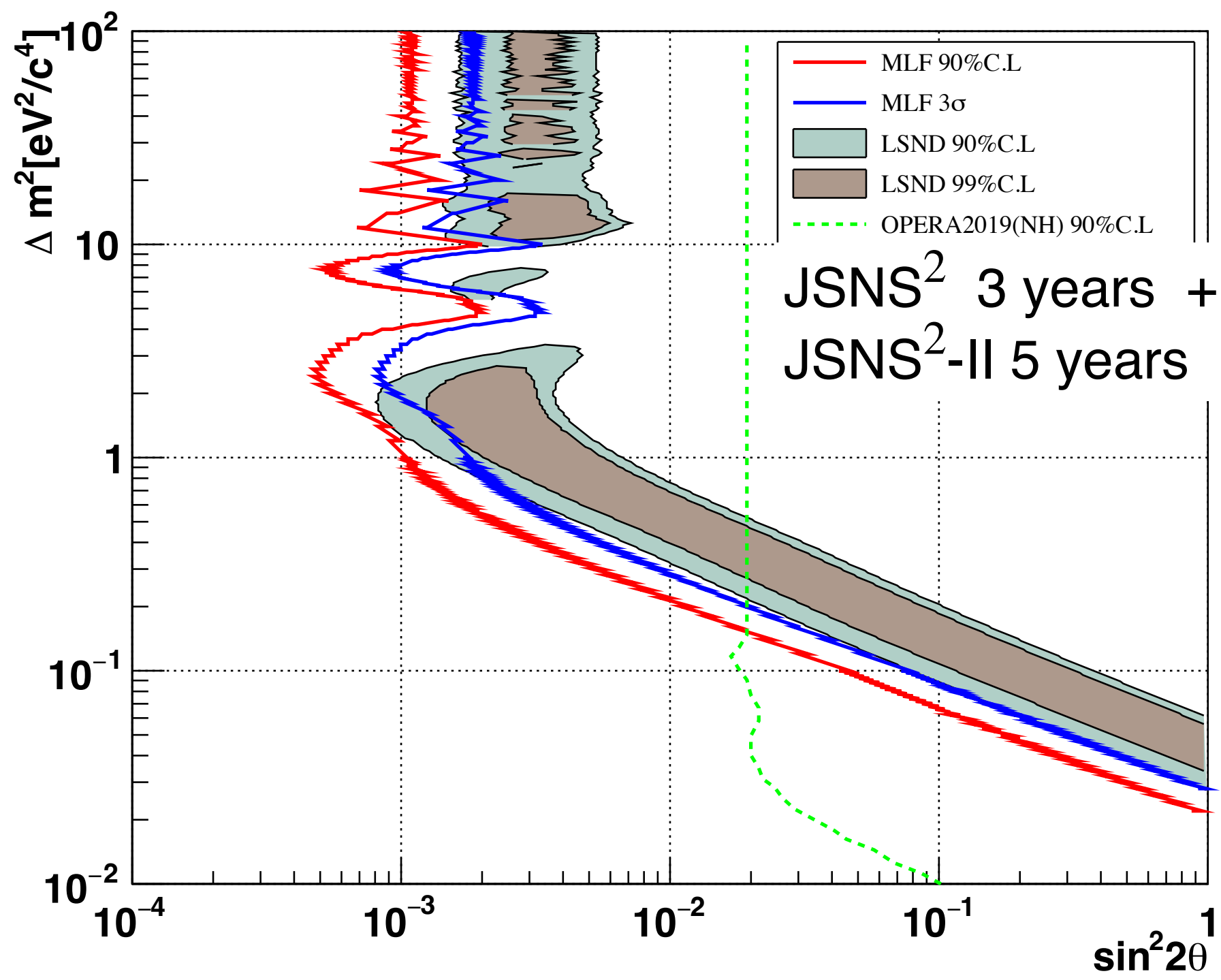
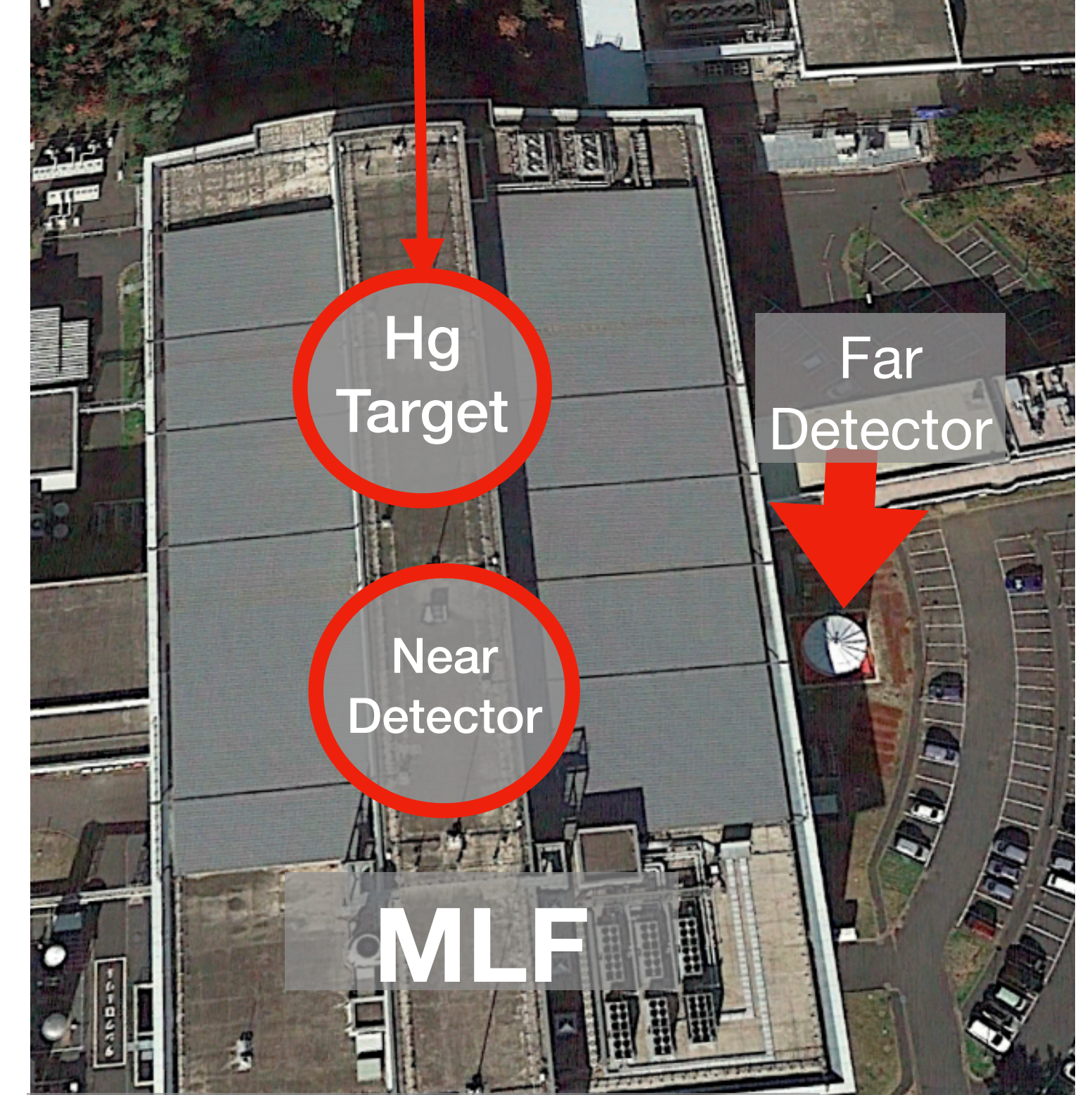
[2] Spectral Function: doi.org/10.1016/0375-9474(94)90920-2

[3] INC: arxiv:2007.15570

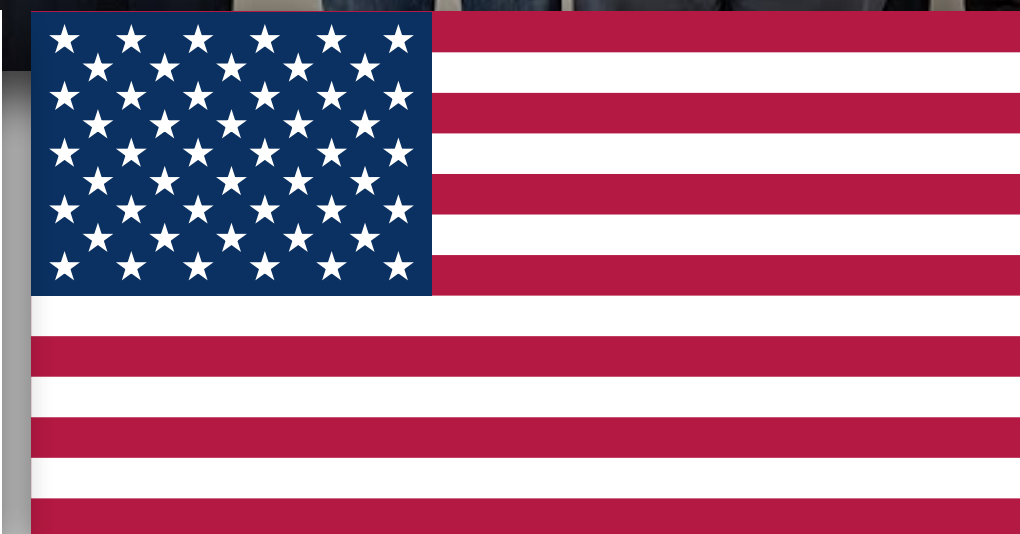
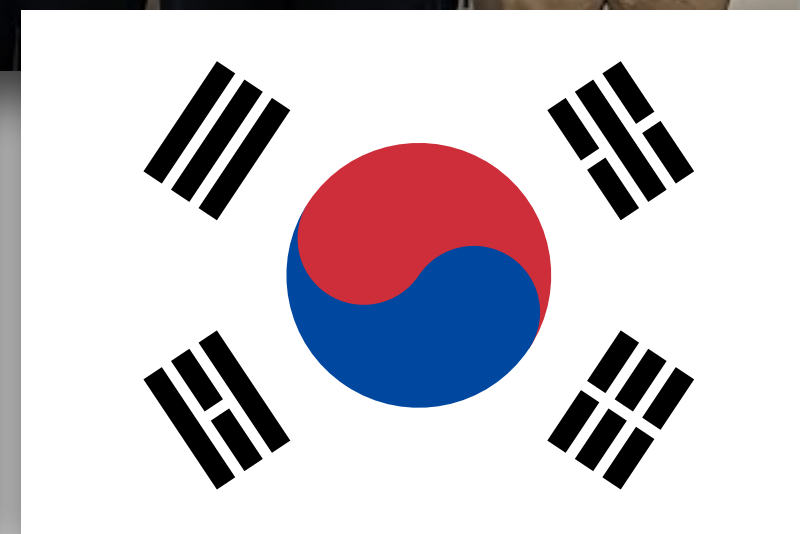
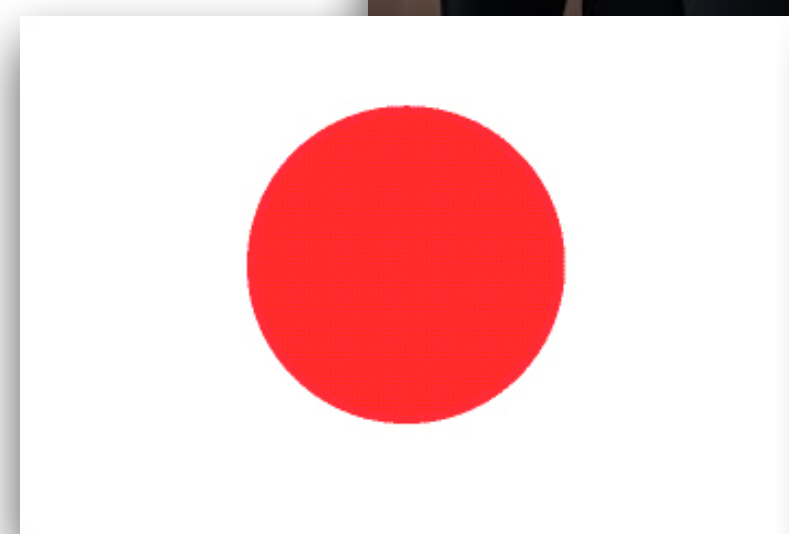
[4] Achilles: arxiv:2205.06378

(arXiv:2012.10807)

- Target volume 32-tonnes, **baseline is 48m**
- Construction began in 2021, expected to begin **data taking in the 2024-2025 run period**
- A second baseline allows observation of “smoking gun” oscillation signature
- Improves sensitivity across parameter space, especially at low  $\Delta m^2$  values



# Thank you for your attention



# Backup

# JSNS<sup>2</sup> vs. LSND

	<b>LSND</b>	<b>JSNS<sup>2</sup></b>	<b>Notes</b>
<b>Detector Mass</b>	167t	17t	-
<b>Baseline</b>	30m	24m	-
<b>Beam Proton Energy</b>	0.8GeV	3GeV	Allows for KDAR measurement. Expect ~10x higher pion production
<b>Beam Power</b>	800kW	1MW	-
<b>Beam Duty Factor</b>	600μs x120Hz	100ns(x2) x25Hz	Expect ~300x fewer ambient IBD backgrounds
<b>Detector Medium</b>	Dilute LS	Gd-LS	-
<b>Neutron Capture</b>	Hydrogen, ~0.2ms, 2.2MeV	Gadolinium, ~26μs, 8MeV	Shorter capture time & higher energy mean fewer backgrounds
<b>Particle ID</b>	Cherenkov	PSD	-

# CNGs Efficiencies and Cross Section

- Neutrino flux estimate requires we carefully estimate our detection efficiency
  - **Estimated average efficiency:  $0.0588 \pm 0.0021$**
- We use a combined Karmen+LSND cross section measurement

$$N_{CNGS} = \frac{\Phi_{\nu_e} \text{POT}}{4\pi r^2} \epsilon \sigma N_C$$

Observed Event Rate  $\rightarrow N_{CNGS}$

Neutrino Flux  $\rightarrow \Phi_{\nu_e}$

Selection Efficiency  $\rightarrow \epsilon$

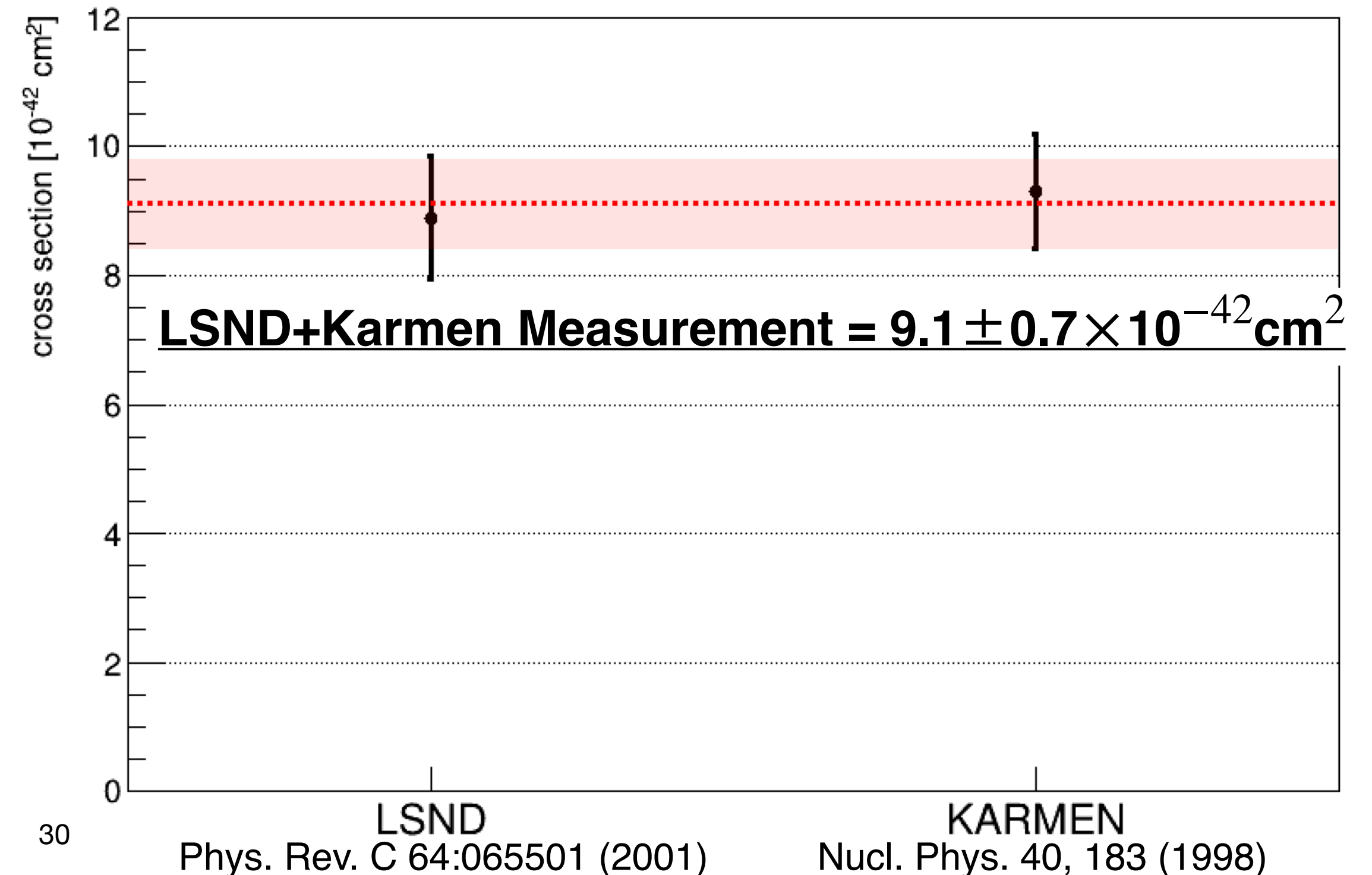
Interaction Cross Section  $\rightarrow \sigma$

Number of Carbons  $\rightarrow N_C$

## Efficiency Estimates

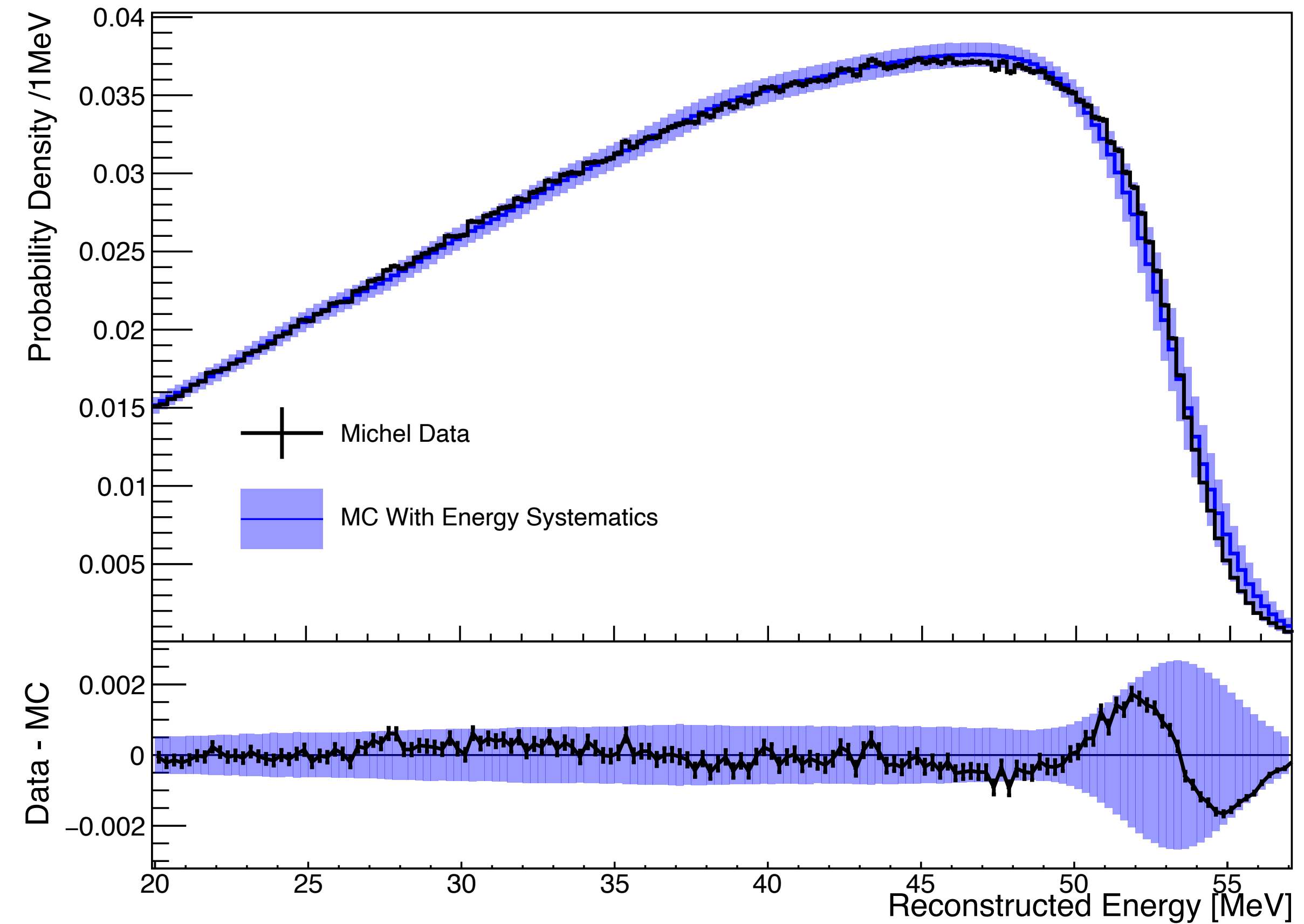
item	efficiency	error (%)	
bkg estimation (MST vs SS)	-	10.7 %	
Energy	2021: 0.255 2022: 0.255	2021: 4.6 % 2022: 5.0 %	
(FADC) Timing	prompt	2021: 0.498 2022: 0.468	2021: 0.4 % 2021: 0.2 %
	$\Delta_t$	2021: 0.5180 2022: 0.7805	2021: 0.08 % 2022: 0.05 %
muon veto	2021: 0.885 2022: 0.901	2021: 0.4 % 2022: 0.3 %	
Michel electron veto	prompt	0.9930	0.01 %
	delayed	0.9768	0.03 %
$\Delta_{VTX}$	0.881	0.8 %	
fiducial	- (related to the number of $^{12}\text{C}$ )	8 %	
item	value and error		
The number of $^{12}\text{C}$	$(4.68 \pm 0.37) \times 10^{29}$ (PSD fiducial volume)		
Cross section	$(9.1 \pm 0.7) \times 10^{-42} \text{ cm}^2$		

## Cross Section Estimates



# KDAR Energy Calibration

- Energy reconstruction is calibrated & systematics are determined using Michel MC & data.
- Additional systematics from the comparison of the energy scale observed from neutron capture (8MeV) and from Michel (~53MeV).
- The energy scale systematics are constrained to 0.68% uncertainty.



# $^{252}\text{Cf}$ Calibration

arXiv:2404.04153

- We developed a 3D calibration arm for our  $^{252}\text{Cf}$  source
- It was used to improve our optical model used for reconstruction

