

Other present and future $0\nu\beta\beta$ experiments

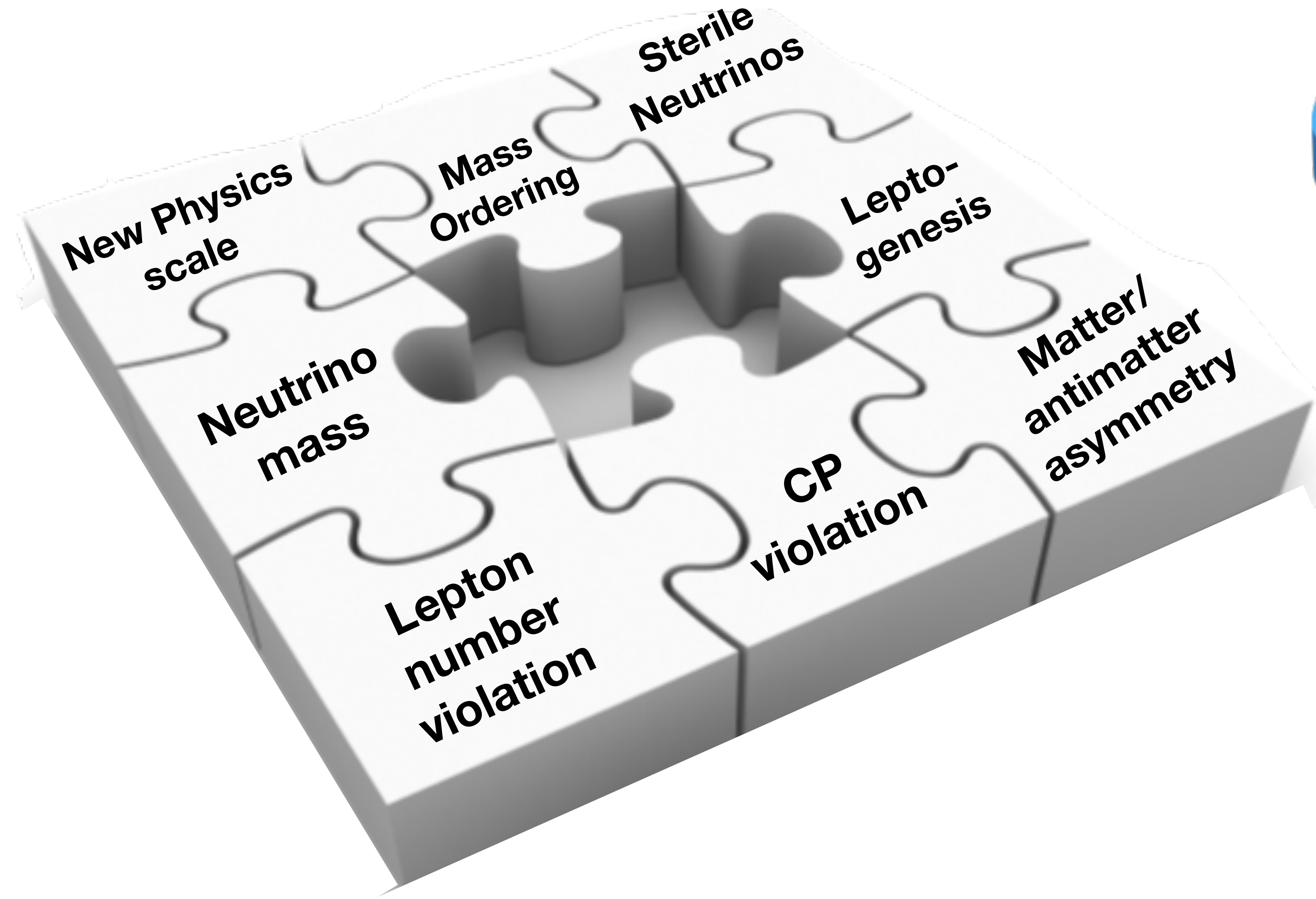
Roxanne Guénette

MANCHESTER
1824

The University of Manchester

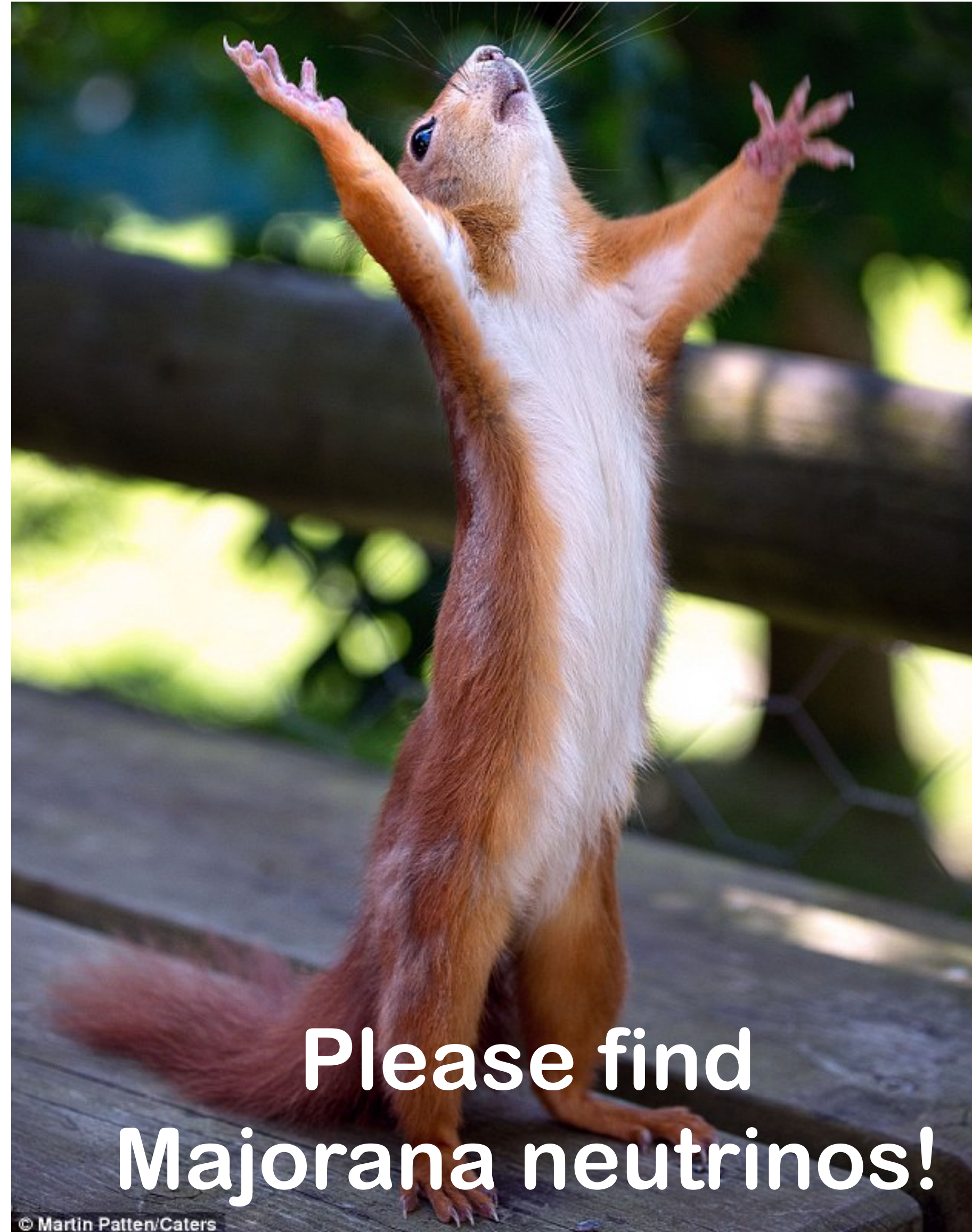
NEUTRINO 2024
18 June 2024

Importance of searching for $0\nu\beta\beta$

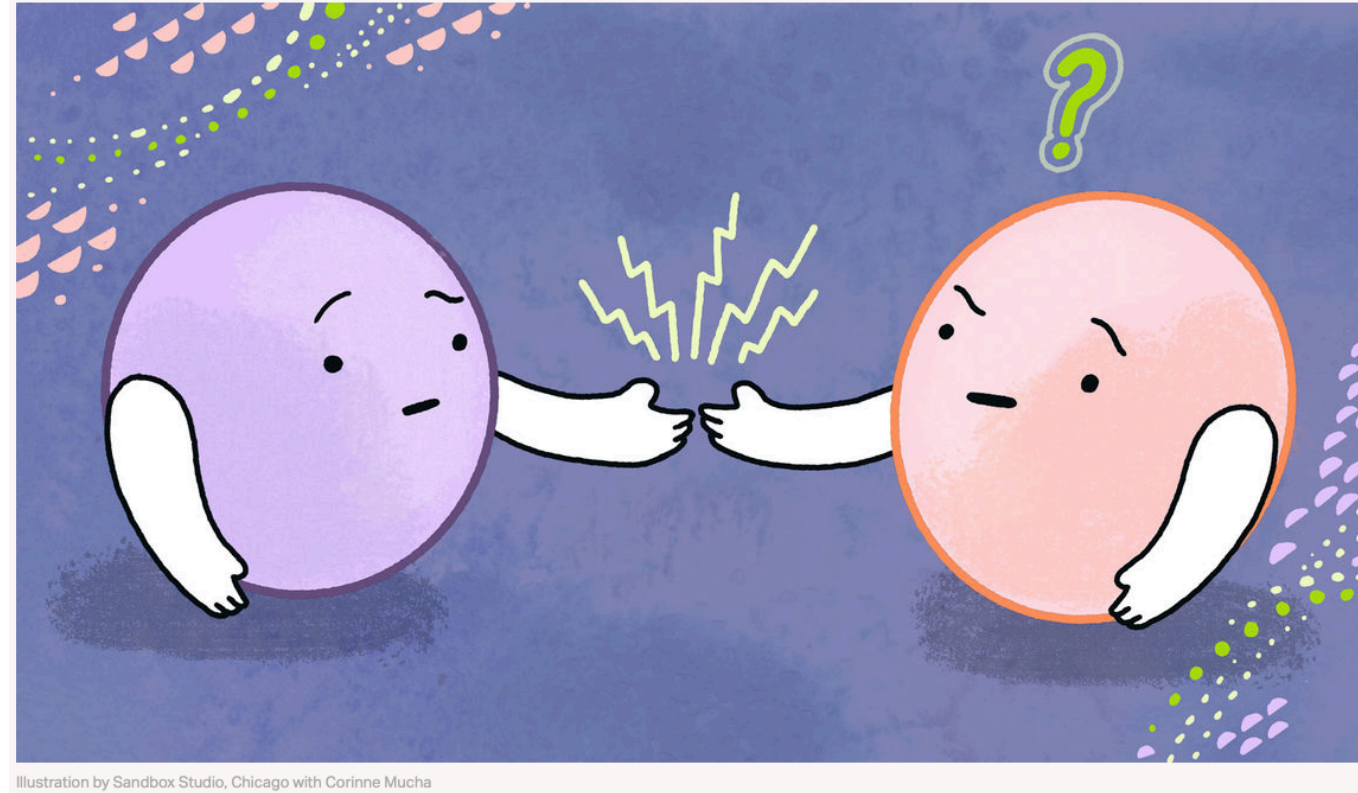


See J. Menendez's talk

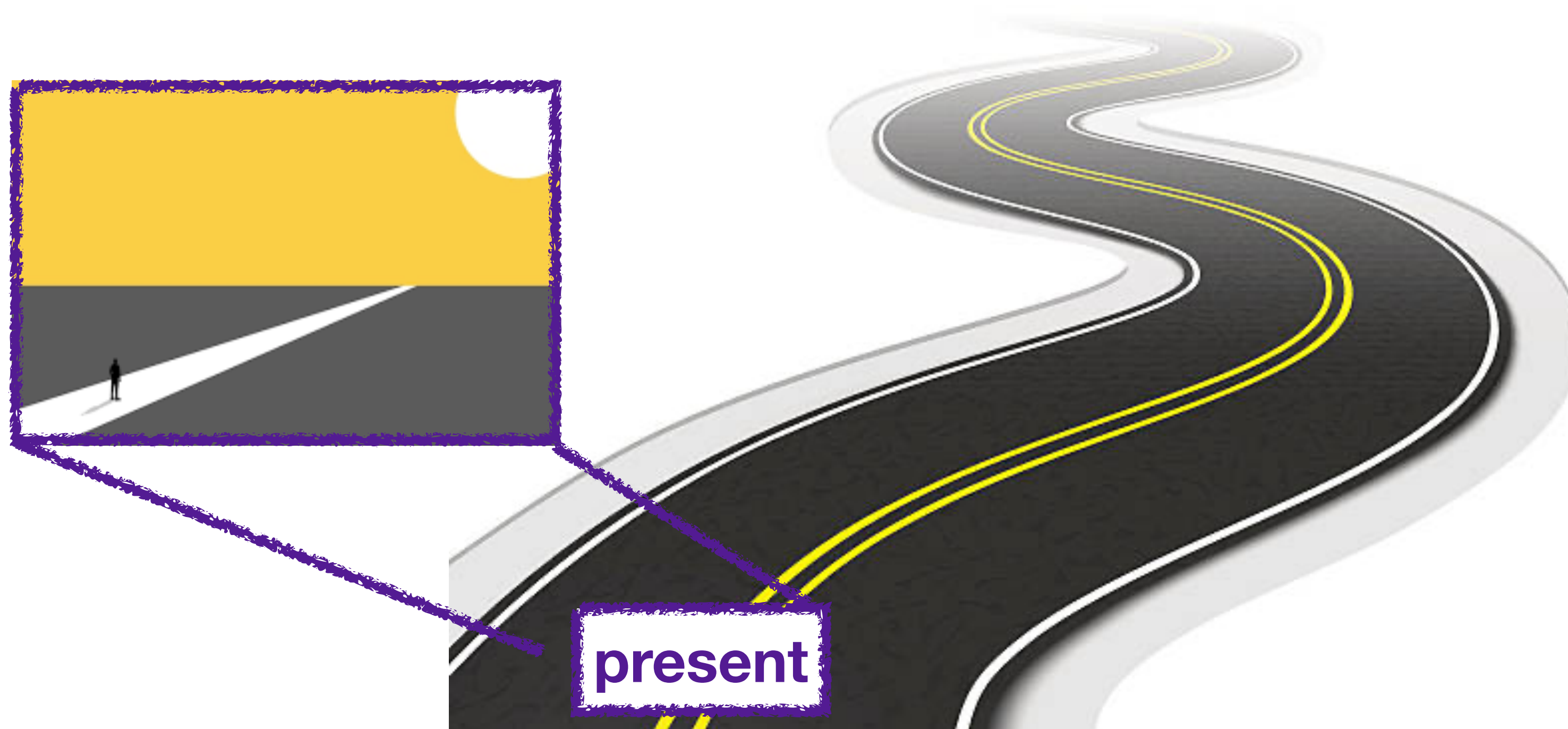
Importance of searching for $0\nu\beta\beta$



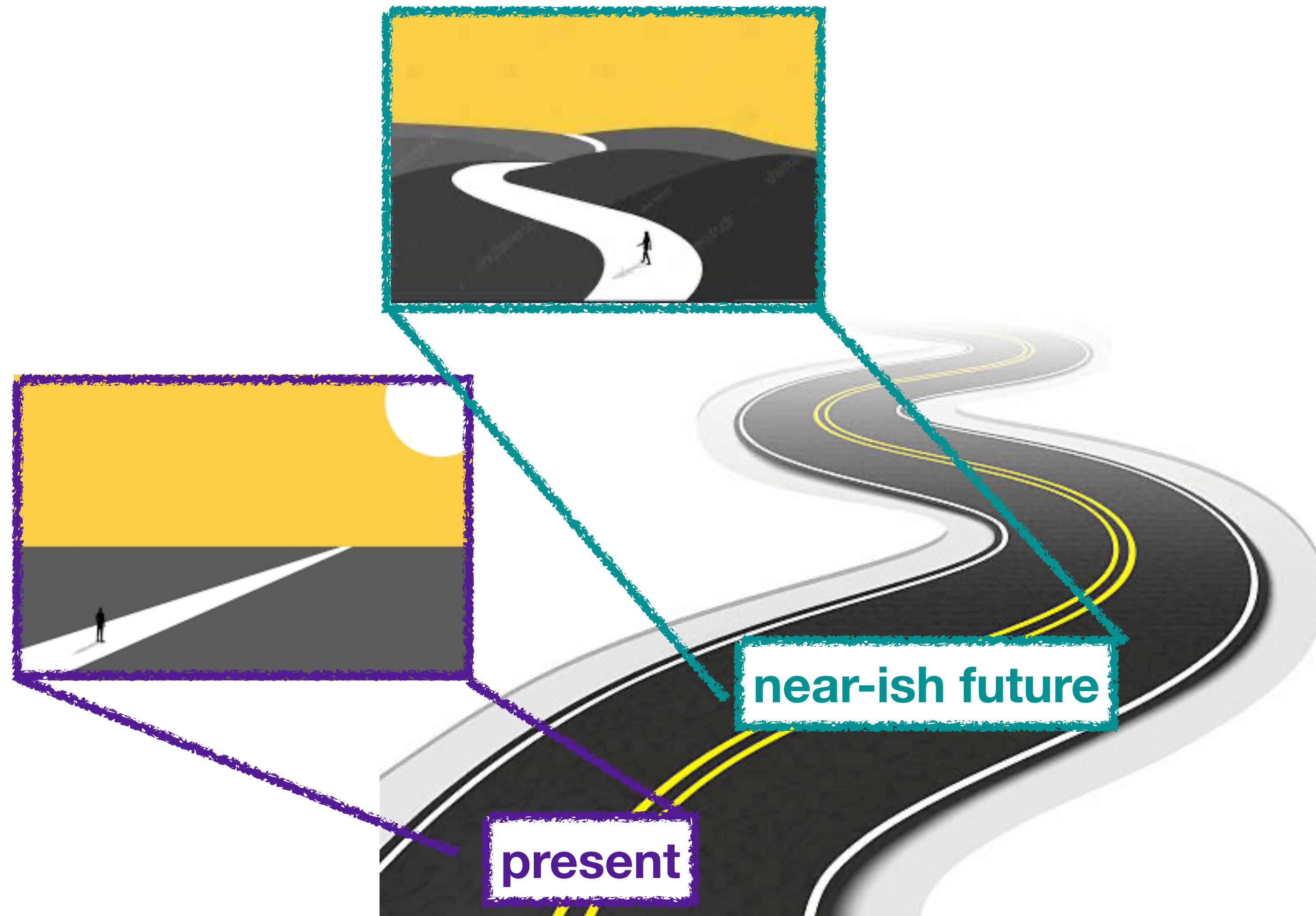
The road to a Majorana neutrino



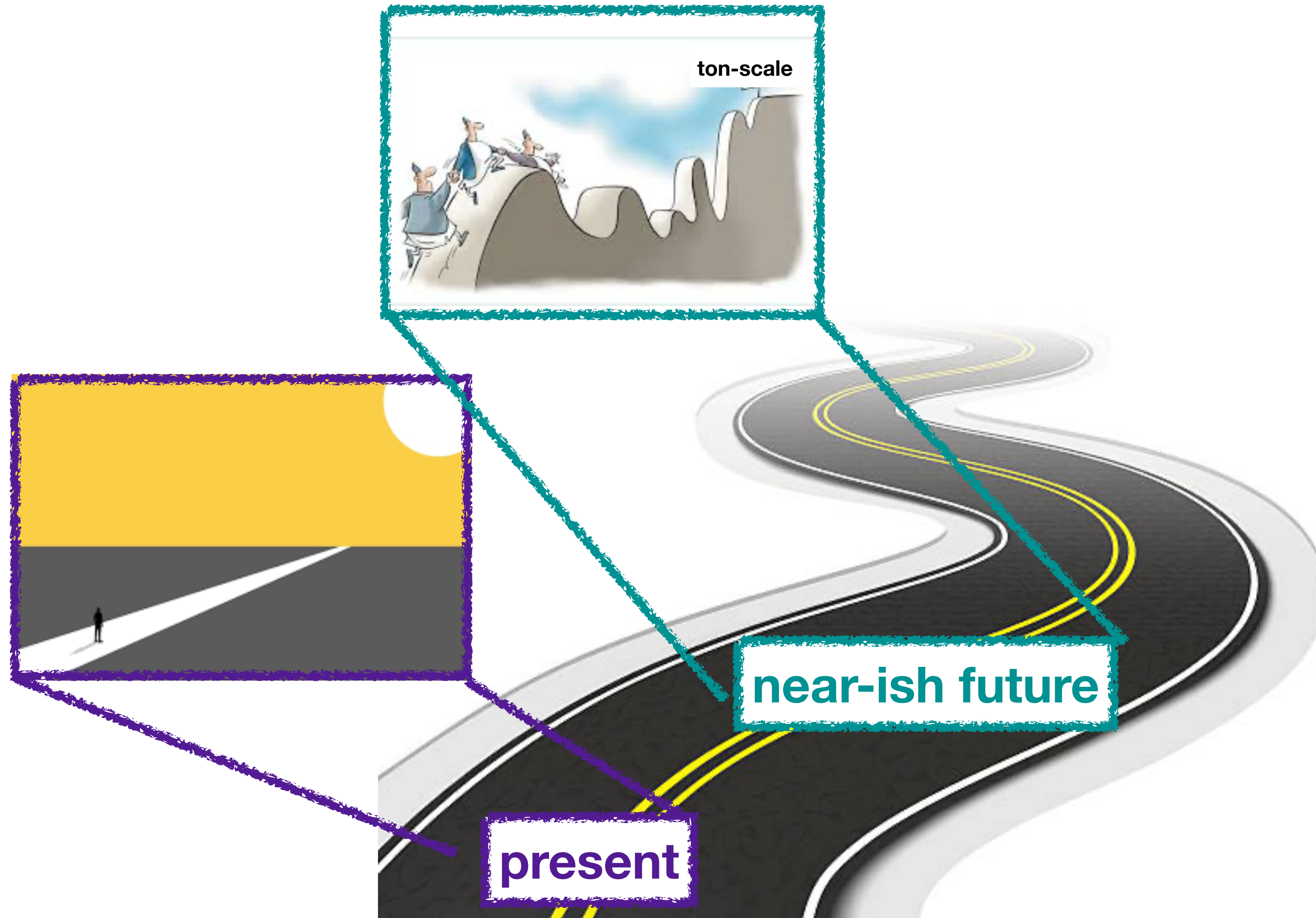
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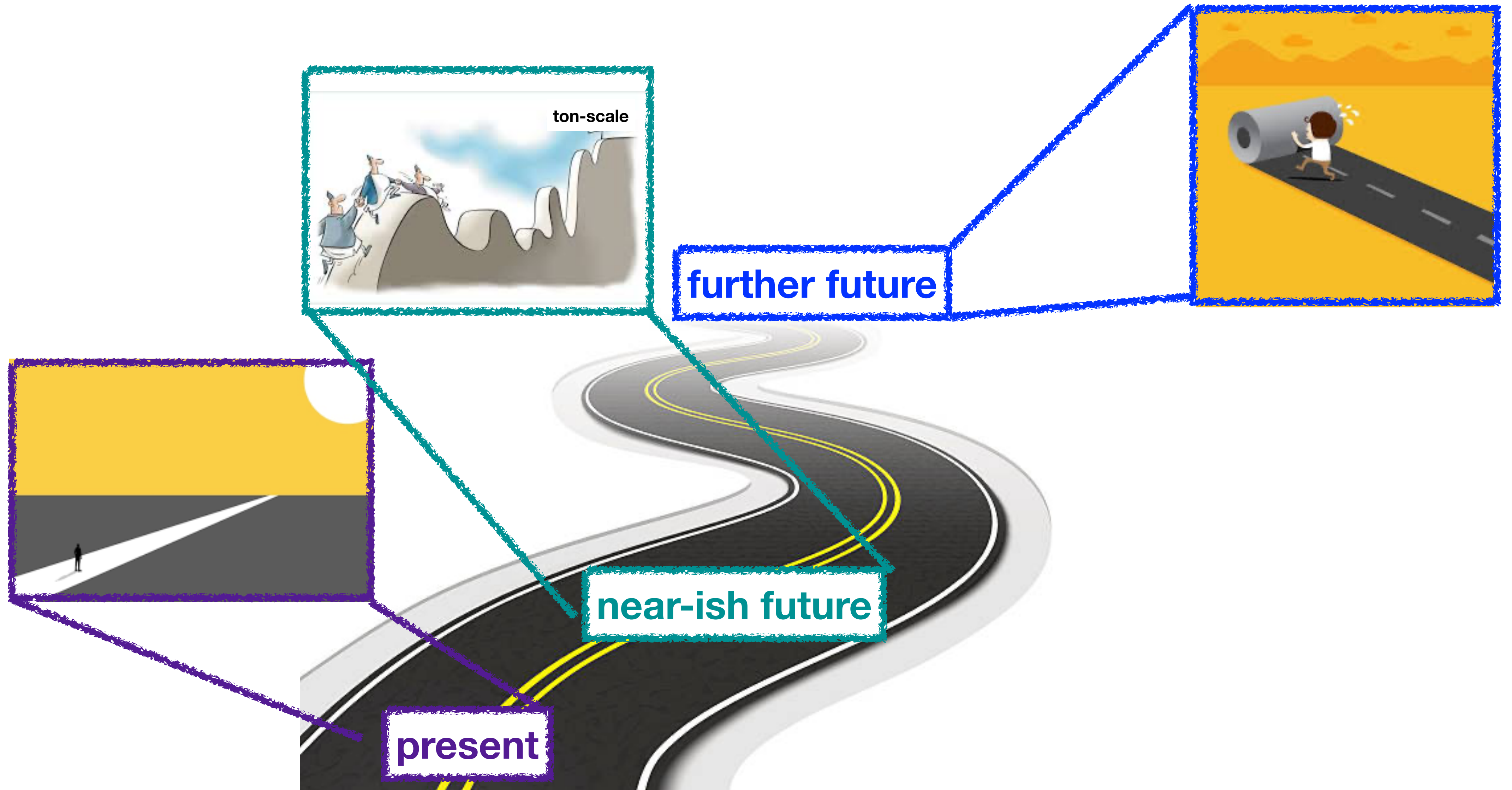
The road to a Majorana neutrino



The road to a Majorana neutrino



The road to a Majorana neutrino



The road to a Majorana neutrino



present

How to search for $0\nu\beta\beta$

- Isotope choice:
 - Highest $Q_{\beta\beta}$ value (lower backgrounds)
 - Highest abundance (lower cost)

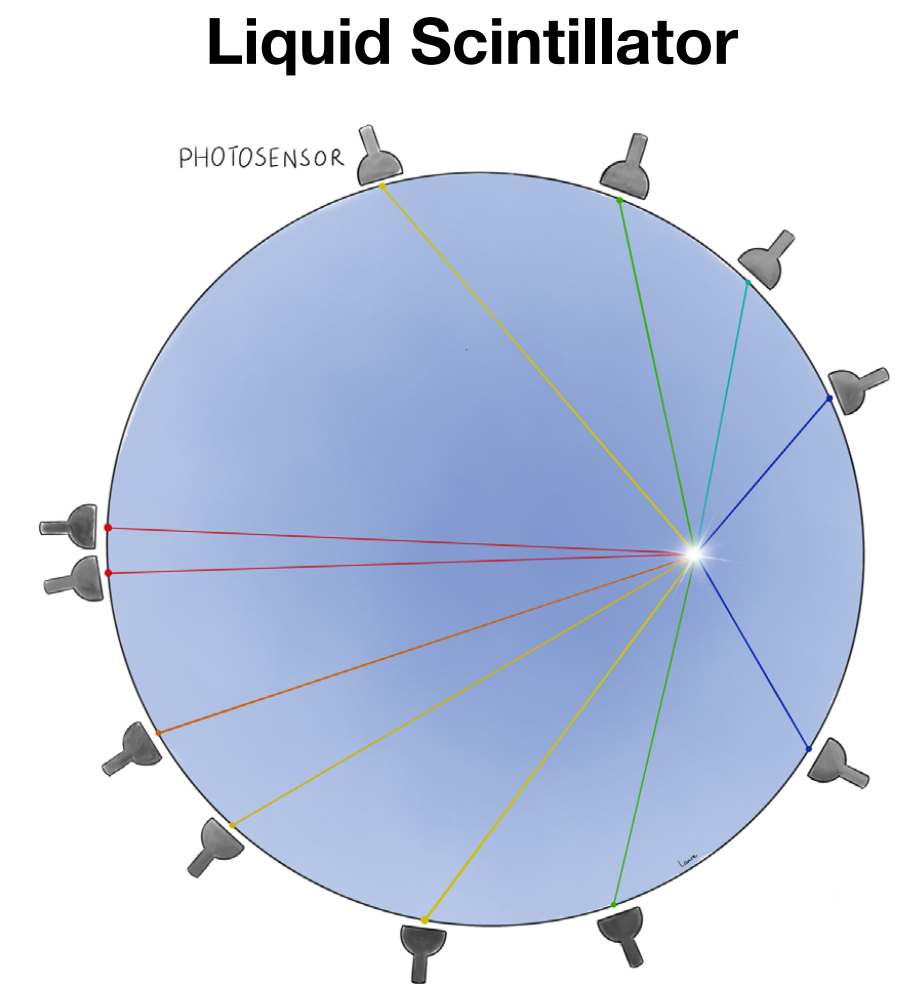
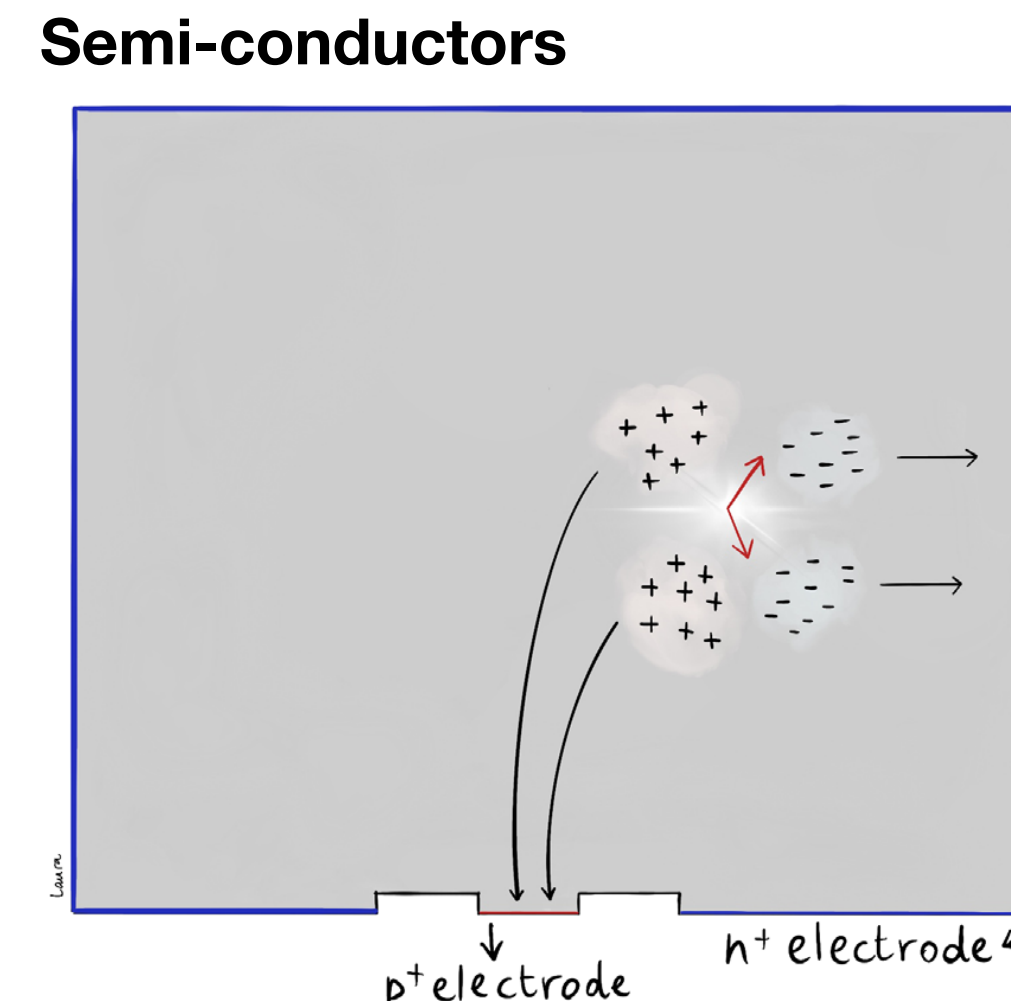
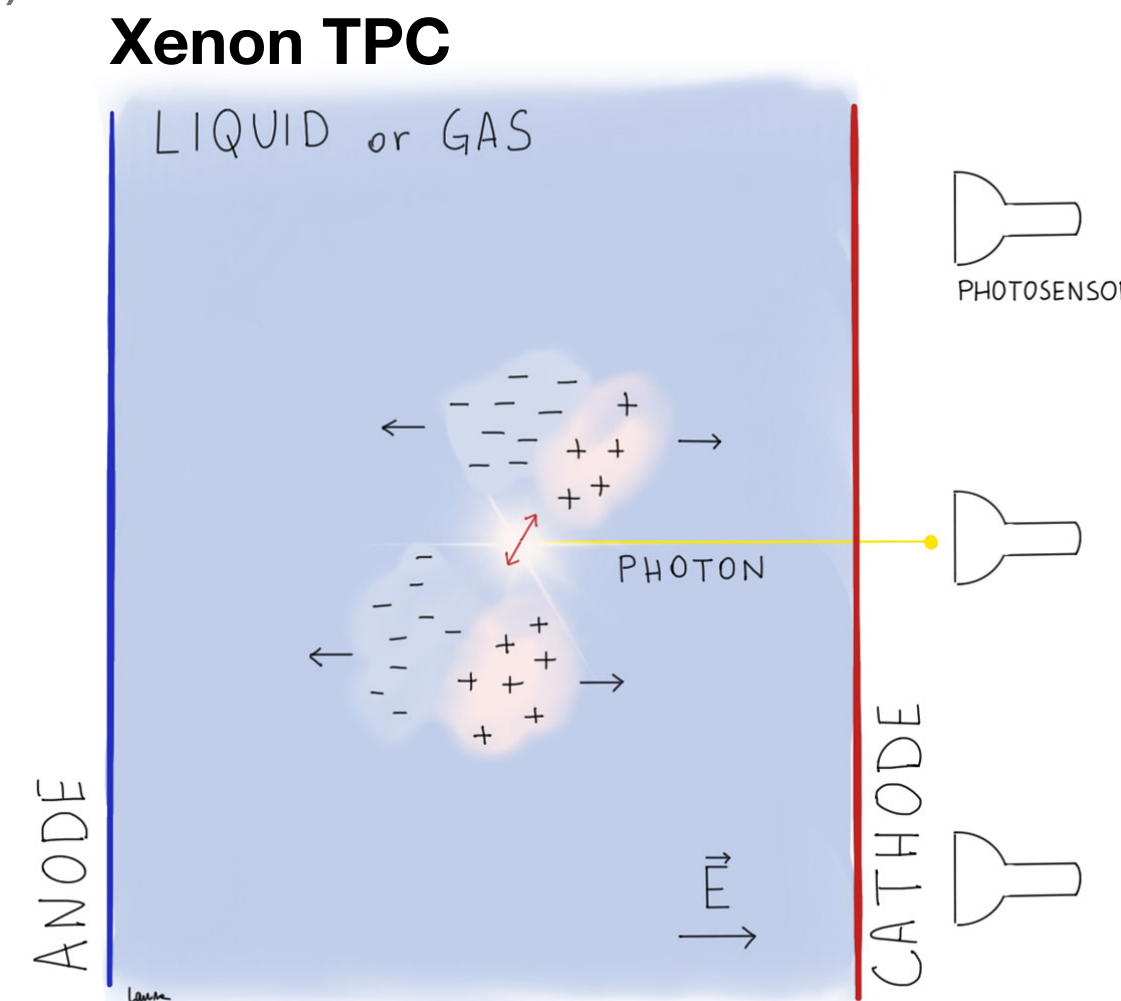
Isotope	Daughter	$Q_{\beta\beta}^a$ [keV]	f_{nat}^b [%]
^{48}Ca	^{48}Ti	4 267.98(32)	0.187(21)
^{76}Ge	^{76}Se	2 039.061(7)	7.75(12)
^{82}Se	^{82}Kr	2 997.9(3)	8.82(15)
^{96}Zr	^{96}Mo	3 356.097(86)	2.80(2)
^{100}Mo	^{100}Ru	3 034.40(17)	9.744(65)
^{116}Cd	^{116}Sn	2 813.50(13)	7.512(54)
^{130}Te	^{130}Xe	2 527.518(13)	34.08(62)
^{136}Xe	^{136}Ba	2 457.83(37)	8.857(72)
^{150}Nd	^{150}Sm	3 371.38(20)	5.638(28)

[M. Agostini, et al., \(2023\) Rev. Mod. Phys. 95, 025002](#)

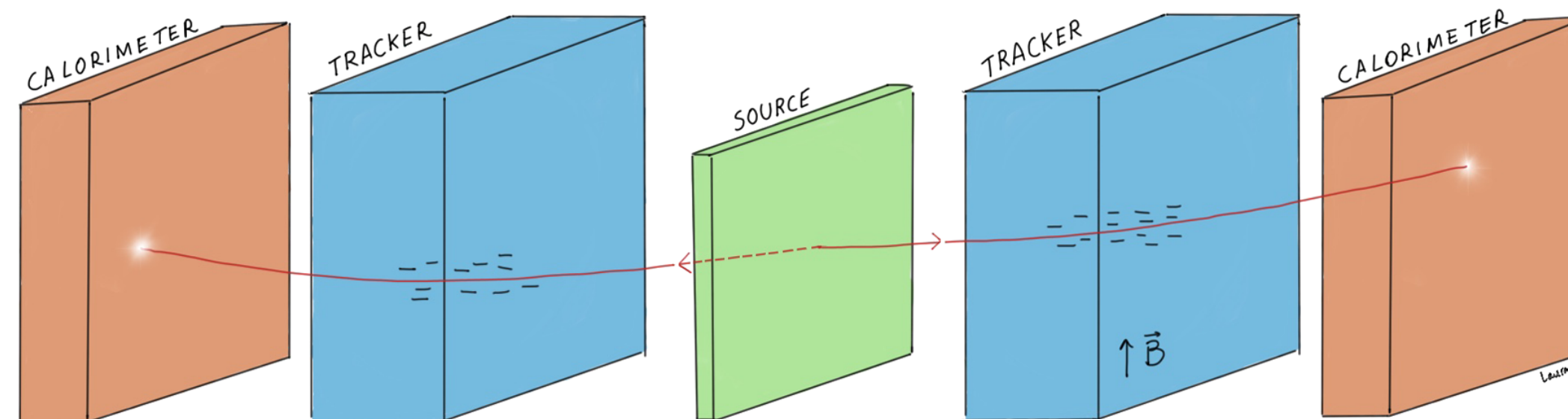
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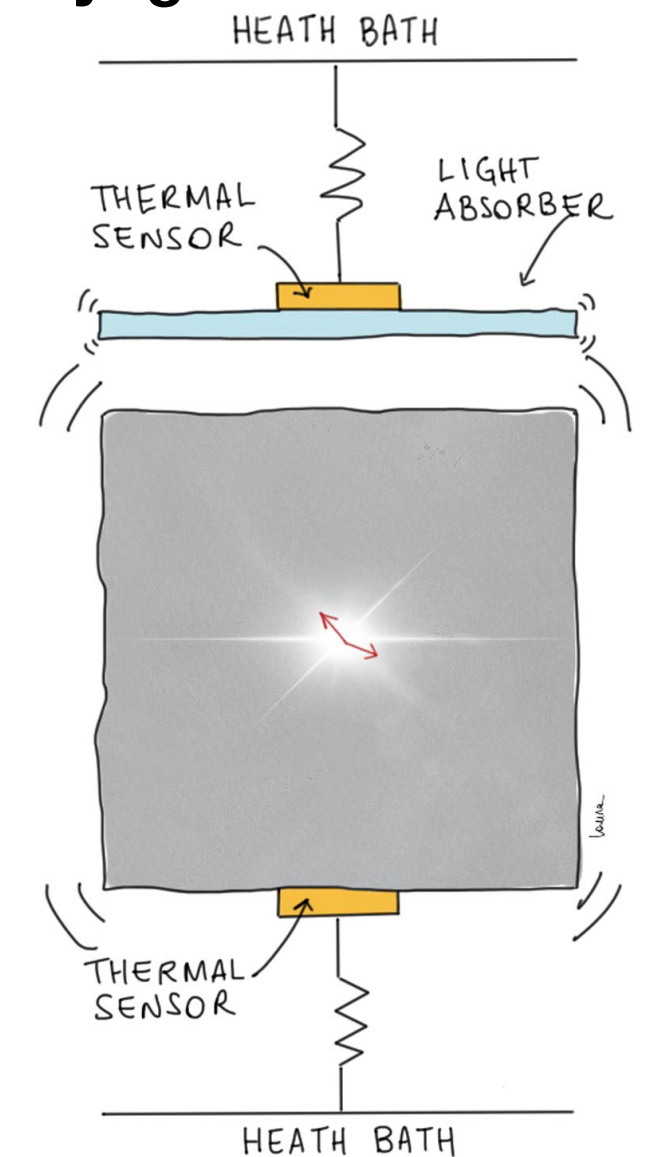
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Tracking calorimeters



Cryogenic bolometers



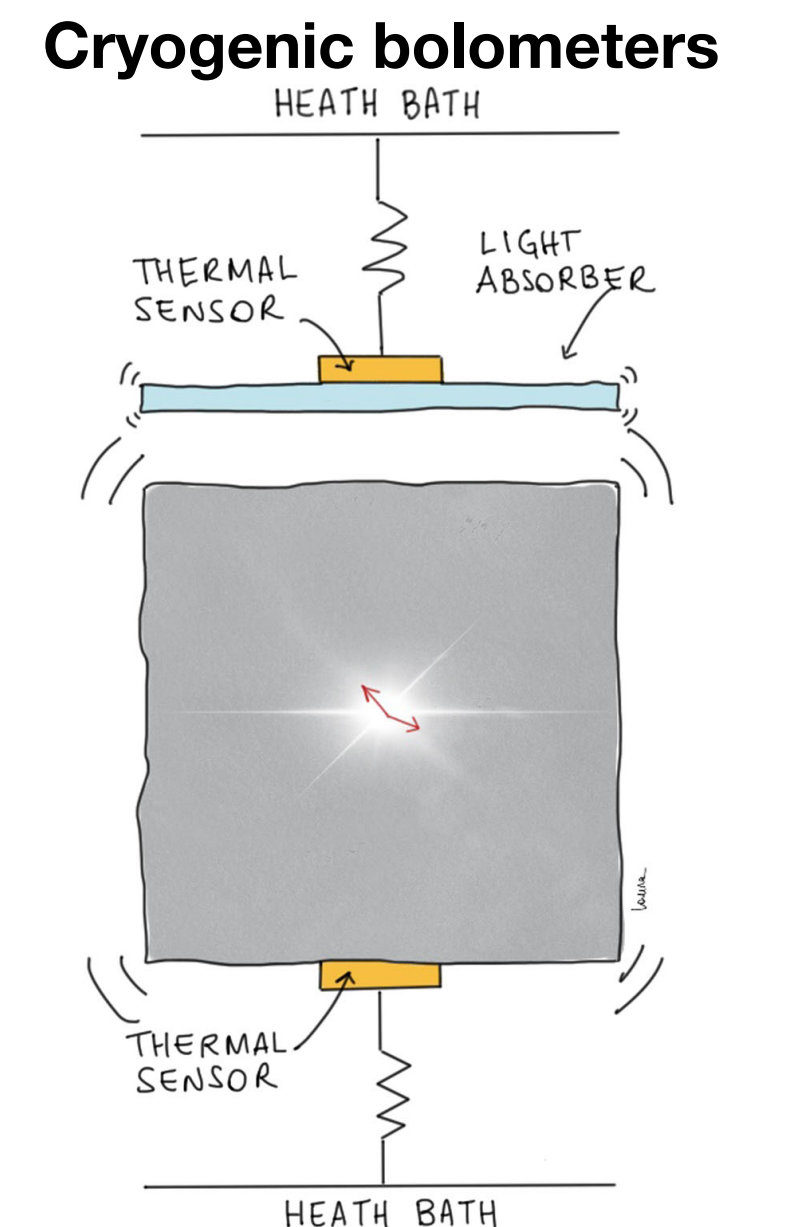
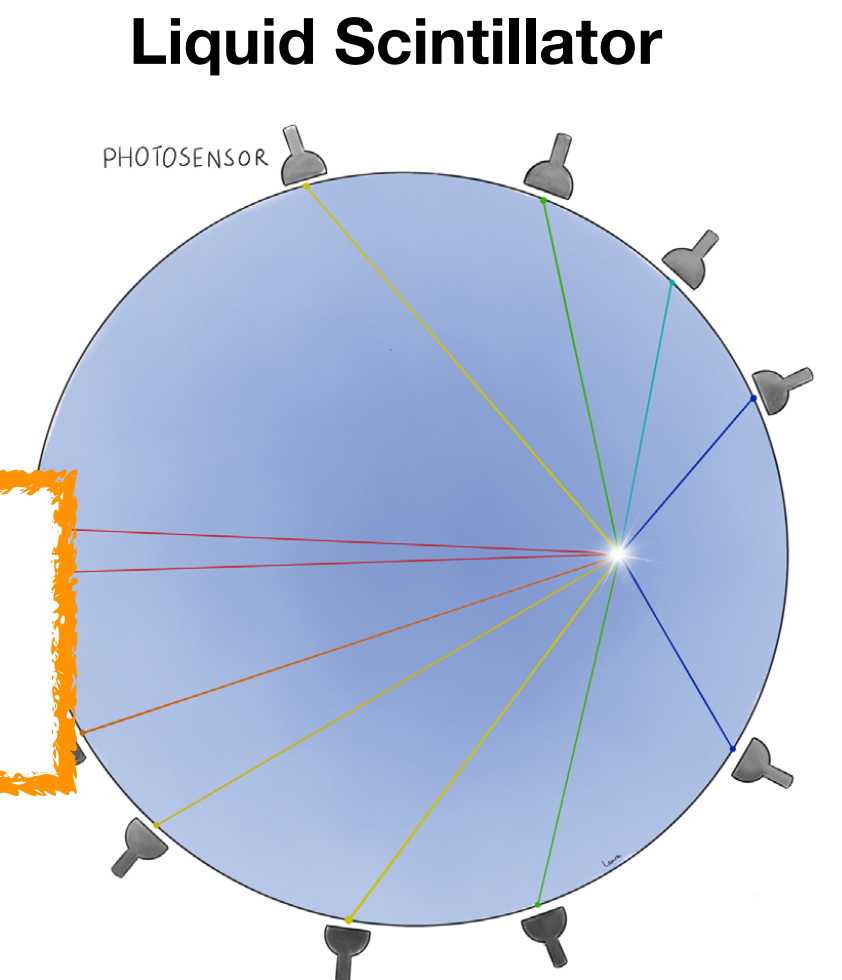
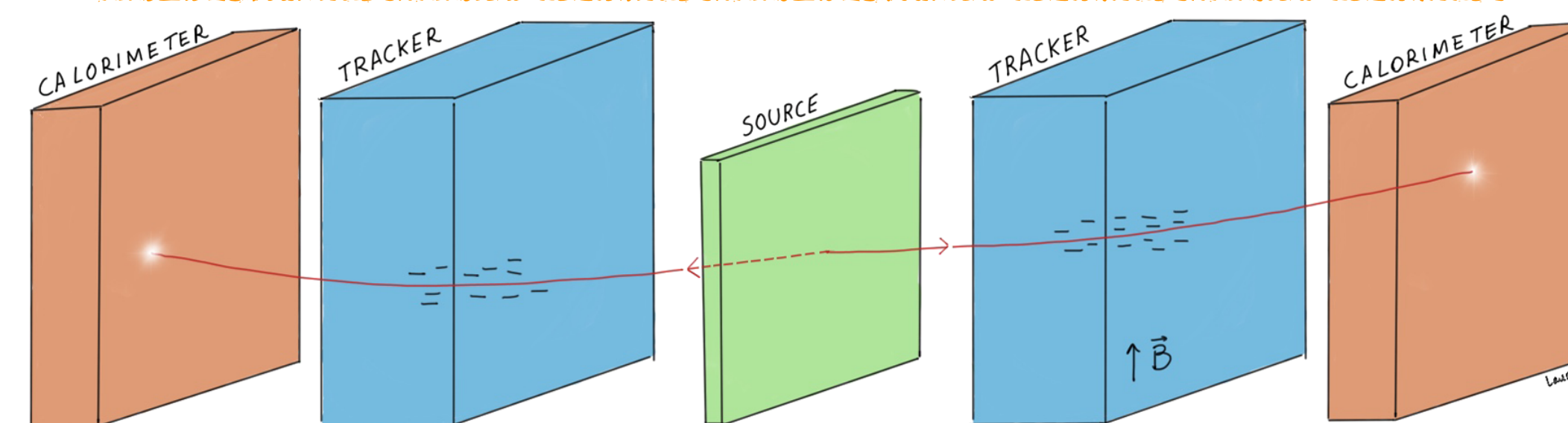
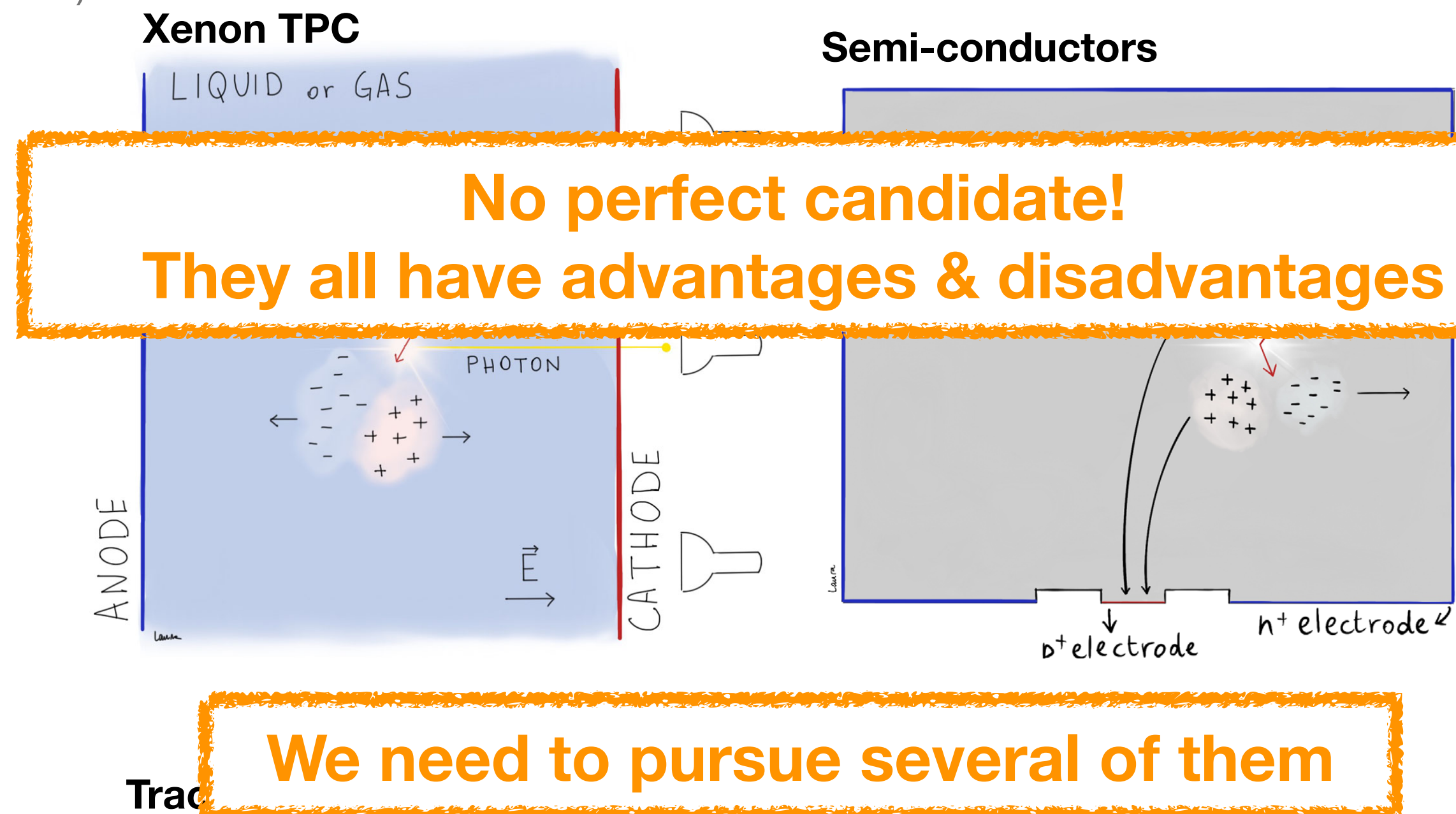
- Detector technology choice:
 - Best energy resolution
 - Lowest backgrounds
 - Most scalable

How to search for $0\nu\beta\beta$

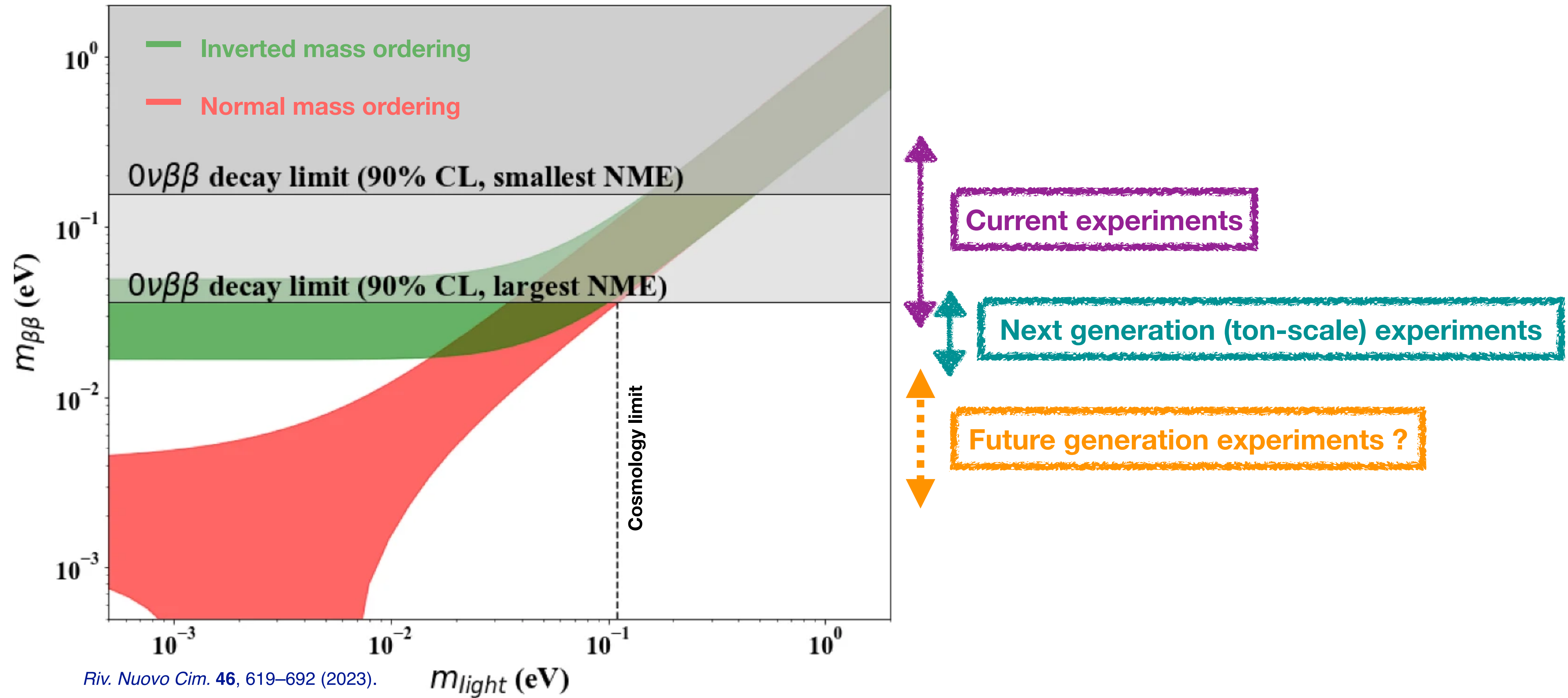
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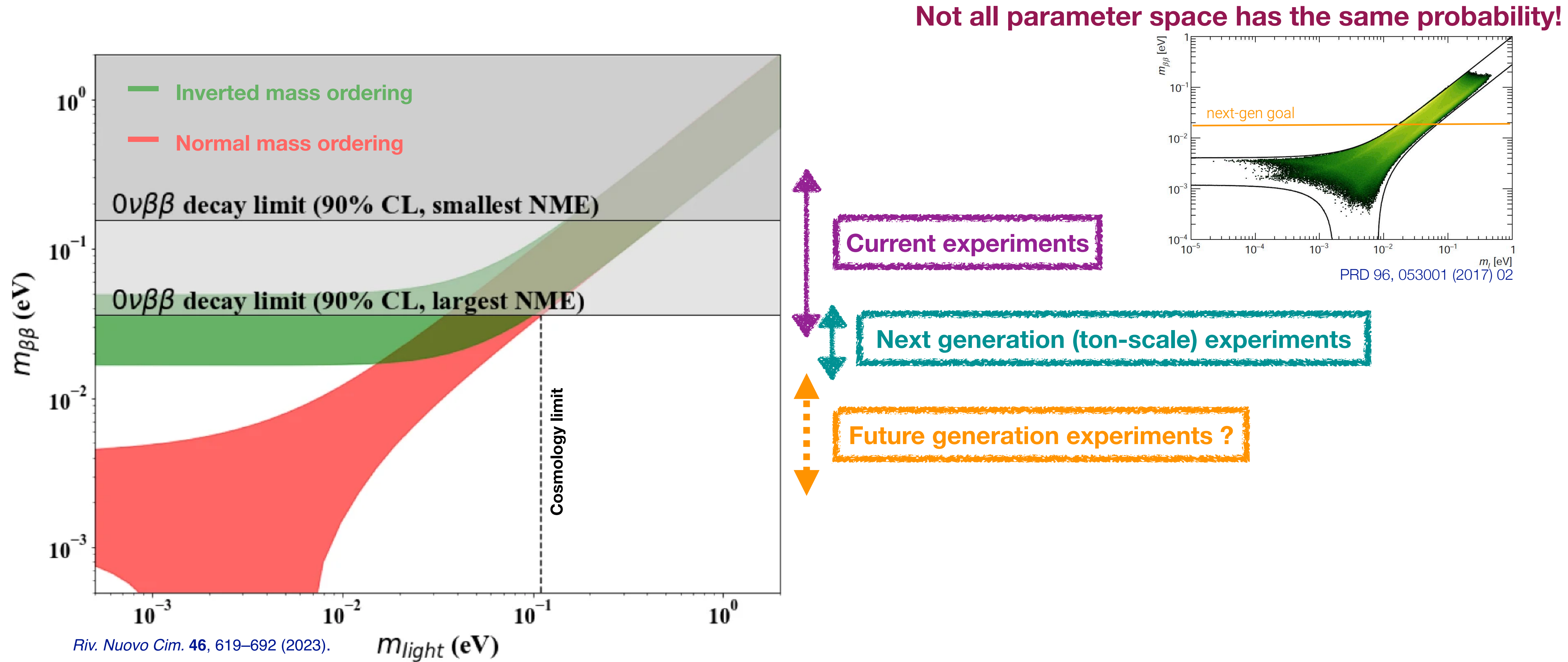
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 - Best energy resolution
 - Lowest backgrounds
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Where to search for $0\nu\beta\beta$



Where to search for $0\nu\beta\beta$



**** This plot is just one example of parameter space... It does not mean that is exactly where the Majorana neutrino is! ****

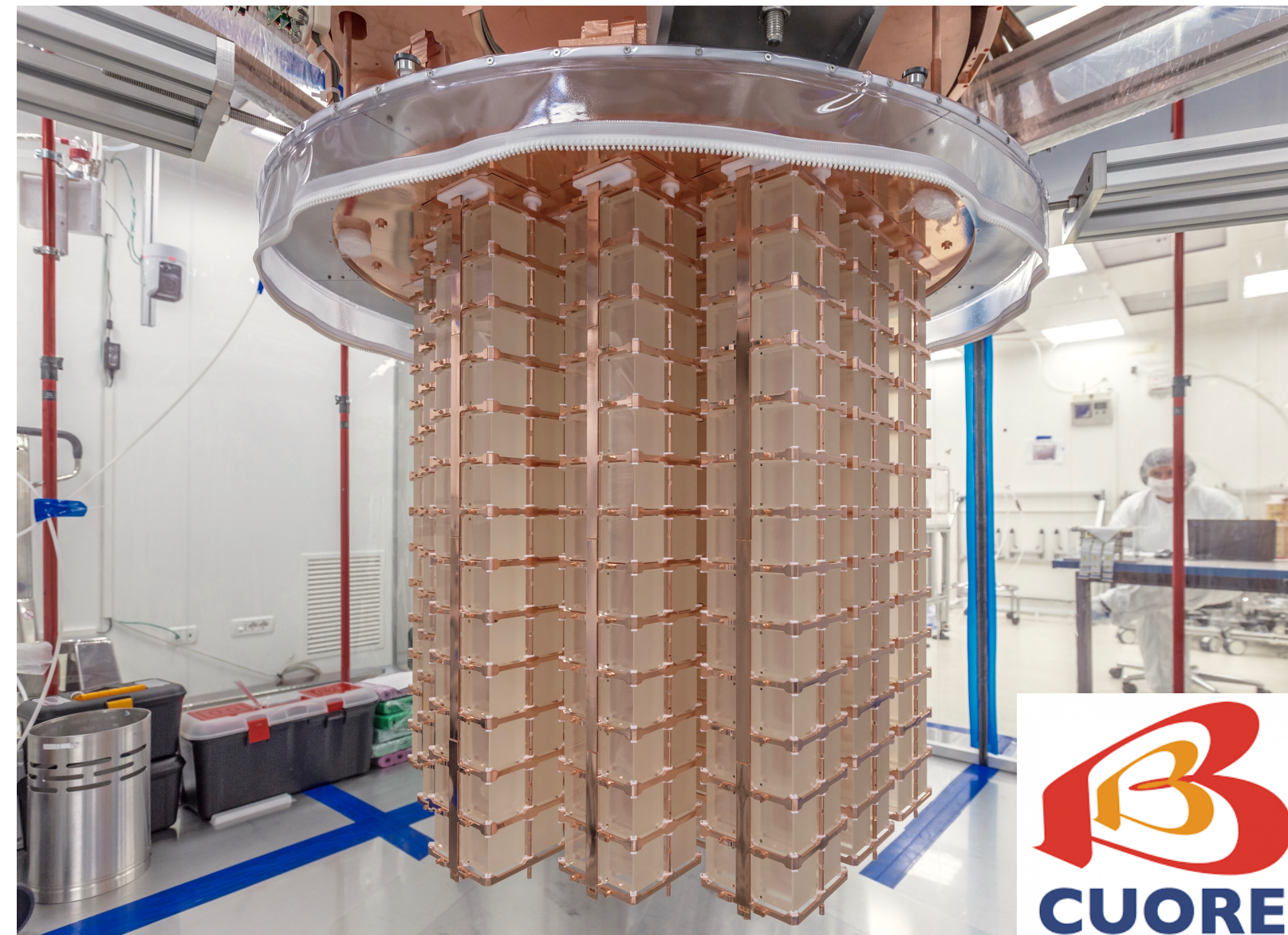
Current running experiments

KamLAND-Zen



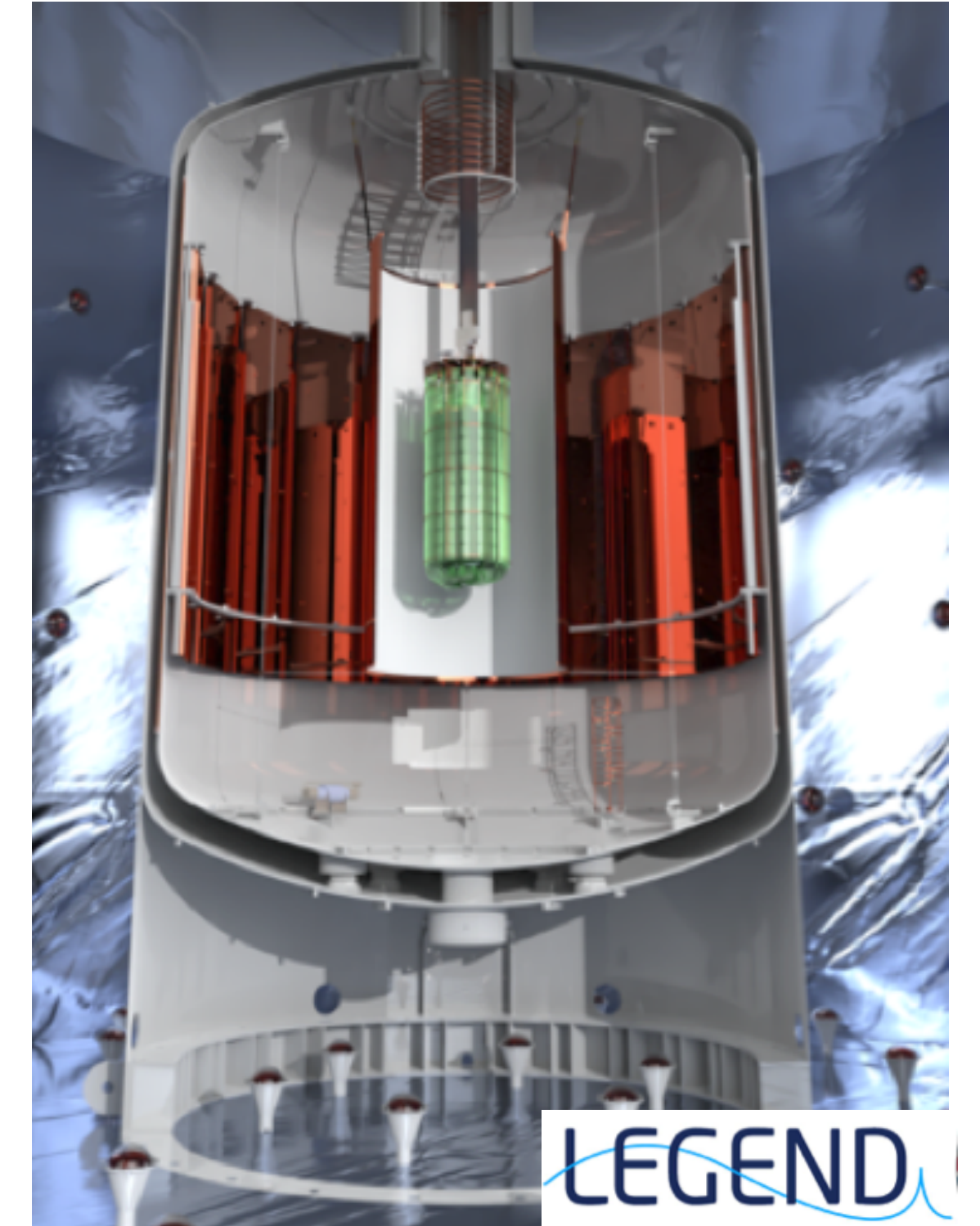
See I. Shimizu's talk

CUORE



See C. Bucci's talk

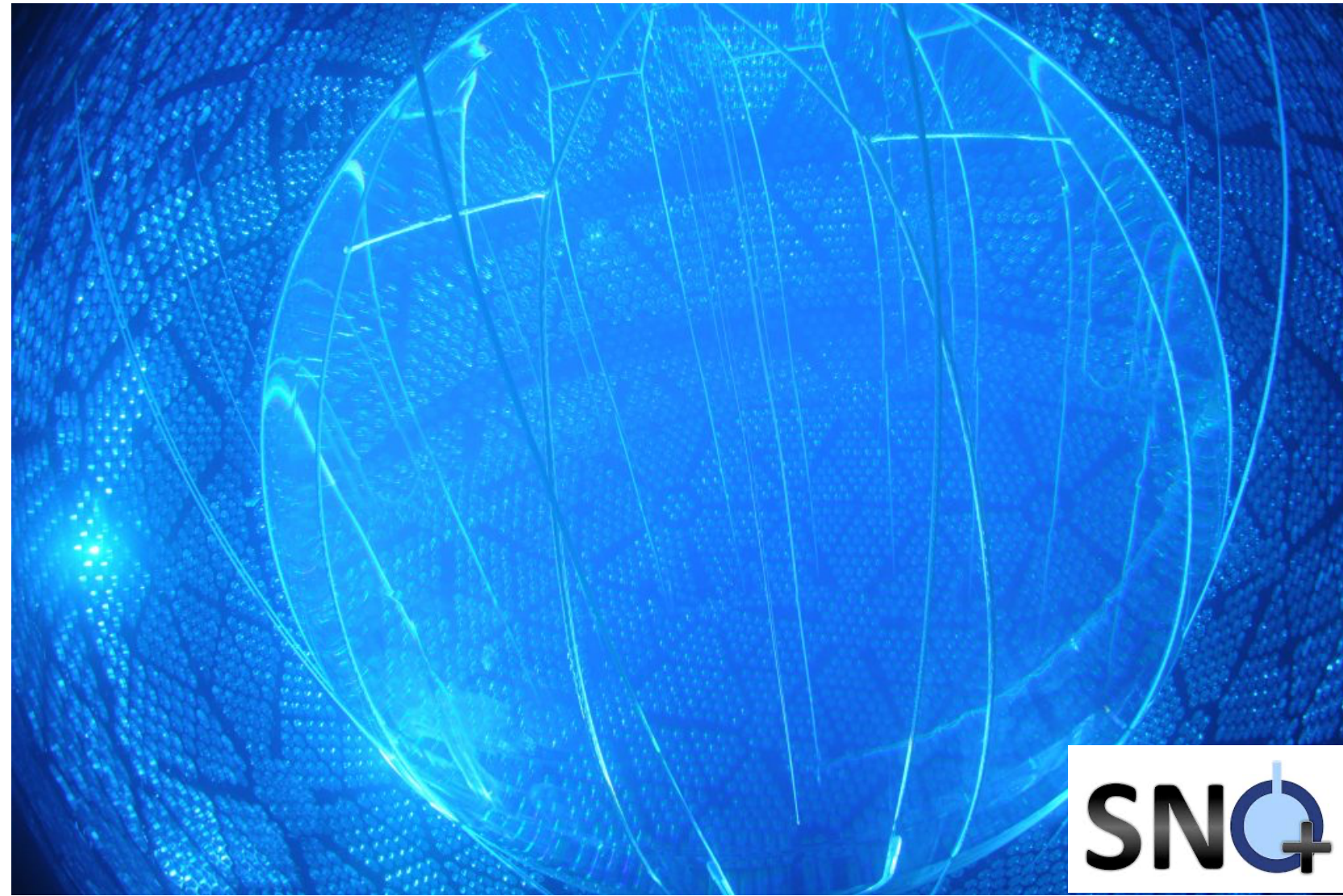
LEGEND-200



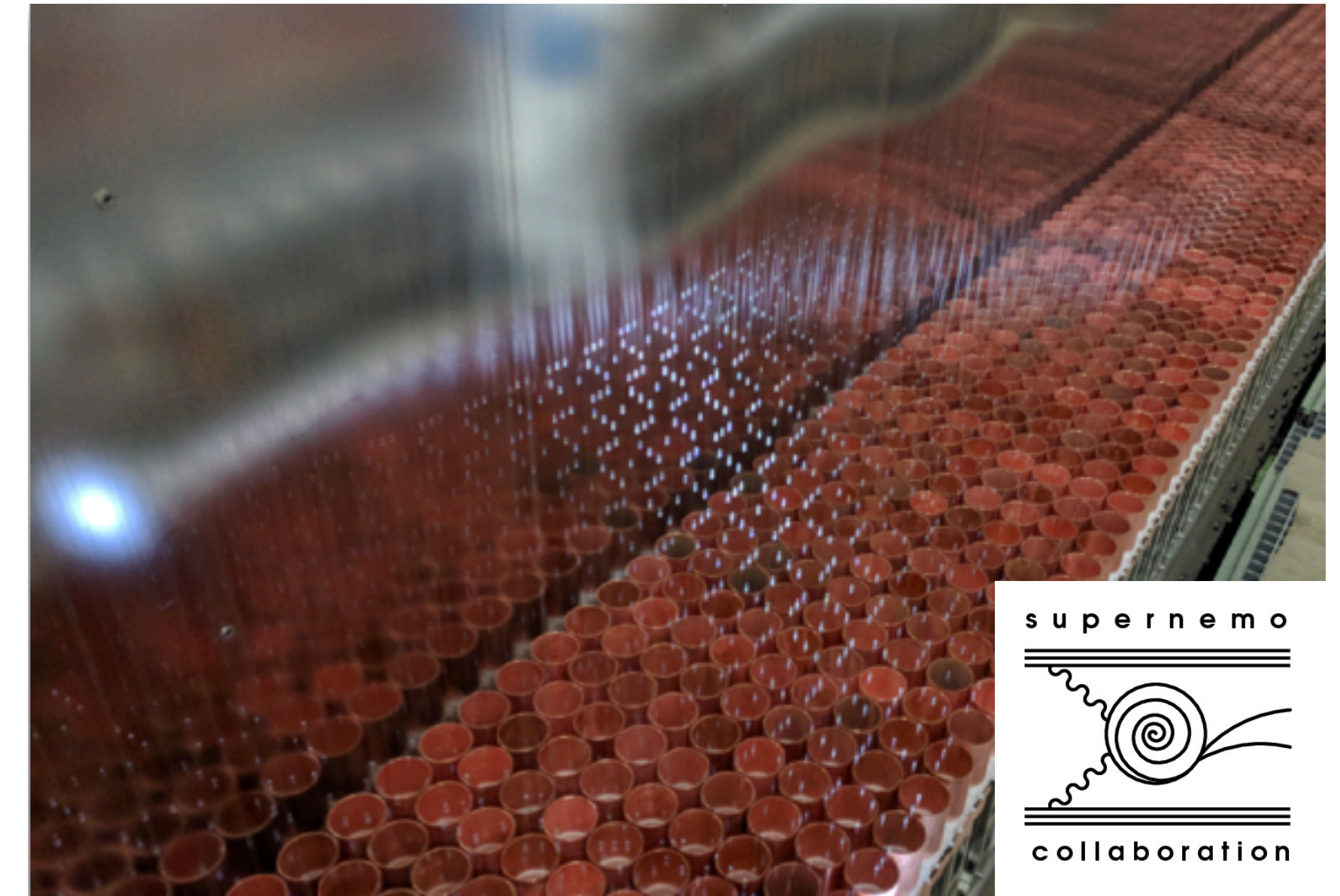
See L. Pertoldi's talk

Current (*almost*) running experiments

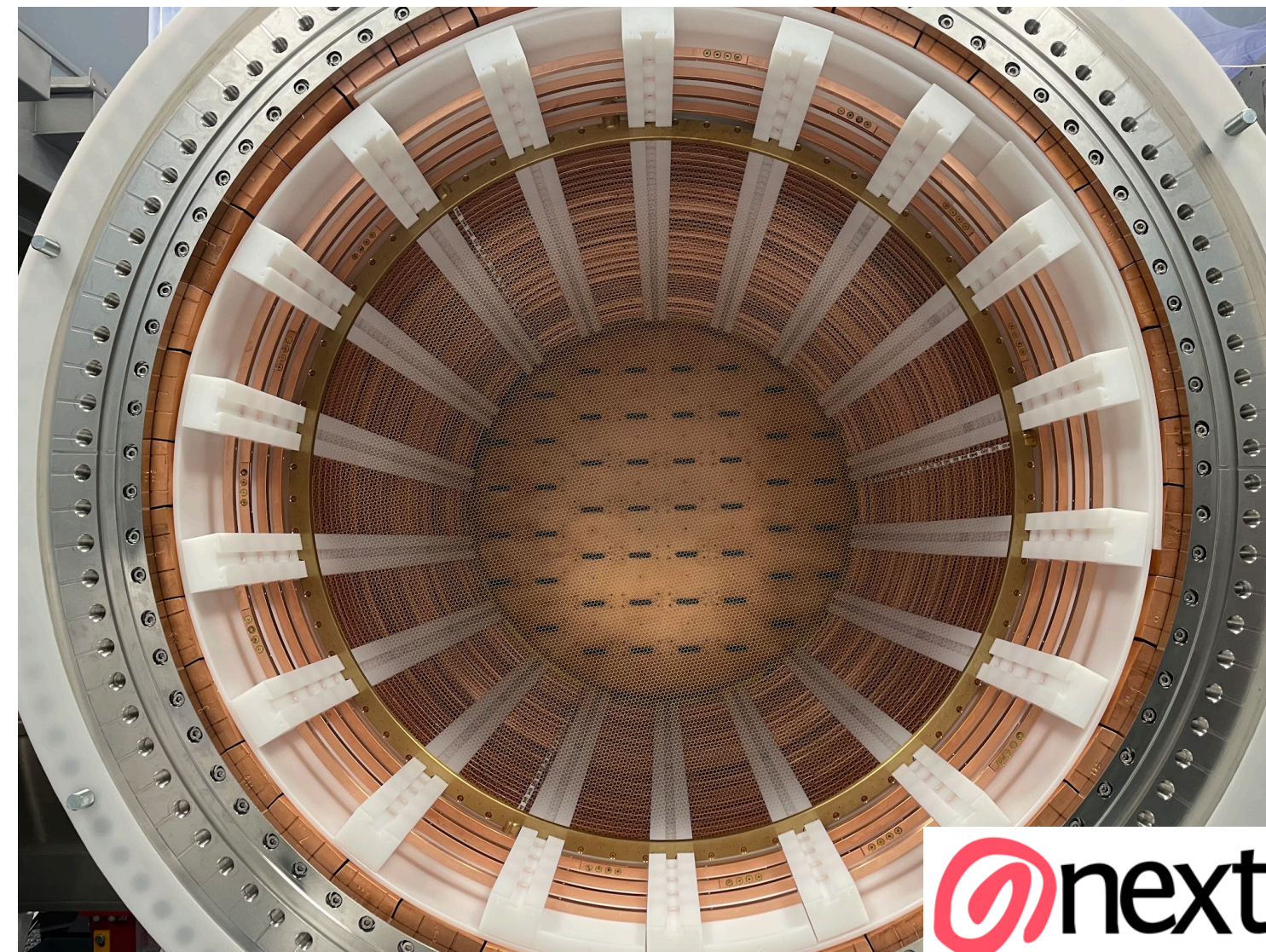
SNO+



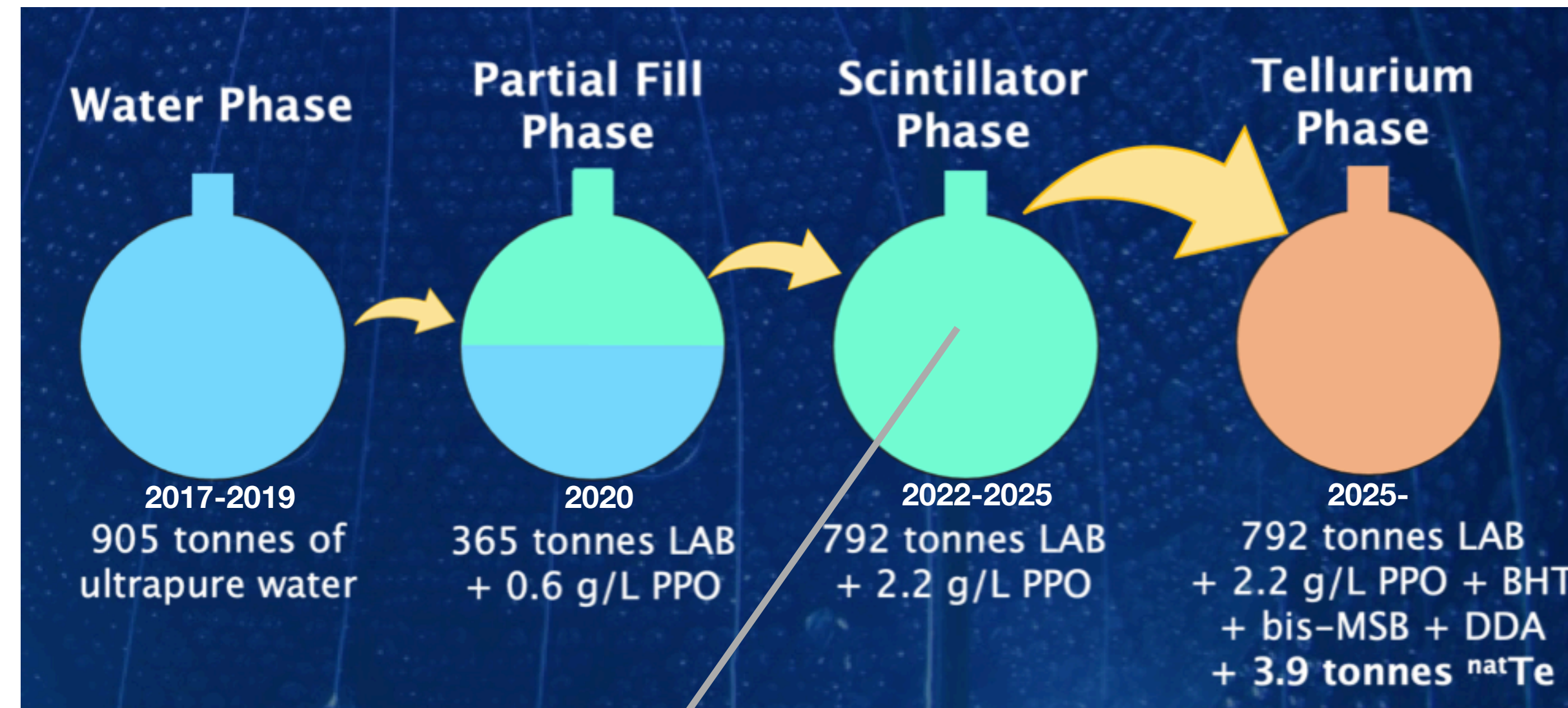
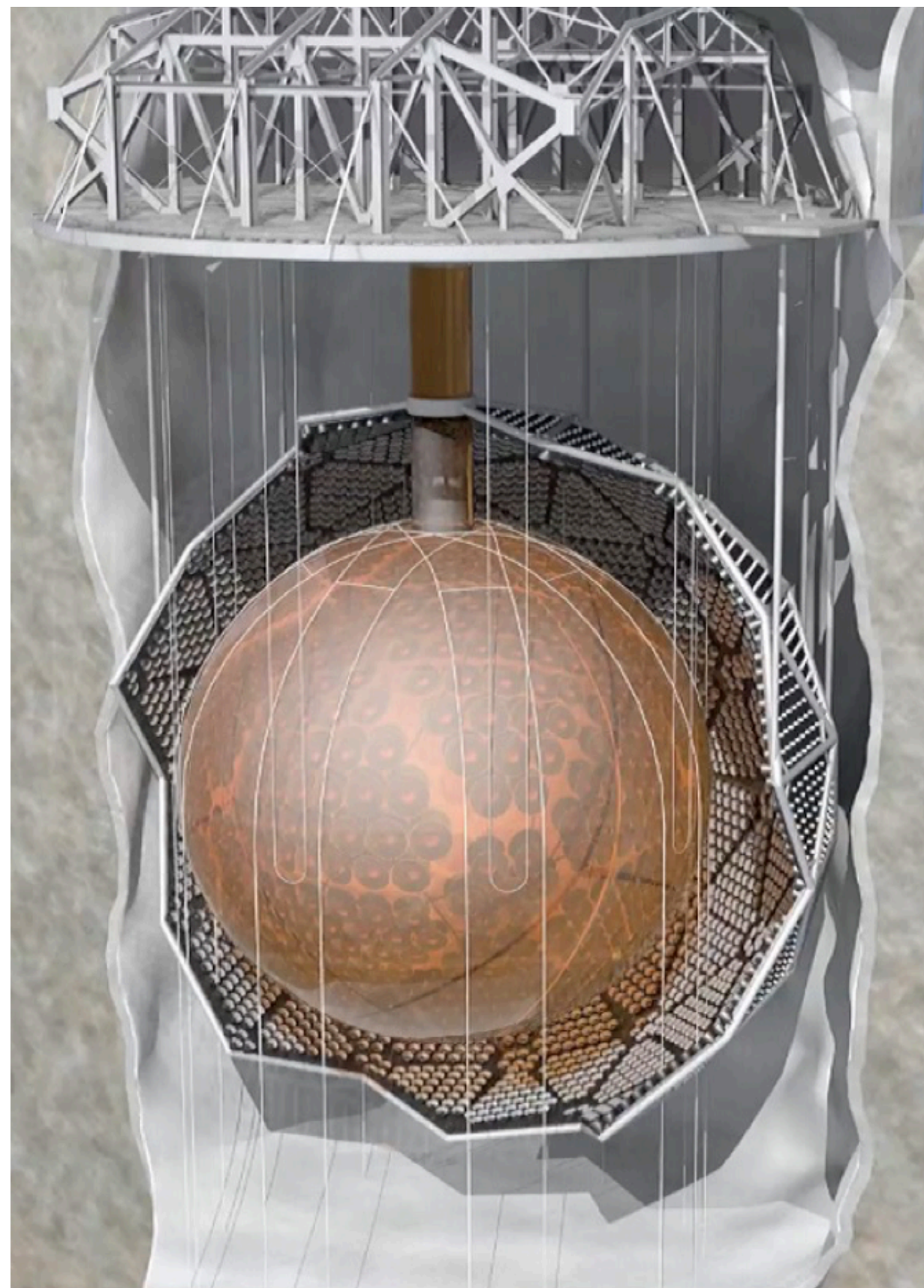
SUPERNEMO



NEXT-100

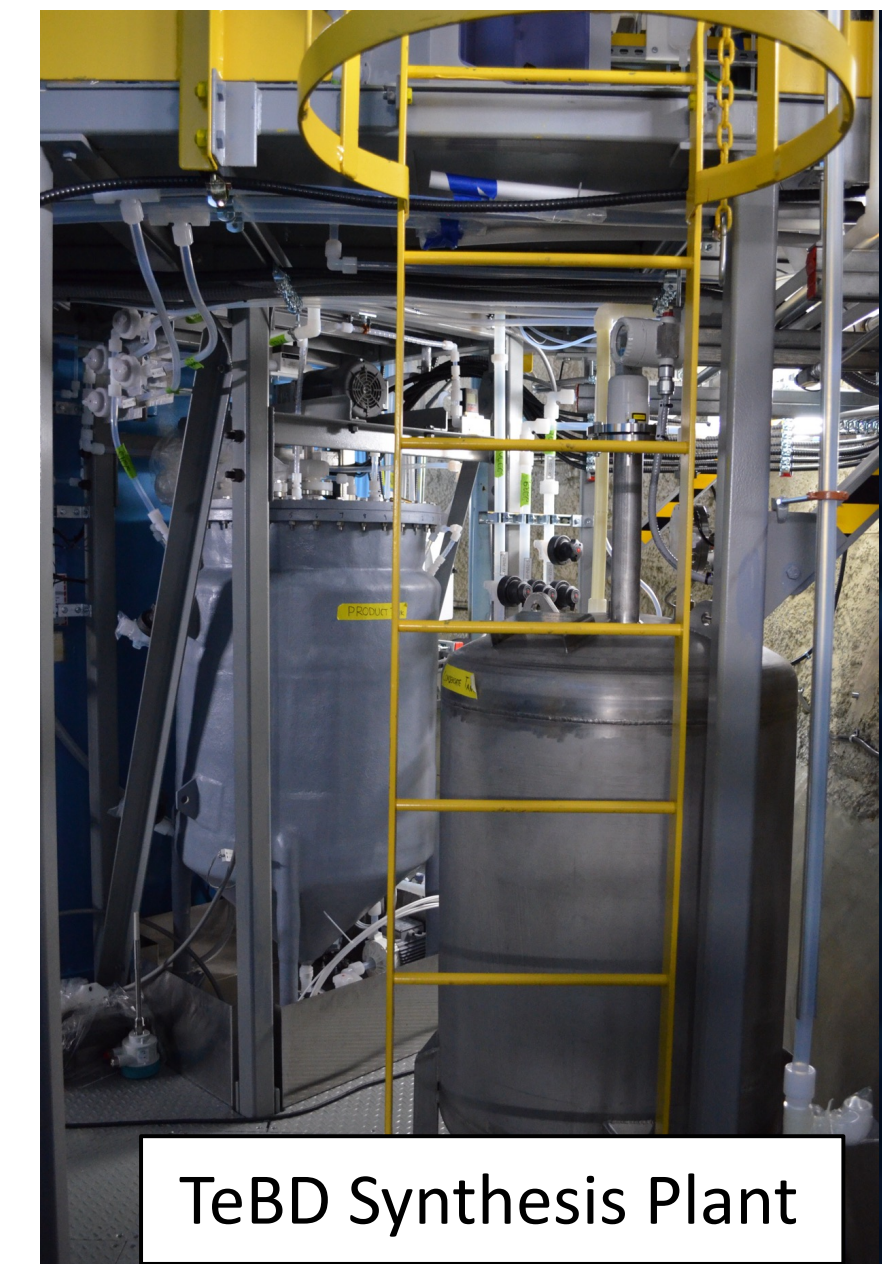


- Isotope of ^{130}Te → highest abundance, good $Q_{\beta\beta}$ (2.53MeV)



Great advantage to study backgrounds without the target

Te loading:
 Phase 1: 0.5%
 Phase 2: 1.5%
 Phase 3: 3%



TeBD Synthesis Plant under commissioning

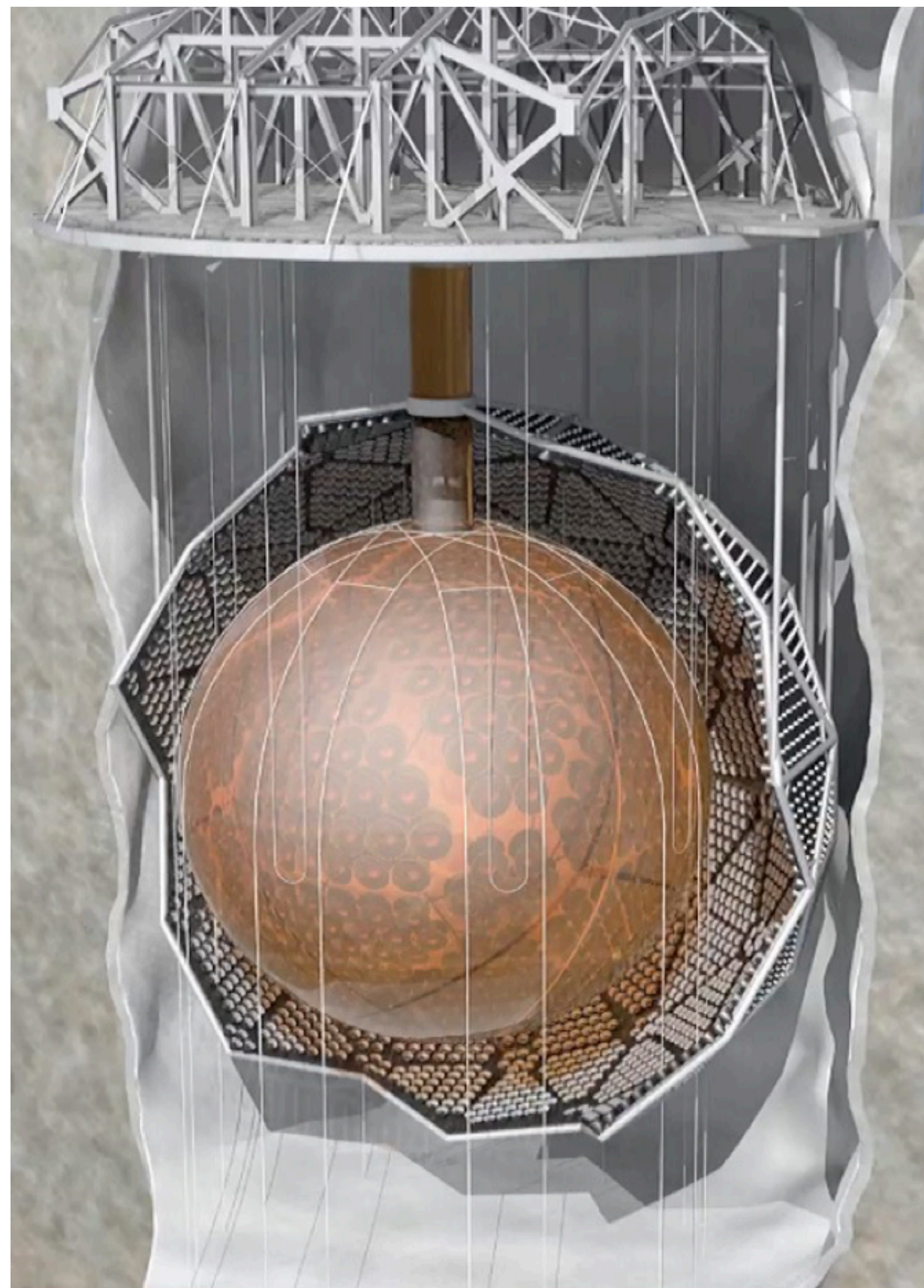
Poster 581:

The SNO+ Tellurium Deployment Programme

Thanks to J. Maneira, S.Biller and M.Chen for input

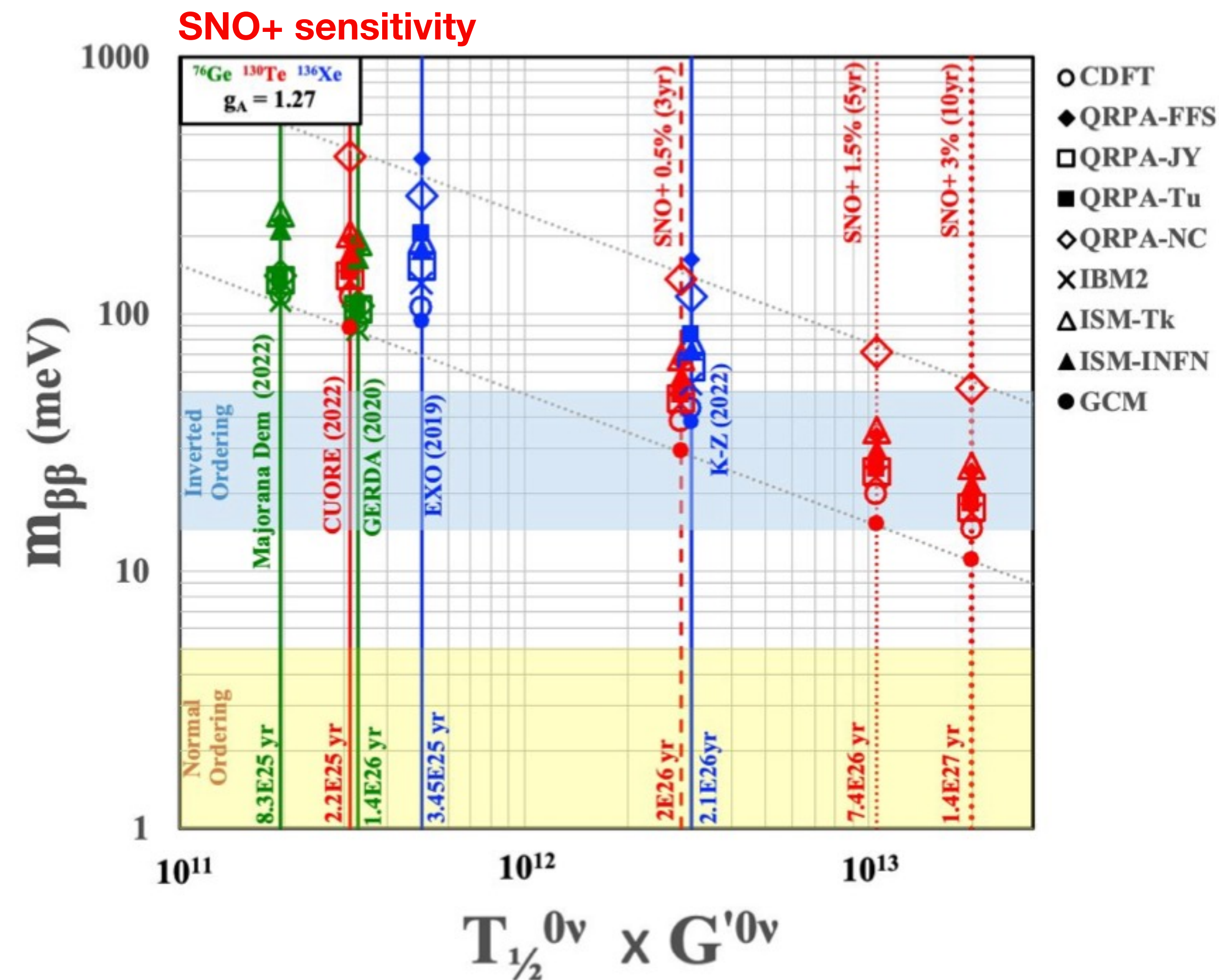
See J. Maneira's talk on Thursday for more details on SNO+ capabilities

- Isotope of ^{130}Te \rightarrow highest abundance, good $Q_{\beta\beta}$ (2.53MeV)

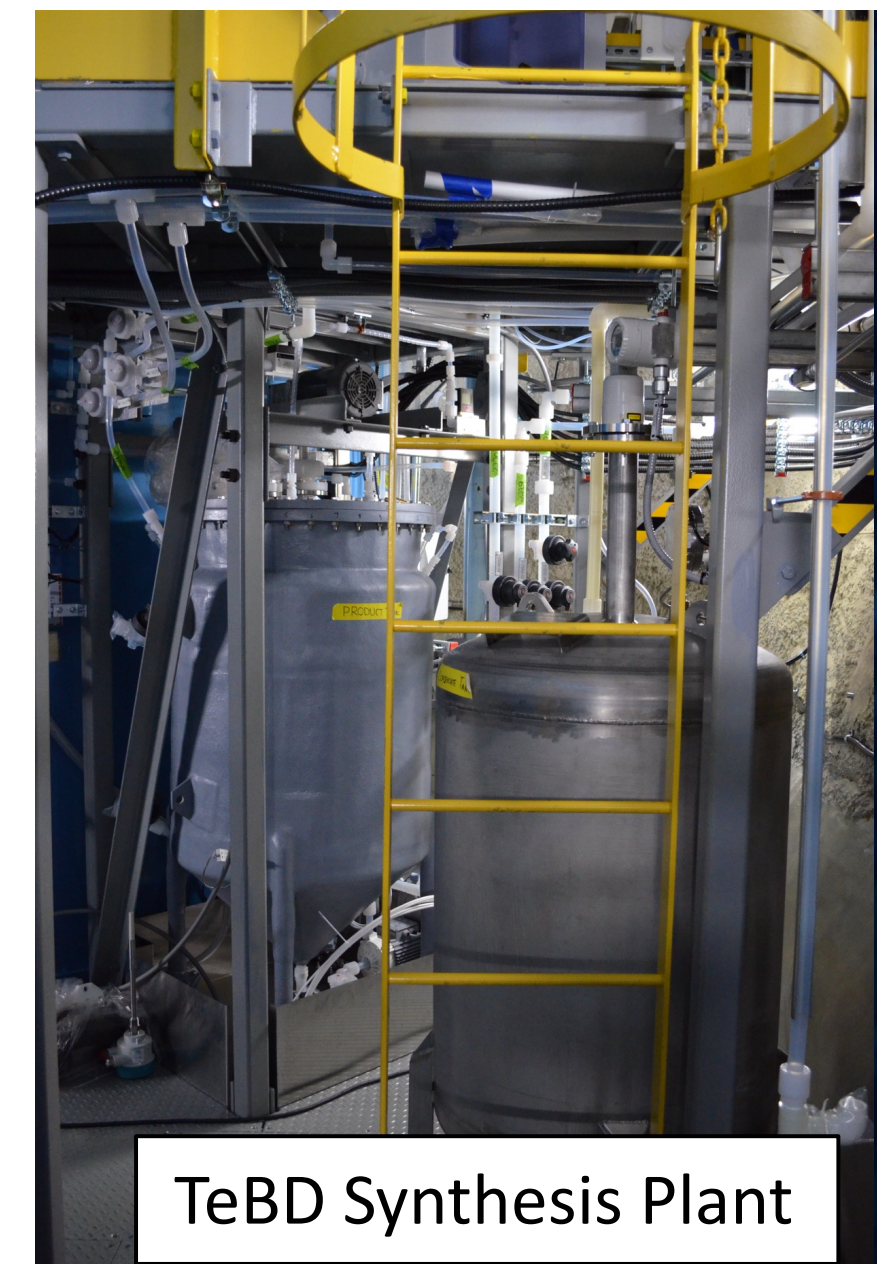


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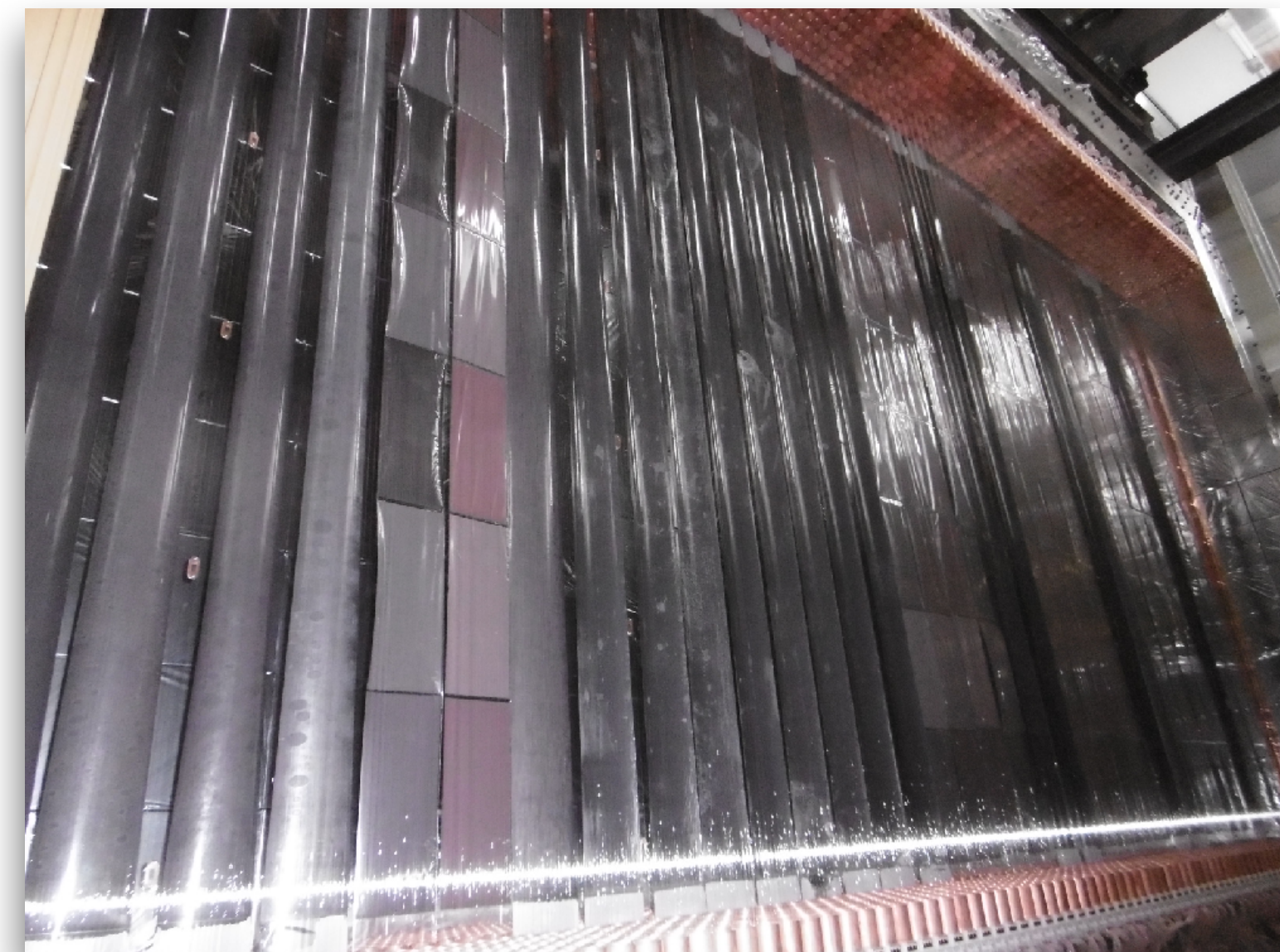
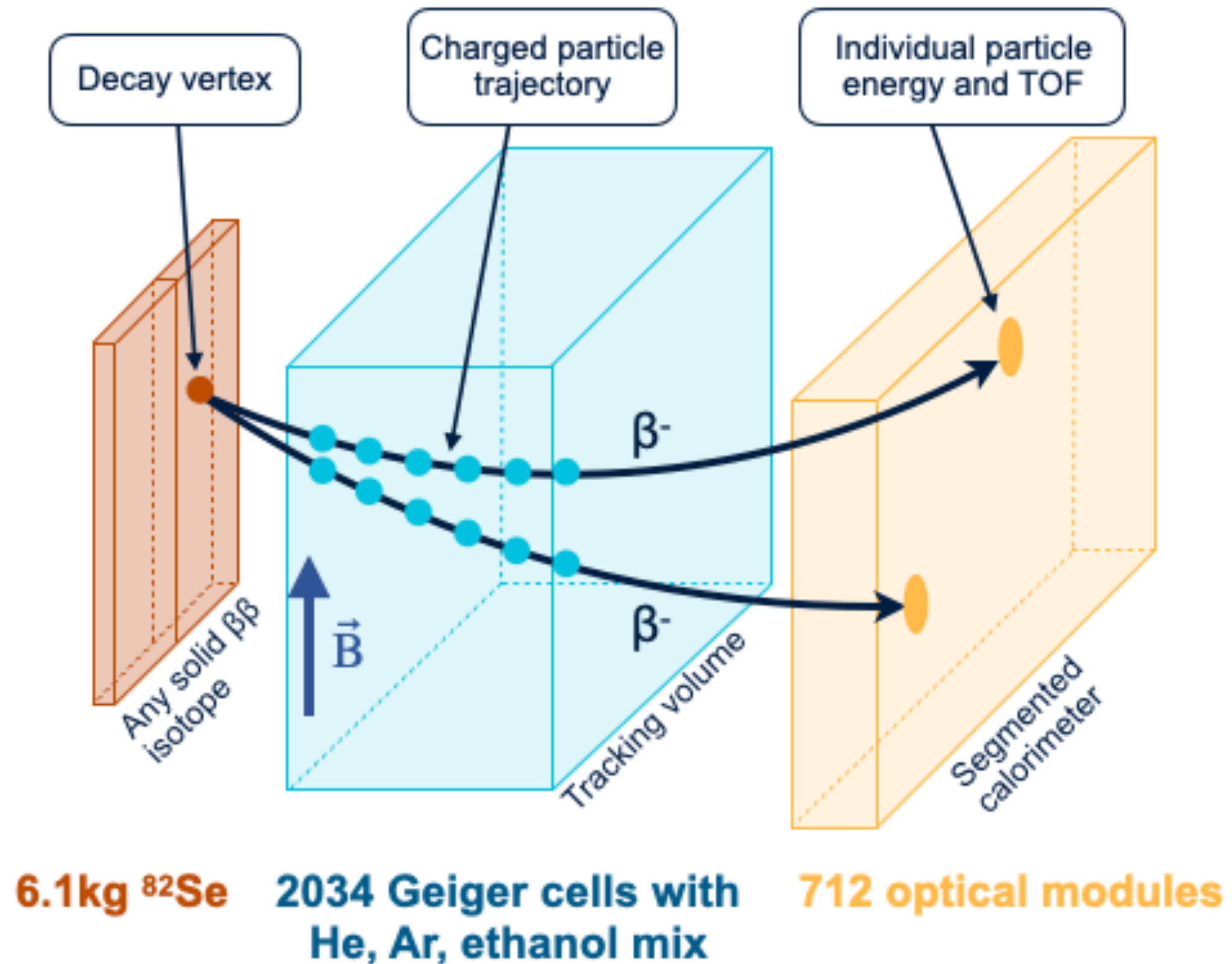
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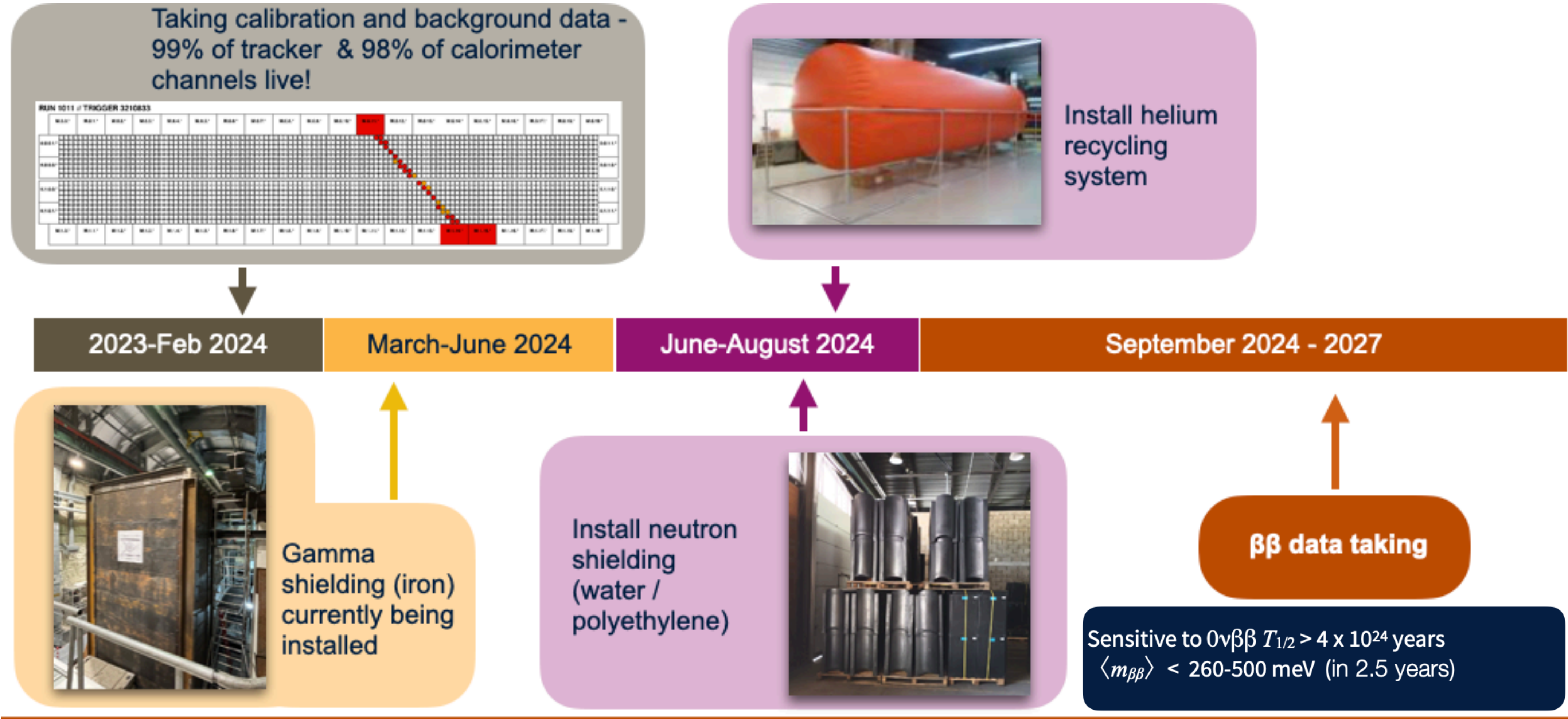
SuperNEMO (demonstrator)

- Modular concept that builds on predecessor NEMO-3 located at LSM (France)

- (Nearly) isotope-agnostic
- Full topological reconstruction and particle ID
- Unique $2\nu\beta\beta$ measurements:
 - nuclear effects
 - exotic decays & new physics
- Could probe $0\nu\beta\beta$ mechanism if discovered
- Proof of concept for future tracking detectors



SuperNEMO (demonstrator)



Poster #41

Identification, measurement, and mitigation strategy

Poster #382

Reconstructing particle energies from measured charges and track topologies

Poster #397

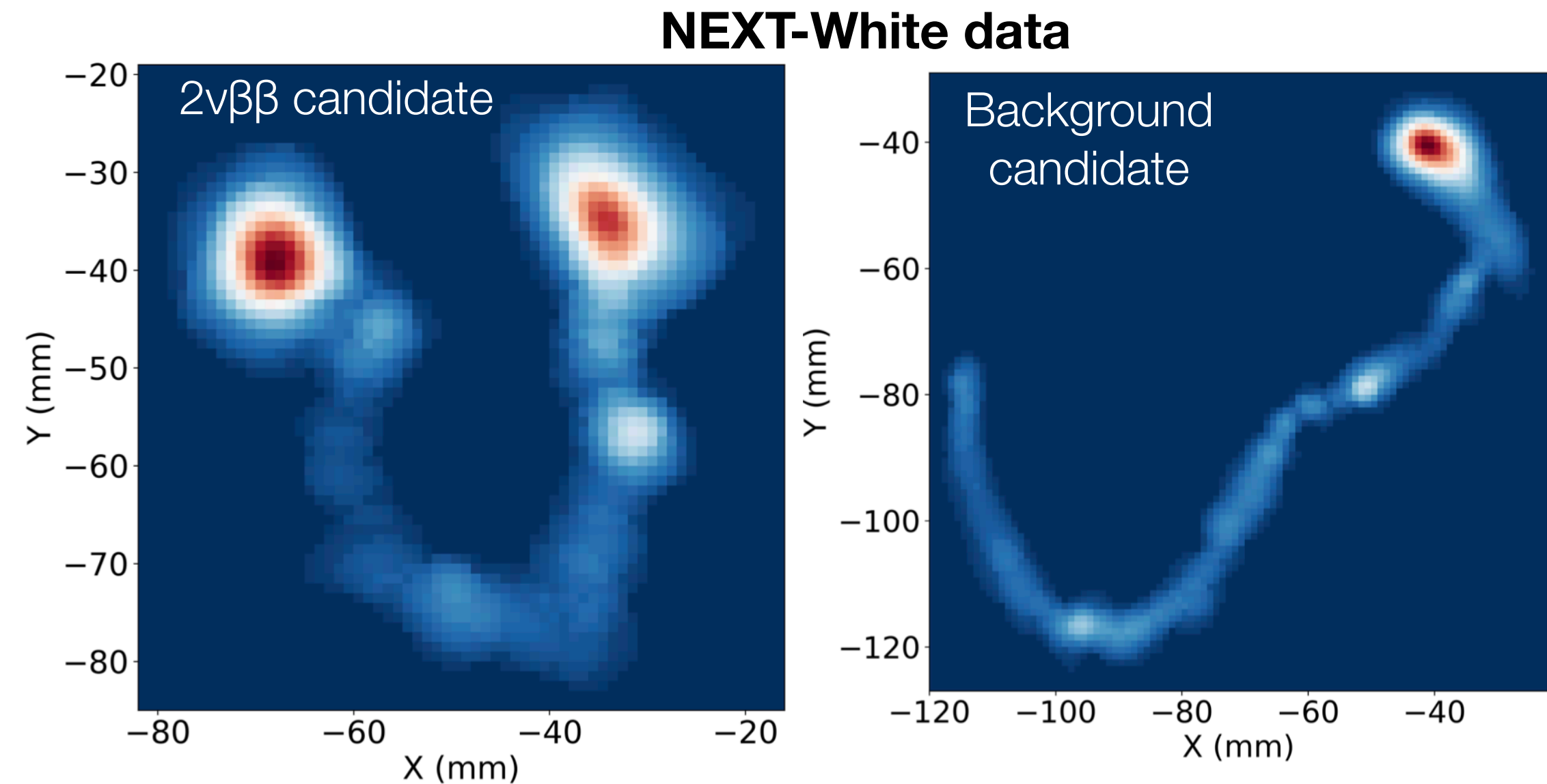
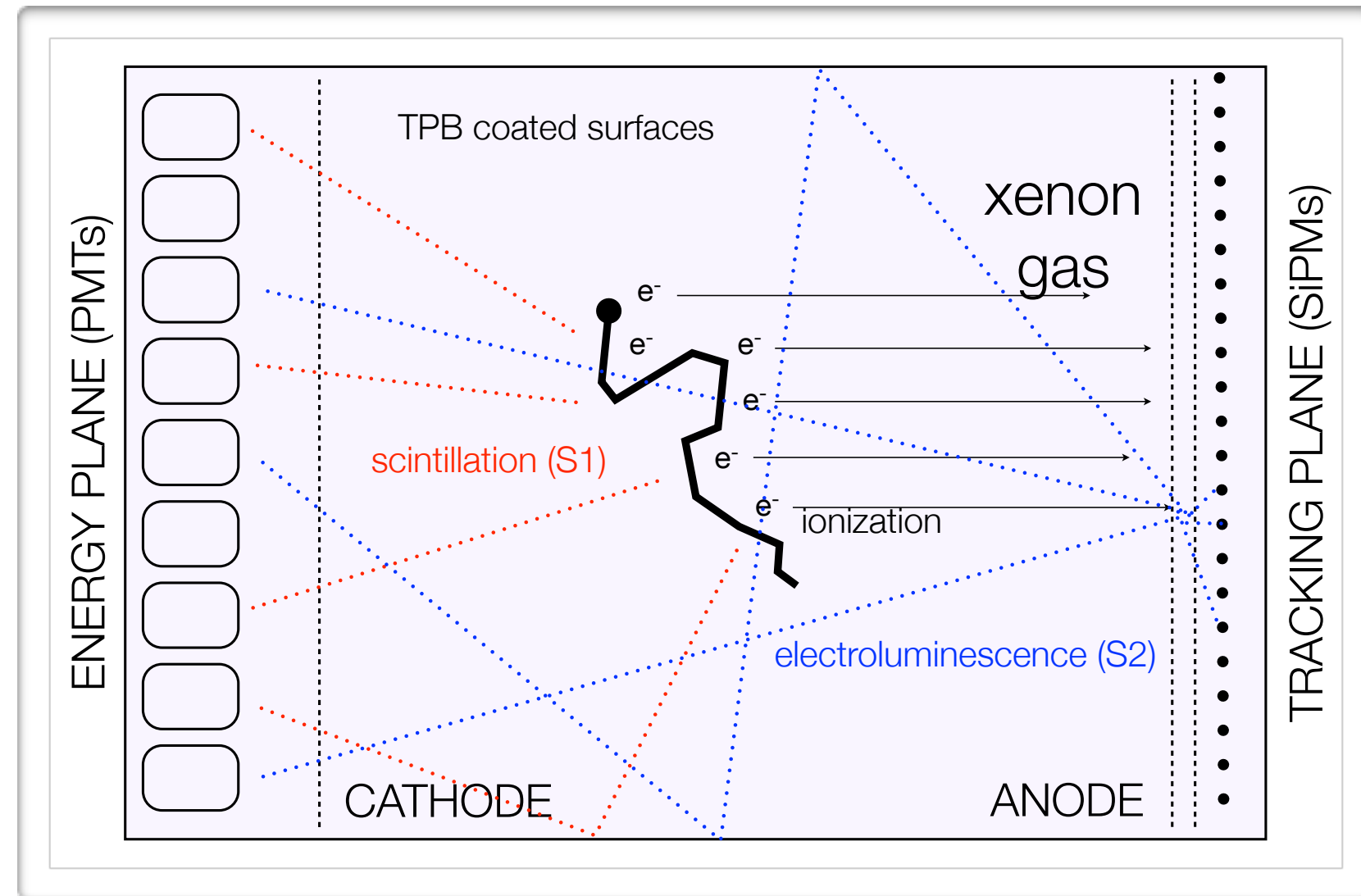
SuperNEMO's background model, and a new measurement of the γ background at LSM using SuperNEMO

Poster #451

More detail on SuperNEMO's capabilities and status

Thanks to Cheryl Patrick for the slides!

- High pressure xenon gas TPC with electroluminescence



- ✓ Gas → great energy resolution (sub-percent FWHM) *JHEP 10 (2019) 230*
- ✓ Kr calibration *JINST 13 (2018) P10014*
- ✓ Tracking plane → topological signal/background separation *JHEP 07 (2021) 146, JHEP 01 (2021) 189, JHEP 10 (2019) 052*
- ✓ Background control → runs with depleted/enriched Xe *JHEP 10 (2019) 51, Phys. Rev. C 105, 055501 (2022), JHEP 09 (2023) 190*

All demonstrated by NEXT-White

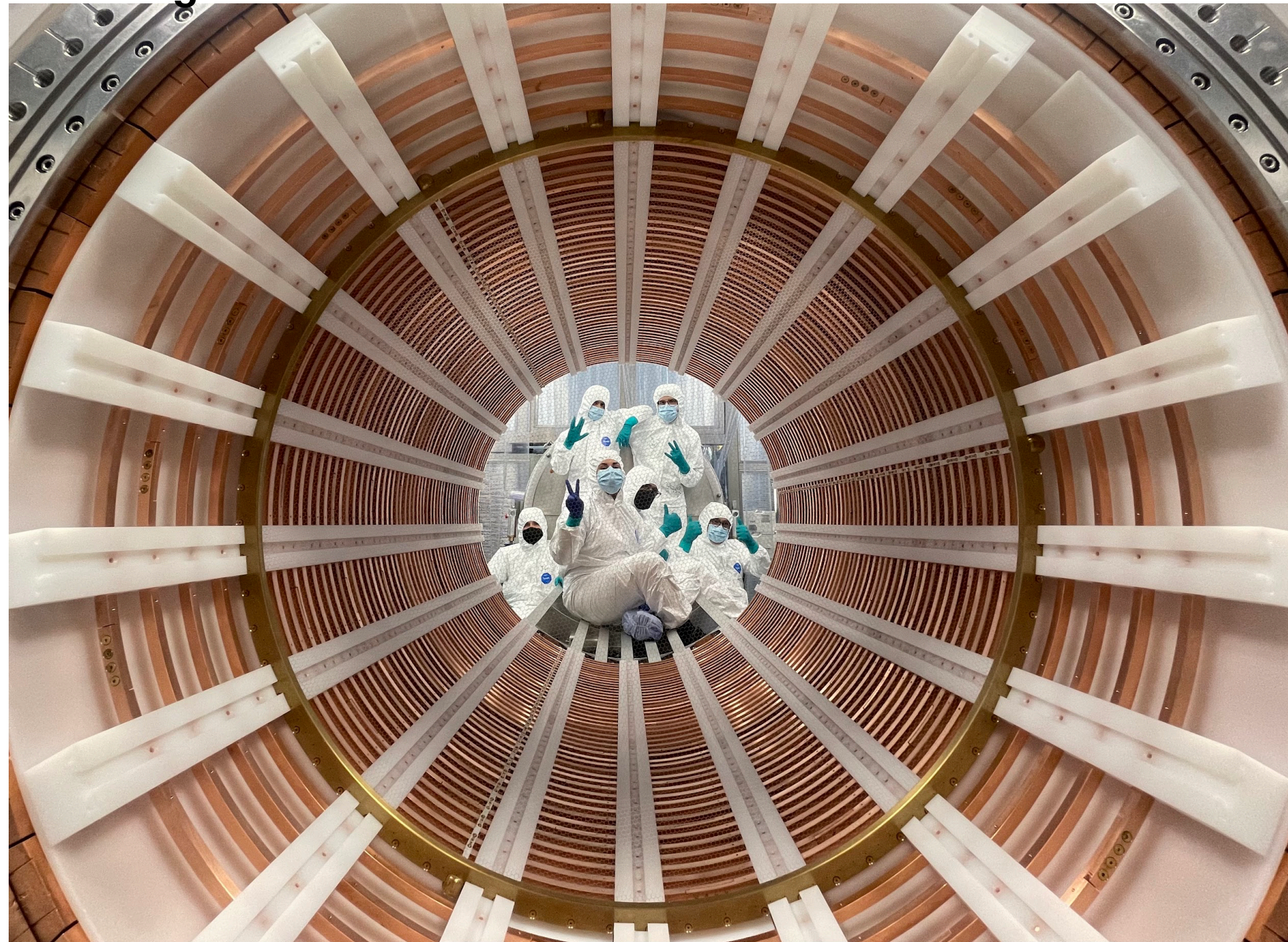
- ✓ Scalability **Goal of NEXT-100**

NEXT-100 sensitivity: 4.1×10^{25} yr
JHEP 1605 (2015) 159

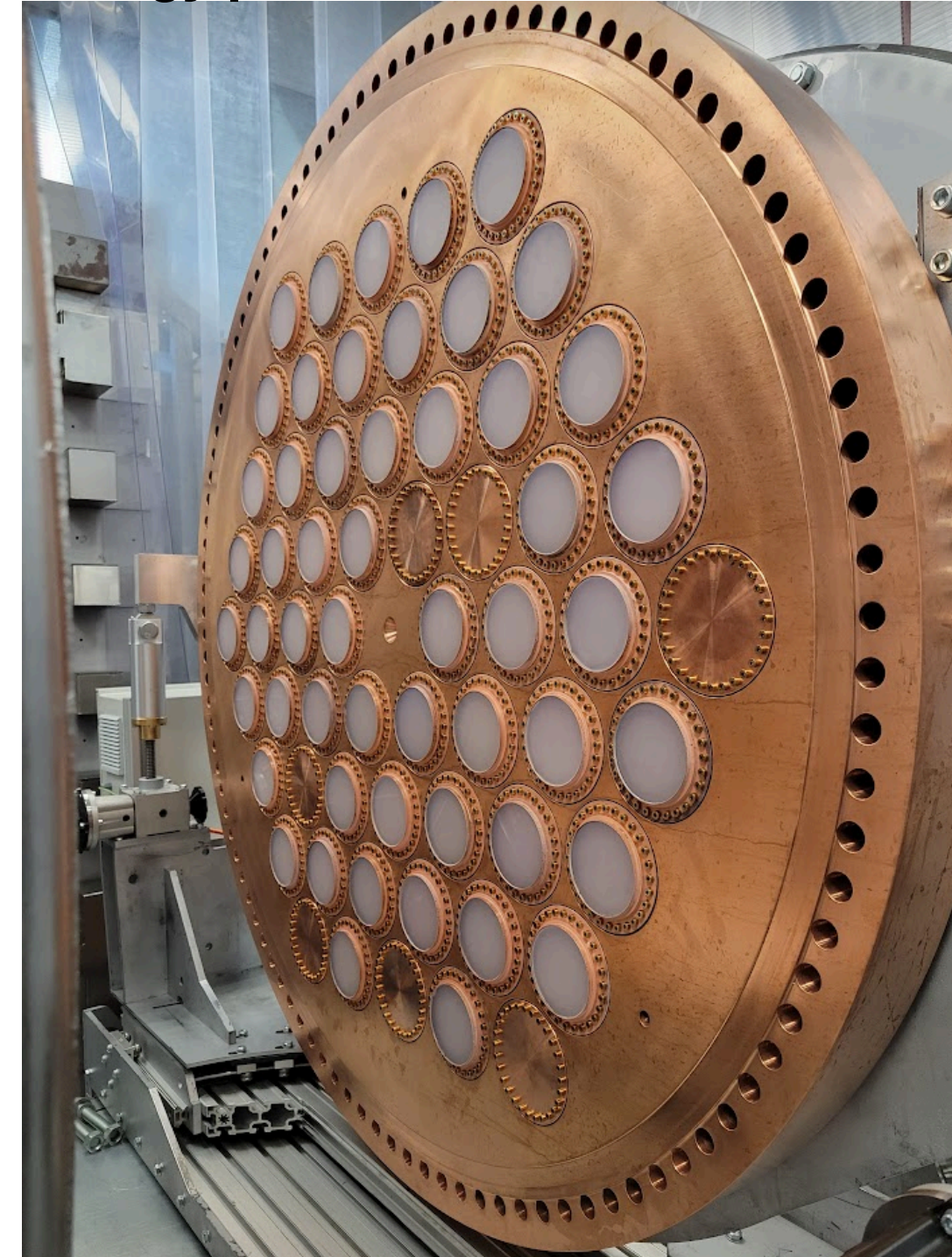
Status of NEXT-100

- Detector fully built and under commissioning!
- First runs in Ar gas in May 2024
- Xenon runs to start shortly

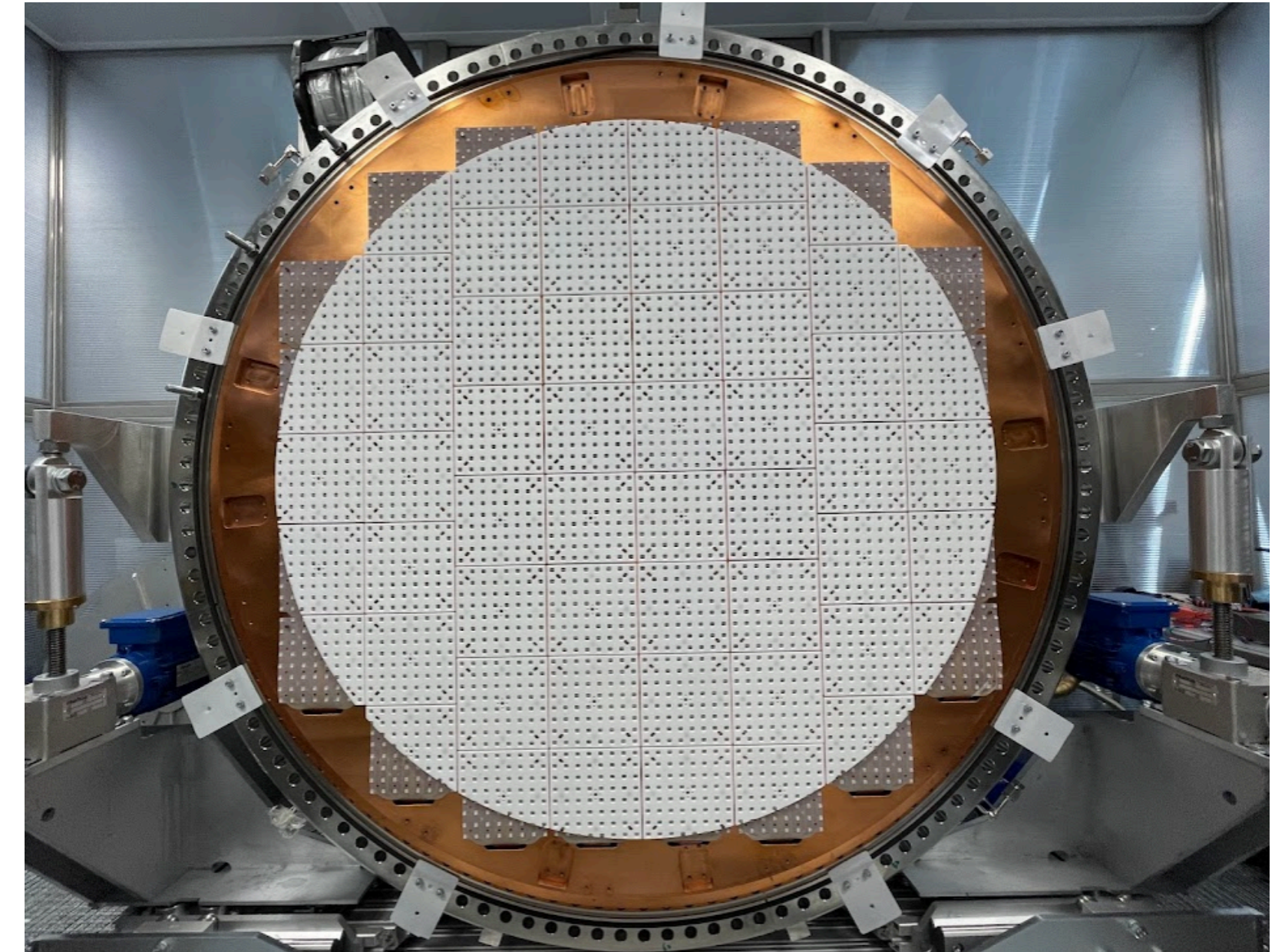
Field cage



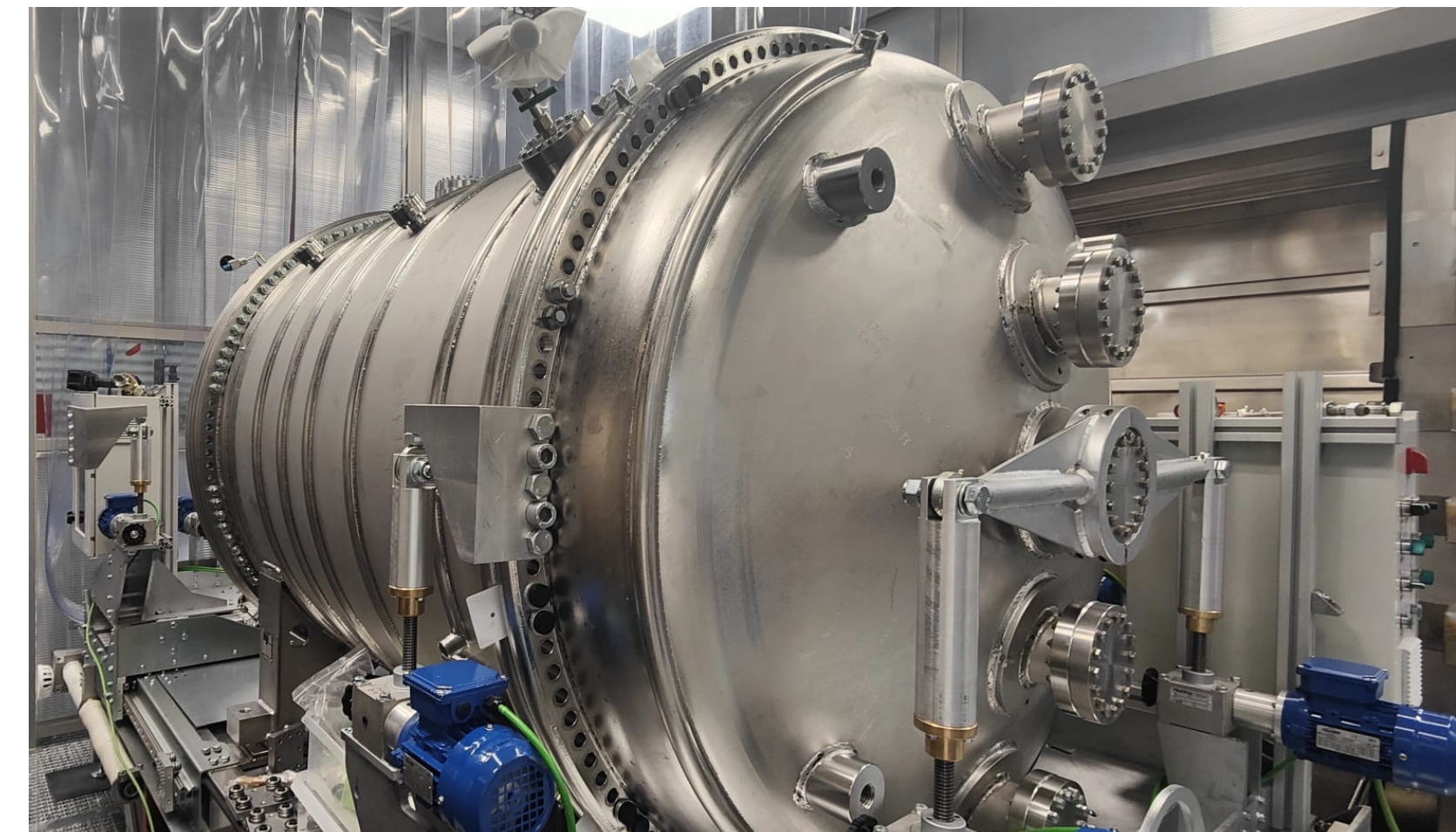
Energy plane



Tracking plane



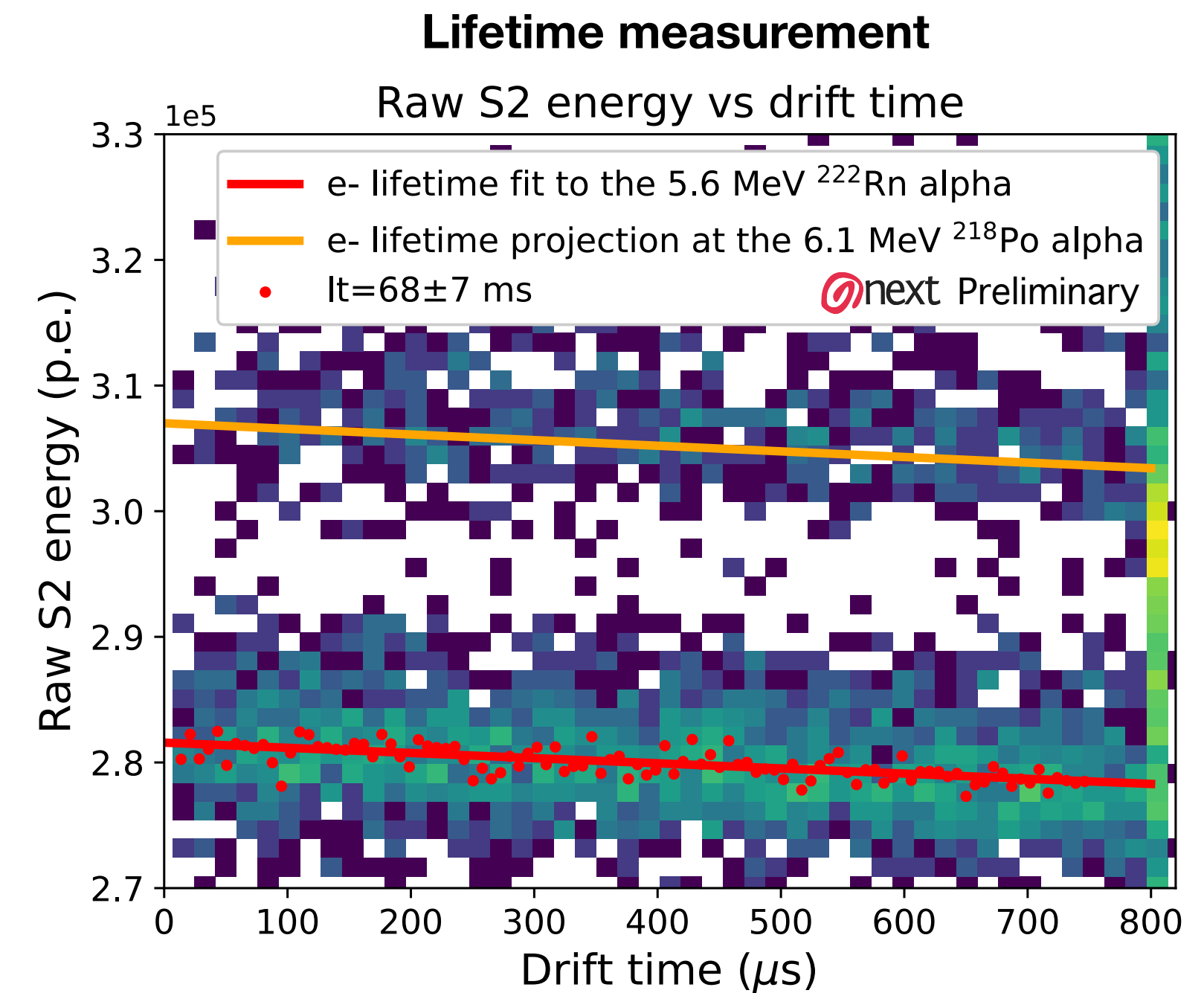
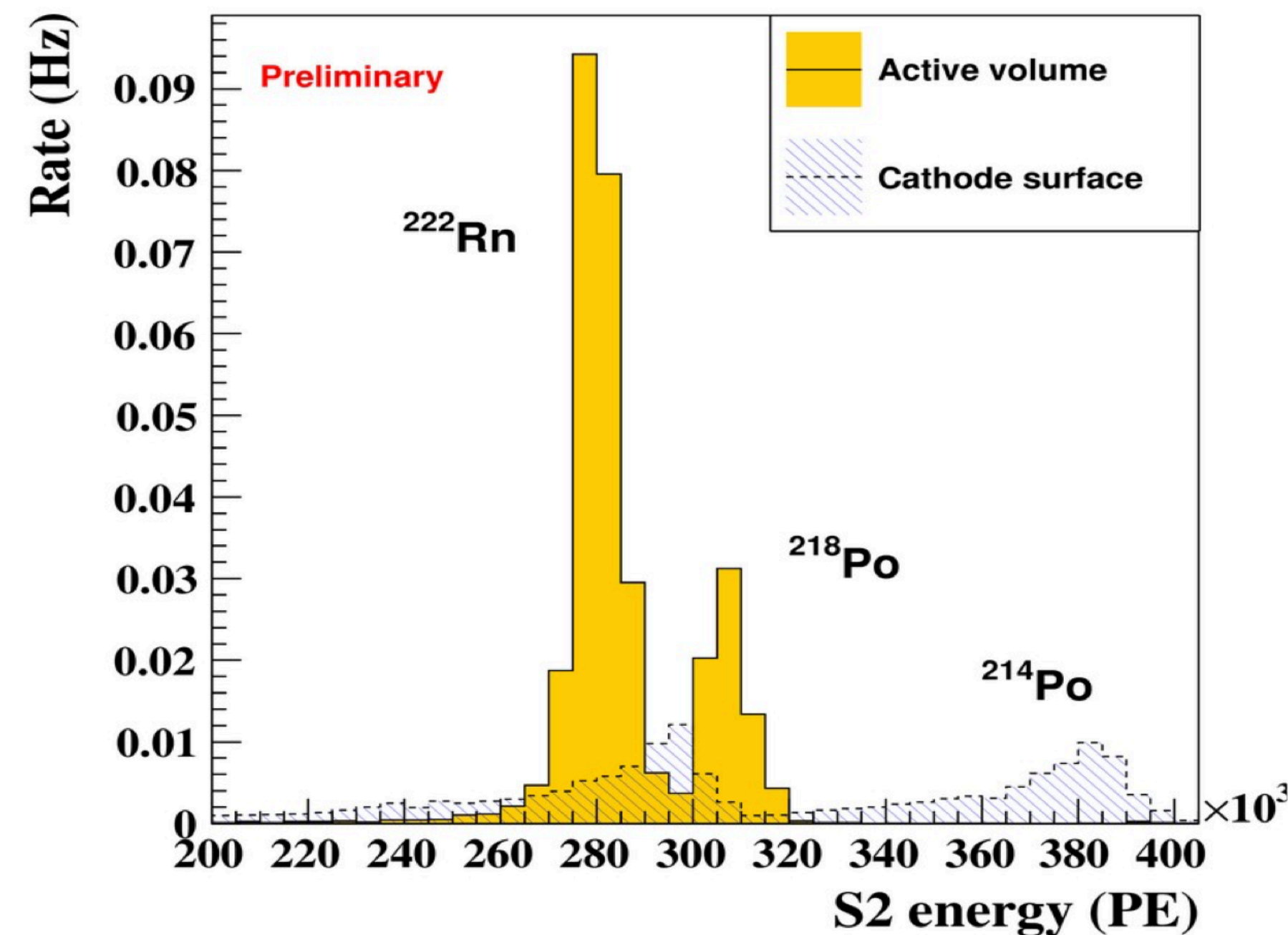
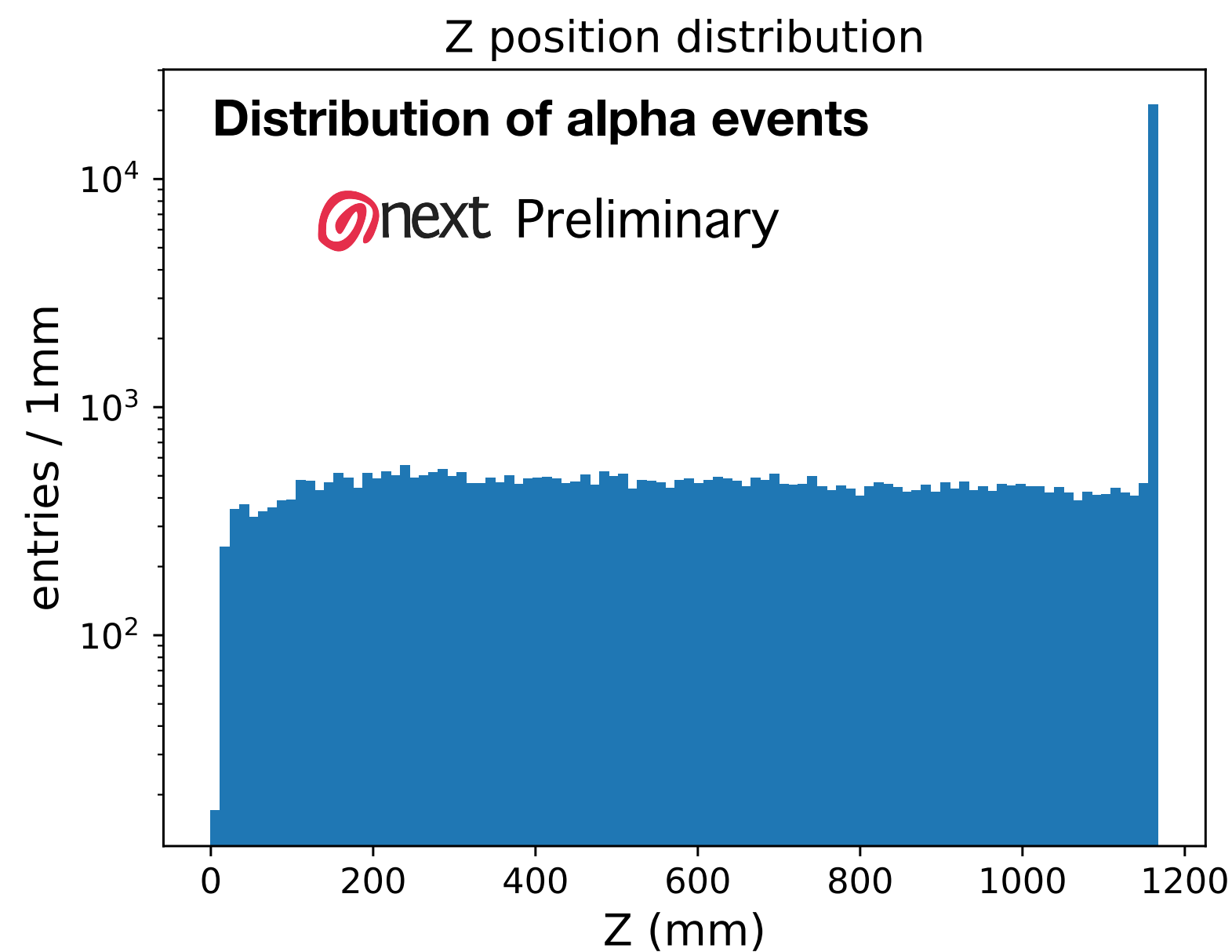
Vessel



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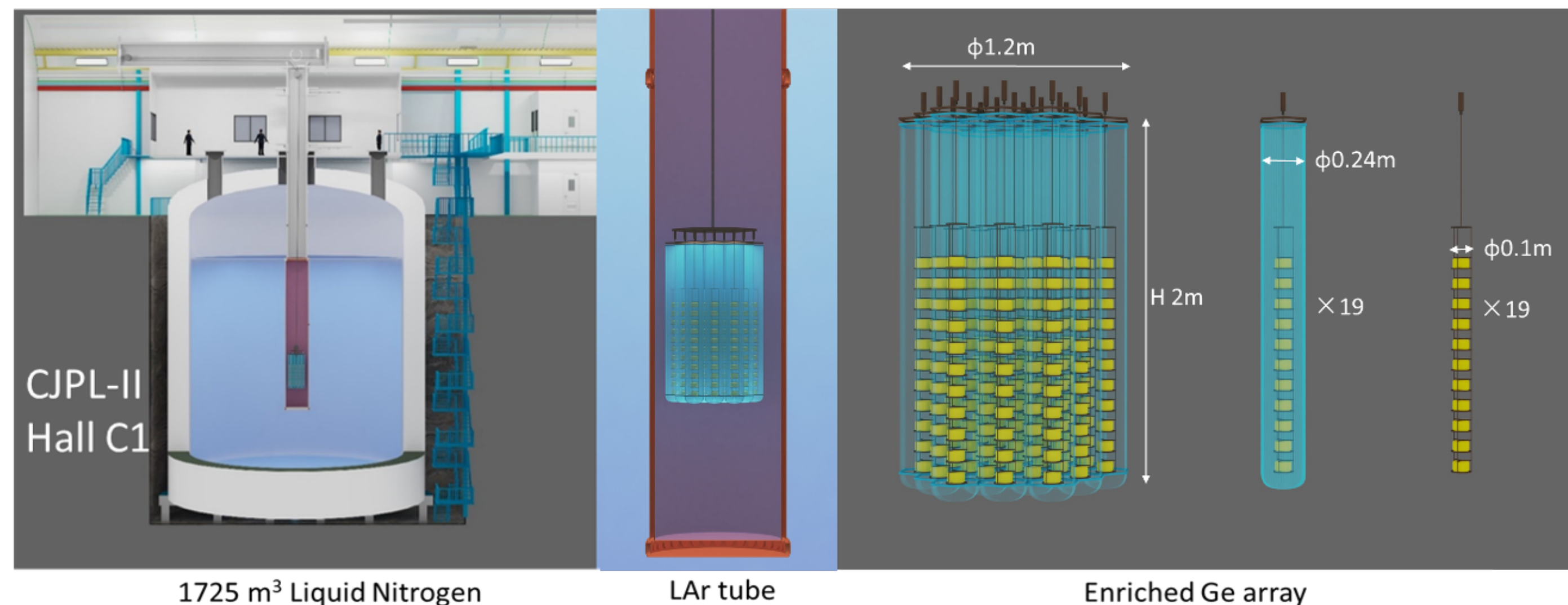
Very recent detector response on alpha particles from ^{222}Rn





Coming soon: CDEX-300v

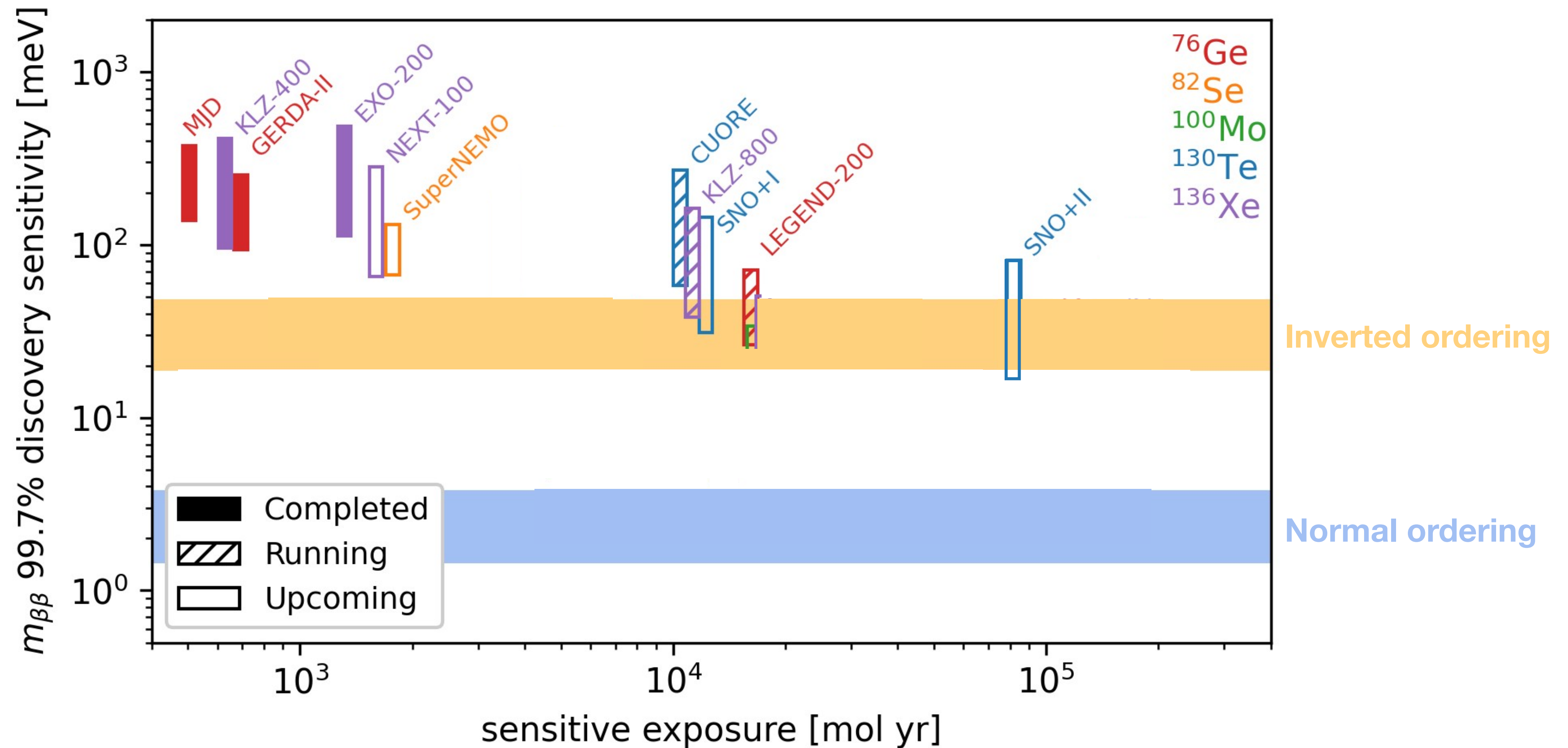
- Enr-Ge detectors test started in 2022 @ CJPL-I
- Test and operate LAr test facility in early 2024
- Hall C1 ready for experiment this June
- Experimental setup in 2024
- First batch of Ge detector installation and test in 2024
- Expected 1.3×10^{27} yr half-life sensitivity with 5 yr running



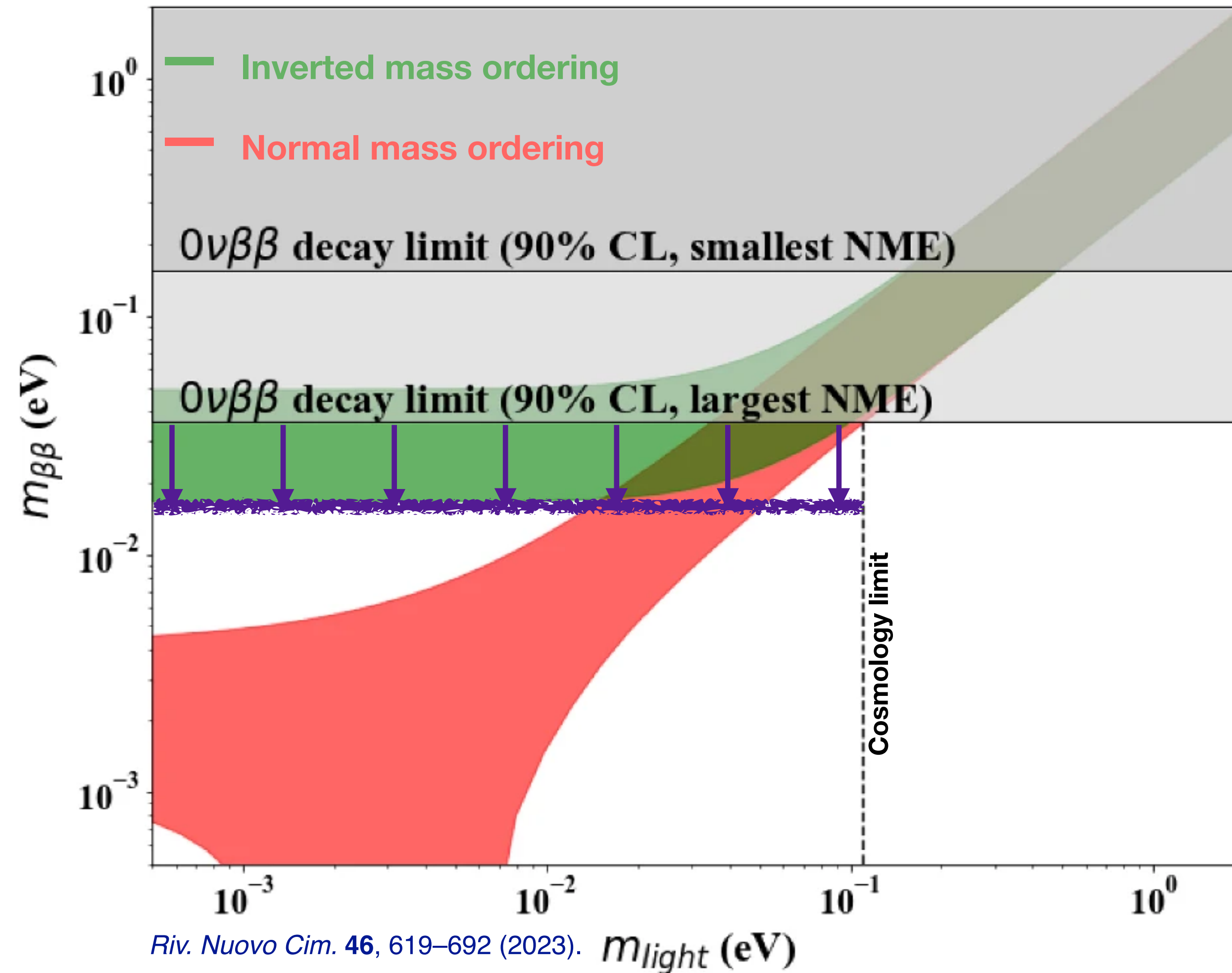
Thanks to Hao Ma for the slide!

Current generation sensitivity

- Unless we are lucky, this generation may not get there...



Next step: Conquer the Inverted Ordering



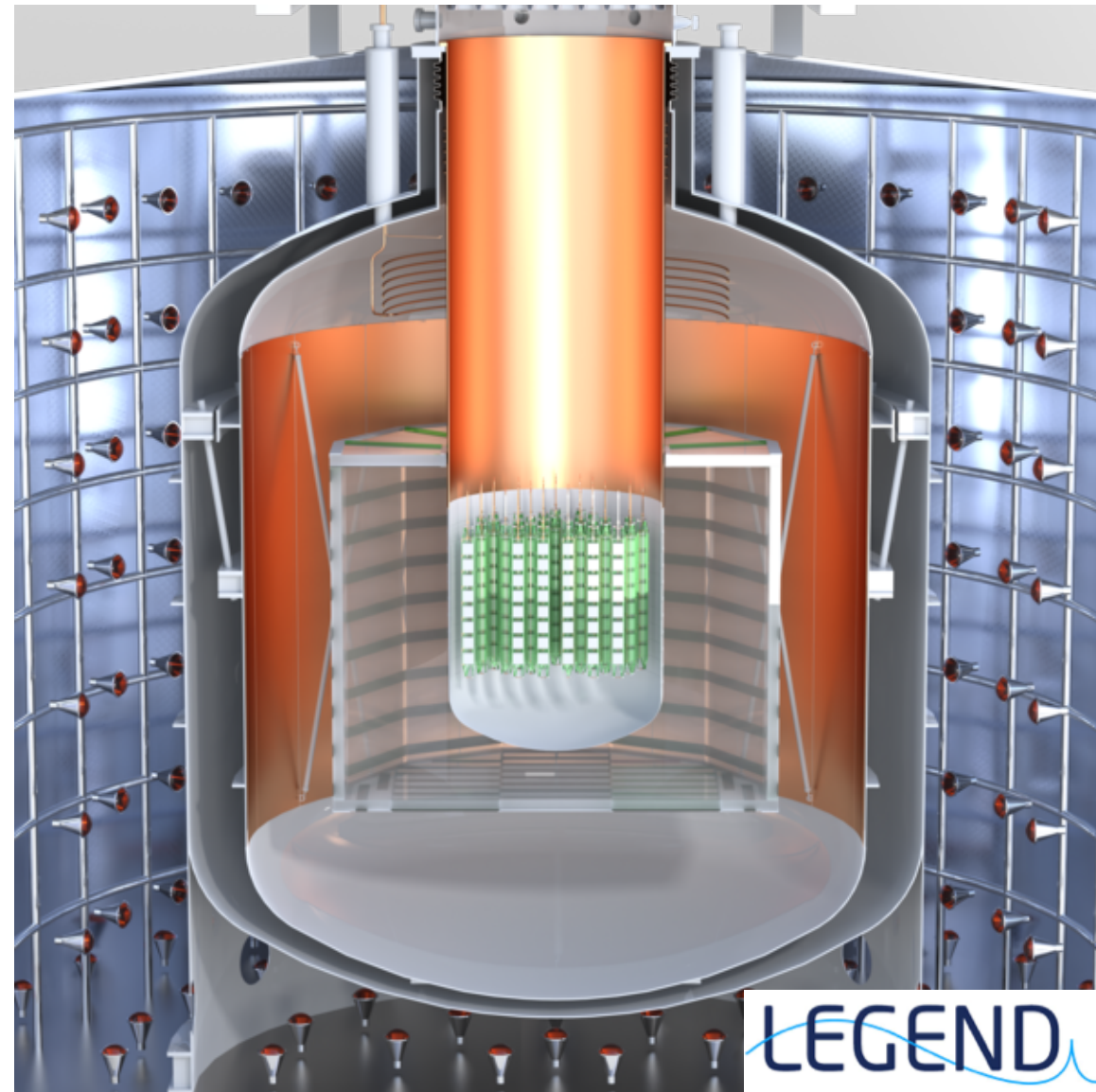
Need tonne-scale experiments!

**** This plot is just one example of parameter space... It does not mean that is exactly where the Majorana neutrino is! ****

Conquering the inverted ordering

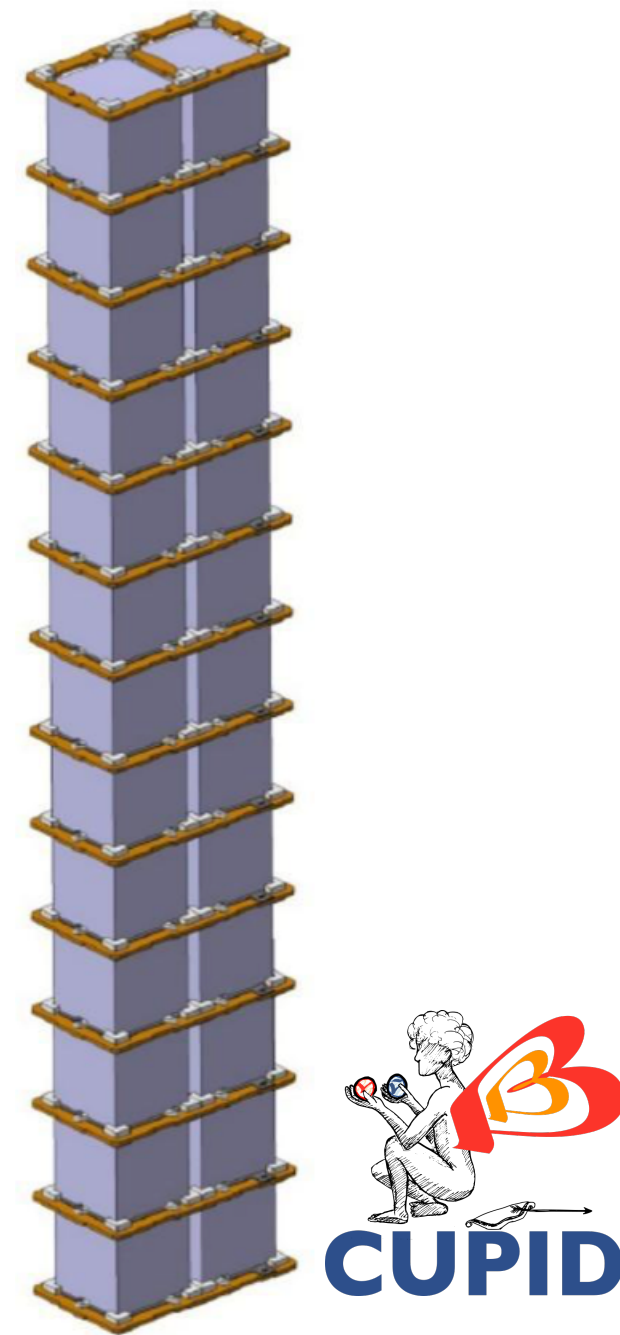
- Currently proposed tonne-scale experiments:

LEGEND-1000



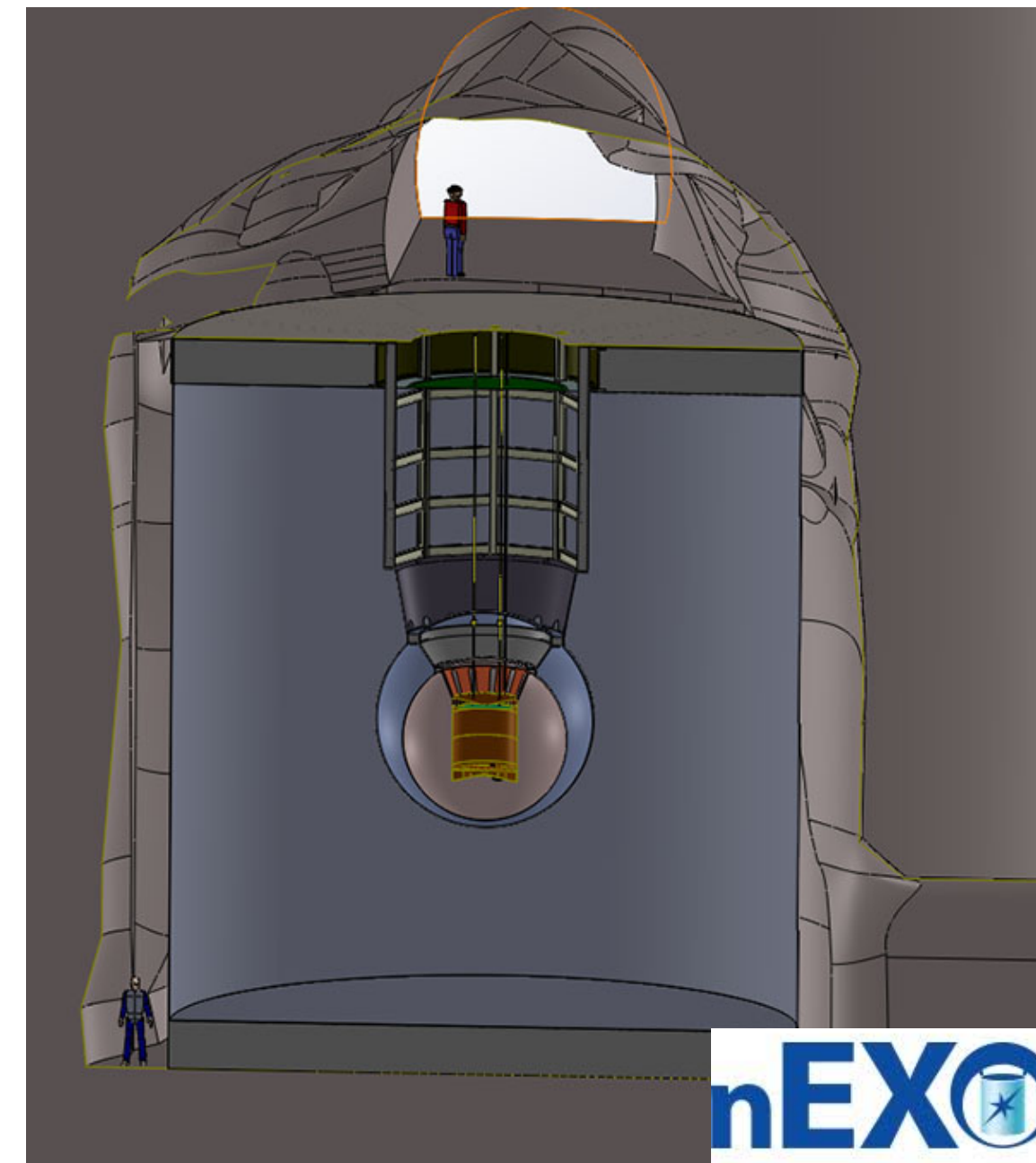
See L. Pertoldi's talk

CUPID



See C. Bucci's talk

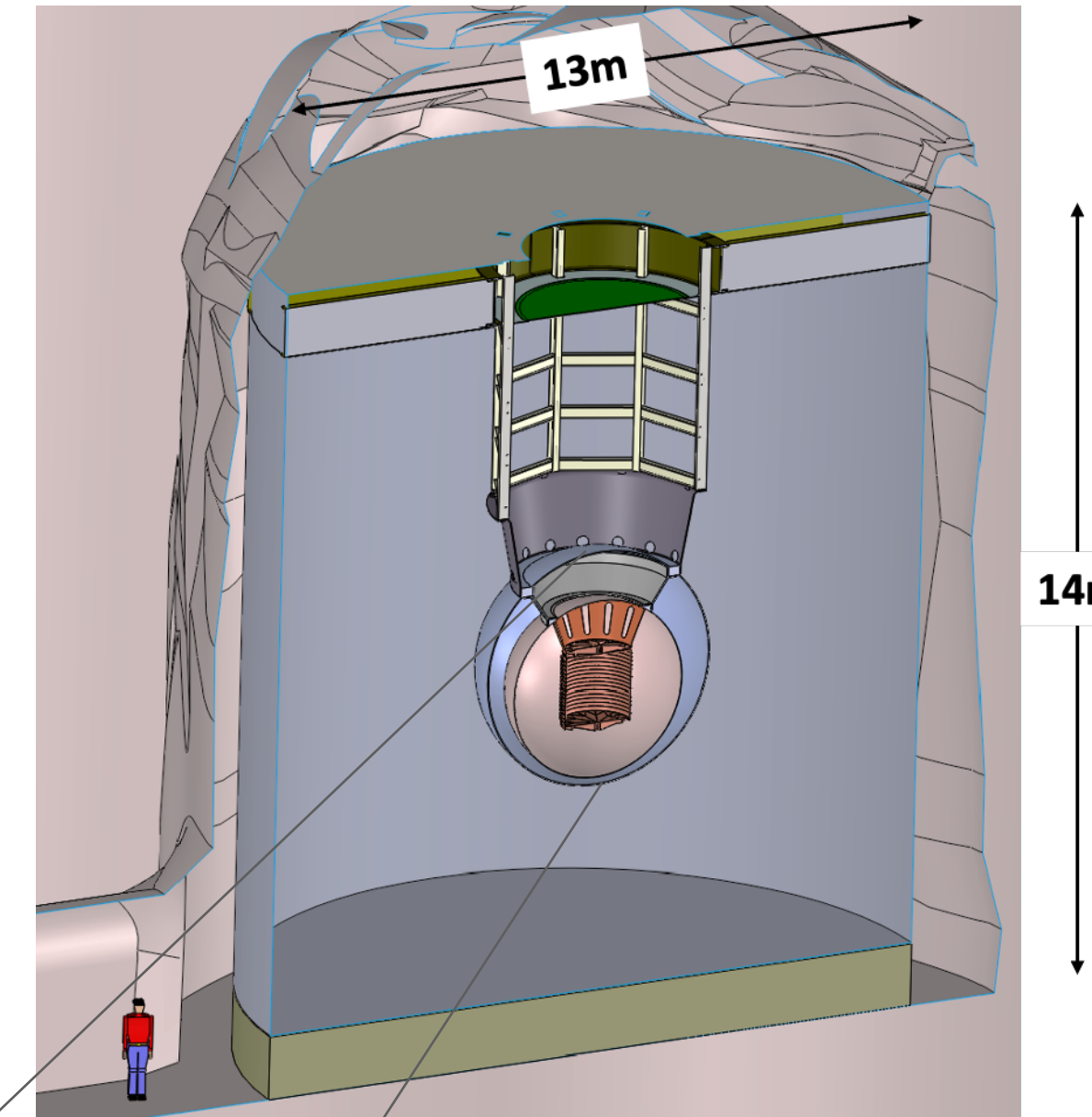
nEXO



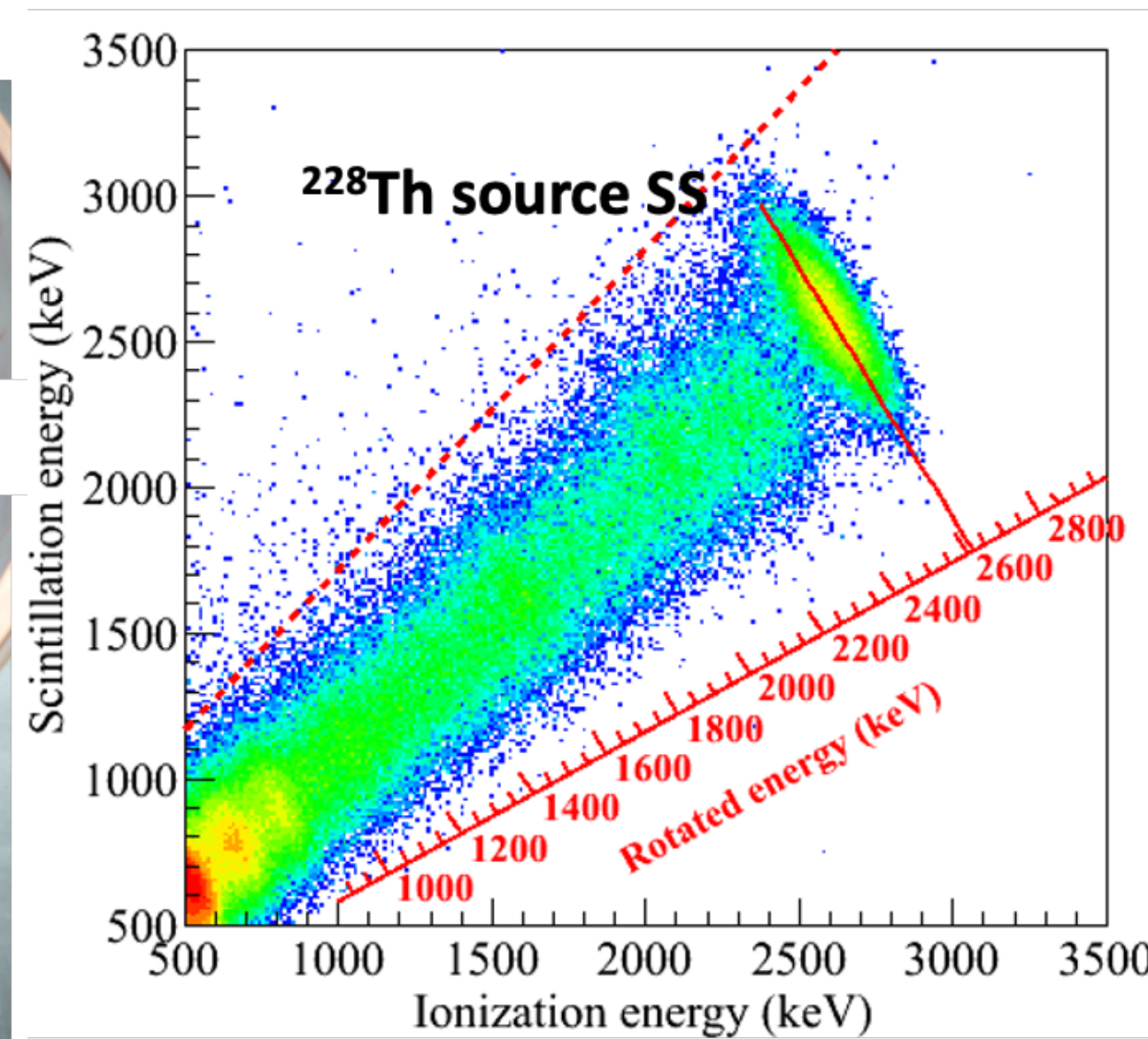
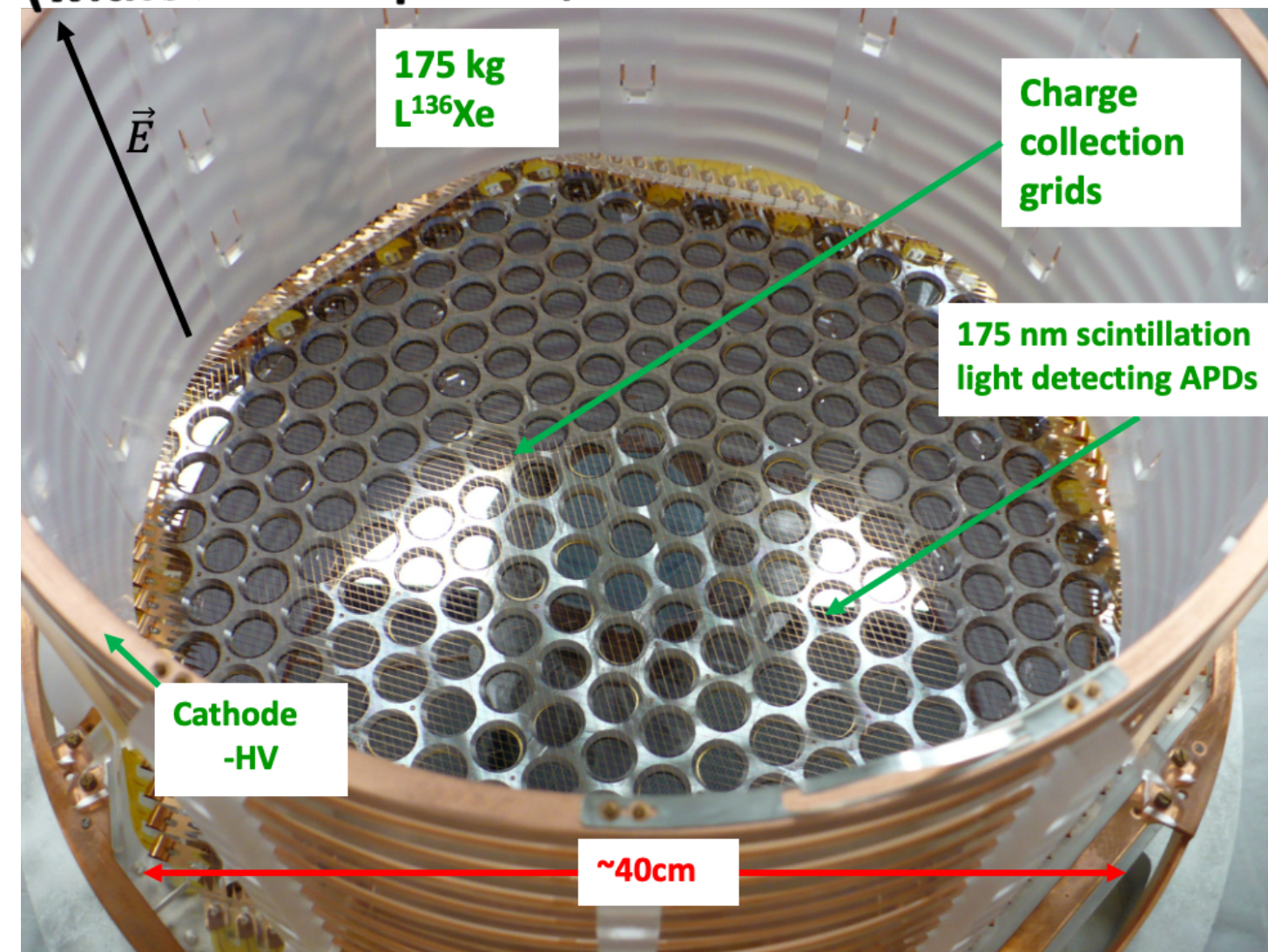
NEXT-HD



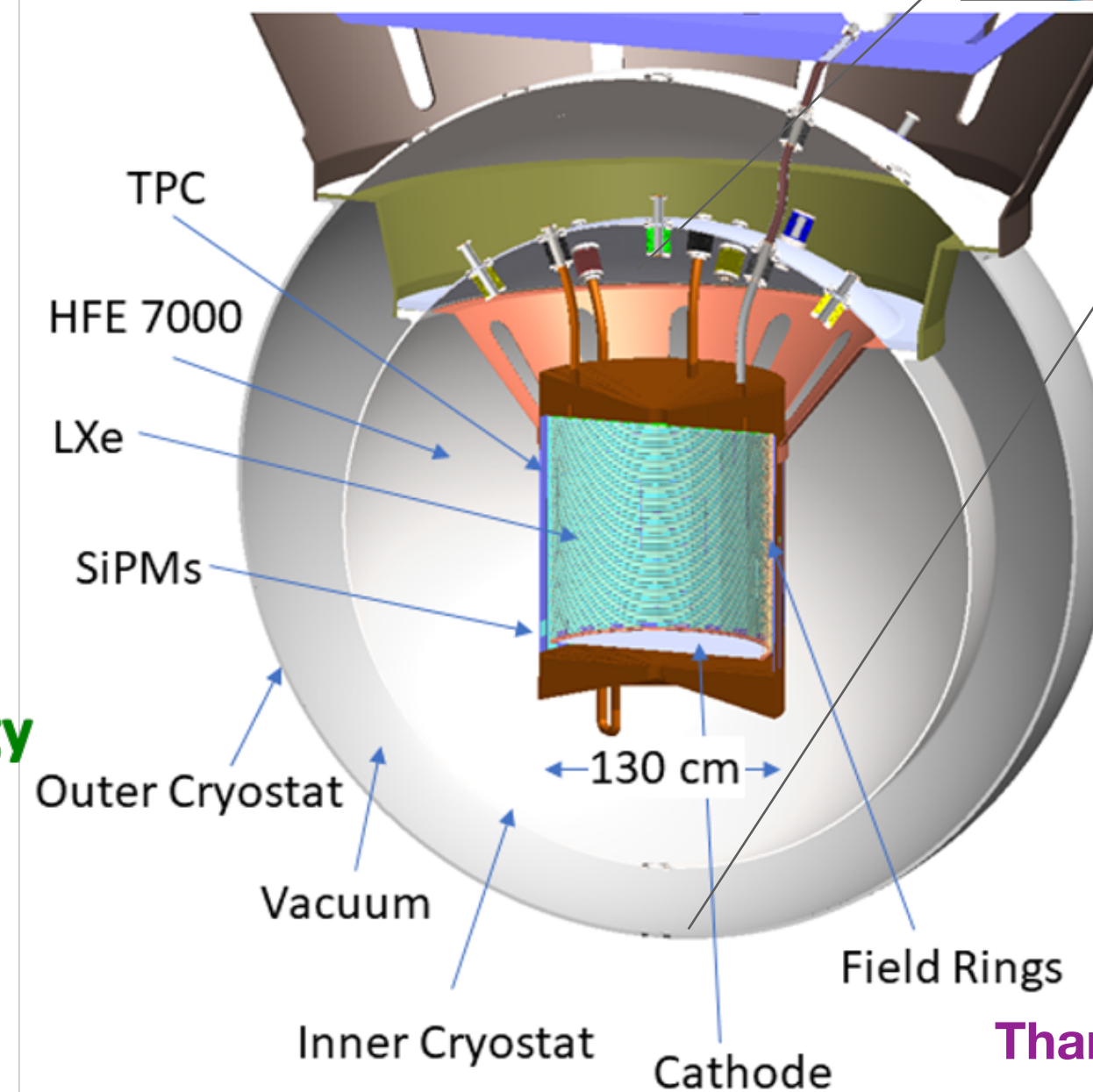
- Builds (directly extrapolated) on successful EXO-200 predecessor
- 5 tonnes of enriched LXe TPC (^{136}Xe)



(1/2 of) the EXO-200 TPC
(without the liquid Xe)



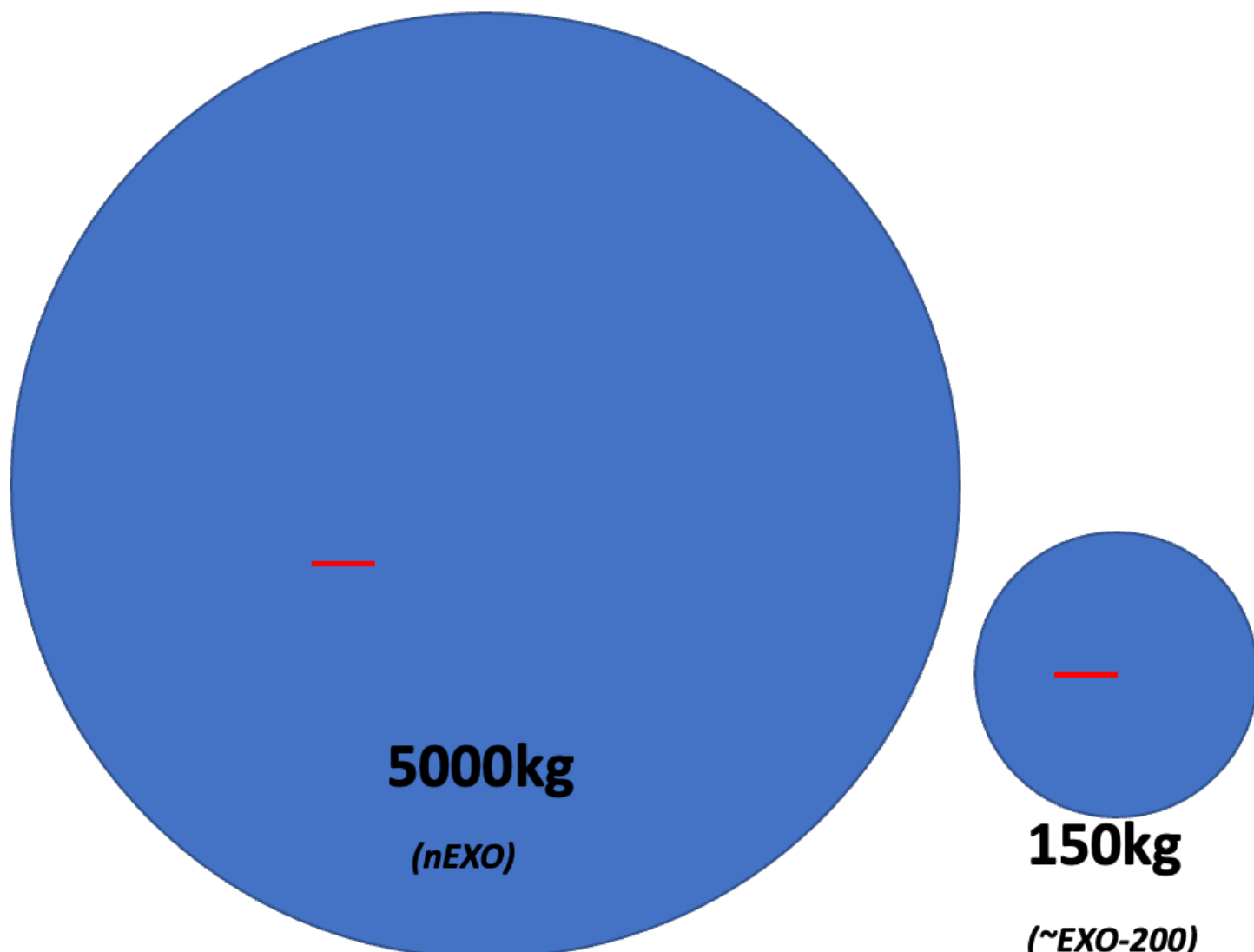
Rotation angle chosen to optimize energy resolution near $Q_{\beta\beta}=2457.83$ keV



Posters # 478, 447, 294

2012: *Phys.Rev.Lett.* 109 (2012) 032505
 2014: *Nature* 510 (2014) 229-234
 2018: *Phys. Rev. Lett.* 120, 072701 (2018)
 2019: *Phys. Rev. Lett.* 123 (2019) 161802

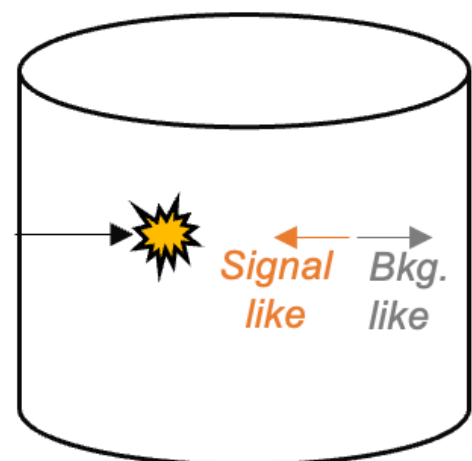
Great self-shielding power with LXe



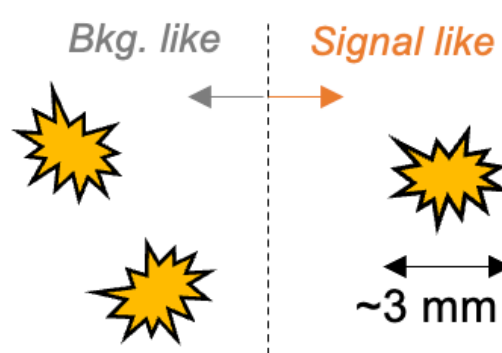
2.5 MeV γ attenuation length 8.7cm =

Signal vs background

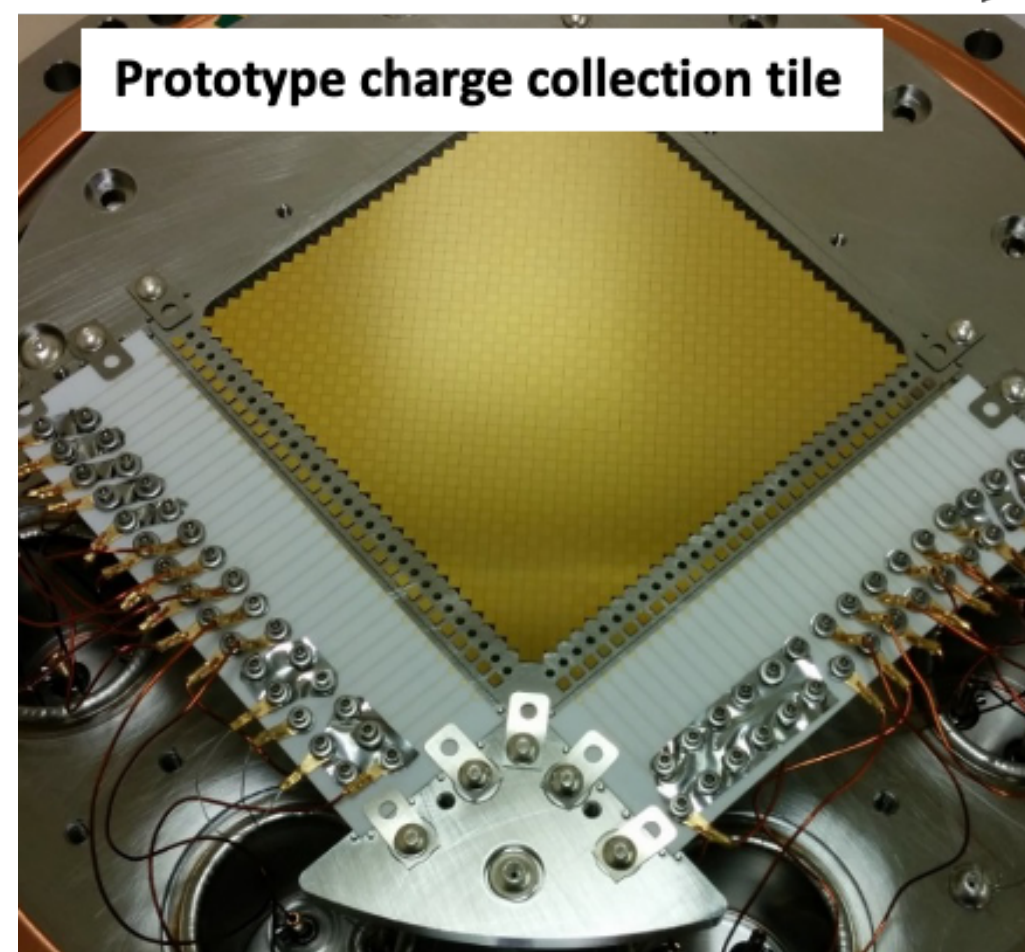
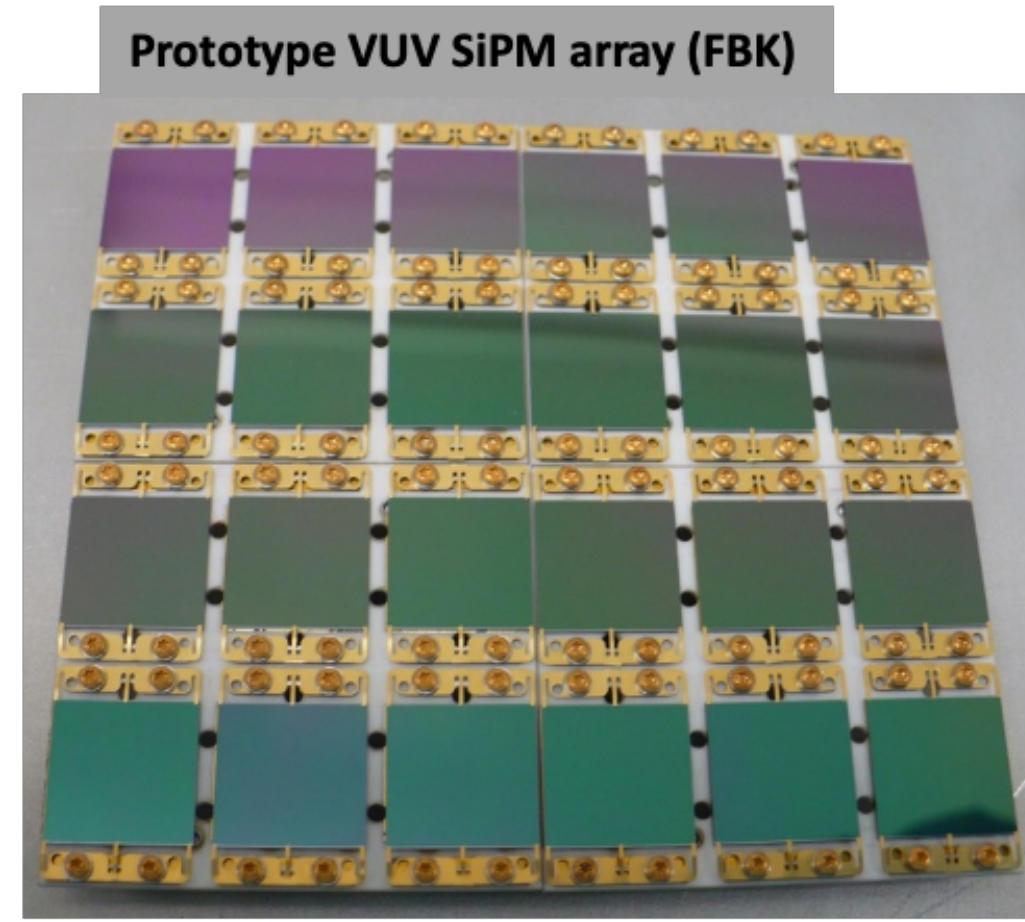
Standoff:



Topology:



New readout technologies



	EXO-200:	nEXO:	Improvements:
Vessel and cryostat	Thin-walled commercial Cu w/HFE	<i>Thin-walled electroformed Cu w/HFE</i>	Lower background
High voltage	Max voltage: 25 kV (end-of-run)	<i>Operating voltage: 50 kV</i>	Full scale parts tested in LXe prior to installation to minimize risk
Cables	Cu clad polyimide (analog)	<i>Cu clad polyimide (digital)</i>	Same cable/feedthrough technology, R&D identified 10x lower bkg substrate and demonstrated digital signal transmission
e⁻ lifetime	3-5 ms	<i>5 ms (req.), 10 ms (goal)</i>	Minimal plastics (no PTFE reflector), lower surface to volume ratio, detailed materials screening program
Charge collection	Crossed wires	<i>Gridless modular tiles</i>	R&D performed to demonstrate charge collection with tiles in LXe, detailed simulation developed
Light collection	APDs + PTFE reflector	<i>SiPMs around TPC barrel</i>	SiPMs avoid readout noise, R&D demonstrated prototypes from two vendors
Energy resolution	1.2%	<i>1.2% (req.), 0.8% (goal)</i>	Improved resolution due to SiPMs (negligible readout noise in light channels)
Electronics	Conventional room temp.	<i>In LXe ASIC-based design</i>	Minimize readout noise for light and charge channels, nEXO prototypes demonstrated in R&D and follow from LAr TPC lineage
Background control	Measurement of all materials	<i>Measurement of all materials</i>	RBC program follows successful strategy demonstrated in EXO-200
Larger size	>2 atten. length at center	<i>>7 atten. length at center</i>	Exponential attenuation of external gammas and more fully contained Comptons

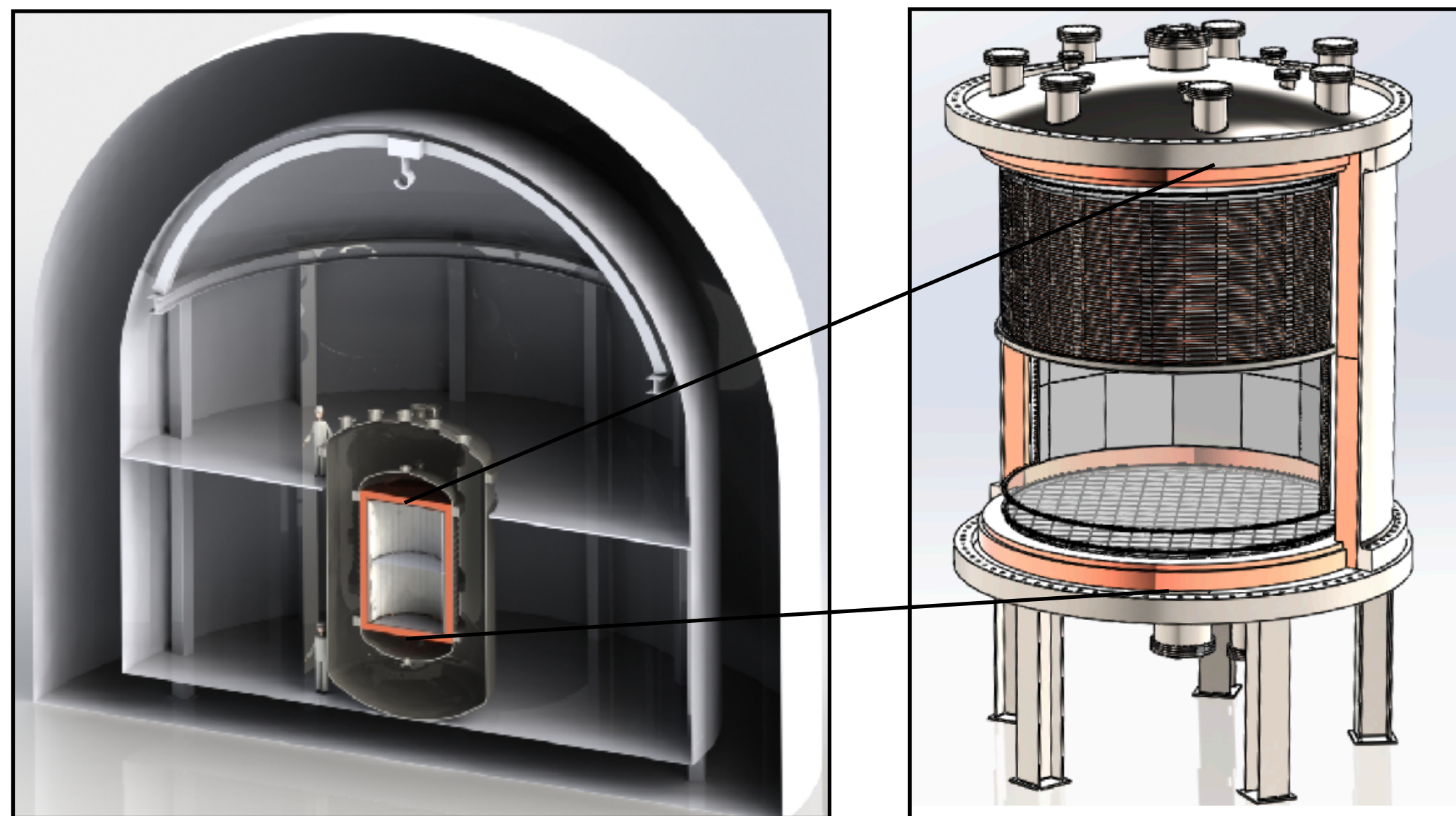
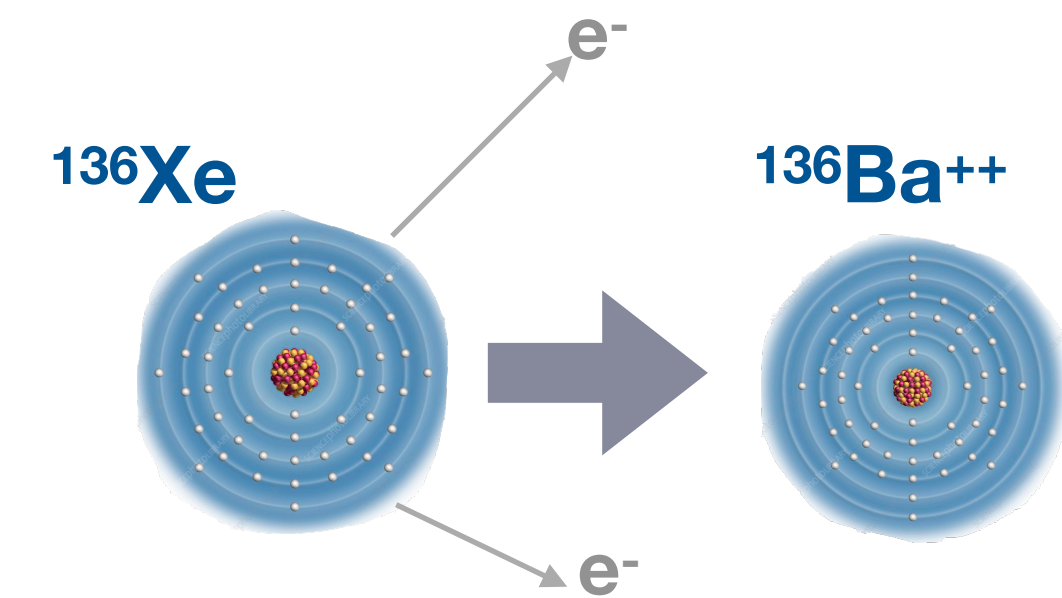
nEXO sensitivity reaches 10²⁸ yr in 6.5 yr data taking

NEXT-HD (and NEXT-BOLD)



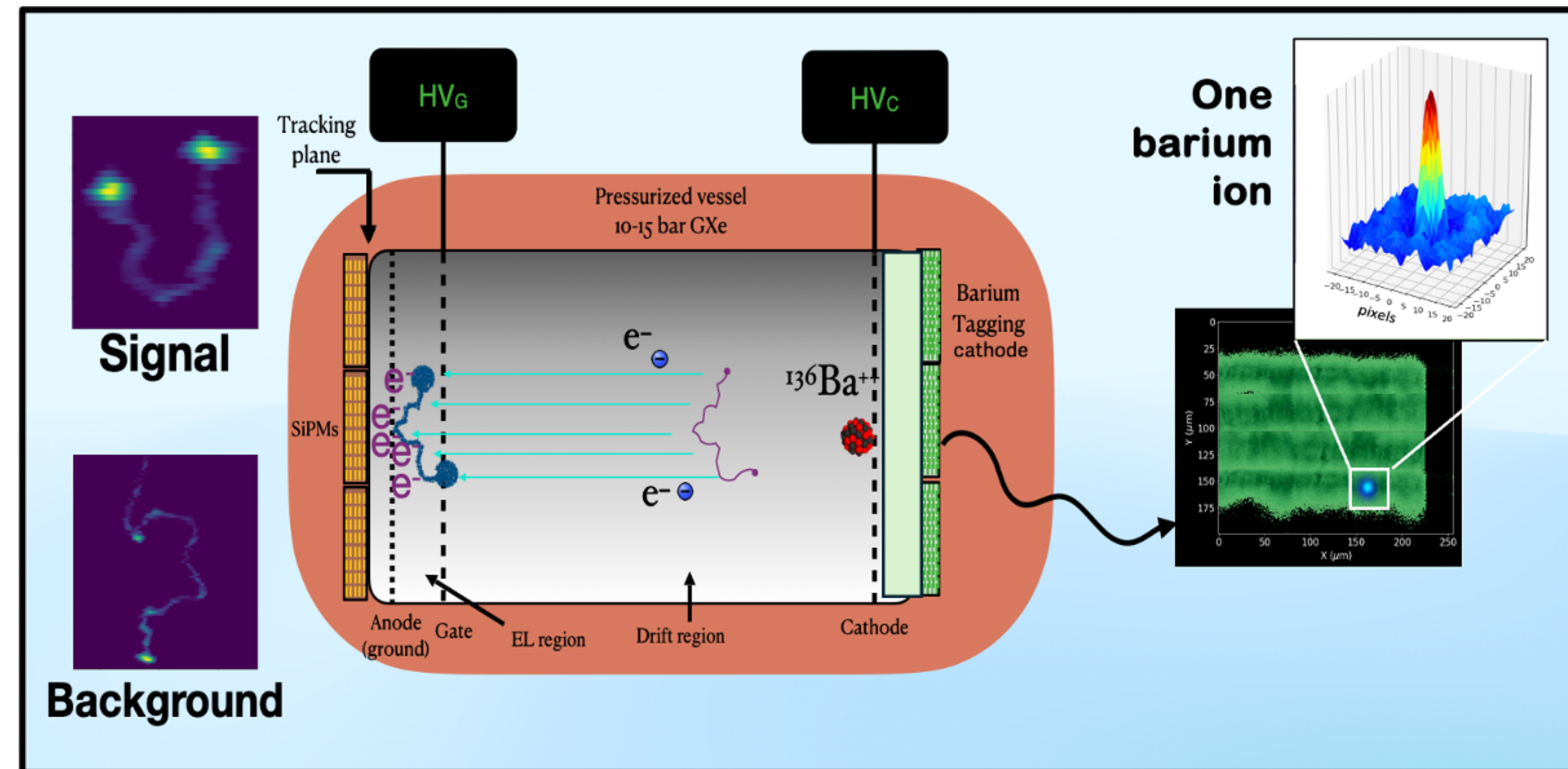
- Module(s) of 1 tonne of ^{136}Xe building in NEXT technologies (NEXT-HD)
- Designed to accommodate future Ba tagging (NEXT-BOLD)

Daughter tagging:



NEXT Collaboration, *JHEP* 164 (2021) 08

Letter of Intent to be submitted in 2025 to LSC

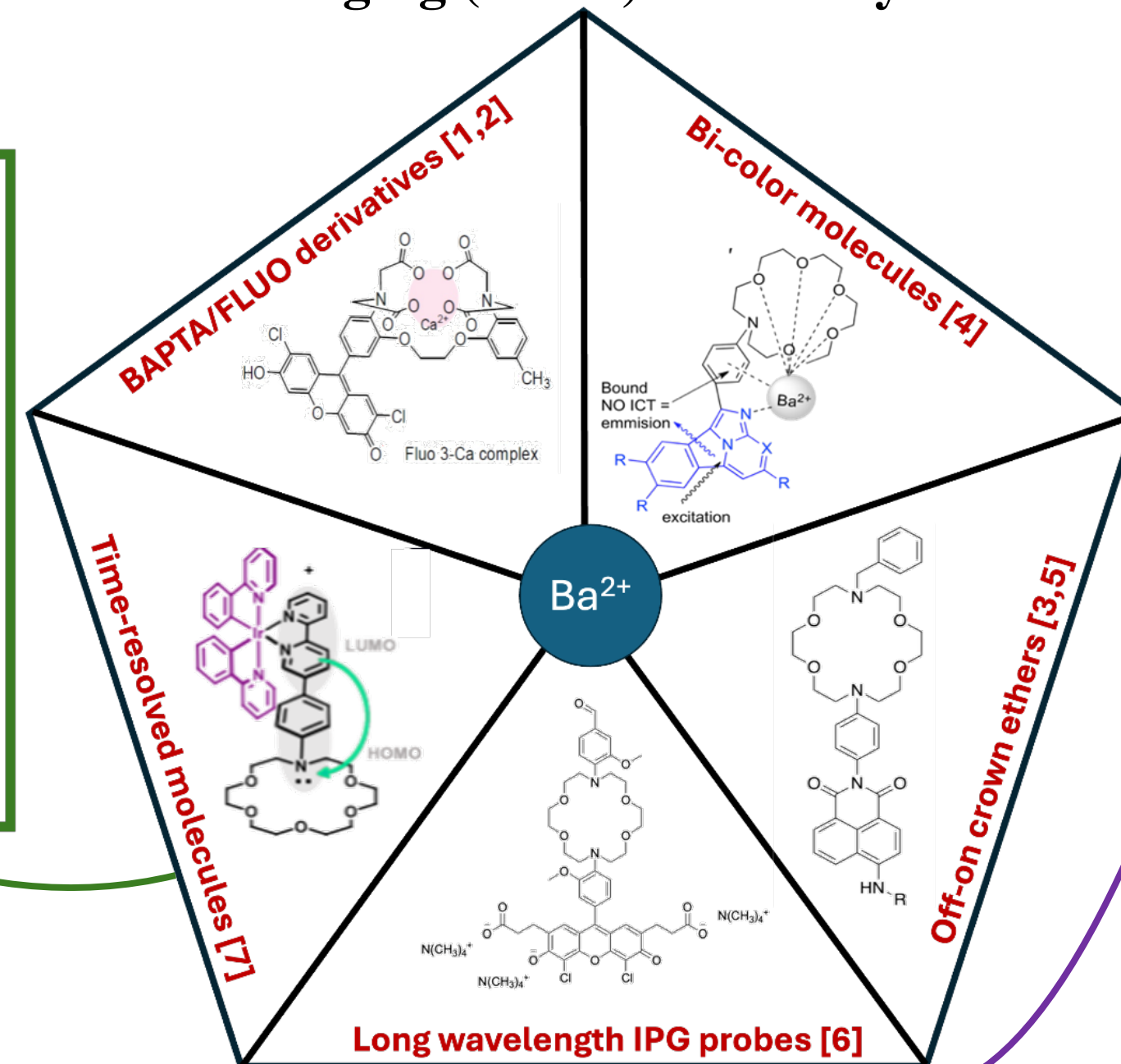
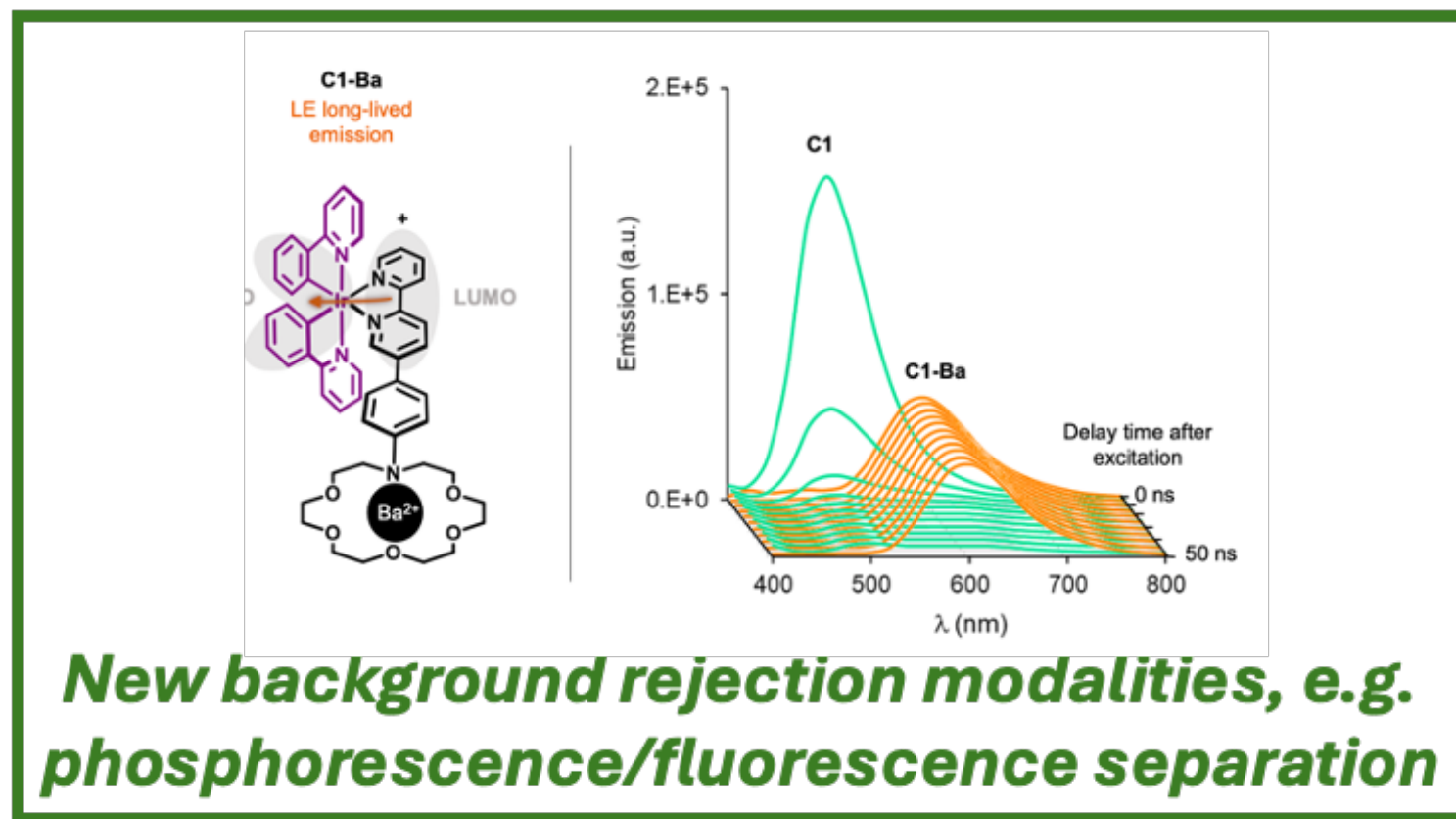


Barium ion is only produced in a true $\beta\beta$ decay, not in any other radioactive event \rightarrow Identification of Ba ion with $\sim 1\%$ FWHM energy measurement would give a **background-free experiment**.

Poster 430: Optical Time Projection Chamber for the Realization of a Ton-Scale Neutrinoless Double Beta Decay Demonstrator by L. Rogers

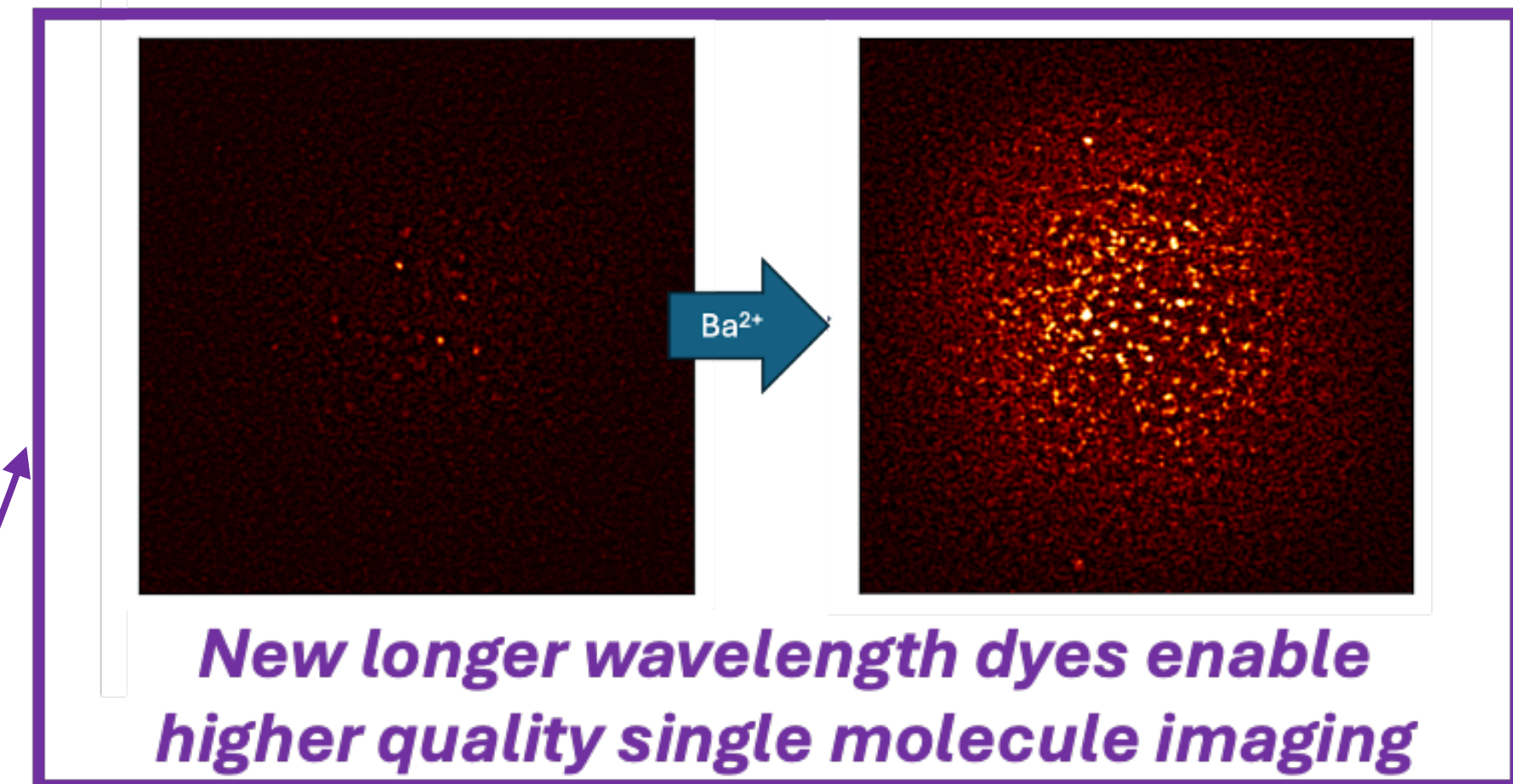
Very recent progress in Ba tagging

The NEXT approach uses single molecule fluorescence imaging (SMFI) to identify barium interacting the specially designed organic molecules.

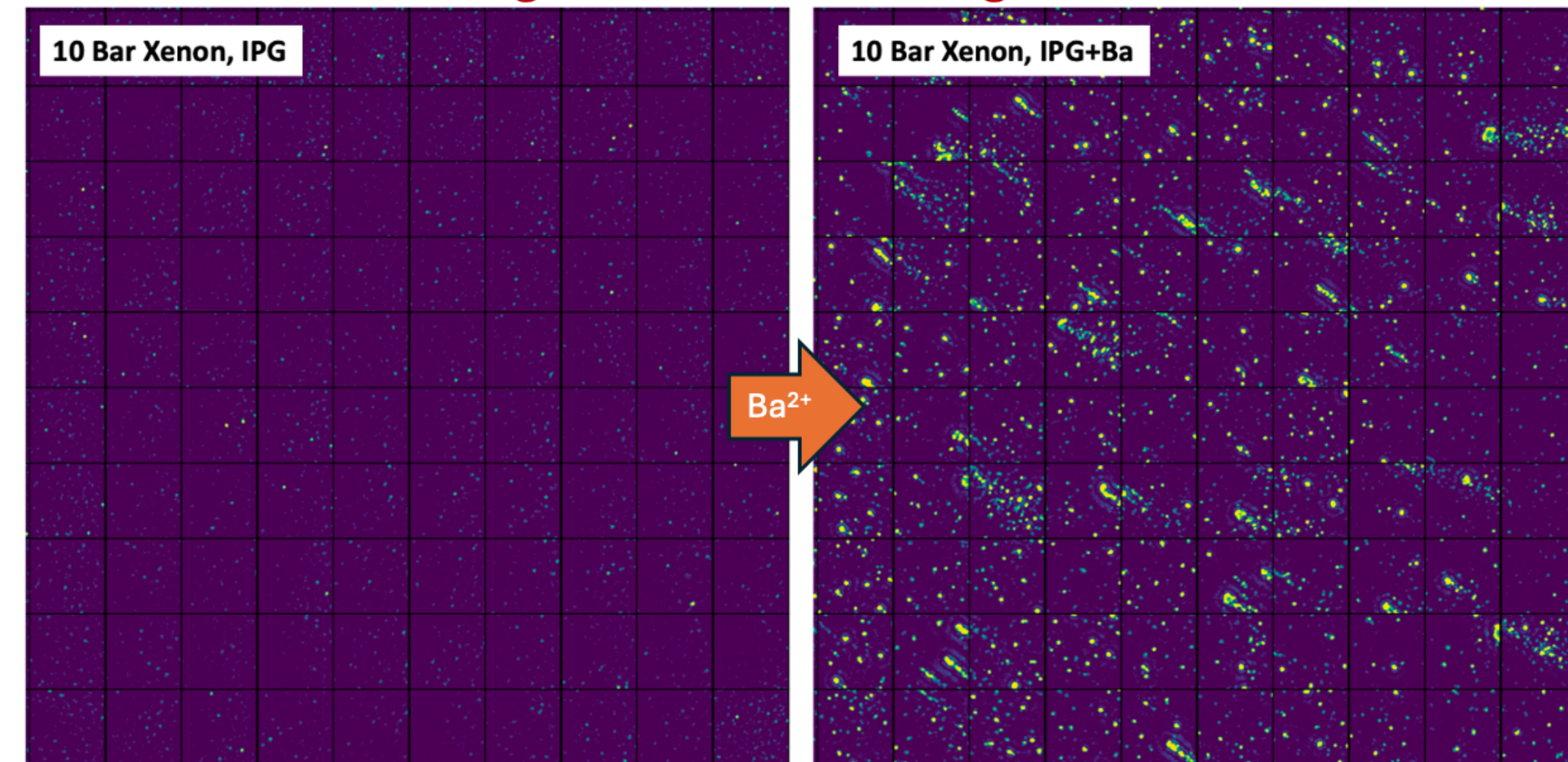


[1]. JINST 11 P12011 (2016)
 [2]. Phys. Rev. Lett. 120, 132504 (2018)
 [3]. Sci Rep 9: 15097 (2019)
 [4]. Nature 583, 48 (2020)

[5]. ACS Sensors 6, 1, 192-20 (2021)
 [6]. [10.26434/chemrxiv-2023-wxpbh](https://doi.org/10.26434/chemrxiv-2023-wxpbh) (2023)
 [7]. Publication in Preparation (2024)



Individual Ba²⁺ ions imaged in 10 bar of xenon gas



NEXT has built microscopy systems capable of imaging individual barium ions in high pressure xenon gas environments.

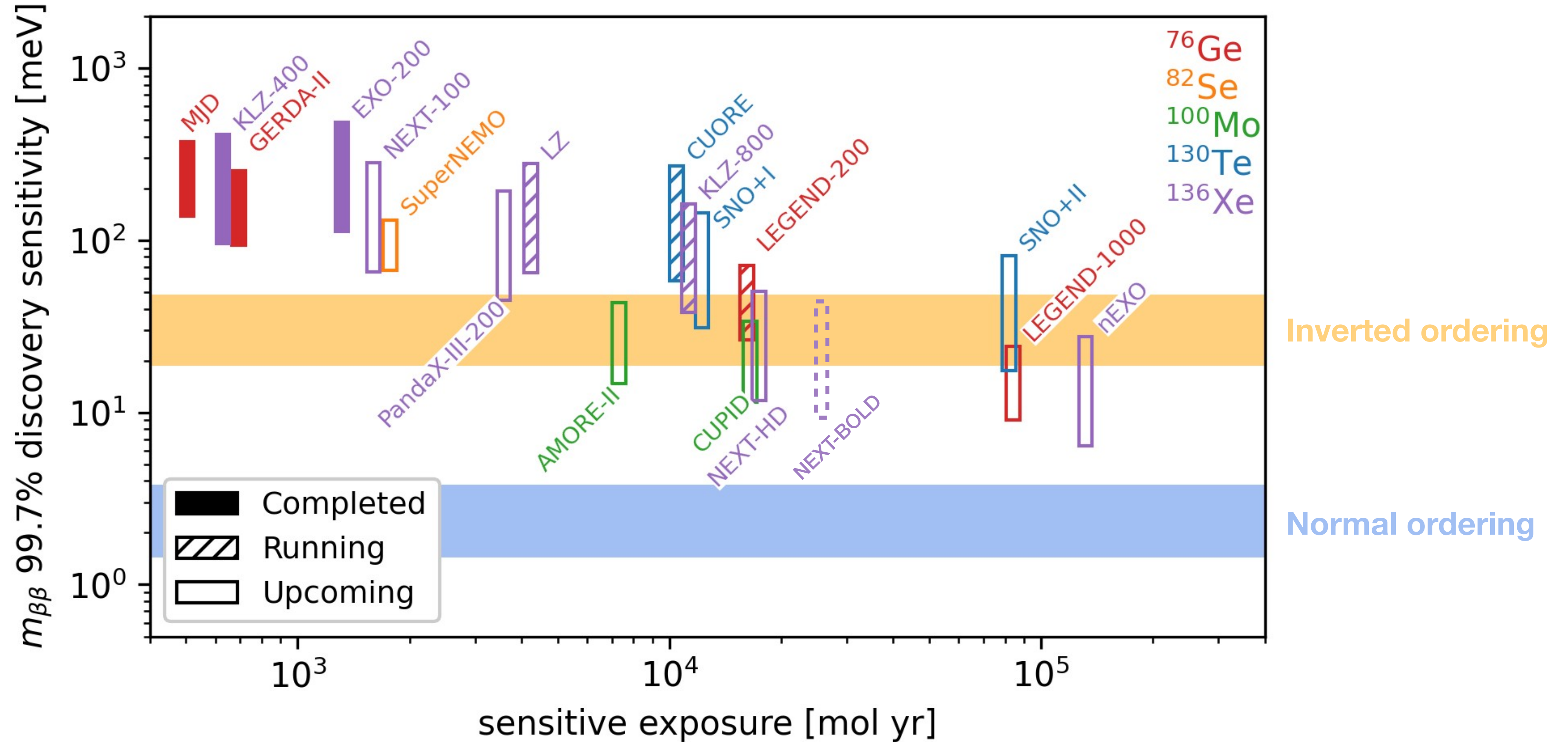
Multidisciplinary effort with physicists, chemists, material scientists

Poster 377: BOLD: three strategies for detecting single Ba ions in NEXT using molecular indicators by P. Herrero Gomez

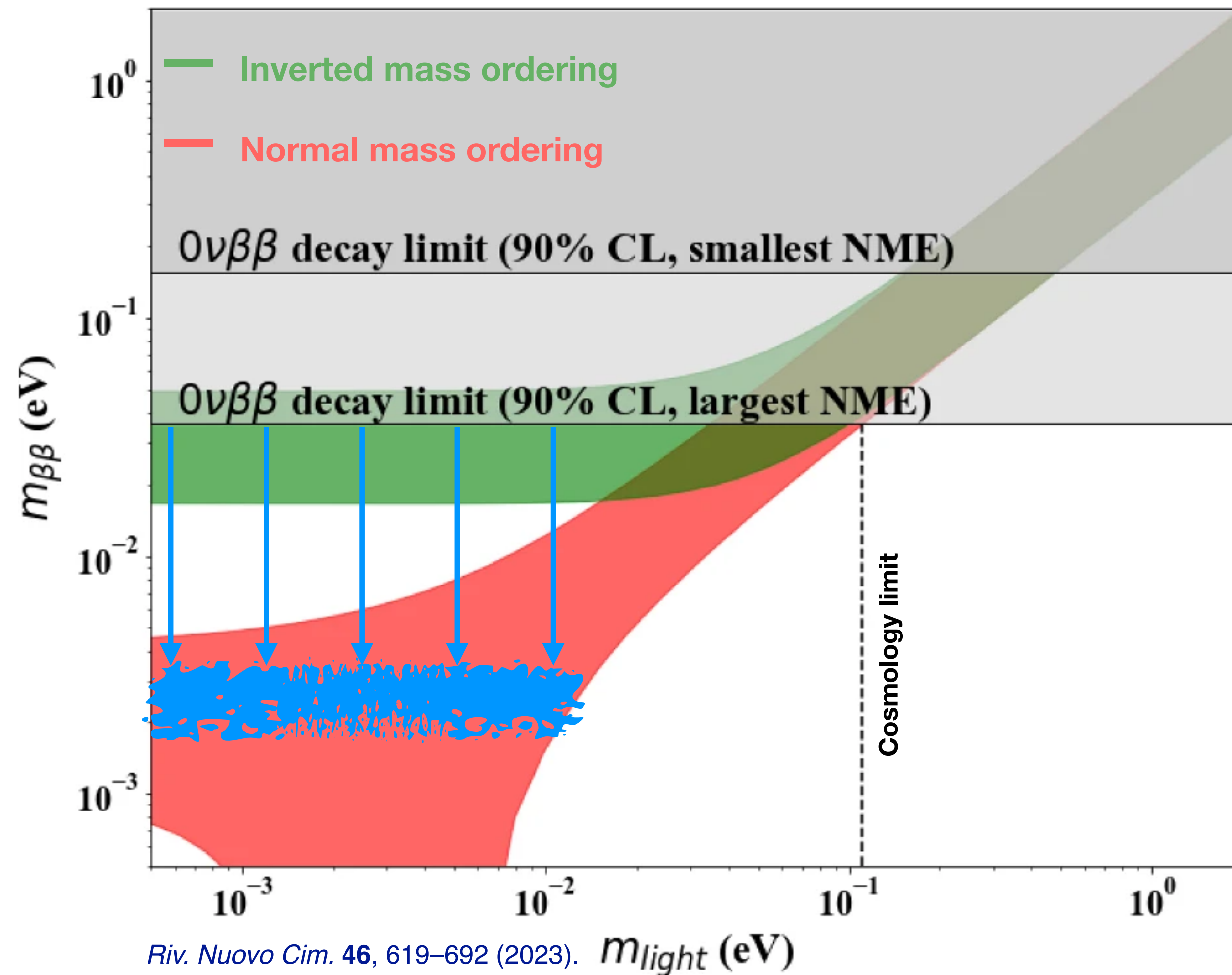
Poster 391: Advancements in Single Barium Ion Capture and Imaging for Barium Tagging Sensors in NEXT Neutrinoless Double Beta Decay Studies by K. Navarro

Conquering the inverted ordering

- Next generation will *cover* the inverted ordering



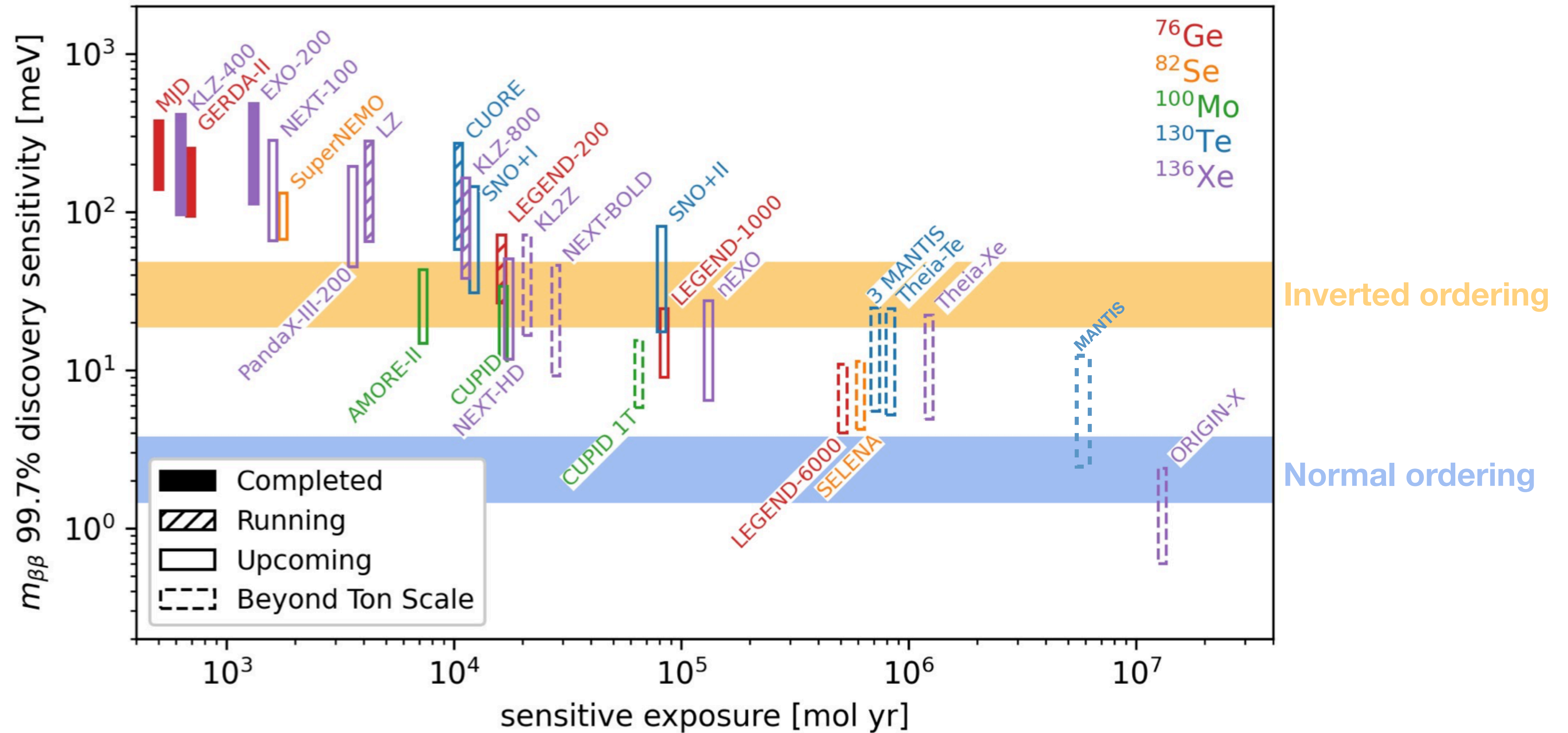
Next next step: Attempt the Normal Ordering



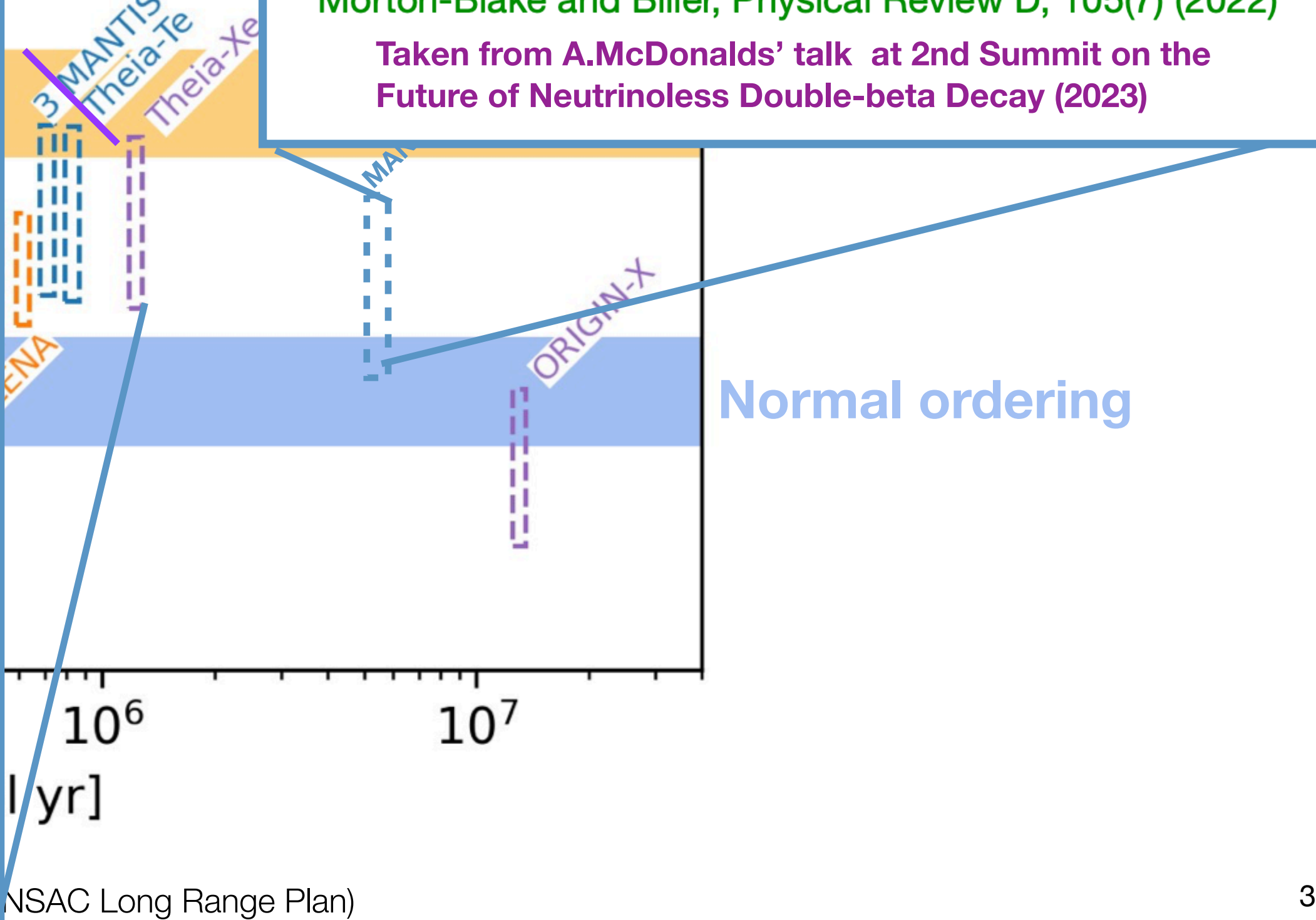
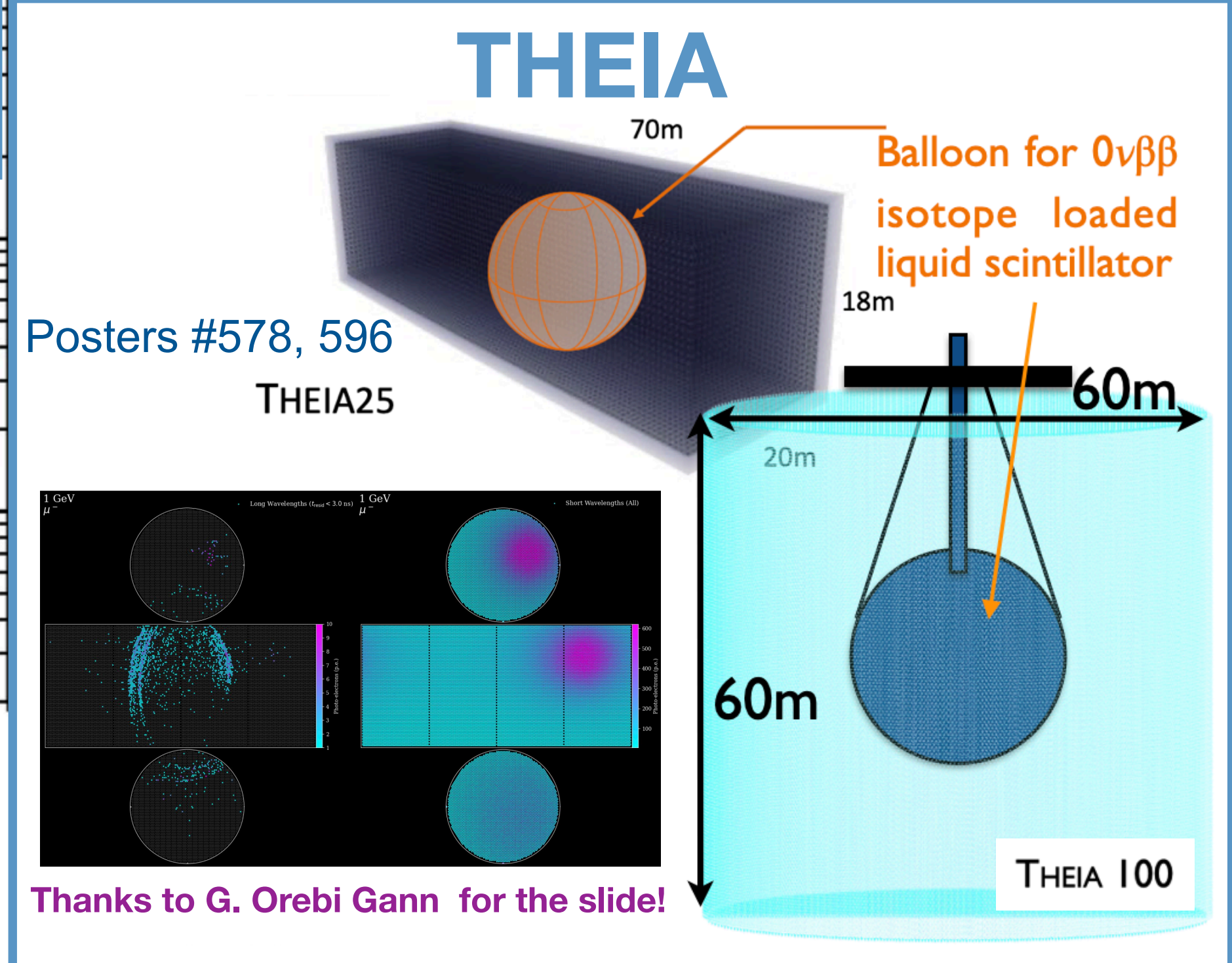
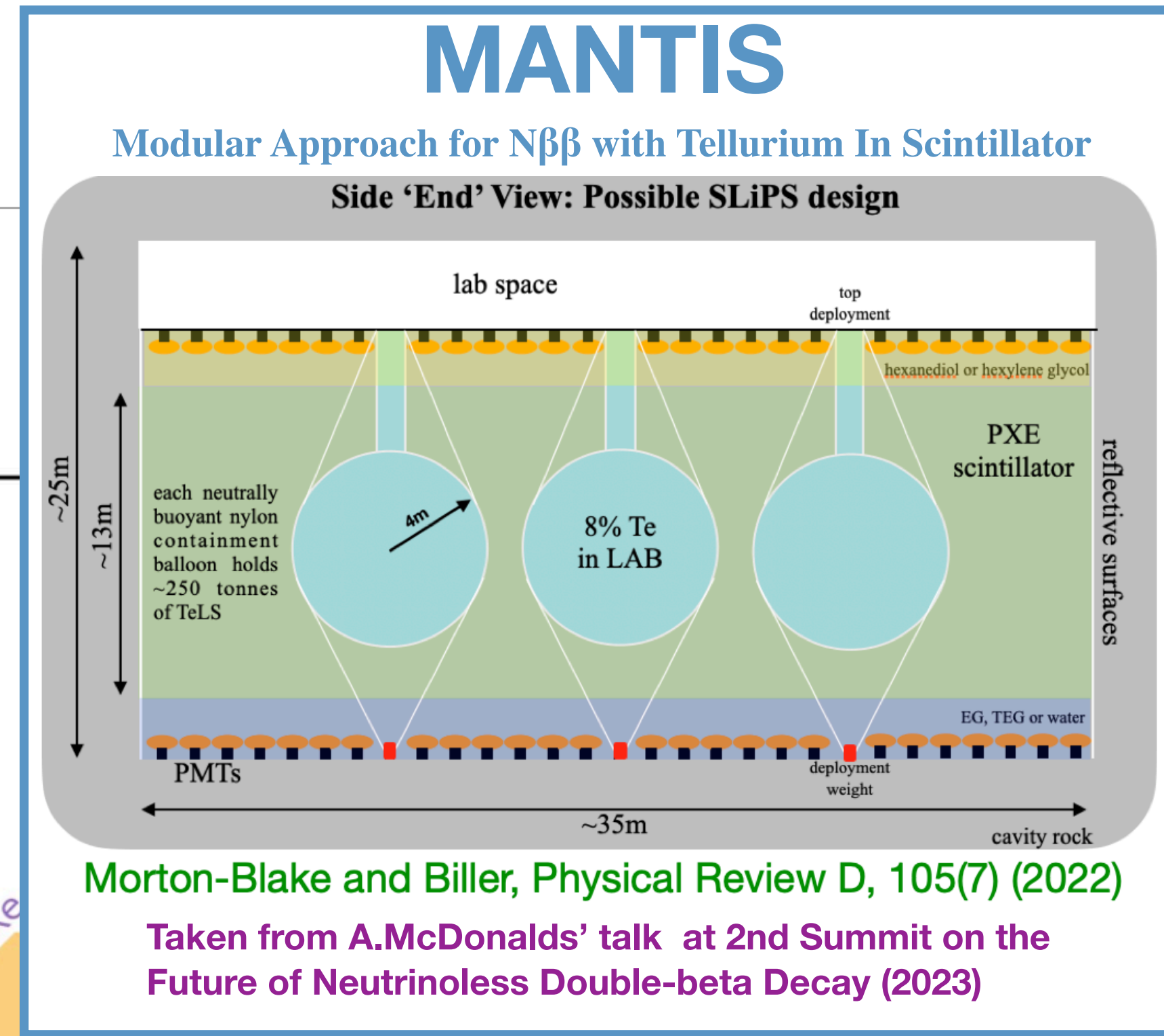
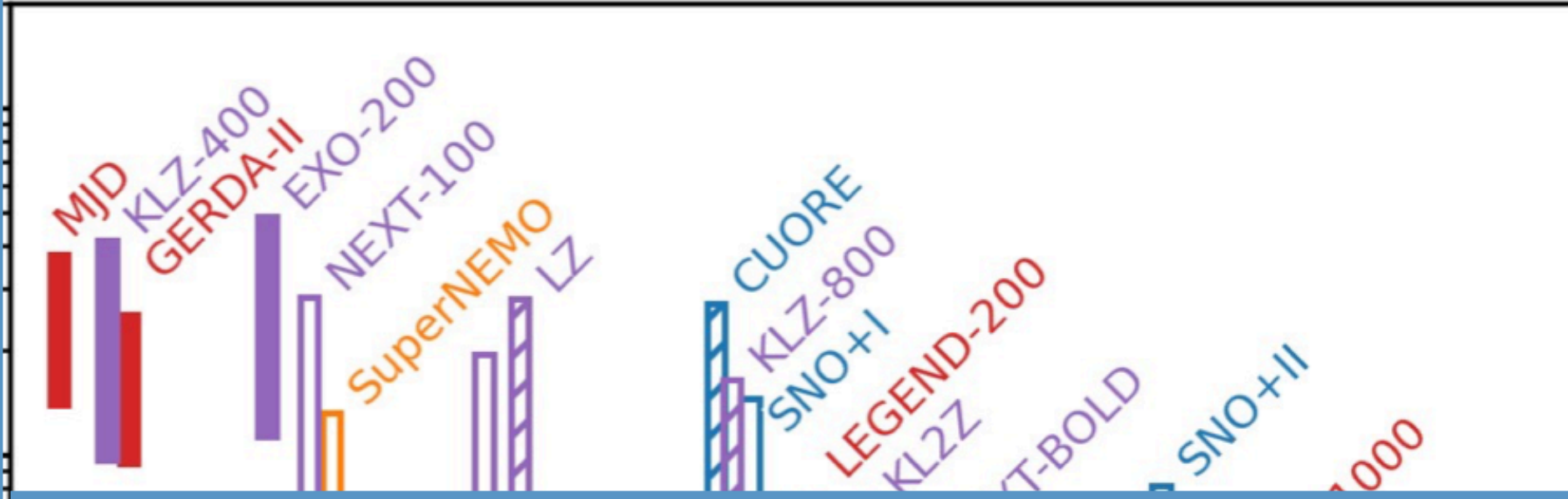
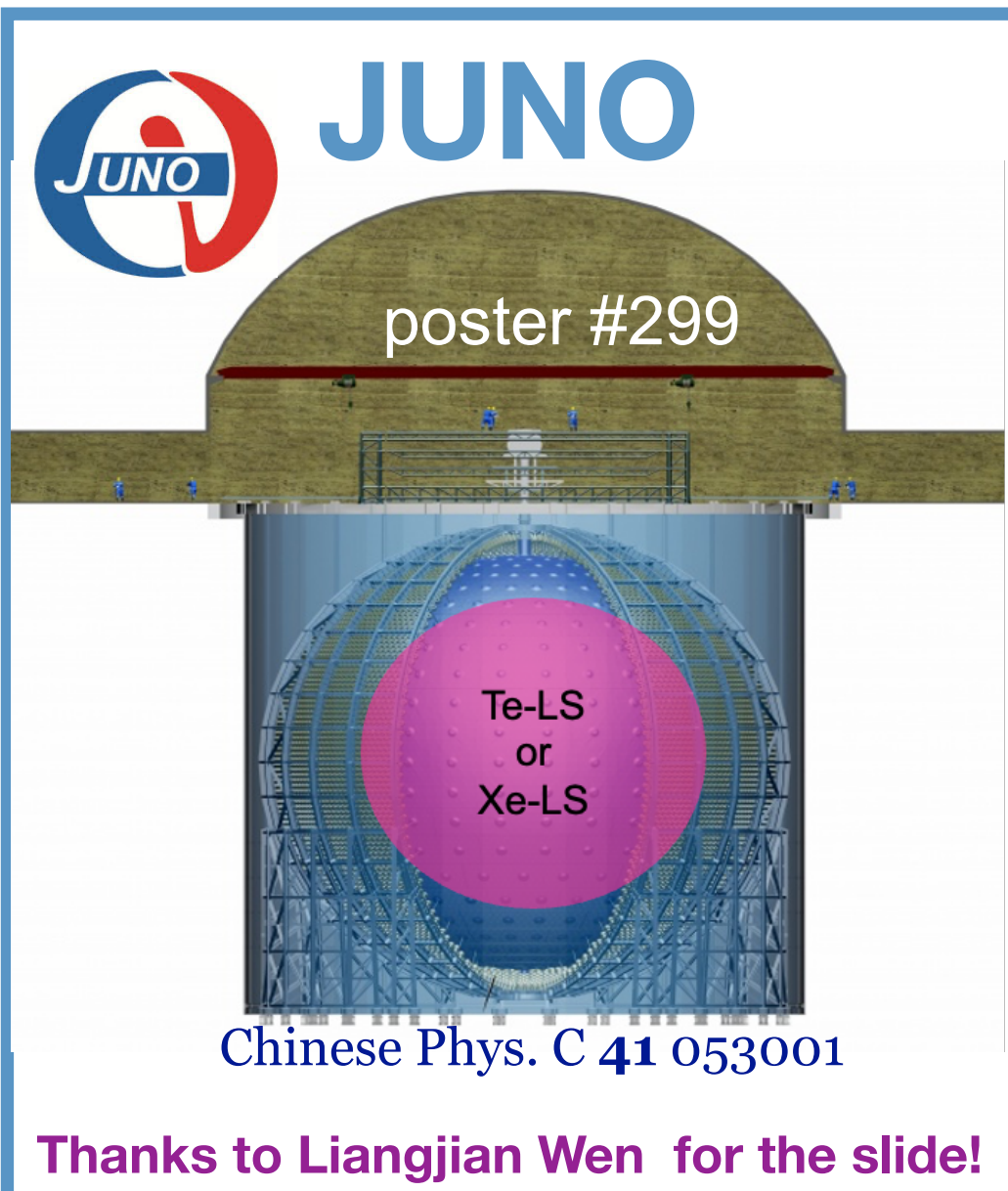
**Need
(significant)
technological advances!**

**** This plot is just one example of parameter space... It does not mean that is exactly where the Majorana neutrino is! ****

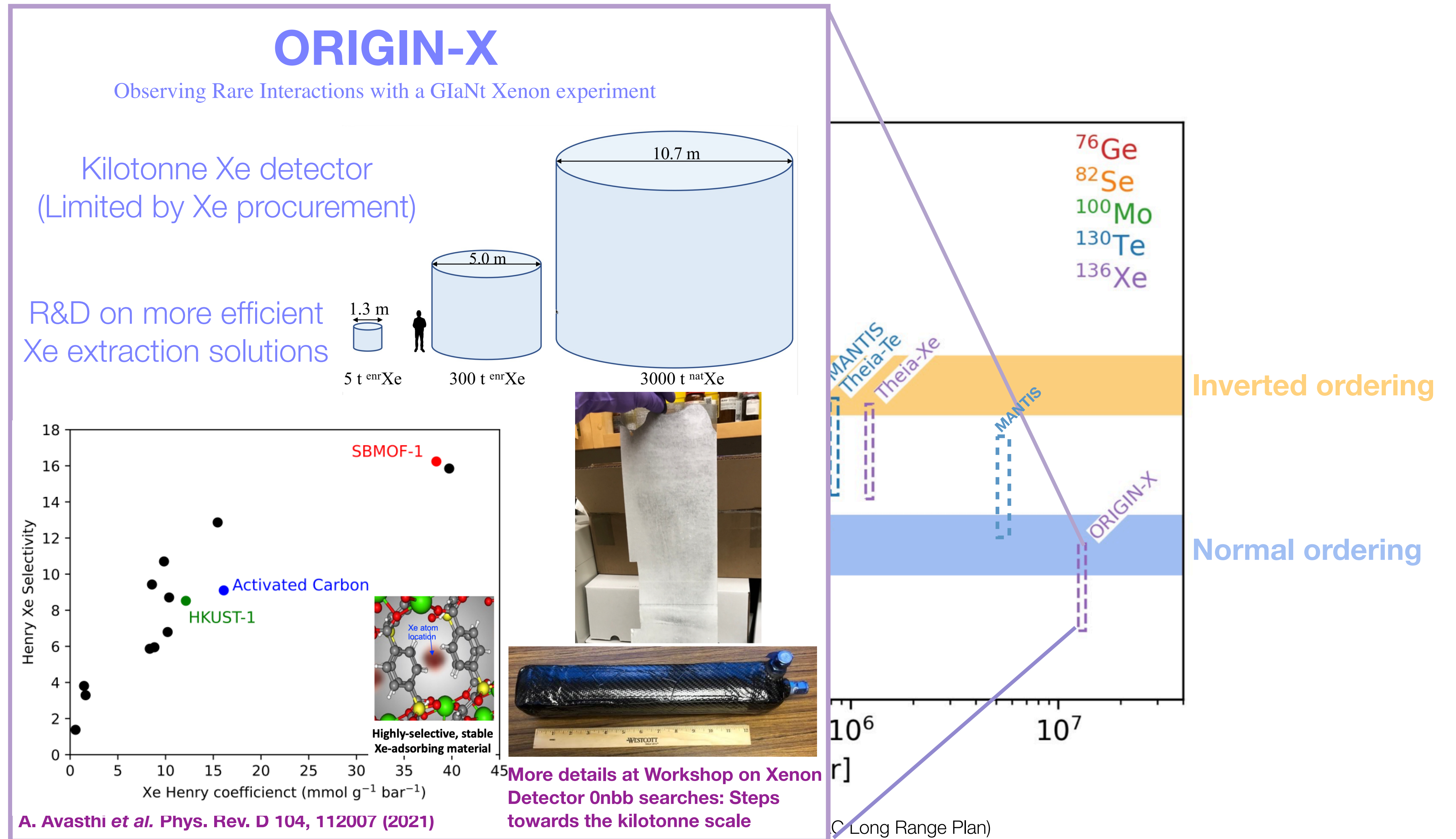
Next next step: Attempt the Normal Ordering



Next next step: Attempt the Normal Ordering



Next next step: Attempt the Normal Ordering



And many more ideas

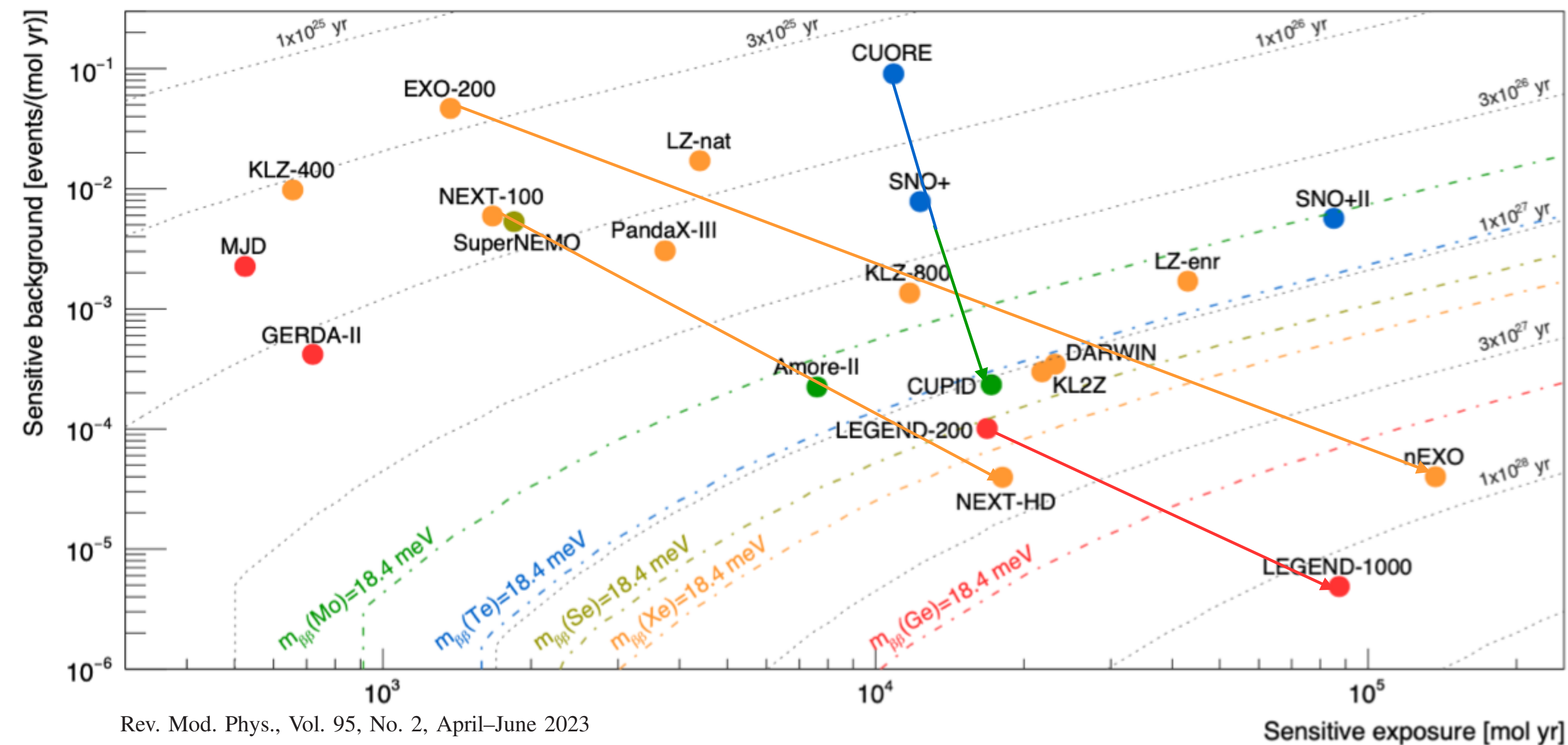
Experiment	Isotope	Mass	Technique	Present Status	Location
CANDLES-III [124]	^{48}Ca	305 kg	$^{nat}\text{CaF}_2$ scint. crystals	Operating	Kamioka
CDEX-1 [125]	^{76}Ge	1 kg	^{enr}Ge semicond. det.	Prototype	CJPL
CDEX-300 ν [125]	^{76}Ge	225 kg	^{enr}Ge semicond. det.	Construction	CJPL
LEGEND-200 [16]	^{76}Ge	200 kg	^{enr}Ge semicond. det.	Commissioning	LNGS
LEGEND-1000 [16]	^{76}Ge	1 ton	^{enr}Ge semicond. det.	Proposal	
CUPID-0 [19]	^{82}Se	10 kg	Zn^{enr}Se scint. bolometers	Prototype	LNGS
SuperNEMO-Dem [126]	^{82}Se	7 kg	^{enr}Se foils/tracking	Operation	Modane
SuperNEMO [126]	^{82}Se	100 kg	^{enr}Se foils/tracking	Proposal	Modane
Selena [127]	^{82}Se		^{enr}Se . CMOS	Development	
IFC [128]	^{82}Se		ion drift SeF_6 TPC	Development	
CUPID-Mo [17]	^{100}Mo	4 kg	$\text{Li}^{enr}\text{MoO}_4$, scint. bolom.	Prototype	LNGS
AMoRE-I [129]	^{100}Mo	6 kg	$^{40}\text{Ca}^{100}\text{MoO}_4$ bolometers	Operation	YangYang
AMoRE-II [129]	^{100}Mo	200 kg	$^{40}\text{Ca}^{100}\text{MoO}_4$ bolometers	Construction	Yemilab
CROSS [130]	^{100}Mo	5 kg	$\text{Li}_2^{100}\text{MoO}_4$, surf. coat bolom.	Prototype	Canfranc
BINGO [131]	^{100}Mo		$\text{Li}^{enr}\text{MoO}_4$	Development	LNGS
CUPID [28]	^{100}Mo	450 kg	$\text{Li}^{enr}\text{MoO}_4$, scint. bolom.	Proposal	LNGS
China-Europe [132]	^{116}Cd		$^{enr}\text{CdWO}_4$ scint. crystals	Development	CJPL
COBRA-XDEM [133]	^{116}Cd	0.32 kg	^{nat}Cd CZT semicond. det.	Operation	LNGS
Nano-Tracking [134]	^{116}Cd		$^{nat}\text{CdTe}$. det.	Development	
TIN.TIN [135]	^{124}Sn		Tin bolometers	Development	INO
CUORE [10]	^{130}Te	1 ton	TeO_2 bolometers	Operating	LNGS
SNO+ [136]	^{130}Te	3.9 t	0.5-3% ^{nat}Te loaded liq. scint.	Commissioning	SNOLab
nEXO [29]	^{136}Xe	5 t	Liq. ^{enr}Xe TPC/scint.	Proposal	
NEXT-100 [137]	^{136}Xe	100 kg	gas TPC	Construction	Canfranc
NEXT-HD [137]	^{136}Xe	1 ton	gas TPC	Proposal	Canfranc
AXEL [138]	^{136}Xe		gas TPC	Prototype	
KamLAND-Zen-800 [13]	^{136}Xe	745 kg	^{enr}Xe dissolved in liq. scint.	Operating	Kamioka
KamLAND2-Zen [41]	^{136}Xe		^{enr}Xe dissolved in liq. scint.	Development	Kamioka
LZ [139]	^{136}Xe	600 kg	Dual phase Xe TPC, nat./enr. Xe	Operation	SURF
PandaX-4T [119]	^{136}Xe	3.7 ton	Dual phase nat. Xe TPC	Operation	CJPL
XENONnT [140]	^{136}Xe	5.9 ton	Dual phase Xe TPC	Operating	LNGS
DARWIN [141]	^{136}Xe	50 ton	Dual phase Xe TPC	Proposal	LNGS
R2D2 [142]	^{136}Xe		Spherical Xe TPC	Development	
LAr TPC [143]	^{136}Xe	kton	Xe-doped LR TPC	Development	
NuDot [144]	Various		Cherenkov and scint. in liq. scint.	Development	
THEIA [145]	Xe or Te		Cherenkov and scint. in liq. scint.	Development	
JUNO [146]	Xe or Te		Doped liq. scint.	Development	
Slow-Fluor [147]	Xe or Te		Slow Fluor Scint.	Development	

I did not have time to discuss these, but they all have unique features....

Opportunistic searches with Dark Matter experiments: L. Baudis' talk on Friday

Reminder of the challenges ahead

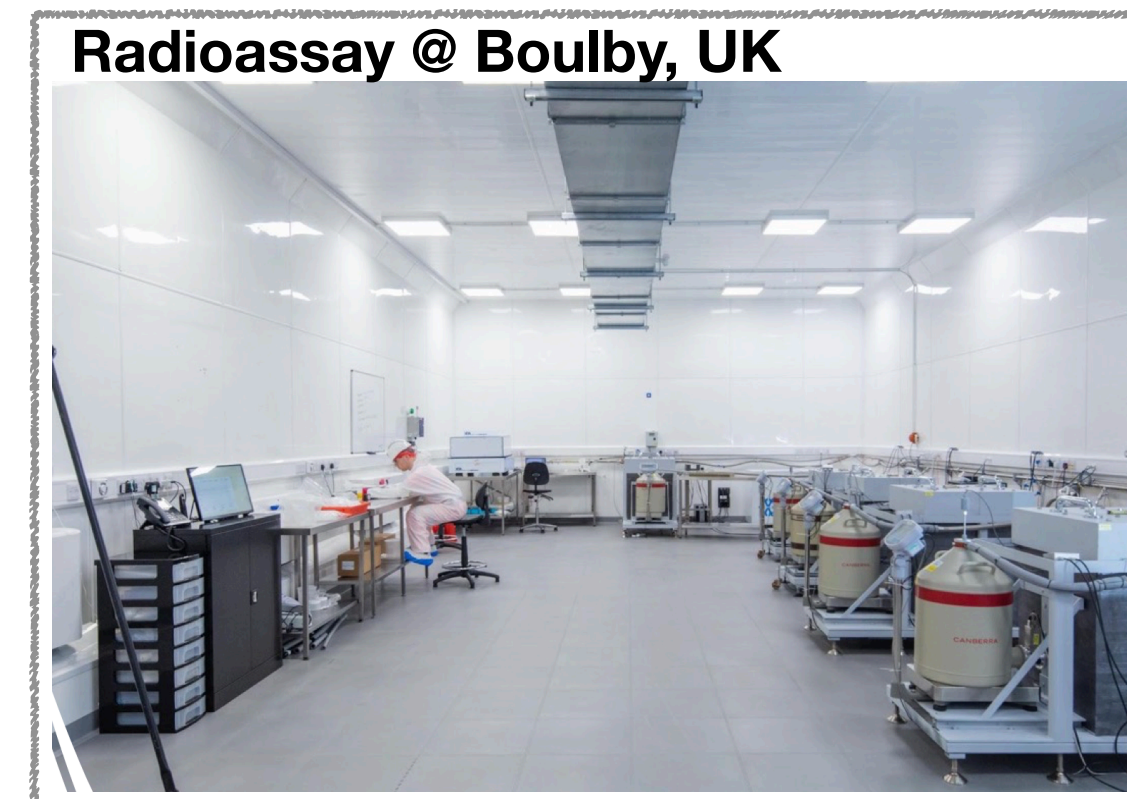
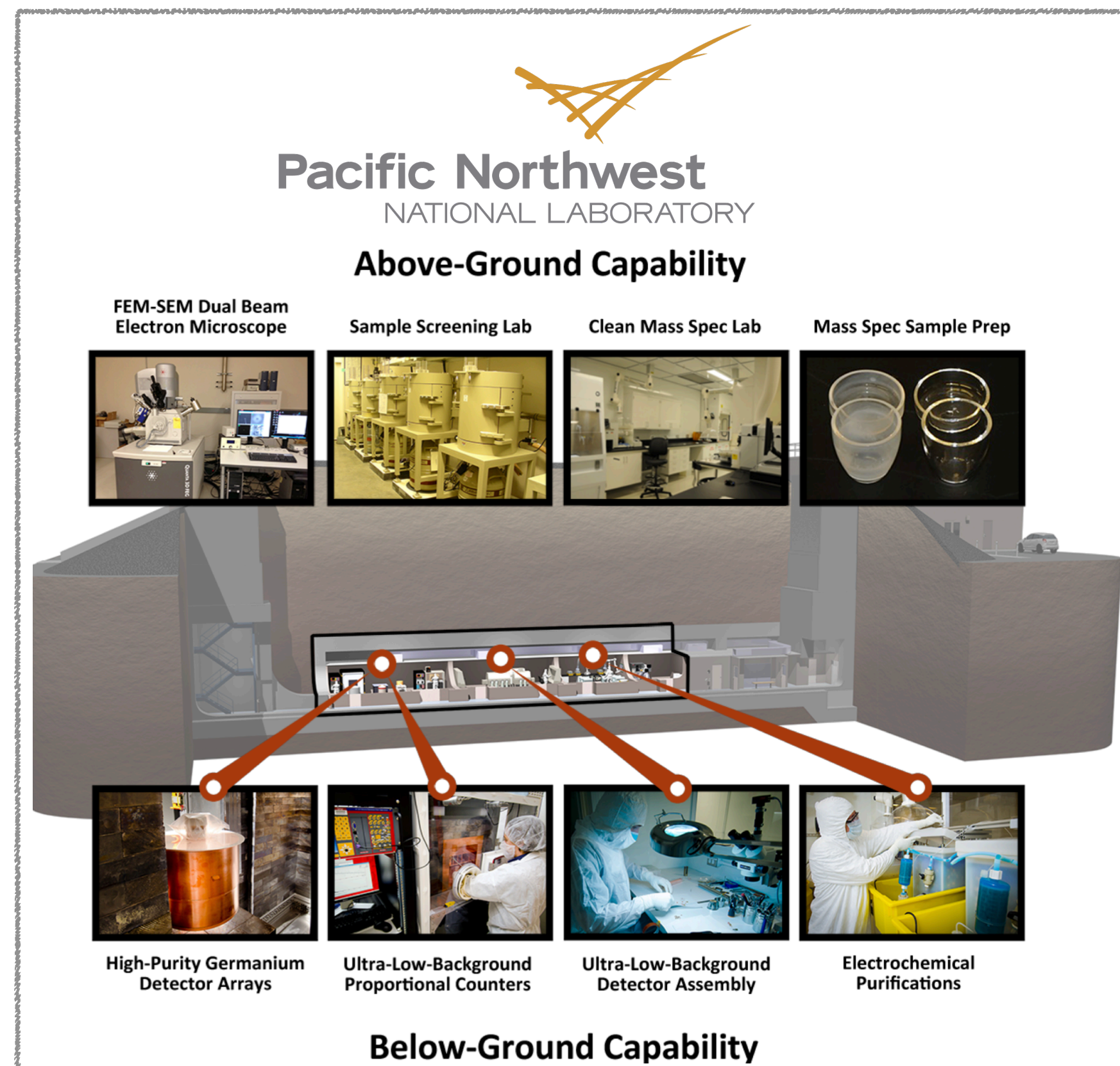
- Very ambitious (technical) goals for the ton-scale experiments



We need to continue our R&D efforts to develop mature technologies

Reminder of the challenges ahead

- Very ambitious (technical) goals for the ton-scale experiments
- Radiopurity requirements are pushing the limits of radioassay capabilities



We can build a strong network of facilities

Reminder of the challenges ahead

- Very ambitious (technical) goals for the ton-scale experiments
- Radiopurity requirements are pushing the limits of radioassay capabilities
- Working with industry is becoming more and more challenging (and expensive)
- Funding situations for fundamental research across the globe is much lower than our scientific ambitions



Let's avoid to duplicate efforts

Reminder of the challenges ahead

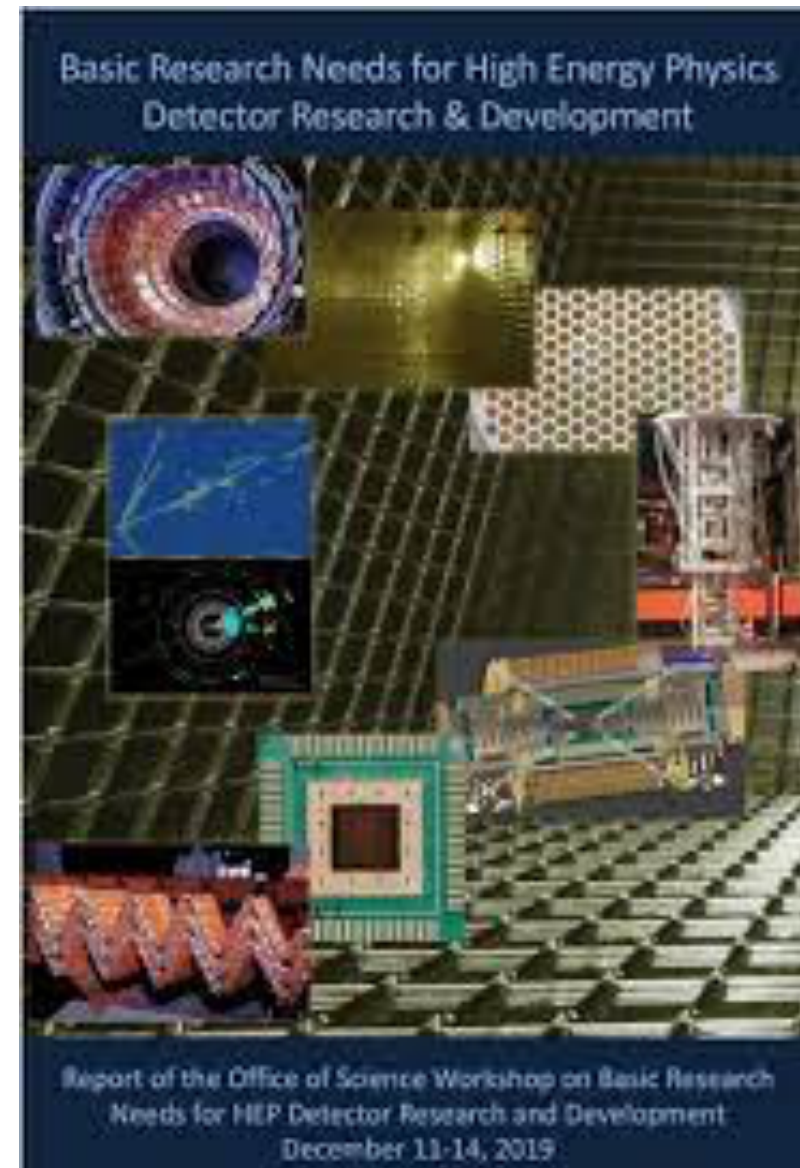
- Very ambitious (technical) goals for the ton-scale experiments
- Radiopurity requirements are pushing the limits of radioassay capabilities
- Working with industry is becoming more and more challenging (and expensive)
- Funding situations for fundamental research across the globe is much lower than our scientific ambitions
- Community(ies) need to come together to be stronger, while pushing on new technology development



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Pushing the instrumentation frontier together

- Need for instrumentation development recognized worldwide



Basic Research Needs for HEP Detector R&D (2019)



ECFA Detector R&D Roadmap (2021)



**SNOWMASS Process (2021)
→ P5 Report (2023)**

And more....

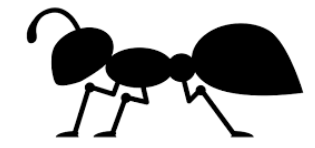
Pushing the instrumentation frontier together

- Need for instrumentation development recognized worldwide
- Creation of new Detector R&D Collaborations (2023) (approved by CERN council)
- Example: DRD2 - Liquid Detectors

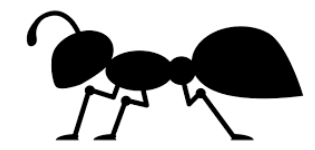


Consider joining the DRD efforts!

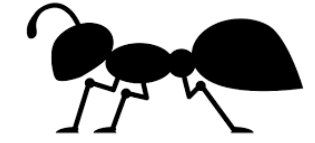
Closing remarks



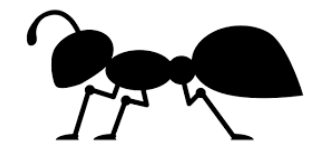
Understanding the nature of neutrinos is of the utmost importance



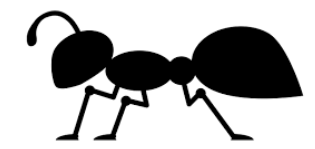
The next generation of experiments will *cover* the inverted neutrino mass ordering, but it will be technologically challenging



To reach the normal neutrino mass ordering, transformative technologies are needed



We need to continue to cast a wide net and explore different isotopes and technologies



Given the technological limitations, the community needs to work together to avoid duplication of scarce resources

Closing remarks



THANK YOU !

