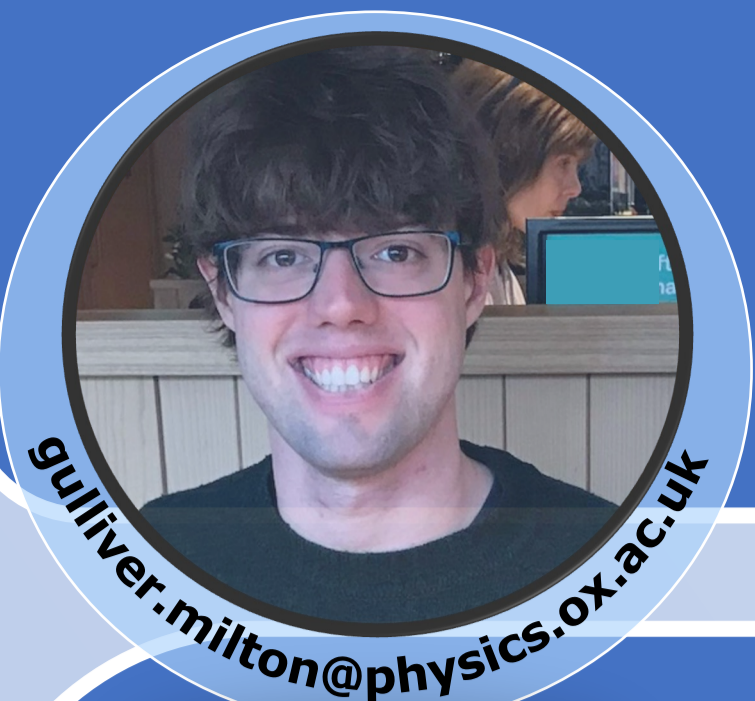


# First Indications of CC Solar Neutrino Interactions on Carbon-13

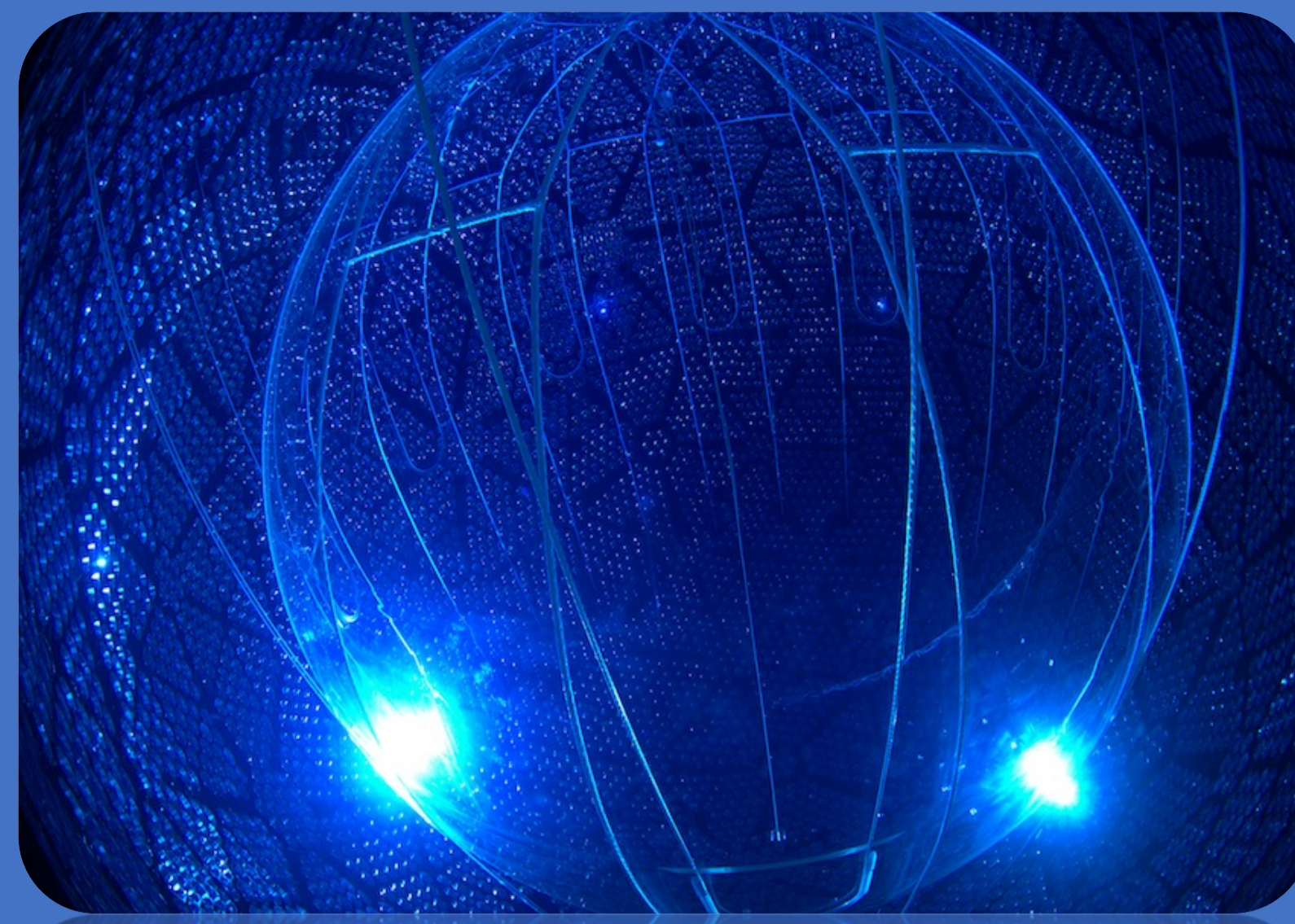


Gulliver Milton on behalf of the SNO+ collaboration



## SNO+ experiment

- **Neutrino detector** located 2 km underground at SNOLAB, in Canada
- Acrylic vessel filled with **780 tonnes of liquid scintillator**
- Ionising radiation generates **scintillation light**
- The light is recorded by the 9362 PMTs
- **The muon rate is ~3 per hour**



Physics goals include the search for **neutrinoless double beta decay** from tellurium-130 and measurements of the reactor, solar, supernovae and, geo neutrinos

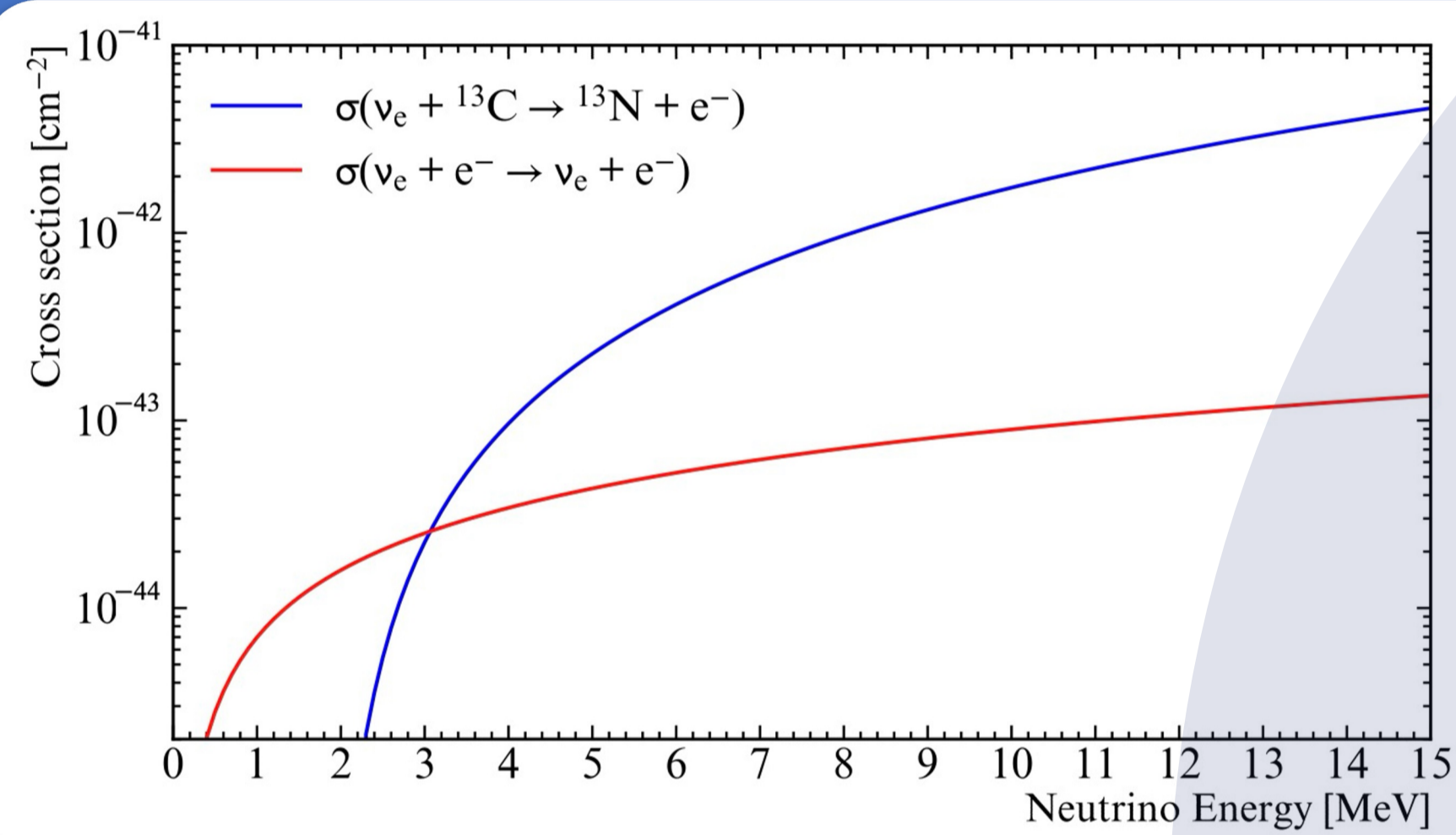
## CC on Carbon-13

- The **1.1% natural abundance** of  $^{13}\text{C}$  in the organic scintillator can undergo a charged current interaction with neutrinos
- The **CC reaction** on  $^{13}\text{C}$  has a threshold of 2.2 MeV and produces a  $^{13}\text{N}$ , which then decays
- The **coincidence nature of the event** significantly **reduces the background**
- **Solar neutrinos** produced in  $^8\text{B}$  decays from the pp chain in the Sun arrive at Earth with a **flux of**:

$$\Phi_{^8\text{B}} = (5.20_{-0.1}^{+0.1}) \times 10^6 \text{ cm}^{-2}\text{s}^{-1} [1]$$

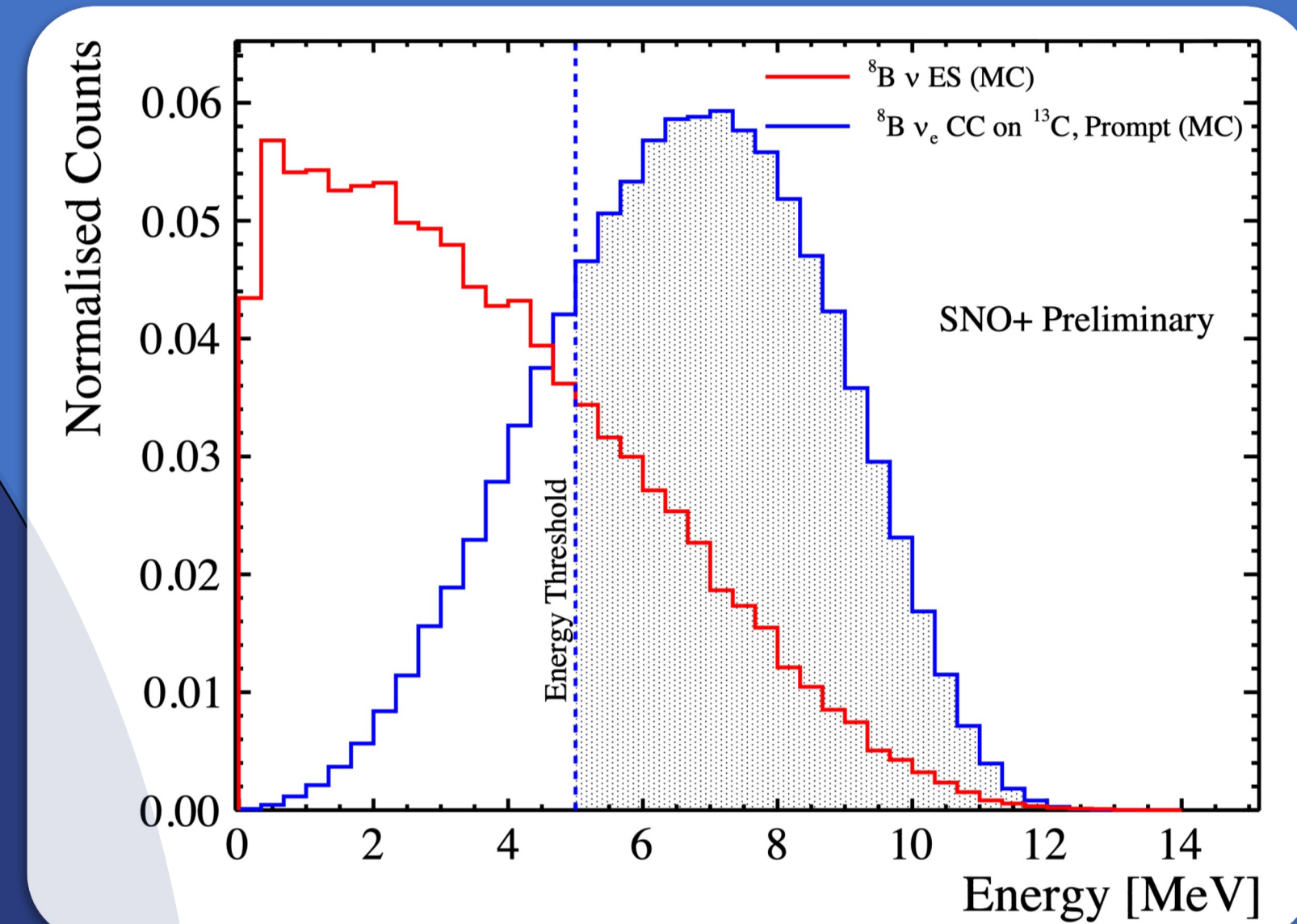
## Signal

- The **cross-section** of the interaction is **orders of magnitude larger** than the **electron ES process** [2]
- In SNO+ there are expected **22 ev/yr/kT**  $^8\text{B}$  neutrino CC interactions with  $^{13}\text{C}$
- This is obtained by integrating over the **theoretical cross-section** [2] and **SSM  $^8\text{B}$  neutrino flux** [1], assuming the **globally fit neutrino oscillations parameters** [3]



## Prompt Event

- **Electron** with energy =  $E(v_e) - 2.2 \text{ MeV}$
- Imposing a **5 MeV cut** removes most background, for example, the **Thallium-208 decay** ( $Q = 5 \text{ MeV}$ )
- The remaining **prompt background** is  $^8\text{B}$  **Elastic scattering**



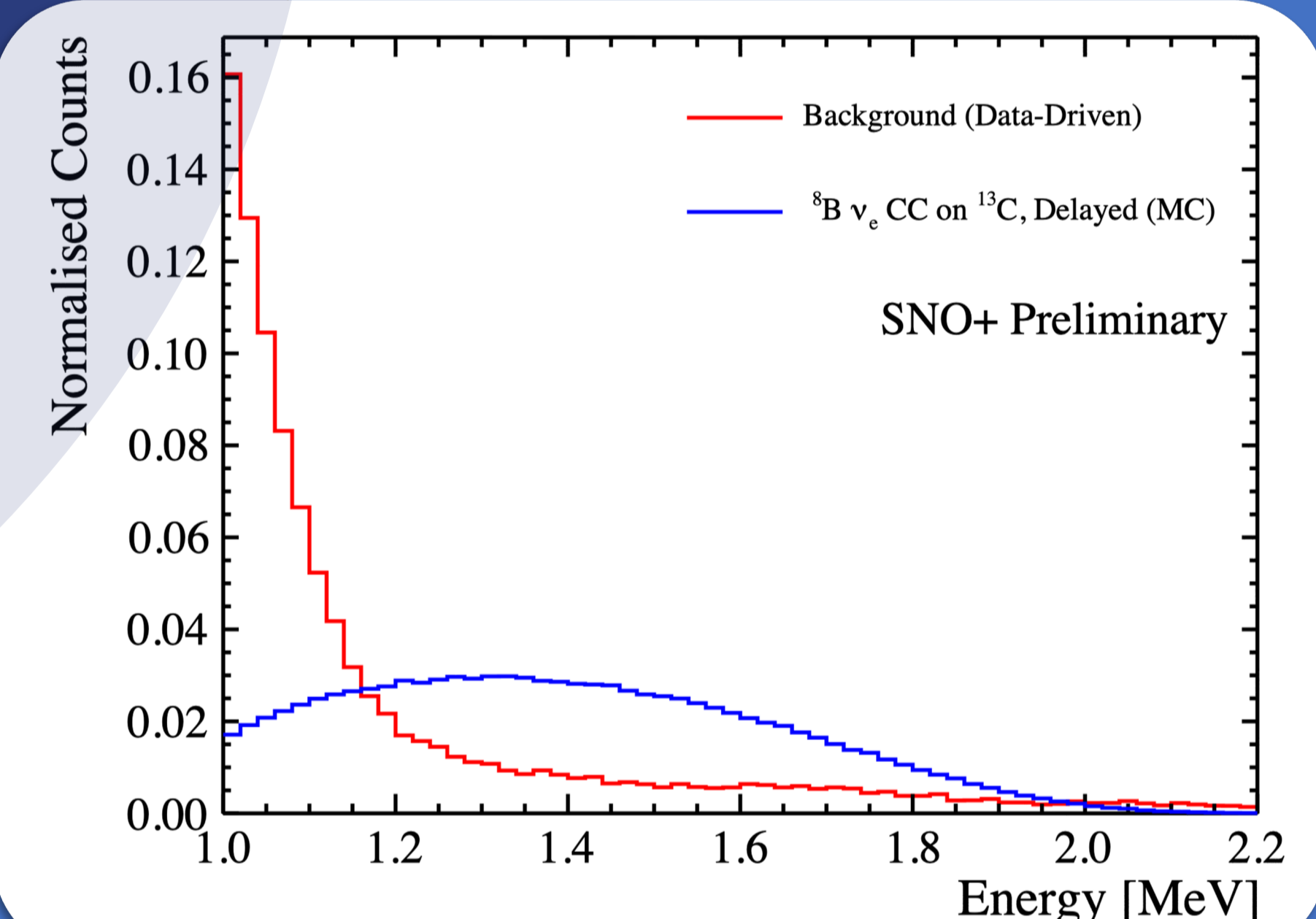
## Backgrounds

- The **dominant background** is from **accidental coincidences**
- Spurious prompt events were used to produce **"fake coincidences"** with **data events** satisfying delayed event cuts
- The **fraction of events** resulting in fake coincidences allows the **random coincidence rate** to be determined
- Given the **low muon flux** at SNO+, modest muon followers cuts **reduce the cosmogenic background** ( $^{11}\text{Be}$  prompt +  $^{11}\text{C}$  delayed) to a **negligible  $7.0 \times 10^{-4}$  ev/yr/kT**

Livetime: 170 days (pre-cuts), 150 days (post-cuts)

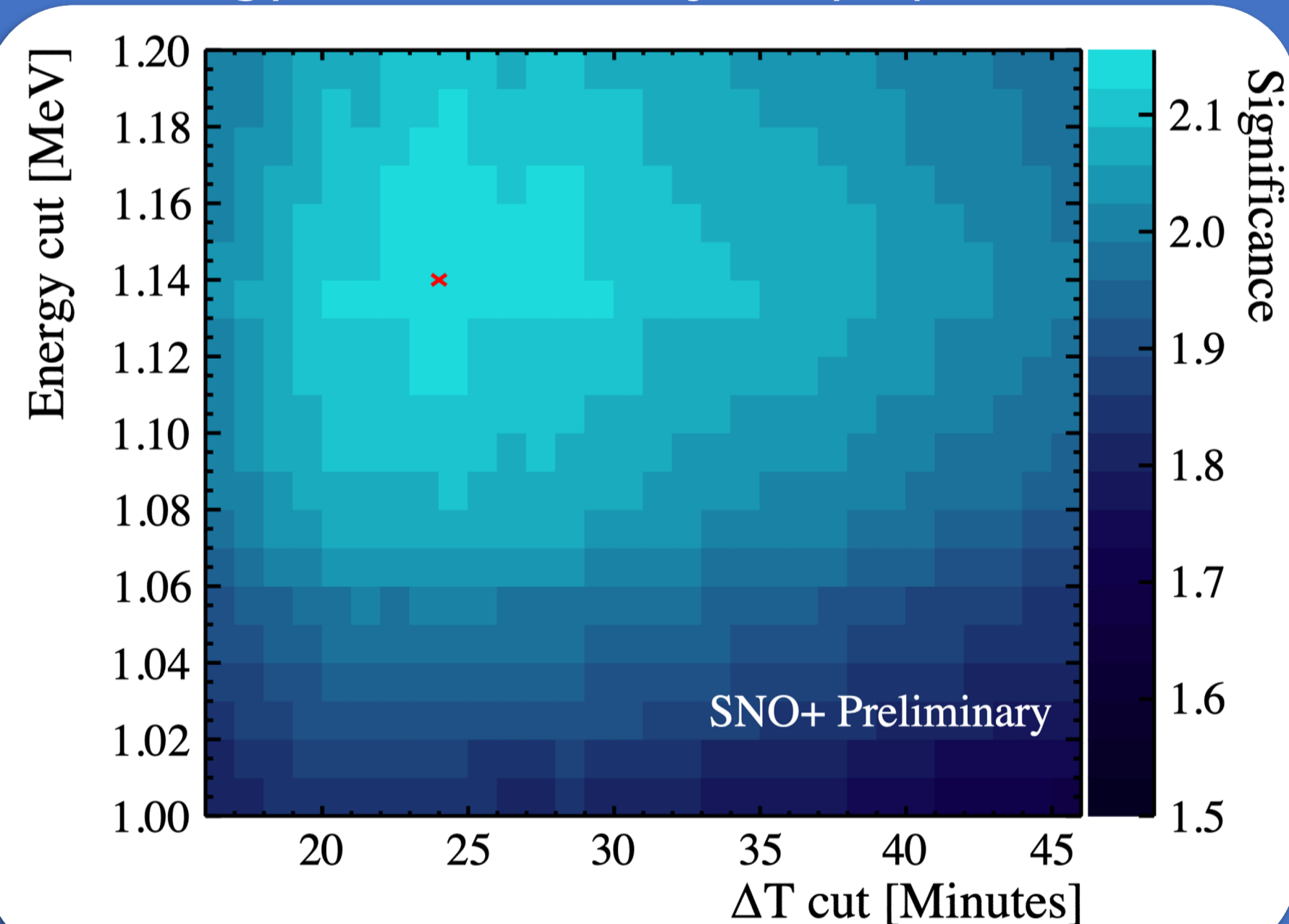
## Delayed Event

- A **positron** with energy: [1.0 to 2.2] MeV
- The annihilation produces two gammas which are detected



## Cuts-Based Analysis

- **Data-driven background rates**
- The fiducial volume,  $\Delta R$ ,  $\Delta T$  and delayed energy window were jointly optimised



- Predetermined cuts**
- $\text{FV} < 5.3 \text{ m}$
  - $\Delta R < 0.36 \text{ m}$
  - $0.01 < \Delta T \text{ (min)} < 24$
  - $1.14 < E_{e^+} \text{ (MeV)} < 2.2$
- These cuts give an **expected background number of 0.31** and a **signal of 1.83**
- Two events observed**

## Likelihood Analysis

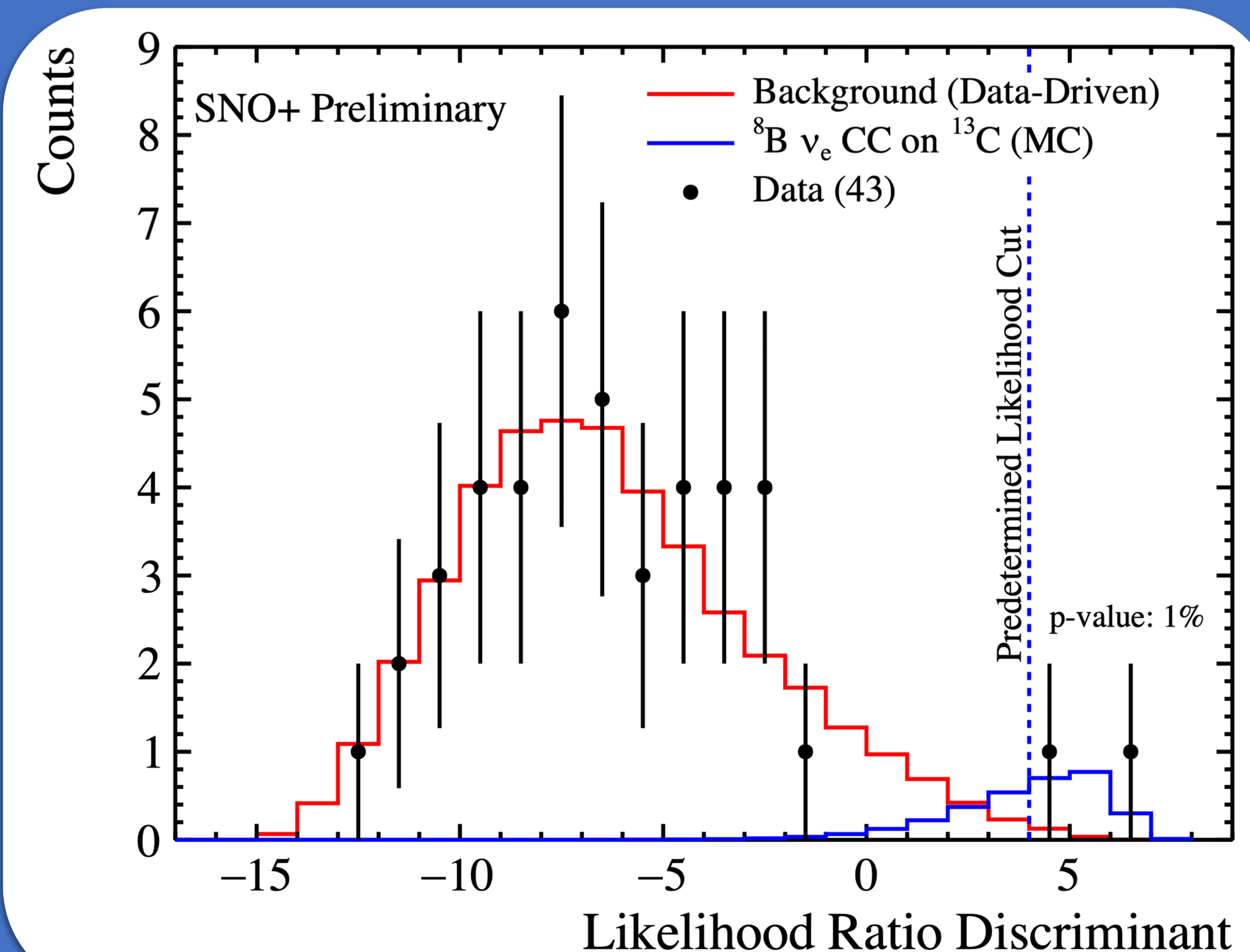
Using **PDFs** of the **delayed energy** (1 to 2.2 MeV),  $\Delta T$  (0.01 to 60 min), and  $\Delta R$  (<1 m), the Likelihood ratio can be constructed

A **predetermined cut** that maximise significance gives an **expected background number of 0.17** and a **signal of 1.79**

**Two events observed**, giving a **background fluctuation probability of 1%**

The successfully **validated background model** gives **confidence in the extrapolation of sensitivity**

Analysis is being continued on an expanded data set



[1] Borexino Collaboration, M. Agostini et al., Comprehensive measurement of pp-chain solar neutrinos, Nature 562, 2018

[2] T. Suzuki, A. B. Balantekin, and T. Kajino, Neutrino capture on  $^{13}\text{C}$  Phys. Rev. C 86, 015502, 2012

[3] P. A. Zyla et al. (Particle Data Group), Review of Particle Physics, Prog. Theor. Exp. Phys, 083C01, 2020