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# Event by Event Classification of $(\alpha, n)$ and IBD Interactions at SNO+



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on behalf of the SNO+ collaboration

## The SNO+ Experiment

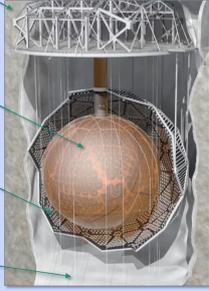
Multi-purpose liquid scintillator detector. Main objective to detect  $0\nu\beta\beta$  decay [1]. Also sensitive to long baseline neutrino oscillation from reactor  $\bar{\nu}_e$  [2, 3]

2 km overburden

6 m radius acrylic vessel (AV) filled with liquid scintillator

Over 9000 PMTs + outward looking PMTs

Surrounded by ultra-pure water

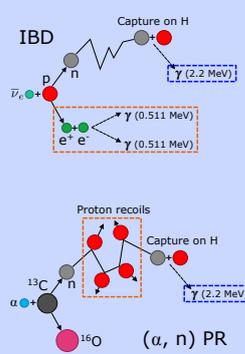


Based on the Sudbury Neutrino Observatory (SNO)

Ultra low radioactive and cosmogenic backgrounds

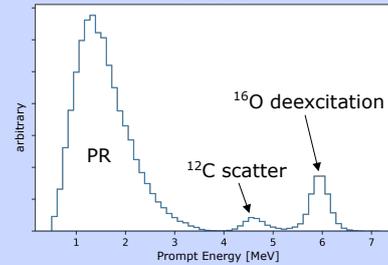
Load  $^{130}\text{Te}$  to produce  $0\nu\beta\beta$ !

## IBD vs $(\alpha, n)$



Reactor  $\bar{\nu}_e$  detected via inverse beta decays (IBD) on protons. Prompt-delayed event structure:  $e^+$  annihilation and neutron capture. Coincidence tagging removes almost all backgrounds.

Main remaining background is  $^{13}\text{C}(\alpha, n)^{16}\text{O}$  events, mimicking this prompt-delayed signature [2, 3]. Three possible prompt events:



- **Proton recoil (PR):** neutron recoils on protons.
- **$^{12}\text{C}$  scatter:** neutron scatters off a  $^{12}\text{C}$ , which releases a  $\gamma$ .
- **$^{16}\text{O}$  deexcitation:**  $^{16}\text{O}$  produced in excited state, releases a  $\gamma$  or  $e^+e^-$  pair.

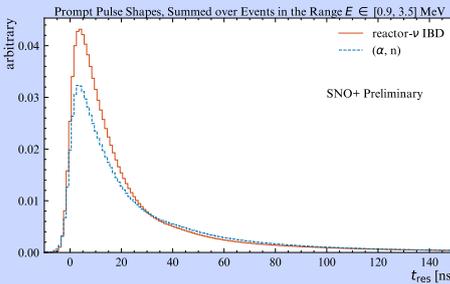
## Pulse Shapes

The pulse shape of an event is its **light emission over time**. At SNO+ it is constructed from the PMT residual hit times:

$$t_{\text{res}} \equiv t_{\text{hit}} - t_{\text{evt}} - t_{\text{TOF}}$$

$t_{\text{hit}}$  is a single PMT hit time,  $t_{\text{evt}}$  the event time, and  $t_{\text{TOF}}$  the straight line time-of-flight between the event and the PMT [4].

**PR events** last longer, so the pulse shape has a **longer tail**. The **scintillation timing** is also different for  $\beta$ 's and protons.



The  $\beta$  timing was calibrating using in-situ tagged  $^{214}\text{Bi}$  and  $^{214}\text{Po}$  decay pairs [5].

The **proton timing** is being calibrated with a combination of tagged low-energy  $(\alpha, n)$  events, and an  $^{241}\text{Am}$   $^9\text{Be}$  (AmBe) source [1]. This can produce similar PR events via the  $^9\text{B}(\alpha, n)^{12}\text{C}$  interaction.

## The Classifier

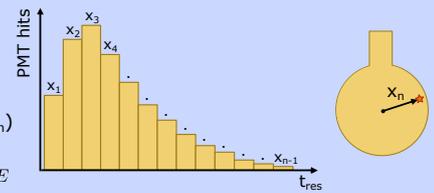
The **Pulse shape** is **correlated** with the event **energy** (E) and radial **position** (R) in the detector, in non-trivial ways. A likelihood ratio based on averaged PDFs does not capture this.

A **Fisher discriminant** is used, which finds the **projection vector** that best separates two multi-dimensional datasets, accounting for their covariance matrices.

Construct a vector  $\vec{x}$  for each event from **binned  $t_{\text{res}}$**  and its **radial position**:

$$\vec{x} = (x_1, x_2, \dots, x_{n-1}, x_n)$$

$$\sum_{i=1}^{n-1} x_i \approx N_{\text{hit}} \approx E$$



Select samples of simulated IBD signal (S) and  $(\alpha, n)$  background (B) events, with similarly uniform E and  $R^3$  distributions.

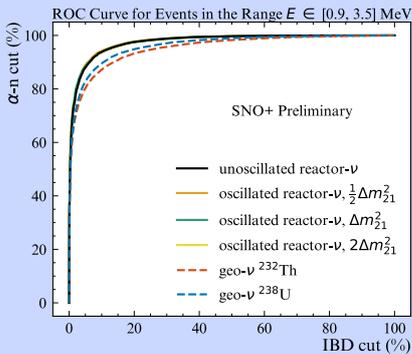
Compute projection vector  $\vec{a}$ , from sample means  $\mu$  and the "within-class" covariance matrix W [6]:

$$\vec{a} \propto W^{-1} (\vec{\mu}_S - \vec{\mu}_B)$$

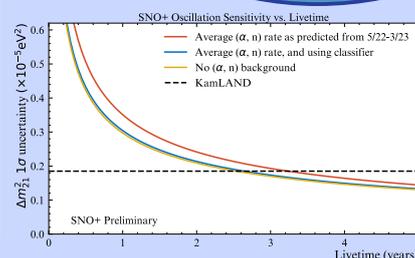
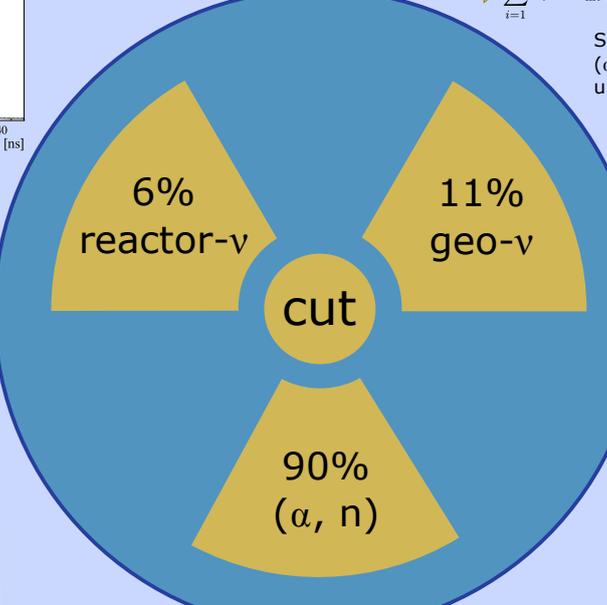
Classify each event via the test statistic:

$$\mathcal{F} = \vec{a} \cdot \vec{x}$$

Classifier applied to simulated IBD and  $(\alpha, n)$  events from expected distributions, assuming different neutrino oscillation scenarios:



Performance is independent of oscillation. **Out-performs log-likelihood classifier**. Example cut: 90%  $(\alpha, n)$ , 6% reactor- $\nu$  and 11% geo- $\nu$ .

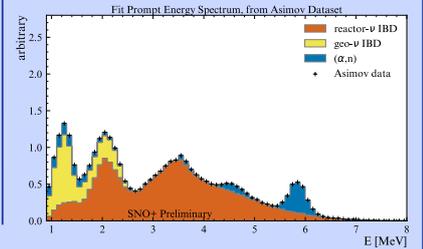
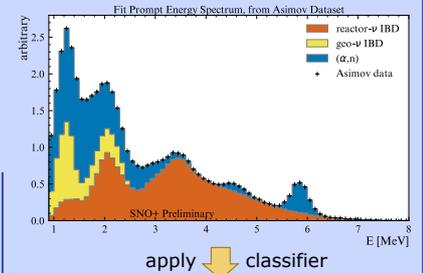


Practically eliminates the impact of  $(\alpha, n)$  on  $\Delta m_{21}^2$  measurements.

Work still preliminary, simulation-based.

See Sofia Andringa's poster on Friday for SNO+ oscillation analysis results!

## Impact on prompt energy spectrum:



## Results

[1] SNO+ collaboration. *The SNO+ experiment*. Journal of Instrumentation, 16(08):P08059, 2021.  
 [2] SNO+ Collaboration. *Observation of antineutrinos from distant reactors using pure water at SNO+*, arXiv preprint arXiv:2210.14154, 2022.  
 [3] SNO+ Collaboration. *Initial measurement of reactor antineutrino oscillation at SNO+*, arXiv preprint arXiv:2405.19700, 2024.  
 [4] C. Mills. *Improved sensitivity to  $\Delta m_{21}^2$  by classification of the  $^{13}\text{C}(\alpha, n)^{16}\text{O}$  background in the SNO+ antineutrino analysis*, Ph.D. thesis, U. Sussex (2022).  
 [5] A. Zummo. *Measurement of reactor antineutrino oscillation with SNO+*, Ph.D. thesis, U. Pennsylvania (2024).  
 [6] Glen Cowan. *Statistical data analysis*. Oxford university press, 1998.

