



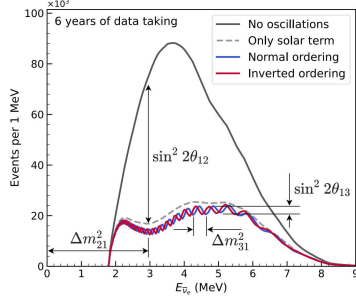
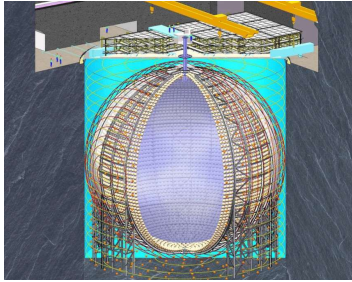
Calibration Strategy of JUNO

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JUNO Experiment

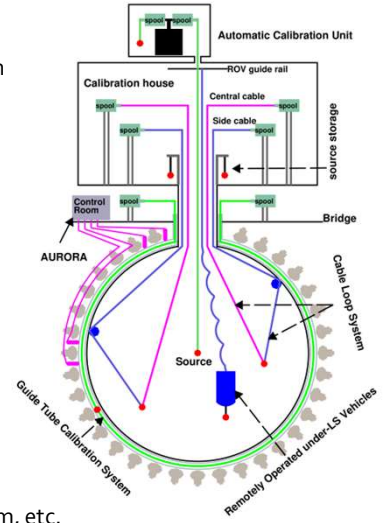
- Jiangmen Underground Neutrino Experiment (JUNO)
 - ✓ neutrino mass ordering measurement
 - ✓ precision measurement of $|\Delta m_{31}^2|$, Δm_{21}^2 and θ_{12}
- calibration goal: for uniformly distributed e+ events
 - ✓ better than 1% energy calibration precision
 - ✓ better than 3% energy resolution at 1 MeV
- 17612 20-inch PMTs, 25600 3-inch PMTs, 78 % photocathode coverage



Calibration System

- Automatic Calibration Unit
- Cable Loop System
- Guide Tube Calibration System
- Remotely Operated Vehicle

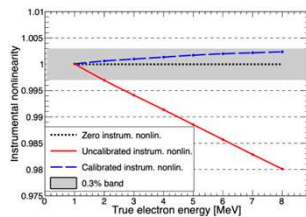
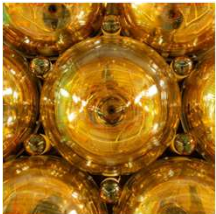
source	type	energy
^{137}Cs	γ	0.662 MeV
^{54}Mn	γ	0.835 MeV
^{60}Co	γ	$1.173 + 1.333$ MeV
^4K	γ	1.461 MeV
^{68}Ge	γ	$0.511 + 0.511$ MeV
$^{241}\text{Am-Be}$	n, γ	$n + 4.43$ MeV ($^{12}\text{C}^*$)
$^{241}\text{Am-}^{13}\text{C}$	n, γ	$n + 6.13$ MeV ($^{16}\text{O}^*$)
$(n, \gamma)p$	γ	2.22 MeV
$(n, \gamma)^{12}\text{C}$	γ	4.94 MeV or $3.68 + 1.26$ MeV



+ low energy sources: ^{226}Rn , ^{241}Am , etc.

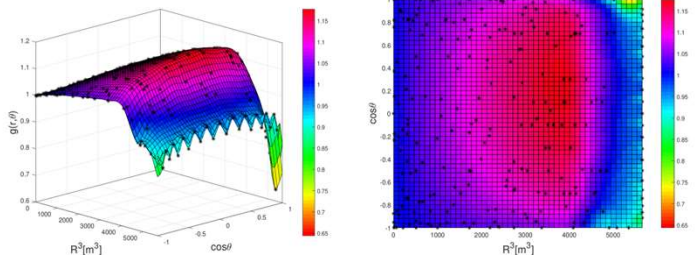
20-inch PMT Calibration

- dark noise, time alignment, gain calibration
- first hit time vs. charge
- relative photon detection efficiency
- 20-inch PMT non-linearity: electronics, waveform reconstruction
 - ✓ use 266 nm laser to inject light on a Teflon ball at detector center
 - ✓ 3-inch PMTs working in digital mode: fired or non-fired
 - ✓ Poisson-zero method: use the number of non-fired 3-inch PMTs to calibrate laser intensity



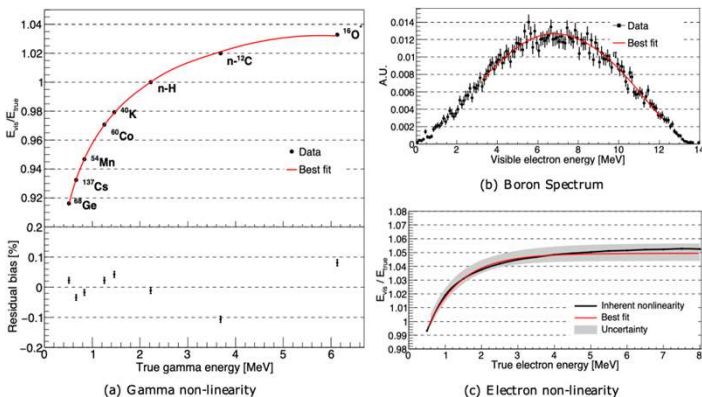
Non-Uniformity Calibration

- detector non-uniformity: PMT acceptance, light attenuation, total internal reflection at the acrylic-water interface
- calibration sources deployed by the calibration system at different positions in the detector
- cosmogenic neutrons, about 1 M / week
- develop reconstruction algorithms that use these calibration data as input to largely remove the detector non-uniformity



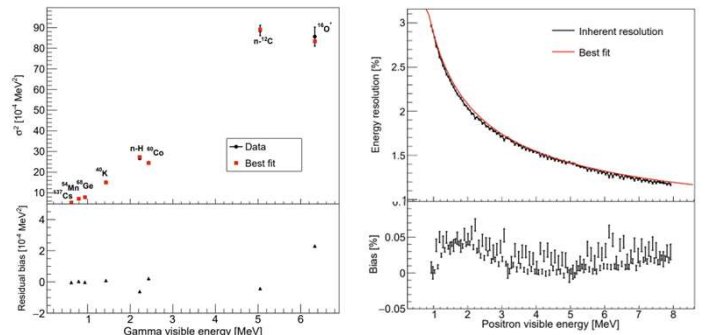
Liquid Scintillator Non-Linearity Calibration

- liquid scintillator non-linearity: Birks' law, Cherenkov light
- gamma calibration sources at different energies
 - ✓ treat a gamma event as a collection of e- and e+ using Geant4
 - ✓ build a non-linearity model for e-/e+, fit to calibration data
 - ✓ predict the non-linearity for e+
- constraint from continuous spectra from cosmogenic B12 and C11



Positron Energy Resolution Calibration

- use the e+/e- nonlinearity model to predict resolution curve for e+
- further parametrize the resolution curve with $\sigma/E = \sqrt{a^2/E + b^2 + c^2/E^2}$
- contributing factors: a: Poisson, b: non-uniformity and Cherenkov, c: mainly dark current and e+ annihilation



References: *JHEP* 03 (2021) 004, F. Zhang PhD thesis (2021), *Chin.Phys.C* 46 (2022) 12, 123001, arXiv:2405.18008