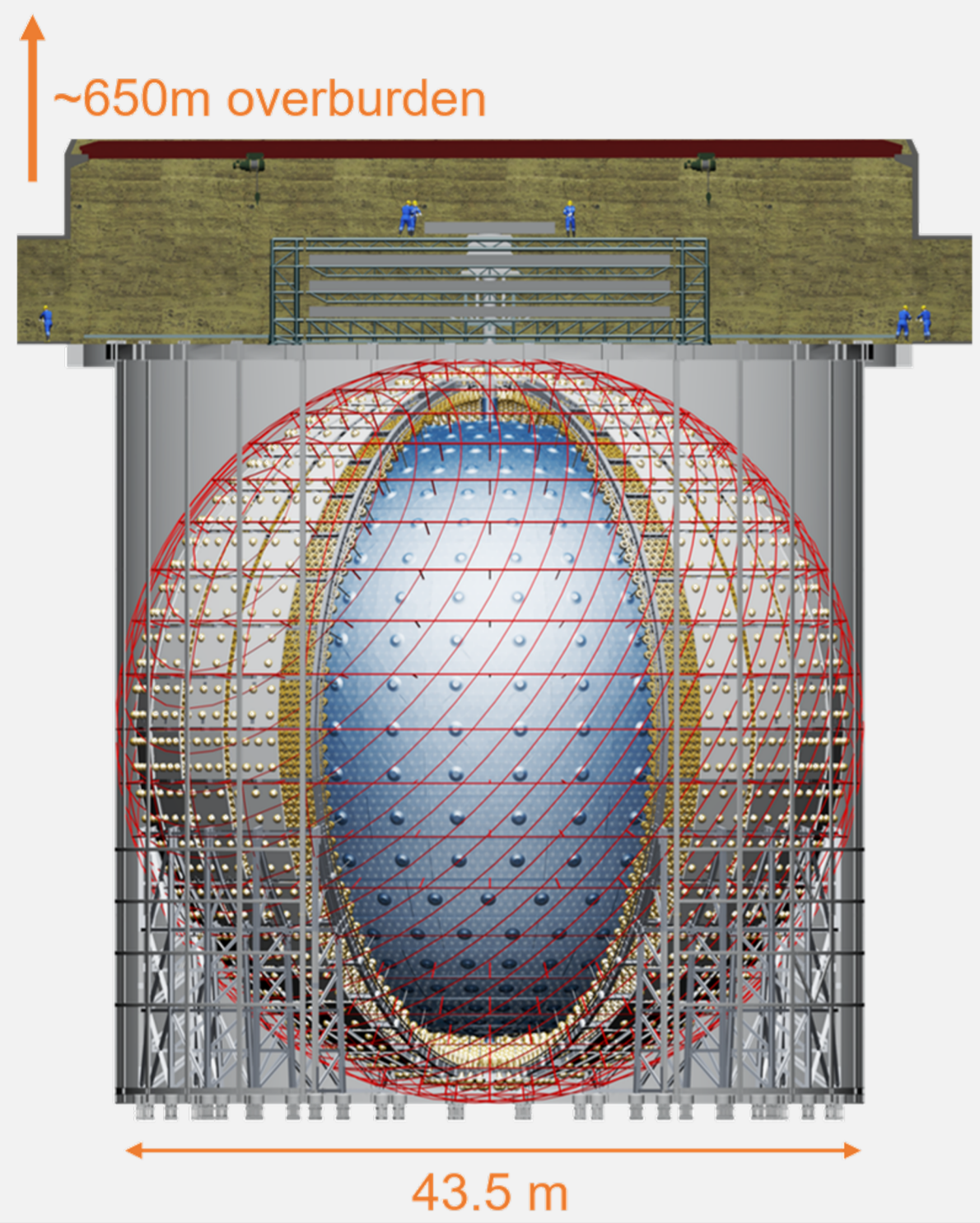
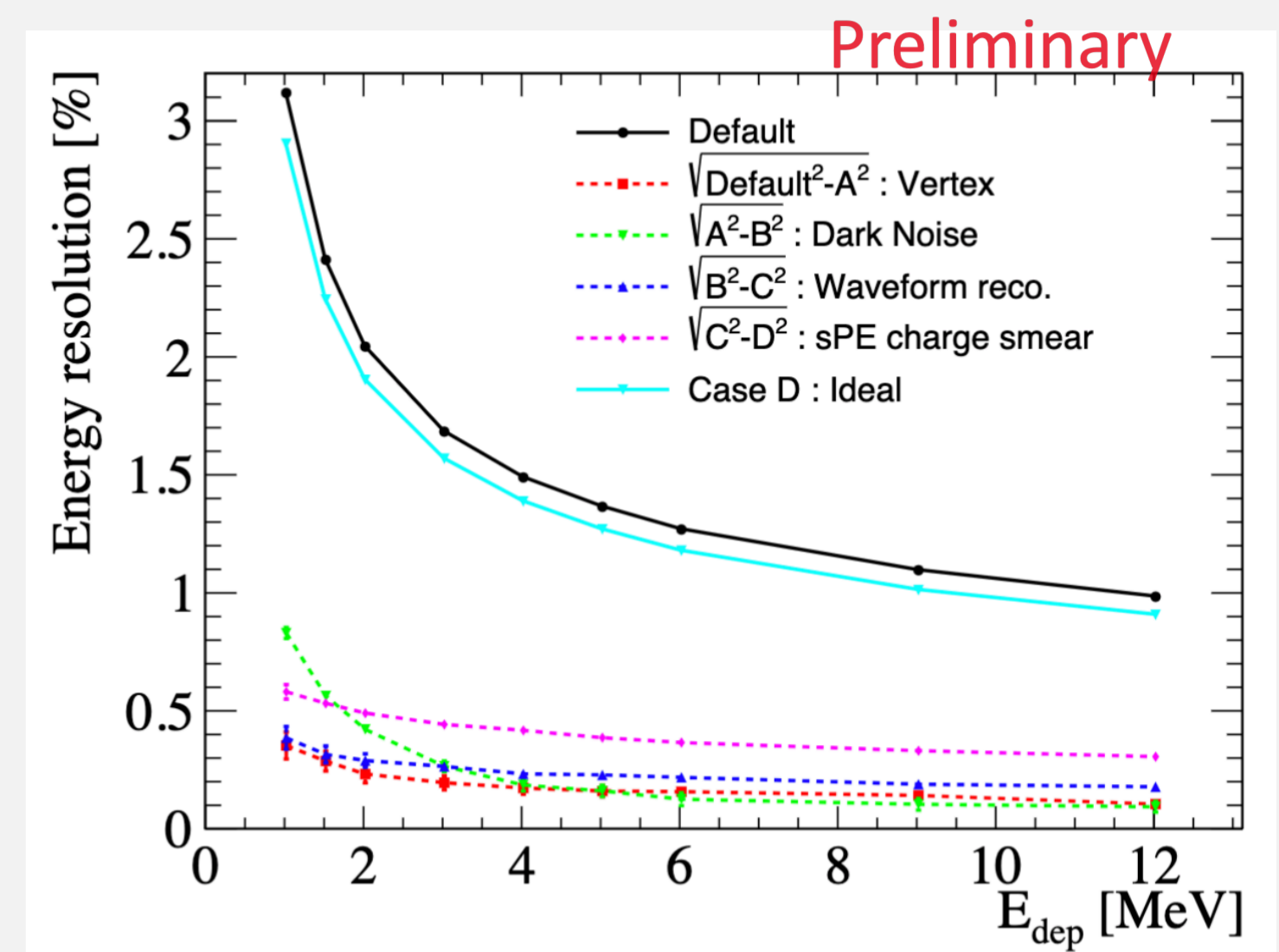
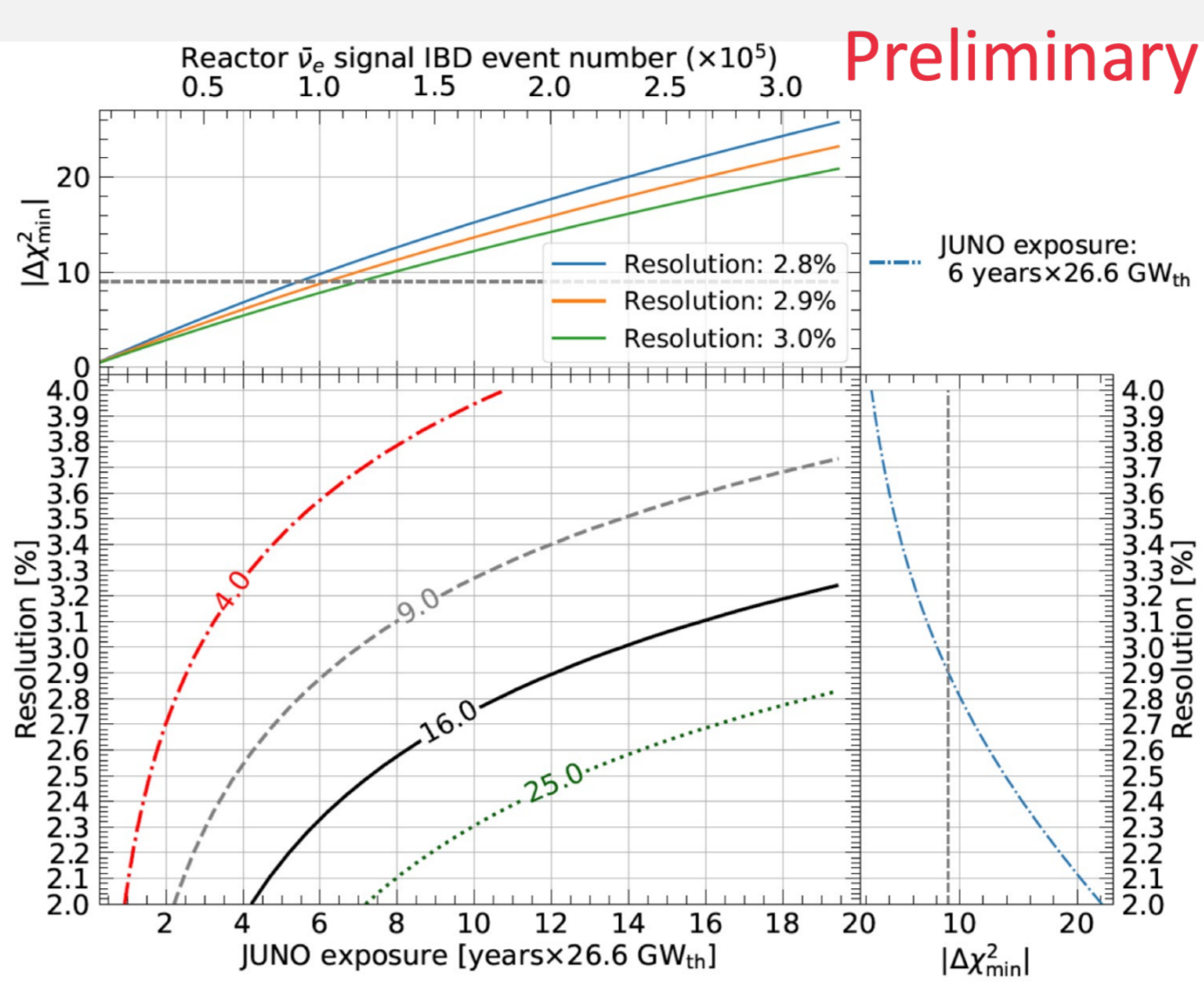


## 1 JUNO Experiment



- World's largest liquid-scintillator detector
- Central Detector: 20kton LS, 17'612 20" PMTs and 25'600 3" PMTs
- Unprecedented energy resolution 3%@1MeV
- Main physics goal: determination of neutrino mass ordering (NMO)

## 2 Energy Resolution



- Energy resolution is crucial for NMO sensitivity in JUNO
- PMT charge smearing is one of the dominant factors

## 3 ML based Photon Counting

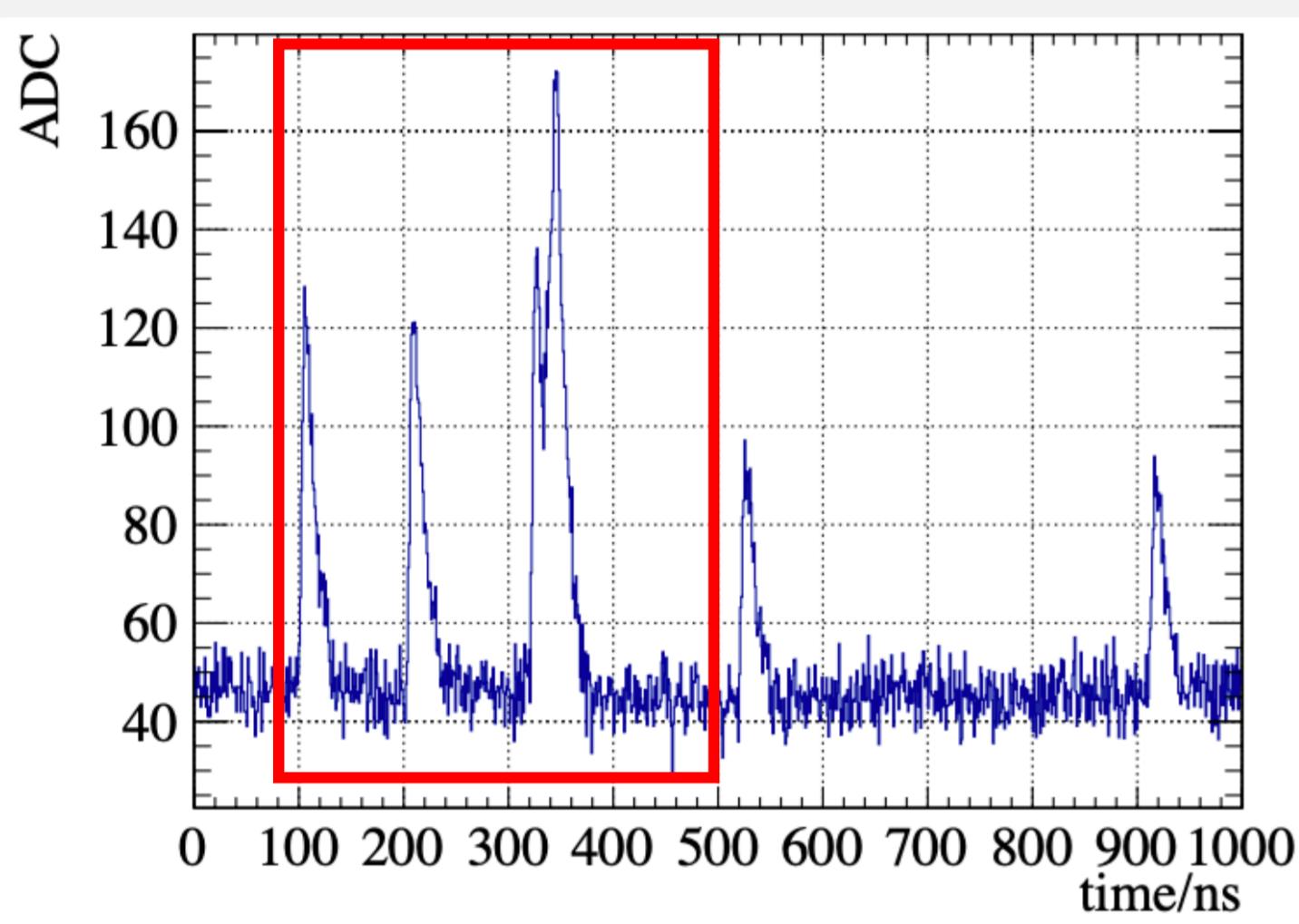
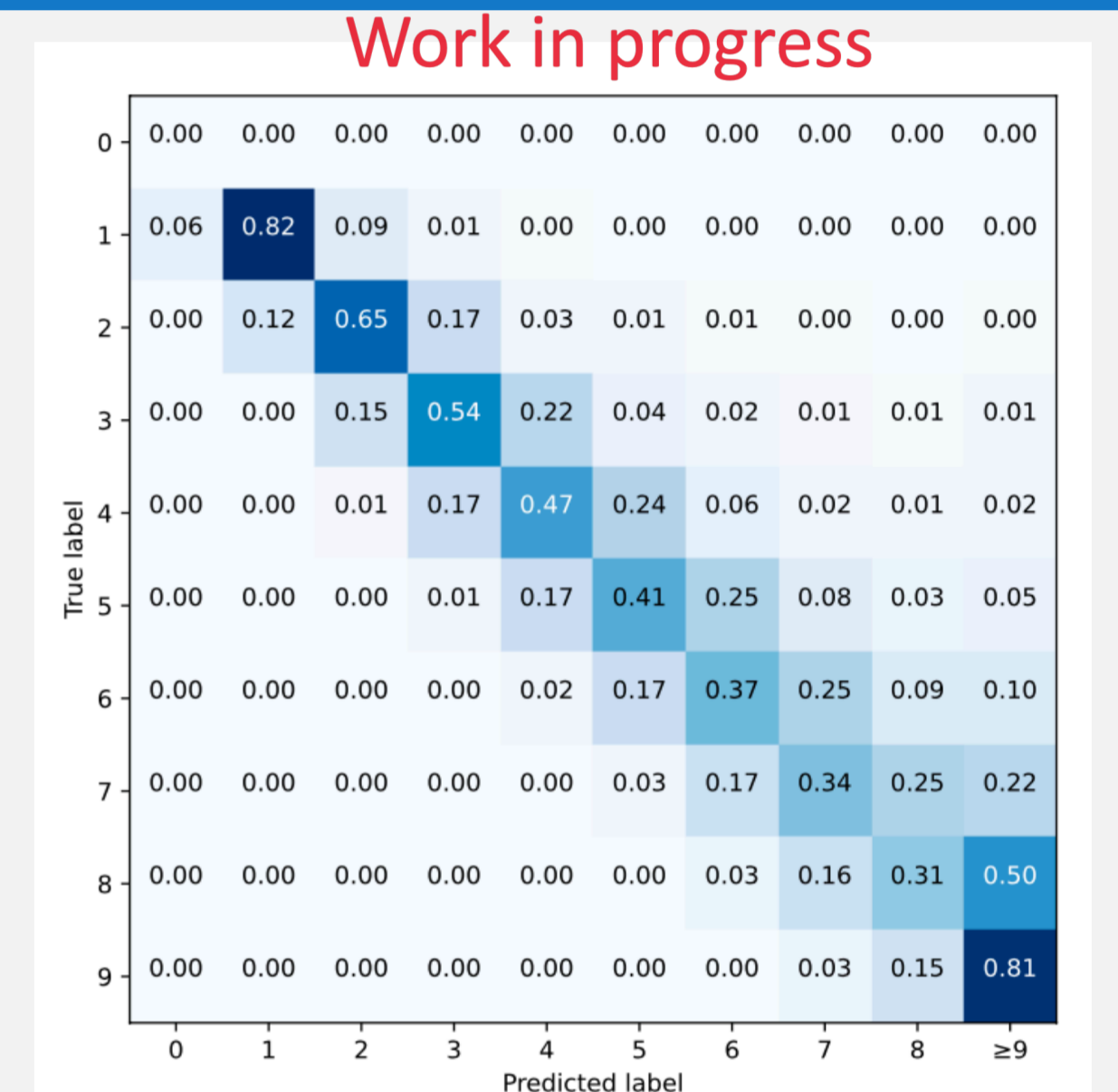
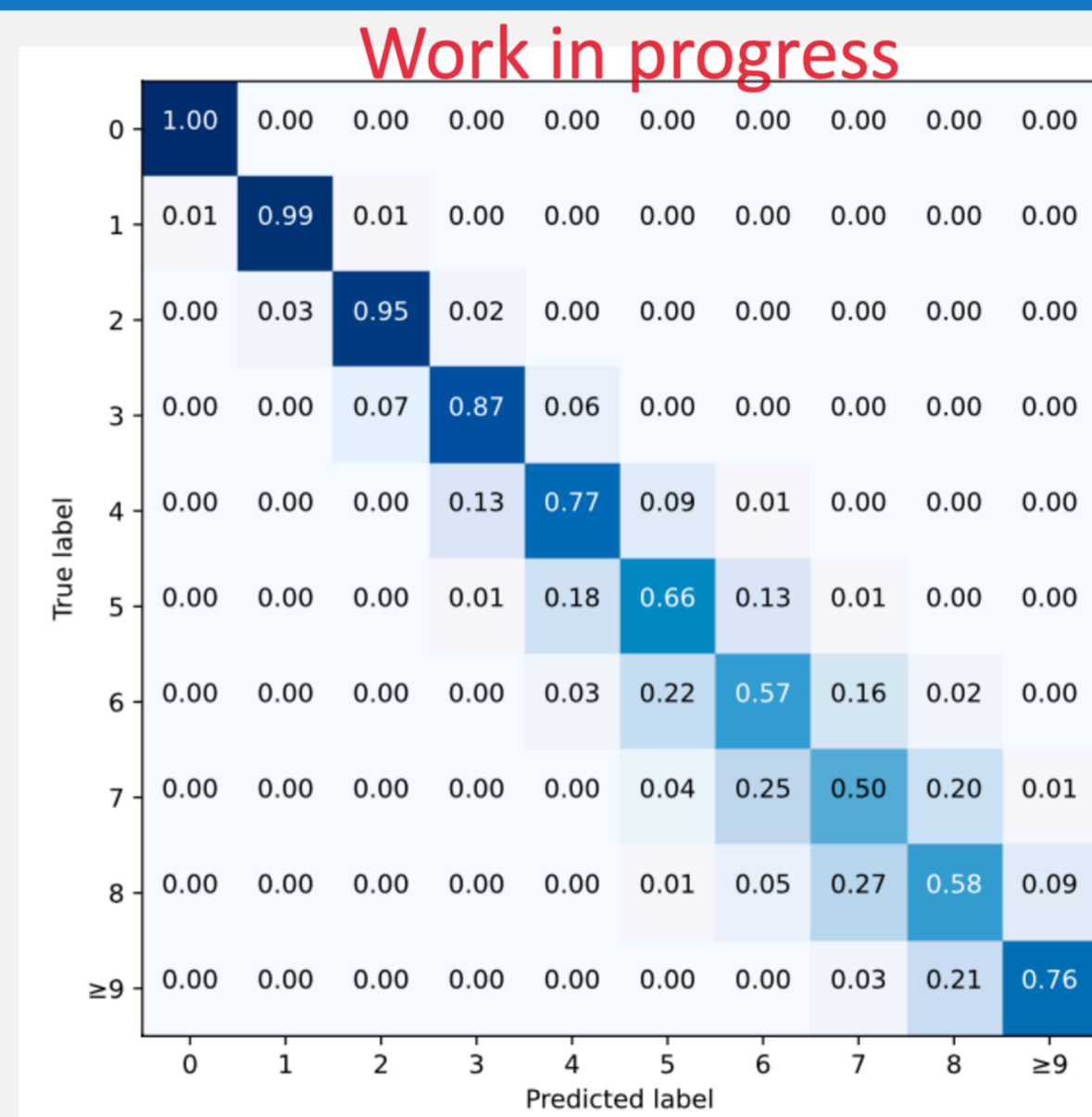


Table 2: Modified RawNet architecture. For convolutional layers, numbers inside parentheses refer to filter length, stride size, and number of filters. For gated recurrent unit (GRU) and fully-connected layers, numbers inside the parentheses indicate the number of nodes.

Layer	Input	Output shape
Strided -conv	Conv(3,3,128) BN LeakyReLU	(128, 140)
Res block	Conv(3,1,128) BN LeakyReLU Conv(3,1,128) BN LeakyReLU MaxPool(3)	×2 (128, 46)
Res block	Conv(3,1,256) BN LeakyReLU Conv(3,1,256) BN LeakyReLU MaxPool(3)	×2 (256, 1)
GRU	GRU(1024)	(1024,)
Speaker embedding	FC(128)	(128,)
Output	FC(10)	(10,)

- Input: pre-processed PMT waveform within 420ns signal window
- Model: Customized RawNet
- Output:  $\{p_k\}$  the probability for predicting ( $k=0,1, \dots \geq 9$ ) PEs

## 4 Photon Counting Performance



- Left: Confusion matrix of RawNet
- 99% (95%, 87%) accuracy for 1PE (2PEs, 3PEs)
- Accuracy decreases rapidly as nPEs increases
- Right: Confusion matrix based on charge classification
- The accuracy is markedly inferior to that of RawNet

## 5 Energy Reconstruction

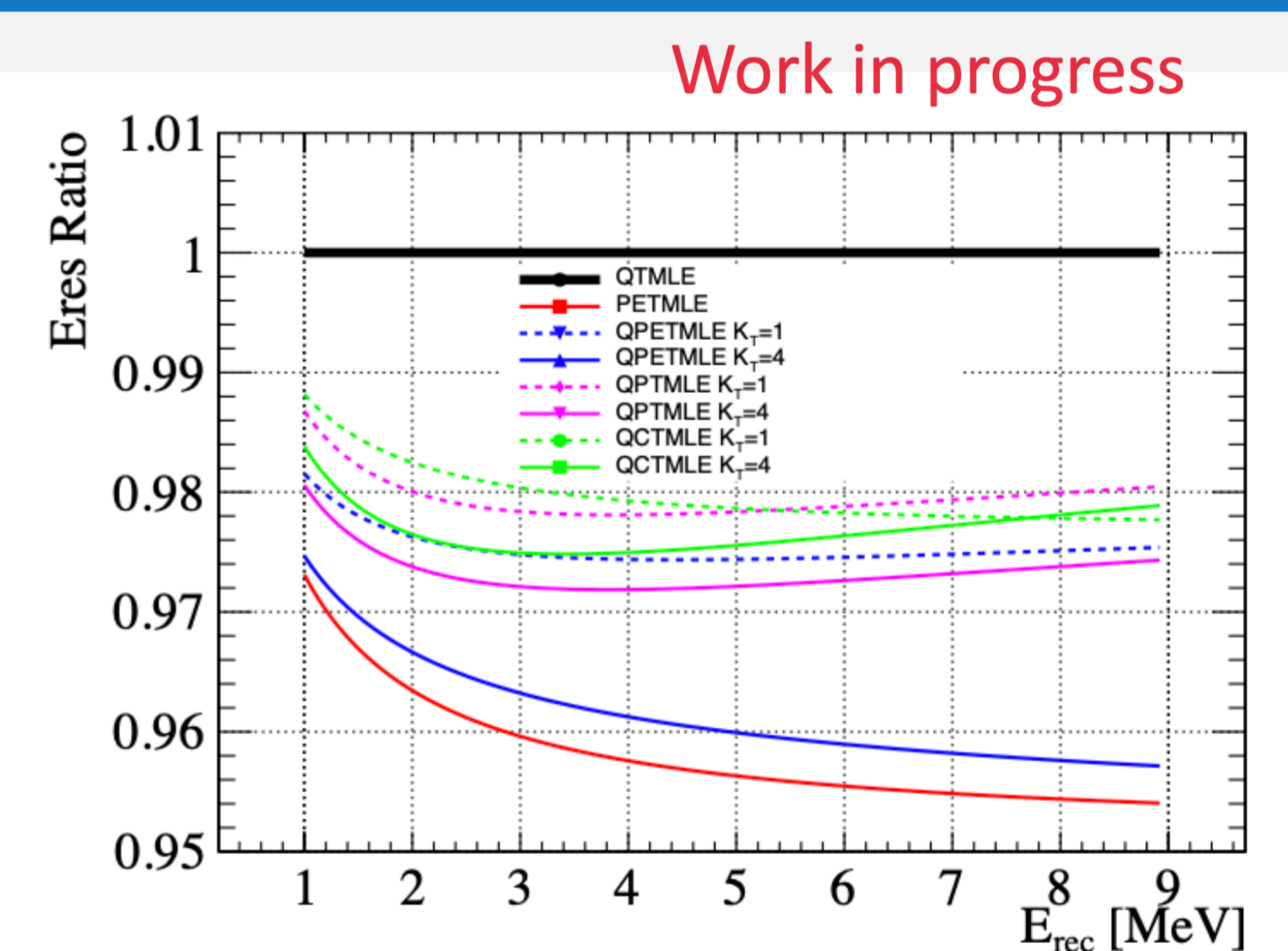
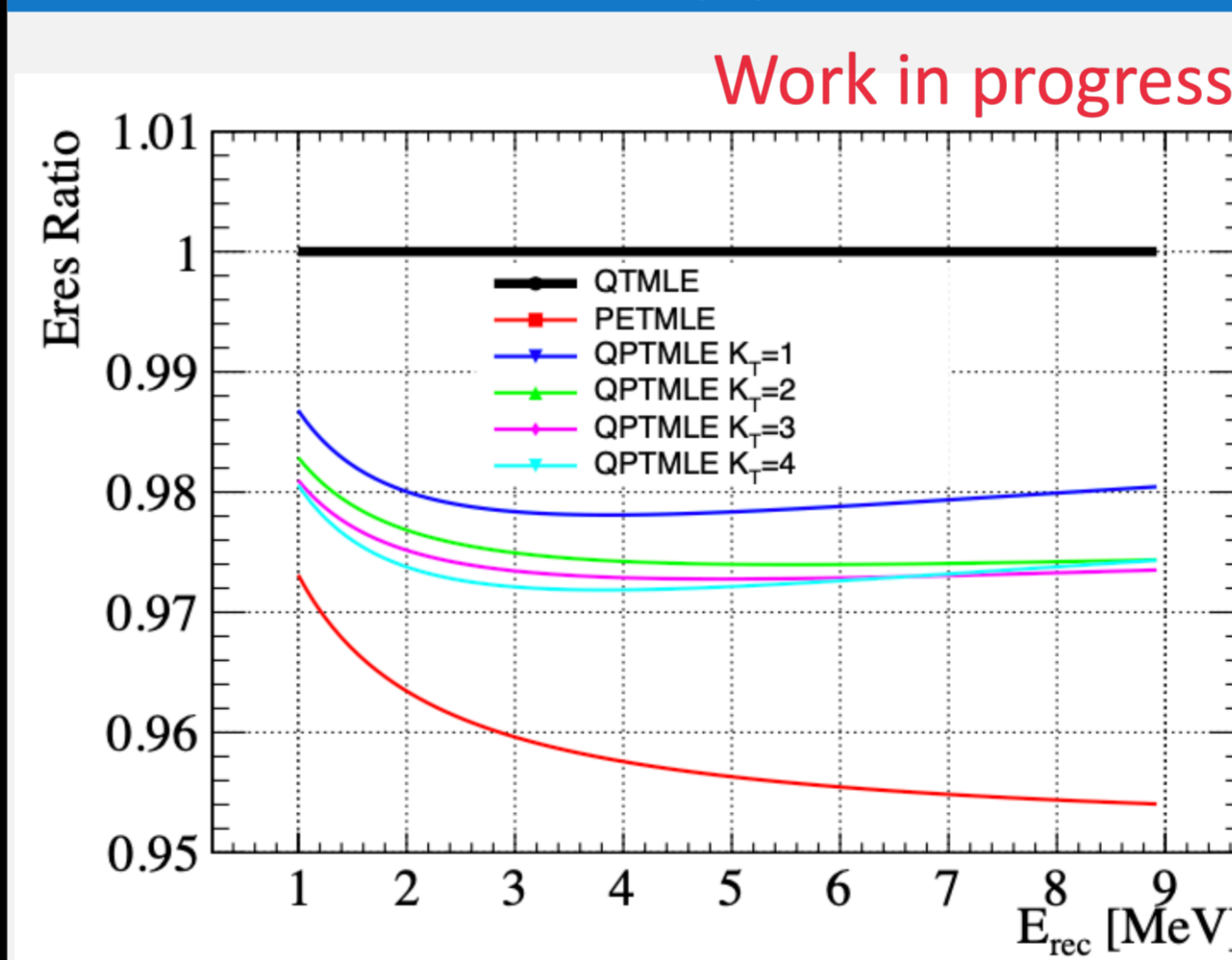
Algo. Name	Observable	Likelihood: $\kappa \leq K_T$	Likelihood: $\kappa > K_T$
QTMLE (reference)	q (charge)	$\mathcal{L}(q_i \mu_i) = \sum_{k=1}^{+\infty} P_Q(q_i k)P(k, \mu_i)$	
PETMLE (ideal)	k (true PEs)	$\mathcal{L}(k_i \mu_i) = P(k_i, \mu_i)$	
QPTMLE (realistic)	$\{p_k\}, q$	$\mathcal{L}(\{p_k\} \mu_i) = \sum_{k=0}^9 R_{K_T k} p_k^i P(k, \mu_i)$	
QPETMLE (100% accuracy)	$k(p_{\kappa=1}), q$	$\mathcal{L}(k_i \mu_i) = P(k_i, \mu_i)$	$\mathcal{L}(q_i \mu_i) = \sum_{k=1}^{+\infty} P_Q(q_i k)P(k, \mu_i)$
QCTMLE	$\kappa (p_{\kappa:\max}), q$	$\mathcal{L}(\kappa_i \mu_i) = \sum_{k=0}^9 C_{k\kappa_i} \times P(k, \mu_i)$	

where  $\mu_i$  is the expected nPEs for the i-th PMT,  $P(k, \mu_i)$  is just the Poisson probability of observing k p.e. given  $\mu_i$  and  $P_Q(q_i|k)$  is the charge pdf for k p.e.

$$R_{K_T k} = \sum_{\kappa=0}^{K_T} C_{k\kappa}$$

confusion matrix  $C_{kk}$

## 6 Energy Resolution Performance



- Using the photon counting information for PMTs with ( $\kappa \leq K_T$ ) PEs can improve the energy resolution
- The improvement becomes smaller as  $K_T$  increases due to the dropping accuracy for high PEs
- Additional checks were done to validate the results

## Summary

- Energy resolution is crucial for the NMO sensitivity in JUNO, while PMT charge smear is one of the dominant factors
- A Machine Learning based photon counting method was developed for PMT waveforms, which can achieve high accuracy at low PEs
- Integration of the photon counting information in the energy reconstruction can partially mitigate the impact of PMT charge smearing, leading to 2% to 2.8% relative improvement on the energy resolution

## References

1. GuiHong Huang et al, Data-driven simultaneous vertex and energy reconstruction for large liquid scintillator detectors, Nucl.Sci.Tech. 34 (2023) 6, 83
2. JUNO Collaboration, Prediction of Energy Resolution in the JUNO Experiment, arXiv:2405.17860
3. Wei Jiang et al, Machine-Learning based photon counting for PMT waveforms and its application to the improvement of the energy resolution in large liquid scintillator detectors, arXiv:2405.18720