



The trigger and data acquisition system for the **TRIDENT phase-1 detector**

Weilun Huang¹, Liang Li², Xinchen Li¹, Hualin Mei^{1,2}, Cen Mo² for the TRIDENT collaboration (Contact: mei.hualin@sjtu.edu.cn) 1. Tsung-Dao Lee Institute, Shanghai Jiao Tong University 2. School of Physics and Astronomy, Shanghai Jiao Tong University





Level 1: Local coincidence trigger ($CL \ge 2$)

Use information from individual hDOM

rate [Hz]

- Perform at hDOM main board's FPGA off-shore
- Number of PMTs having over-threshold signals within 20 ns, mostly removing K40 background
- Based on information from the whole detector volume (e.g. total NPE, space-time correlation) Perform at computing cluster on-shore, filter data and further classification (track, cascade, BSM ...)
- Example above (assuming 3KHz random noise per PMT from K40/dark noise):
 - Define primary hDOM (most hits: t^0, x^0) \rightarrow Calculate $dR^2 = (t_0 t)^2 (x_0 x)^2 \rightarrow Score = -log(1 + |dR^2|)$



[1] A multi-cubic-kilometer neutrino telescope in the western Pacific Ocean, Nature Astronomy volume 7, pages 1497–1505 (2023) [2] Fan Hu, Zhuo Li, Donglian Xu, Exploring a PMT+SiPM hybrid optical module for next generation neutrino telescopes, PoS (ICRC 2021), 1043 [3] Qichao Chang, Fan Hu, Iwan Morton-Blakec, Donglian Xu, Optimizing the optical array geometry for TRIDENT, PoS (ICRC 2023)1203 [4] https://daqling.docs.cern.ch/

