





# Data vs. MC comparison of light signal from cosmic rays in the ICARUS detectors

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## 1 – The ICARUS detectors system

ICARUS<sup>§,\*</sup> is collecting data exposed to BNB and Numi off-axis beam within the SBN program at Fermilab; due to its operations at shallow depths, it is also exposed to a huge flux of cosmic rays, which is exploited for detectors calibration. It is composed of two identical cryostats, surronded by the Cosmic Ray Taggers ( $\sim$ 95% efficiency tagging cosmics rays).

In each cryostat two Liquid Argon TPC with a common cathode are placed. The electrons ionized in TPC are continuously detected by 3 non-destructive readout planes with different orientation ( $0^{\circ}$ ,  $\pm 60^{\circ}$ ).

## 1.1 – Scintillation light detection system

Behind the wires plane, 90 PMTs per TPC (5% coverage, 15 ph.e./MeV) provide the scintillation detection system <sup>§,†</sup> to detect vacuum ultraviolet photons produced by ionizing particles in LAr and allowing to:

Drift coordinate Schematic view of the ICARUS-T600 readout principle, shown for one TPC§

The ICARUS PMTs mounted behind the wires of one TPC§



- identify the interaction time (*time resolution*  $\sim ns$ ).
- Localize events in PMT plane (*spatial resolution <50cm*).
- Roughly determine the event topologies.
- Generate a trigger signal:
  - ICARUS main trigger signal\* = light signals from PMTs in coincidence with beam spills.
    - Beam events are collected requiring at least 5 fired PMT pairs (Mj = 5) inside one of 6 m longitudinal slices equipped with 30+30 opposite PMTs.
  - MinBias : minimum-bias triggers with out requesting scintillation light a priori; the timing is provided by CRT. It provide the sample for trigger efficiency study: the trigger is emulated starting from recorded PMT waveforms, and the logic is evaluated for each stopping muon.

#### 2 – MC simulation of the light signal

The simulation of the scintillation photons are generated with a Monte Carlo<sup>+</sup>: (i) photons are generated based on energy deposition and particle type, and (ii) are propagated through the liquid argon; (iii) all their information are stored; (iv) photon by photon, the sigle photon response is added. (v) The simulated noise is added to the waveforms. (vi) If the signal exceed a threshold (~0.6 ph.e.) on a channel, the waveform is recorded in a  $4\mu$ s time window.



The trigger efficiency as function of muon energy for a highly-pure sample of stopping cosmic muons (MinBias data from Run2) selected based on topology and calorimetry (Bragg peak). Courtesy of Riccardo Triozzi.

The PMTs associated with a cosmic ray muon crossing the cathode §

#### 3 – Preliminary study of light signal: comparison data vs. MC

The MC simulation has been data-based optimized tuning the simulation's parameters related to the gain and to the quantum efficiency. The data sample is a run of about 15 k-events from the Run2 collected with BNB-majority.

- 3.1 Samples' selections: the brighest light signal in coincidence with cathode crossing vertical tracks
  - 1. The cathode crossing vertical tracks were selected  $\rightarrow$  sample completely under control in time and position;
  - 2. Only the first flash (i.e. collection of light signals in the time window of 40ns in at least 5 PMT) in coincidence (in time and in spatial barycenter along beam direction) with selected tracks was consider;
  - 3. the first optical hits (i.e. light signal) looking along the time for each PMT are recognized
  - 4. the 10 with the highest amplitude are selected: brightest signals.

### 3.2 – Good agreement between data and MC amplitude

The tuned MC well reproduce the data amplitude of the brightest light signals.





4.1 – Validation of the trigger efficiency. The data trigger efficiency (Mj=5) for single track is well matched using data-based tuning of the MC parameters.



4.2 – Validation of the light position The data flash barycentre along the beam direction is quite well reproduced by MC one. It is important in the analysis <sup>2</sup> to select the track-flash match (and assign a time to noncathde-crossing tracks).

References:

- <sup>§</sup> P. Abratenko et al., ICARUS at the Fermilab Short-Baseline Neutrino program: initial operation. EPJ C 83, 467 (2023).
- \* B. Ali-Mohammadzadeh et al., Design and implementation of the new scintillation light detection systemofICARUST600. Jol 15, T10007 (2020).
- \* C. Farnese et al., Implementation of the trigger system of the ICARUS-T600 detector at Fermilab. NIM A 1045, 167498 (2023).
- + E. Snider and G.Petrillo, LArSoft: toolkit for simulation, reconstruction and analysis of liquid argon TPC neutrino detectors. JoP Conf. Series 898, 042057 (2017). S. Agostinelli et al., NIM A 506, 250 (2003). C. Andreopoulos et al., NIM A 614, 87 (2010). C. Andreopoulos et al., preprintarXiv:1510.05494 (2015).

For more details: \*ICARUS at the Short-Baseline Neutrino program: first results D. Gibin plenary talk. Neutrino reconstruction analysis at ICARUS detector M. Artero Pons poster #51