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## Lorentz invariance violation (LIV)

- Lorentz invariance states that the outcome of an experiment is:
  - the same for two inertial observers watching the same experiment
  - independent of the inertial laboratory in which it is performed
- LIV preserves observer-independence, but violates the second condition
- LIV is allowed in many theories beyond the Standard Model

$$(i\gamma^\mu \partial_{x^\mu} - m)\psi(x) = 0 \quad (i\gamma^\mu \partial_{x^\mu} - m)\psi'(x) = 0$$



Observer cannot use the same equations to describe identical experiments in different laboratories

## Standard Model Extension (SME)

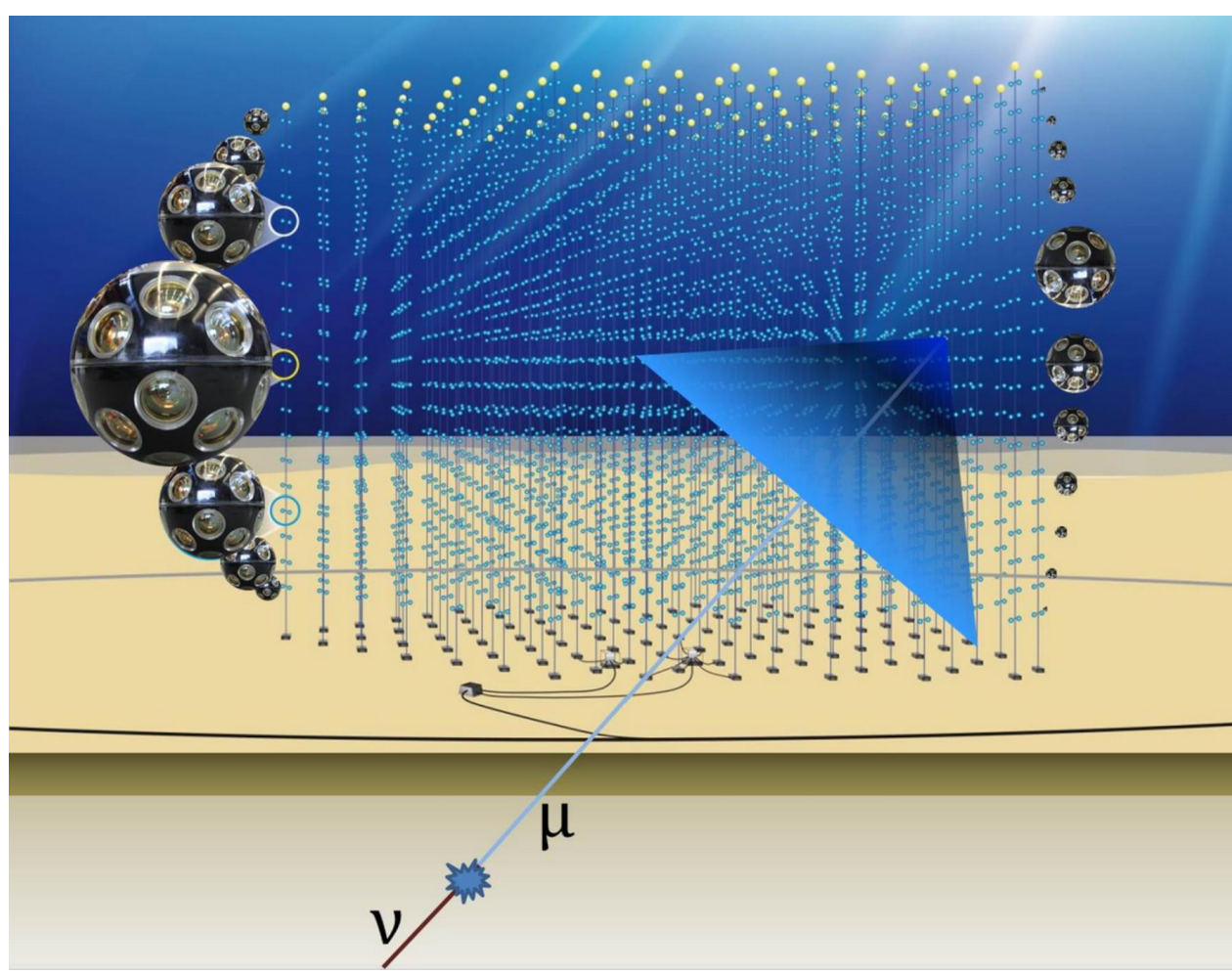
- An extension of the Standard Model including all possible LIV operators<sup>1</sup>
- Focus on isotropic LIV models that preserve rotational invariance in a preferred frame
- General LIV Hamiltonian for neutrinos introducing complex-valued LIV coefficients<sup>2</sup>:

$$H_{LIV} = \begin{pmatrix} \hat{a}_{ee}^{(3)} & \hat{a}_{e\mu}^{(3)} & \hat{a}_{e\tau}^{(3)} \\ \hat{a}_{e\mu}^{(3)*} & \hat{a}_{\mu\mu}^{(3)} & \hat{a}_{\mu\tau}^{(3)} \\ \hat{a}_{e\tau}^{(3)*} & \hat{a}_{\mu\tau}^{(3)*} & \hat{a}_{\tau\tau}^{(3)} \end{pmatrix} - \frac{4}{3}E \begin{pmatrix} \hat{c}_{ee}^{(4)} & \hat{c}_{e\mu}^{(4)} & \hat{c}_{e\tau}^{(4)} \\ \hat{c}_{e\mu}^{(4)*} & \hat{c}_{\mu\mu}^{(4)} & \hat{c}_{\mu\tau}^{(4)} \\ \hat{c}_{e\tau}^{(4)*} & \hat{c}_{\mu\tau}^{(4)*} & \hat{c}_{\tau\tau}^{(4)} \end{pmatrix} + E^2 \hat{a}^{(5)} - E^3 \hat{c}^{(6)} + \dots$$

- Nonzero LIV coefficient implies deviations from standard oscillations
- Mass dimension determines oscillation dependence on baseline L and energy E
- Focus on mass dimension up to six

Coefficient	Unit	CPT	Oscillation effect
$\hat{a}^{(3)}$	GeV	odd	$\propto L$
$\hat{c}^{(4)}$	-	even	$\propto LE$
$\hat{a}^{(5)}$	GeV <sup>-1</sup>	odd	$\propto LE^2$
$\hat{c}^{(6)}$	GeV <sup>-2</sup>	even	$\propto LE^3$

## Experiment descriptions



Artist impression of completed KM3NeT/ORCA

- KM3NeT/ORCA**
  - Cherenkov neutrino telescope under construction in the Mediterranean Sea
  - It will consist of 115 Detection Units (DUs)
  - Focus here on subdetector KM3NeT/ORCA6 operated with six DUs from Jan. 2020 – Nov. 2021
- ANTARES**
  - Cherenkov neutrino telescope operated in the Mediterranean Sea from 2007 to 2022
  - It consisted of 12 lines

## Analysis method

- Analysis uses 2D binning in reconstructed energy and zenith angle
- Models assuming LIV are fitted to data by minimizing negative log-likelihood

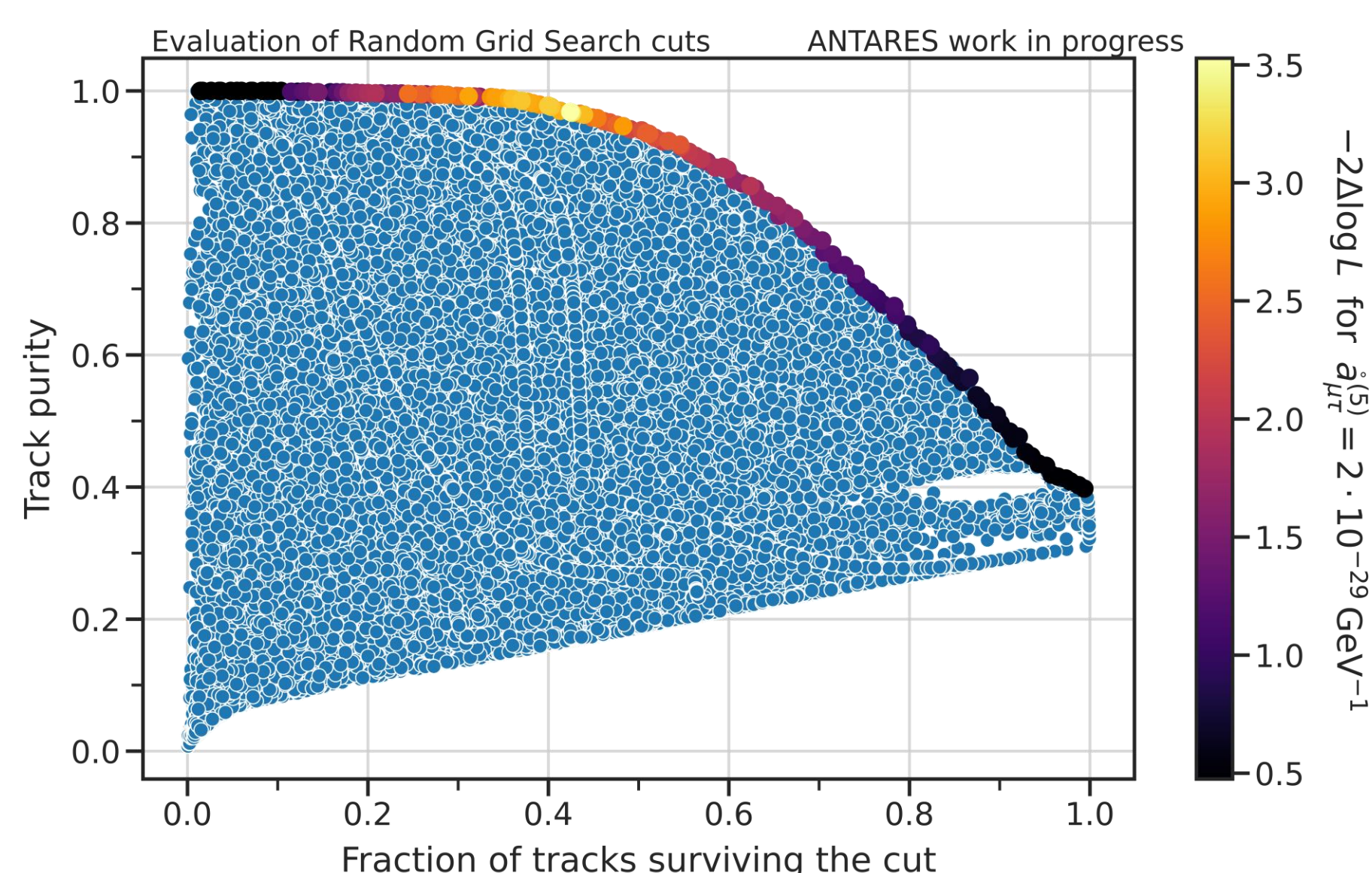
- Standard oscillation parameter values from NuFit v5.0 global fit (with SK)
- Systematic flux parameters for KM3NeT/ORCA6 on the right
- Parameters with a \* are currently kept fixed in ANTARES

Fit parameter	Prior uncertainty
Energy Scale*	9%
Overall Norm	Free
Shower Norm*	Free
HP Track Norm	Free
Spectral Index	0.3
HE Light Simulation*	20%
Muon Norm*	Free
NC Norm	20%
$\nu_\tau$ -CC Norm	20%
$\nu_e/\bar{\nu}_e$	7%
$\nu_\mu/\nu_e$	2%
$\bar{\nu}_\mu/\nu_\mu$	5%
$\nu_{\text{hor}}/\nu_{\text{ver}}$	2%

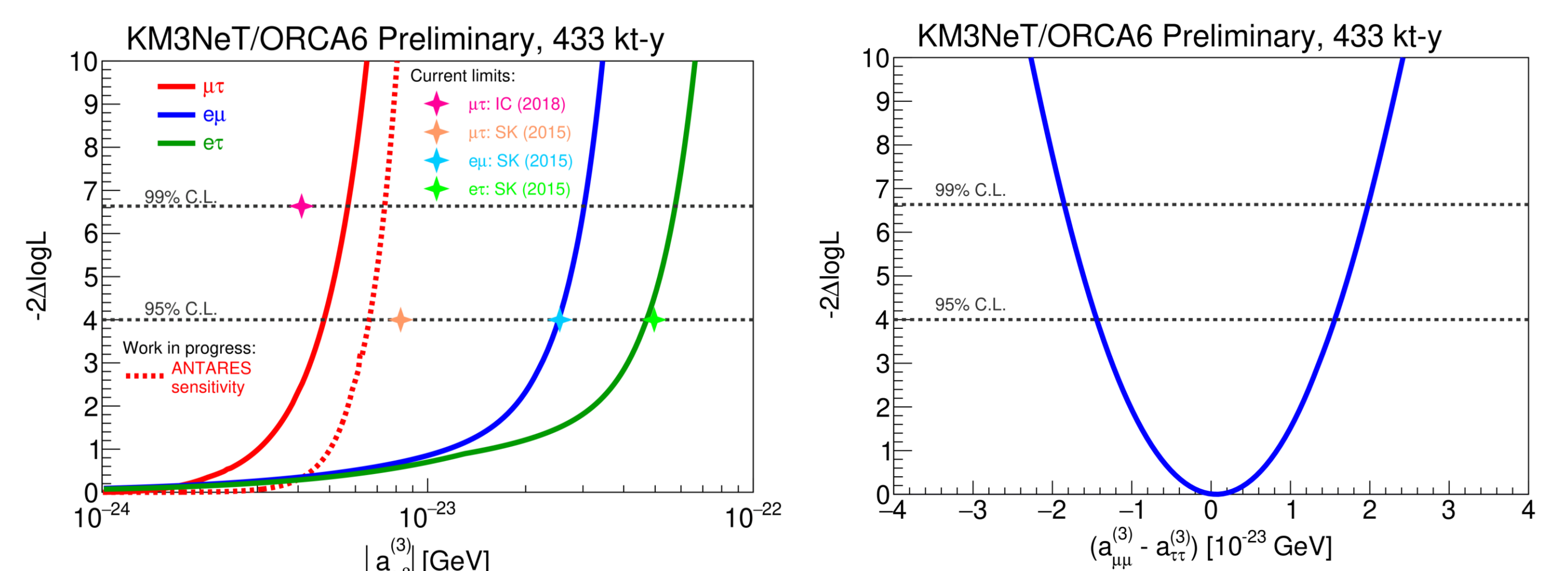
- Off-diagonal coefficients are tested one by one, starting with the lowest mass dimension
- On-diagonal coefficients have strongly correlated oscillation effects

## Event selection

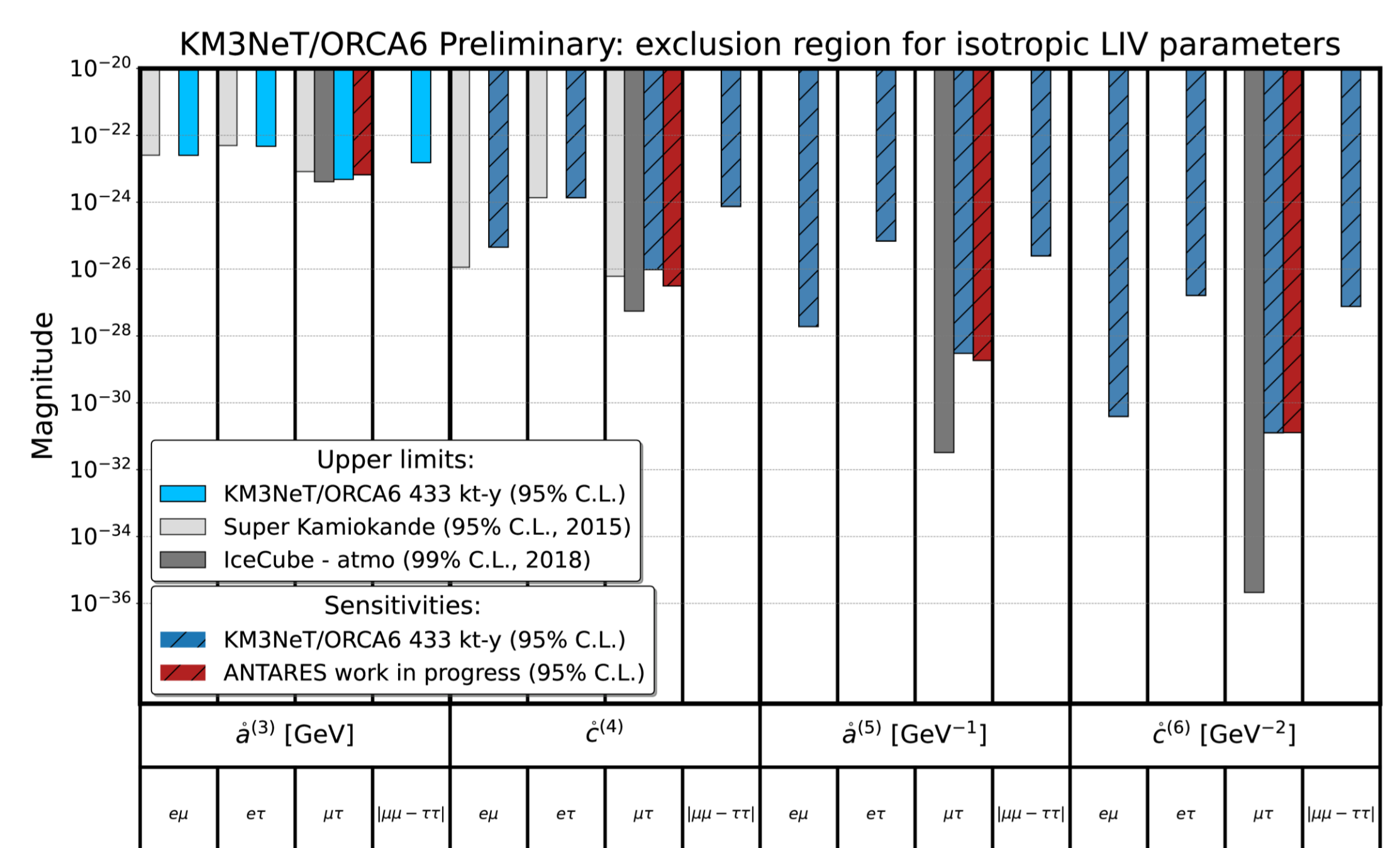
- KM3NeT/ORCA6**
  - Events are sorted into three different classes: high-purity neutrino-induced muon tracks, low-purity tracks, and showers
  - Boosted decision tree scores determine the class of an event
  - Most events have energies between 2 GeV and 100 GeV
- ANTARES**
  - Only high-purity track class is used so far
  - Event class determined by set of cuts
  - Random grid search<sup>3</sup> tries out and evaluates 100k cuts
- Best cuts form Pareto front
- Color of Pareto front cut indicates  $\chi^2$  resulting from fit to LIV model with  $\hat{a}_{\mu\tau}^{(5)} = 2 \cdot 10^{-29} \text{ GeV}^{-1}$
- Pareto front cut achieving the highest  $\chi^2$  defines high-purity tracks
- Most events have energies between 100 GeV and 1 TeV



## Results



We aim to put the first constraints on several LIV coefficients!



## References

<sup>1</sup>Colladay, Kostelecký (1998): Phys. Rev. D 58, 116002

<sup>2</sup>Kostelecký, Mewes (2012): Phys. Rev. D 85, 096005

<sup>3</sup>Bhat et al. (2018): Computer Physics Communications 228, 245–257

SK (2015): Phys. Rev. D 91, 052003

IC (2018): Nature Phys 14, 961–966