



Cailian Jiang

cailianjiang@smail.nju.edu.cn

Poster ID: 324

Sensitivity to invisible modes of neutron decay on JUNO

Cailian Jiang¹, Yuxiang Hu² On behalf of the JUNO collaboration¹NanJing University, ²Institute of High Energy Physics

1. JUNO experiment

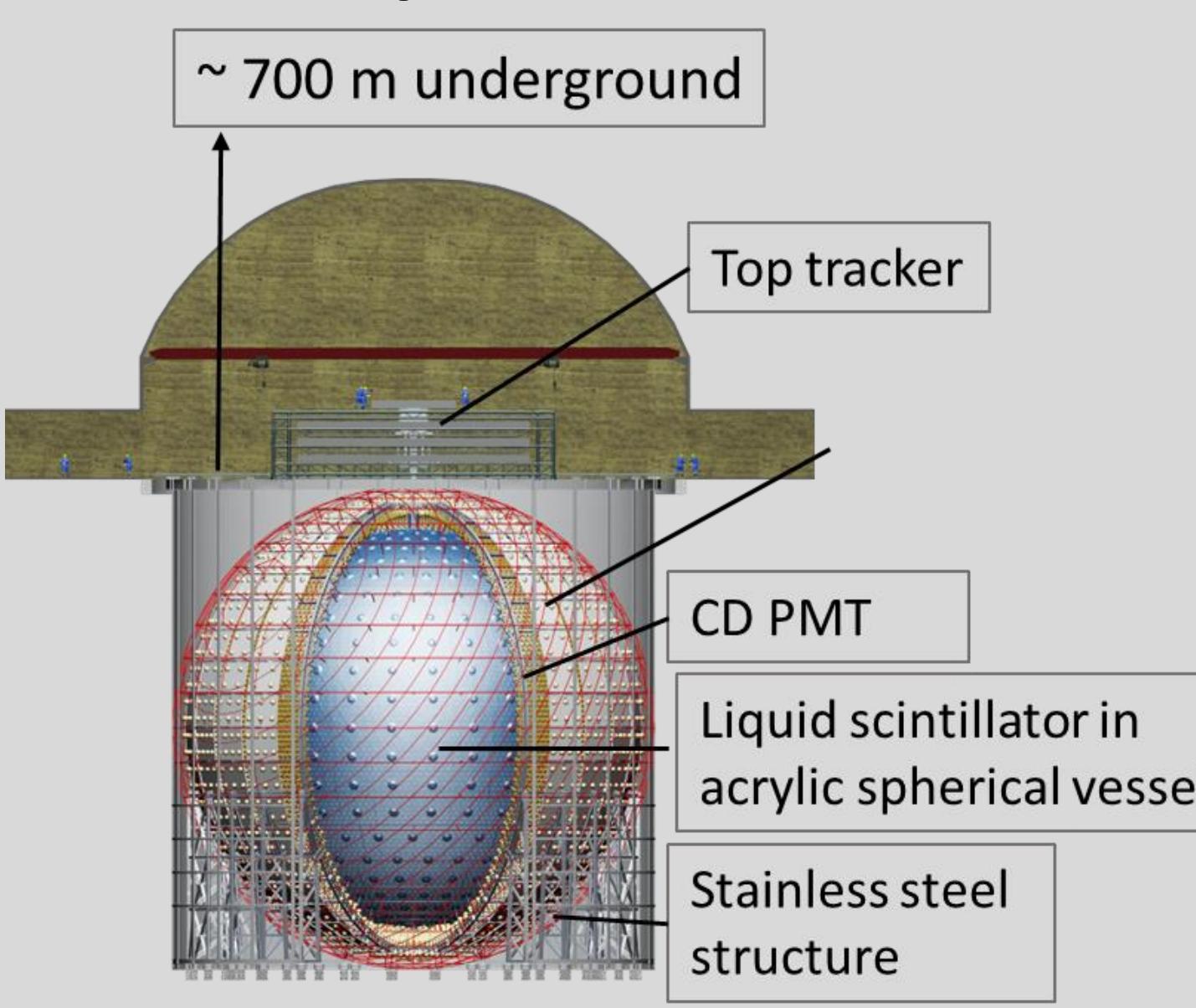
Jiangmen Underground Neutrino Observatory (JUNO)

Multi-purpose medium baseline reactor neutrino experiment^[1]

• Physics goals:

- Neutrino mass ordering
- Solar neutrino, atmospheric neutrino, supernova neutrino...
- new physics
- nucleon decay searches

Photon Statistics	1665 p.e./MeV
PMT coverage	77%
LS transparency	> 20 m
Light yield(anthracene)	45%
Detection Eff.(QE×CE)	30%
Target mass	20 kt



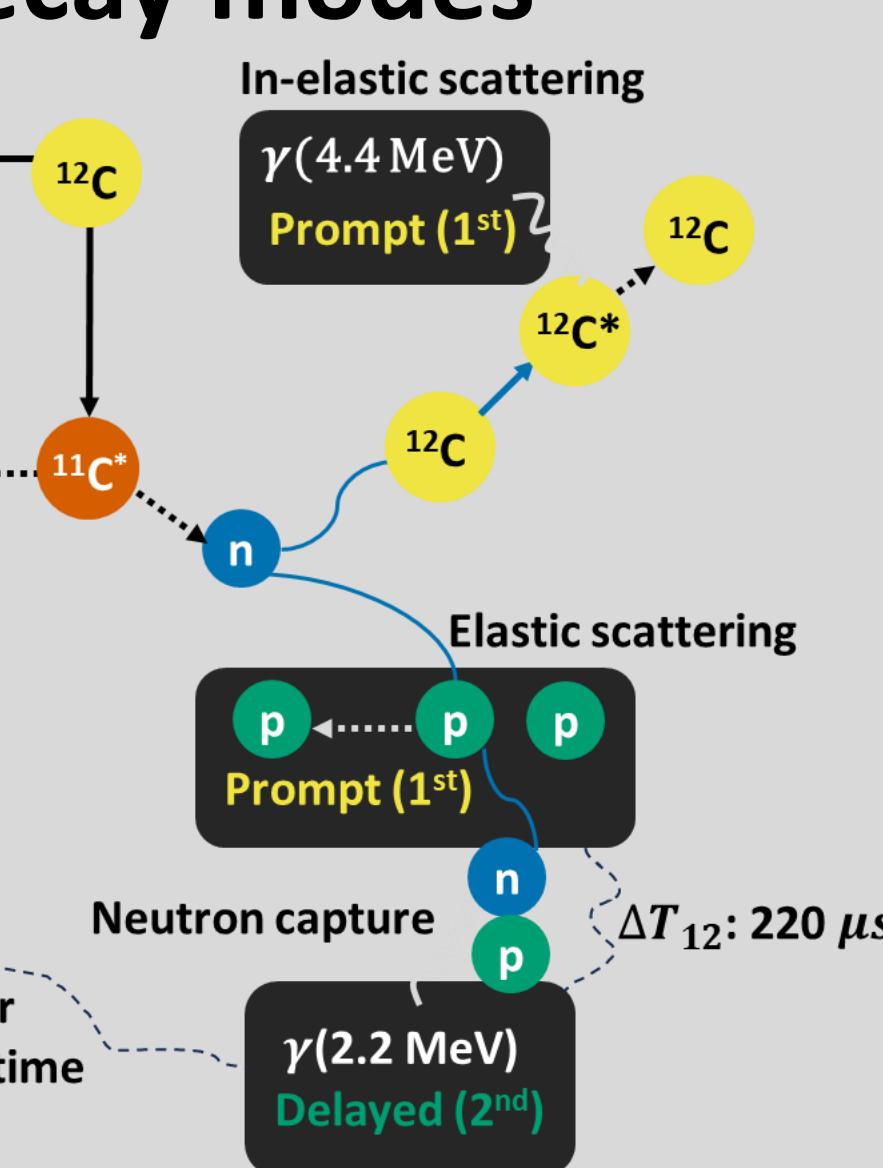
3. Invisible Neutron Decay Search in JUNO

• Bounded neutrons in ^{12}C : two invisible decay modes

- $n \rightarrow \text{invisible}$ ($^{12}\text{C} \rightarrow ^{11}\text{C}^*$)
- $nn \rightarrow \text{invisible}$ ($^{12}\text{C} \rightarrow ^{10}\text{C}^*$)

• De-excitation modes in LS

- $^{11}\text{C}^* \rightarrow n + ^{10}\text{C}$ ($\text{Br}_{n1} = 3.0\%$)
- $^{11}\text{C}^* \rightarrow n + \gamma + ^{10}\text{C}$ ($\text{Br}_{n2} = 2.8\%$)
- $^{10}\text{C}^* \rightarrow n + ^9\text{C}$ ($\text{Br}_{nn1} = 6.2\%$)
- $^{10}\text{C}^* \rightarrow n + p + ^8\text{B}$ ($\text{Br}_{nn2} = 6.0\%$)

Above modes can form triple coincidence signals^[4]

5. Backgrounds

Background source

• Single

- Radioactivity, Long-lived isotopes

• Single + Single + Single

Triple

• Correlated (Prompt-Delayed)

- IBD (Inverse Beta decay), Li9/He8, Fast neutron, Alpha-n, Atm – ν NC

• Triple (Prompt-Delayed-Decay)

- Atm – ν NC \rightarrow $\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + n + ^{11}\text{C}$, $\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + 2n + ^{10}\text{C}$, $\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + 3p + n + ^8\text{Li}$

$\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + n + p + ^{10}\text{B}$,
 $\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + n + p + \alpha + ^6\text{Li}$,
 $\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + n + 2p + ^9\text{Be}$,
 $\nu/\bar{\nu} + ^{12}\text{C} \rightarrow \nu/\bar{\nu} + n + p + d + ^8\text{Be}$.

Background suppression

• Combine two suppression method

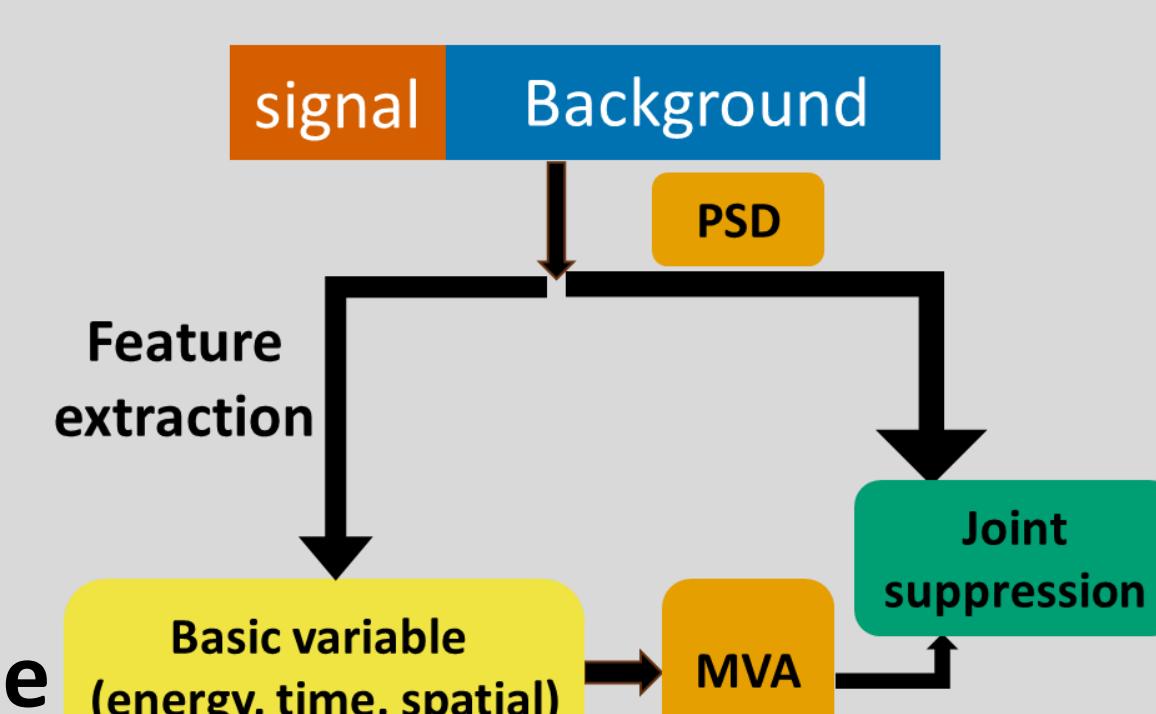
- Pulse Shape Discrimination

- Particle's emission photon time are different

- Multi Variate Analysis

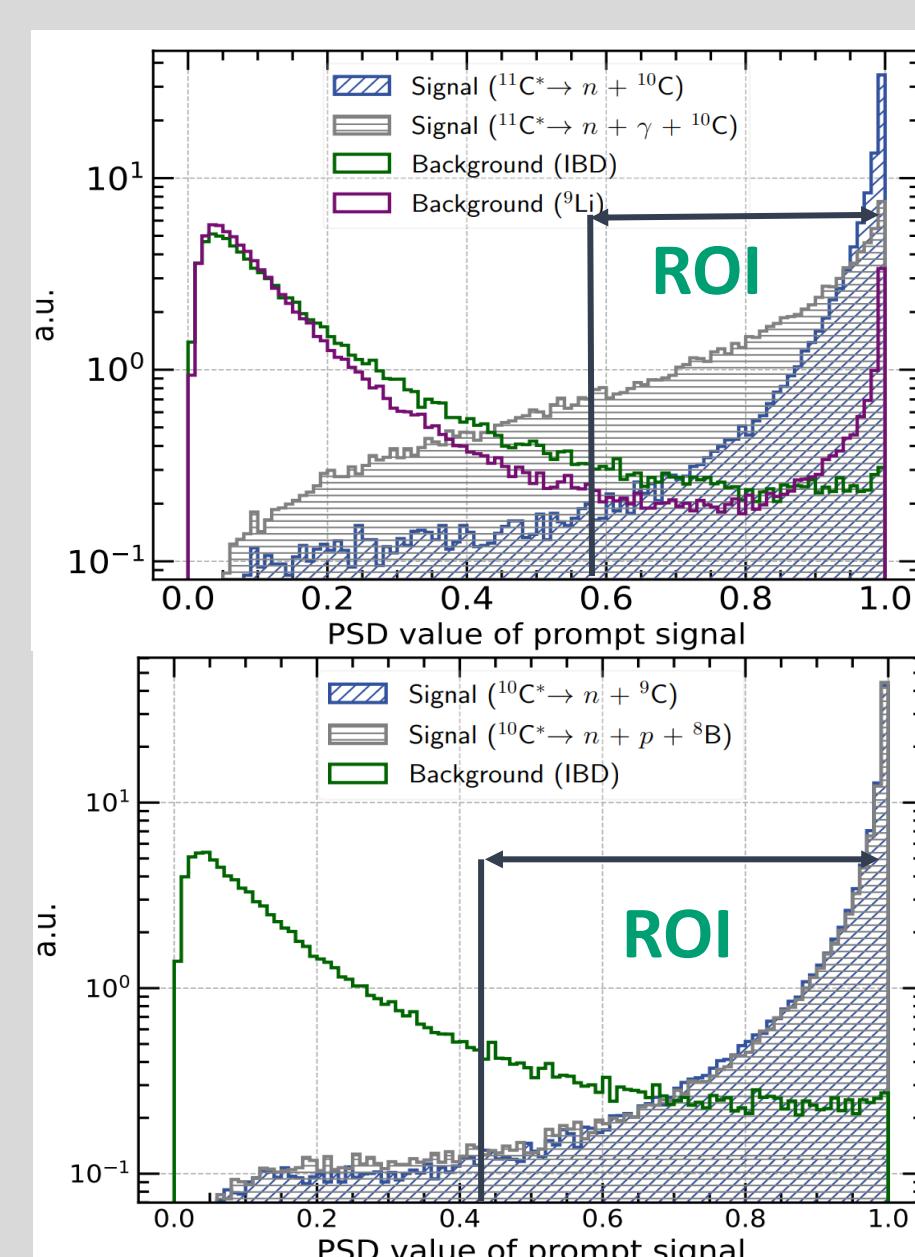
- Combine multidimensional features

- Both PSD and MVA have good performance
- Effectively suppress background



PSD

MVA

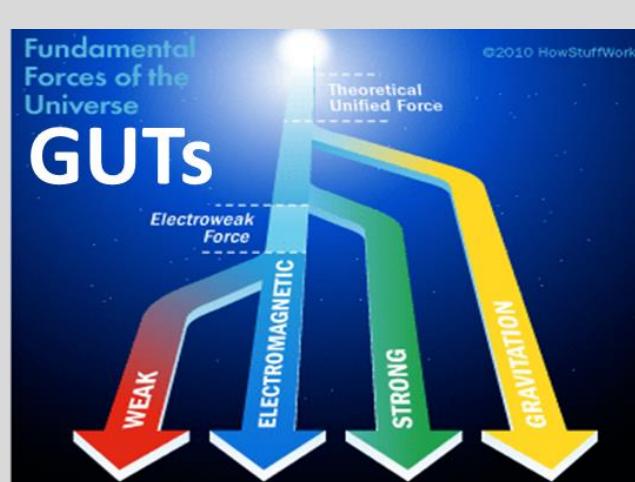
 $n \rightarrow \text{inv}$ $nn \rightarrow \text{inv}$

ROI: Region of interest

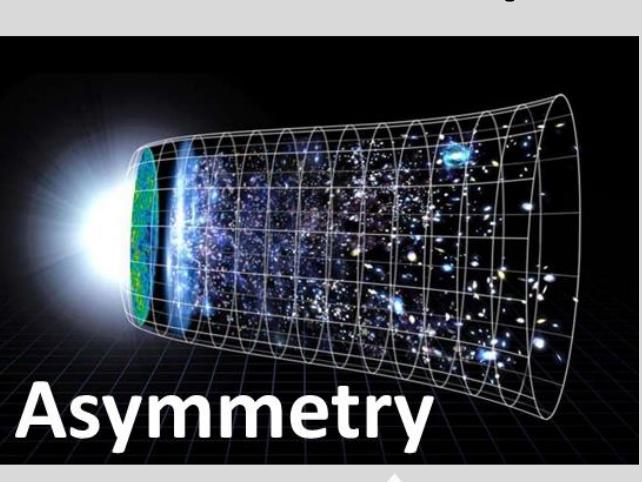
2. Motivation

• Predict the instability of nucleon^[2]

- Nucleon can decay into lighter subatomic particles
- Invisible modes of neutron decay
- neutron decay into undetected particles
- Search nucleon decay to test GUTs



Predicted



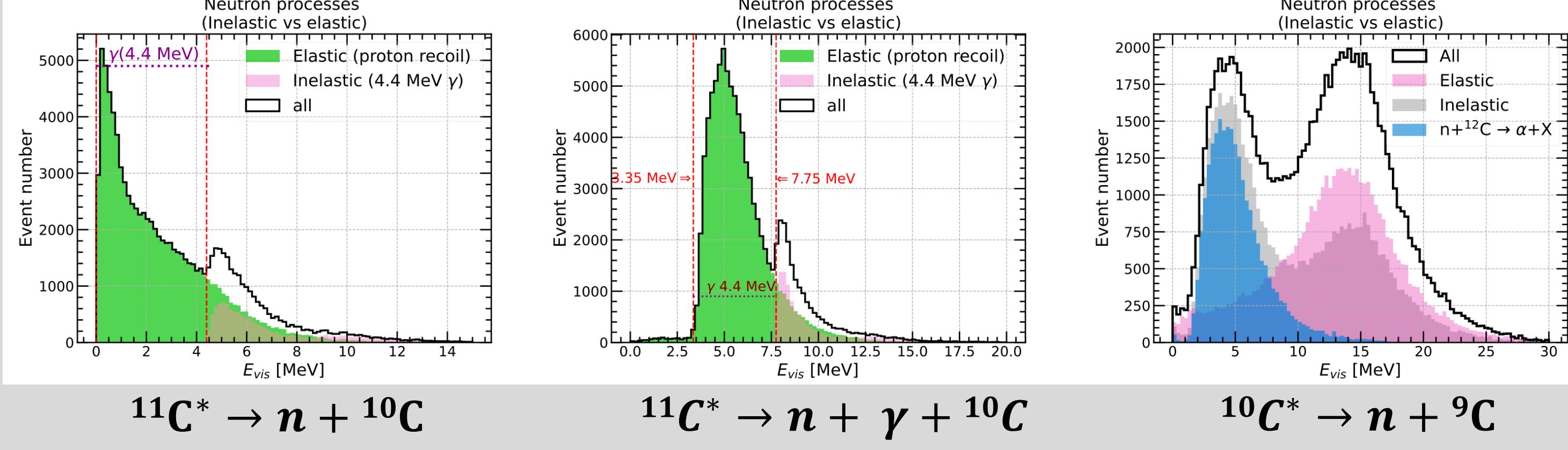
Basic ingredients

- Matter-antimatter is asymmetric in universe

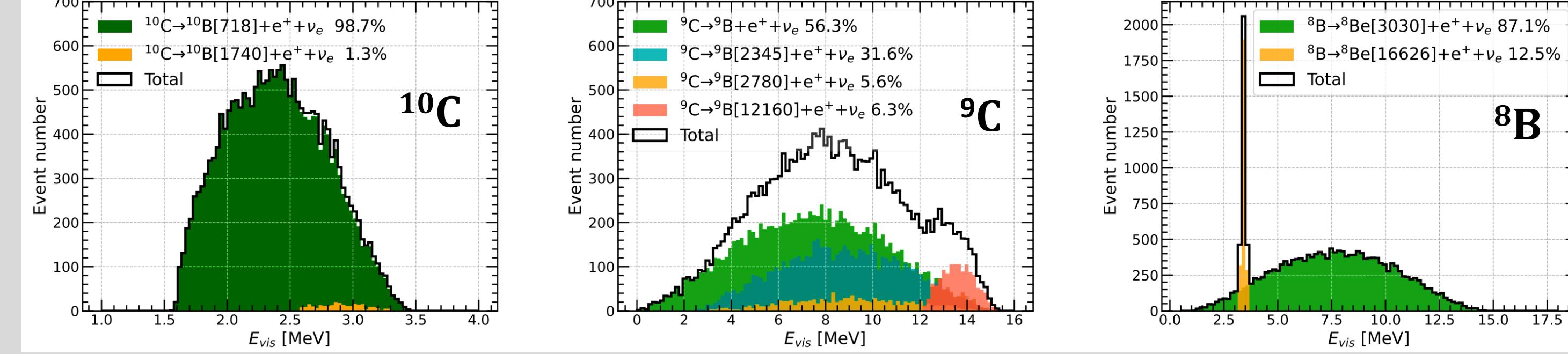
- Sakharov conditions^[3]
 - explain the asymmetry
 - Baryon number violation $\Delta B \neq 0$
 - C and CP violation
 - departure from thermal equilibrium

4. Signal

• Visible energy of prompt signal

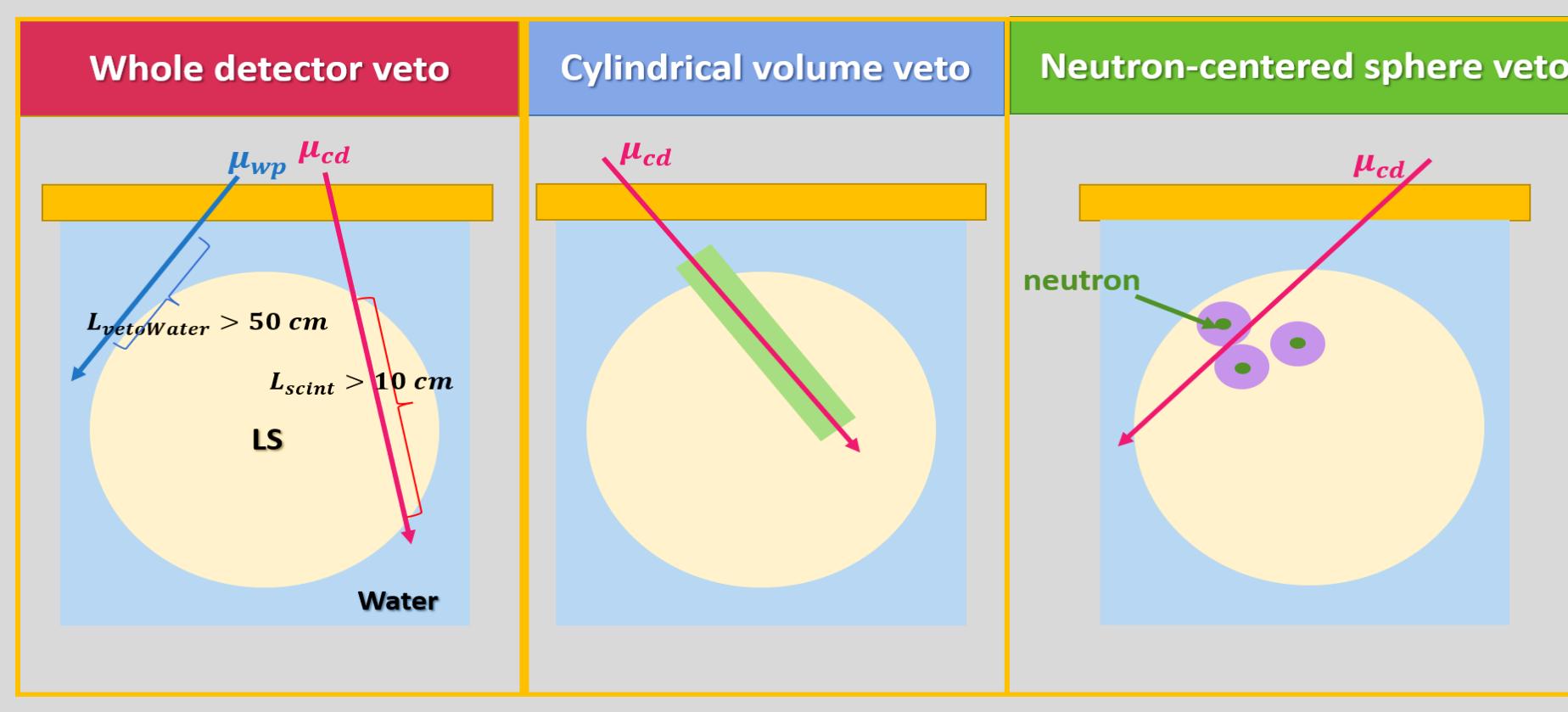


• Visible energy of decay signal



• Event selection

• Muon veto and Basic event selection

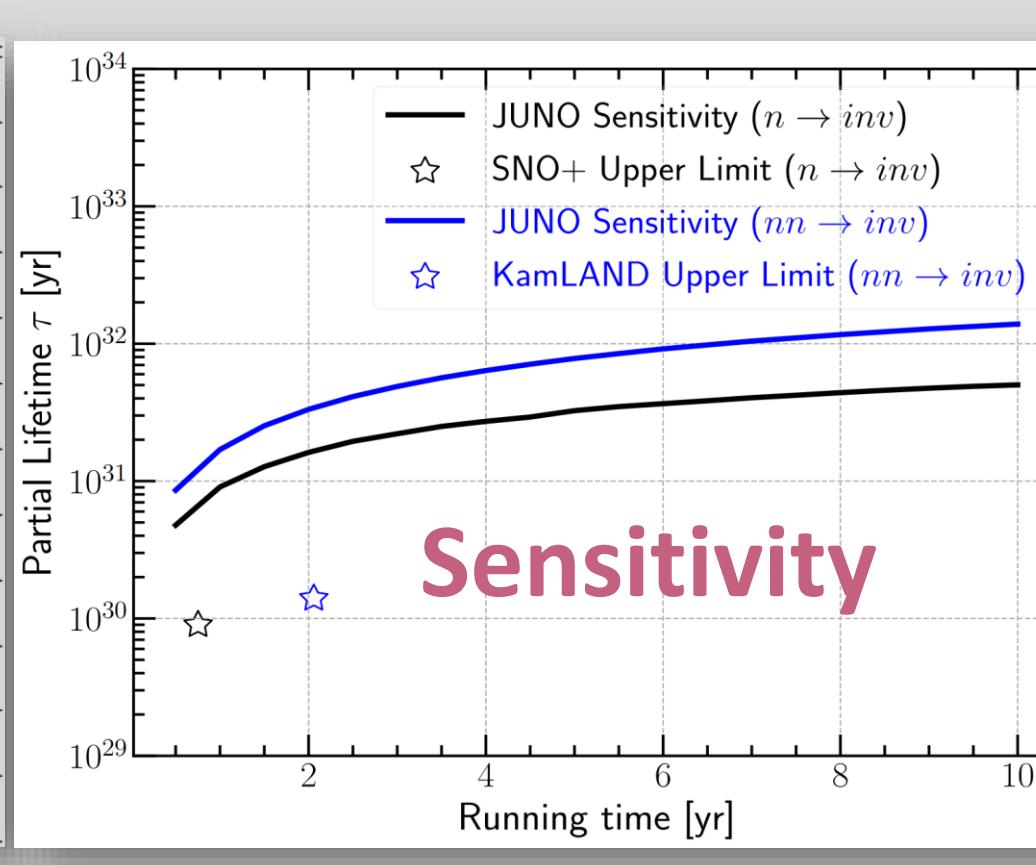


Quantity	$n \rightarrow \text{inv}$	$nn \rightarrow \text{inv}$
$R_{1,2,3}$ [m]	< 16.7	< 16.7
E_1 [MeV]	0.7-12	0.7-30
E_2 [MeV]	1.9-2.5	1.9-2.5
E_3 [MeV]	3.0-3.5	3.0-16.0
ΔT_{12} [ms]	< 1	< 1
ΔT_{23} [s]	0.002-100	0.002-3.0
ΔR_{12} [m]	< 1.5	< 1.5
ΔR_{23} [m]	< 1.5	< 1.5
ΔR_{13} [m]	< 1.0	< 1.0

6. Sensitivity and Conclusion

• Consider 10 years' JUNO data taking

Backgrounds (10 years)	$n \rightarrow \text{inv}$		$nn \rightarrow \text{inv}$	
	Basic selection	PSD + MVA	Basic selection	PSD + MVA
IBD + Single	1235 ± 50	2.72 ± 0.10	3.01 ± 0.09	0.0110 ± 0.0003
Atm- ν NC	3.0 ± 1.1	0.93 ± 0.67	4.3 ± 3.5	0.55 ± 0.63
$^{13}\text{C}(\alpha, n)^{12}\text{C}$ + Single	3.4 ± 1.4	0.036 ± 0.013	—	—
$^{9}\text{Li}/^{8}\text{He}$ + Single	1.55 ± 0.39	0.29 ± 0.17	0.13 ± 0.13	0.13 ± 0.13
Accidental	1.46 ± 0.05	0.095 ± 0.004	—	—
Total	1244 ± 50	4.07 ± 0.68	7.4 ± 3.5	0.69 ± 0.64
Signal efficiency (%)	$n \rightarrow \text{inv}$		$nn \rightarrow \text{inv}$	
ϵ_{nn1}	35.6 ± 0.2	23.5 ± 0.2	54.0 ± 0.3	48.2 ± 0.3
ϵ_{nn2}	43.6 ± 0.3	30.3 ± 0.3	49.2 ± 0.3	36.3 ± 0.3

 $\tau/B(n \rightarrow \text{inv}) > 5.0 \times 10^{31}$ year at 90% C.L. $\tau/B(nn \rightarrow \text{inv}) > 1.4 \times 10^{31}$ year at 90% C.L.

An order of magnitude improvement to current best limits with 2 years data taking

7. References

- [1] F. An et al. [JUNO] J. Phys. G 43, no.3, 030401 (2016).
- [2] P. Nath and P. Fileviez Perez, Phys. Rept. 441, 191-317 (2007).
- [3] A. D. Sakharov, Sov. Phys. Usp. 34, 392 (1991).
- [4] Y. A. Kamshkov and E. Kolbe, Phys. Rev. D 67, 076007 (2003).

