



Understanding gravitationally induced decoherence parameters in neutrino oscillations using a microscopic quantum mechanical model

Alba Domi^{1,3}, Thomas Eberl¹, Max Joseph Fahn^{1,2}, Kristina Giesel^{1,2}, Lukas Hennig¹, Ulrich Katz¹, Roman Kemper^{1,2} and Michael Kobler^{1,2}

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¹ Erlangen Centre for Astroparticle Physics (ECAP), Friedrich-Alexander-Universität Erlangen-Nürnberg, Nikolaus-Fiebiger-Str. 2, 91058 Erlangen, Germany

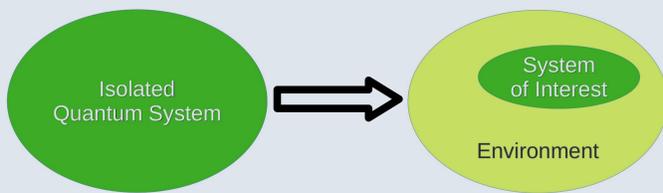
² Institute for Quantum Gravity, Theoretical Physics III, Friedrich-Alexander-Universität Erlangen-Nürnberg, Staudtstr. 7, 91058 Erlangen, Germany

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Neutrinos and Gravity

- Neutrinos interact with gravity \implies influenced by gravitational waves \implies changes oscillation behaviour (gravitationally induced decoherence)
- Determining this influence...
 - ... helps to better understand neutrino oscillations
 - ... yields information on stochastic gravitational waves that are not directly detectable, produced by several sources after the Big Bang
- Many works using phenomenological models (e.g. [1, 7, 8, 10]) \longleftrightarrow connection to underlying microscopic physics not always immediate \implies Here: Study a microscopic quantum mechanical model to connect to phenomenological models

Neutrinos as an open quantum system



- Investigate effective dynamics of a neutrino (system of interest $\hat{\rho}_S$) interacting with gravitational waves (environment) without the necessity to track their detailed dynamics
- Master equation for effective evolution of a neutrino:

$$\frac{\partial}{\partial t} \hat{\rho}_S(t) = -\frac{i}{\hbar} [\hat{H}_S + \hat{H}_{\text{add}}, \hat{\rho}_S(t)] + \mathcal{D}[\hat{\rho}_S(t)]$$

\implies **Red terms** have to be postulated (phenomenological models) or derived

- New processes compared to an isolated quantum system encoded in red terms:
 - **Energy shifts/renormalisation** of the energies of $\hat{\rho}_S$
 - **Dissipation** (energy flux from the system into the environment)
 - **Decoherence** (information flux into the environment) \implies diagonalisation ("classicalisation") of $\hat{\rho}_S$ in a certain basis

Microscopic quantum mechanical model

- Based on [11], extended for neutrinos in [3]
- System of interest: Neutrino, Hamiltonian \hat{H}_S
- Environment: Gravitational waves modeled by a bath of Harmonic oscillators with frequencies ω_i , positions (configuration variables) \hat{q}_i and canonically conjugated momenta \hat{p}_i
- Coupling: motivated by General Relativity, where the energy-momentum tensor of matter couples to the metric (= configuration variable) of the gravitational field; coupling constants g_i
- Total Hamiltonian of the quantum mechanical model:

$$\hat{H} = \underbrace{\hat{H}_S^{(0)} + \hat{H}_S^{(C)}}_{\hat{H}_S} \otimes \hat{\mathbb{1}}_\mathcal{E} + \hat{\mathbb{1}}_S \otimes \underbrace{\frac{1}{2} \sum_{i=1}^N [\hat{p}_i^2 + \omega_i^2 \hat{q}_i^2]}_{\hat{H}_\mathcal{E}} - \underbrace{\hat{H}_S \otimes \sum_{i=1}^N g_i \hat{q}_i}_{\hat{H}_{\text{int}}}$$

with counter term $\hat{H}_S^{(C)}$ that renormalises a Lamb-shift-like contribution to the master equation

Derivation of the master equation

- Assumptions:
 - Interaction weak compared to the free evolution of the neutrino
 - Gravitational waves follow a Bose-Einstein distribution moderated by a temperature parameter T
 - Correlation functions in the environment decay on time scales much shorter than the state of the neutrino varies (Markov assumption; holds here, see [3]) \implies coarse-graining, send $t_0 \rightarrow -\infty$
- Master equation (Lindblad form):

$$\frac{d}{dt} \hat{\rho}_S(t) = -\frac{i}{\hbar} [\hat{H}_S^{(0)}, \hat{\rho}_S(t)] + \frac{8\eta^2 k_B T}{\hbar^2} \left(\hat{H}_S^{(0)} \hat{\rho}_S(t) \hat{H}_S^{(0)} - \frac{1}{2} \left\{ (\hat{H}_S^{(0)})^2, \hat{\rho}_S(t) \right\} \right)$$

- Free parameters: T, η

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Neutrino setup

- Consider neutrinos that propagate through the Earth
- System Hamiltonian in vacuum mass basis:

$$\hat{H}_S^{(0)} = \hat{H}_{\text{vac}} + \hat{U}^\dagger \hat{H}_{\text{mat}} \hat{U},$$

with

$$\hat{H}_{\text{vac}} = E \mathbb{1}_3 + \frac{c^4}{6E} \begin{pmatrix} -\Delta m_{21}^2 - \Delta m_{31}^2 & 0 & 0 \\ 0 & \Delta m_{21}^2 - \Delta m_{32}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 + \Delta m_{32}^2 \end{pmatrix}, \quad \hat{H}_{\text{mat}} = \pm \sqrt{2} G_f N_e \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

- Three bases:
 - Flavour basis
 - Vacuum mass basis: $\hat{H}_S^{(0)}$ is diagonal for $N_e = 0$, i.e. for vacuum. Connected to flavour basis by PMNS matrix \hat{U} .
 - Effective mass basis: $\hat{H}_S^{(0)}$ is diagonal (with eigenvalues \tilde{H}_i) for a fixed N_e , i.e. for a specific layer of density of the Earth. Connected to vacuum mass basis by $\tilde{V}(N_e)$.
- Solution of the master equation in effective mass basis:

$$\tilde{\rho}_{ij}(t) = \tilde{\rho}_{ij}(0) \cdot e^{-\frac{i}{\hbar}(\tilde{H}_i - \tilde{H}_j)t - \frac{\gamma_{ij}^2 k_B T}{\hbar^2}(\tilde{H}_i - \tilde{H}_j)^2 t}$$

and in vacuum mass basis:

$$\rho_{ij}(t) = \rho_{mn}(0) \tilde{V}_{km}^\dagger \tilde{V}_{nl} \tilde{V}_{ij}^\dagger e^{-\frac{i}{\hbar}(\tilde{H}_k - \tilde{H}_l)t - \frac{\gamma_{ij}^2 k_B T}{\hbar^2}(\tilde{H}_k - \tilde{H}_l)^2 t}$$

Results

- Connection to phenomenological models:
 - Phenomenological models: specific Lindblad form for dissipator $\mathcal{D}[\hat{\rho}_S(t)]$ \implies Solution of the master equation:

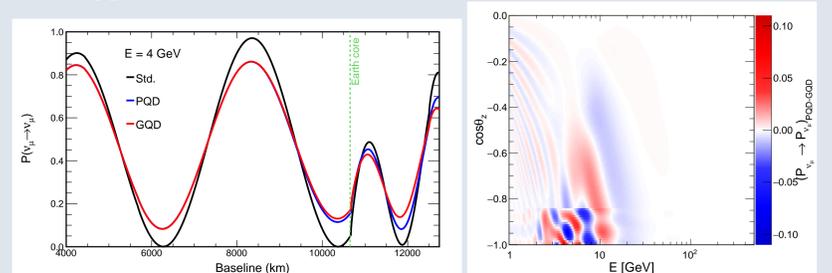
$$\tilde{\rho}_{ij}(t) = \tilde{\rho}_{ij}(0) \cdot e^{-\frac{i}{\hbar}(\tilde{H}_i - \tilde{H}_j)t - \Gamma_{ij} t} \quad \text{with } \Gamma_{ij} = \gamma_{ij} E^n$$

Then restrict number of parameters by considering specific cases where γ_{ij} are zero or equal to each other

- Implications from our quantum mechanical model in vacuum:

$$\gamma_{ij} = \frac{\eta^2 c^8 k_B T}{\hbar^3} (\Delta m_{ij}^2)^2 \quad n = -2$$

- In matter: No match possible, as for the phenomenological models γ_{ij} is constant while in the quantum mechanical model it depends on the matter density $N_e \implies$ different oscillation probabilities (for $T = 0.9K, \eta = 10^{-8}s, n = -2$ and fitting values for γ_{ij} using PREM [4] and OscProb [2]):



\implies Analyses using the phenomenological ansatz from above can only constrain the free parameters in vacuum, not in matter

- Lamb-shift contribution: Renormalised using a counter term. Without renormalisation: dependence on unphysical arbitrary cutoff-frequency Ω of gravitational waves background, diverges for $\Omega \rightarrow \infty$ \implies **Interpretation of the Lamb-shift contribution without a renormalisation problematic**
- Free parameters:
 - T : "Temperature" parameter characterising the gravitational waves environment
 - η : Coupling strength between neutrino system and gravitational environment. Should be determined by the coupling of matter to gravity by General Relativity. Using a field-theoretical model from ([6, 9]), a naive comparison (see [3, 5]) yields a direct relation to the Planck length, in particular $\eta \approx 10^{-42}s$.

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