

Real-time Supernova Neutrinos Detection in XENONnT



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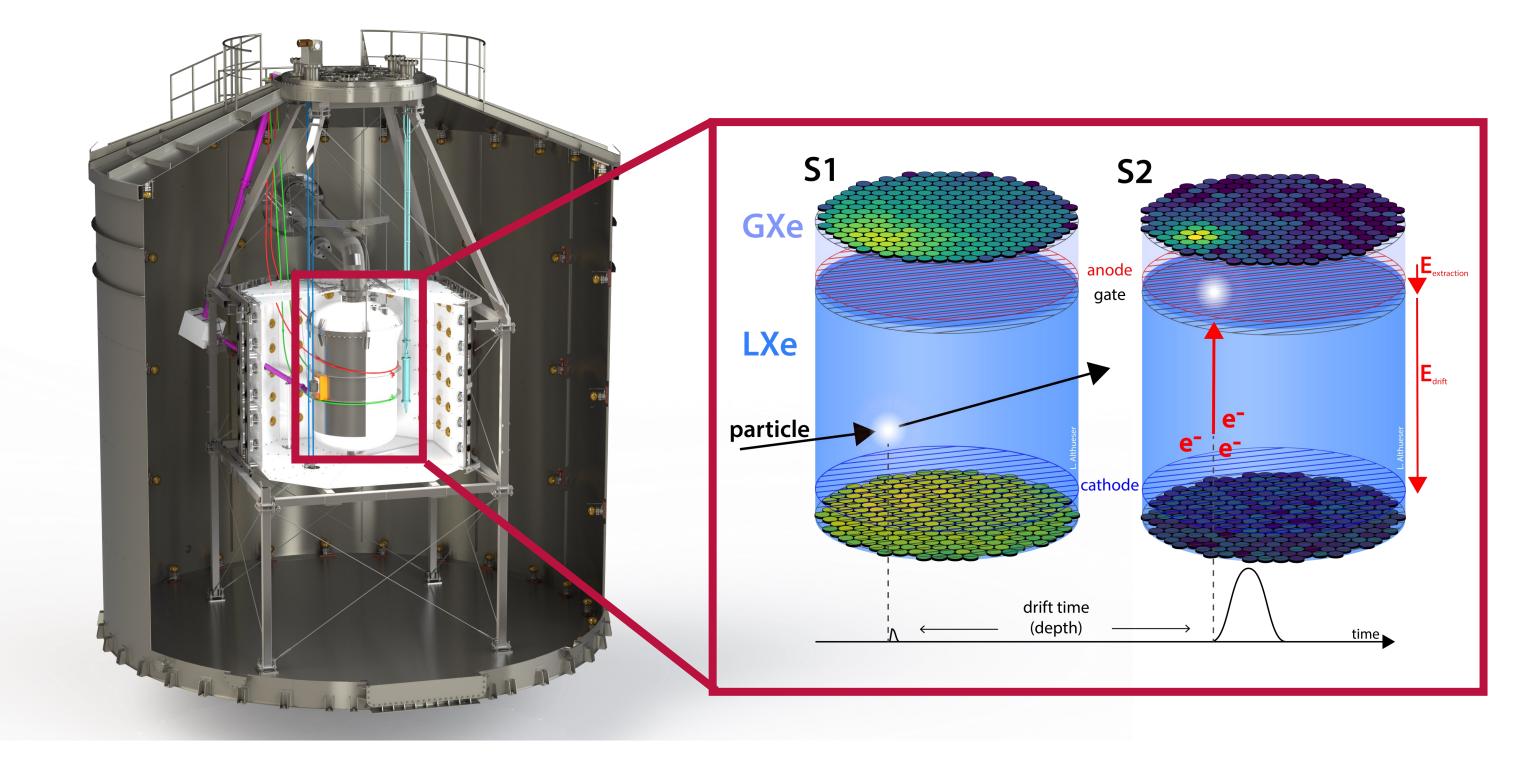


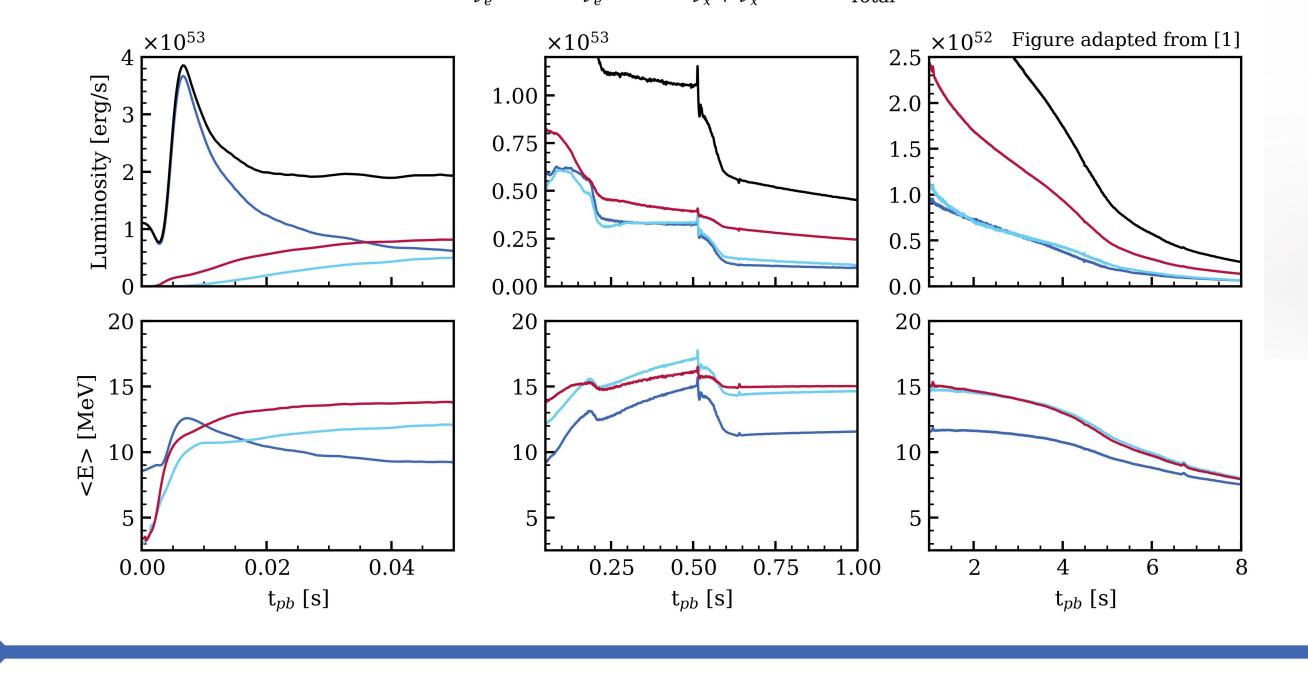
Supernova neutrinos

The XENONnT detector

- Type II core-collapse supernovae emit ~99% of their progenitor gravitational energy as neutrinos
- Neutrinos of O(10) MeV
- Burst-type event mostly in a 10 s window



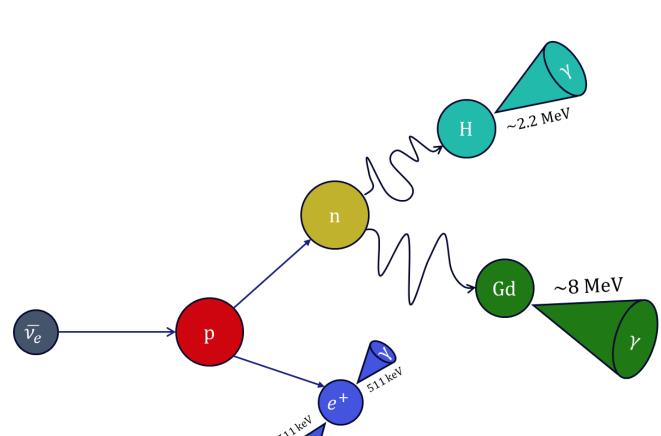




In the dual-phase Xe TPC

- Main interaction through coherent elastic neutrinonucleus scattering (CEvNS)
- Flavour-independent measurement

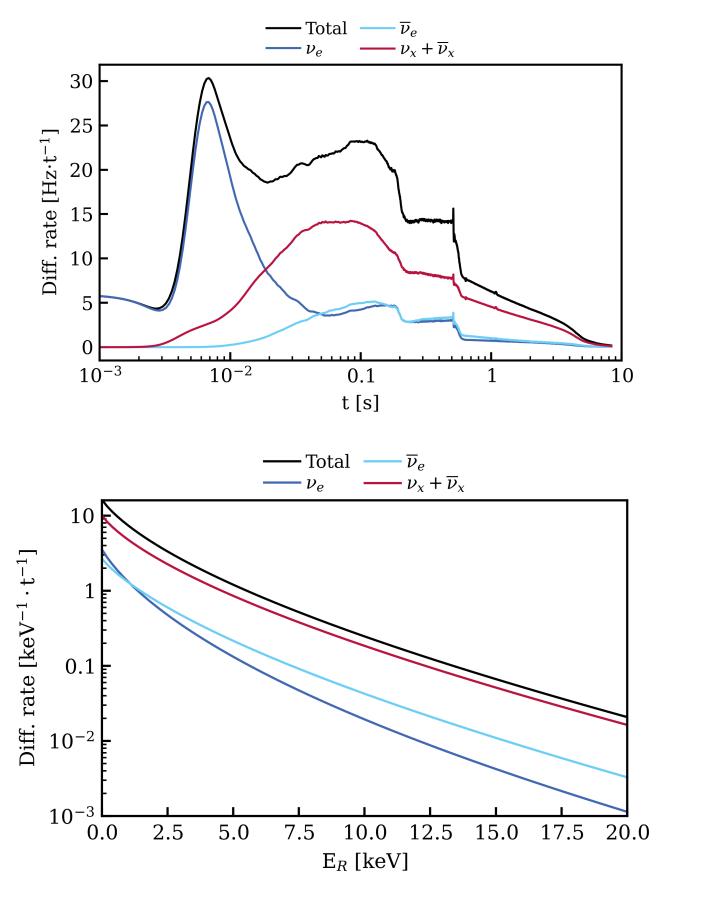
In veto detectors

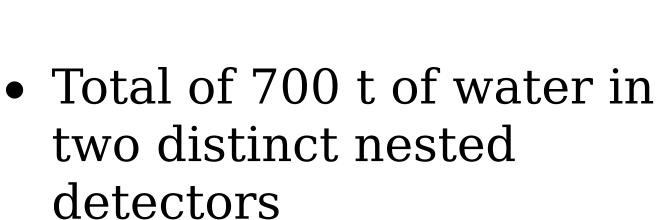


- Located at Laboratori Nazionali del Gran Sasso (LNGS) - 3600 m.w.e.
- Dual-phase xenon time projection chamber (TPC) with a 5.9 t active target [2]
 - Energy reconstruction
 - 3D position reconstruction
 - Particle discrimination (ER vs NR)
- Optimized for WIMP search and low-energy recoils O(10) keV

Multi-messenger astrophysics

• Low energy nuclear recoils of O(1) keV

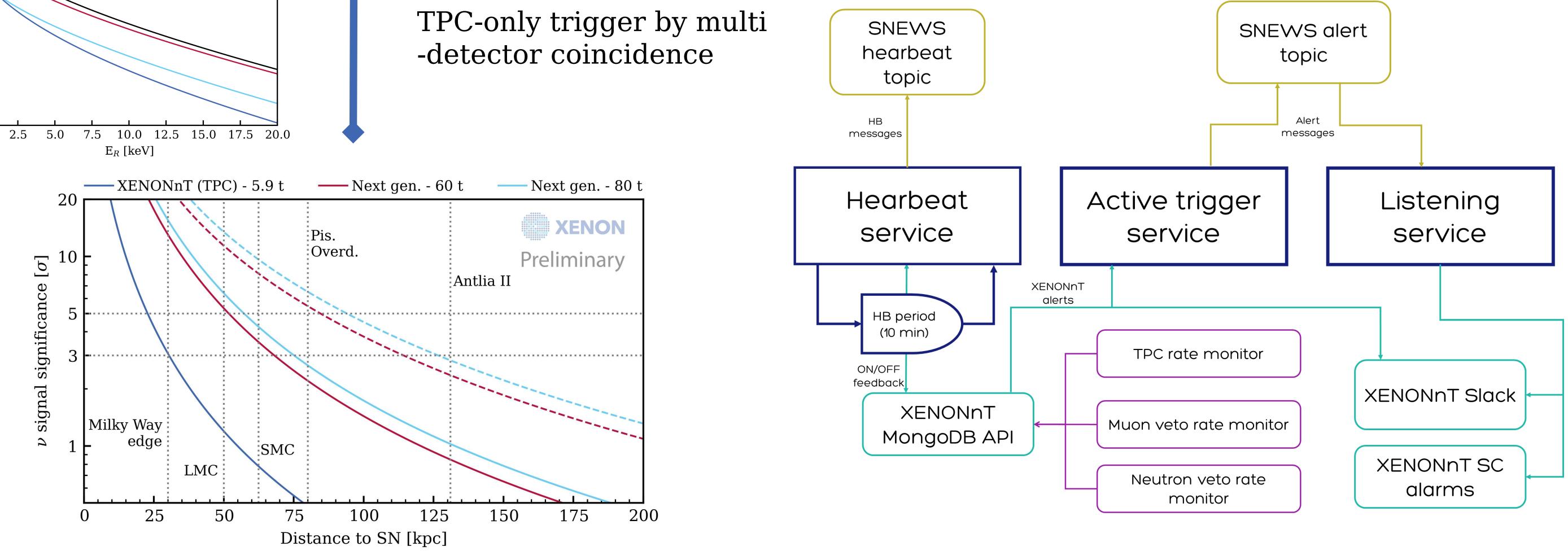




- Main interaction through Inverse Beta Decay (IBD), producing Cerenkov light
- Expecting 100-200 interactions
- Considerably reduce the false alarm rate of the -detector coincidence



- The SuperNova Early Warning System (SNEWS) looks for coincident triggers from neutrino experiments [3]
- XENONnT is the first DM-based experiment integrated in the upcoming SNEWS 2.0 [4]



[1] - A. Mirizzi, et al, Supernova Neutrinos: Production, Oscillations and Detection, Riv. Nuovo Cim. 39 (2016)

[2] - E. Aprile, et al. (XENON Collaboration), *The XENONnT Dark Matter Experiment*, e-print: 2402.10446

[3] - P. Antonioli, et al., SNEWS: The Supernova Early Warning System, New J.Phys. 6 (2004)

[4] - S. Al Kharusi, et al. (SNEWS collaboration), SNEWS 2.0: a next-generation supernova early warning system for multimessenger astronomy, New J.Phys. 23 (2021)

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