



An excursion into dense matter with Marcello



Arnau Rios Huguet

Ramón y Cajal Fellow

Institute of Cosmos Sciences

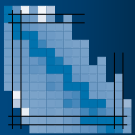
Universitat de Barcelona

arnau.rios@fga.ub.edu

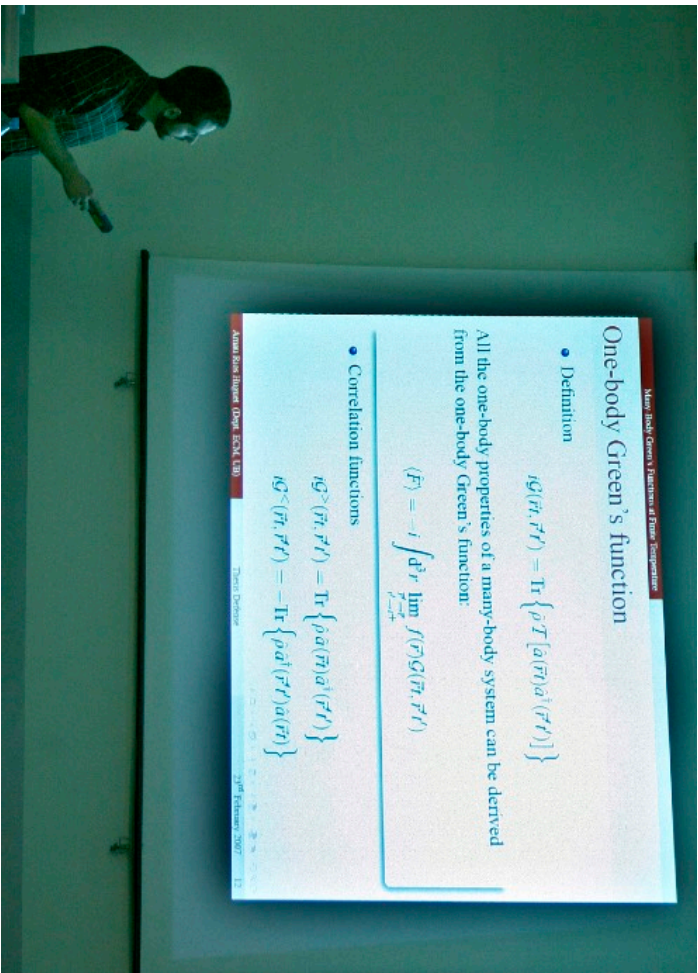
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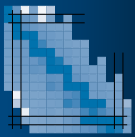


Celebrating Dr. Marcello
Baldo's 80th Birthday
16 October 2023



23 February 2007



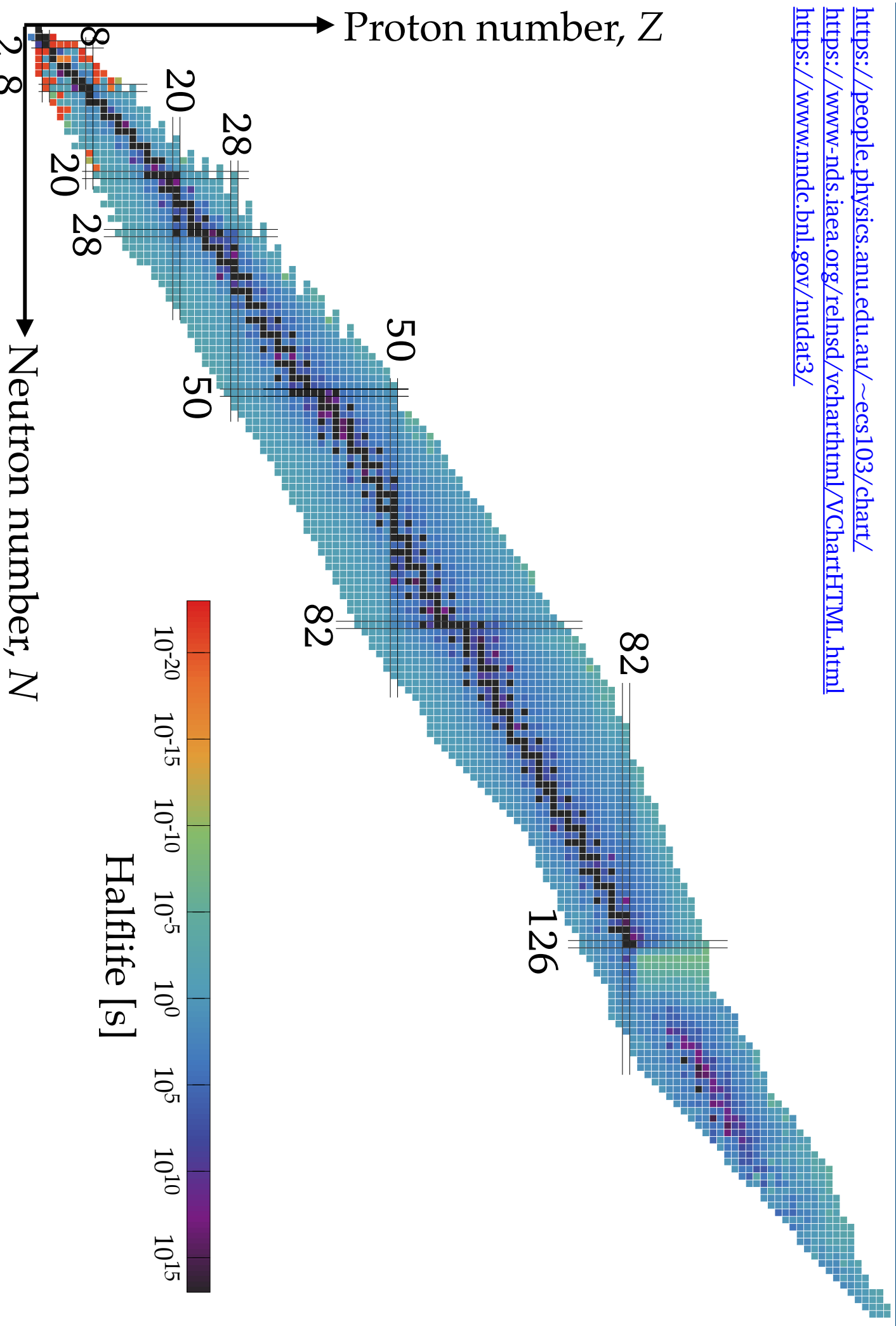


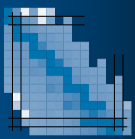
Isotope chart

<https://people.physics.anu.edu.au/~ecs103/chart/>

<https://www-nds.iaea.org/relnsd/vcharthtml/VChartHTML.html>

<https://www.nndc.bnl.gov/nudat3/>



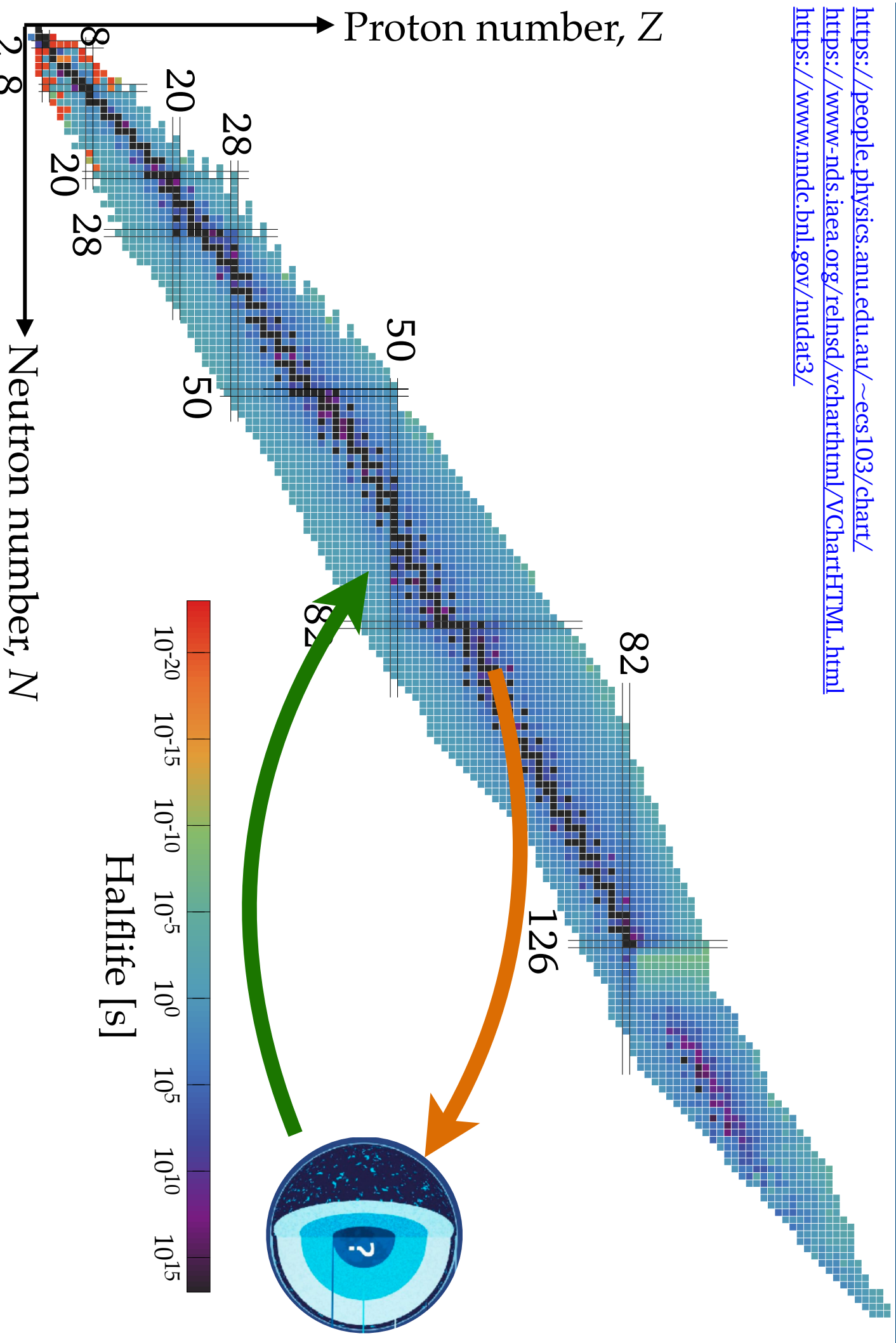


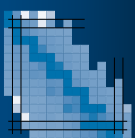
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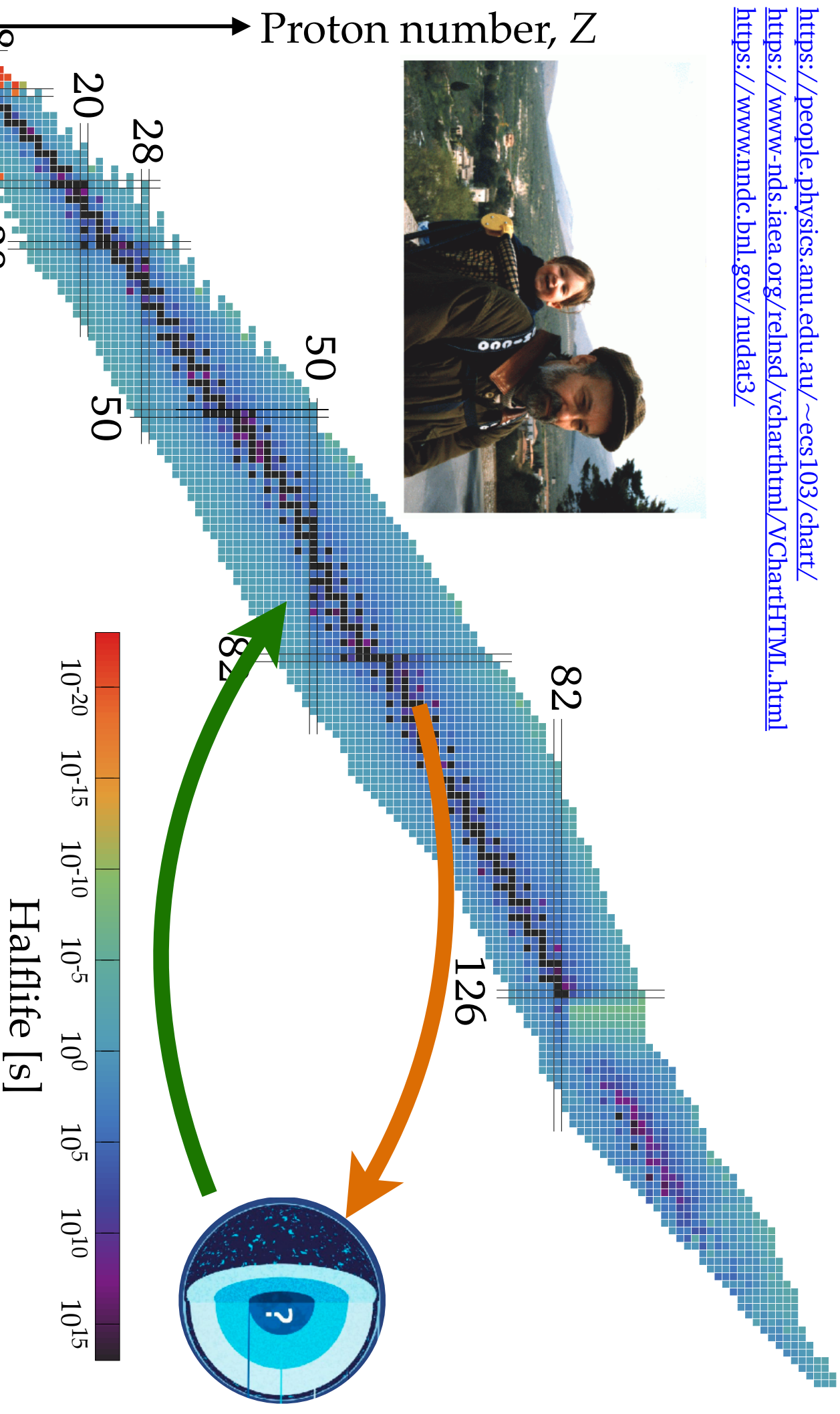


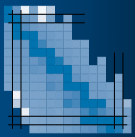
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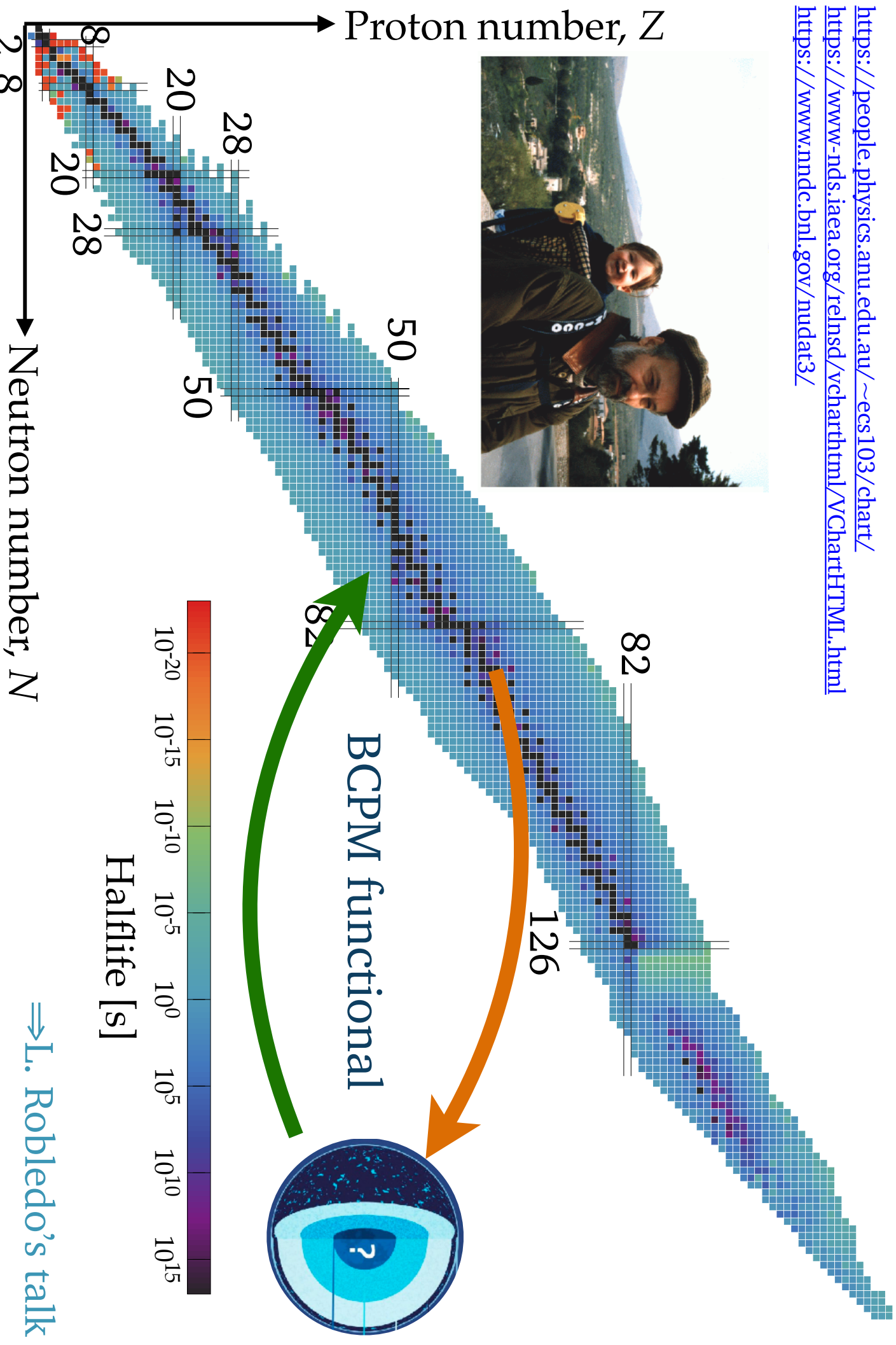


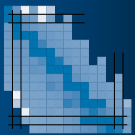
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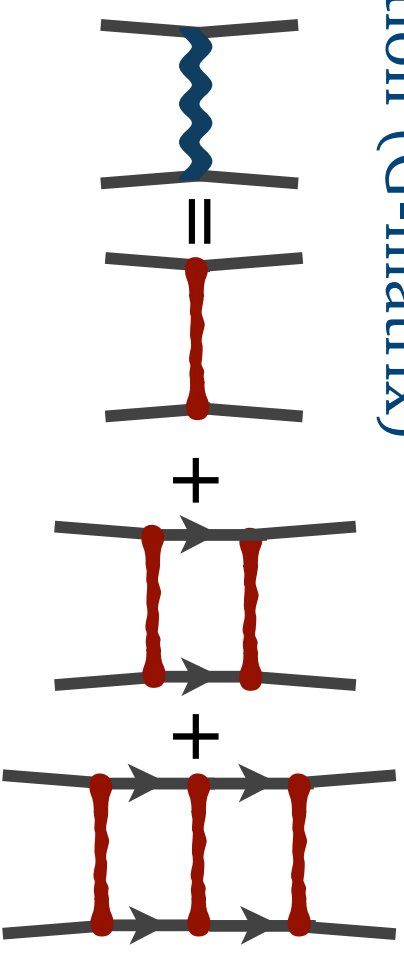


Brueckner-Bethe-Goldstone

- Infinite **resummation** of two-hole line diagrams (energy expansion)
- **pp** Pauli blocked **in-medium** interaction (G-matrix)

Effective interaction

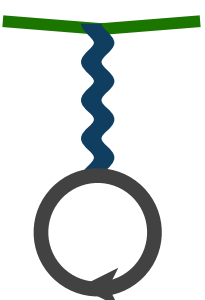
$$G(\omega) = V + V \frac{Q}{\omega - \epsilon - \epsilon' + i\eta} G(\omega)$$



Single-particle spectrum

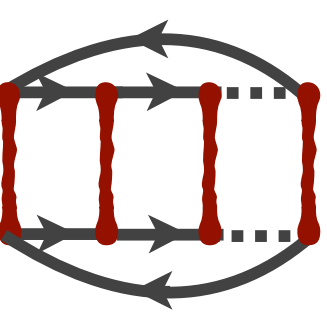
$$U(k) = \sum_{|\vec{k}'| < k_F} \langle \vec{k}\vec{k}' | G(\omega = \epsilon(k) + \epsilon(k')) | \vec{k}\vec{k}' \rangle_A$$

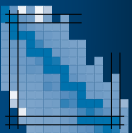
$$\epsilon(k) = \frac{\hbar^2 k^2}{2m_\tau} + \text{Re}[U(k)]$$



Energy

$$\frac{E}{A}(\rho, \beta) = \frac{1}{A} \sum_\tau \sum_{|\vec{k}| < k_{F\tau}} \left(\frac{\hbar^2 k^2}{2m_\tau} + \frac{1}{2} \text{Re}[U_\tau(\vec{k})] \right)$$





3-hole line “epics”

Three hole-line contribution in nuclear matter revisited

H.Q. Song^{a,b,c}, M. Baldo^{d,1}, G. Giansiracusa^e, U. Lombardo^e

^a Institute of Nuclear Research, Academia Sinica, Shanghai 201800, China

^b CCAST (World Laboratory), P.O. Box 8730, Beijing 100080, China

^c T.D. Lee Physics Laboratory, Fudan University, Shanghai 200433, China

^d INFN, Sezione di Catania, 57 Corso Italia, I-95129 Catania, Italy

^e Dipartimento di Fisica and INFN, 57 Corso Italia, I-95129 Catania, Italy

Song, Baldo, Giansiracusa, Lombardo,

Phys. Lett. B 411 237 (1997)

1, NUMBER 8

PHYSICAL REVIEW LETTERS

24 AUG

Bethe-Brueckner-Goldstone Expansion in Nuclear Matter

H. Q. Song,^{1,2,3} M. Baldo,^{4,*} G. Giansiracusa,⁵ and U. Lombardo⁵

¹Institute of Nuclear Research, Academia Sinica, Shanghai 201800, China

²CCAST (World Laboratory), P.O. Box 8730, Beijing 100080, China

³T. D. Lee Physics Laboratory, Fudan University, Shanghai 200433, China

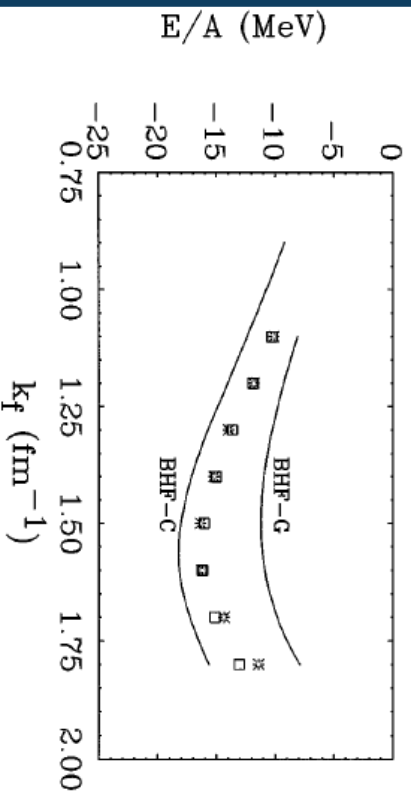
⁴INFN, Sezione di Catania, 57 Corso Italia, I-95129 Catania, Italy

⁵Dipartimento di Fisica and INFN, 57 Corso Italia, I-95129 Catania, Italy

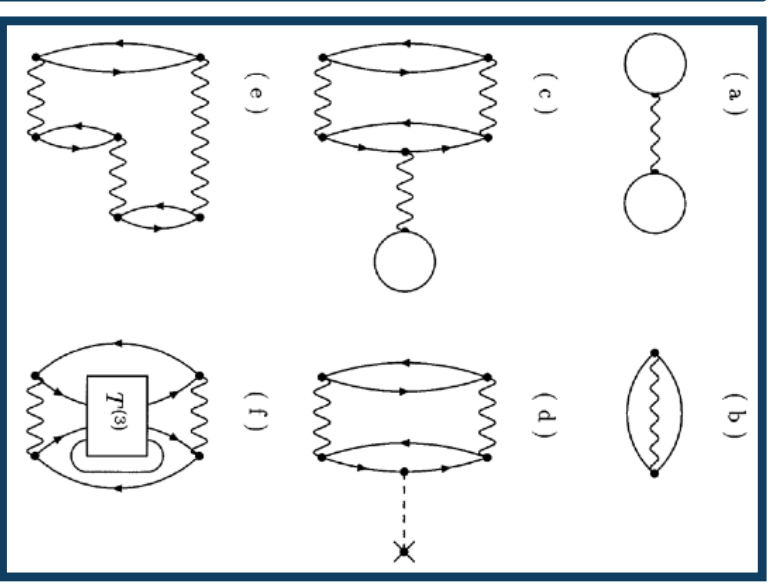
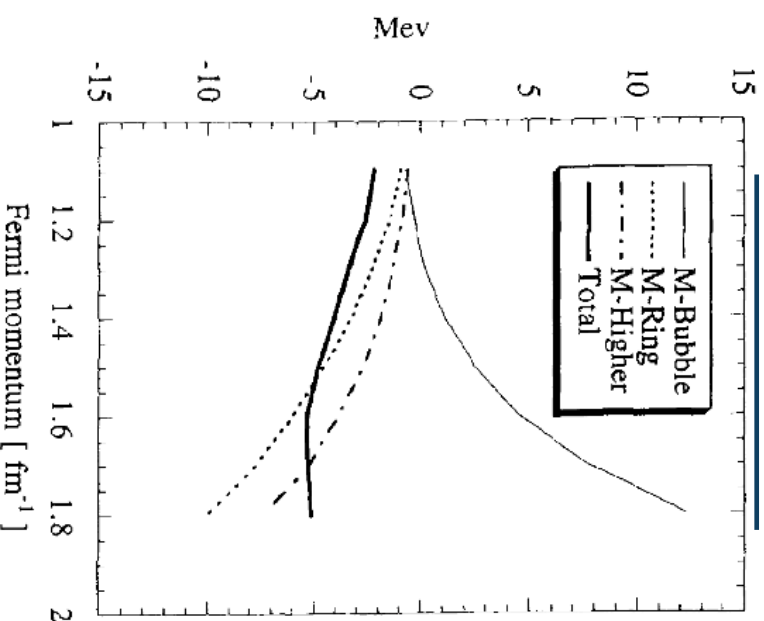
Song, Baldo, Giansiracusa, Lombardo,

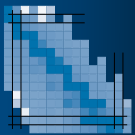
Phys. Rev. Lett. 81 1584 (1998)

Total EOS



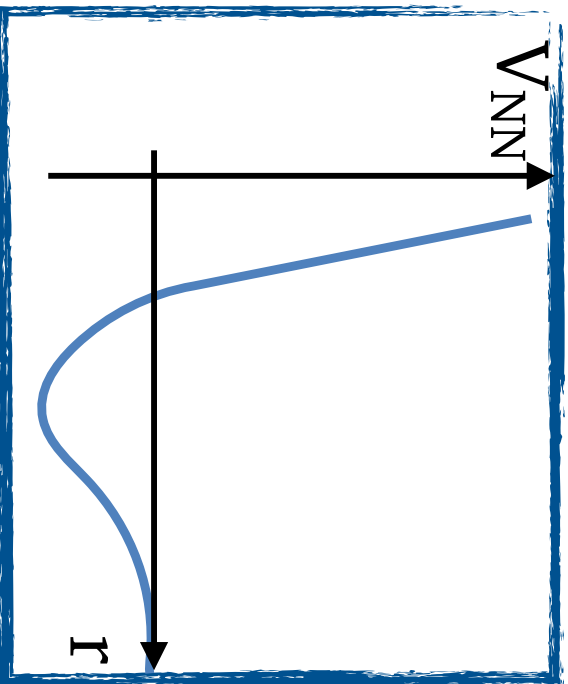
3-hole line



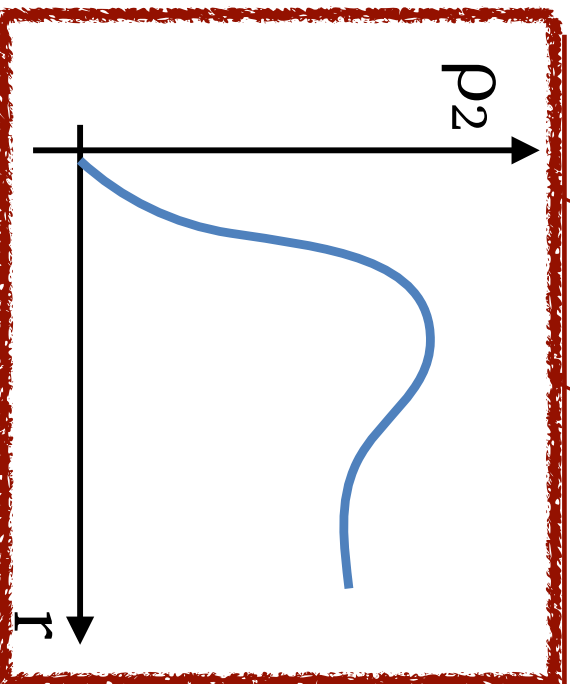


Nuclear predictions 19xx style

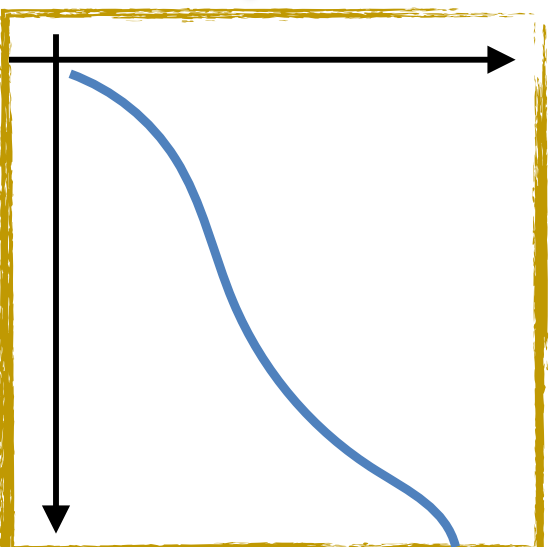
Hamiltonian



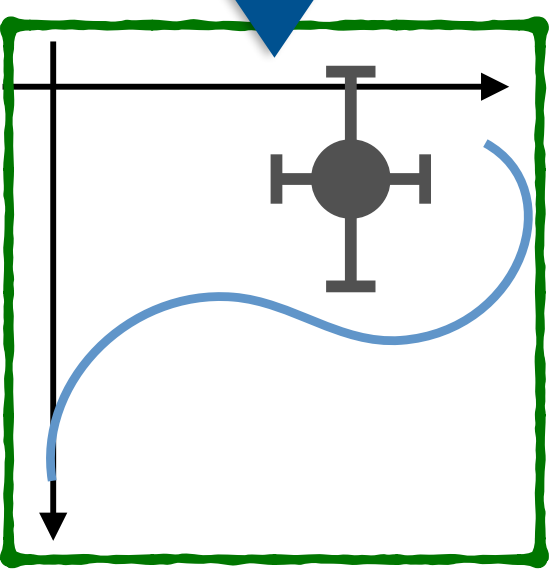
Many-body method



Astronuclear
property



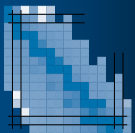
Neutron star
observations



$\Psi +$

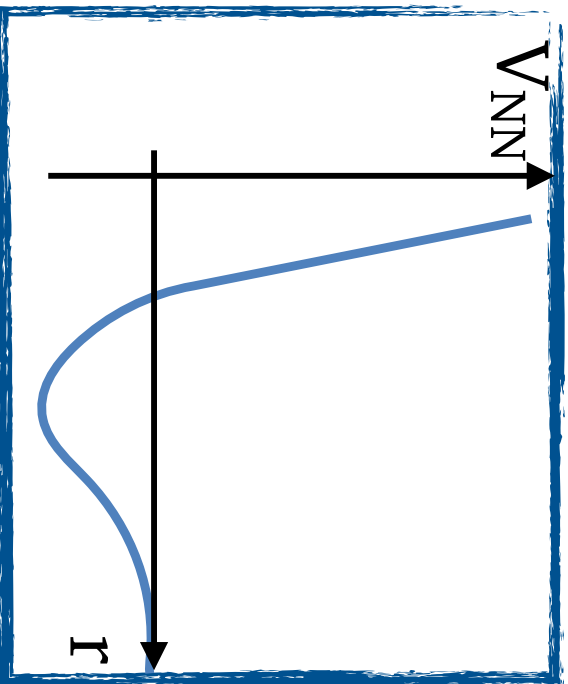


Astro
"black box"

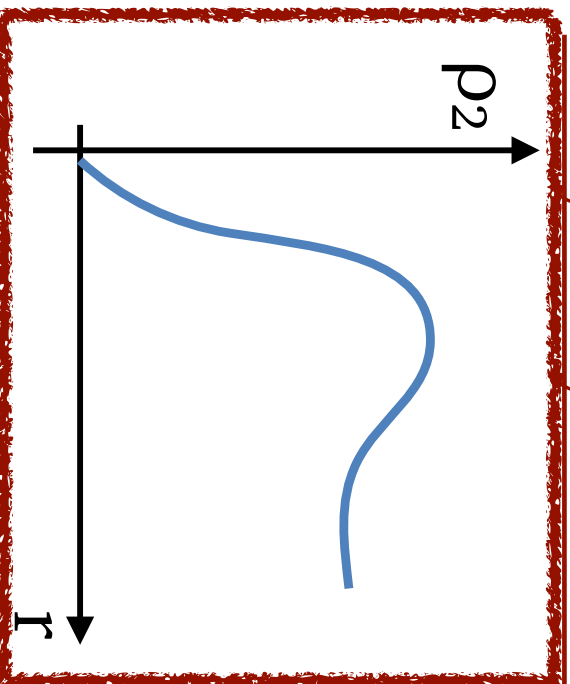


Nuclear predictions 19xx style

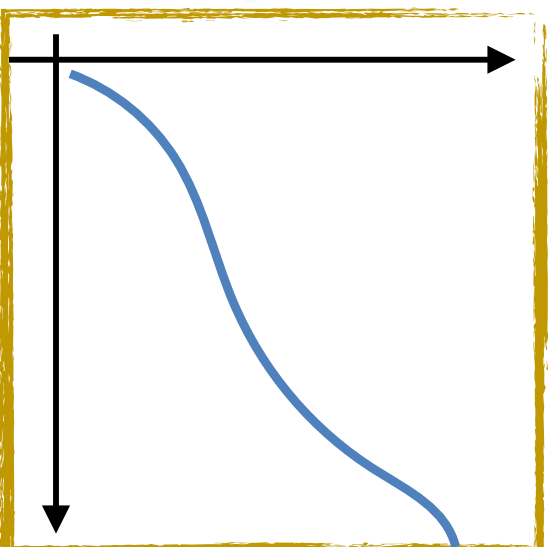
Hamiltonian



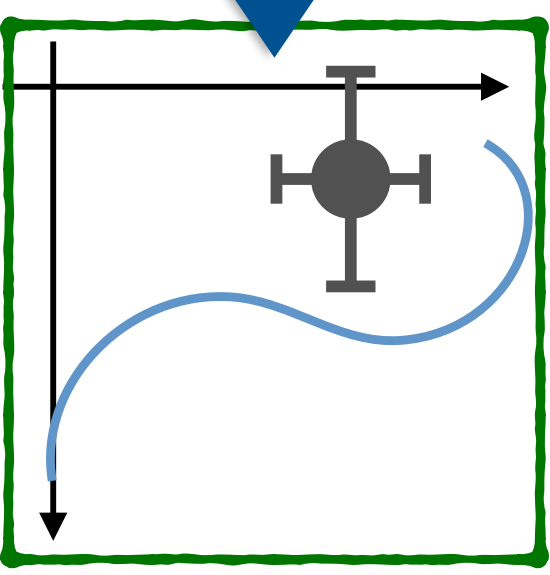
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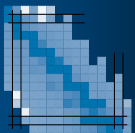


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observations



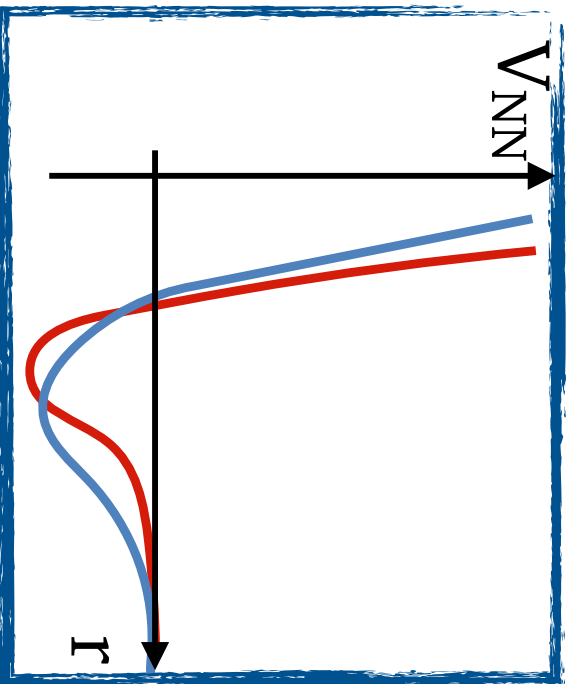
ψ +

Astro
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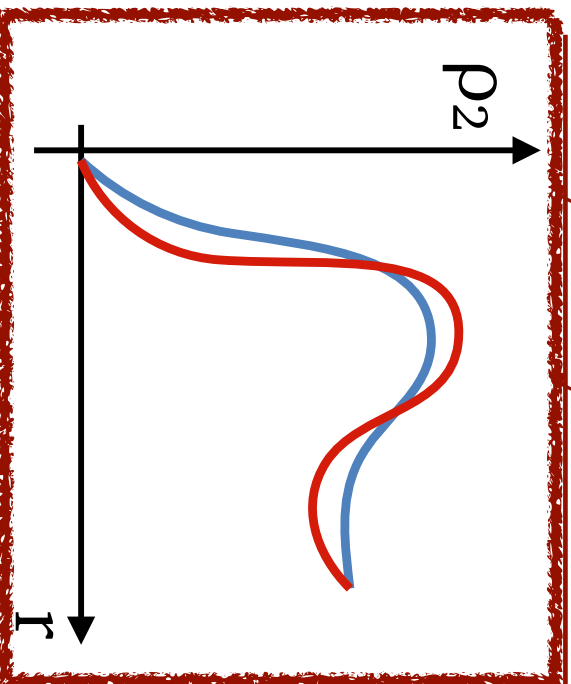


Nuclear predictions 19xx style

Hamiltonian

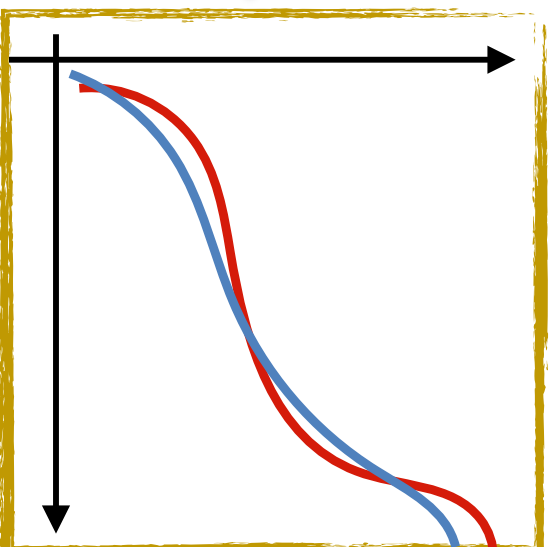


Many-body method



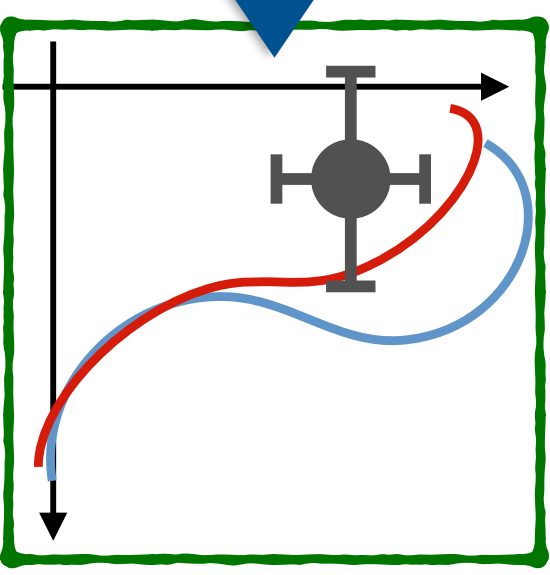
Astronuclear

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Neutron star

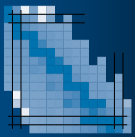
observations



$\Psi +$

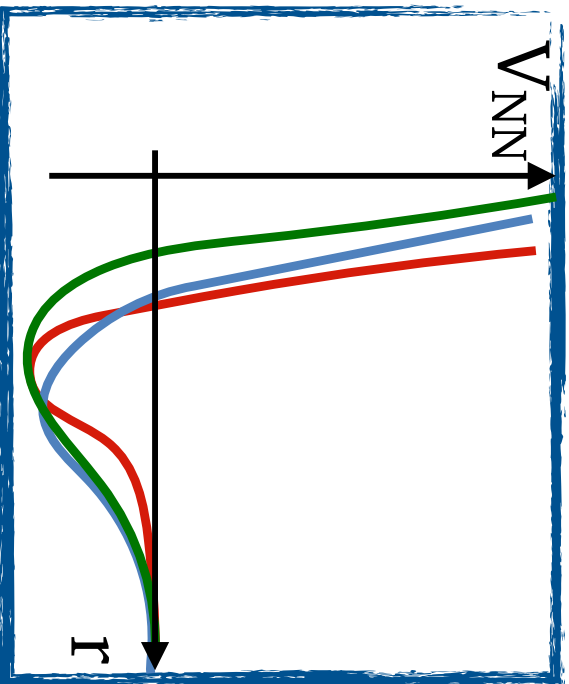


Astro
“black box”

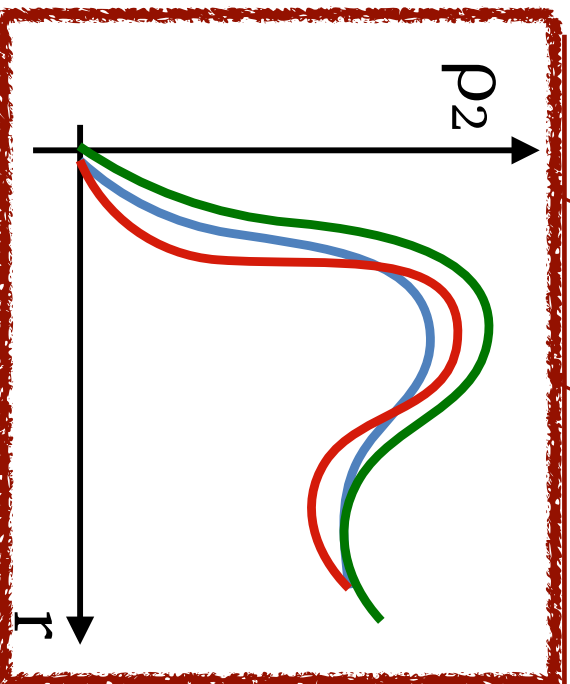


Nuclear predictions 19xx style

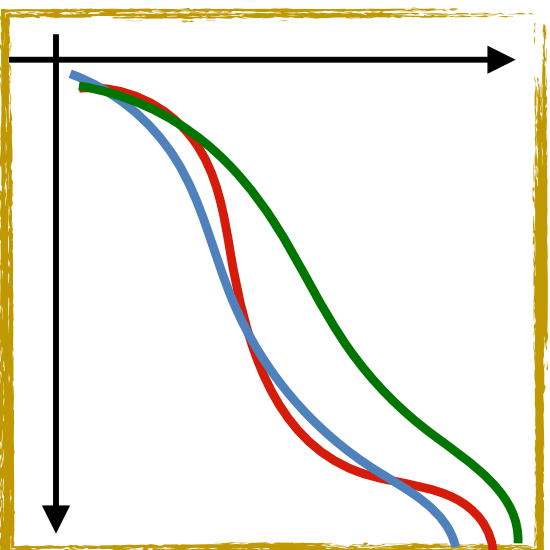
Hamiltonian



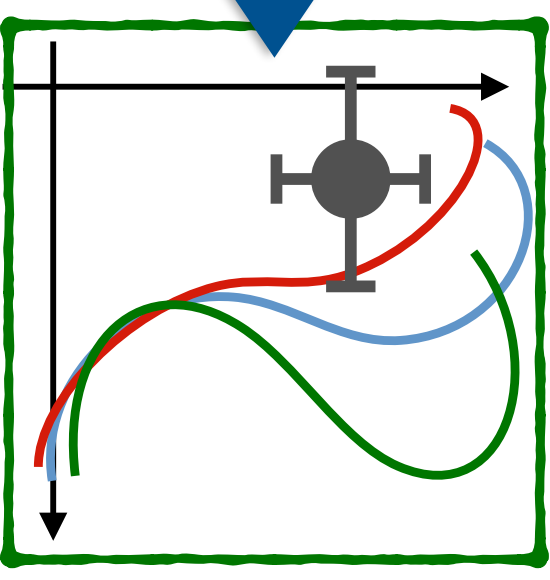
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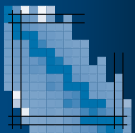


Neutron star
observations



$\Psi +$

Astro
"black box"



Our common paper

PHYSICAL REVIEW C **86**, 064001 (2012)

Comparative study of neutron and nuclear matter with simplified Argonne nucleon-nucleon potentials

M. Baldo,¹ A. Polls,² A. Rios,³ H.-J. Schulze,¹ and I. Vidaña⁴

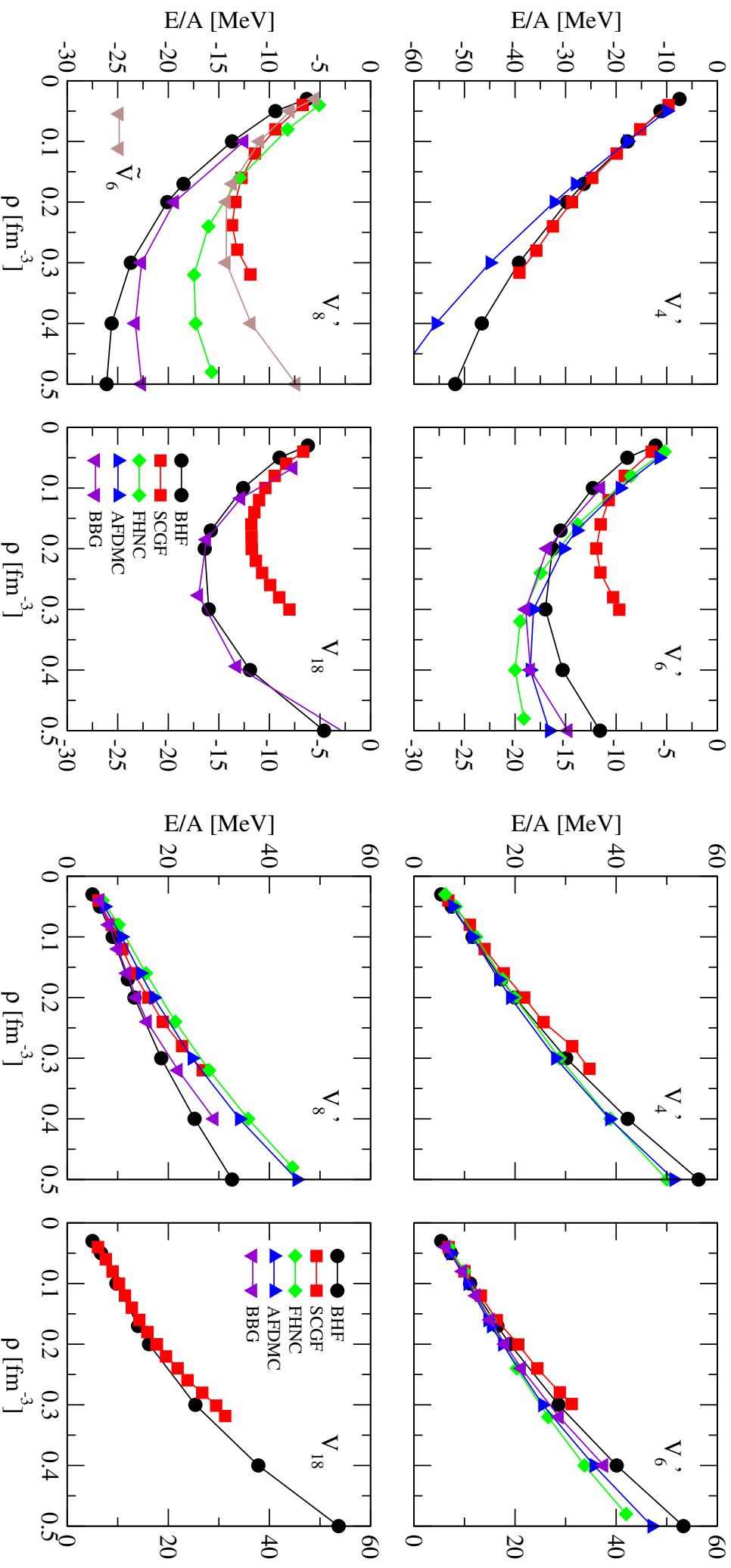
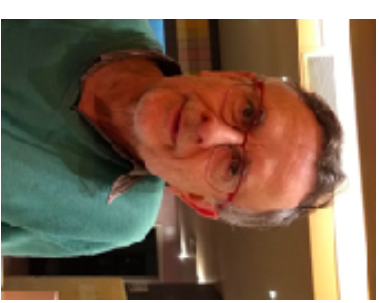
¹INEN Sezione di Catania, Dipartimento di Fisica, Via Santa Sofia 64, I-95123 Catania, Italy

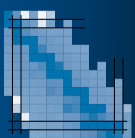
²Departament d'Estructura i Constituents de la Matèria, Universitat de Barcelona, E-08028 Barcelona, Spain

³Department of Physics, Faculty of Engineering and Physical Sciences, University of Surrey, Guildford, Surrey GU2 7XH, United Kingdom

⁴Centro de Física Computacional, Department of Physics, University of Coimbra, PT-3004-516 Coimbra, Portugal

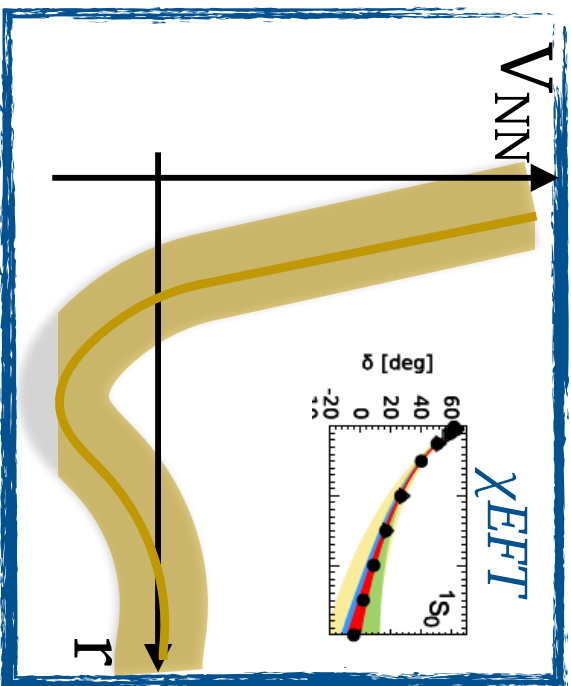
(Received 26 July 2012; revised manuscript received 26 October 2012; published 3 December 2012)



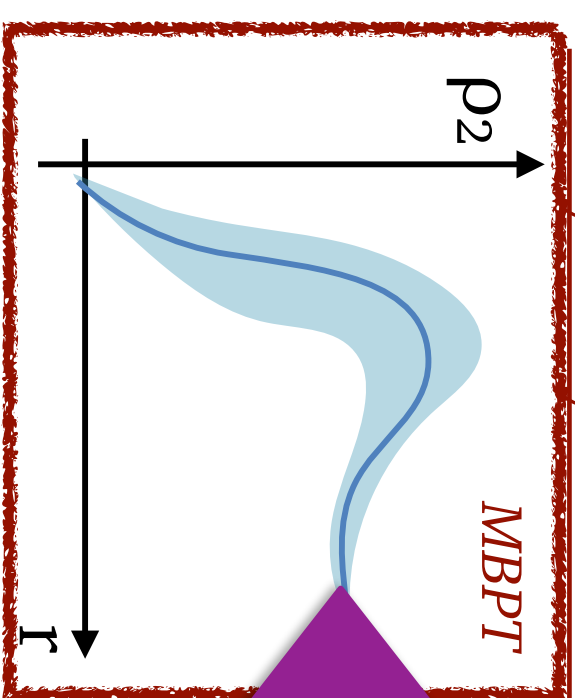


Nuclear error quantification

Hamiltonian

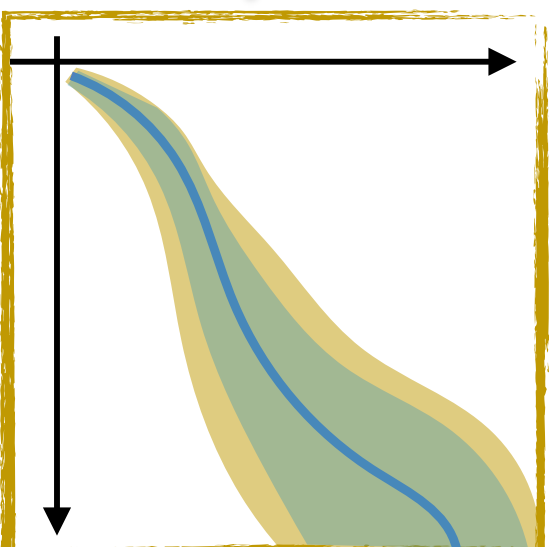


Many-body method



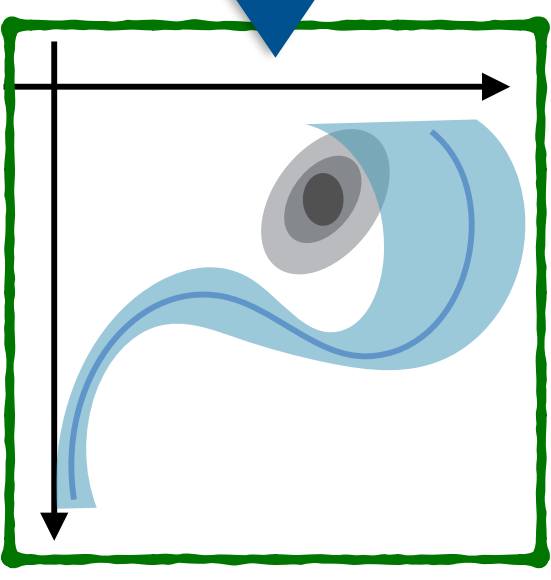
Backwards

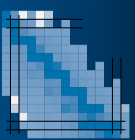
Astronuclear
property



Backwards

Neutron star
observations





History matching

nature
physics

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<https://doi.org/10.1038/s41567-022-01715-8>



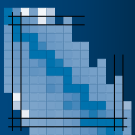
OPEN

Ab initio predictions link the neutron skin of ^{208}Pb to nuclear forces

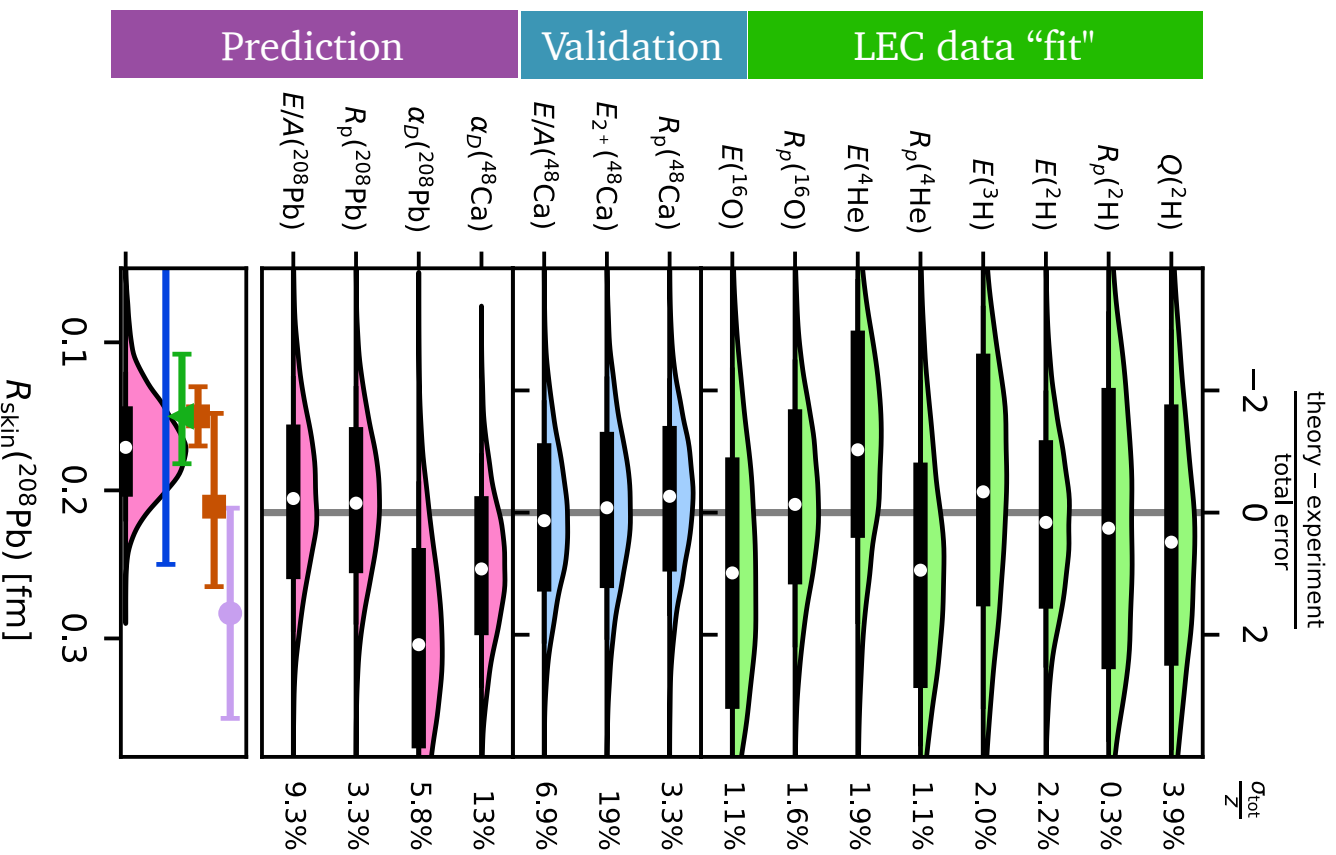
Baishan Hu^{1†}, Weiguang Jiang^{2†}, Takayuki Miyagi^{1,3,4†}, Zhonghao Sun^{5,6†}, Andreas Ekström², Christian Forssén²✉, Gaute Hagen^{1,5,6}, Jason D. Holt^{1,7}, Thomas Papenbrock^{1,5,6}, S. Ragnar Stroberg^{8,9} and Ian Vernon¹⁰

- 17 IECs from $\chi^2_{\text{EFF}}\text{T}$
- Start from 10^9 realisations
- History matching reduces to 34

Hu, Jiang, Miyagi, Holt, *Nature Physics* **18** 1196 (2022)



History matching



nature
physics

ARTICLES

<https://doi.org/10.1038/s41567-022-01715-8>

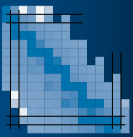


OPEN Ab initio predictions link the neutron skin of ^{208}Pb to nuclear forces

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- 17 LECs from $\chi^2_{\text{EFF}}\text{T}$
- Start from 10^9 realisations
- **History matching** reduces to 34
- Validate data in ^{48}Ca (importance weights)
- Predict ^{208}Pb & nuclear matter

Hu, Jiang, Miyagi, Holt, *Nature Physics* 18 1196 (2022)

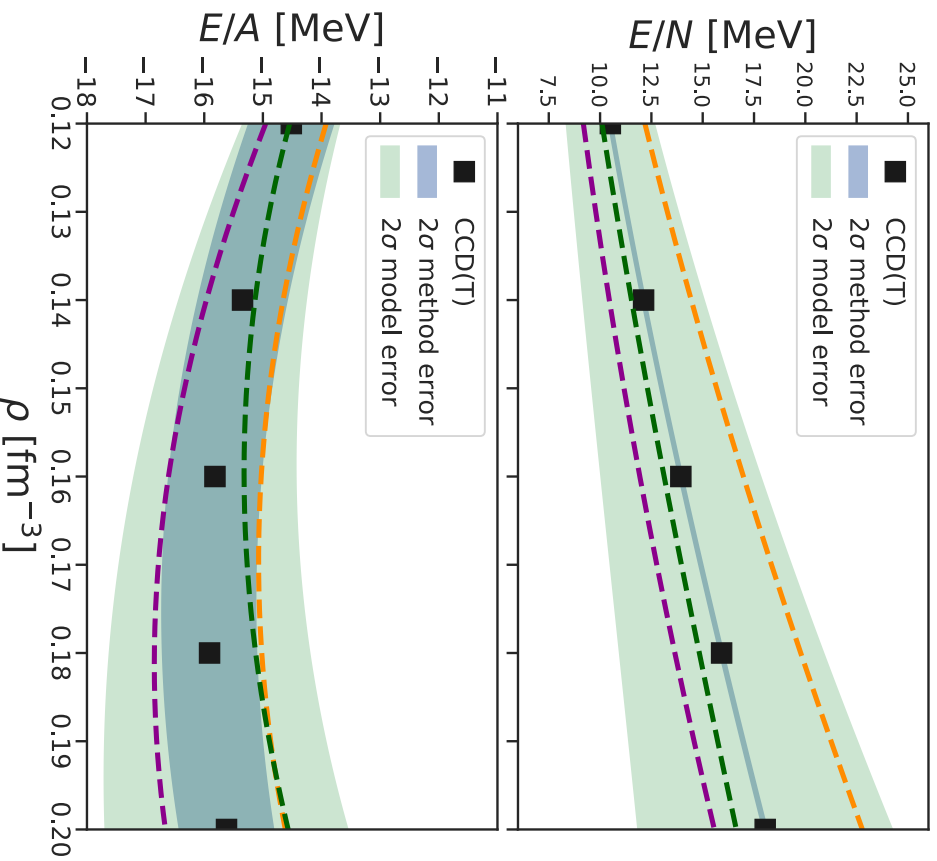


OPEN

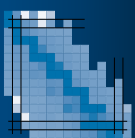
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Hu, Jiang, Miyagi, Holt, *Nature Physics* 18 1196 (2022)



History matching

nature
physics

<https://doi.org/10.1038/s41567-022-01715-8>

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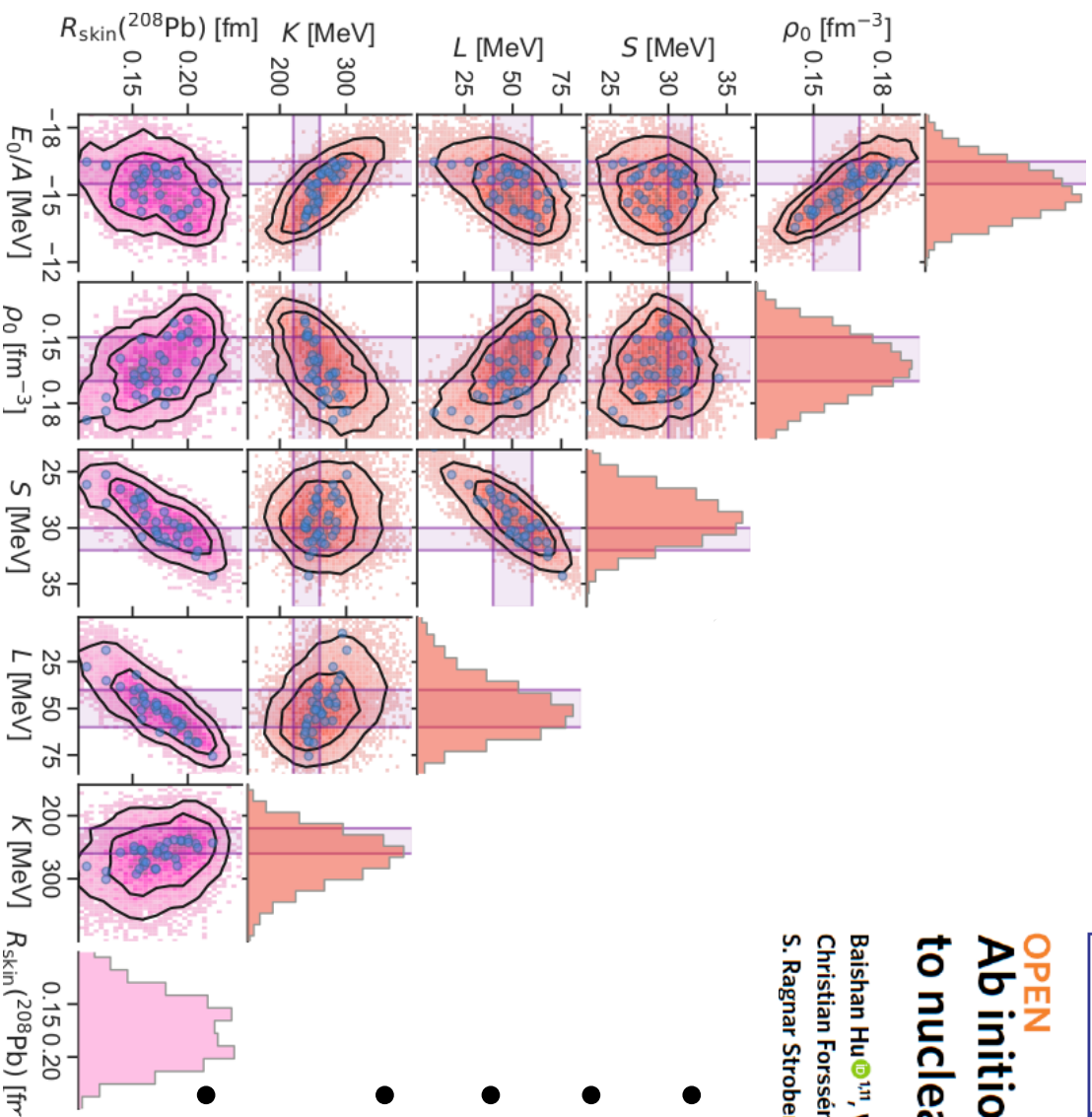


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Ab initio predictions link the neutron skin of ^{208}Pb to nuclear forces

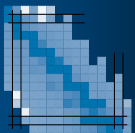
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E_0/A [MeV] ρ_0 [fm⁻³] S [MeV] L [MeV] K [MeV] $R_{\text{skin}}(^{208}\text{Pb})$ [fm]

Hu, Jiang, Miyagi, Holt, *Nature Physics* 18 1196 (2022)



An idea that lasted!

PHYSICAL REVIEW C **101**, 045801 (2020)

Editors' Suggestion

Benchmark calculations of pure neutron matter with realistic nucleon-nucleon interactions

M. Piarulli,¹ I. Bombaci,^{2,3} D. Logoteta,^{2,3} A. Lovato,^{4,5} and R. B. Wiringa⁴

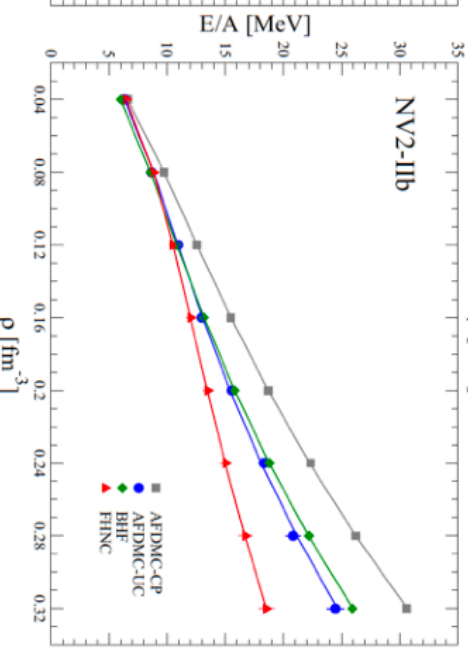
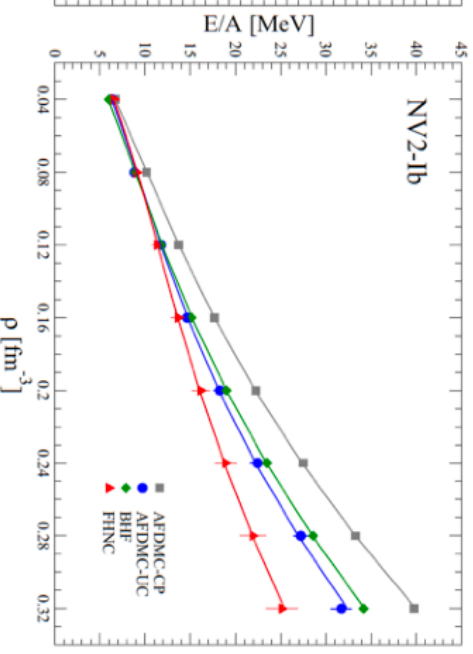
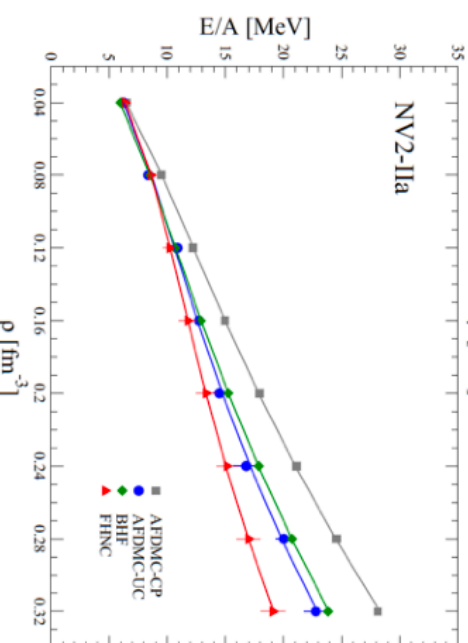
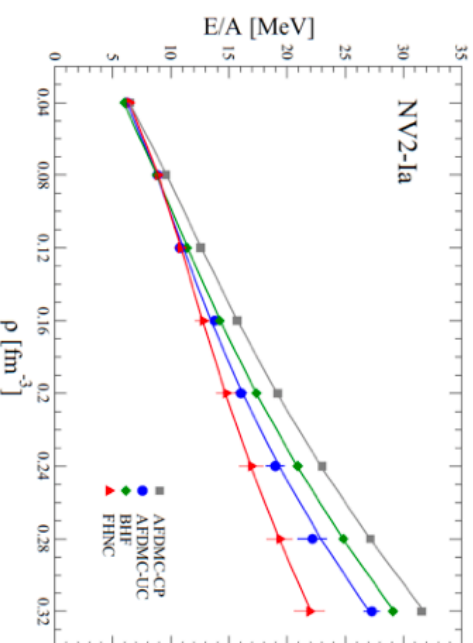
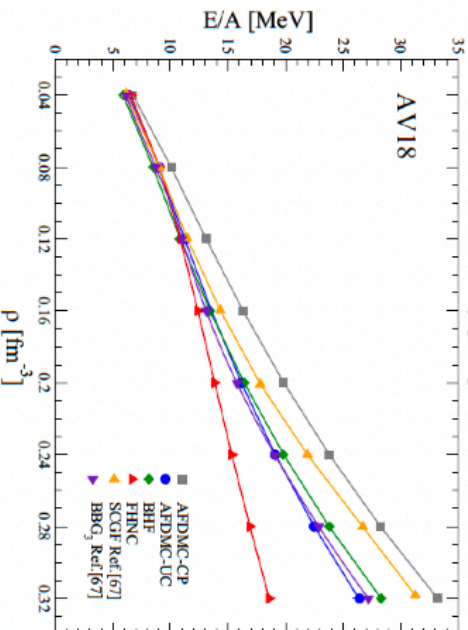
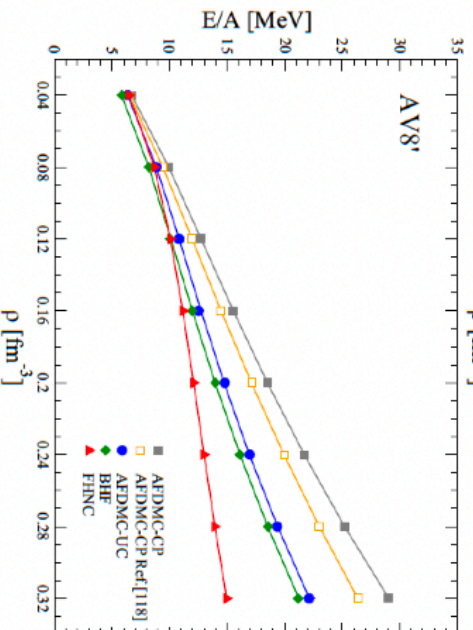
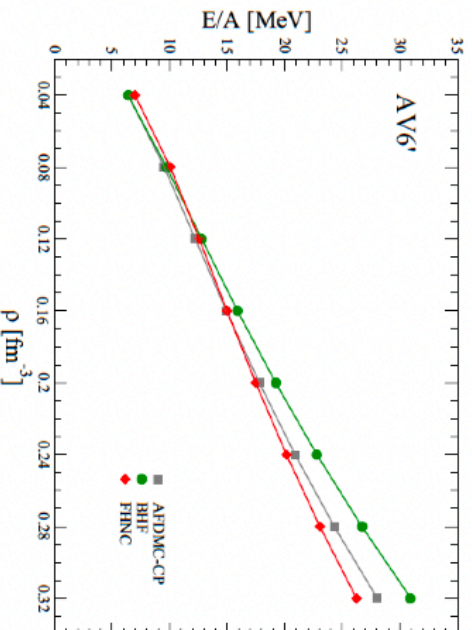
¹Physics Department, Washington University, St Louis, Missouri 63130, USA

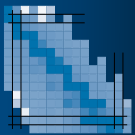
²Dipartimento di Fisica "E. Fermi", Università di Pisa, Largo B. Pontecorvo 3, I-56127 Pisa, Italy

³INFN, Sezione di Pisa, Largo B. Pontecorvo 3, I-56127 Pisa, Italy

⁴Physics Division, Argonne National Laboratory, Argonne, Illinois 60439, USA

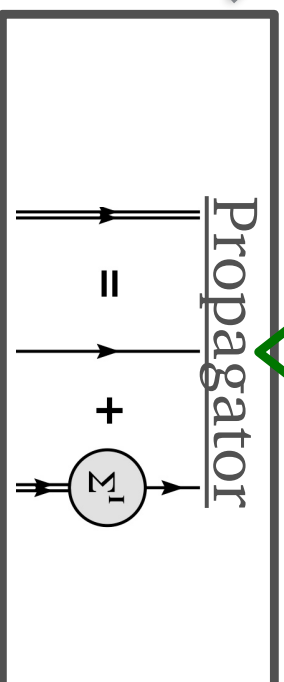
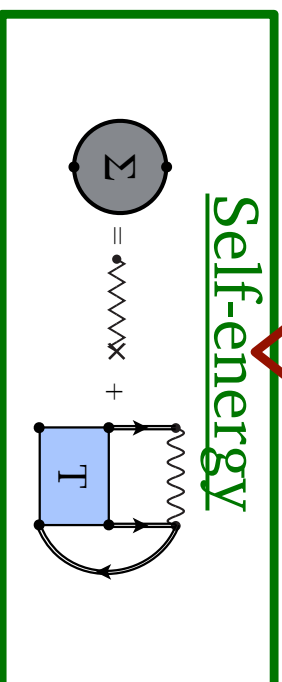
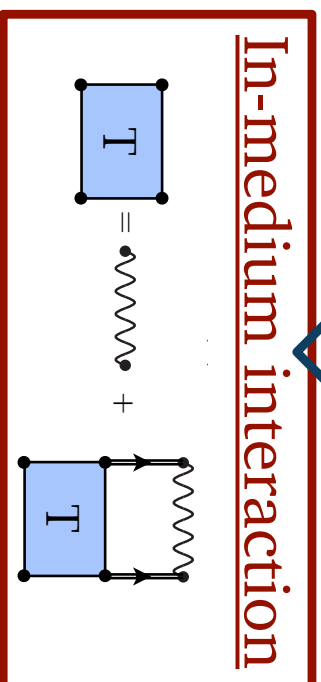
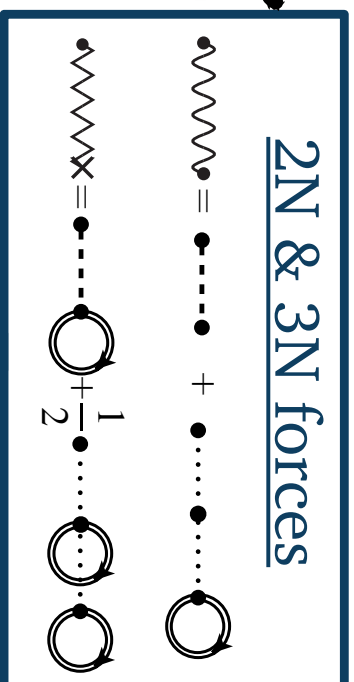
⁵INFN-TIPEA, Trento, Institute of Fundamental Physics and Applications, 38123, Trento, Italy



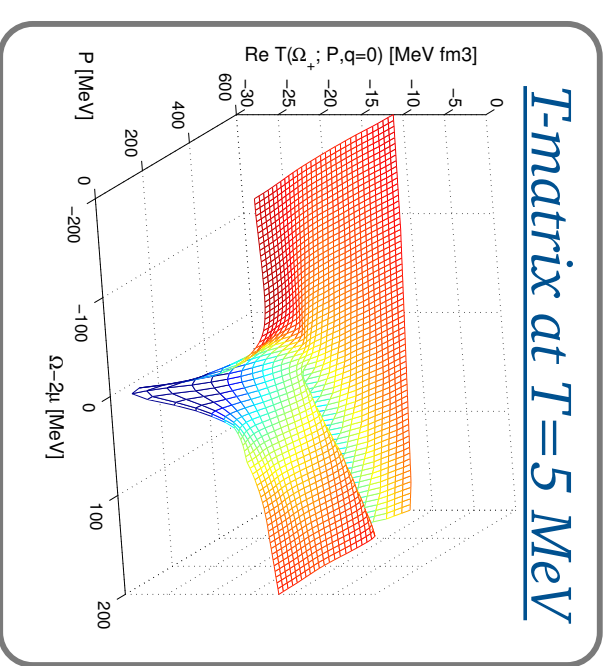


Self-Consistent Green's Functions

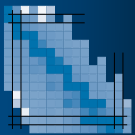
(ρ, T)



Carbone, Rios & Polls PRC 88 044302 (2013);
 PRC 90, 054322 (2014);
 Carbone PhD Thesis



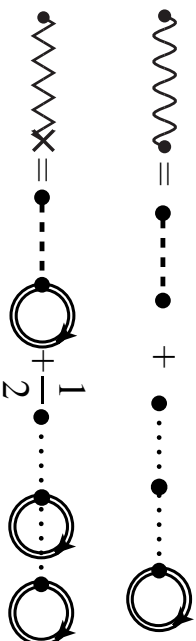
Ramos, Polls & Dickhoff, NPA 503 1 (1989)
 Alm *et al.*, PRC 53 2181 (1996)
 Dewulf *et al.*, PRL 90 152501 (2003)
 Frick & Muther, PRC 68 034310 (2003)
 Rios, PhD Thesis, U. Barcelona (2007)
 Soma & Bozek, PRC 78 054003 (2008)
 Rios & Soma PRL 108 012501 (2012)



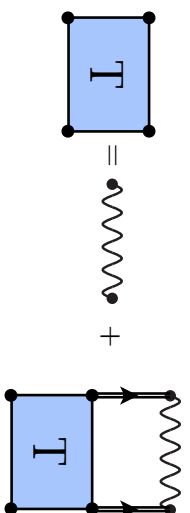
Self-Consistent Green's Functions

(ρ, T)

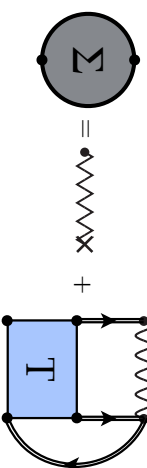
2N & 3N forces



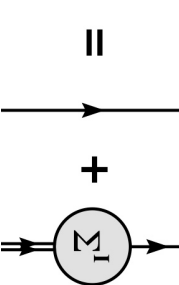
In-medium interaction



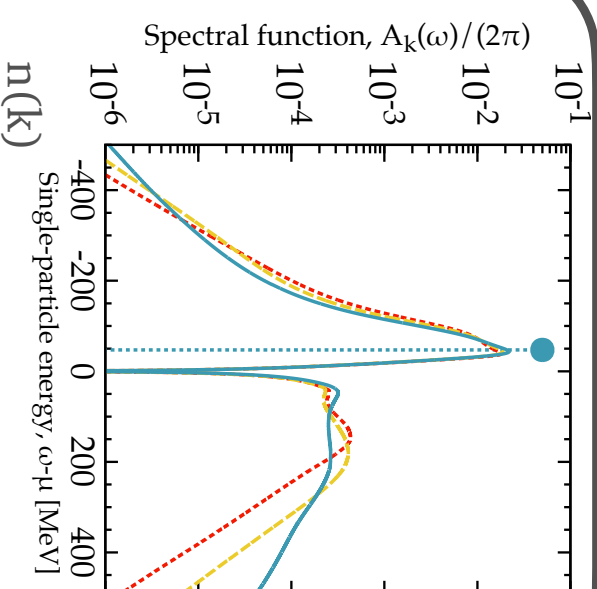
Self-energy



Propagator



Carbone, Rios & Polls PRC 88 044302 (2013);
 PRC 90, 054322 (2014);
 Carbone PhD Thesis



$n(k)$
 Thermodynamics & EoS
 Transport

Ramos, Polls & Dickhoff, NPA 503 1 (1989)

Alm *et al.*, PRC 53 2181 (1996)

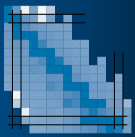
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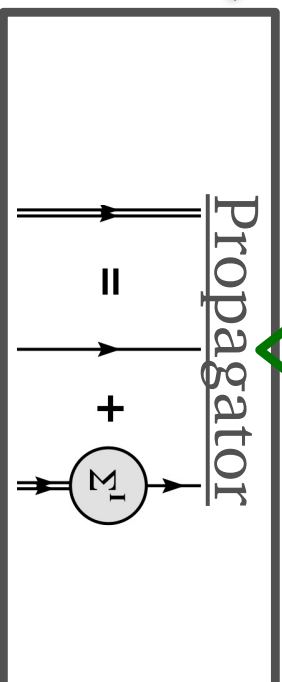
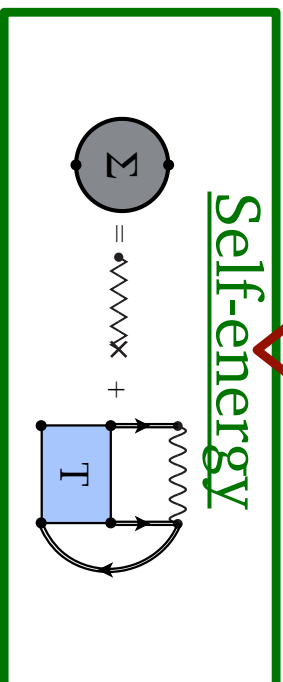
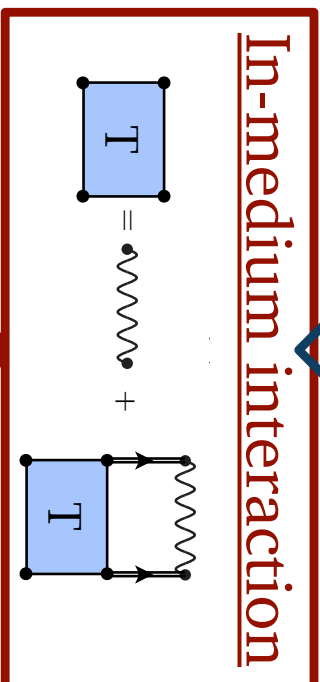
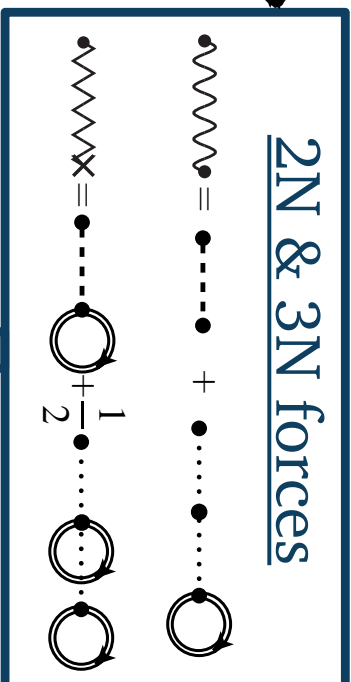
Soma & Bozek, PRC 78 054003 (2008)

Rios & Soma PRL 108 012501 (2012)

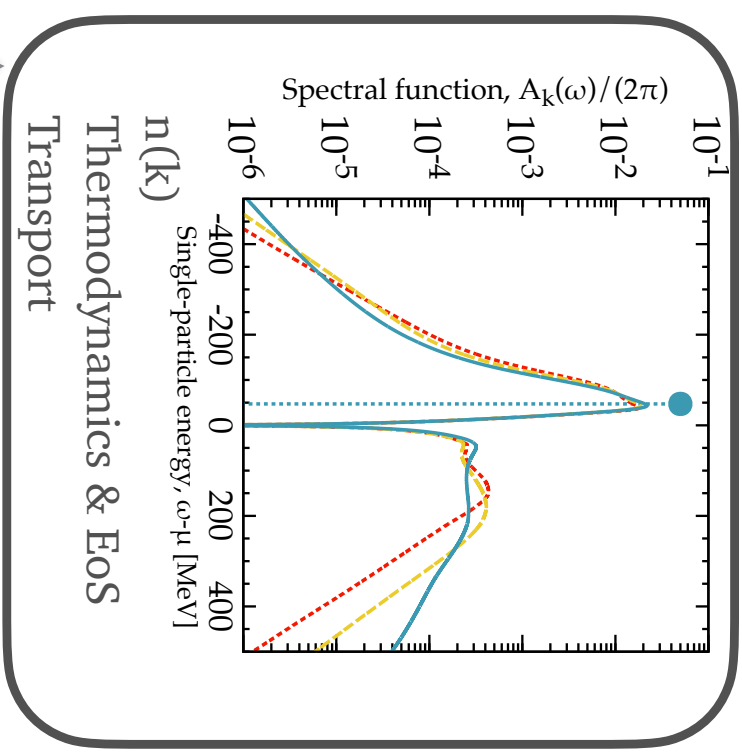


Self-Consistent Green's Functions

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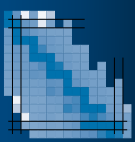


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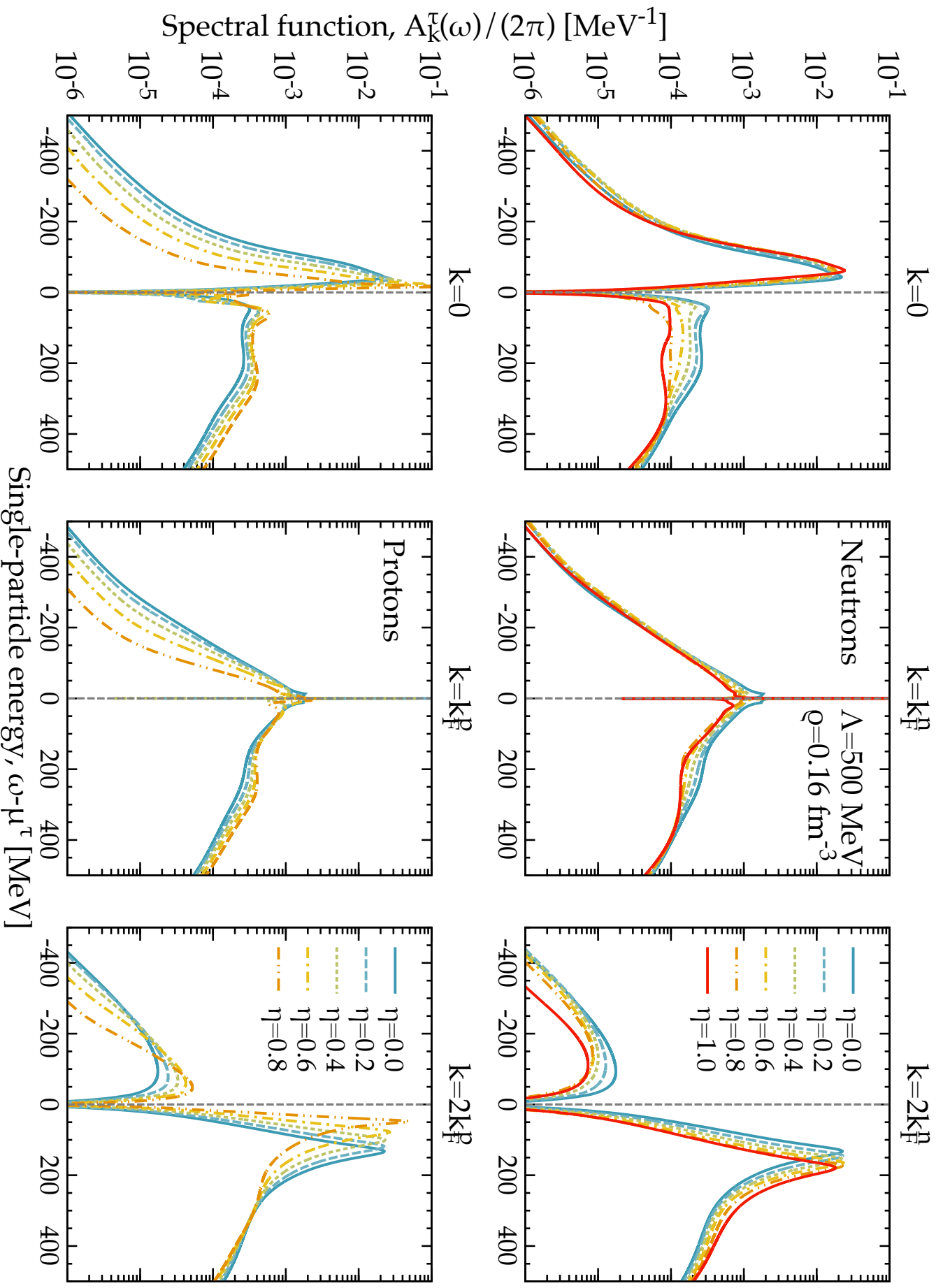


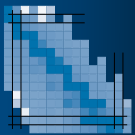
- Off-shell ✓
- Matsubara formalism ✓
- Φ -derivable ✓

Ramos, Polls & Dickhoff, NPA 503 1 (1989)
 Alm *et al.*, PRC 53 2181 (1996)
 Dewulf *et al.*, PRL 90 152501 (2003)
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 Rios, PhD Thesis, U. Barcelona (2007)
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Asymmetry dependence in spectral functions





Spectral functions à la Baldo

Off-the-energy-shell properties of the mass operator and spectral functions in nuclear matter

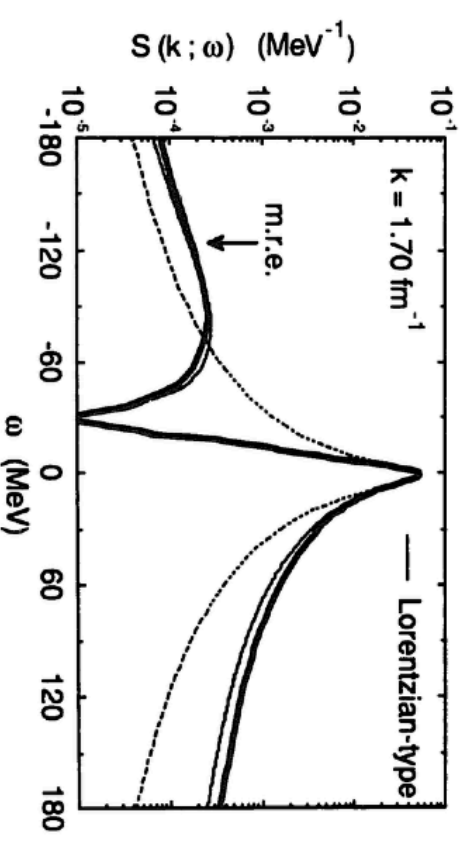
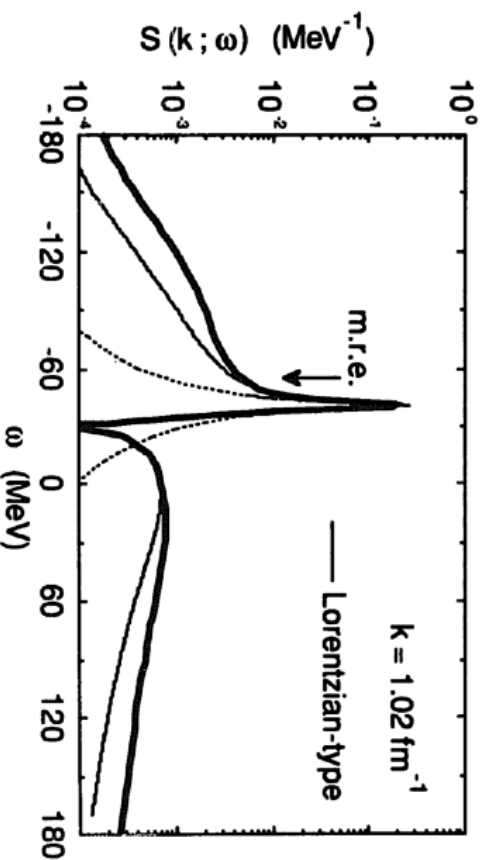
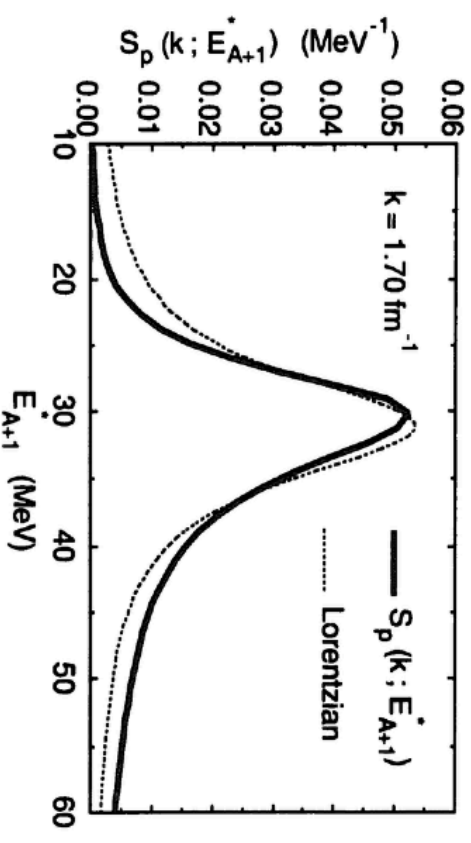
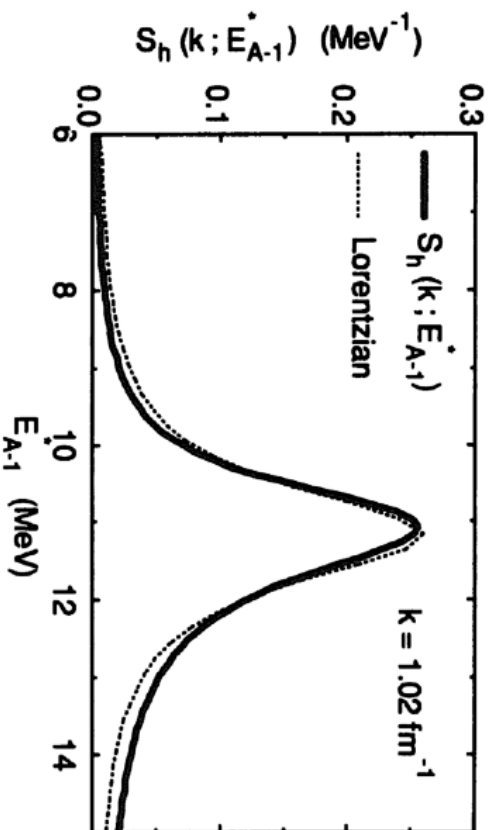
M. Baldo, I. Bombaci, G. Giansiracusa and U. Lombardo

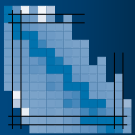
Dipartimento di Fisica, Università di Catania, and Istituto Nazionale di Fisica Nucleare, Sezione di Catania, Corso Italia 57, I-95129 Catania, Italy

C. Mahaux and R. Sartor

Institut de Physique B5, Université de Liège, B-4000, Liège 1, Belgium

Baldo *et al*, Nucl Phys A 545 741 (1992)





Finite temperature

Nuclear liquid-gas phase transition

M. Baldo¹ and L. S. Ferreira^{1,2}

¹Istituto Nazionale di Fisica Nucleare, Sezione di Catania and Dipartimento di Fisica, Università di Catania

Corso Italia 57, I-95129 Catania, Italy

²Centro de Física das Interações Fundamentais and Departamento de Física, Instituto Superior Técnico,

Av. Rovisco Pais, 1096 Lisboa, Portugal

(Received 8 June 1998)

Baldo & Ferreira, Phys Rev C 59 682 (1999)

Cylindrical diagrams!

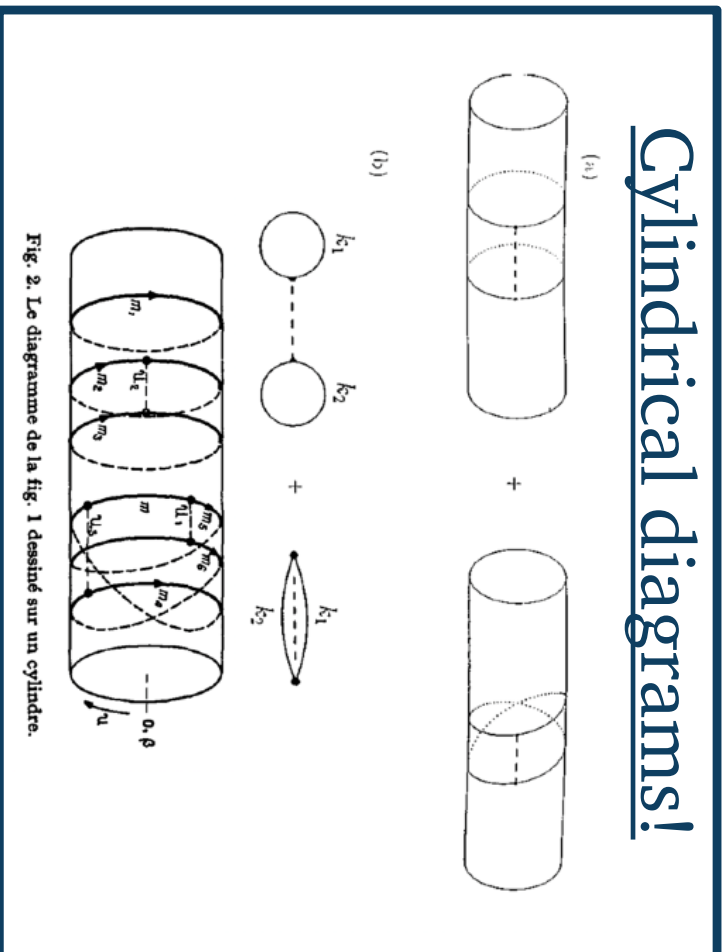


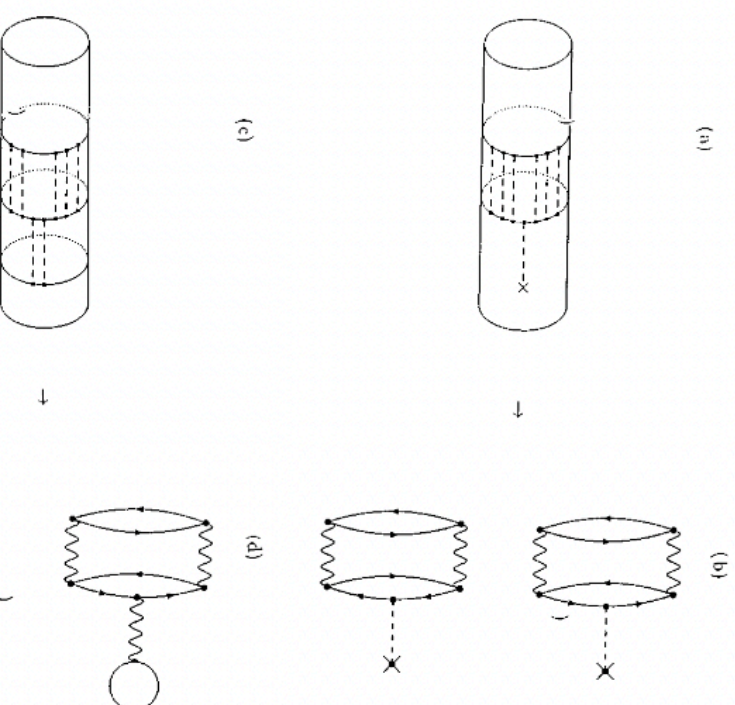
Fig. 2. Le diagramme de la fig. 1 dessiné sur un cylindre.

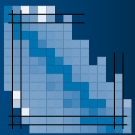
Bloch, Nucl. Phys. 7 451 (1958)

Bloch and De Dominicis, Nucl. Phys. 7 459 (1958)

Nucl. Phys. 10 181 (1959)

Nucl. Phys. 10 509 (1959)





Nuclear liquid-gas phase transition

M. Baldo¹ and L. S. Ferreira^{1,2}

¹*Istituto Nazionale di Fisica Nucleare, Sezione di Catania and Dipartimento di Fisica, Università di Catania*

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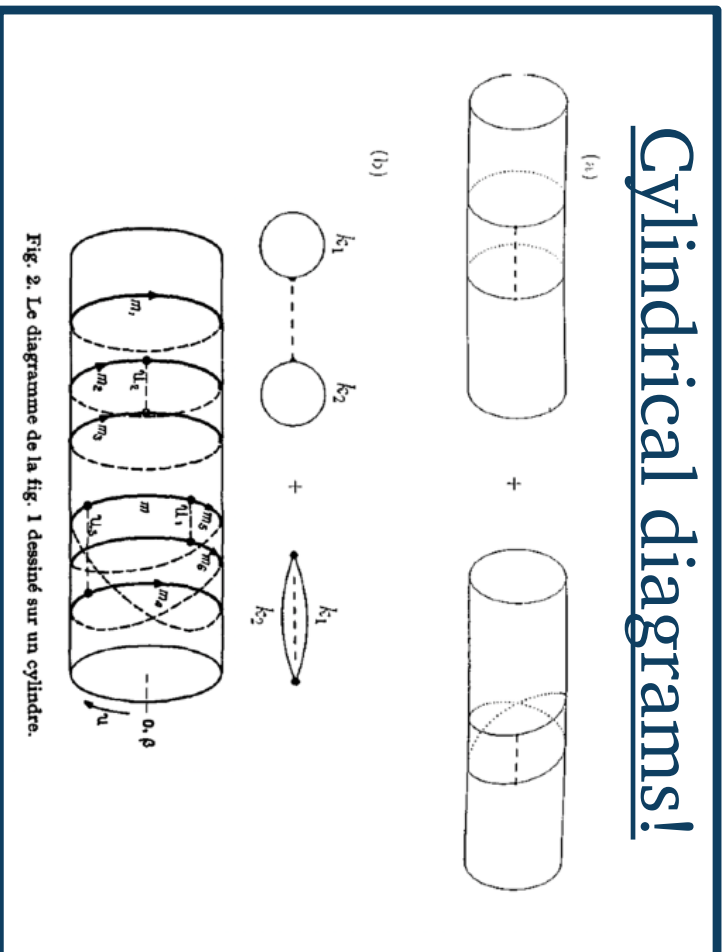


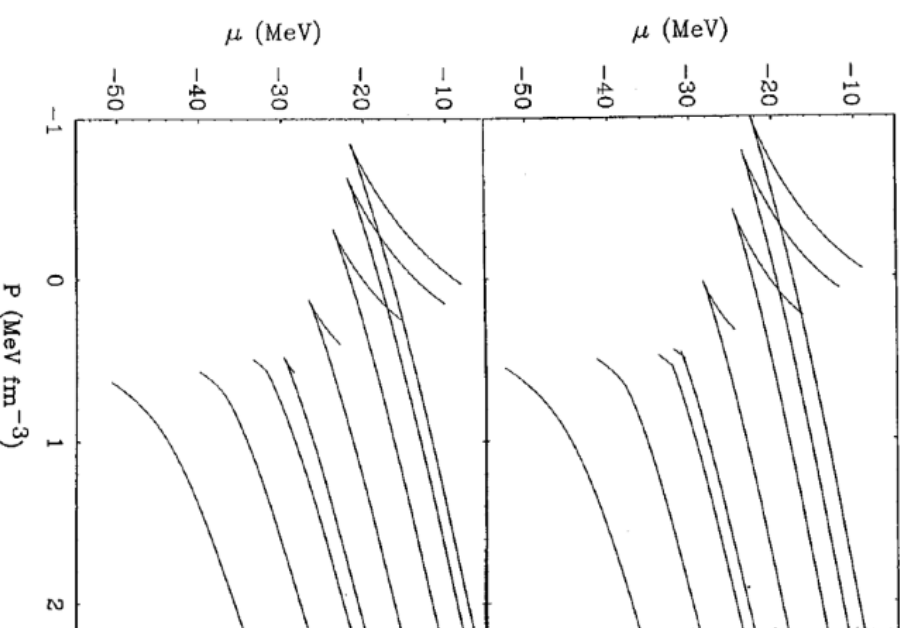
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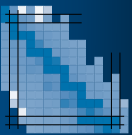
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Standing on the shoulders...

PHYSICAL REVIEW C 78, 044314 (2008)

Liquid-gas phase transition in nuclear matter from realistic many-body approaches

A. Rios*

National Superconducting Cyclotron Laboratory and Department of Physics and Astronomy, Michigan State University,
East Lansing, 48824-1321 Michigan, USA

A. Polls and A. Ramos

Departament d'Estructura i Constituents de la Matèria and Institut de Ciències del Cosmos, Universitat de Barcelona,
Avda. Diagonal 647, E-08028 Barcelona, Spain

H. Mülher

Institut für Theoretische Physik, Universität Tübingen, D-72076 Tübingen, Germany
(Received 15 May 2008; published 20 October 2008)

PHYSICAL REVIEW C 98, 025804 (2018)

Microscopic predictions of the nuclear matter liquid-gas phase transition

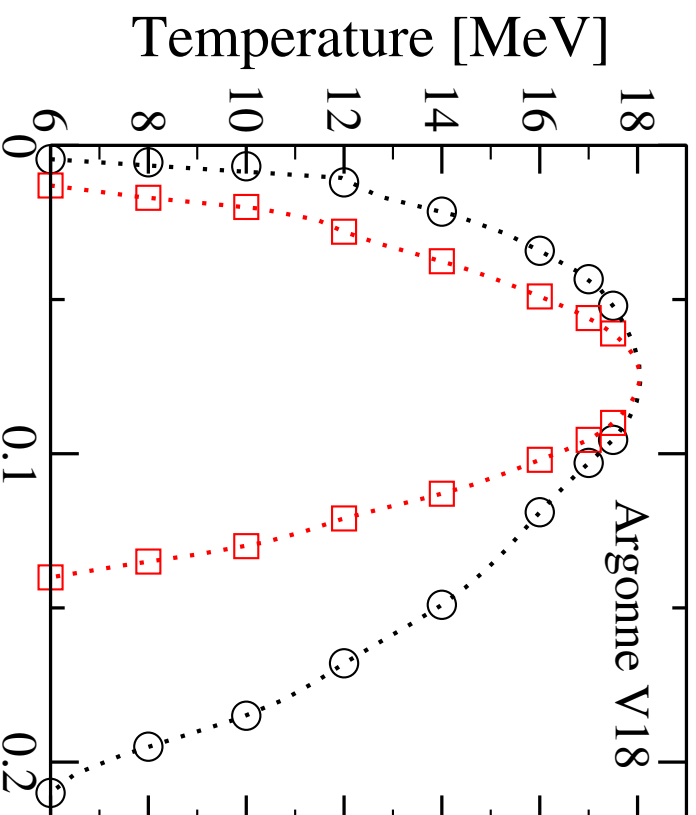
Arianna Carbone,^{1,*} Artur Polls,^{2,†} and Arnau Rios^{3,‡}

European Centre for Theoretical Studies in Nuclear Physics and Related Areas (ECT*) and Fondazione Bruno K
Sirtida delle Tabarelle 286, I-38123 Villazzano (TN), Italy

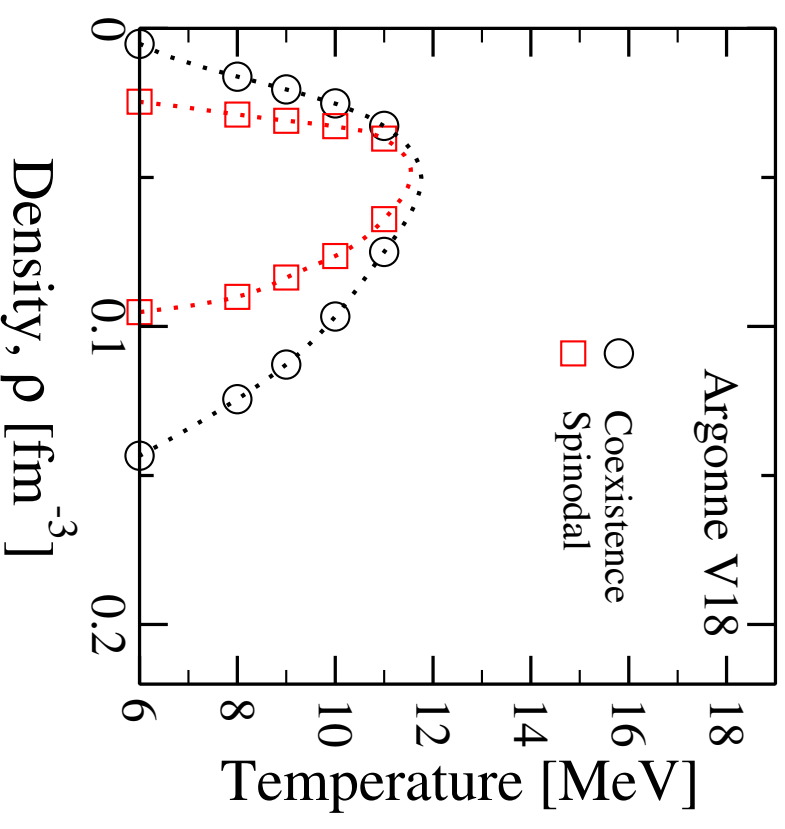
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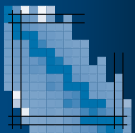
Physics, Faculty of Engineering and Physical Sciences, University of Surrey, Guildford, Surrey GU2 7XH,

BHF



SCGF





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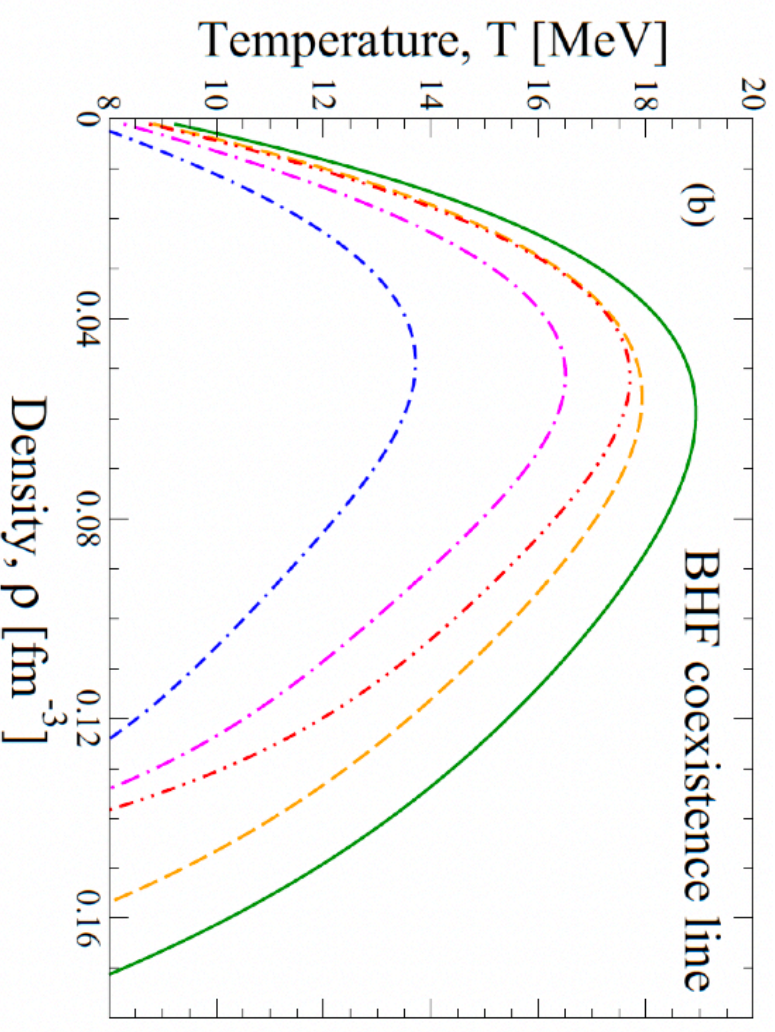
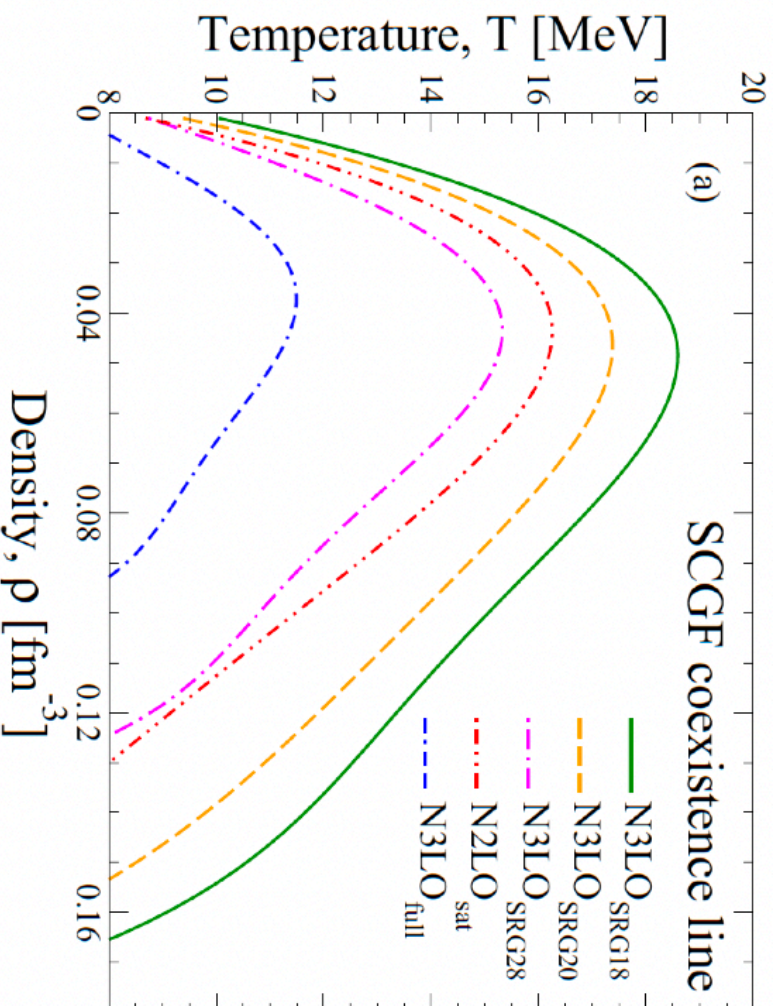
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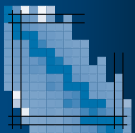
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Physics, Faculty of Engineering and Physical Sciences, University of Surrey, Guildford, Surrey GU2 7XH,



Carbone, Polls, Rios, Phys. Rev. C 98 025804 (2018)

[arXiv:2006.10610]



Hypernuclear matter in the Brueckner-Hartree-Fock approximation

H.-J. Schulze^a, A. Lejeune^a, J. Cugnon^a, M. Baldo^b, U. Lombardo^b

^a Université de Liège, Institut de Physique B5, B-4000 Liège 1, Belgium

^b Dipartimento di Fisica, Università di Catania, Corso Italia, 57, I-95129 Catania, Italy

Received 25 April 1995

Editor: G.F. Bertsch

PHYSICAL REVIEW C, VOLUME 61, 055801

Hyperon stars in the Brueckner-Bethe-Goldstone theory

M. Baldo and G. F. Bertsch

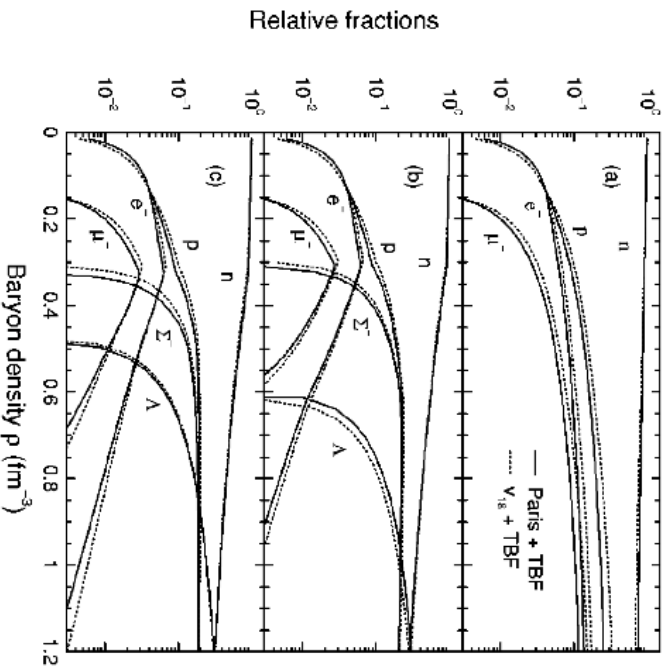
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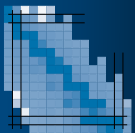
(Received 3 July 1997)

Description of hyperonic matter and hypernuclei within the Brueckner-Hartree-Fock theory

Isaac Vidaña Haro

Departament d'Estructura i Constituents de la Matèria

Universitat de Barcelona



Hyperons

R

Physica Letters B 355 (1995) 21–26

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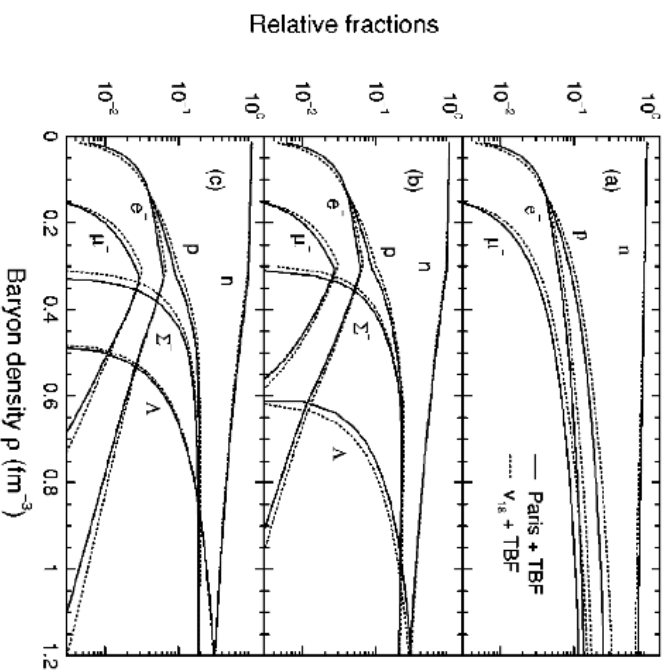
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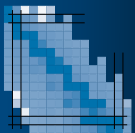
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⇒D. Blaschke's talk



Properties of the nuclear medium

M Baldo and G F Burgio

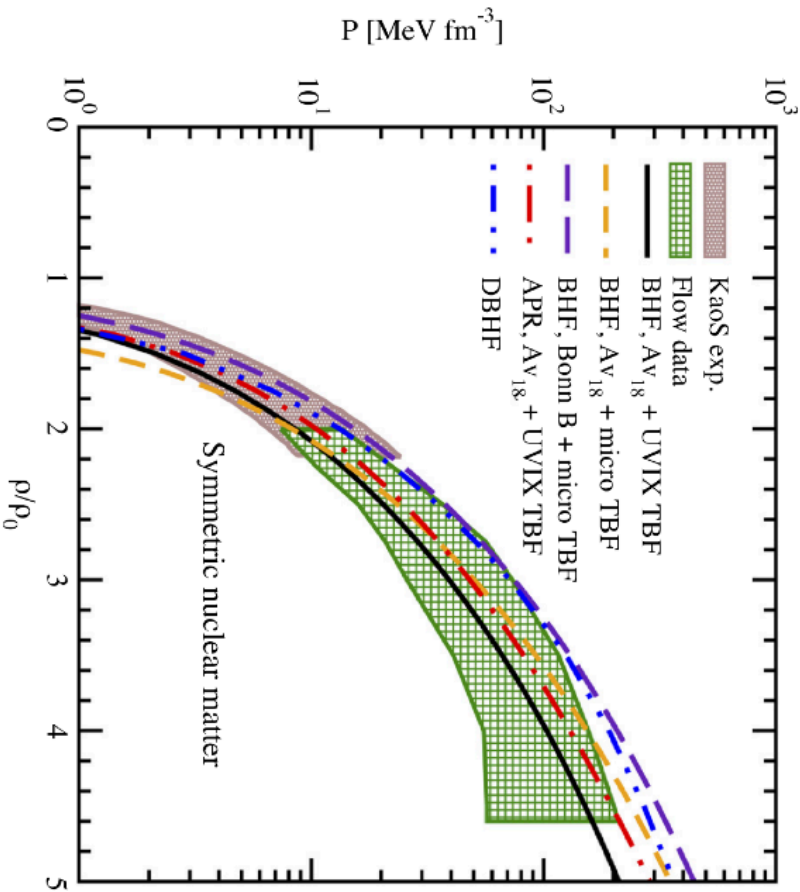
Istituto Nazionale di Fisica Nucleare, Sez. di Catania, Via S Sofia 64 95123 Catania, Italy

E-mail: marcello_baldo@ct.infn.it and giorella_burgio@ct.infn.it

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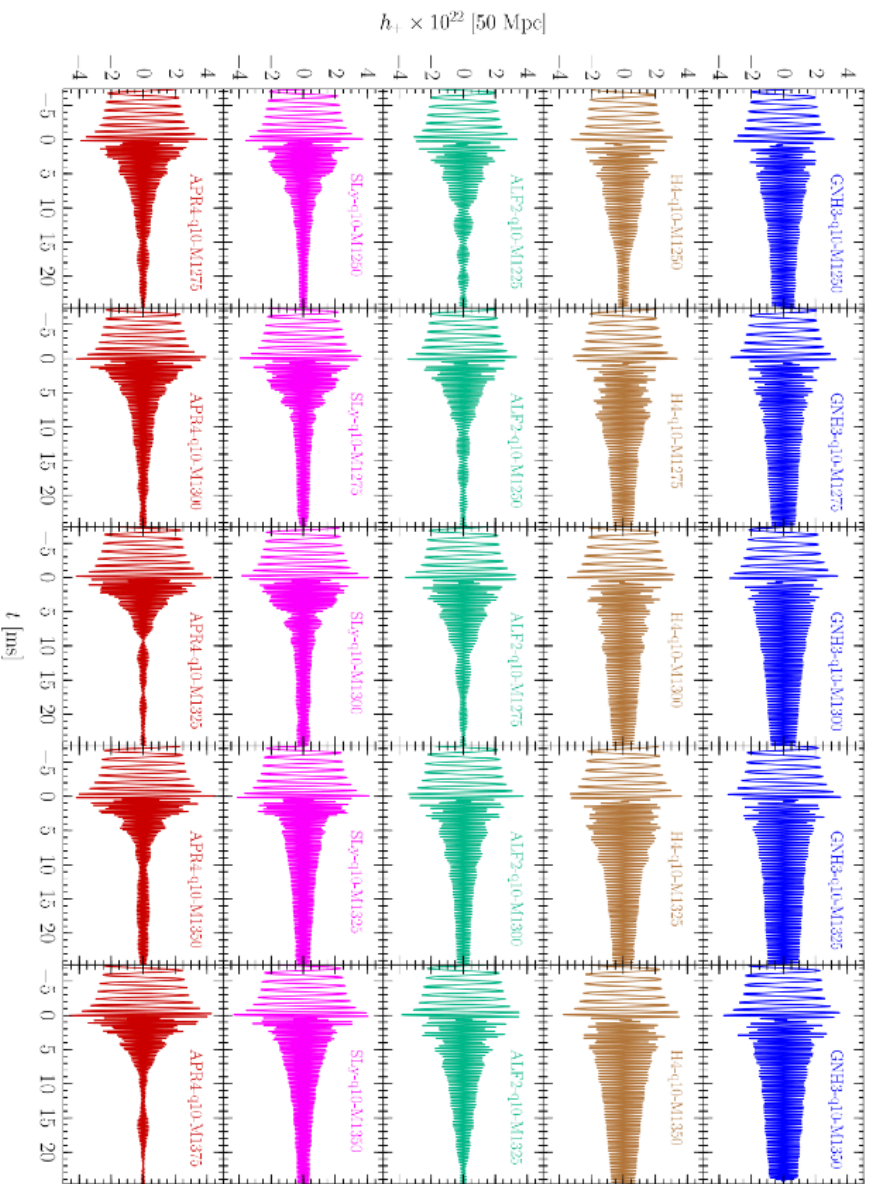


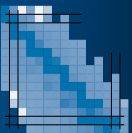
Review

The nuclear symmetry energy

M. Baldo, G.F. Burgio *

Istituto Nazionale di Fisica Nucleare, Sez. Catania, via S. Sofia 64, Catania, Italy





Beyond-BCS pairing: overview

SUPERFLUIDITY IN NEUTRON MATTER AND NUCLEAR MATTER WITH REALISTIC INTERACTIONS

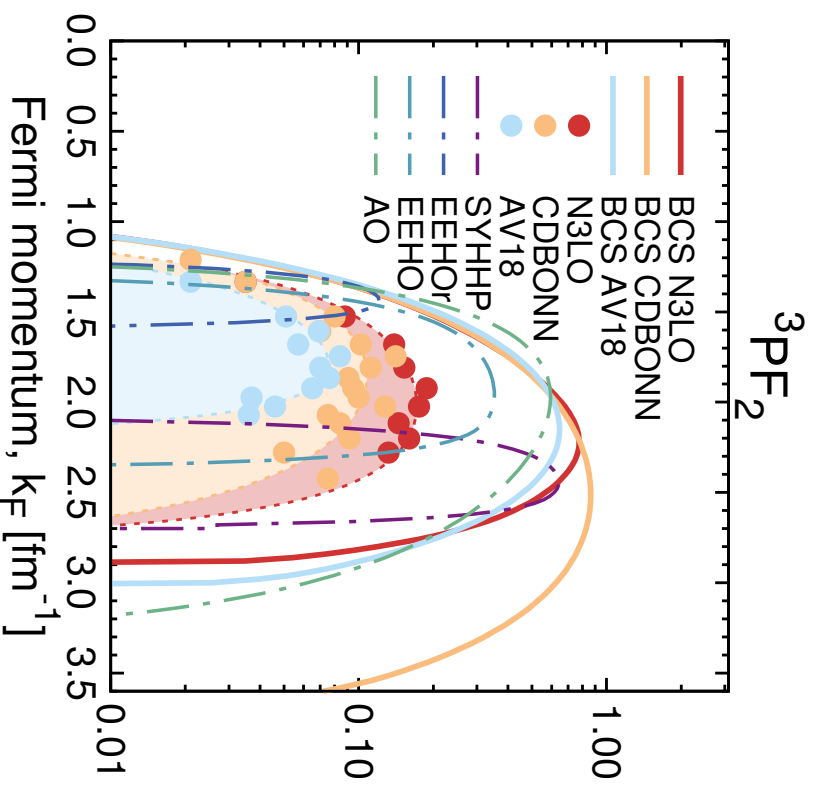
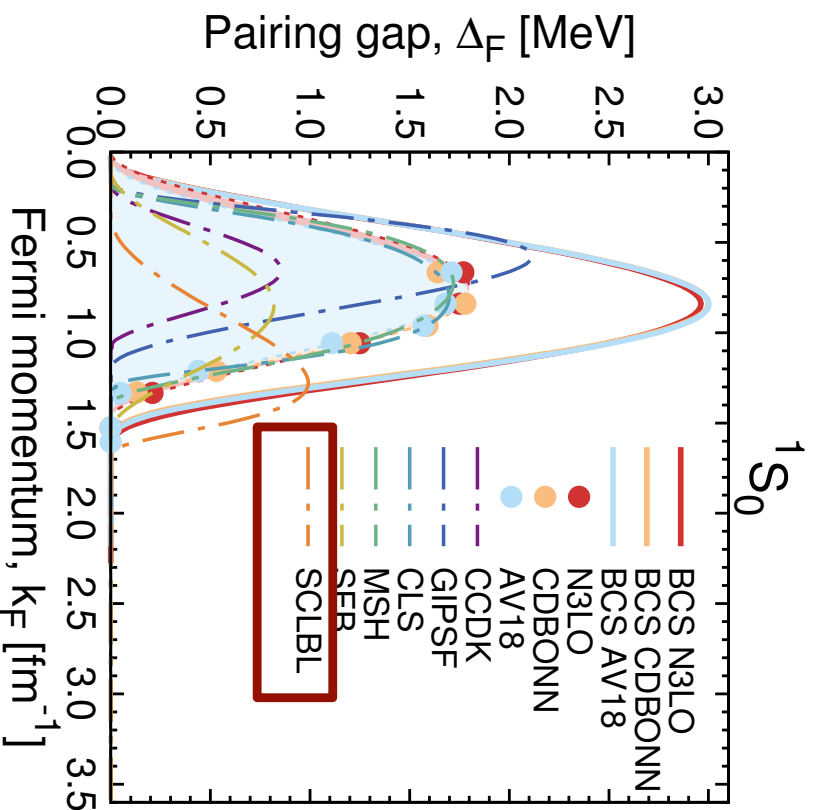
M. BALDO¹, J. CUGNON², A. LEJEUNE² and U. LOMBARDO¹

¹ Dipartimento di Fisica, Univ. di Catania and INFN sezione di Catania, Corso Italia 57, I-95129 Catania, Italy

² Université de Liège, Physique Nucléaire Théorique, Institut de Physique au Sart Tilman, Bâtiment B.5, B-4000 Liège 1, Belgium

Received 22 February 1990

(Revised 19 April 1990)



³P₂-³F₂ pairing in neutron matter with modern nucleon-nucleon potentials

M. Baldo,¹ Ø. Elgarøy,² L. Engvik,² M. Hjorth-Jensen,³ and H.-J. Schulze¹

¹Sezione INFN, Università di Catania, Corso Italia 57, I-95129 Catania, Italy

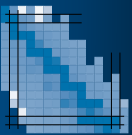
²Department of Physics, University of Oslo, N-0316 Oslo, Norway

³Nordita, Blegdamsvej 17, DK-2100 Copenhagen Ø, Denmark

(Received 29 June 1998)

Baldo et al, Phys. Rev. C 58 1921 (1996)

SCLBL=Schulze, Cugnon, Lejeune, Baldo & Lombardo, Phys. Lett. B 375 1 (1996)



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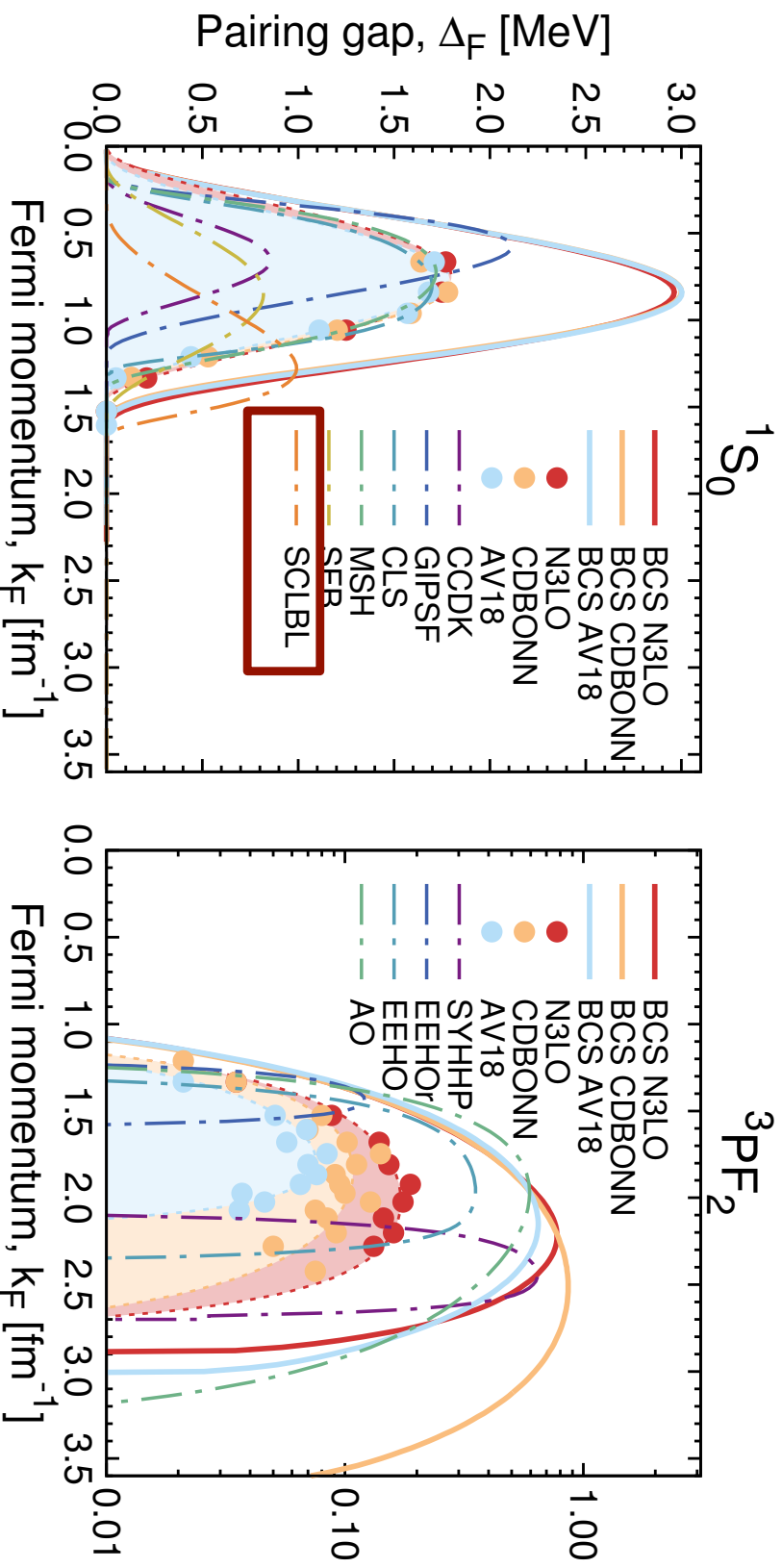
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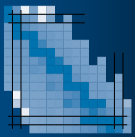
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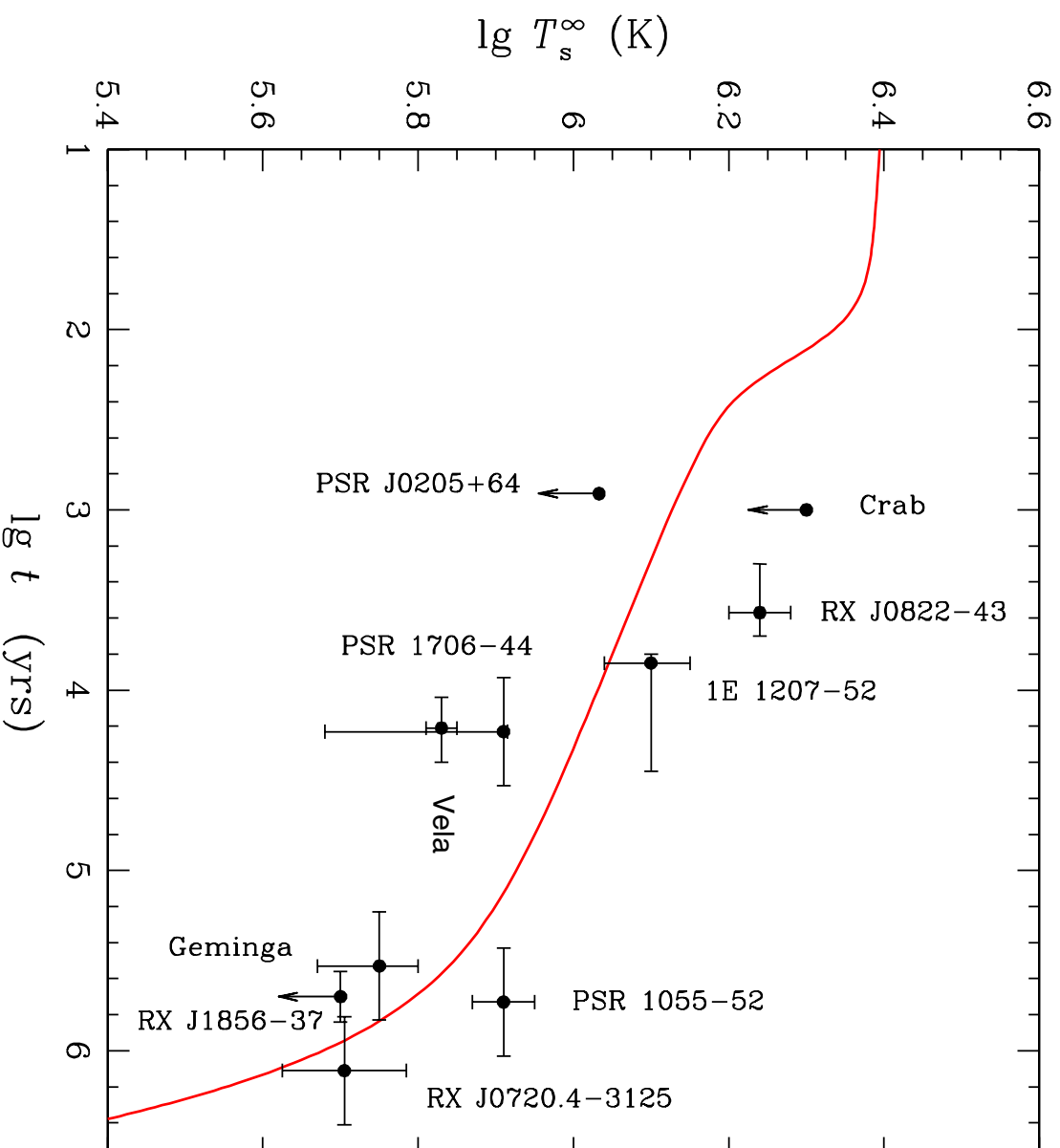


SCLBL=Schulze, Cugnon, Lejeune, Baldo & Lombardo, Phys. Lett. B 375 1 (1996)

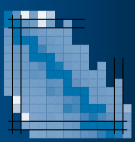
⇒U. Lombardo's talk



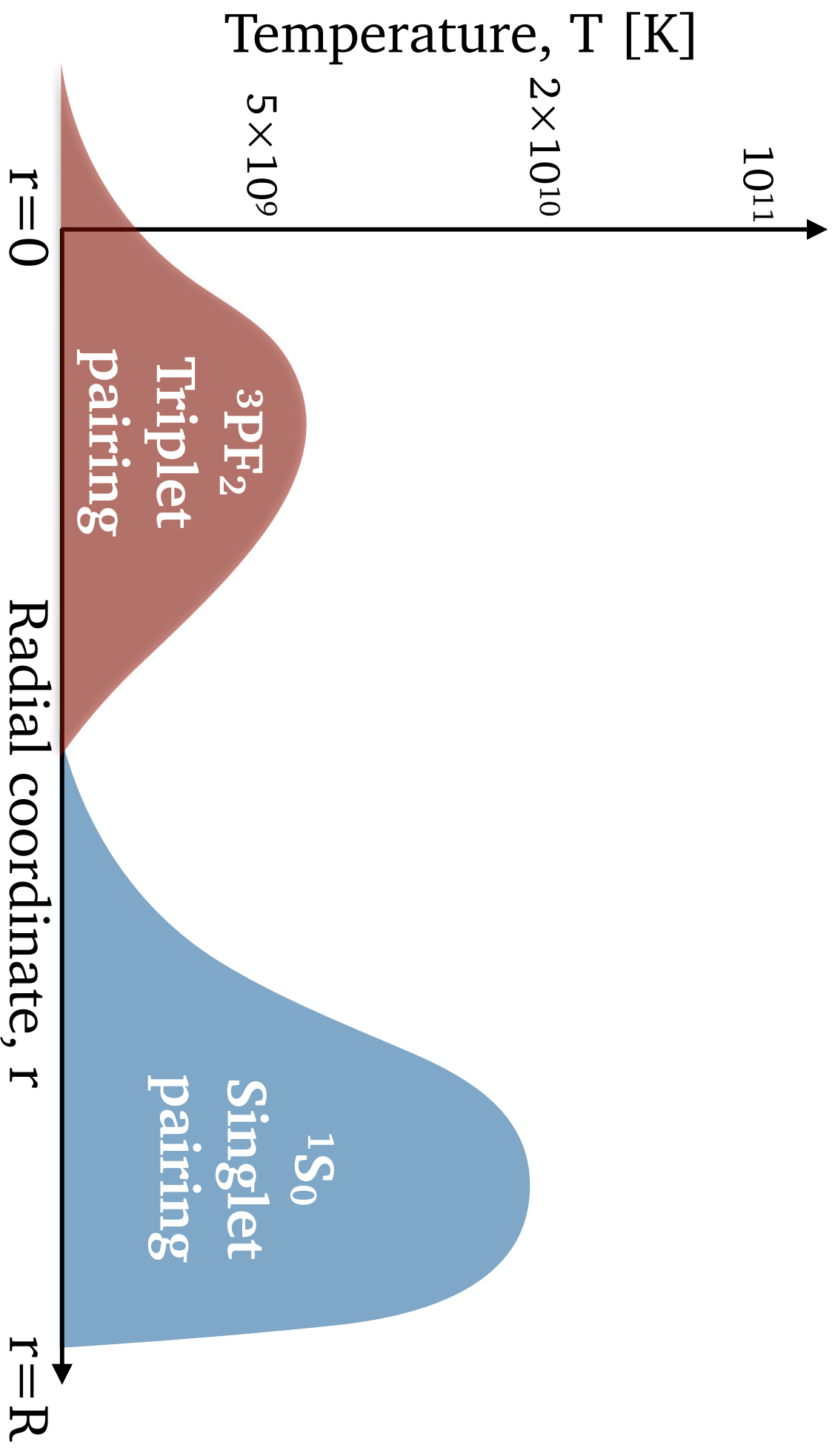
Cooling curve of neutron stars

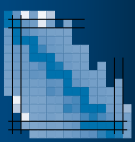


- Observational data available for a handful of NS
- Sensitive to interior physics (mostly pairing)

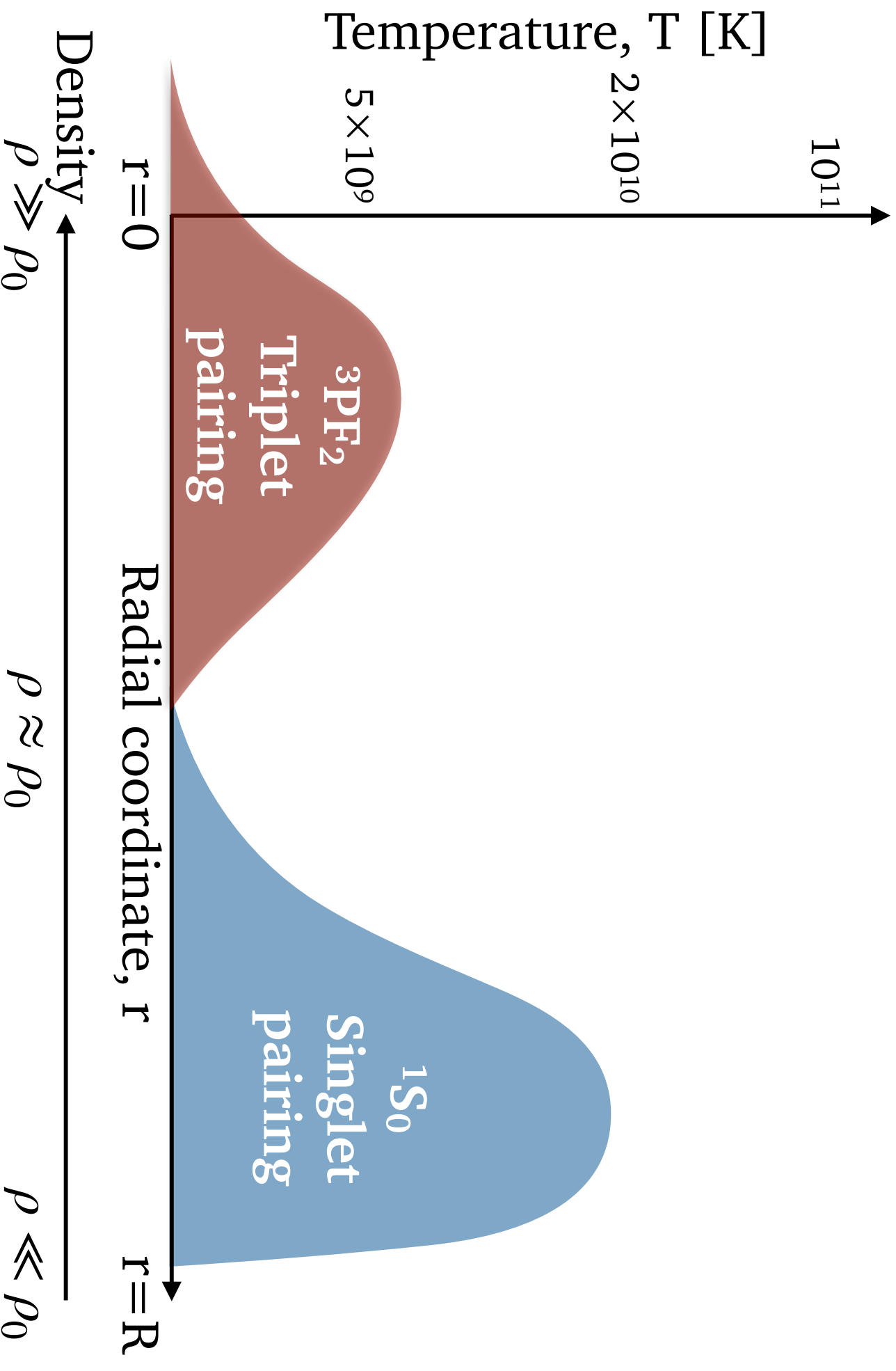


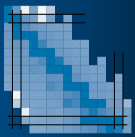
Pairing gaps & cooling





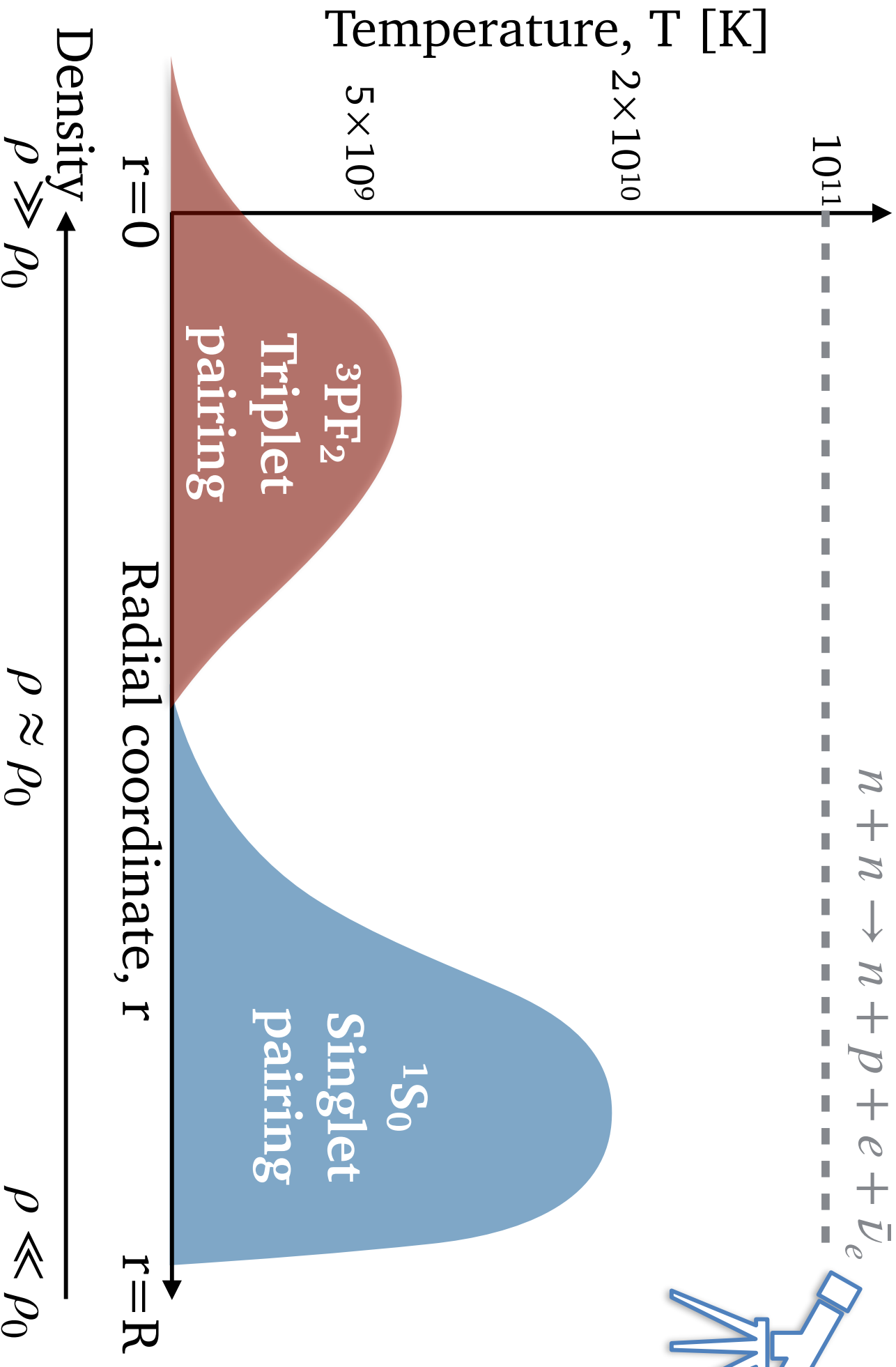
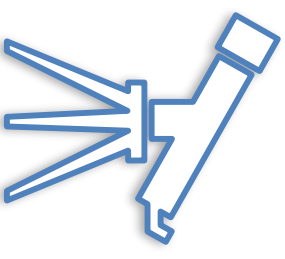
Pairing gaps & cooling

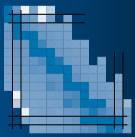




Pairing gaps & cooling

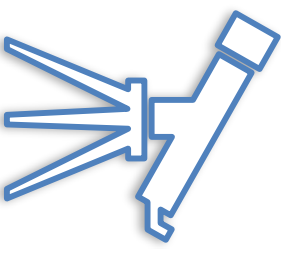
$t=0$, “hot” protonneutron star



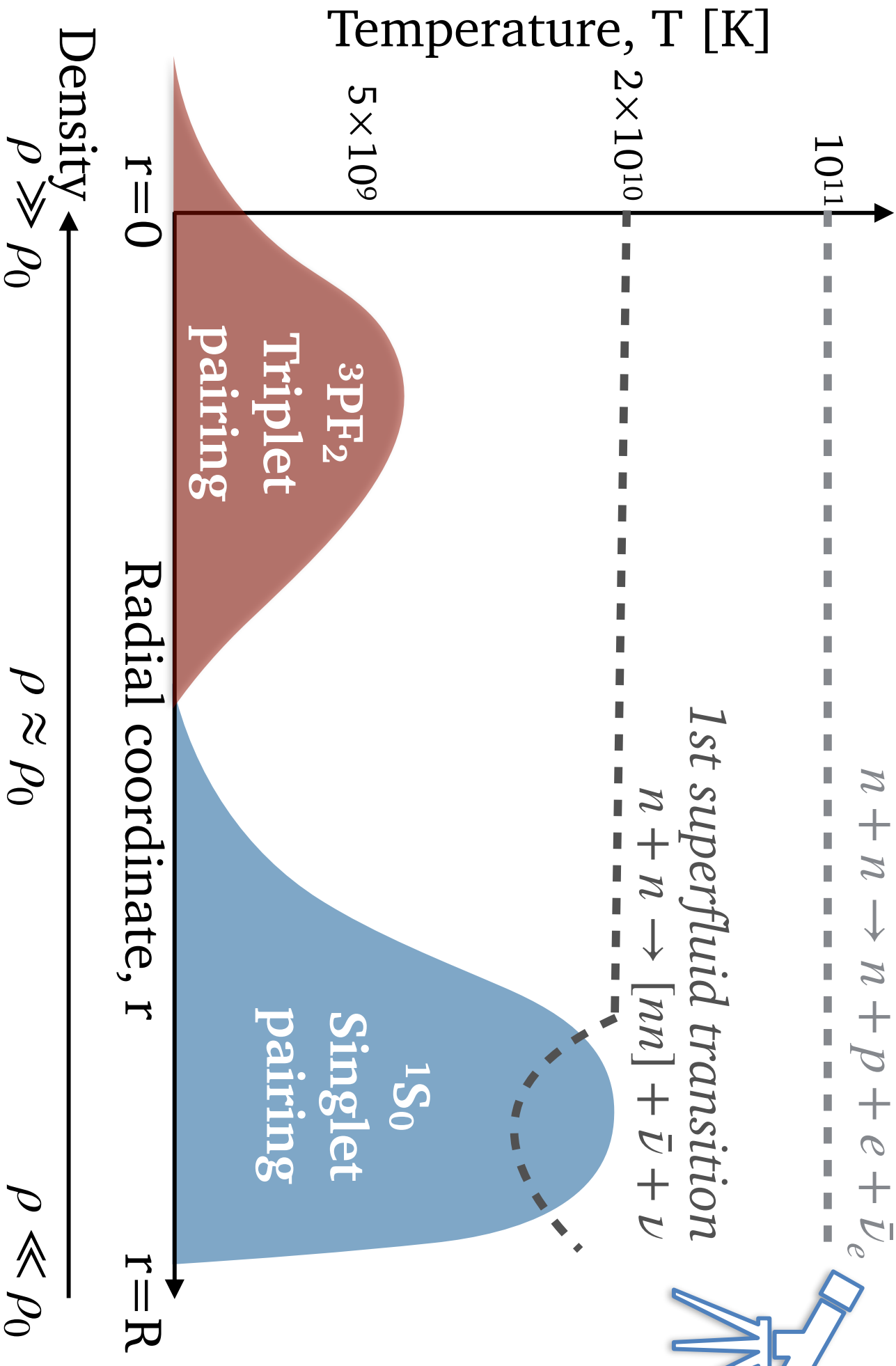


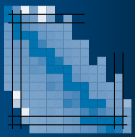
Pairing gaps & cooling

$t=0$, “hot” protonneutron star



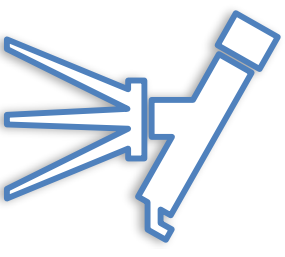
1st superfluid transition





Pairing gaps & cooling

$t=0$, “hot” protoneutron star



1st superfluid transition



Temperature, T [K]

10^{11}

2×10^{10}

5×10^9

$r=0$

Radial coordinate, r

$r=R$

3P_F_2
Triplet
pairing

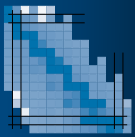
FAST COOLING!
 1S_0
Singlet
pairing

Density

$\rho \gg \rho_0$

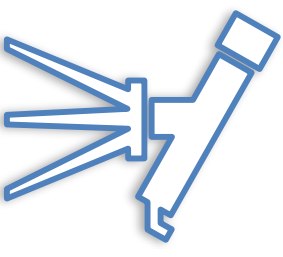
$\rho \approx \rho_0$

$\rho \ll \rho_0$



Pairing gaps & cooling

$t=0$, “hot” protoneutron star



1st superfluid transition



2nd superfluid transition



$3PF_2$

FAST COOLING!
 $1S_0$
Singlet
pairing

Triplet
pairing

Temperature, T [K]

10^{11}

2×10^{10}

5×10^9

$r=0$

Radial coordinate, r

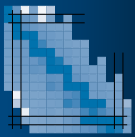
$r=R$

Density

$\rho \gg \rho_0$

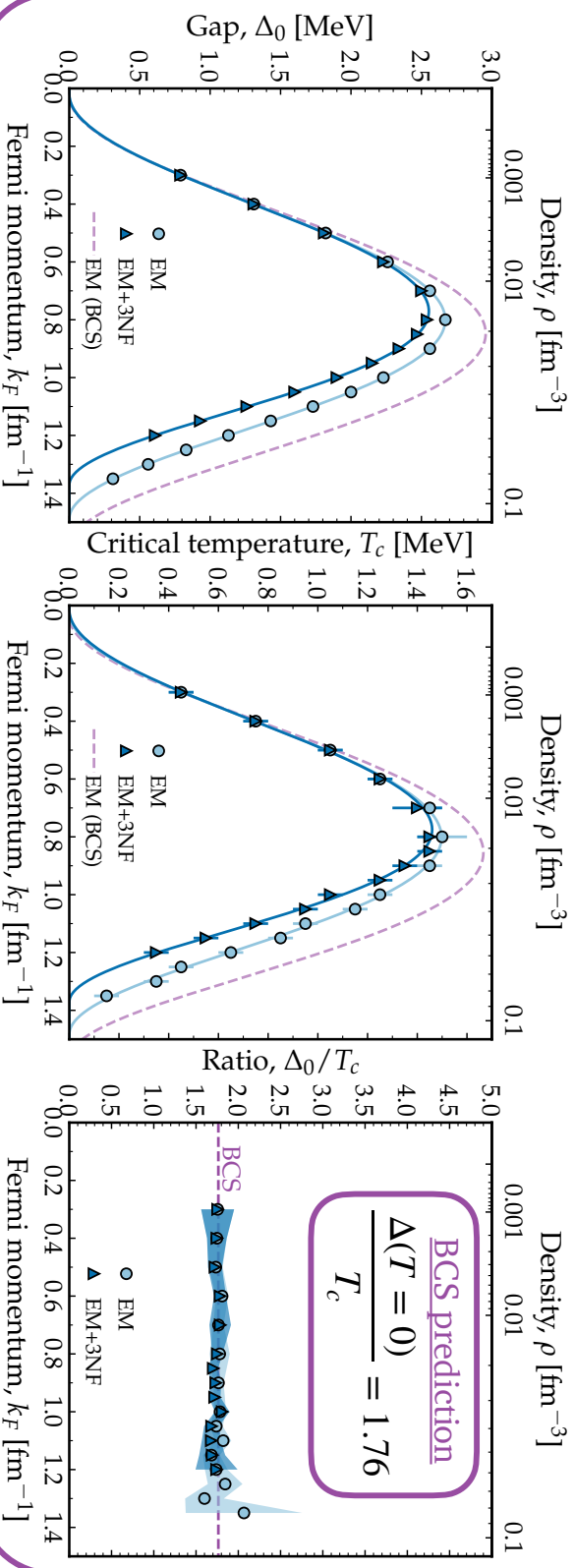
$\rho \approx \rho_0$

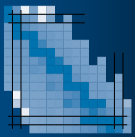
$\rho \ll \rho_0$



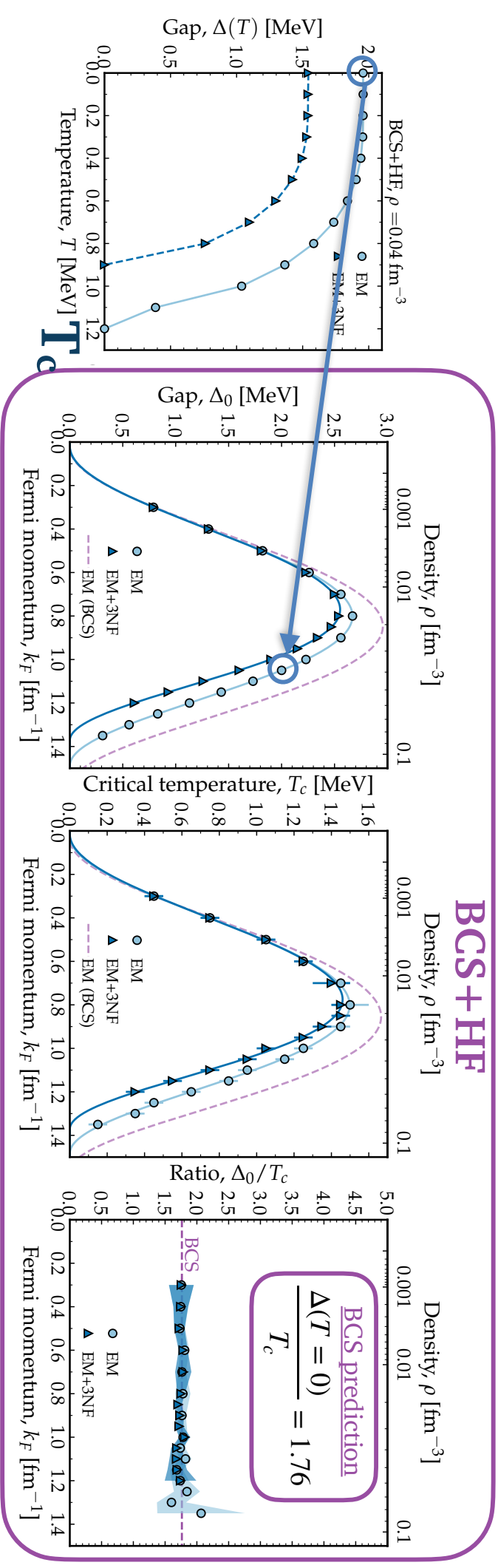
Why does it matter?

BCS+HF

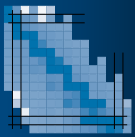




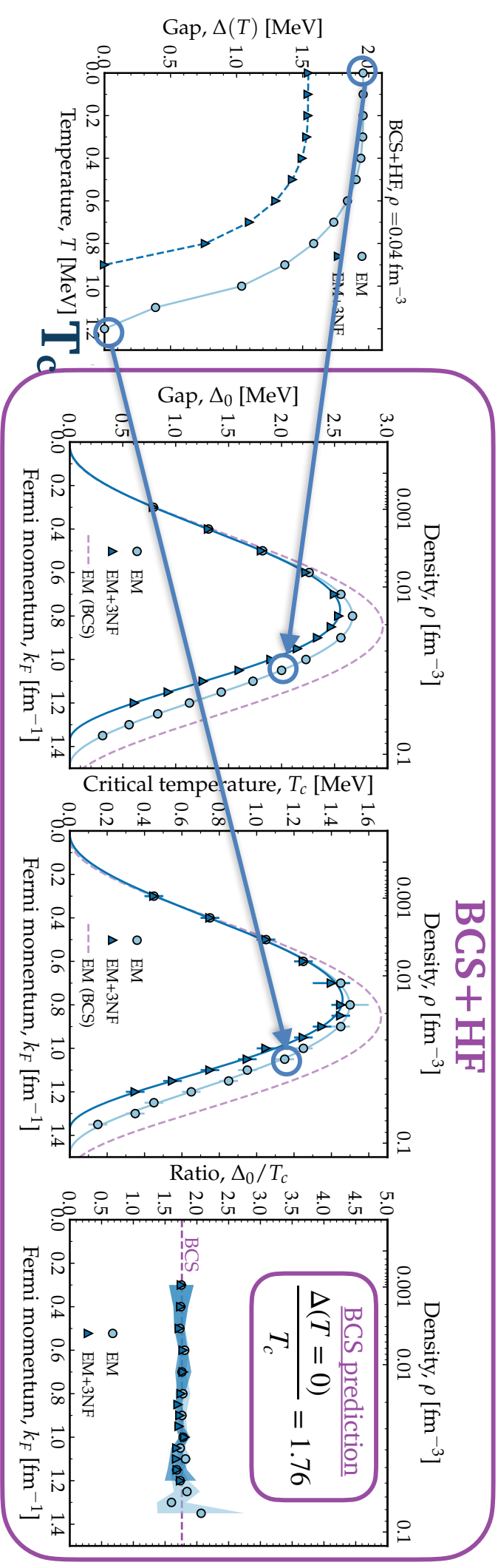
Why does it matter?

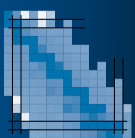


BCS+HF

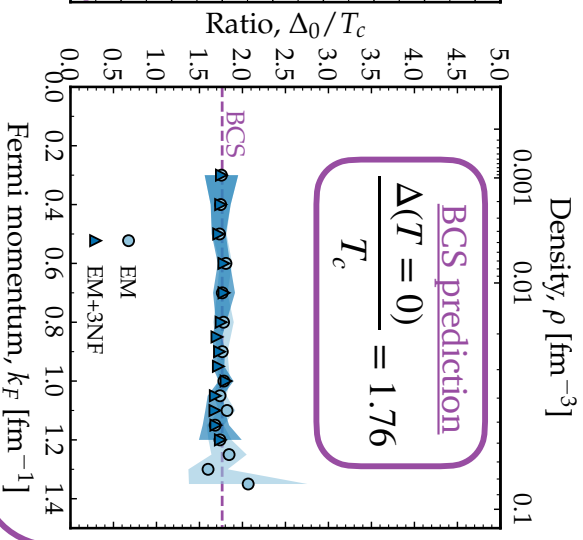
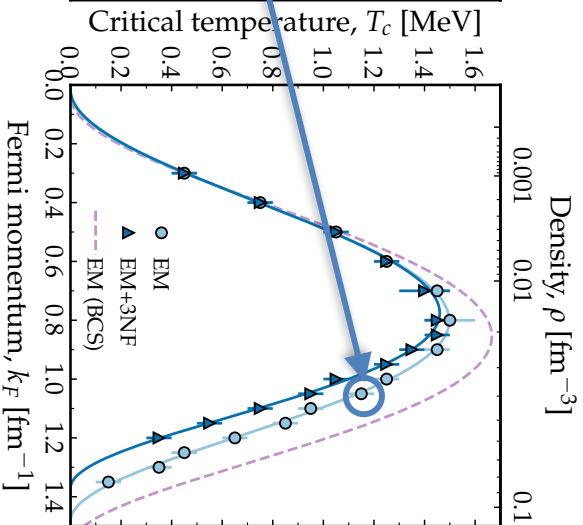
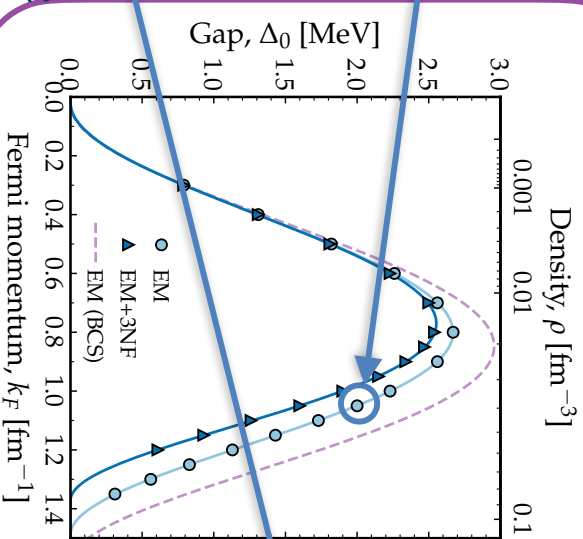
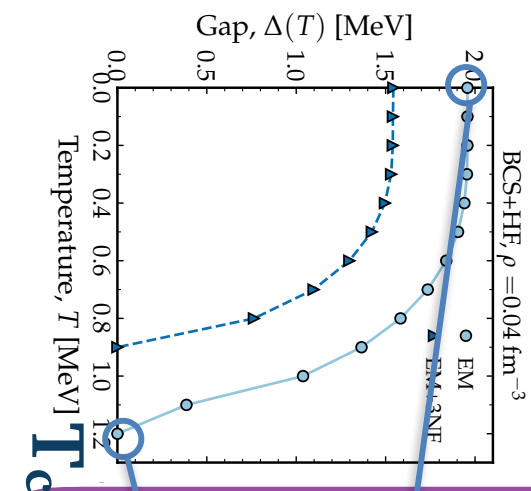


Why does it matter?



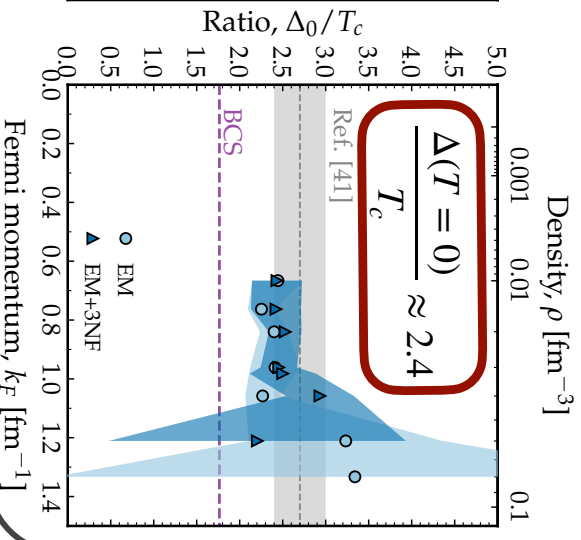
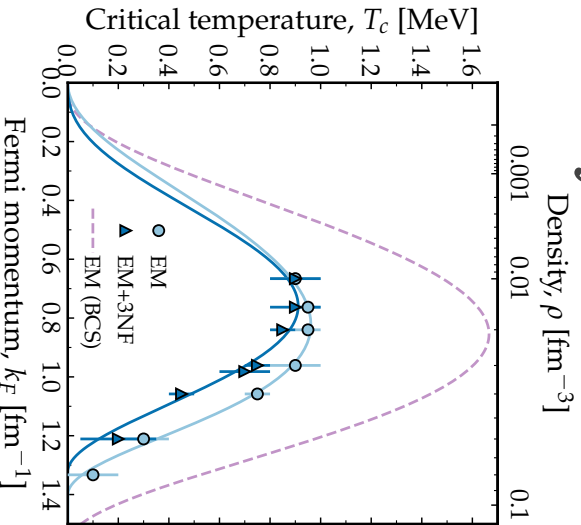
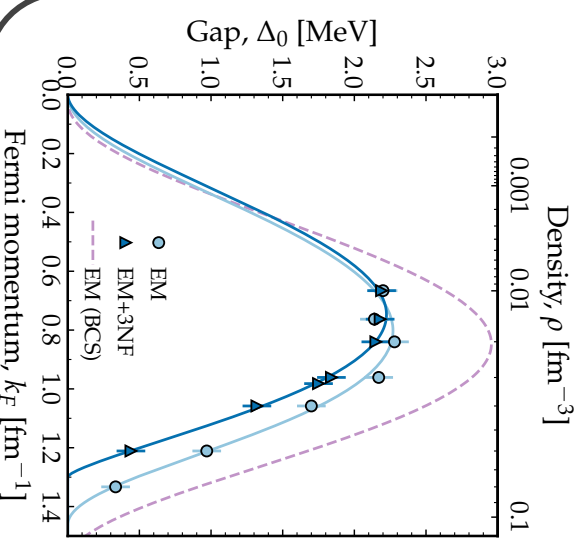


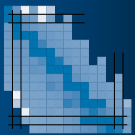
Why does it matter?



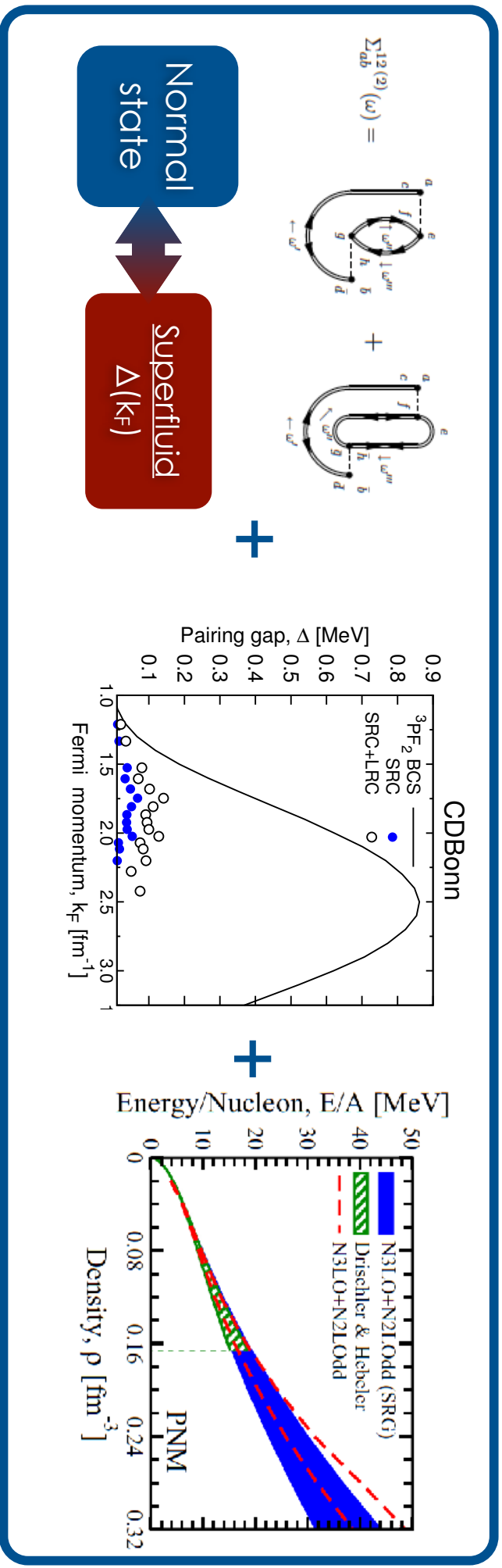
BCS+HF

“Beyond”-BCS





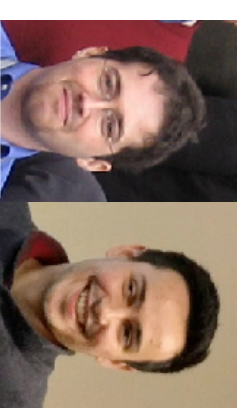
How to go beyond?

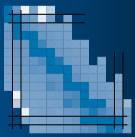


Existing frameworks difficult to generalise

Nambu-covariant **SCGF** technique

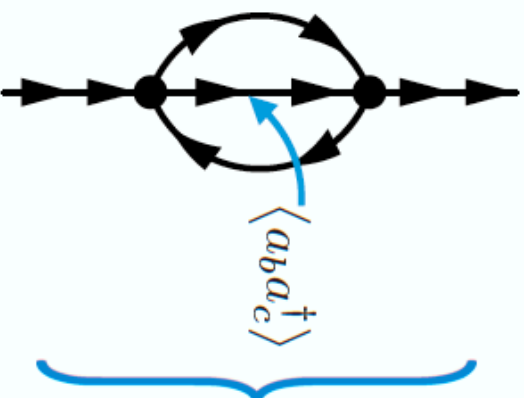
- Symmetry breaking ✓
- Finite temperature ✓
- Systematic expansion w diagrams ✓
- 3 nucleon forces ✓



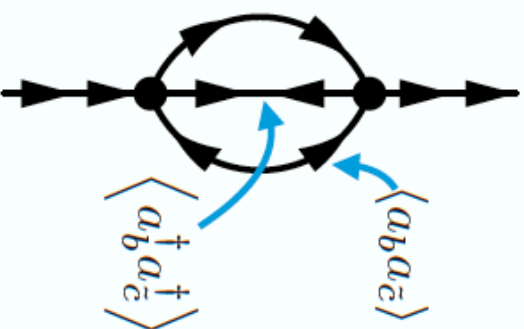


What was the issue before?

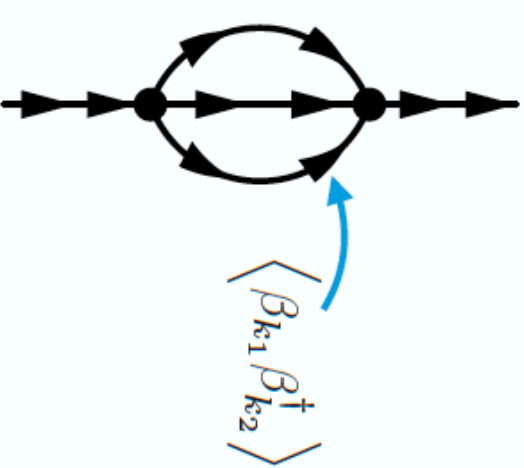
Standard PT



PT à la Gor'kov



PT à la Bogoliubov

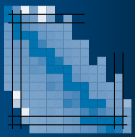


Bogoliubov transformation



Reformulation

- Based on Nambu fields
- Propagators transform contravariantly
- Vertices transform covariantly
- Un-oriented diagrammatic



Nambu-Covariant Perturbation Theory

Double dimension H space

$$\mathcal{H}_1^e \equiv \mathcal{H}_1 \times \mathcal{H}_1^\dagger$$

Basis

$$\mathcal{B}^e \equiv \mathcal{B} \cup \bar{\mathcal{B}}$$

$$|b\rangle$$

$$\langle \bar{b} |$$

Elements

$$\begin{pmatrix} |\Psi_1\rangle \\ \langle \Psi'_1| \end{pmatrix}$$

$$\mu = (b, l)$$

$$\bar{\mu} \equiv (b, \bar{l})$$

$$\therefore 1 \mapsto \bar{1} = 2$$

$$2 \mapsto \bar{2} = 1$$

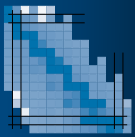
$$|b, 1\rangle \equiv \begin{pmatrix} |b\rangle \\ 0 \end{pmatrix}$$

$$|b, 2\rangle \equiv \begin{pmatrix} 0 \\ \langle \bar{b} | \end{pmatrix}$$

Product & metric tensor

$$g \left(\begin{pmatrix} |\Psi_1\rangle \\ \langle \Psi'_1| \end{pmatrix}, \begin{pmatrix} |\Psi_2\rangle \\ \langle \Psi'_2| \end{pmatrix} \right) \equiv \langle \Psi'_2 | \Psi_1 \rangle + \langle \Psi'_1 | \Psi_2 \rangle$$

$$g_{\mu\nu} \equiv g(|\mu\rangle, |\nu\rangle) = \delta_{\mu\bar{\nu}}$$



Nambu fields

Nambu fields

$$\begin{aligned} A^{(b,1)} &\equiv a_b, \\ A^{(b,2)} &\equiv \bar{a}_b, \\ \bar{A}^{(b,1)} &\equiv \bar{a}_b, \\ \bar{A}^{(b,2)} &\equiv a_b. \end{aligned}$$

$$A^\mu \equiv A^{(b,g)} = \begin{pmatrix} a_b \\ a_b^\dagger \end{pmatrix}$$

$$\bar{A}_\mu \equiv \bar{A}^{(b,g)} = (a_b^\dagger \quad a_b)$$

$$\bar{a}_b = a_b^\dagger \neq a_b^\dagger$$

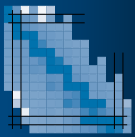
Commutator relations

(On extended indices!)

$$\begin{aligned} \{ A^\mu, A^\nu \} &= g^{\mu\nu}, \\ \{ A^\mu, \bar{A}_\nu \} &= g^{\mu\nu}, \\ \{ \bar{A}_\mu, A^\nu \} &= g_{\mu\nu}, \\ \{ \bar{A}_\mu, \bar{A}_\nu \} &= g_{\mu\nu} \end{aligned}$$

Co-or contravariant

$$\begin{aligned} \bar{A}_\mu &= \sum_\nu g_{\mu\nu} A^\nu, \\ \bar{A}_\mu &= \sum_\nu g_{\mu\nu} A^\nu, \\ A^\mu &= \sum_\nu g^{\mu\nu} \bar{A}_\nu, \\ A^\mu &= \sum_\nu g^{\mu\nu} \bar{A}_\nu. \end{aligned}$$

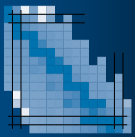


Why Bogoliubov tensor algebra?

$$\text{Tensor product: } r_{\nu_1}^{\mu_1} \mu_2 \mu_3 = s_{\nu_1}^{\mu_1} t^{\mu_2 \mu_3}$$

$$\text{Tensor contraction: } r_{\nu}^{\mu} = \sum_{\alpha} s_{\alpha}^{\mu} t^{\alpha}_{\nu}$$

- Co (contra-)variance under Bogoliubov transforms provide **invariant** expressions in **any basis**
- Potential to **optimise** the extended basis
- **Tensor-network** structure becomes **transparent**
- Leads to **diagrammatic** expansion (à la de Dominicis-Martin or Haussmann)
- Other **formalisms** through specific basis or metric



Hamiltonian partitioning

$$\Omega = \Omega_0 + \Omega_1$$

$$\Omega_0 = \frac{1}{2} \sum_{\mu\nu} U_{\mu\nu} A^\mu A^\nu$$

$$\Omega_1 = \sum_{k=1}^n \frac{1}{(2k)!} \sum_{\mu_1 \dots \mu_{2k}} \nu_{\mu_1 \dots \mu_{2k}}^{(k)} A^{\mu_1} \dots A^{\mu_{2k}}$$

Covariant k-body vertices

Green's functions

Contravariant k-body Green's function

$$(-1)^k \mathcal{G}^{\mu_1 \dots \mu_{2k}}(\tau_1, \dots, \tau_{2k}) \equiv \left\langle T [A^{\mu_1}(\tau_1) \dots A^{\mu_{2k}}(\tau_{2k})] \right\rangle$$

with $\langle \cdot \rangle = \text{Tr}(\cdot \rho)$ and $\rho \equiv \frac{e^{-\beta\Omega}}{\text{Tr}(e^{-\beta\Omega})}$

Unperturbed case: $\Omega \longleftrightarrow \Omega_0$

Fully antisymmetric vertex

Definition

$$\nu_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}]}^{(k)} \equiv \frac{1}{(2k)!} \sum_{\sigma \in S_{2k}} \epsilon(\sigma) \nu_{\mu_{\sigma(1)} \mu_{\sigma(2)} \dots \mu_{\sigma(2k-1)} \mu_{\sigma(2k)}}^{(k)}$$

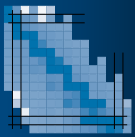
Antisymmetrisation defines a new $(0, 2k)$ -tensor

Not the case in a mixed representation

Propagators

$$-\mathcal{G}^{\mu\nu}(\omega_p) = \begin{array}{c} \mu \\ \parallel \\ \nu \end{array} \uparrow \omega_p$$

$$-(\mathcal{G}^{(0)})^{\mu\nu}(\omega_p) = \begin{array}{c} \mu \\ | \\ \nu \end{array} \uparrow \omega_p$$



Perturbative expansion

Hamiltonian partitioning

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Covariant k-body vertices

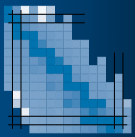
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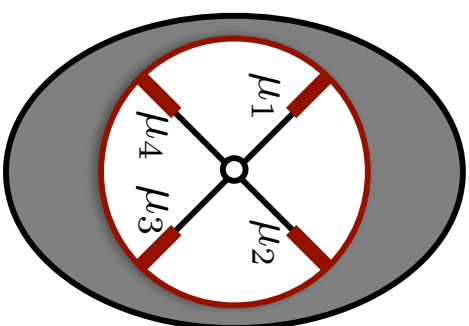


Why antisymmetric vertices?

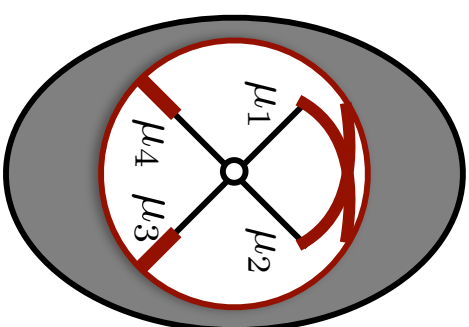
Un-symmetrised

vertex

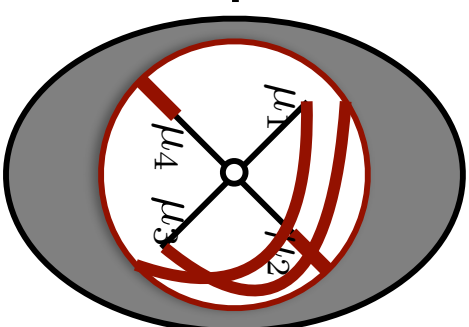
$$v_{\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}}^{(k)} =$$



+



+



+ ...

Antisymmetrized

vertex

$$v_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}]}^{(k)} =$$

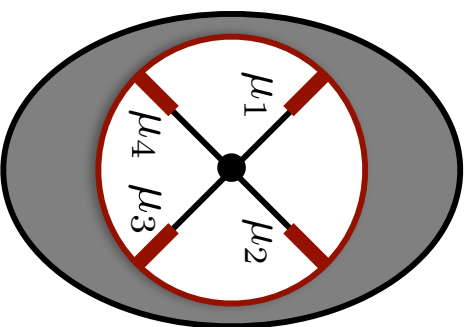
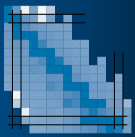


Diagram factorisation

- Derivations rely on
 - ▶ Wick theorem \Rightarrow sum over pairing
 - ▶ Sum over single-particle and Nambu indices
- **Extends Hugenholtz antisymmetrisation**
- Antisym is a **one-off pre-computing cost**



Perturbative expansion

Order n graphical rules

- Draw all topologically distinct connected unlabelled diagrams
- with $2k$ external legs
- with n vertices (for order n contributions)

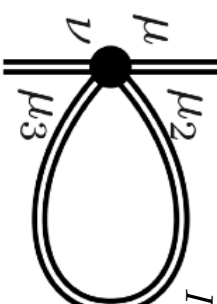
Feynman rules

- Label vertices from 1 to n
 - S is the number of vertex labels permutations leaving the diagram invariant
- For each line multiply by $-(\mathcal{G}^{(0)})^{\mu\nu}(\omega_e)$
- For each k -body vertex multiply by $v_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}]^{(k)}}$
- Sum over each internal μ index and each independent ω_e frequency
- Multiply by $\frac{(-1)^{n+L}}{S \times 2^T \prod_{l=2}^{l_{\max}} (l!)^m}$

Gaudin rules

- These simplify Matsubara sums
- Require spanning trees

Tadpoles are exceptional



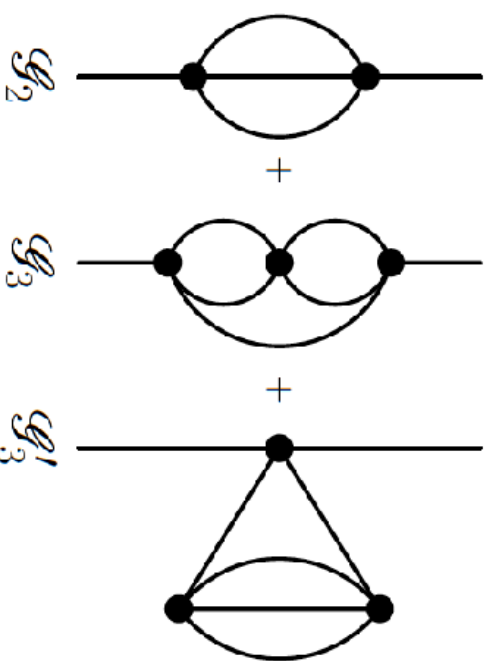
$$I_{\mu\nu} = \sum_{\mu_2 \dots \mu_{2k-1}} \frac{(-1)^k}{2^{k-1}(k-1)!} v_{[\mu \mu_2 \mu_3 \nu]}^{(k=2)} \times \frac{1}{\beta} \sum_{\omega_e} -\mathcal{G}^{\mu_2 \mu_3}(\omega_e) e^{-i\omega_e \eta p}$$

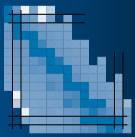
- Partially antisymmetrized vertices needed:

$$v_{[\mu_1 \dots \mu_x \dots \mu_y \dots \mu_{2k}]^{(k)}} \equiv \frac{2^p p!}{(2k)!} \sum_{\sigma \in S_{2k}/S_2^p \times S_p} \epsilon(\sigma) v_{\mu_{\sigma(1)} \dots \mu_{\sigma(x)} \dots \mu_{\sigma(y)} \dots \mu_{\sigma(2k)}}$$

- p internal lines are fixed
- k -body generalisation works

HFB partitioning 3rd order





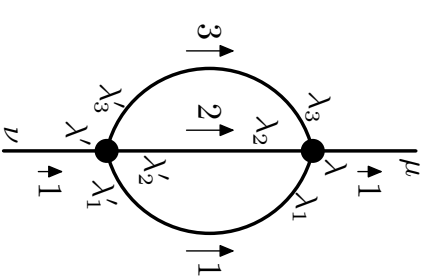
Advantages vs Gorkov

Gorkov GF

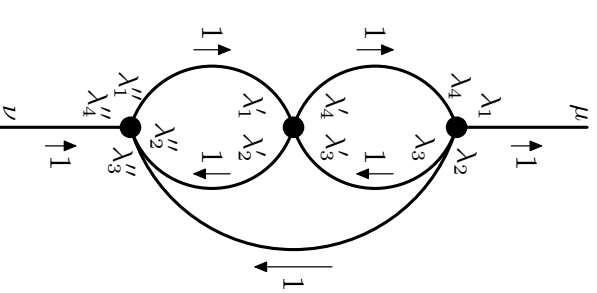
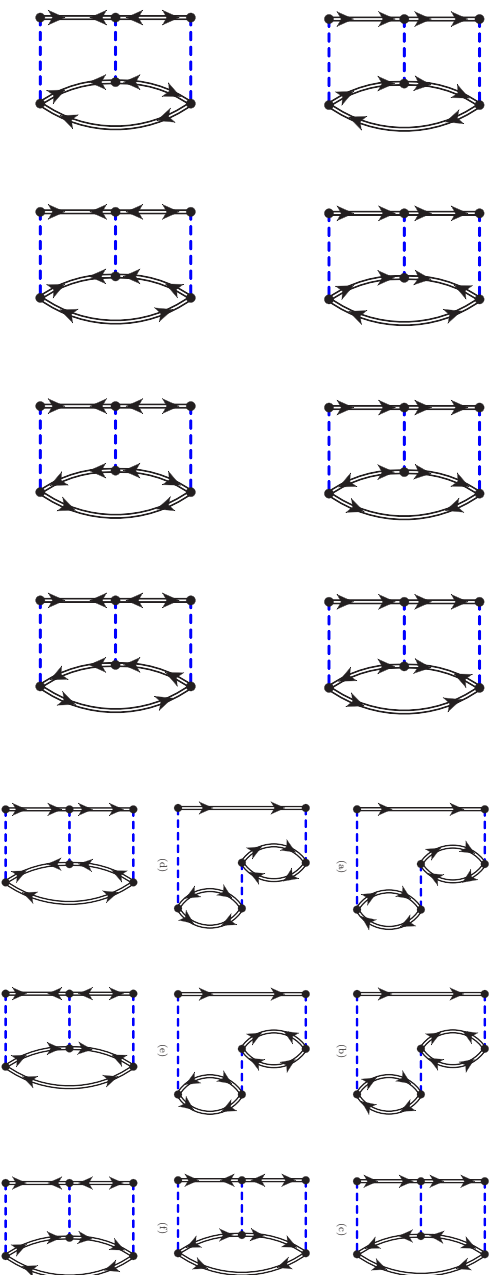
NCGF

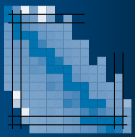
Order 2

$$\Sigma_{ab}^{11(2)}(\omega) =$$



Order 3





Self-consistent Green's function resummation

Dyson equation

Partitioning considered

$$\Omega = \underbrace{\frac{1}{2!} \sum_{\mu\nu} U_{\mu\nu} A^\mu A^\nu}_{\Omega_0} + \underbrace{\frac{1}{4!} \sum_{\alpha\beta\gamma\delta} v_{\alpha\beta\gamma\delta}^{(2)} A^\alpha A^\beta A^\gamma A^\delta}_{\Omega_1}$$

Dyson equation

$$\mathcal{G}^{\mu\nu}(\omega_n) = \mathcal{G}^{(0)\mu\nu}(\omega_n) + \sum_{\lambda_1\lambda_2} \mathcal{G}^{(0)\mu\lambda_1}(\omega_n) \Sigma_{\lambda_1\lambda_2}(\omega_n) \mathcal{G}^{\lambda_2\nu}(\omega_n)$$

Diagrammatic expansion of $\Sigma_{\mu\nu}(\omega_n)$

with unperturbed propagators

$$\Sigma_{\mu\nu}(\omega_n) = \frac{\mathcal{F}_{\mu\nu}(\omega_n) - \mathcal{F}_{\nu\mu}(-\omega_n)}{2}$$

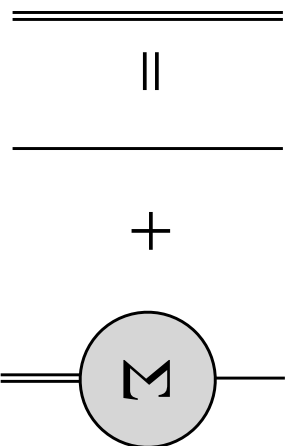
$\mathcal{F}_{\mu\nu}(\omega_n) = \sum$ 1PI diagrams with $\mathcal{G}^{(0)}$

with self-consistent propagators

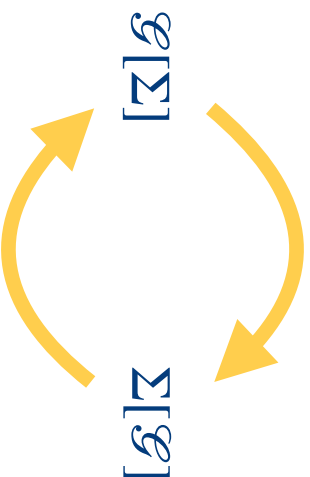
$$\Sigma_{\mu\nu}(\omega_n) = \frac{\mathcal{F}_{\mu\nu}(\omega_n) - \mathcal{F}_{\nu\mu}(-\omega_n)}{2}$$

$$\mathcal{F}_{\mu\nu}(\omega_n) = \sum \text{2PI diagrams with } \mathcal{G} \quad (= \mathcal{F}_{\mu\nu}(\omega_n))$$

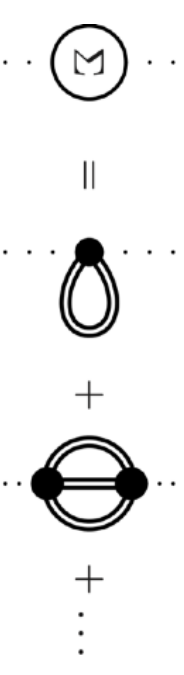
Diagrammatic representation

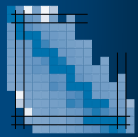


SCGF cycle



Self-energy expression

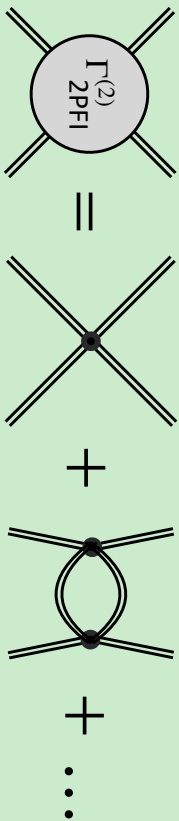




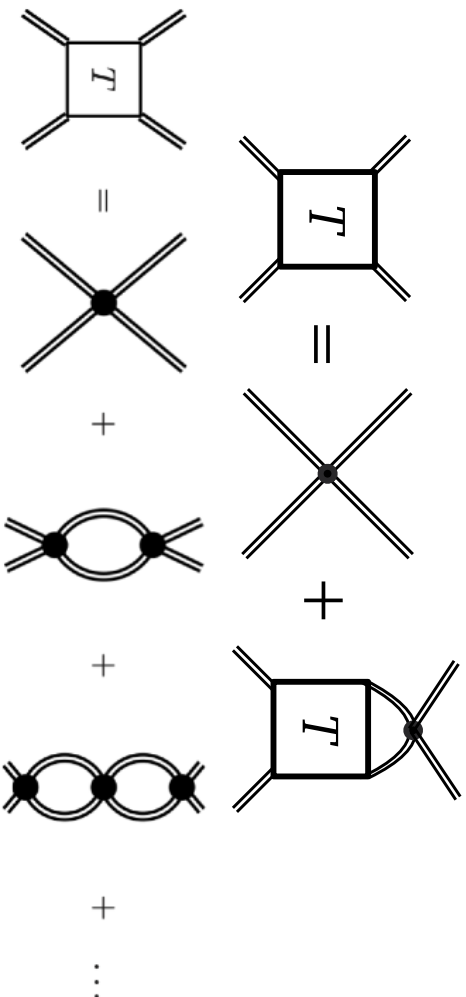
T-matrix: ladders

Approximations on $\Gamma_{2\text{PFI}}^{(2)}$

- Sum of all possible rungs



T-matrix $\equiv \Gamma^{(2)}$ in ladder approximation



Ladder approximation

- Analytic/Retarded/Advanced/Sp function \Rightarrow as usual

- T-matrix equation

$$T_{MN}(Z) = V_{MN}^{(2)} + \frac{1}{2} \sum_{LL'} V_{ML}^{(2)} \Pi^{LL'}(Z) T_{L'N}(Z)$$

where $V_{MN}^{(2)} \equiv v_{[\mu_1 \mu_2 \nu_1 \nu_2]}^{(2)}$, $M \equiv (\mu_1, \mu_2)$ & $N \equiv (\nu_1, \nu_2)$

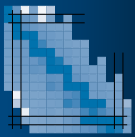
Solving the ladder

- Spectral representation

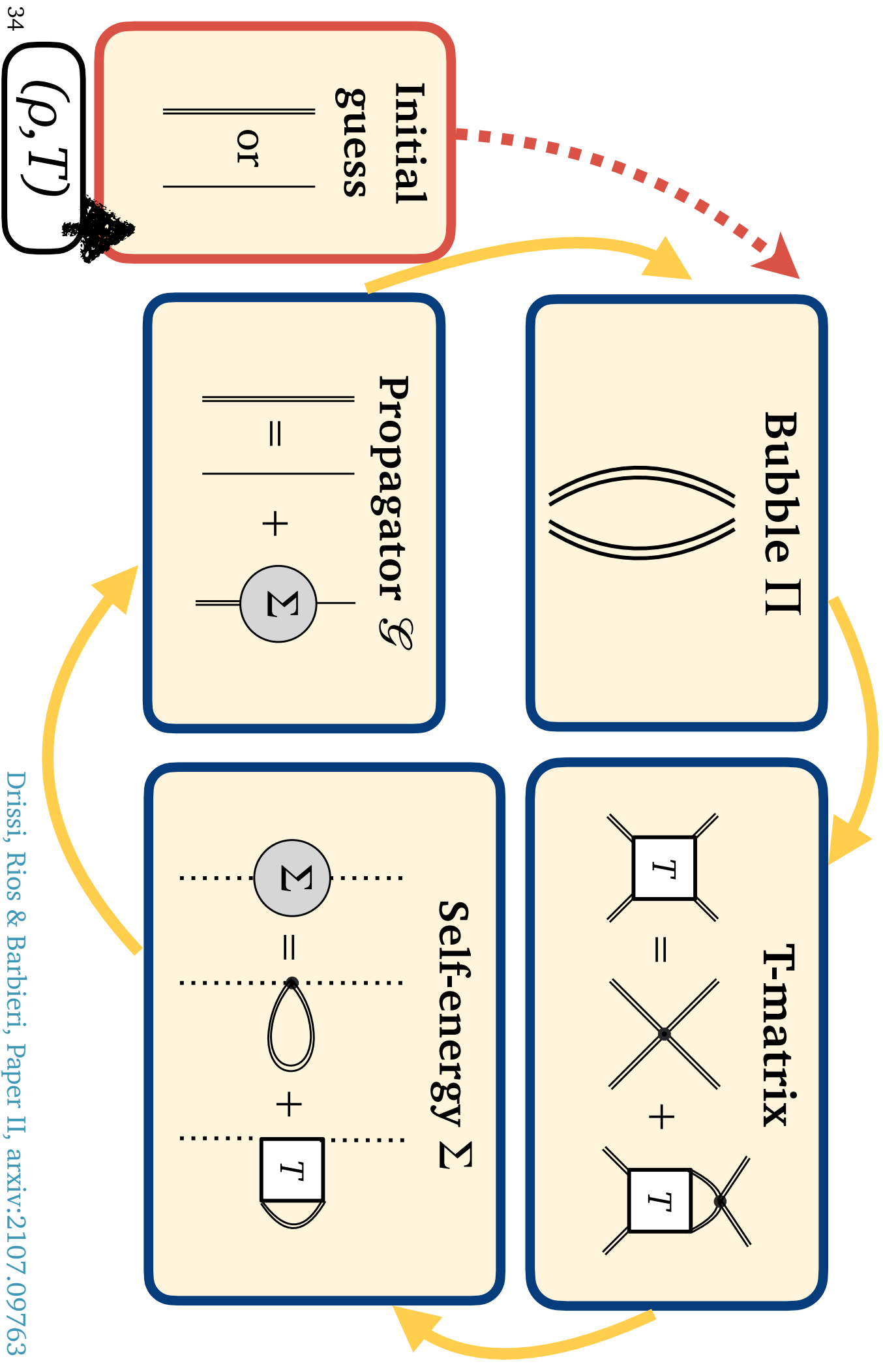
$$T_{MN}(Z) \equiv V_{MN}^{(2)} + \int_{-\infty}^{+\infty} \frac{d\Omega}{2\pi} \frac{T_{MN}(\Omega)}{Z - \Omega}$$

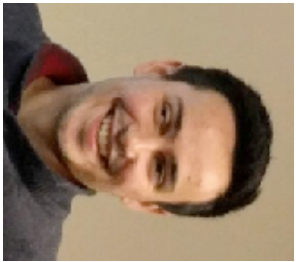
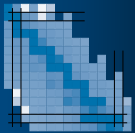
- Solution

$$T(\Omega) = iV^{(2)} \left\{ \left(gg - \frac{1}{2} \Pi^R(\Omega) V^{(2)} \right)^{-1} - \left(gg - \frac{1}{2} \Pi^A(\Omega) V^{(2)} \right)^{-1} \right\}$$



Nambu-Covariant Ladders





M. Drissi
UNIVERSITY OF
SURREY
TRIUMF

A. Carbone (Milano) + C. Barbieri (Milano)
W. H. Dickhoff (St Louis) + A. Polls



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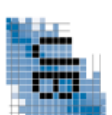


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