# Stríng Theory on Tíme-Dependent Orbífolds: from breakdown to breakthrough?



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20<sup>th</sup> October 2023, Sapienza, Università di Roma 4th PRIN meeting "String Theory as a bridge between Gauge Theories and Quantum Gravity"

A.A., R.Finotello, I.Pesando - On the Origin of the Divergences in Time-Dependent Orbifolds - Eur.Phys.J.C 80, (2020)
A.A., I.Pesando - Light-Cone Quantization of a Scalar Field on Time-Dependent Backgrounds - Eur.Phys.J.C 82, (2022)
A.A., R.J.Szabo, I.Pesando - The Role of Noncommutativity on the Null Boost Orbifold - To appear, (2023)

## Talk Outline

- Introduction and motivations
- The Null-Boost Orbifold (NBO)
- Scattering amplitudes divergences
- The origin of the divergences

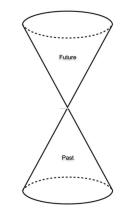
The role of noncommutativity

#### The cosmological point of view

- The **inflationary scenario** is the current paradigm of early universe cosmology, as it offers a successful solution to several problems (flatness, horizon, structure formation...), but it doesn't address the singularities matter.
- **Bouncing cosmologies**, like the Big Crunch/Big Bang scenario or the Ekpyrotic (cyclic) scenario, are more ambitious as they admit cosmological singularities.
- Besides the "origin of the universe" issue, **spacetime singularities** are anyway an open problem in General Relativity: the theory breaks down at the Planck scale and quantum gravity is needed.

### How do we study the Big Bang?

- Conceptual issues:
  - Is it the beginning of time?
  - If there is a bounce, how are the information before and after the singularity related?
  - How do we define observables and the S-matrix?

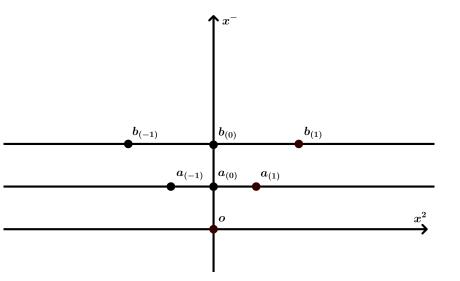


- We want to construct toy models capable of reproducing a **spacelike** singularity which appears in a certain region of space at a specific value of the time coordinate and then disappears.
- String theory, as a promising candidate for a theory of quantum gravity, is supposed to provide a satisfactory description of Big Bang type singularities.

#### The geometry of the NBO/1

The NBO is defined on  $(\mathbb{M}^3/\Gamma) \otimes \mathbb{R}^{D-3}$  where  $\Gamma$  is the subgroup of ISO(2,1) whose generator in lightcone coordinates is the Poincaré Killing vector  $k = 2\pi i \Delta J_{+2}$  with  $k^3 = 0$ .

Spacetime points are identified as 
$$x = \begin{pmatrix} x^- \\ x^2 \\ x^+ \\ \vec{x} \end{pmatrix} \sim e^{nk} x = \begin{pmatrix} x^- \\ x^2 + n(2\pi\Delta)x^- \\ x^+ + n(2\pi\Delta)x^2 + \frac{1}{2}n^2(2\pi\Delta)^2x^- \\ \vec{x} \end{pmatrix}, n \in \mathbb{Z}.$$



#### Scattering Amplitudes and Divergences: Closed Strings

Among the various issues which we can encounter when dealing with strings on timedependent backgrounds, the most urgent of this model concerns **interactions**.

It was shown that the four tachyons closed string  $2 \rightarrow 2$  tree-level scattering amplitude, the analogous of Virasoro-Shapiro amplitude in flat spacetime, exhibits unusual divergences in a specific kinematical regime (the Regge limit of high energy and small fixed angle):

$$A_{4T}^{closed} \sim \int^{q \sim \infty} \frac{dq}{|q|} q^{4-\alpha' \vec{p}_{\perp t}^2} \,,$$

[H.Liu, G.Moore, N.Seiberg (2002) - M.Berkooz, B.Craps, D.Kutasov, G.Rajesh (2005)]

which is divergent when  $\alpha' \vec{p}_{\perp t}^2 < 4$ .

#### Scattering Amplitudes and Divergences: Open Strings

This pathological behaviour has been interpreted in the literature as "the result of a large gravitational backreaction of the incoming matter into the singularity due to the exchange of a single graviton".

[H.Liu, G.Moore, N.Seiberg (2002) - M.Berkooz, B.Craps, D.Kutasov, G.Rajesh (2005) - L.Cornalba, M.S.Costa (2005)]

But if we perform an analogous computation for the four tachyons **open string "Veneziano"** amplitude we find:

$$A_{4T}^{open} \sim \int^{q \sim \infty} \frac{dq}{|q|} q^{1 - \alpha' \vec{p}_{\perp t}^2} \operatorname{tr} \left( \{T_1, T_2\} \{T_3, T_4\} \right), \quad [AA, R. Finotello, I. Pesando (2020)]$$

which is divergent when  $\alpha' \vec{p}_{\perp t}^2 < 1!$ 

Our claim is therefore that **string perturbation theory fails**, but the nature of the divergences is not strictly related to a gravitational backreaction, since here we are dealing with open string at tree level.

#### The geometry of the NBO/2

"Good" NBO coordinates

 $\begin{cases} x^- &= u \\ x^2 &= \Delta uz \\ x^+ &= v + \frac{1}{2} \Delta^2 uz^2 \end{cases}$ 

- New metric  $ds^2 = -2 \, du \, dv + (\Delta u)^2 (dz)^2 + d\vec{x}^2$
- New identifications  $(u, v, z, \vec{x}) \sim (u, v, z + 2\pi n, \vec{x})$

What happens if we study simple effective QFT models on this background?

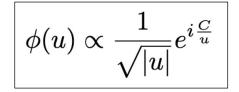
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#### Scalar QED on the NBO

$$S_{sQED} = \int_{\Omega} d^D x \sqrt{-\det g} \left( -(D^{\mu}\phi)^* D_{\mu}\phi - M^2 \phi^* \phi - \frac{1}{4} f^{\mu\nu} f_{\mu\nu} - \frac{g_4}{4} |\phi|^4 \right)$$
  
with  $D_{\mu}\phi = (\partial_{\mu} - i e a_{\mu})\phi$   $f_{\mu\nu} = \partial_{\mu}a_{\nu} - \partial_{\nu}a_{\mu}$ 

First we derive the eigenfunctions of the scalar d'Alembertian, which near the singularity u = 0 reads:



with  $C \sim l^2$ , where *l* is the discrete momentum associated to the compact direction *z*.

Now if we compute the interaction contact term of N scalar  $\phi_i$  and consider the specific case where all  $l_i = 0$ , we obtain:

$$\mathcal{A}_N \propto \int_{u \sim 0} du \, |u|^{-rac{N}{2}+1}$$

This result is technically not integrable for  $N \ge 4$ . The absence of a continuous exponential term (the  $C_i$ 's have isolated zeros due to the discrete momenta) does not allow for a distributional interpretation of  $A_N$  near the singularity.

#### The origin of the divergences/1

- We have shown that **fourth and higher order contact terms are ill defined** and we therefore claim that the root of the divergences lies in the **non existence of a well-defined underlying effective QFT**.
- If we think of String Theory as a UV completion of QFT amplitudes, we expect to find analogous behaviours in the string scattering amplitudes.
- That's exactly what happens, for example, with the open string 3-point amplitude with two tachyons and the first massive state, since it may be divergent when some physical polarizations are chosen and all discrete momenta  $l_i = 0$ :

$$\mathcal{A}_{TTM}^{open} \propto \int_{u \sim 0} du \, \frac{1}{|u|^{5/2}} \, \mathrm{tr}\left(\{T_{(1)}, T_{(2)}\}T_{(3)}\right)$$

[AA, R.Finotello, I.Pesando (2020)]

### The origin of the divergences/2

• We found an analogous result on the boost orbifold (BO).

• The gravitational backreaction may however reappear at the one loop level in the open string case.

• A recent analysis of these models at the quantum mechanical level suggests that the full Hamiltonian theory is well-defined. [*I.Pesando (2022)*]

• Is it possible to recover a meaningful perturbative formulation on this background?

#### Seiberg-Witten Map

Inspired by the resolution of some static orbifold singularities, we add a Kalb-Ramond background B-field:

$$S = \frac{1}{4\pi\alpha'} \int_{\Sigma} d^2\sigma \, \left( g_{\mu\nu} \partial_a x^{\mu} \partial^a x^{\nu} - 2\pi i \alpha' \epsilon^{ab} B_{\mu\nu} \partial_a x^{\mu} \partial_b x^{\nu} \right) \quad \text{with} \quad B = B_{-2} \, dx^- \wedge dx^2.$$

Applying the Seiberg-Witten map  $\frac{1}{(g+2\pi\alpha' B)_{\mu\nu}} = (G+\frac{\theta}{2\pi\alpha'})^{\mu\nu}$  we get:

$$G^{\mu\nu} = \begin{pmatrix} 0 & -1 & 0 \\ -1 & -(2\pi\alpha'b)^2 & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \theta^{\mu\nu} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & (2\pi\alpha')^2b \\ 0 & -(2\pi\alpha')^2b & 0 \end{pmatrix} \quad \text{with} \quad b = B_{-2}.$$

In the decoupling zero slop limit  $\alpha' \to 0$  with  $b \sim (\alpha')^{-2}$ :

- $\theta = (2\pi \alpha')^2 b$  stays finite
- $G^{\mu\nu} = \eta^{\mu\nu}$  (after a coordinate change)

• Boundary propagator:  $\langle x^{\mu}(\tau)x^{\nu}(\tau')\rangle = -\alpha' G^{\mu\nu}\log(\tau-\tau')^2 + \frac{i}{2}\theta^{\mu\nu}\epsilon(\tau-\tau')$ 

#### Noncommutative QFT

In the ordinary QFT, the N-point scalar function reads

$$\mathcal{A}_N \propto \int_{u \sim 0} du \, |u|^{-rac{N}{2}+1}$$

In the noncommutative QFT, the N-point scalar function reads

$$\mathcal{A}_N \propto \int_{u \sim 0} du \, |u|^{-rac{N}{2}+1} e^{i rac{ heta f(p_{-i})}{u}}$$

with  $f(p_{-i})$  a combination of the continuous momenta  $p_{-i}$  along the *u* direction

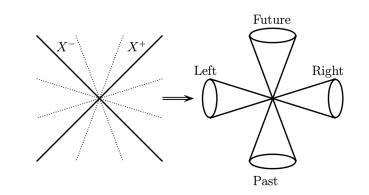
This means that we are avoiding the presence of isolated zeros and therefore noncommutativity makes plausible the interpretation of  $A_N$  as a distribution: in particular, this integral satisfies the properties of *distributionally integrable functions* defined by R.Estrada and J.Vindas in 2012.

### Time-dependent orbifolds: the other models

Up to ISO(2,1) conjugations, there exist only four different  $\mathbb{M}^3$  time-dependent orbifolds:

Orbifold	Generator
Boost (BO)	$2\pi i\Delta J_{+-}$
Null-Boost (NBO)	$2\pi i\Delta J_{+2}$
Shifted-Boost (SBO)	$2\pi i (\Delta J_{+-} + RP_2)$
O-Plane (OPO)	$2\pi i (\Delta J_{+2} + RP_{-})$

[L.Cornalba, M.S.Costa (2005)]



The Boost Orbifold

### To sum up

- Time-dependent orbifolds are useful toy models to reproduce and study cosmological singularities
- The origin of the unexpected divergences which appear in string scattering amplitudes can be traced back to the failure of the standard Feynman diagrammatic approach of the underlying effective QFT
- The noncommutative QFT which arises as a decoupling limit after the introduction of a *B*-field seems promisingly well-defined

#### **Conclusions and Outlooks**

- Understand and define better the noncommutative QFT [AA, I.Pesando, R.J.Szabo (2023)]
- Extend this singularity resolution to the other time-dependent orbifolds
- Discuss the relation of the noncommutative QFT with different decoupling limits which lead to Noncommutative Open String Theory (S-duality? OM theory?)
- Verify if we can recover a full string perturbation theory in presence of the *B*-field, even beyond tree-level

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### Thank you for your attention!