

Black holes as point particles: from amplitudes to self-force

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THE UNIVERSITY
of EDINBURGH

"Gravitational Wave Probes of Fundamental Physics" workshop

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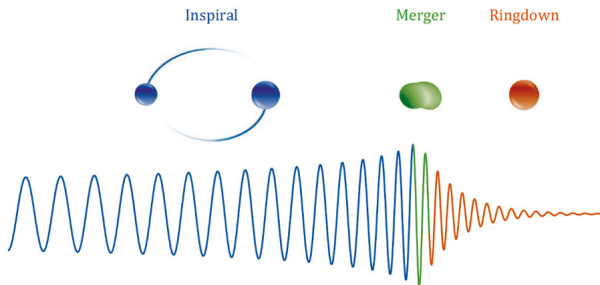
- 1 Motivation and introduction
- 2 The Post-Minkowskian expansion and classical amplitudes
- 3 From scattering to bound observables
- 4 From amplitudes to self-force resummation
- 5 Conclusion

Motivation and introduction (I)

- The recent discovery of gravitational waves **calls for new analytical techniques** to study **the two-body problem**.

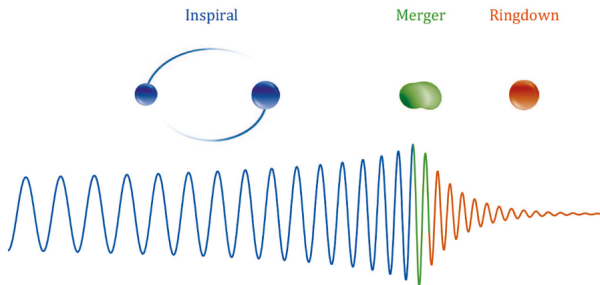
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- Today: focus on the **inspiral phase**, where we can model **compact objects as point particles** in the spirit of **effective field theory** [Goldberger, Rothstein]

Motivation and introduction (II)

- Idea: use **particle field theory tools** (\rightarrow **scattering amplitudes**)

Real world	EFT of point particles
Compact objects of mass M	Point particles of mass M
Spin effects of magnitude a	Spinning particles of classical spin a
Tidal effects, GR curvature corrections	Higher-dimensional operators
Absorption effects	Non-unitary absorption dofs

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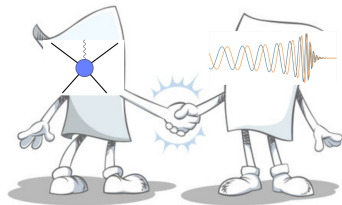
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- Why amplitudes? (adapted to **scattering orbits**... **bound orbits?** **Stay tuned!**)

Amplitudes are **gauge-invariant**, universal objects which encode in a compact way the **perturbative scattering** dynamics for **point particles** in a QFT.

New perspective on GR!



KMOC formalism for the two-body problem (I)

- **Two-body scattering in GR**: Consider as initial state **two massive particles** separated by an **impact parameter** b^μ [Kosower,Maybee,O'Connell=KMOC]

$$|\psi_{\text{in}}\rangle = \int d\Phi(p_1, p_2) \psi_1(p_1) \psi_2(p_2) e^{i(b \cdot p_1)/\hbar} |p_1 p_2\rangle$$

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- The dynamics of the evolution is determined by the action

$$S = -\frac{1}{16\pi G_N} \int d^4x \sqrt{-g} R + \sum_{j=1}^2 \frac{1}{2} \int d^4x \sqrt{-g} (g^{\mu\nu} \partial_\mu \phi_j \partial_\nu \phi_j - m_j^2 \phi_j^2) + S_{\text{GF}}$$

where we perform the perturbative expansion

$$g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}, \quad \kappa = \sqrt{32\pi G_N} \rightarrow \text{Post-Minkowskian expansion in } G_N.$$

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- See **Gambino's** talk if you are interested in the metric from point particles!

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- We can compute **classical observables** \mathcal{O} with expectation values

$$\langle \psi_{\text{in}} | \mathcal{S}^\dagger \mathcal{O} \mathcal{S} | \psi_{\text{in}} \rangle \Big|_{\hbar \rightarrow 0} = 2\Re i \langle \psi_{\text{in}} | \mathcal{O} T | \psi_{\text{in}} \rangle \Big|_{\hbar \rightarrow 0} + \langle \psi_{\text{in}} | T^\dagger \mathcal{O} T | \psi_{\text{in}} \rangle \Big|_{\hbar \rightarrow 0}$$

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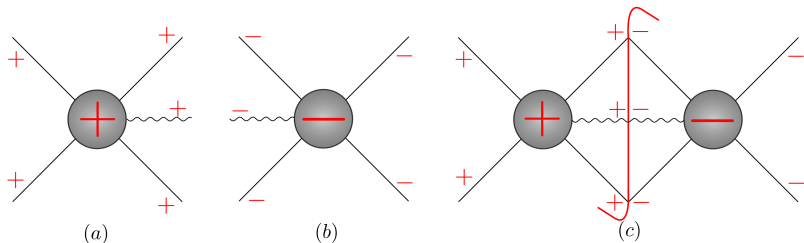
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- Connection with the “classical” on-shell reduction of the in-in approach in the (+)/(-) basis [Britto, RG, Jehu]

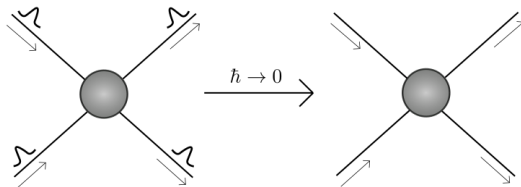


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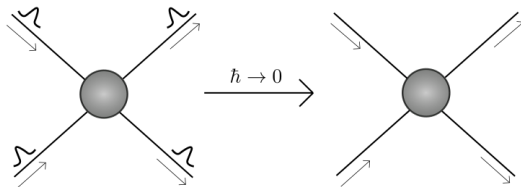
$$\psi(p) = \mathcal{N} m^{-1} \exp\left[-\frac{p \cdot u}{\hbar l_c / \ell_w^2}\right] \xrightarrow{\text{rest frame}} \mathcal{N}' \exp\left(-\frac{p^2}{2m^2 \ell_c^2 / \ell_w^2}\right)$$

where p^μ is the momentum, $\ell_{c,j} = \hbar/m_j$ is the Compton wavelength, ℓ_w the intrinsic spread of the wavefunction. If b^μ is the impact parameter we require,

$$\ell_{c,j} \ll \ell_w \ll b = \sqrt{-b^2}.$$

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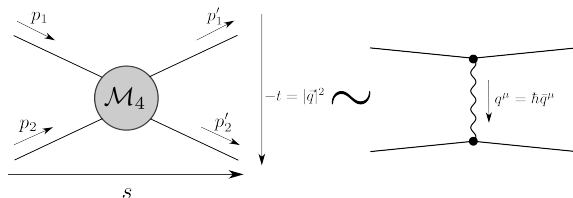
- **Massless particles**: use **coherent states**! [Cristofoli, RG, Kosower, O'Connell]

Classical limit of scattering amplitudes

- Conservative 4-pt amplitude $\mathcal{M}_4(p_1, p_2; p'_1, p'_2)$: in the classical limit $\hbar \rightarrow 0$

$$\begin{aligned} p_1^\mu &:= p_A^\mu + \hbar \frac{\bar{q}^\mu}{2}, & (p'_1)^\mu &:= p_A^\mu - \hbar \frac{\bar{q}^\mu}{2}, & s &= (p_A + p_B)^2, \\ p_2^\mu &:= p_B^\mu - \hbar \frac{\bar{q}^\mu}{2}, & (p'_2)^\mu &:= p_B^\mu + \hbar \frac{\bar{q}^\mu}{2}, & t &= -\hbar^2 |\vec{q}|^2, \end{aligned}$$

where p_A, p_B are the classical momenta and q is the momentum transfer.



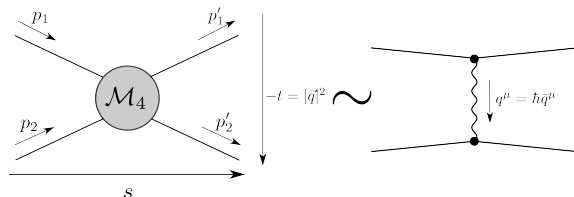
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- Generalization for the $4 + M$ -pt amplitude $\mathcal{M}_{4+M}(p_1, p_2; p'_1, p'_2, k_1, \dots, k_M)$

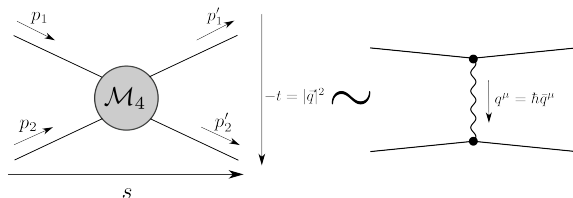
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- Main lesson:** only wavevectors $\bar{q}_{1,2}^\mu, \bar{k}_j$ are classical, need to restore \hbar !

Waveforms from KMOC formalism: an example

- How is the **waveform** derived from **scattering amplitudes**?

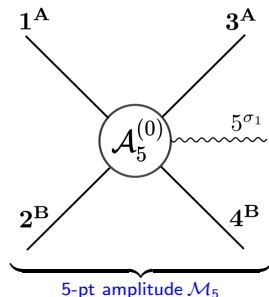
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- The on-shell expectation value of the **time-domain waveform** relevant for the inspiral phase is [Cristofoli,RG,Kosower,O'Connell]

$$\langle \psi_{\text{in}} | \mathcal{S}^\dagger h_{\mu\nu}(x) \mathcal{S} | \psi_{\text{in}} \rangle = \frac{1}{\hbar^{\frac{1}{2}}} 2\Re \sum_{\sigma=\pm} \int d\Phi(k) \varepsilon_\mu^{*(\sigma)}(k) \varepsilon_\nu^{(\sigma)}(k) \tilde{j}(b; k^\sigma) e^{-ik \cdot x / \hbar}$$

where at leading Post-Minkowskian order only the **5-pt amplitude** is relevant

$$\tilde{j}(b; k^{\sigma_1}) \equiv \int d\Phi(p'_1 p'_2 p_1 p_2) \psi^*(p'_1, p'_2) \psi(p_1, p_2) e^{-ib \cdot \bar{q}_1}$$

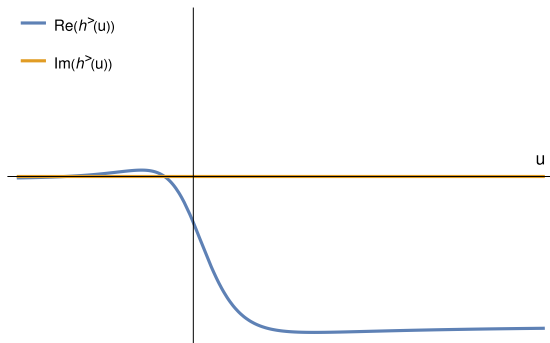


The Kovacs-Thorne waveform for hyperbolic encounters

- At tree-level: [simple analytic time-dependent expressions](#) available! [De Angelis, RG, Novichkov; Jakobsen, Mogull, Plefka, Steinhoff]

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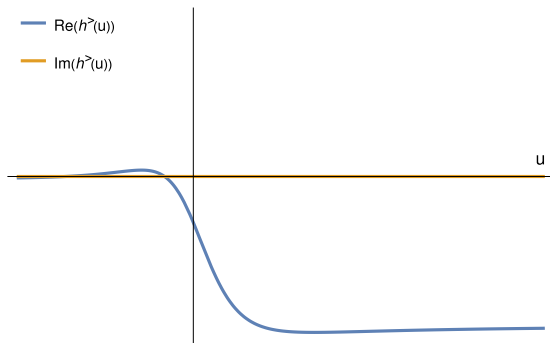
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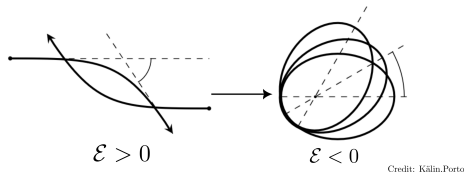
- How about **bound waveforms for compact binaries coalescence**?

From scattering to bound observables (I)

- Classical scattering amplitudes describe hyperbolic encounters. If we define

$$\mathcal{E} := \frac{E - m_1 - m_2}{\mu}, \quad p_\infty^2 = -\tilde{p}_\infty^2 = \frac{E^2 - (m_1 + m_2)^2}{2m_1 m_2},$$

we have $\mathcal{E}, p_\infty^2 > 0$ for scattering orbits and $\mathcal{E}, p_\infty^2 < 0$ for bound orbits.

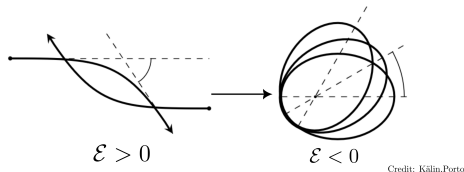


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- Two powerful methods to **extract bound state physics from amplitudes**:
 - 1) Extract perturbatively the classical potential (\sim **Hamiltonian**) valid for **arbitrary orbits** [Niell,Rothstein;Cheung,Rothstein,Solon]
 - 2) **Gauge invariant map between scattering and bound observables** [Kälin,Porto]

$$\mathcal{O}^>(\mathcal{E} > 0, J, c_X, a_1, a_2, m_1, m_2) \rightarrow \mathcal{O}^<(\mathcal{E} < 0, J, c_X, a_1, a_2, m_1, m_2)$$

which can be derived from the **Bethe-Salpeter eq.** [Adamo,RG; Adamo,RG,Ilderton].

From scattering to bound observables (II)

- For **aligned-spin binaries** where the motion remains on the equatorial plane there is a **conjectural dictionary** [Kälin,Porto;Saketh,Vines,Steinhoff, Buonanno;Cho,Kälin,Porto;Adamo,RG; Heissenberg;Adamo,RG,Ilderton]

Bound observable	Scattering observable
$\Delta\Phi(\tilde{p}_\infty, L, a, c_X)$	$\chi(-i\tilde{p}_\infty, L, a, c_X) + \chi(+i\tilde{p}_\infty, L, a, c_X)$
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- New waveform map** (up to **1PN and tree-level**)[Adamo,RG,Ilderton]

$$h^{<\text{dyn}}(u; \tilde{p}_\infty, L, a, c_X) = h^{>\text{dyn}}(u; +i\tilde{p}_\infty, L, a, c_X)$$

in agreement with the prescription for the orbital elements [Damour,Deruelle]

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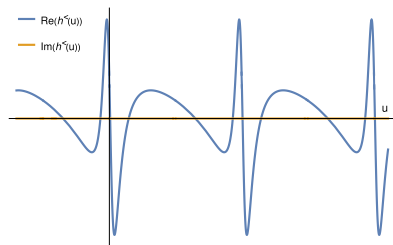
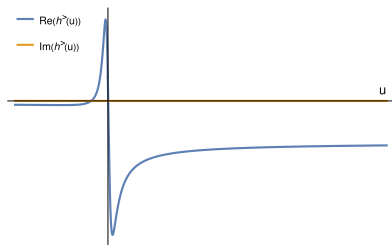
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- Need to study tail effects** appearing at higher orders! [Cho,Kälin,Porto]

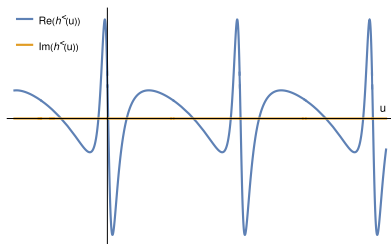
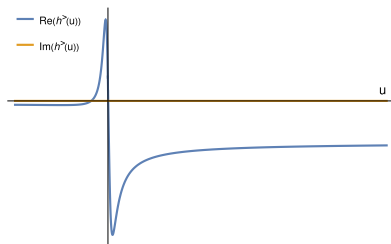
From scattering to bound observables (III)

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- At 1PN [Junker, Schäfer] this corresponds to a **map between PN multipoles** [Adamo, RG, Ilderton]

$$h^>(u; p_\infty, L) = \frac{4G_N}{c^2} \left(U_2^> + \frac{1}{c}(V_2^> + U_3^>) + \frac{1}{c^2}(V_3^> + U_4^>) \right),$$

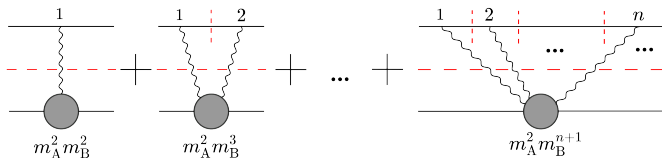
$$U_j^<(u; \tilde{p}_\infty) = U_j^>(u; +i\tilde{p}_\infty) \text{ for } j \leq 4, \quad V_k^<(u; \tilde{p}_\infty) = V_k^>(u; +i\tilde{p}_\infty) \text{ for } k \leq 3.$$

Need for resummation: from amplitude to self-force

- Need for **resummation of perturbative contributions**! Can we look at simpler examples? [Del Duca, RG, Sasank]
- Consider the **geodesic equation**,

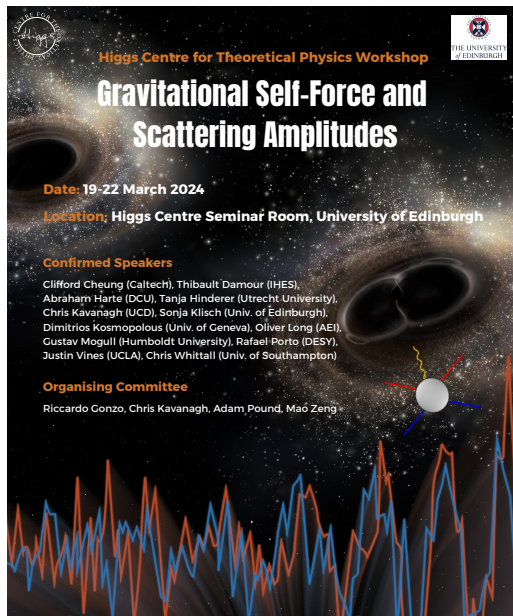
$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\alpha\beta}^\mu \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0,$$

corresponding to the **leading order in the mass ratio expansion**: this requires to **resum an infinite set of diagrams** [Cheung, Solon, Shah; Brandhuber, Chen, Travaglini, Wen; Bjerrum-Bohr, Planté, Vanhove]



Interestingly, the **scattering-to-bound map works for Kerr geodesics**! [Shi, RG]

- **Open problem**: What can we learn by studying the **self-force expansion**? [Kosmopolous, Solon; Cheung, Parra-Martinez, Rothstein, Shah, Wilson-Gerow]



Higgs Centre for Theoretical Physics Workshop

Gravitational Self-Force and Scattering Amplitudes

Date: 19-22 March 2024

Location: Higgs Centre Seminar Room, University of Edinburgh

Confirmed Speakers

Clifford Cheung (Caltech), Thibault Damour (IHES), Abraham Harte (DCU), Tanja Hinderer (Utrecht University), Chris Kavanagh (UCD), Sonja Klisch (Univ. of Edinburgh), Dimitrios Kosmopolous (Univ. of Geneva), Oliver Long (AEI), Gustav Mogull (Humboldt University), Rafael Porto (DESY), Justin Vines (UCLA), Chris Whittall (Univ. of Southampton)

Organising Committee

Riccardo Gonzo, Chris Kavanagh, Adam Pound, Mao Zeng

Summary and future directions

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- **Future directions**: extend the scattering-to-bound map to include tail effects for the waveform and other observables, explore the resummation of PM contributions, extend the analytical continuation to generic spin alignments, ...

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